# Microscopic study of Ca + Ca fusion

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We investigate the fusion barriers for reactions involving Ca isotopes  ${}^{40}$ Ca +  ${}^{40}$ Ca,  ${}^{40}$ Ca +  ${}^{48}$ Ca, and  ${}^{48}$ Ca +  ${}^{48}$ Ca using the microscopic time-dependent Hartree-Fock (TDHF) theory coupled with a density constraint. In this formalism the fusion barriers are directly obtained from TDHF dynamics. We also study the excitation of the preequilibrium GDR for the  ${}^{40}$ Ca +  ${}^{48}$ Ca system and the associated  $\gamma$ -ray emission spectrum. Fusion cross sections are calculated using the incoming-wave boundary condition approach. We examine the dependence of fusion barriers on collision energy as well as on the different parametrizations of the Skyrme interaction.

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## I. INTRODUCTION

The microscopic study of nuclear many-body problem and the understanding of the nuclear interactions that reproduce the observed structure and reaction properties are the underlying challenges of low-energy nuclear physics. In this context, detailed investigations of the fusion process will lead to a better understanding of the interplay among the strong, Coulomb, and weak interactions as well as the enhanced correlations present in these many-body systems.

Recently, particular experimental attention has been given to fusion reactions involving Ca isotopes [1-4]. These new experiments supplemented the older fusion data [5] and extended it to lower sub-barrier energies. Comparison of the sub-barrier cross sections with those calculated using standard coupled-channel calculations suggested a hindrance of the fusion crosssections at deep sub-barrier energies [1-3]. One of the underlying reasons for the failure of the standard coupled-channel approach is the use of frozen densities in the calculation of double-folding potentials, resulting in potentials that behave in a completely unphysical manner for deep sub-barrier energies. While the outer part of the barrier is largely determined by the early entrance channel properties of the collision, the inner part of the potential barrier is strongly sensitive to dynamical effects such as particle transfer and neck formation. This has been remedied in part by extensions of the coupled-channel approach to include a repulsive core [6] or the incorporation of neck degrees of freedom [7,8]. More recent calculations [4,9,10] using the coupled-channel approach with a repulsive core have provided much improved fits to the data. A detailed microscopic study of the fusion process for Ca-based reactions  ${}^{40}Ca + {}^{40}Ca + {}^{40}Ca + {}^{48}Ca$ , and  $^{48}$ Ca +  $^{48}$ Ca could provide further insight into the reaction dynamics as well as a good testing ground for the theory, since these isotopes are commonly used in fitting the parameters of the effective nuclear interactions, such as the Skyrme force.

During the past several years, we have developed a microscopic approach for calculating heavy-ion interaction potentials that incorporates all of the dynamical entrance channel effects included in the time-dependent Hartree-Fock (TDHF) description of the collision process [11,12]. The method is based on the TDHF evolution of the nuclear system coupled with density-constrained Hartree-Fock calculations

(DC-TDHF) to obtain the ion-ion interaction potential. The formalism was applied to study fusion cross sections for the systems <sup>132</sup>Sn + <sup>64</sup>Ni [13], <sup>64</sup>Ni + <sup>64</sup>Ni [14], <sup>16</sup>O + <sup>208</sup>Pb [15], and <sup>132,124</sup>Sn + <sup>96</sup>Zr [16], as well as to the study of the entrance channel dynamics of hot and cold fusion reactions leading to superheavy element Z = 112 [17], and dynamical excitation energies [18]. In all cases, we have found good agreement between the measured fusion cross sections and the DC-TDHF results. This is rather remarkable given the fact that the only input in DC-TDHF is the Skyrme effective N-N interaction, and there are no adjustable parameters.

In Sect. II we outline the main features of our microscopic approach, the DC-TDHF method. In Sec. II we also discuss the calculation of ion-ion separation distance, coordinatedependent mass, calculation of fusion cross sections, and giant dipole resonance (GDR) formalism. In Sec. III we present interesting aspects of the reaction dynamics and compare our results with experiment and other calculations. In Sec. V we summarize our conclusions.

#### **II. FORMALISM**

## A. DC-TDHF method

In the DC-TDHF approach [11] the TDHF time-evolution takes place with no restrictions. At certain times during the evolution the instantaneous density is used to perform a static Hartree-Fock minimization while holding the neutron and proton densities constrained to be the corresponding instantaneous TDHF densities [19,20]. In essence, this provides us with the TDHF dynamical path in relation to the multidimensional static energy surface of the combined nuclear system. The advantages of this method in comparison to other mean-field based microscopic methods such as the constrained Hartree-Fock (CHF) method are obvious. First, there is no need to introduce artificial constraining operators which assume that the collective motion is confined to the constrained phase space: second, the static adiabatic approximation is replaced by the dynamical analog, where the most energetically favorable state is obtained by including sudden rearrangements, and the dynamical system does not have to move along the valley of the potential energy surface. In short, we have a self-organizing system which selects its evolutionary path by itself following the microscopic dynamics. All of the dynamical features included in TDHF are naturally included in the DC-TDHF calculations. These effects include neck formation, mass exchange, internal excitations, deformation effects to all order, as well as the effect of nuclear alignment for deformed systems. In the DC-TDHF method the ion-ion interaction potential is given by

$$V(R) = E_{\rm DC}(R) - E_{\rm A_1} - E_{\rm A_2}, \tag{1}$$

where  $E_{DC}$  is the density-constrained energy at the instantaneous separation R(t), while  $E_{A_1}$  and  $E_{A_2}$  are the binding energies of the two nuclei obtained with the same effective interaction. In writing Eq. (1) we have introduced the concept of an adiabatic reference state for a given TDHF state. The difference between these two energies represents the internal energy. The adiabatic reference state is the one obtained via the density constraint calculation, which is the Slater determinant with lowest energy for the given density with vanishing current and approximates the collective potential energy [19]. We would like to emphasize again that this procedure does not affect the TDHF time evolution, and contains no free parameters or normalization.

In addition to the ion-ion potential it is also possible to obtain coordinate dependent mass parameters. One can compute the "effective mass" M(R) using the conservation of energy

$$M(R) = \frac{2[E_{\rm c.m.} - V(R)]}{\dot{R}^2},$$
(2)

where the collective velocity  $\dot{R}$  is directly obtained from the TDHF evolution and the potential V(R) from the density constraint calculations. In calculating fusion cross sections, this coordinate-dependent mass is used to obtain a transformed ion-ion potential as described below.

#### **B.** Excitation energy

The calculation of the excitation energy is achieved by dividing the TDHF motion into a collective and intrinsic part. The major assumption in achieving this goal is to assume that the collective part is primarily determined by the density  $\rho(\mathbf{r}, t)$  and the current  $\mathbf{j}(\mathbf{r}, t)$ . Consequently, the excitation energy can be formally written as [18]

$$E^*(t) = E_{\text{TDHF}} - E_{\text{coll}}(\rho(t), \mathbf{j}(t)), \qquad (3)$$

where  $E_{\text{TDHF}}$  is the total energy of the dynamical system, which is a conserved quantity, and  $E_{\text{coll}}$  represents the collective energy of the system. In the next step we break up the collective energy into two parts,

$$E_{\text{coll}}(t) = E_{\text{kin}}(\rho(t), \mathbf{j}(t)) + E_{\text{DC}}(\rho(t)), \qquad (4)$$

where  $E_{kin}$  represents the kinetic part and is given by

$$E_{\rm kin}(\rho(t),\mathbf{j}(t)) = \frac{m}{2} \int d^3r \, \mathbf{j}^2(t)/\rho(t), \tag{5}$$

which is asymptotically equivalent to the kinetic energy of the relative motion,  $\frac{1}{2}\mu \dot{R}^2$ , where  $\mu$  is the reduced mass and R(t) is the ion-ion separation distance. The dynamics of the ion-ion

separation R(t) is provided by an unrestricted TDHF run, thus allowing us to deduce the excitation energy as a function of the distance parameter,  $E^*(R)$ .

## C. Calculation of R

In practice, TDHF runs are initialized with energies above the Coulomb barrier at some large but finite separation. The two ions are boosted with velocities obtained by assuming that the two nuclei arrive at this initial separation on a Coulomb trajectory. Initially the nuclei are placed such that the point x = 0 in the x-z plane is the center of mass. During the TDHF dynamics the ion-ion separation distance is obtained by constructing a dividing plane between the two centers and calculating the center of the densities on the left and right halves of this dividing plane. The coordinate R is the difference between the two centers. The dividing plane is determined by finding the point at which the tails of the two densities intersect each other along the x axis. Since the actual mesh used in the TDHF calculations is relatively coarse, we use a cubic-spline interpolation to interpolate the profile in the x direction and search for a more precise intersection value. This procedure has been recently described in Ref. [21] in great detail.

The standard procedure for calculating *R* as described above starts to fail after a substantial overlap is reached (for *R* values smaller than the ones studied in this manuscript). We instead define the ion-ion separation as  $R = R_0 \sqrt{|Q_{20}|}$ , where  $Q_{20}$  is the mass quadrupole moment for the entire system, calculated by using the collision axis as the symmetry axis, and  $R_0$  is a scale factor determined to give the correct initial separation distance at the start of the calculations. Calculating *R* this way yields almost identical results to the previous procedure until that procedure begins to fail, and continues smoothly after that point. Of course the minimum value of *R* calculated this way is never zero but is determined by the quadrupole moment of the composite system.

#### D. Fusion cross section

We now outline the calculation of the total fusion cross section using an arbitrary coordinate-dependent mass M(R). Starting from the classical Lagrange function

$$L(R, R) = \frac{1}{2}M(R)R^2 - V(R),$$
(6)

we obtain the corresponding Hamilton function

.

$$H(R, P) = \frac{P^2}{2M(R)} + V(R),$$
(7)

where the canonical momentum is given by P = M(R)R. Following the standard quantization procedure for the kinetic energy in curvilinear coordinates [22]

$$T = \frac{-\hbar^2}{2} \left[ g^{-\frac{1}{2}} \frac{\partial}{\partial q^{\mu}} g^{\frac{1}{2}} g^{\mu\nu} \frac{\partial}{\partial q^{\nu}} \right], \tag{8}$$

where  $g_{\mu\nu}(q)$  denotes the metric tensor and  $g^{\mu\nu}(q)$  the reciprocal tensor, one obtains the quantized Hamiltonian

$$H(R, \hat{P}) = \frac{1}{2} \left[ M(R)^{-\frac{1}{2}} \hat{P} M(R)^{-\frac{1}{2}} \hat{P} \right] + V(R).$$
(9)

with the momentum operator  $\hat{P} = -i\hbar d/dR$ . The total fusion cross section

$$\sigma_f = \frac{\pi}{k^2} \sum_{L=0}^{\infty} (2L+1)T_L,$$
 (10)

can be obtained by calculating the potential barrier penetrabilities  $T_L$  from the Schrödinger equation for the relative motion coordinate *R*, using the Hamiltonian (9) with an additional centrifugal potential

$$\left[H(R, \hat{P}) + \frac{\hbar^2 L(L+1)}{2M(R)R^2} - E_{\text{c.m.}}\right]\psi_L(R) = 0.$$
(11)

Alternatively, instead of solving the Schrödinger equation with coordinate dependent mass parameter M(R) for the heavy-ion potential V(R), we can instead use the constant reduced mass  $\mu$  and transfer the coordinate dependence of the mass to a scaled potential  $U(\bar{R})$  using the well known coordinate scale transformation [23]

$$d\bar{R} = \left(\frac{M(R)}{\mu}\right)^{\frac{1}{2}} dR.$$
 (12)

Integration of Eq. (12) yields

$$\bar{R} = f(R) \iff R = f^{-1}(\bar{R}).$$
 (13)

As a result of this point transformation, both the classical Hamilton function, Eq. (7), and the corresponding quantum mechanical Hamiltonian, Eq. (9), now assume the form

$$H(\bar{R}, \bar{P}) = \frac{\bar{P}^2}{2\mu} + U(\bar{R}),$$
(14)

and the scaled heavy-ion potential is given by the expression

$$U(\bar{R}) = V(R) = V(f^{-1}(\bar{R})).$$
(15)

The fusion barrier penetrabilities  $T_L(E_{c.m.})$  are obtained by numerical integration of the two-body Schrödinger equation

$$\left[\frac{-\hbar^2}{2\mu}\frac{d^2}{d\bar{R}^2} + \frac{L(L+1)\hbar^2}{2\mu\bar{R}^2} + U(\bar{R}) - E\right]\psi = 0, \quad (16)$$

using the *incoming wave boundary condition* (IWBC) method [24]. IWBC assumes that once the minimum of the potential is reached fusion will occur. In practice, the Schrödinger equation is integrated from the potential minimum  $R_{\min}$ , where only an incoming wave is assumed, to a large asymptotic distance, where it is matched to incoming and outgoing Coulomb wave functions. The barrier penetration factor  $T_L(E_{c.m.})$  is the ratio of the incoming flux at  $R_{\min}$  to the incoming Coulomb flux at large distance. Here, we implement the IWBC method exactly as it is formulated for the coupled-channel code CCFULL described in Ref. [25]. This gives us a consistent way of calculating cross sections at above and below the barrier energies.

#### E. GDR excitation

Let us now consider a central collision in the x direction and introduce the quantity

$$D(t) = \frac{NZ}{A} \left[ \frac{1}{Z} \sum_{p=1}^{Z} \langle x_p(t) \rangle - \frac{1}{N} \sum_{n=1}^{N} \langle x_n(t) \rangle \right], \quad (17)$$

which represents the expectation value of the *x* component of the dipole operator  $d_x/e$  taken with the time-dependent TDHF Slater determinant  $|\Phi(t)\rangle$ . Following the bremsstrahlung approach developed by Baran *et al.* [26,27], we define the dipole acceleration

$$D''(t) = \frac{d^2 D(t)}{dt^2}$$
(18)

and introduce its Fourier transform

$$D''(\omega) = \int_{t_{\min}}^{t_{\max}} D''(t)e^{i\omega t}dt.$$
 (19)

Alternatively, for nearly harmonic vibrations one can use the expression  $|D''(\omega)|^2 = \omega^4 |D(\omega)|^2$ , with

$$D(\omega) = \int_{t_{\min}}^{t_{\max}} D(t) e^{i\omega t} \sin^4 \left( \pi \frac{t - t_{\min}}{t_{\max} - t_{\min}} \right) dt.$$
(20)

The time-filtering  $\sin^4$  is used to smooth out peaks coming from finite integration time. The "power spectrum" of the electric dipole radiation is given by [26]

$$\frac{dP}{dE_{\gamma}} = \frac{2\alpha}{3\pi E_{\gamma}} \left| \frac{1}{c} D''(\omega) \right|^2, \qquad (21)$$

where  $\alpha = e^2/(\hbar c) \approx 1/137$  denotes the fine-structure constant. Recently, preequilibrium GDR excitation has also been studied in the context of TDHF [28].

#### **III. RESULTS**

Calculations were done in three-dimansional (3D) geometry and using the full Skyrme interaction, including all of the time-odd terms in the mean-field Hamiltonian [29]. The primary Skyrme parametrization used was SLy4 [30], but we have also tested the new UNEDF0 [31] and UNEDF1 [32] parametrizations. For the reactions studied here, the lattice spans 48 fm along the collision axis and 15 fm in the other two directions. Derivative operators on the lattice are represented by the basis-spline collocation method. One of the major advantages of this method is that we may use a relatively large grid spacing of 1.0 fm and nevertheless achieve high numerical accuracy. The initial separation of the two nuclei is 18 fm for central collisions. The time propagation is carried out using a Taylor series expansion (up to orders 10-12) of the unitary mean-field propagator, with a time step  $\Delta t = 0.4$  fm/c. We have performed density constraint calculations every 10-20 time steps. The accuracy of the density constraint calculations is commensurate with the accuracy of the static calculations.

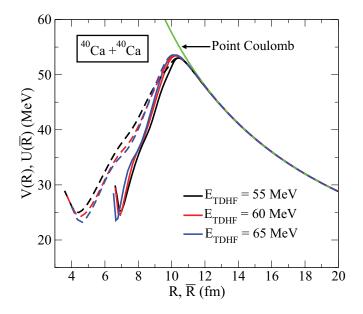


FIG. 1. (Color online) Potential barriers V(R) for the  ${}^{40}\text{Ca} + {}^{40}\text{Ca}$  system obtained from density constrained TDHF calculations using Eq. (1) at three different energies,  $E_{\text{TDHF}} = 55$  MeV (black solid curve),  $E_{\text{TDHF}} = 60$  MeV (red solid curve), and  $E_{\text{TDHF}} = 65$  MeV (blue solid curve). The three dashed curves correspond to the transformed potential in Eq. (15) using the coordinate-dependent masses. Also shown is the point Coulomb potential.

# A. <sup>40</sup>Ca + <sup>40</sup>Ca system

In this subsection we will present our results for the  ${}^{40}Ca + {}^{40}Ca$  system. Most of our general discussions will be provided here as they would be the same for other systems. Specific points about individual systems and comparison of results for the three systems will be taken up in the subsequent sections.

In Fig. 1 we show the microscopic ion-ion potential barriers obtained using Eq. (1) calculated at three different collision energies,  $E_{\text{TDHF}} = 55$  MeV (black solid curve),  $E_{\text{TDHF}} = 60 \text{ MeV}$  (red solid curve), and  $E_{\text{TDHF}} = 65 \text{ MeV}$ (blue solid curve). We observe a relatively small dependence on the collision energy. At lower energies the system has more time available for rearrangements to take place through the formation of a neck, whereas this is less and less the case at higher energies, and the potential barrier approaches the frozen-density limit [21]. This produces the observed trend in Fig. 1, where the lowest barrier peak corresponds to the lowest energy, and the barrier height increases with increasing collision energy. The barrier heights are 53.02, 53.43, and 53.57 MeV, respectively. Similar energy dependence was also observed in the DD-TDHF calculations of Ref. [21], and becomes more prevalent for heavier systems. Similarly, the position of the barrier peak moves toward smaller Rvalues, albeit very slowly in this case, with increasing energy. Corresponding R values for the barrier maximum are 10.41, 10.32, and 10.23 fm. What is also shown in Fig. 1 is the Coulomb potential, assuming the two nuclei to be point particles with Z = 20. During the approach phase the microscopically calculated DC-TDHF potential traces the

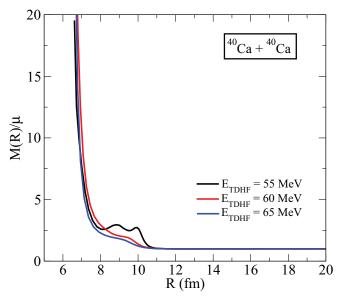


FIG. 2. (Color online) Coordinate dependent mass M(R) scaled by the constant reduced mass  $\mu$ , obtained from Eq. (2), at three different TDHF energies.

point Coulomb potential, differing by less than 150 keV, which provides a test for the numerical accuracy.

The energy dependence of potential barriers can also be understood if we examine the coordinatedependent mass of Eq. (2) shown in Fig. 2 for three different energies,  $E_{\text{TDHF}} = 55 \text{ MeV}$  (black solid curve),  $E_{\text{TDHF}} = 60 \text{ MeV}$ (red solid curve), and  $E_{\text{TDHF}} = 65$  MeV (blue solid curve). The R dependence of this mass at lower energies is very similar to the one found in CHF calculations [23]. On the other hand, at higher energies the coordinate-dependent mass essentially becomes flat, which is again a sign that most dynamical effects are contained at lower energies. The peak at small R values is due to the fact that the center-of-mass energy is above the barrier, and the denominator of Eq. (2) becomes small due to the slowdown of the ions. We have used the coordinate dependent masses shown in Fig. 2 to obtain the scaled potentials  $U(\bar{R})$  of Eq. (15). These potentials are shown as the dashed curves in Fig. 1. As we see, the coordinate-dependent mass only changes the inner parts of the barriers for all energies. Furthermore, the effect is largest for the lowest energy collision and diminishes as we increase the collision energy.

We have obtained the fusion cross sections by numerical integration of Eq. (16). The resulting cross sections are shown in Fig. 3. We observe that all of the scaled barriers give a very good description of the experimental fusion cross sections. The high-energy part of the fusion cross sections are primarily determined by the barrier properties in the vicinity of the barrier peak. On the other hand sub-barrier cross sections are influenced by what happens in the inner part of the barrier, and here the dynamics and consequently the coordinate dependent mass becomes very important. As we observe in Fig. 3 we get a very good agreement with experiment for the barriers obtained at the lowest two collision energies, whereas the  $E_{c.m.} = 65$  MeV curve slightly underestimates the

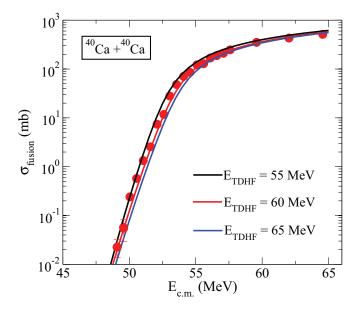


FIG. 3. (Color online) Total fusion cross section as a function of  $E_{\rm c.m.}$ . Three separate theoretical cross section calculations are shown, based on the energy-dependent DC-TDHF heavy-ion potentials V(R) at energies  $E_{\rm TDHF} = 55, 60, \text{ and } 65 \text{ MeV}$ . The experimental data (filled circles) are taken from Ref. [4].

cross section at lower energies. Although not shown in Fig. 3, the cross sections obtained using the the unscaled potentials V(R) and a constant reduced mass  $\mu$  also agree well with the data at higher energies, but either significantly overestimate or underestimate the cross section at lower energies due to the absence of the coordinate-dependent mass. In summary, the calculated fusion cross sections for the <sup>40</sup>Ca + <sup>40</sup>Ca system reproduce the experimental cross sections reasonably well, which is a testament that TDHF with Skyrme force provides a good description for this collision.

# B. <sup>48</sup>Ca + <sup>48</sup>Ca system

In Fig. 4 we show the microscopic ion-ion potential barriers for the stable neutron-rich  ${}^{48}Ca + {}^{48}Ca$  system calculated at three different collision energies,  $E_{\text{TDHF}} = 55 \text{ MeV}$ (black solid curve),  $E_{\text{TDHF}} = 65 \text{ MeV}$  (red solid curve), and  $E_{\text{TDHF}} = 70 \text{ MeV}$  (blue solid curve). The barrier heights, for increasing collision energy, are 50.98, 51.37, and 51.52 MeV, respectively, with corresponding R values of 10.80, 10.63, and 10.52 fm. The comparison of these barriers with those of the  ${}^{40}$ Ca +  ${}^{40}$ Ca system shows that the barrier heights are reduced by about 2 MeV, and the location of the barrier maximum is at a slightly larger R value. This is due to the fact that two two  $^{48}$ Ca nuclei are larger than the corresponding <sup>40</sup>Ca nuclei, and thus their outer skins come into contact at a larger R value. After this point the nuclear interaction sets in, causing the trajectory to deviate from the point-Coulomb one, and producing a peak at a lower energy value.

Figure 5 shows the fusion cross sections for the  ${}^{48}Ca + {}^{48}Ca$  system calculated using the barriers of Fig. 4. While the overall quality of the agreement with the experimental fusion cross sections is good, specially for the last two collision energies,

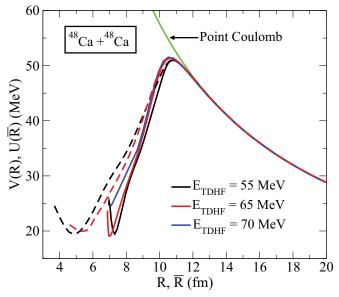


FIG. 4. (Color online) Potential barriers V(R) for the <sup>48</sup>Ca + <sup>48</sup>Ca system obtained from density-constrained TDHF calculations using Eq. (1) at three different energies,  $E_{\text{TDHF}} = 55$  MeV (black solid curve),  $E_{\text{TDHF}} = 65$  MeV (red solid curve), and  $E_{\text{TDHF}} = 70$  MeV (blue solid curve). The dashed curves correspond to the transformed potential in Eq. (15) using the coordinate dependent masses. Also shown is the point Coulomb potential.

the curve corresponding to the lowest collision energy of 55 MeV overestimates the experiment for lower  $E_{c.m.}$  values. This illustrates the sensitivity of the results to the height of the potential barrier, the difference in this case between the two barriers being around 0.4 MeV.

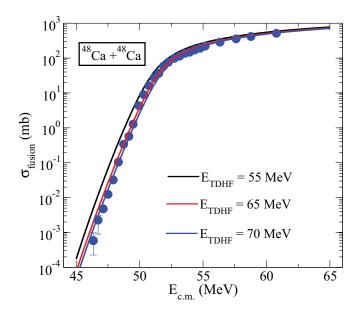


FIG. 5. (Color online) Fusion cross sections for the <sup>48</sup>Ca + <sup>48</sup>Ca system as a function of  $E_{c.m.}$ . Three separate theoretical cross-section calculations are shown, based on the energy-dependent DC-TDHF heavy-ion potentials V(R) at energies  $E_{TDHF} = 55$ , 65, and 70 MeV. The experimental data (filled circles) are taken from Ref. [4].

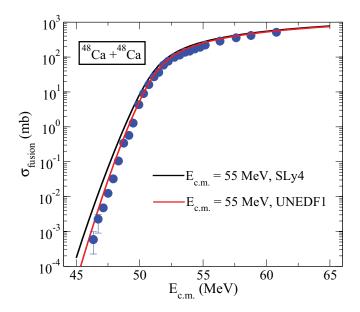


FIG. 6. (Color online) Total fusion cross section as a function of  $E_{\text{c.m.}}$  for two different parametrizations of the Skyrme force, SLy4 and UNEDF1, at the collision energy of  $E_{\text{TDHF}} = 55$ . The experimental data (filled circles) are taken from Ref. [4].

While the quality of the DC-TDHF results for the <sup>48</sup>Ca + <sup>48</sup>Ca system is very good for a parameterfree microscopic approach, we have decided to investigate this further. One of the problems is that the Skyrme fits do badly in reproducing the single-particle properties [33], particularly the neutron single-particle states. For example, with SLy4 parametrization the neutron levels  $1d_{5/2}$ ,  $2s_{1/2}$ , and  $1d_{3/2}$ in <sup>48</sup>Ca nucleus are about 7-3 MeV lower in energy than the corresponding experimental values [31]. In one of the parametrizations of the nuclear density functional, UNEDF0, these single-particle energies are raised to values closer to experimental ones. However, in this case the magic gap at N = 28 vanishes [31]. Recently, a new parametrization, UNEDF1, was introduced with a better incorporation of large deformations among other improvements [32]. Another possible advantage of the UNEDF1 parametrization for TDHF calculations is that no center-of-mass correction term was used in the functional, which is also the case in TDHF. We have tried this parametrization in our DC-TDHF calculation for the  $^{48}$ Ca +  $^{48}$ Ca system at  $E_{c.m.} = 55$  MeV. In practice, we had to reduce our time step from  $\Delta t = 0.4 \text{ fm}/c$  to  $\Delta t = 0.1 \text{ fm}/c$ , and our density-constraint convergence parameters by a factor of 10 or more. This is probably due to the larger power of the density in the  $t_3$  term of the Skyrme interaction. The peak of the fusion barrier is slightly broader, and higher in energy by 187 keV for UNEDF1. The use of UNEDF1 results in an improvement of the fusion cross sections as shown in Fig. 6. We have also tried using the parametrization SLy5, but this resulted in no appreciable difference from the SLv4 case.

# C. <sup>40</sup>Ca + <sup>48</sup>Ca system

In this section we will examine the fusion of the asymmetric  ${}^{40}$ Ca +  ${}^{48}$ Ca system. For highlighting the differences among all

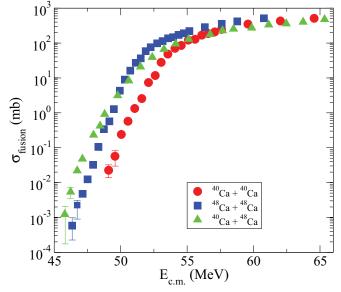


FIG. 7. (Color online) Experimental fusion cross sections as a function of  $E_{c.m.}$  for the three systems studied. The experimental data are taken from Ref. [4].

three systems we have plotted only the experimental fusion cross sections in Fig. 7. Relative to the  ${}^{40}Ca + {}^{40}Ca$  and  ${}^{48}Ca + {}^{48}Ca$  systems the  ${}^{40}Ca + {}^{48}Ca$  fusion cross sections show a different systematic behavior. At higher bombarding energies the  ${}^{40}Ca + {}^{48}Ca$  cross sections fall below the  ${}^{40}Ca + {}^{40}Ca$  data points, whereas at sub-barrier energies they rise above the  ${}^{48}Ca + {}^{48}Ca$  data. Historically, the theoretical description of the fusion cross sections for the  ${}^{40}Ca + {}^{48}Ca$  system have been complicated by the couplings to various transfer channels with positive Q values [34,35]. A strong enhancement of fusion cross sections at lower energies seen in Fig. 7 was attributed to this effect [9].

In Fig. 8 we show the microscopic ion-ion potential barriers for the <sup>40</sup>Ca + <sup>48</sup>Ca system calculated at three different collision energies,  $E_{\text{TDHF}} = 55$  MeV (black solid curve),  $E_{\text{TDHF}} = 60$  MeV (red solid curve), and  $E_{\text{TDHF}} = 70$  MeV (blue solid curve). The barrier heights, for increasing collision energy, are 51.11, 51.45, and 51.81 MeV, respectively, with corresponding *R* values of 10.75, 10.65, and 10.51 fm. We have used the coordinate-dependent masses to obtain the scaled potentials  $U(\bar{R})$  of Eq. (15). These potentials are also shown as the dashed curves in Fig. 8. As before, the coordinate-dependent mass only changes the inner parts of the barriers for all energies, and the effect diminishes as we increase the collision energy.

Figure 9 shows the fusion cross sections for the  ${}^{40}$ Ca +  ${}^{48}$ Ca system calculated using the barriers of Fig. 8. The observed trend for sub-barrier energies is typical for DC-TDHF calculations when the underlying microscopic interaction gives a good representation of the participating nuclei. Namely, the potential barrier corresponding to the lowest collision energy gives the best fit to the sub-barrier cross sections, since this is the one that allows for more rearrangements to take place and grows the inner part of the barrier. Considering the fact that historically the low-energy sub-barrier cross sections of

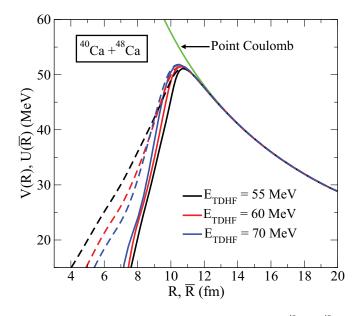


FIG. 8. (Color online) Potential barriers for the  ${}^{40}\text{Ca} + {}^{48}\text{Ca}$  system obtained from density-constrained TDHF calculations at three different energies,  $E_{\text{TDHF}} = 55$  MeV (black solid curve),  $E_{\text{TDHF}} = 60$  MeV (red solid curve), and  $E_{\text{TDHF}} = 70$  MeV (blue solid curve). The three dashed curves correspond to the transformed potential in Eq. (15) using the coordinate-dependent masses. Also shown is the point Coulomb potential.

the  ${}^{40}Ca + {}^{48}Ca$  system have been the ones not reproduced well by the standard models, the DC-TDHF results are quite satisfactory, indicating that the dynamical evolution of the nuclear density in TDHF gives a good overall description of the collision process. The shift of the cross-section curve with

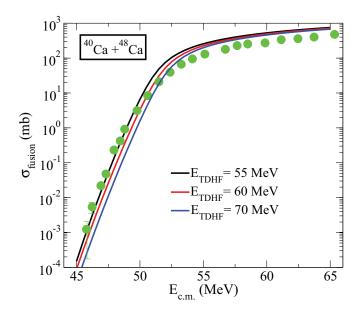


FIG. 9. (Color online) Fusion cross sections for the  ${}^{40}Ca + {}^{48}Ca$  system as a function of  $E_{c.m.}$ . Three separate theoretical cross section calculations are shown, based on the energy-dependent DC-TDHF heavy-ion potentials V(R) at energies  $E_{TDHF} = 55$ , 60, and 70 MeV. The experimental data (filled circles) are taken from Ref. [4].

increasing collision energy is typical. In principle one could perform a DC-TDHF calculation at each energy above the barrier and use that cross section for that energy. However, this would make the computations extremely time consuming and may not provide much more insight.

The trend at higher energies is atypical. The calculated cross sections are larger than the experimental ones by about a factor of 2. Such lowering of fusion cross sections with increasing collision energy is commonly seen in lighter systems where various inelastic channels, clustering, and molecular formations are believed to be the contributing factors [36]. In the recent coupled-channel approach this is solved by the addition of a small imaginary potential near the minimum of the repulsive core [37]. Such an imaginary part was found not to be necessary for the  ${}^{40}Ca + {}^{40}Ca$  and  $^{48}$ Ca +  $^{48}$ Ca systems [9,10]. At this time we do not have access to coupled-channel results for the  ${}^{40}Ca + {}^{48}Ca$  system. We have repeated our DC-TDHF calculations using the UNEDF1 interaction, which resulted in a small improvement, reducing the difference with the experimental values to a factor of about 1.5. This issue will be discussed further in the next section, where we examine the excitation properties obtained from TDHF calculations.

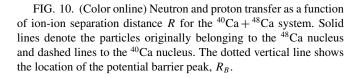
### **IV. EXCITATIONS**

Excitations are believed to have a significant impact on the outcome of the fusion reactions. The excitations can range from the entrance channel quantal excitations of the projectile and target, as in the coupled-channel approach, to collective excitations of the preequilibrium system, to compound nucleus excitations. These can be further influenced by particle transfer, preequilibrium emissions, evaporation, and other events. Theoretically such effects are commonly introduced by hand into various reaction models. However, the influence of excitations on nuclear reaction dynamics remains a difficult and unresolved problem, as it combines both nuclear structure and dynamics under nonequilibrium conditions.

In Sec. II B we have outlined the calculation of the dynamical excitation energy within the DC-TDHF formalism. This approach was used to calculate excitation energies for various systems, including the excitation energy of the heavy systems leading to superheavy formations [17]. Here, we have shown that what is also important is the excitation of the system at the point of capture, which is a deciding factor for forming a composite system or fusion-fission. This is different from excitation of the system from quasifission. The dynamical excitation energies calculated from DC-TDHF are relative to an unequilibrated composite system rather than a true compound nucleus.

The systems studied in this manuscript present interesting target-projectile combinations for studying excitations. For example, the nonsymmetric  ${}^{40}Ca + {}^{48}Ca$  system will have a pre-equilibrium GDR excitation in comparison to the other two symmetric systems, as well as particle transfer. Naturally, some of these effects are included in the fusion barriers obtained via the DC-TDHF method, as they influence the change in

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the TDHF density used in the density-constraint calculation. The quantity  $E^*(R(t))$  calculated via TDHF and DC-TDHF represents the average of dynamical excitations present in the mean-field theory, with the exception of excitations not determined by the nuclear density and current (such as the spin currents). Another point to consider is the pairing interaction. It has been a general consensus that during a heavy-ion collision the effects of pairing are washed away due to the high excitations. In this particular case paring effects are minimal for the initial Ca nuclei studied. However, even if the pairing can be ignored during the collision process, in many cases it plays an important role for obtaining good initial HF states, as well as obtaining realistic density constraint solutions when a single composite is formed. Fortunately, the latter case corresponds to the region of barrier minimum and not the region around the barrier peak, where most fusion cross sections are measured. This may in some way explain the success of the DC-TDHF barriers for fusion.

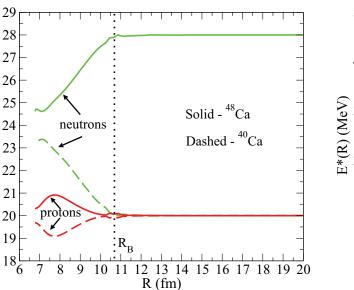
In Fig. 10 we plot the average number of neutrons and protons transferred during the early stages of the TDHF collision of the  ${}^{40}\text{Ca} + {}^{48}\text{Ca}$  system at  $E_{\text{TDHF}} = 60$  MeV. The solid lines denote the neutrons and protons asymptotically belonging to the <sup>48</sup>Ca nucleus, and dashed lines denote those belonging to the <sup>40</sup>Ca nucleus. A number of interesting things can be observed from the plot. The first is the fact that most of the transfer seems to start after we pass the potential barrier peak. This indicates that particle transfer primarily modifies the inner part of the barrier and not so much the barrier height. The other observation is that on average about three neutrons are transferred from <sup>48</sup>Ca to <sup>40</sup>Ca, but there is also a small amount of proton transfer in the opposite direction.

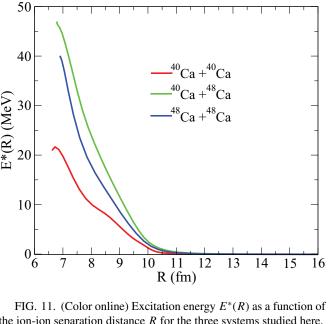
Figure 11 shows the excitation energy  $E^*(R)$  for the three systems studied here. The excitation energy was calculated for

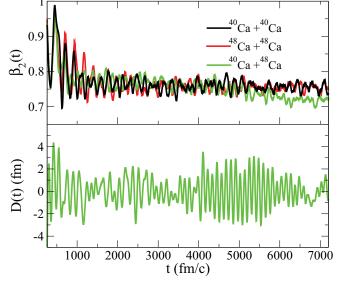
the ion-ion separation distance R for the three systems studied here.

the same value of  $\varepsilon = E_{\text{c.m.}}/\mu = 2.75$  MeV for all systems, which corresponds to collision energies of 55, 60, and 66 MeV, respectively. All curves initially behave in a similar manner: at large distances the excitation is zero, and as the nuclei approach the barrier peak the excitations start and monotonically rise for larger overlaps. The interesting observation is that the excitations for the intermediary  ${}^{40}Ca + {}^{48}Ca$  system start at a slightly earlier time and rise above the other two systems. This may be largely due to the fact that an asymmetric system has some additional modes of excitation in comparison to the

FIG. 12. (Color online) Time evolution of isoscalar deformation parameter  $\beta_2$  for the head-on collision of all three systems at a collision energy of  $E_{\text{TDHF}} = 55$  MeV (top panel), and the time evolution of the isovector dipole amplitude D(t) for the <sup>40</sup>Ca + <sup>48</sup>Ca system (bottom panel).







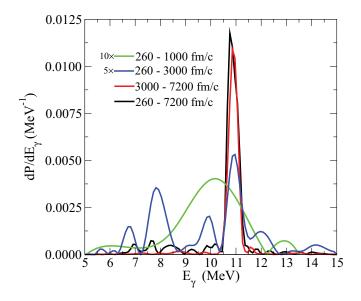


FIG. 13. (Color online) Power spectrum of the isovector dipole amplitude D(t) for the <sup>40</sup>Ca + <sup>48</sup>Ca system is shown for various time-intervals.

other two symmetric systems. It may be plausible to consider the direct influence of the excitation energy  $E^*(R)$  on the fusion barriers by making an analogy with the coupled-channel approach, and constructing a new potential  $V^*(R) = V(R) +$  $E^*(R)$ , which has all the excitations added to the ion-ion potential V(R) that should be calculated at higher energies to minimize the nuclear rearrangements (frozen-density limit). The resulting potentials somewhat resemble the repulsive-core coupled-channel potentials of Ref. [9]. This approach does lead to improvements in cases where most of the excitation energy is in the form of collective excitations rather than irreversible stochastic dissipation (true especially for lighter systems). For the cases studied here we do see a small improvement for the fusion cross sections of the  ${}^{40}Ca + {}^{48}Ca$  system. The viability of this approach requires further examination and will be studied in the future.

One of the excitation modes for the  ${}^{40}\text{Ca} + {}^{48}\text{Ca}$  system, namely the particle transfer, was already discussed above. The others are various isovector modes such as the preequilibrium GDR excitation. In Fig. 12 we show the time development of the isoscalar deformation parameter  $\beta_2 = \frac{4\pi}{5} \langle r^2 Y_{20} \rangle / Ar_{\text{rms}}$ (top panel) and the isovector dipole amplitude D(t) of Eq. (17) (bottom panel). The isovector amplitudes for the two symmetric systems would appear as a zero-line on this plot. We have left out the initial approach phase of the collision (about 260 fm/c) from these plots. The time evolution of the deformation  $\beta_2$  is very similar for all three systems, except at larger times the  ${}^{40}Ca + {}^{48}Ca$  system moves toward smaller deformation in comparison to the other two systems. This is again most likely due to increased excitations that drive the system toward a more compact shape. The evolution of the dipole operator D(t) shows little damping over a long time interval. The power spectrum associated with this time evolution is shown in Fig. 13. We have evaluated the spectrum for different time intervals. The broadest curve corresponds to a very early stage of the collision, 260-1000 fm/c, but has the lowest amplitude (multiplied by 10 in Fig. 13). In the second interval, 260–3000 fm/c, we see the development of various peaks and an increase in amplitude. The spectrum for the entire time interval, 260–7200 fm/c, is dominated by a sharp peak around 11 MeV primarily originating from the later stages of the collision, which is evident if one compares it to the spectrum from the interval 3000-7200 fm/c.

#### V. SUMMARY

In this article we have provided a microscopic study of Ca + Ca fusion using the DC-TDHF approach. These reactions have recently been of considerable interest for the fusion community, with a flurry of phenomenological analyses of the fusion data. Here, we have provided a microscopic alternative to these analyses. We have shown that microscopically obtained ion-ion potentials do give a reasonably good description of the fusion cross sections.

The fully microscopic TDHF theory has shown itself to be rich in nuclear phenomena and continues to stimulate our understanding of nuclear dynamics. The time-dependent mean-field studies seem to show that the dynamic evolution builds up correlations that are not present in the static theory. While modern Skyrme forces provide a much better description of static nuclear properties in comparison to the earlier parametrizations, there is a need to obtain even better parametrizations that incorporate deformation and reaction data into the fit process.

#### ACKNOWLEDGMENTS

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- [1] A. M. Stefanini et al., Phys. Lett. B 679, 95 (2009).
- [2] C. L. Jiang et al., Phys. Rev. C 82, 041601 (2010).
- [3] G. Montagnoli et al., Nucl. Phys. A 834, 159c (2010).
- [4] G. Montagnoli et al., Phys. Rev. C 85, 024607 (2012).
- [5] H. A. Aljuwair, R. J. Ledoux, M. Beckerman, S. B. Gazes, J. Wiggins, E. R. Cosman, R. R. Betts, S. Saini, and Ole Hansen, Phys. Rev. C 30, 1223 (1984).
- [6] Ş. Mişicu and H. Esbensen, Phys. Rev. Lett. 96, 112701 (2006).
- [7] T. Ichikawa, K. Hagino, and A. Iwamoto, Phys. Rev. C 75, 057603 (2007).
- [8] T. Ichikawa, K. Hagino, and A. Iwamoto, Phys. Rev. C 75, 064612 (2007).
- [9] H. Esbensen, C. L. Jiang, and A. M. Stefanini, Phys. Rev. C 82, 054621 (2010).
- [10] Ş. Mişicu and F. Carstoiu, Phys. Rev. C 84, 051601(R) (2011).

- [11] A. S. Umar and V. E. Oberacker, Phys. Rev. C 74, 021601(R) (2006).
- [12] A. S. Umar and V. E. Oberacker, J. Phys. G 36, 025101 (2009).
- [13] A. S. Umar and V. E. Oberacker, Phys. Rev. C 76, 014614 (2007).
- [14] A. S. Umar and V. E. Oberacker, Phys. Rev. C 77, 064605 (2008).
- [15] A. S. Umar and V. E. Oberacker, Eur. Phys. J. A 39, 243 (2009).
- [16] V. E. Oberacker, A. S. Umar, J. A. Maruhn, and P.-G. Reinhard, Phys. Rev. C 82, 034603 (2010).
- [17] A. S. Umar, V. E. Oberacker, J. A. Maruhn, and P.-G. Reinhard, Phys. Rev. C 81, 064607 (2010).
- [18] A. S. Umar, V. E. Oberacker, J. A. Maruhn, and P.-G. Reinhard, Phys. Rev. C 80, 041601(R) (2009).
- [19] R. Y. Cusson, P.-G. Reinhard, M. R. Strayer, J. A. Maruhn, and W. Greiner, Z. Phys. A **320**, 475 (1985).
- [20] A. S. Umar, M. R. Strayer, R. Y. Cusson, P.-G. Reinhard, and D. A. Bromley, Phys. Rev. C 32, 172 (1985).
- [21] K. Washiyama and D. Lacroix, Phys. Rev. C 78, 024610 (2008).
- [22] M. R. Spiegel, *Vector Analysis*, Schaum's Outline Series (McGraw-Hill, New York, 1959).
- [23] K. Goeke, F. Grümmer, and P.-G. Reinhard, Ann. Phys. 150, 504 (1983).
- [24] G. H. Rawitscher, Phys. Rev. 135, B605 (1964).
- [25] K. Hagino, N. Rowley, and A. T. Kruppa, Comput. Phys. Commun. 123, 143 (1999).

- [26] V. Baran, M. Colonna, M. Di Toro, A. Guarnera, and A. Smerzi, Nucl. Phys. A 600, 111 (1996).
- [27] V. Baran, C. Rizzo, M. Colonna, M. DiToro, and D. Pierroutsakou, Phys. Rev. C 79, 021603(R) (2009).
- [28] C. Simenel, Ph. Chomaz, and G. de France, Phys. Rev. C 76, 024609 (2007).
- [29] A. S. Umar and V. E. Oberacker, Phys. Rev. C 73, 054607 (2006).
- [30] E. Chabanat, P. Bonche, P. Haensel, J. Meyer, and R. Schaeffer, Nucl. Phys. A 635, 231 (1998); 643, 441(E) (1998).
- [31] M. Kortelainen, T. Lesinski, J. Moré, W. Nazarewicz, J. Sarich, N. Schunck, M. V. Stoitsov, and S. Wild, Phys. Rev. C 82, 024313 (2010).
- [32] M. Kortelainen, J. McDonnell, W. Nazarewicz, P.-G. Reinhard, J. Sarich, N. Schunck, M. V. Stoitsov, and S. M. Wild, Phys. Rev. C 85, 024304 (2012).
- [33] M. Kortelainen, J. Dobaczewski, K. Mizuyama, and J. Toivanen, Phys. Rev. C 77, 064307 (2008).
- [34] S. Landowne, C. H. Dasso, R. A. Broglia, and G. Pollarolo, Phys. Rev. C 31, 1047 (1985).
- [35] H. Esbensen, S. H. Fricke, and S. Landowne, Phys. Rev. C 40, 2046 (1989).
- [36] K. T. R. Davies, K. R. S. Devi, S. E. Koonin, and M. R. Strayer, in *Treatise on heavy ion Science*, Vol. 3, edited by D. A. Bromley (Plenum, New York, 1985), p. 3.
- [37] H. Esbensen, Phys. Rev. C 77, 054608 (2008).