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Search for three- α states around an ¹⁶O core in ²⁸Si

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(Received 10 June 2012; revised manuscript received 3 August 2012; published 10 September 2012)

We investigate the existence of weakly coupled gaslike states comprised of three α particles around an ¹⁶O core in ²⁸Si. We calculate the excited states in ²⁸Si using the multiconfiguration mixing method based on the ¹⁶O + 3 α cluster model. We also include the ¹⁶O + ¹²C and ²⁴Mg + α basis wave functions prepared by the generator coordinate method. To identify the gaslike states, we calculate the isoscalar monopole transition strengths and the overlap of the obtained states with the geometrical cluster wave function and the Tohsaki-Horiuchi-Schuck-Röpke (THSR) wave function. The results show that the obtained fourth and twelfth states significantly overlap with the THSR wave function. These two states clearly coexist with the ¹⁶O + ¹²C cluster states, emerging at similar energies. The calculated isoscalar monopole strengths between those two states are significantly large, indicating that the states are members of the excitation mode. Furthermore, the calculated root-mean-square radii for these states also suggest that a layer of gaslike three α particles could exist around the surface of the ¹⁶O core, which can be described as a "two-dimensional gas" in the intermediate state before the Hoyle-like three- α states emerge.

DOI: 10.1103/PhysRevC.86.031303

PACS number(s): 21.10.-k, 21.60.Gx, 27.20.+n

The investigations of excited states in light-mass nuclei provide a good opportunity to study the rich variety of nuclear structures of quantum many-body systems. One of these states is the well known Hoyle state in ¹²C [1]. The Hoyle state is the second 0⁺ state in ¹²C, with an excitation energy of 7.65 MeV, just above the three- α decay threshold energy. However, it is difficult to reproduce its energy and geometrical properties by calculations based only on the shell model. On the other hand, the microscopic three- α cluster models successfully describe the properties of the Hoyle state (such as the observed α -decay width) and indicate that the three- α state develops well in the 0^+_2 state [2,3].

Recent calculations using the α cluster model suggest that in the Hoyle state, the three "gaslike" α particles are weakly coupled with each other near the three- α -decay threshold energy [4,5]. Based on this picture, the Tohsaki-Horiuchi-Schuck-Röpke (THSR) wave function was proposed. Using this wave function, Tohsaki et al. proposed the extremely large rms nuclear radius for the 0_2^+ state in ¹²C [4]. Those studies have triggered interest in whether such gaslike states exist in heavier-mass nuclei [6,7]. In this connection, the existence of weakly coupled gaslike α cluster states around a *core* has been recently suggested [8–12]. To study this, the Monte-Carlo technique for the THSR wave function was proposed [13]. The possibility of a gaslike three- α cluster state around ⁴⁰Ca in ⁵²Fe has been discussed using this technique [14]. To identify such gaslike states experimentally, the measurement of an isoscalar monopole transition strength has been proposed [15,16]. The enhancement of the transition strength would correspond to the development of such a cluster structure. In this respect, gaslike two- α states around an ¹⁶O core in ²⁴Mg have been investigated both theoretically [15,17–19] and experimentally [16,20,21].

The aim of the present paper is to search for the weakly coupled gaslike three- α cluster states around an ¹⁶O core in ²⁸Si. In analogy to the three- α -particle Hoyle state in ¹²C, we

can expect the existence of states for three α particles, similar to the Hoyle state, around an ¹⁶O core. However, in ²⁸Si, many strongly coupled cluster states would emerge at and below/above the (¹⁶O + 3 α) decay threshold energy [22–24]. Three α particles around an ¹⁶O core would form the strongly coupled ¹²C state, to gain the intercluster potential energy. An important feature is that these weakly and strongly coupled systems would coexist and that their emergences compete with each other. To discuss the existence of the weakly coupled gaslike three- α state around an ¹⁶O core, it is important to completely understand the occurrence of all strongly coupled cluster states within the orthogonality condition among the states.

To investigate the existence of the gaslike three- α state around the ¹⁶O core, we calculate the excited states of ²⁸Si using the ¹⁶O + 3 α cluster model. We superpose many randomly generated Slater determinants, which correspond to the multiconfiguration mixing (MCM) method. To describe the molecular states, we also include the basis wave functions with ¹⁶O + ¹²C and ²⁴Mg + α configurations prepared by the generator coordinate method (GCM). To identify the relevant states, we calculate the overlap of the obtained results with the THSR and the geometrical cluster wave functions. We also calculate the rms radius and the isoscalar monopole transition strength. We show below how the gaslike three- α states emerge and how they connect with the ground state.

In the present study, we use Brink's α cluster model [25]. In this approach the wave function for ²⁸Si is described by the Slater determinant of the seven α clusters. The wave function of the α cluster, ϕ_{α} , is described by the direct product of four wave functions for each nucleon (proton and neutron with spin down and up). The spacial part of the wave function for the α cluster is given by $\phi_{\alpha}(\vec{R}) \propto \prod_{i=1}^{4} \exp[-\nu(\vec{r}_i - \vec{R})^2]$, where \vec{R} denotes the center position vector of the Gaussian function, \vec{r}_i denotes the spatial coordinate of each nucleon, and ν denotes



FIG. 1. (Color online) Density plots of the basis wave functions for the ¹⁶O and ²⁴Mg cores. The contours correspond to multiple steps of 0.2 fm⁻². The color is normalized by the largest density in each plot.

the size parameter. We take $v = 1/2b^2$ with b = 1.46 fm in the calculations. To describe the α cluster, we take the same \vec{R} for all four nucleons.

To calculate the three- α states around the ¹⁶O core in ²⁸Si, we superimpose many basis wave functions for ¹⁶O + 3 α cluster configurations. To take into account the molecular states, we also prepare the basis wave functions using ¹⁶O + ¹²C and ²⁴Mg + α configurations based on the GCM.

We first prepare the cluster wave functions for the ¹⁶O, ¹²C, and ²⁴Mg configurations. The cluster wave functions of ¹⁶O, ¹²C, and ²⁴Mg with the center-of-mass position \vec{R} are denoted by $\Phi_{^{16}O}(\vec{R})$, $\Phi_{^{12}C}(\vec{R})$, and $\Phi_{^{24}Mg}(\vec{R})$, respectively. For $\Phi_{^{16}O}(\vec{R})$, we take a tetrahedron configuration with four α clusters. Figure 1(a) shows the total density of $\Phi_{^{16}O}$ used in this study. The α - α distance between each center position of the Gaussian functions of the four α 's in the ¹⁶O core is 0.5 fm.

For $\Phi_{^{12}C}(R)$, we take the equilateral triangle configuration of the three α clusters with a length of a fm. The triangle is placed on the x-z plane and a side (two of the α clusters) is perpendicular to the z axis in the center-of-mass flame of ^{12}C . The residual α cluster is placed on the negative z axis. We also take three different configurations by rotating the α cluster triangle on the x axis with angles of 0°, 45°, and 90°. Those configurations are denoted by $^{12}C_{\parallel}$, $^{12}C_{\perp}$, and $^{12}C_{\perp}$, respectively.

The wave function of $\Phi_{^{24}Mg}(\vec{R})$ is an isosceles configuration with $\Phi_{^{16}O}(\vec{R}_0)$ and two α clusters. Figure 1(b) shows the density plot of $\Phi_{^{24}Mg}$ used in this study. The configuration of $\Phi_{^{16}O}(\vec{R}_0)$ is the same as mentioned above. The crosses denote the positions of the two α clusters. The distance between the two α clusters is denoted by $l_{2\alpha}$. The distance between the two α clusters is denoted by $l_{2\alpha}$. The distance between $\Phi_{^{16}O}(\vec{R}_0)$ and the *z* position of the base is denoted by $z_{2\alpha}$. We determine $l_{2\alpha}$ and $z_{2\alpha}$ by minimizing the expectation value of the total energy for $\Phi_{^{24}Mg}$. We here take $l_{2\alpha} = 2.5$ fm and $z_{2\alpha} = 3.0$ fm.

We next prepare the basis wave functions using ${}^{16}\text{O} + 3\alpha$, ${}^{16}\text{O} + {}^{12}\text{C}$, and ${}^{24}\text{Mg} + \alpha$ cluster configurations. The basis wave function using the ${}^{16}\text{O} + 3\alpha$ cluster configuration, $\Psi^{(i)}_{{}^{16}\text{O}+3\alpha}$, is given by $\Psi^{(i)}_{{}^{16}\text{O}+3\alpha} = [\mathcal{A}\Phi_{{}^{16}\text{O}}(\vec{R}_0)\phi_{\alpha}(\vec{R}_1)\phi_{\alpha}(\vec{R}_2)\phi_{\alpha}(\vec{R}_3)]_i$, where \mathcal{A} denotes the

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antisymmetrization operator. We first take $\vec{R}_0 = 0$ and the positions of the three- α clusters \vec{R}_1 , \vec{R}_2 , and \vec{R}_3 are randomly generated under the condition of the distributed probability $w(\vec{R}_i)$. In this study, we use $w(\vec{R}_i) \propto \exp[-\vec{R}_i^2/\sigma^2]$, where σ is the size parameter. Subsequently, the center-of-mass correction for $\Psi_{16O+3\alpha}^{(i)}$ was performed.

The basis wave functions using the ¹⁶O + ¹²C configuration, $\Psi_{^{16}O^{+12}C}^{(i)}$, are given by $\Psi_{^{16}O^{+12}C}^{(i)} = [\mathcal{A}\Phi_{^{16}O}(\vec{R}_0)\Phi_{^{12}C}(\vec{R})]_i$. The ¹²C cluster is shifted to the positive *z* direction. Here, the distance between the center-of-mass positions of the ¹⁶O and ¹²C clusters is denoted by $d_{^{16}O^{-12}C}$. The basis wave functions using ²⁴Mg + α configurations, $\Psi_{^{24}Mg+\alpha}^{(i)}$, are given by $\Psi_{^{24}Mg+\alpha}^{(i)} = [\mathcal{A}\Phi_{^{24}Mg}(\vec{R}_0)\phi_{\alpha}(\vec{R})]_i$. The wave function of $\phi_{\alpha}(\vec{R})$ is placed on the *x*, *y*, or *z* axis. Those configurations are denoted by ²⁴Mg + α_x , ²⁴Mg + α_y , and ²⁴Mg + α_z . The distance between the center-of-mass positions of the ²⁴Mg and α clusters is denoted by $d_{^{24}Mg-\alpha}$.

In the calculation, we generate 1000 wave functions for $\Psi_{^{16}O+3\alpha}^{(i)}$. For $\Psi_{^{16}O+^{12}C}^{(i)}$, we take four and eight values for *a* and $d_{^{16}O-^{12}C}$ in each angle of $^{12}C_{\parallel}$, $^{12}C_{\perp}$, and $^{12}C_{\perp}$, respectively. Those correspond to a = (1.0, 2.0, 3.0, 4.0) fm and $d_{^{16}O-^{12}C} = (1.0, 2.0, \ldots, 8.0)$ fm. For $\Psi_{^{24}Mg+\alpha}^{(i)}$, we take eight values for $d_{^{24}Mg-\alpha}$ corresponding to $d_{^{24}Mg-\alpha} = (1.0, 2.0, \ldots, 8.0)$ fm in each direction of α_x , α_y , and α_z .

Thus, the total number of the basis wave functions is 1120. The total wave function Ψ_{Total} is described with the linear combination of the basis wave functions given by $\Psi_{\text{Total}} =$ $\sum_{i} c_i P^+ P_{00}^0 \Psi^{(i)}$, where $\Psi^{(i)}$ denotes each basis wave function. The coefficient c_i is determined by diagonalizing the total Hamiltonian. The symbols P^+ and P_{00}^0 denote the projection operator for the parity and angular momentum (J = M =K = 0) to the 0⁺ state, respectively. For the calculations of the angular-momentum projection, $24 \times 32 \times 24$ grids points are taken with respect to the α , β , and γ directions of the Euler angle. For the total Hamiltonian, we use the Volkov No. 2 effective N-N interaction [26]. In the calculation, the Majorana exchange parameter M is chosen to be 0.645 so as to reproduce the binding energy of ²⁸Si. In the calculation, the cluster decay threshold energies for ${}^{16}O + 3\alpha$, ${}^{16}O + {}^{12}C$, 20 Ne + 2 α , and 24 Mg + α are -189.29, -186.48, -192.29, and -197.23 MeV, respectively. We also performed a calculation with a stronger interaction (Majorana parameter M = 0.62), which gives a bound state of ${}^{12}C$, but the qualitative behavior of the calculated results is not changed.

To identify the gaslike α states, we calculate the overlap between the obtained states and the THSR wave function. The THSR wave function with the gaslike three α particles around an ¹⁶O core, $\Psi_{\text{THSR}}^{(\sigma)}$, is given by

$$\Psi_{\text{THSR}}^{(\sigma)} = \int d\vec{R}_1 d\vec{R}_2 d\vec{R}_3 \cdot \exp\left[-\left(\vec{R}_1^2 + \vec{R}_2^2 + \vec{R}_3^2\right)/\sigma^2\right] \\ \cdot \left[\mathcal{A}\phi_{^{16}\text{O}}(\vec{R}_0)\phi_{\alpha}(\vec{R}_1)\phi_{\alpha}(\vec{R}_2)\phi_{\alpha}(\vec{R}_3)\right] \\ = \mathcal{A}\phi_{^{16}\text{O}}(\vec{R}_0)\prod_{i=1}^3 \int d\vec{R}_i \ \phi_{\alpha}(\vec{R}_i)\exp\left[-\vec{R}_i^2/\sigma^2\right], \quad (1)$$

where σ is the size parameter of the gaslike α state. We can perform the Monte-Carlo integration for Eq. (1) (see details in [13]). Then, the Monte-Carlo THSR wave function $\Psi_{\text{THSR}}^{(\sigma)}$ is given by $\Psi_{\text{THSR}}^{(\sigma)} = \sum_{i} P^+ P_{00}^0 \Psi_{16O+3\alpha}^{(i)}$. The integral points for \vec{R}_1, \vec{R}_2 , and \vec{R}_3 are randomly generated under the condition of the weight factor $w(\vec{R}_i)$ given by $w(\vec{R}_i) \propto \exp[-\vec{R}_i^2/\sigma^2]$. In this study, we calculate the THSR wave function with $\sigma =$ 3, 4, and 5 fm. For the Monte-Carlo calculations, we use 900 randomly generated basis wave functions for each σ . The expectation values of the total energy for those with $\sigma = 3$, 4, and 5 converge well at energies of -191.39, -188.48, and -182.33 MeV, respectively.

To investigate the properties of the states obtained in this procedure, we also calculate the isoscalar monopoletransition strength and the rms radius. Here, the unit of the transition strength $B_0(ISO)$ is taken as the value obtained from the 1s to the 2s states, described by the single-particle wave function of the three-dimensional harmonic oscillator. The value of $B_0(ISO)$ is then given by $B_0(ISO) = \sqrt{5}b^2$ fm² [15]. In the present study, we take $B_0(ISO) = 4.77$ fm² with b = 1.46.

Figure 2(a) shows the convergence behavior of the states measured from the ${}^{16}\text{O} + 3\alpha$ decay threshold energy versus the number of the basis wave functions. We plot here only the results below the 15th state. In Fig. 2(a), the gray area (i) denotes the results calculated with the subspace for the ${}^{16}\text{O} + {}^{12}\text{C}_{\perp}$, ${}^{12}\text{C}_{\parallel}$, and ${}^{12}\text{C}_{\perp}$ configurations. The gray area (ii) also denotes those for the ${}^{24}\text{Mg} + \alpha_x$, α_y , and α_z configurations in addition to the ${}^{16}\text{O} + {}^{12}\text{C}$ components. The lowest energy in the calculated states with the total wave

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functions is -24.64 MeV, which is in good agreement with the experimental value of -24.03 MeV.

We first investigate the structure of the ground (first) state. This state corresponds to the prolate shape with a significant ${}^{16}\text{O} + {}^{12}\text{C}_{\perp}$ cluster component, which originates from the lowest state of the results of the gray area (i) in Fig. 2(a). The rms radius obtained for this state is 2.80 fm. To clarify the cluster structure, we calculate the overlap of the obtained states with the wave functions using ${}^{16}\text{O} + {}^{12}\text{C}$ and ${}^{24}\text{Mg} + \alpha$ configuration. Figure 3 shows the density plots of the wave functions maximally overlapping with the calculated states. The lowest state overlaps maximally with the ${}^{16}\text{O} + {}^{12}\text{C}_{\perp}$ configuration of 77.8% at a = 2.0 fm and $d_{16\text{O}-12\text{C}} = 3.0$ fm [see Fig. 3(a)].

The second state corresponds to the one-body oblate shape. The obtained rms radius is 2.77 fm, which is somewhat smaller than that of the first state. The second state overlaps maximally with the ${}^{16}O + {}^{12}C_{\parallel}$ configuration of 17.3% at a = 3.0 fm and $d_{{}^{16}O^{-12}C} = 1.0$ fm, but many other components also have similar overlaps at small $d_{{}^{16}O^{-12}C}$. Thus, this state is the compact one-body system rather than a specific cluster state. In Brink's α cluster model without the spin-orbit force, the shape configurations for the first and second 0⁺ states are prolate and oblate (pentagon), respectively. This shows that the results of our calculations are consistent with previous studies, although the experimental data suggest that the ground and the third 0⁺ states correspond to the oblate and the prolate shapes, respectively [27].

We next investigate the structure of the excited states. Figures 2(b) and 2(c) show the calculated isoscalar monopole-transition strength from the first and second states in the unit



FIG. 2. (Color online) (a) Convergence behavior of the calculated energies versus the number of basis wave functions measured from the ${}^{16}O + 3\alpha$ decay threshold energy. The dashed lines denote the decay threshold energies for ${}^{16}O + 3\alpha$, ${}^{20}Ne + 2\alpha$, and ${}^{24}Mg + \alpha$. In the gray areas denoted by (i) and (ii), we use the basis wave functions with ${}^{16}O + {}^{12}C$ and ${}^{24}Mg + \alpha$ configurations, respectively. (b) and (c) Isoscalar monopole-transition strength from the first and second states in the unit of the single-particle (S.P.) excitation from the 1*s* to the 2*s* states, respectively. (d), (e), and (f) Overlap between each state and the THSR wave function with $\sigma = 3$, 4, and 5 fm, respectively.



FIG. 3. (Color online) Density plots of the basis wave function overlapping maximally with the calculated states. The contours correspond to multiples of 0.2 fm^{-2} . The color is normalized by the largest density in each plot.

of $B_0(IS0)$, respectively. Figures 2(d)-2(f) show the calculated overlaps between the obtained states and the THSR wave function with $\sigma = 3$, 4, and 5 fm, respectively. Below, we only discuss the characteristic results which can be identified by the analysis.

In Figs. 2(d)–2(f), we can see that the fourth and twelfth states significantly overlap with the THSR wave function. These states are candidates for the gaslike three- α state discussed in this study. The overlap between the fourth state and the THSR wave functions with $\sigma = 3$, 4, and 5 fm are 16.7%, 9.8%, and 3.3%, respectively. Those overlaps for the twelfth state are 20.0%, 13.7%, and 4.9%. The rms radii obtained for the fourth and the twelfth states are 2.92 and 3.02 fm, respectively. These values are somewhat spatially extended rather than that of the first and second states.

In addition, we obtain a remarkably strong transition strength from the fourth to twelfth states by $B(IS0)/B_0(IS0) = 3.20$, indicating that those states strongly correlate with each other due to the cluster excitation. We thus consider that the fourth and twelfth states are a member of the new excitation mode. We also obtain a relatively large transition strength from the first to the fourth states by $B(IS0)/B_0(IS0) = 2.21$ [see Fig. 3(b)], suggesting that this new excitation mode is built on the first state with the prolate shape.

The sixth state has the strongest transition strength from the ground state by $B(IS0)/B_0(IS0) = 3.67$. This state overlaps maximally with the ${}^{16}O + {}^{12}C_{\perp}$ configuration of 42.2% at a = 2.0 fm and $d_{16O^{-12}C} = 4.0$ fm [see Fig. 3(b)]. The obtained rms radius is 2.95 fm. This corresponds to the ${}^{16}O + {}^{12}C$ molecular state where the relative motion between two clusters is excited from the ground (first) to the higher-nodal states. The seventh state also has strong transition strength from the second state by $B(IS0)/B_0(IS0) = 3.30$. This state is the cluster excitation of the ${}^{24}Mg + \alpha_x$ configuration. The maximum overlap between the seventh state and the ${}^{24}Mg + \alpha_x$ configuration is 23.2% at $d_{{}^{24}Mg^{-}\alpha} = 5.0$ fm [see Fig. 3(c)]. The obtained rms radius is 2.93 fm.

An important observation is that the fourth and the twelfth states do not largely overlap with any geometrical configurations of the α clusters. Despite this, the rms radii for these states are considerably extended to similar extents as other cluster states. The fourth and twelfth excited states

significantly overlap with the THSR wave function not only with $\sigma = 3$ but also $\sigma = 4$ fm. This suggests the existence of gaslike three α particles spreading to the outside of the ¹⁶O core. However, the calculated rms radii for the fourth and twelfth states are not so extended (2.92 and 3.02 fm, respectively). That is, the three α particles still exist around the ¹⁶O core.

Therefore, we consider that the fourth and twelfth states are members of a new type of excitation mode built on the prolate state, which has not yet been well established in experiments. In the first excited state of this mode (fourth state), the weakly coupled gaslike three α particles, without the angular correlations, emerge around the surface of the ¹⁶O core. That would be regarded as a "two-dimensional (2D) gas." This is because the rms radii for those three α particles are not so widely spread and those states significantly overlap mainly with the THSR wave function of $\sigma = 3$ fm. On the other hand, the Hoyle state in ¹²C overlaps with the THSR wave function with larger σ values (3–4 fm). That is, the radial component of the wave functions for those states are confined around the surface of the ¹⁶O core.

The emergence of this weakly coupled 2D gaslike state significantly competes with that of the ${}^{16}O + {}^{12}C$ molecular state (sixth state), which has strongly coupled three α 's of ¹²C, as their energies are almost comparable to each other. This indicates that the energy necessary for the excitation from the strongly coupled ¹²C to the weakly coupled three α 's around the ¹⁶O core is comparable to the relative motion between ¹⁶O and ¹²C. Despite this competition, the coexistence of the 2D gaslike and the strongly coupled states is highly possible, as shown in this study. In the second excited state of this new mode (twelfth state), the gaslike layer is well developed and more spatially expanded, because the isoscalar monopole strength only from the first excited (fourth) state is extremely large. However, this state is also a 2D gaslike one, because the spatial component of this state still distributes around the ¹⁶O core.

In more highly excited states, the Hoyle-like weakly coupled 3α state, that is a "three-dimensional (3D) gaslike" state around the ¹⁶O core, may emerge. Then, the second excited 2D gaslike state would be the intermediate state before the 3D gaslike state emerges and would be directly connected

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to the 3D gaslike state by the large isoscalar monopole transition strength. However, it is difficult to identify such 3D gaslike states in the present calculations, because many continuum states are coupled with other states in the high excitation energies. The development of analyses, such as the complex scaling and the pseudopotential methods, is necessary when searching for such states.

In summary, we have investigated the existence of weakly coupled gaslike three- α states around an ¹⁶O core. To study this, we have calculated the excited states of ²⁸Si using the multiconfiguration mixing method with the basis wave functions randomly generated by the ¹⁶O + 3 α cluster model. We have also included the ¹⁶O + ¹²C and ²⁴Mg + α basis wave functions, prepared by the GCM method to describe well the molecular states. To identify the states, which we have obtained, we have calculated their overlap with the geometrical cluster wave functions and the THSR wave function. Furthermore we have also calculated the rms radii and the isoscalar monopole transition strengths for the obtained states.

We have found that the fourth and twelfth excited states largely overlap with the THSR wave function with $\sigma = 3$ and

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4 fm. The calculated isoscalar monopole strengths are also significantly large from the first to the fourth and from the fourth to the twelfth states, indicating those may be members of a new excitation mode. At around the energy of the fourth excited state, the ¹⁶O + ¹²C cluster state, which has three strongly coupled α 's of ¹²C, also emerges, but the fourth state coexists with this state. The calculated rms radii for the fourth and twelfth states suggest that a layer of dilute three α particles may exist around the ¹⁶O core. This gaslike structure in the layer would be the intermediate state, before the complete 3D Hoyle-like gas state emerges.

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A part of this research has been funded by the MEXT HPCI Strategic Program. Tz.K. is grateful to the Daphne Jackson Trust and the STFC for their support. Numerical computation in this work has been carried out at the Yukawa Institute Computer Facility using the SR16000 system. This work was performed as part of the Yukawa International Project for Quark-Hadron Sciences (YIPQS). It was also supported by a Grant-in-Aid for the Global COE Program "The Next Generation of Physics, Spun from Universality and Emergence."

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