

Effects of initial state fluctuations in the final state elliptic flow measurements using the NeXSPheRIO model

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We present a systematic study of the effects from initial condition fluctuations in systems formed by heavy-ion collisions using the hydrodynamical simulation code NeXSPheRIO. The study was based on a sample of events generated simulating Au + Au collisions at center-of-mass energy of 200 GeV per nucleon pair with an impact parameter ranging from most central to peripheral collisions. The capability of the NeXSPheRIO code to control and save the initial condition (IC) as well as the final state particles after the three-dimensional hydrodynamical evolution allows for the investigation of the sensitivity of the experimental observables to the characteristics of the early IC. Comparisons of results from simulated events generated using fluctuating initial conditions and a smooth initial condition are presented for the experimental observable elliptic flow parameter (v_2) as a function of the transverse momentum p_t and centrality. We compare v_2 values estimated using different methods, and how each method responds to the effects of fluctuations in the initial condition. Finally, we quantify the flow fluctuations and compare them to the fluctuations of the initial eccentricity of the energy density distribution in the transverse plane.

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I. INTRODUCTION

In many dynamical systems, the evolution of the system affects its physical characteristics and the information of the early stages is lost when observing only the probes from the final stages. However, by studying how much of the initial condition information is retained and survives the dynamical evolution it is possible to obtain valuable insight not only on the characteristics of the initial conditions but also on the interactions that occur in the evolution. These conditions can be applied in many physical systems such as in the study of the evolution of our universe and the anisotropy observed in the cosmic microwave background [1], or in the study of the nuclear reactions in high energy [2].

In relativistic heavy-ion collisions an important experimental observable is known as the elliptic flow. Because of geometrical anisotropy created in the initial condition by the two colliding nuclei and the evolution of the system, a space-momentum correlation develops and an anisotropy in the azimuthal distribution of the final particles can be measured. The elliptic flow or the Fourier expansion second harmonic (v_2) corresponds to the amplitude of this azimuthal anisotropy. Hydrodynamic models predict that indeed elliptic flow should be sensitive to the eccentricity of the initial conditions [3]. Model calculations of v_2 are also in good agreement with much of the experimental data [4–12]. Furthermore, two or more particle correlations may carry important information

on the initial state, hence on the mechanism of quantum chromodynamics at the very early stage of the collision. Recent studies [13–21] show that a scenario including fluctuations of the initial condition in addition to the geometrical eccentricity describes better the experimental data than without the inclusion of initial state fluctuations. These fluctuations in the initial condition will affect the eccentricity by changing the average value and also by introducing an additional fluctuation which will then affect the final observed v_2 values. To explore the effects of these fluctuations, we need a more precise study, quantifying the effects of these initial condition fluctuations to the final experimental observables. For this purpose we have used the simulation code NeXSPheRIO [22] that allows the generation of events using smooth or fluctuating initial conditions (ICs) and obtain the final particles after the hydrodynamical evolution. The procedures used in experimental data analysis for v_2 determination are applied to the simulated data and the effects of the initial fluctuations are investigated through the differences in the v_2 estimates using different methods. NeXSPheRIO allows for the control and study of both the initial condition before the hydrodynamical evolution and the final particles after hadronization and freeze-out. Thus, it allows for a detailed study, on an event-by-event analysis, of the correlation between initial condition parameters such as eccentricity and fluctuations to the final state observables such as v_2 . This simulation code has already been extensively tested and has presented reasonable agreement with experimental data [15,23,24]. In the next section we describe how the initial energy density is obtained in the NeXSPheRIO code and how the eccentricity is calculated. In Sec. III we describe

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the methods used to estimate the elliptic flow. In Sec. IV we present our results and, finally, in Sec. V we summarize our conclusions.

II. INITIAL ENERGY DENSITY

The NeXSPheRIO code allows the use of smooth and fluctuating initial conditions, which is particularly suited and convenient to study the effects of fluctuations in the elliptic flow calculations. The initial conditions are generated by a microscopic model called NeXuS [22] which produces, on an event-by-event basis, detailed space distributions of the energy-momentum tensor, baryon-number, strangeness, and charge densities [22]. The initial condition is then used as input to the hydrodynamical model and the system evolution is computed up to a given point where a decoupling mechanism is applied and final particles are produced [22]. Considering the transverse profile of the initial energy density distribution around $\eta = 0$, it is possible to calculate the eccentricity of the initial geometry. In the situation where there is no fluctuation and the colliding nuclei are considered to have smooth distributions, the major axis of the almond shape of the overlap area at the moment of the collision is perpendicular to the plane defined by the impact parameter and the beam axis, generally referred to as the reaction plane, while the minor axis coincides with the direction of the impact parameter.

The eccentricity calculated with respect to this reaction plane, defined by ε_{RP} , is given by [25]

$$\varepsilon_{\text{RP}} = \frac{\sigma_y^2 - \sigma_x^2}{\sigma_y^2 + \sigma_x^2}, \quad (1)$$

where σ_x^2 and σ_y^2 are the variances of the distribution along the x and y directions, respectively. Because of fluctuations, the actual distribution of the hot material created by collisions of participants can have the principal axis different from those of the smooth initial condition and the minor axis deviates from the reaction plane direction defined above [25]. The plane that maximizes the eccentricity determined from the initial energy distribution is generally called the participant plane [26]. The participant plane eccentricity ε_{PP} is given by [25]

$$\varepsilon_{\text{PP}} = \frac{\sqrt{(\sigma_y^2 - \sigma_x^2)^2 + 4\sigma_{xy}^2}}{\sigma_x^2 + \sigma_y^2}, \quad (2)$$

where σ_{xy} is the covariance of the transverse energy density distribution. Note that the reaction plane is determined just by the initial geometry of the collision for the whole system uniquely, whereas the participant plane depends on the dynamics of the collisions in each event. In Fig. 1, we show the distributions of the reaction plane eccentricity and the participant plane eccentricity as a function of the collision impact parameter. In the x axis of Fig. 1, we also show the equivalent number of charged particle density produced at $\eta = 0$, $dN_{\text{ch}}/d\eta$, which is obtained after the full event simulation procedure is completed. The top panel shows the event-by-event reaction plane eccentricity ε_{RP} and the bottom panel shows the participant plane eccentricity ε_{PP} . By construction, the smooth IC reaction plane eccentricity

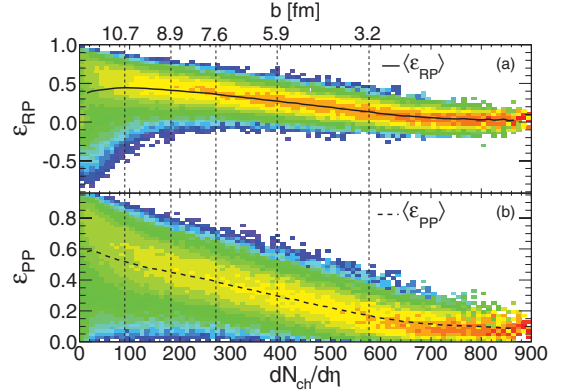


FIG. 1. (Color online) (a) Reaction plane eccentricity ε_{RP} and (b) participant plane eccentricity ε_{PP} as a function of the number of charged particles produced at $\eta = 0$. The black line (a) represents the mean of the ε_{RP} distribution, which is equal to the ε_{RP} and ε_{PP} of the smooth IC case. The dashed line (b) represents the average of the ε_{PP} for the fluctuating IC case.

corresponds to the mean eccentricity of the fluctuating IC case, represented by the solid black line in Fig. 1(a). In addition, the participant plane eccentricity matches the reaction plane eccentricity for the smooth IC case. For the most peripheral collisions the reaction plane eccentricity can assume negative values because of fluctuation that will force the asymmetry plane to be perpendicular to the reaction plane. The dashed line (b) represents the mean of the participant plane eccentricity for the fluctuating IC case only.

III. ELLIPTIC FLOW DETERMINATION

The momentum anisotropy of the final particle azimuthal distributions is generally described by a Fourier expansion of the azimuthal angle ϕ of each particle with respect to the reaction plane angle Ψ_{RP} [27]. The most relevant contribution comes from the second coefficient of the expansion v_2 , which is from the almond shape of the initial overlap area of the incident nuclei in the transverse plane. However, the reaction plane angle cannot be directly measured experimentally and, therefore, several methods and techniques have been developed to estimate the v_2 coefficient from the experimental data [27–30]. For this work, we have computed the elliptic flow coefficients in several different ways: (a) using the true reaction plane ($v_2\{\text{RP}\}$) known in the simulation; (b) estimating a reference plane with the event plane method ($v_2\{\text{EP}\}$) [27,28]; (c) calculating the participant plane from the initial energy density distribution profile around $\eta = 0$; and (d) using the formalism of the cumulant method ($v_2\{2\}$ and $v_2\{4\}$) [30]. Below we summarize how each method works.

The reaction plane elliptic flow, $v_2\{\text{RP}\}$, is calculated with respect to the true reaction plane as

$$v_n\{\text{RP}\} = \langle \cos[n(\phi - \Psi_{\text{RP}})] \rangle, \quad (3)$$

where ϕ is the azimuthal angle of each particle, Ψ_{RP} is the true reaction plane angle, and the angle brackets denotes averages first over all selected particles in each event, and then over all events. In NeXSPheRIO, for fluctuating IC, an event is first

generated by generating nucleons inside the colliding nuclei. Thus, before the collision, the colliding matter already loses its symmetry with respect to the reaction plane, as defined by the impact parameter vector and the beam axis. In this sense, the concept of the reaction plane loses its physical meaning for event-by-event calculations. However, we keep the calculations of $v_2\{\text{RP}\}$ as a reference for comparisons with the flow results obtained with other methods.

The event plane method is used to estimate the true reaction plane using the flow anisotropy itself. It is calculated by defining the flow vector as follows [28]:

$$\Psi_n = \frac{\tan^{-1}(Y_n/X_n)}{n}, \quad \begin{cases} X_n = \sum_i w_i \cos(n\phi_i), \\ Y_n = \sum_i w_i \sin(n\phi_i), \end{cases} \quad (4)$$

where w_i is a weight and ϕ_i is the azimuthal angle of particle i . The flow coefficients are then determined by [28]

$$v_n\{\text{EP}\} = \frac{\langle \cos[n(\phi - \Psi_n)] \rangle}{\langle \cos[n(\Psi_n - \Psi_{RP})] \rangle}, \quad (5)$$

where the denominator in the right-hand side of Eq. (5) is the resolution of the event plane, and the angle brackets denote averages taken over all selected particles in each event, and then over all events. Experimentally, Ψ_{RP} is not known and thus the resolution must be inferred in some other way. The technique applied here and extensively used in experimental data analysis consists of dividing the event being analyzed into two subevents (event a and event b) of the same multiplicity and calculate the resolution of the subevent. It was shown that the resolution of the subevents is given by [27]

$$\begin{aligned} \langle \cos [n(\Psi_n^a - \Psi_{RP})] \rangle &= \langle \cos [n(\Psi_n^b - \Psi_{RP})] \rangle \\ &= \sqrt{\langle \cos [n(\Psi_n^a - \Psi_n^b)] \rangle}, \end{aligned} \quad (6)$$

where Ψ_n^a and Ψ_n^b are the event plane angles for the subevents a and b , respectively. From the resolution of the subevents one is able to estimate the resolution of the event plane for the full event (see details in Refs. [27,28]). The event plane angle can deviate from the true reaction plane angle because of the fluctuations in the initial conditions and also because of the resolution caused by the limited statistics of the measured particles. Neglecting the angle shift from the statistical resolution, it is expected that Ψ_{EP} would follow Ψ_{PP} , the participant plane azimuthal angle. So, to check this we also calculate v_2 with respect to the participant plane angle:

$$v_n\{\text{PP}\} = \langle \cos(n[\phi - \Psi_{PP}]) \rangle, \quad (7)$$

where Ψ_{PP} is calculated from the initial energy density transverse distribution profile as [25]

$$\Psi_{PP} = \tan^{-1} \left(\frac{\pm \sigma_{xy}}{\sigma_y^2 - \lambda^\mp} \right), \quad (8)$$

$$\lambda^\pm = \frac{1}{2} (\sigma_y^2 + \sigma_x^2 \pm \sqrt{(\sigma_y^2 - \sigma_x^2)^2 + 4\sigma_{xy}}). \quad (9)$$

A recent study [31] showed that the event plane angle Ψ_{EP} as estimated from the azimuthal particle distribution itself seems to be more correlated to the participant plane angle Ψ_{PP} than to the reaction plane angle Ψ_{RP} . Following this study we have also computed the distributions of the correlations

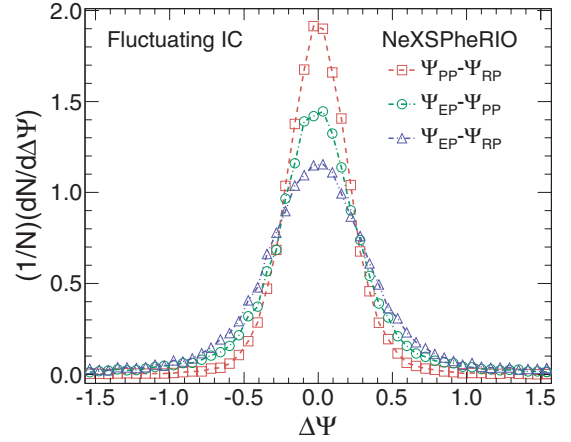


FIG. 2. (Color online) Correlation between different reference plane angles for mid-central NeXSPheRIO events. Red open squares (dashed line) show the difference between Ψ_{PP} and Ψ_{RP} , green open circles (dot-dashed line) show the difference between Ψ_{EP} and Ψ_{PP} , and blue open triangles (double-dot-dashed line) show the difference between Ψ_{EP} and Ψ_{RP} .

between Ψ_{RP} , Ψ_{EP} , and Ψ_{PP} . In Fig. 2 we show the distribution of the difference between the different angles, calculated for NeXSPheRIO mid-central events. The difference $\Psi_{PP} - \Psi_{RP}$ (shown in red open squares) provides the magnitude of the variation in the participant plane direction caused by event-by-event fluctuations in the initial energy density profile. The distribution of $\Psi_{EP} - \Psi_{RP}$ (shown in blue open triangles) represents the magnitude of the event plane dispersion caused by the IC fluctuations convoluted to the effect of resolution caused by the limited number of particles used in the Ψ_{EP} determination. The distribution of $\Psi_{EP} - \Psi_{PP}$ (shown in open green circles) is formed by the statistical resolution and the width of the true correlation between Ψ_{EP} and Ψ_{PP} , where by true correlation we mean the part that comes from the initial state fluctuation. For instance, assuming that the participant plane and the event plane are totally correlated (as in an event with smooth IC), the distribution of the difference between these two angles would be from the statistical resolution only. The width (σ) of the distribution of the angle difference as a function of the collision centrality is summarized in Fig. 3. From this plot we can see a decrease of the angular resolution for the very peripheral collisions and also for the very central collisions. The increase of the width for central collision is caused by the decrease of the initial eccentricity as well as the increase of the lumpiness in the initial energy density distribution, therefore, enhancing the effect of fluctuations.

An alternative method used in experimental data analysis to estimate the elliptic flow is through the cumulant formalism. The cumulant method calculates the flow coefficients directly from particle correlations, without the explicit need for a reference plane. The prescription used in this work is called Q cumulants or direct cumulants, presented in Ref. [30]. The procedure is divided into two parts where in the first part, the reference flow is calculated using all particles inside the selected range. In the second part, the differential flow (p_i dependent) is then calculated for the particles of interest. By

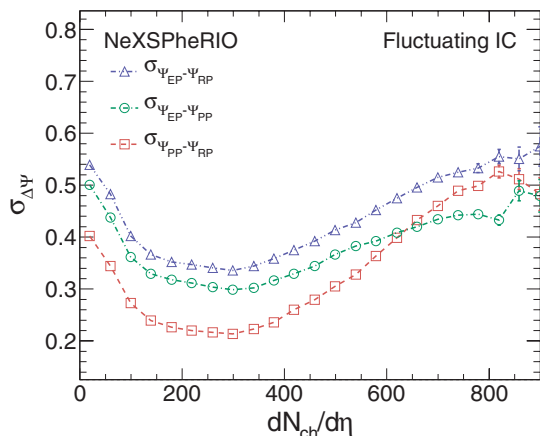


FIG. 3. (Color online) Width of the correlation between different reference plane angles as a function of the number of charged particles at $\eta = 0$. Red open squares (dashed line) show the width of $\Psi_{PP} - \Psi_{RP}$, green open circles (dot-dashed line) show the width of $\Psi_{EP} - \Psi_{PP}$, and blue open triangles (double-dot-dashed line) show the width of $\Psi_{EP} - \Psi_{RP}$.

following the notation used in Ref. [30], the flow coefficients for second- and fourth-order cumulants can be written as

$$\text{Reference flow: } \begin{cases} v_n\{2\} = \sqrt{\langle\langle 2 \rangle\rangle} \\ v_n\{4\} = \sqrt[4]{-\langle\langle 4 \rangle\rangle - 2 \cdot \langle\langle 2 \rangle\rangle^2}, \end{cases} \quad (10)$$

$$\text{Differential flow: } \begin{cases} v'_n\{2\} = \frac{\langle\langle 2' \rangle\rangle}{\sqrt{\langle\langle 2 \rangle\rangle}} \\ v'_n\{4\} = \frac{-[\langle\langle 4' \rangle\rangle - 2 \cdot \langle\langle 2' \rangle\rangle \langle\langle 2 \rangle\rangle]}{\sqrt[4]{-\langle\langle 4 \rangle\rangle - 2 \cdot \langle\langle 2 \rangle\rangle^2}}, \end{cases} \quad (11)$$

where the double brackets denote weighted averages of two- and four-particle correlations, first over the particles and then over the events. The weights are the total number of combinations from two- or four-particle correlations, respectively, and they are used to minimize the effects from multiplicity fluctuations. The advantage of the Q -cumulant method is that it is not necessary to perform nested loops to compute all possible combinations in multiparticle correlations. Instead, it uses the flow vectors to calculate directly the multiparticle cumulants (see details in Ref. [30]). The derivation of the expressions for higher order cumulants is straightforward. We present here only up to the fourth order because the sixth and higher orders do not seem to differ much from the fourth (see, for instance, Ref. [32]). In the next section we present the details of the simulated sample used, the centrality classes, and the estimates for v_2 obtained using each one of the methods described above.

IV. RESULTS

To compare the elliptic flow estimates obtained with different techniques, we used a sample of simulated events for Au + Au collisions at the center-of-mass energy of 200 GeV per nucleon pair. For the analysis presented here we used only the charged particles produced and the weak decays have been turned off to minimize the nonflow contribution. In our sample each event generated has associated a three-

TABLE I. Centrality classes of the NeXSPheRIO events used in the study.

Centrality	b range (fm)	$\langle dN_{ch}/d\eta \rangle$
0%–10%	0.00 – 4.78	576.5 ± 0.4
10%–20%	4.78 – 6.77	393.8 ± 0.3
20%–30%	6.77 – 8.29	271.5 ± 0.2
30%–40%	8.29 – 9.57	182.3 ± 0.1
40%–60%	9.57 – 11.72	89.9 ± 0.1

dimensional distribution of the energy density in the initial overlap region of the incident nuclei. The energy density distribution is computed from the energy-momentum tensor given by the NeXuS code at an initial stage and can fluctuate on an event-by-event basis, depending on the nuclear distribution determined by the incident nuclei and by the collision impact parameter [22]. Alternatively, it is also possible to input a smooth distribution generated by averaging over many events. For the analysis presented in the following we have divided the simulated sample into five event centrality classes for both fluctuating and smooth initial conditions, as described in Table I. A minimum of 30 000 events was used for each event centrality class, and for both smooth and fluctuating ICs. The total simulated data sample adds to more than a half-million events.

A. Flow comparisons

We present results obtained for the v_2 estimates as a function of the transverse momentum p_t , and centrality of the collision as given by the mean number of charged particles produced at $\eta = 0$. The calculations for all methods were performed using charged particles within the pseudorapidity window $|\eta| < 1.0$ and for $0.15 < p_t < 2.0$ GeV/c. The event plane determination was done using particles from $2.5 < |\eta| < 4.0$, therefore, avoiding autocorrelation on $v_2\{\text{EP}\}$ calculation, and with a requirement of a minimum of 15 charged particles within this pseudorapidity region for each event. We have also included published results reported by the STAR [32], PHENIX [33], and PHOBOS [34] experiments for comparison. In the following plots (Figs. 4–8), the v_2 calculated by the different methods are presented by different symbols. Blue open circles for the reaction plane v_2 , red open squares for the v_2 calculated through the event plane method, gray open diamonds for the participant plane v_2 , and green open triangles and orange open crosses for the v_2 calculated using two- and four-particle cumulant methods, respectively. In Figs. 4 and 5 we show the results for v_2 estimates as a function of transverse momentum for both smooth (left panels) and fluctuating (right panels) initial conditions, for centralities 10%–20% and 40%–60%, respectively. The blue star symbols are experimental results from the STAR experiment [32] and represent v_2 estimates obtained with the event plane method, and the yellow triangles are results from the PHENIX experiment [33] obtained with the second-order cumulant method. The p_t dependence of the v_2 curve from the NeXSPheRIO data generated using fluctuating ICs shows

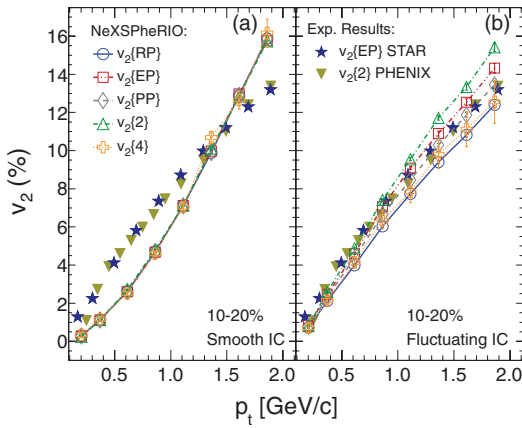


FIG. 4. (Color online) Differential v_2 as a function of the transverse momentum for 10%–20% central events. Open symbols are $v_2\{\text{RP}\}$ (blue circles), $v_2\{\text{EP}\}$ (red squares), $v_2\{\text{PP}\}$ (gray diamonds), $v_2\{2\}$ (green triangles), and $v_2\{4\}$ (orange crosses), for smooth (a) and fluctuating (b) initial conditions. Solid blue stars and yellow triangles are results from the STAR [32] and PHENIX [33] experiments, respectively. The lines are just to guide the eyes.

a better agreement to the experimental data than the results from the smooth IC. The comparisons presented in Figs. 4 and 5 also show that in the case of the smooth IC, the different methods for the v_2 estimate that are based on the measurement of the final state particles provide the same results as the v_2 calculated from the IC participant plane method. However, in the case of fluctuating ICs, there is a discrepancy between the experimentally measurable v_2 methods and the actual elliptic flow from the participant plane calculation. Moreover, the difference between the results from different v_2 methods increases with the transverse momentum. This shows clearly how each method is affected by the fluctuations in the initial condition. Assuming that the participant plane is the plane

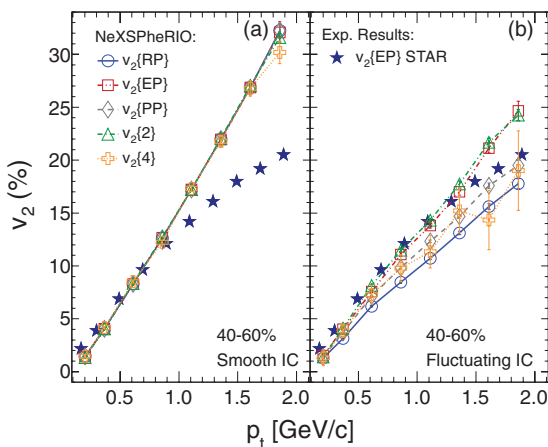


FIG. 5. (Color online) Differential v_2 as a function of the transverse momentum for 40%–60% peripheral events. Open symbols are $v_2\{\text{RP}\}$ (blue circles), $v_2\{\text{EP}\}$ (red squares), $v_2\{\text{PP}\}$ (gray diamonds), $v_2\{2\}$ (green triangles), and $v_2\{4\}$ (orange crosses), for smooth (a) and fluctuating (b) initial conditions. Solid blue stars are results from the STAR [32] experiment. The lines are just to guide the eyes.

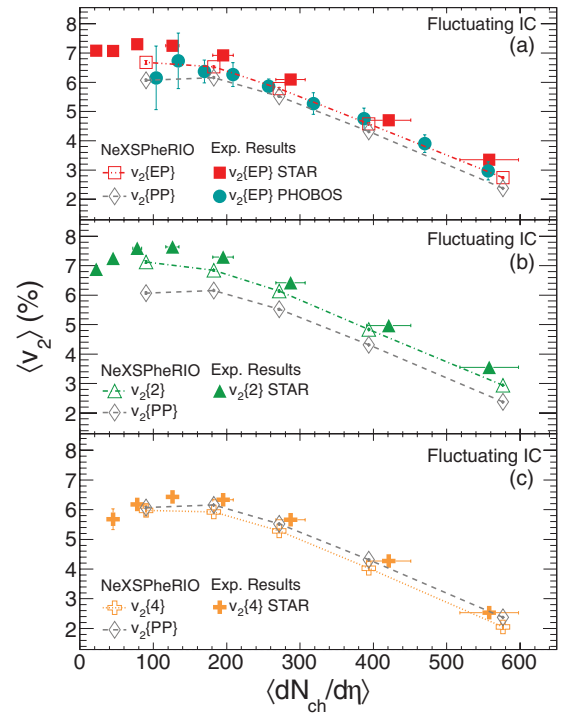


FIG. 6. (Color online) Integrated v_2 as a function of $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$. (a) Shows $v_2\{\text{EP}\}$ for NeXSPheRIO events (red open squares) and results from STAR [32] (red solid squares) and PHOBOS [34] (green solid circles) experiments; (b) shows $v_2\{2\}$ for NeXSPheRIO events (green open triangles) and from STAR experiment (green solid triangles) [32]; (c) shows $v_2\{4\}$ for NeXSPheRIO events (orange open crosses) and from STAR experiment (orange solid crosses) [32]. The v_2 results obtained with respect to the participant plane (gray open diamonds) were also included for reference. The lines are just to guide the eyes.

that defines the direction of the elliptic anisotropy created by the initial state and, therefore, defines the direction of the v_2 in the final state, it is possible to take the values of $v_2\{\text{PP}\}$ as a reference for the real $\langle v_2 \rangle$. Comparing to the results obtained with methods used experimentally (namely $v_2\{\text{EP}\}$, $v_2\{2\}$, and $v_2\{4\}$), we see $v_2\{\text{EP}\}$ above $v_2\{\text{PP}\}$ and below $v_2\{2\}$, consistent with $\langle v_2 \rangle \leq v_2\{\text{EP}\} \leq \sqrt{\langle v_2^2 \rangle}$ [25].

In Fig. 6 we show the dependence of the integrated v_2 with the collision centrality. In the Fig. 6(a) we present the results obtained with the event plane method; Fig. 6(b) shows the results obtained with the second-order cumulant, and Fig. 6(c) shows the results obtained with the fourth-order cumulant. Results from the STAR [32] and PHOBOS [34] experiments are also presented for comparison. In all three panels we also included the participant plane v_2 as a reference. The behavior of the flow obtained with the NeXSPheRIO model is very similar to the experimental results but the values are systematically lower. In Fig. 7 we have combined all the NeXSPheRIO results in the same panel and a clear ordering of the values from the different methods can be observed. This ordering presented by $v_2\{4\}$, $v_2\{\text{EP}\}$, and $v_2\{2\}$ is also observed in experimental results (see Ref. [32]). As expected, the elliptic flow calculated with respect to the reaction plane is

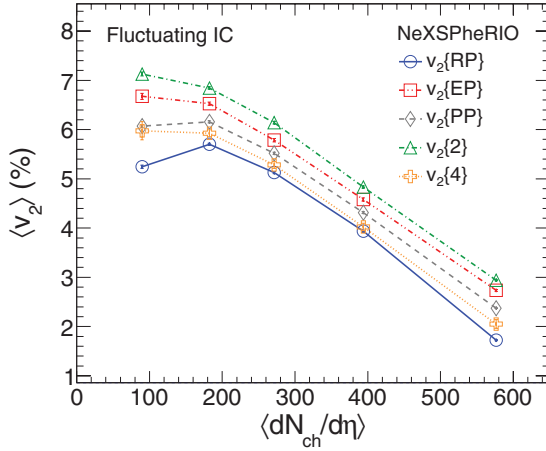


FIG. 7. (Color online) Mean integrated elliptic flow $\langle v_2 \rangle$ as a function of $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$. Blue open circles are $v_2\{RP\}$, red open squares are $v_2\{EP\}$, gray open diamonds are $v_2\{PP\}$, green open triangles are $v_2\{2\}$, and orange open crosses are $v_2\{4\}$. The lines are just to guide the eyes.

lower than all other results, confirming that in a scenario with lumpy initial energy density distribution, the plane defined by the impact parameter vector and the beam axis no longer drives the preferred direction of the flow. To see how the elliptic flow scales with the eccentricity of the initial state, we plotted the mean integrated v_2 obtained with the cumulant method over the respective eccentricity cumulant moment, calculated using the participant plane eccentricity [18,25,32]:

$$\varepsilon\{2\} = \sqrt{\langle \varepsilon_{PP}^2 \rangle}, \quad (12)$$

$$\varepsilon\{4\} = \sqrt[4]{2\langle \varepsilon_{PP}^2 \rangle^2 - \langle \varepsilon_{PP}^4 \rangle}, \quad (13)$$

as a function of the multiplicity density. Comparing these ratios with $v_2\{RP\}/\langle \varepsilon_{RP} \rangle$, $v_2\{PP\}/\langle \varepsilon_{PP} \rangle$, and $v_2\{EP\}/\varepsilon\{2\}$ as shown in Fig. 8, obtained for $|\eta| < 1.0$ (fluctuating ICs), we find a good agreement among the NeXSPheRIO results. The points obtained with different methods seem to fall almost on top of each other, with the only exception of $v_2\{2\}/\varepsilon\{2\}$, which is systematically higher. Such behavior is being investigated and we have already observed that the difference between $v_2\{2\}/\varepsilon\{2\}$ and the other ratios seems to vanish when increasing the pseudorapidity window considered in the calculations to $|\eta| < 6.0$. The observable $v_2\{4\}/\varepsilon\{4\}$ provides quite consistent results with $v_2\{PP\}/\langle \varepsilon_{PP} \rangle$, the later one being a quantity obtained directly from the anisotropy of the IC and, therefore, not accessible experimentally. For comparison, we also show results from the STAR [35] experiment for $v_2\{2\}/\varepsilon\{2\}$ and $v_2\{4\}/\varepsilon\{4\}$, with the eccentricity taken from the fKLN-CGC model, for Au + Au collisions at the center-of-mass energy of 200 GeV per nucleon pair. The experimental results are always higher than the NeXSPheRIO results, but they present similar separation between $v_2\{2\}/\varepsilon\{2\}$ and $v_2\{4\}/\varepsilon\{4\}$. Song *et al.* [10] have recently reported that the inclusion of viscous effects can reduce the baseline of ideal fluid v_2/ε . Moreover, they pointed out that a proper event-by-event treatment can

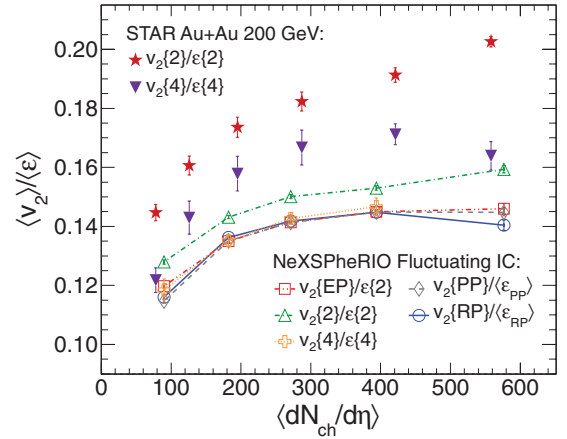


FIG. 8. (Color online) Eccentricity scaled v_2 as a function of $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$. Open symbols are $v_2\{PP\}/\langle \varepsilon_{PP} \rangle$ (gray diamonds), $v_2\{RP\}/\langle \varepsilon_{RP} \rangle$ (blue circles), $v_2\{EP\}/\varepsilon\{2\}$ (red squares), $v_2\{2\}/\varepsilon\{2\}$ (green triangles), and $v_2\{4\}/\varepsilon\{4\}$ (orange crosses), from NeXSPheRIO fluctuating ICs. Solid symbols are results from the STAR experiment [35] for $v_2\{2\}/\varepsilon\{2\}$ (red stars) and $v_2\{4\}/\varepsilon\{4\}$ (blue triangles). The error bars are the quoted statistical and systematic uncertainties added in quadrature. The lines are just to guide the eyes.

also affect this baseline. Thus, a study of the observables in an event-by-event scenario, as presented here, is also important for the comparison between data and viscous hydrodynamical models.

B. Flow fluctuations

The differences observed for the flow estimates indicate that each method responds differently to the fluctuations in the initial condition, which can be used to study these fluctuations with the final observables. In particular, from the definitions of flow estimates obtained with the cumulant formalism, it was suggested by the authors of Ref. [36] that, if only the leading order of $\sigma_{v_2}^2$ and δ are considered, then

$$v_2\{2\}^2 \approx \langle v_2 \rangle^2 + \delta + \sigma_{v_2}^2, \quad (14)$$

$$v_2\{4\}^2 \approx \langle v_2 \rangle^2 - \sigma_{v_2}^2, \quad (15)$$

where δ is the nonflow contribution (in general, correlations other than those related to the reaction plane), and σ_{v_2} is the elliptic flow fluctuation. The approximation in Eq. (15) is valid for $\sigma_{v_2} \ll \langle v_2 \rangle$ and negligible higher order moments. We note that this approximation breaks down for peripheral and central collisions where the skewness and kurtosis of the v_2 distribution and terms related to $\langle v_2 \rangle \sigma_{v_2}$ will contribute significantly to $v_2\{4\}$. From Eqs. (14) and (15), we can extract that the difference between $v_2\{2\}^2$ and $v_2\{4\}^2$ to provide a quantitative measure of the nonflow contribution added to the flow fluctuation as given by

$$v_2\{2\}^2 - v_2\{4\}^2 \approx \delta + 2\sigma_{v_2}^2. \quad (16)$$

We present this quantity as a function of the transverse momentum in Fig. 9. To reduce the statistical uncertainties in

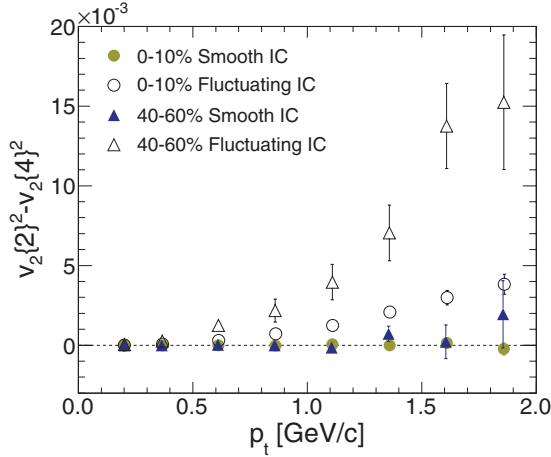


FIG. 9. (Color online) v_2 fluctuations as characterized by Eq. (16) as a function of transverse momentum. Results from smooth ICs are shown by solid symbols for 0%–10% central (yellow circles) and 40%–60% peripheral (blue triangles) events. Results from fluctuating ICs are shown by open symbols for 0%–10% central (circles) and 40%–60% peripheral (triangles) events.

low multiplicity events, the particle pseudorapidity acceptance was increased to $|\eta| < 6.0$. Results from the smooth IC case, shown in Fig. 9 as the solid yellow circles and the solid blue triangles, for event centrality 0%–10% and 40%–60%, respectively, are consistent with negligible fluctuation and nonflow contributions in the model. Results from the fluctuating IC events are presented in Fig. 9 by open circles for the 0%–10% central events and open triangles for the 40%–60% peripheral events. It is clear that in these cases, the quantity defined by Eq. (16) is nonzero, indicating that indeed such a parameter is sensitive to the fluctuations of the IC. The absolute magnitude of the fluctuation is higher in the peripheral 40%–60% events compared to the central events. Figure 10 shows the results of $v_2\{2\}^2 - v_2\{4\}^2$ calculated for the different event centrality classes. Solid squares represent the smooth IC case and open squares the fluctuating IC case, both for $|\eta| < 6.0$. The smooth IC case is consistent with zero, while the fluctuating IC case shows a steady decrease toward the most central collisions. In addition, we have also performed the analysis for fluctuating IC obtained within a tighter pseudorapidity window ($|\eta| < 1.0$), represented by open diamonds in Fig. 10. We observe a similar behavior compared to the fluctuating IC case for a wider η window, but with points systematically higher. The increase of the values observed for this case suggests a dependence of the fluctuations with the pseudorapidity window. Published data from the STAR experiment [35] are also shown for $|\eta| < 1.0$. These points present values always higher than the results from NeXSPheRIO, but the behavior is similar. Moreover, the contribution of nonflow effects is expected to be greater in experimental data.

Assuming a negligible nonflow component in Eq. (16), it is possible to take the difference $v_2\{2\}^2 - v_2\{4\}^2$ as an estimate of the absolute flow fluctuation. An estimate for the relative flow fluctuation can also be defined by using Eqs. (14) and

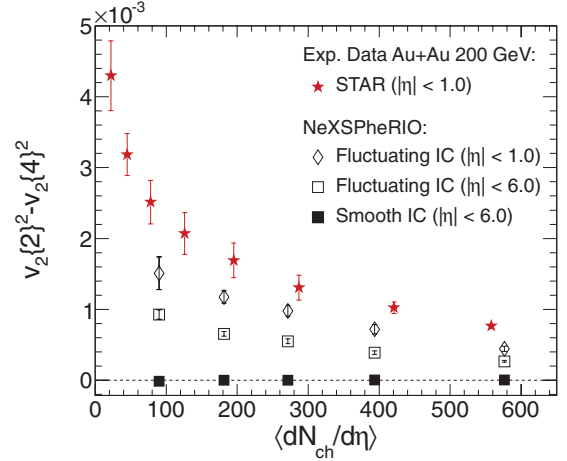


FIG. 10. (Color online) v_2 fluctuations as characterized by Eq. (16) as a function of $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$. The open diamonds are the results for the fluctuating IC case with $|\eta| < 1.0$, and the squares are the results for the fluctuating IC (open squares) and smooth IC (solid squares) cases, both for $|\eta| < 6.0$. Red solid stars are results from the STAR experiment [35]. The error bars are the quoted statistical and systematic uncertainties added in quadrature.

(15) as

$$R_v = \sqrt{\frac{v_2\{2\}^2 - v_2\{4\}^2}{v_2\{2\}^2 + v_2\{4\}^2}}. \quad (17)$$

In the case that the v_2 distribution is Gaussian-like, with the mean much larger than the width, the quantity R_v is a reasonable approximation for relative flow fluctuation $\sigma_{v_2}/\langle v_2 \rangle$ [26,37]. In Fig. 11 the R_v parameter as a function of p_t is presented for the most central 0%–10% and the peripheral 40%–60% event centrality classes. Opposite to what was observed for the absolute flow fluctuations in Fig. 9, in this case the central events show higher values than the peripheral events. The higher value of R_v observed in central events

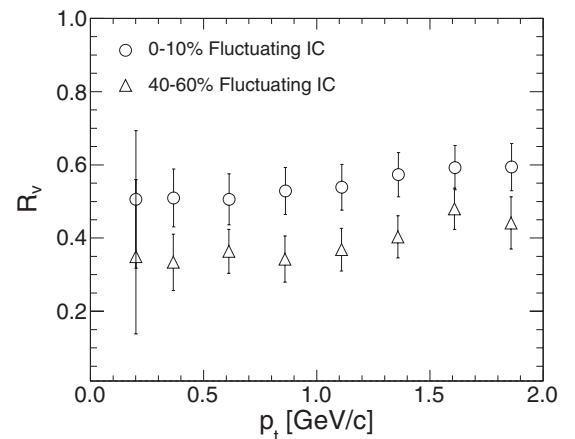


FIG. 11. R_v parameter as a function of the transverse momentum. Open circles represent the 0%–10% most central events and open triangles are for 40%–60% peripheral events.

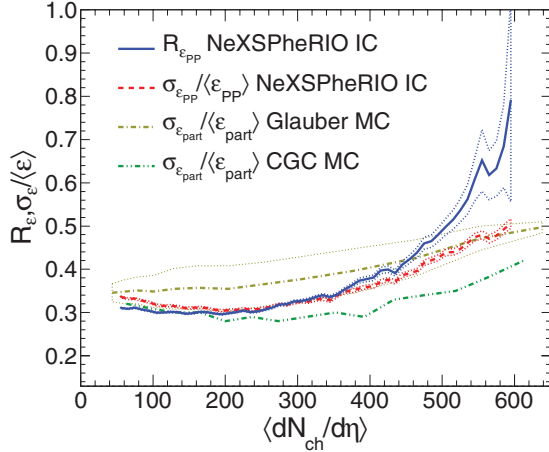


FIG. 12. (Color online) Comparison between the R_ϵ parameter (blue solid line) and the eccentricity relative fluctuation σ_ϵ/ϵ (red dashed line) calculated with respect to the participant plane in the NeXSPheRIO initial condition, as a function of $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$. Also shown is the relative eccentricity fluctuation calculated with the Glauber model (yellow dot-dashed line) and CGC (green double-dot-dashed line) Monte Carlo [38].

compared to the peripheral are from both the increase of the fluctuations in central collisions and also the increase of asymmetry in the v_2 distributions. The dependence of R_v with centrality will be further discussed next. In addition, we observe negligible dependence with transverse momentum up to 2 GeV/c for both central and peripheral events. Similarly to Eq. (17), it is possible to define an equivalent quantity for the eccentricity R_ϵ :

$$R_\epsilon = \sqrt{\frac{\epsilon\{2\}^2 - \epsilon\{4\}^2}{\epsilon\{2\}^2 + \epsilon\{4\}^2}}, \quad (18)$$

where $\epsilon\{2\}$ and $\epsilon\{4\}$ are given by Eqs. (12) and (13), respectively. Because we have access in our simulated events to both the IC and also the final state particles, we can verify if the quantity R_ϵ is a good approximation for the relative eccentricity fluctuation σ_ϵ/ϵ . In Fig. 12 we show the R_ϵ parameter as defined by Eq. (18) (blue solid line) and the relative eccentricity fluctuation σ_ϵ/ϵ (red dashed line), extracted from the participant plane eccentricity distribution of Fig. 1(b). We have also included for comparison results for the Glauber model (yellow dot-dashed line) and color glass condensate (CGC) (green double-dot-dashed line) Monte Carlo calculations [38]. In this case, the references reported these curves as a function of the number of participants in the collision (N_{part}). We used the values presented in Table II of Ref. [39] to convert N_{part} to the equivalent $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$, used in this work to summarize the collision centrality. This result shows that the observable R_v is a good measure of the relative fluctuations in mid-central collisions, but overestimates the fluctuations in central collisions. The higher value of R_v in central events is mainly from the contributions of higher moments of the ϵ_{pp} distribution such as skewness and kurtosis. The Gaussian-like condition is no longer satisfied

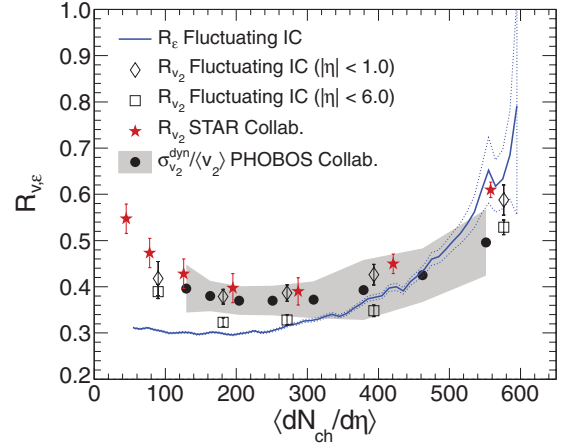


FIG. 13. (Color online) Comparison between the R_v and R_ϵ parameters as a function of $\langle dN_{ch}/d\eta \rangle$ at $\eta = 0$. The symbols represent R_v for the fluctuating IC case with $|\eta| < 6.0$ (open squares) and $|\eta| < 1.0$ (open diamonds). The blue solid line is the R_ϵ parameter calculated with respect to the participant plane in the NeXSPheRIO IC. Solid symbols are published results from PHOBOS [38] (black circles) and STAR [35] (red stars) experiments. The shaded band represent the errors quoted by PHOBOS and the error bars in STAR data are the statistical and systematic uncertainties added in quadrature.

and the approximations done in Eq. (15) that allows for the interpretation that R_v is a good measure of the flow fluctuation is no longer valid.

As an extension of the work already presented by Hama and collaborators [23] and Sorensen [40], we have also plotted the R_v parameter as given by Eq. (17) as a function of the mean number of charged particles produced at $\eta = 0$, hence, calculated for the different event centrality classes. The results are presented in Fig. 13. Open diamonds were obtained for $|\eta| < 1.0$ and the open squares for $|\eta| < 6.0$. In addition to the R_v parameter we have also included the R_ϵ parameter (blue solid line) calculated with respect to the participant plane obtained from the IC. In addition, results obtained for the dynamic relative flow fluctuation, $\sigma_{v_2}^{dyn}/\langle v_2 \rangle$ (black solid circles), reported by the PHOBOS experiment [38], and the R_v parameter (red solid stars) reported by the STAR experiment [35], are shown for comparison. Therefore, because the final state observable that we have is R_v , it is important to compare it with the initial state quantity R_ϵ , and not σ_ϵ/ϵ . The NeXSPheRIO values of R_v for $|\eta| < 1.0$ seem to agree very well with the experimental results. Even though a perfect agreement between the R_ϵ curve and the points calculated with the R_v parameter is not observed, it is remarkable that those quantities, obtained using the properties of the very beginning and the very final stage of the system evolution, still yield such similar values. Thus, we can conclude that the experimentally observable parameter R_v is a good estimate of the fluctuations in the initial state of the collisions, for mid-central events, which cannot be probed directly. Moreover, we can conclude that the R_v parameter is not affected by the hydrodynamic evolution and the freeze-out process.

V. SUMMARY

In this work we have analyzed a large number of simulated events produced with the NeXSPheRIO code, and both the IC and the final state particles were saved and analyzed on an event-by-event basis. Smooth and fluctuating ICs were used to generate the events for five different centrality classes. This allowed us to study the effects of fluctuations in the initial energy density distribution through the v_2 estimates from the final particle azimuthal distributions. We showed for the smooth IC cases that the methods generally used in experimental data analysis to estimate v_2 , namely $v_2\{\text{EP}\}$, $v_2\{2\}$, and $v_2\{4\}$, produce consistent results with flow estimates that use information from the initial condition, namely $v_2\{\text{RP}\}$ and $v_2\{\text{PP}\}$. In addition, for the fluctuating IC we found that the methods start to deviate from each other when going to higher transverse momentum, giving rise to a systematic ordering as a function of centrality. The discrepancy between the different v_2 calculation methods were exploited to be used as a measurement of the IC degree of fluctuation. Our results show that the magnitude of the nonflow plus the flow fluctuations increases with p_t for both central and peripheral

events, being more pronounced for the latter one. On the other hand, the relative fluctuations, as defined by the R_v parameter, show no dependence with transverse momentum. We also observed a dependence of the quantity defined in Eq. (16) with the collision centrality and the pseudorapidity window of the particles used in the analysis. Although the behavior observed for R_v and R_ε as a function of the mean number of charged particles at $\eta = 0$ does not completely agree, the values obtained are remarkably similar, which indicates that indeed the measurements of the final state flow fluctuations may provide important information on the initial state fluctuations.

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- [1] A. A. Penzias and R. W. Wilson, *Astrophys. J.* **142**, 419 (1965); W. Hu, *Lect. Notes Phys.* **470**, 207 (1996); E. Wright, 291 (2003).
- [2] A. P. Mishra, R. K. Mohapatra, P. S. Saumia, and A. M. Srivastava, *Phys. Rev. C* **77**, 064902 (2008).
- [3] J. Y. Ollitrault, *Phys. Rev. D* **46**, 229 (1992).
- [4] P. F. Kolb, U. Heinz, P. Huovinen, K. J. Eskola, and K. Tuominen, *Nucl. Phys. A* **696**, 197 (2001).
- [5] P. Huovinen, P. F. Kolb, U. Heinz, P. V. Ruuskanen, and S. A. Voloshin, *Phys. Lett. B* **503**, 58 (2001).
- [6] C. Nonaka, R. J. Fries, and S. A. Bass, *Phys. Lett. B* **583**, 73 (2004).
- [7] J. Adams *et al.* (STAR Collaboration), *Nucl. Phys. A* **757**, 102 (2005).
- [8] K. Adcox *et al.* (PHENIX Collaboration), *Nucl. Phys. A* **757**, 184 (2005).
- [9] P. Romatschke and U. Romatschke, *Phys. Rev. Lett.* **99**, 172301 (2007).
- [10] H. Song, S. A. Bass, U. Heinz, T. Hirano, and C. Shen, *Phys. Rev. Lett.* **106**, 192301 (2011).
- [11] B. Schenke, S. Jeon, and C. Gale, *Phys. Rev. Lett.* **106**, 042301 (2011).
- [12] B. Schenke, S. Jeon, and C. Gale, *Phys. Rev. C* **82**, 014903 (2010).
- [13] O. Socolowski, F. Grassi, Y. Hama, and T. Kodama, *Phys. Rev. Lett.* **93**, 182301 (2004).
- [14] R. Andrade, F. Grassi, Y. Hama, T. Kodama, and O. Socolowski, *Phys. Rev. Lett.* **97**, 202302 (2006).
- [15] R. Andrade, F. Grassi, Y. Hama, T. Kodama, and O. Socolowski, *Braz. J. Phys.* **37**, 717 (2007).
- [16] R. P. G. Andrade, Y. Hama, F. Grassi, O. Socolowski, and T. Kodama, *Braz. J. Phys.* **37**, 99 (2007).
- [17] R. P. G. Andrade, F. Grassi, Y. Hama, T. Kodama, and W.-L. Qian, *Phys. Rev. Lett.* **101**, 112301 (2008); C. E. Aguiar, Y. Hama, T. Kodama, and T. Osada, *Nucl. Phys. A* **698**, 639 (2002).
- [18] M. Miller and R. Snellings, [arXiv:cond-mat/0312008](https://arxiv.org/abs/cond-mat/0312008).
- [19] J. Takahashi, B. M. Tavares, W. L. Qian, R. Andrade, F. Grassi, Y. Hama, T. Kodama, and N. Xu, *Phys. Rev. Lett.* **103**, 242301 (2009).
- [20] X. Zhu, M. Bleicher, and H. Stöcker, *Phys. Rev. C* **72**, 064911 (2005).
- [21] H. Petersen, C. Coleman-Smith, S. A. Bass, and R. Wolpert, *J. Phys. G* **38**, 045102 (2011).
- [22] H. J. Drescher, F. M. Liu, S. Ostapchenko, T. Pierog, and K. Werner, *Phys. Rev. C* **65**, 054902 (2002); Y. Hama, T. Kodama, and O. Socolowski, *Braz. J. Phys.* **35**, 24 (2005).
- [23] Y. Hama *et al.*, *Phys. Atom. Nucl.* **71**, 1558 (2008).
- [24] F. Grassi, Y. Hama, T. Kodama, and O. Socolowski, *J. Phys. G* **31**, S1041 (2005).
- [25] B. Alver *et al.*, *Phys. Rev. C* **77**, 014906 (2008).
- [26] P. Sorensen, [arXiv:0905.0174](https://arxiv.org/abs/0905.0174).
- [27] A. M. Poskanzer and S. A. Voloshin, *Phys. Rev. C* **58**, 1671 (1998).
- [28] S. A. Voloshin, A. M. Poskanzer, and R. Snellings, [arXiv:0809.2949](https://arxiv.org/abs/0809.2949).
- [29] N. Borghini, P. M. Dinh, and J.-Y. Ollitrault, *Phys. Rev. C* **64**, 054901 (2001).
- [30] A. Bilandzic, R. Snellings, and S. Voloshin, *Phys. Rev. C* **83**, 044913 (2011).
- [31] H. Holopainen, H. Niemi, and K. J. Eskola, *Phys. Rev. C* **83**, 034901 (2011).
- [32] J. Adams *et al.* (STAR Collaboration), *Phys. Rev. C* **72**, 014904 (2005).
- [33] S. Afanasiev *et al.* (PHENIX Collaboration), *Phys. Rev. C* **80**, 024909 (2009).
- [34] B. B. Back *et al.* (PHOBOS Collaboration), *Phys. Rev. C* **72**, 051901 (2005).

- [35] G. Agakishiev *et al.* (STAR Collaboration), [arXiv:1111.5637](#).
- [36] J.-Y. Ollitrault, A. M. Poskanzer, and S. A. Voloshin, *Phys. Rev. C* **80**, 014904 (2009).
- [37] S. A. Voloshin, A. M. Poskanzer, A. Tang, and G. Wang, *Phys. Lett. B* **659**, 537 (2008).
- [38] H.-J. Drescher and Y. Nara, *Phys. Rev. C* **76**, 041903 (2007); B. Alver *et al.*, *Phys. Rev. Lett.* **104**, 142301 (2010).
- [39] B. I. Abelev *et al.* (STAR Collaboration), *Phys. Rev. C* **79**, 034909 (2009).
- [40] P. Sorensen, *J. Phys. G* **34**, S897 (2007).