Isoscalar monopole excitations in ¹⁶O: α -cluster states at low energy and mean-field-type states at higher energy

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Isoscalar monopole strength function in ¹⁶O up to $E_x \simeq 40$ MeV is discussed. We found that the fine structures at the low-energy region up to $E_x \simeq 16$ MeV in the experimental monopole strength function obtained by the ¹⁶O(α , α') reaction can be rather satisfactorily reproduced within the framework of the 4 α cluster model, while the gross three bump structures observed at the higher energy region ($16 \leq E_x \leq 40$ MeV) look likely to be approximately reconciled by the mean-field calculations such as random-phase approximation and quasirandomphase approximation. In this paper, it is emphasized that two different types of monopole excitations exist in ¹⁶O; one is the monopole excitation to cluster states which is dominant in the lower-energy part ($E_x \leq 16$ MeV), and the other is the monopole excitation of the mean-field type such as one-particle, one-hole (1p1h), which is attributed mainly to the higher-energy part ($16 \leq E_x \leq 40$ MeV). It is found that this character of the monopole excitations originates from the fact that the ground state of ¹⁶O with the dominant doubly closed shell structure has a duality of the mean-field type as well as α -clustering character. This dual nature of the ground state seems to be a common feature in light nuclei.

DOI: 10.1103/PhysRevC.85.034315

PACS number(s): 23.20.-g, 21.60.Gx, 27.20.+n

I. INTRODUCTION

Isoscalar monopole excitation in nuclei provides important information on its underlying structure. In the collective liquid drop model, the isoscalar giant monopole resonance (ISGMR), which has been established in medium and heavy nuclei [1], corresponds to a breathing mode of the nucleus arising owing to in-phase oscillations of the proton and neutron fluids. In heavy nuclei, the ISGMR is observed as a single peak in the α inelastic scattering cross sections at small angles, and its excitation energy follows an empirical formula $E_x \simeq 80 A^{-1/3}$ MeV, which is directly related to the compressibility of nuclear matter. A lot of work has been done to extract experimentally the nuclear compressibility by comparing it with microscopic calculations, for example, using the random-phase approximation (RPA). Recently the isoscalar monopole distributions in ⁹⁰Zr, ¹¹⁶Sn, ¹⁴⁴Sm, and ²⁰⁸Pb were measured with greater precision than previously [2]. The results indicated that the compressibility of nuclear matter is $K_{\rm nm} = 231 \pm 5$ MeV.

It is interesting to study what happens for the ISGMR in lighter nuclei. When the nuclear masses decrease from medium nuclei to light nuclei, the surface-energy correction becomes more important and the excitation energy of ISGMR should become lower compared with the empirical formula, indicating a lower nuclear compressibility [3–5]. A lot of theoretical

work has been so far devoted to the study of ISGMR in light nuclei, for example, within the RPA framework [4,6-10] and others [11].

The RPA calculations with the nonrelativistic framework were performed in ¹⁶O, ⁴⁰Ca, ⁹⁰Zr, and ²⁰⁸Pb [4,6-9]. According to the results with the Gogny force and Skyrme forces, etc. [4], it was found that (1) the isoscalar monopole strength in ¹⁶O spreads out over some energy region of $20 \leq E_x \leq 40$ MeV, with its centroid energy being $E_x = 22 \sim$ 29 MeV, the value of which depends on the NN interactions employed; (2) the monopole strength becomes more and more concentrated in a single peak as the nucleus becomes heavier; (3) the percentage of energy-weighted sum rule carried by the resonances increases with the mass number of the nucleus; and (4) the nuclear compressibility becomes smaller as the nucleus becomes lighter. However, Ma et al. investigated the isoscalar monopole modes in ¹⁶O, ⁴⁰Ca, ⁹⁰Zr, and ²⁰⁸Pb using the relativistic RPA (RRPA) method [10]. They found that when going from heavy to lighter nuclei, the single-peak structure of ISGMR in ²⁰⁸Pb changes to a peak with several small humps in ⁴⁰Ca and eventually the monopole strength spreads out widely to form a couple of peaks in ¹⁶O. Hence, it was pointed out that it becomes difficult, in a nucleus such as ⁴⁰Ca, to define theoretically the energy and width of the ISGMR.

The experimental isoscalar monopole strengths with a great precision were recently provided in ¹²C, ¹⁶O, and ²⁴Mg up



FIG. 1. Experimental isoscalar monopole strength function of ¹⁶O [12] is shown by the histogram. The experimental data below $E_x \approx$ 10 MeV are absent because of an energy cut in the experimental condition. The real line is the calculated result by the relativistic RPA calculation [10] multiplied by 0.25 and shifted down in energy by 4.2 MeV. This figure is taken from Ref. [12].

to $E_x \simeq 50$ MeV, using inelastic scattering of α particles, by the Texas group [12]. They found that the isoscalar monopole strength in light nuclei does not concentrate on a single peak and the monopole strength spreads out in several regions of energies. The histogram in Fig. 1 shows the experimental isoscalar monopole strength function in ${}^{16}O[12]$. It is compared with the RRPA calculation by Ma et al. [10]. It was found that the centroid in the RRPA response function is at 25.3 MeV, which is higher than the experimental data $(E_x = 21.13 \pm 0.49 \text{ MeV})$. To match their calculation to the experimental centroid, the calculated strength function was shifted down in energy by 4.2 MeV and furthermore they normalized it to approximately 30% of the isoscalar energy weighted sum rule (EWSR) by multiplying the RRPA curve by a factor of 0.25 [12]. Then, the normalized and shifted curve and the experimental result are in moderately good agreement with each other with respect to the shape of the gross three-peak structure. However, their calculation failed to reproduce the 0^+ states found in the low-energy region (5 $\lesssim E_x \lesssim$ 16 MeV), in particular, at $E_x = 6.05$, 12.05, and 14.1 MeV observed in inelastic α scattering and electron scattering, etc. [12,13]. According to the ${}^{16}O(e, e')$ experiments [13], the three states are excited rather strongly by the (e, e') reaction, and their monopole matrix elements are 3.55 ± 0.21 , 4.03 ± 0.09 , and 3.3 ± 0.7 fm², respectively, comparable to the single-particle monopole strength [14]. The total percentage of the energy weighted strength to the isoscalar monopole EWSR for these three 0^+ states amounts to be as large as over 15% [13,14].

In the nonrelativistic calculation for ^{16}O [4] a significant discrepancy is also revealed as compared with the experimental

data, in particular, in the low-energy region $(5 \leq E_x \leq E_x)$ 16 MeV), although the gross structures at the higher-energy region ($E_x \gtrsim 20$ MeV) in the RPA calculations are in rather good agreement with the experimental data. This discrepancy in the low-energy region can also be seen in Fig. 2 obtained by the recent second random-phase approximation (SRPA) calculations with a Skyrme force for ${}^{16}O$ [15], in which the coupling between 1p1h and 2p2h as well as between 2p2h configurations among themselves are fully taken into account. In particular, their calculation fails to reproduce the monopole transition strength to the 0_2^+ state at $E_x = 6.05$ MeV observed by the ${}^{16}O(e, e')$ experiment. Thus, the monopole strengths in the lower-energy region (5 $\leq E_x \leq$ 16 MeV) are likely to be out of scope in the mean-field theory. These results mean that the monopole strength function of ¹⁶O is not fully understood in the mean-field theory at the present stage, and other degrees of freedom beyond the mean field should be taken into account.

The 0_2^+ and 0_3^+ levels of ¹⁶O including its ground state, together with their monopole strengths, have in the past nicely been reproduced with a semimicroscopic cluster model, that is, the $\alpha + {}^{12}$ C orthogonality condition model (OCM) [16]. The OCM is an approximation of the resonating group method (RGM) [17]. Many successful applications of OCM are reported in Ref. [18]. The $\alpha + {}^{12}$ C OCM calculation as well as the $\alpha + {}^{12}$ C generator-coordinate-method one [19] demonstrates that the 0_2^+ state at $E_x = 6.05$ MeV and the 0_3^+ state at $E_x = 12.05$ MeV have $\alpha + {}^{12}$ C structures, where the α particle orbits around the 12 C(0⁺) core in an *S* wave and around the 12 C(2⁺) core in a *D* wave, respectively.



FIG. 2. (Color online) RPA [dashed (black) lines] and SRPA [solid (red) lines] for the isoscalar monopole strength distributions of ¹⁶O. This figure is taken from Ref. [15].

The 14.1-MeV 0^+ state, however, could not be explained by the $\alpha + {}^{12}$ C model calculations [16,19].

Recently, the structure study of ¹⁶O has made a great advance up to $E_x \simeq 16$ MeV around the 4 α disintegration threshold. The six lowest 0⁺ states of ¹⁶O, up to $E_x \simeq 16$ MeV, including the ground state, have for first time been reproduced very well with the 4 α OCM [20,21]. The 4 α OCM shares 68% of the EWSR value of the isoscalar monopole transition of ¹⁶O, while the $\alpha + {}^{12}$ C OCM shares 31%, as is discussed below. Thus, it is interesting to investigate whether the 4 α OCM can reproduce the experimental isoscalar monopole strength function in the low-energy region up to $E_x \simeq 16$ MeV in ¹⁶O, a region that is difficult to be treated in the mean-field theory. As is discussed below, the five excited 0⁺ states of ¹⁶O up to $E_x \simeq 16$ MeV have α -cluster structures [16,18–21].

The purpose of the present paper is twofold: The first is to show that the isoscalar monopole strength function calculated with the 4α OCM is in good correspondence to the experimental one in the low-energy region up to $E_x \simeq 16 \,\mathrm{MeV}$ shown in Fig. 1, and the second is to emphasize two features in the isoscalar monopole excitation of $^{\bar{1}6}O$, that is, that the monopole excitation to cluster states is dominant in the lowerenergy part ($E_x \leq 16 \,\text{MeV}$) of the monopole strength function, whereas the monopole excitation of the 1p1h type contributes to the higher-energy region ($16 \leq E_x \leq 40$ MeV). The two features have been implied in a mixed model calculation of the $\alpha + {}^{12}C$ cluster and symplectic basis functions [22], although the model space of the simplectic group [23,24] and the effective NN force used in its calculation are significantly different from those in the RPA calculations [4,6-9,15]. This mixed calculation confirmed that the strength carried by the $\alpha + {}^{12}C$ cluster model keeps its basic property. We show that the two features arise from the fact that the ground state of ¹⁶O originally possesses a dual nature, allowing α -type excitations as well as 1p1h-type ones, as is discussed below. In this paper, a shell-model calculation with the model space of 0s-, 0p-, 0d1s-, and 0f1p-shells for ¹⁶O is also performed to investigate the extent to which the shell model works for describing the low-lying 0^+ states.

In Sec. II, the monopole excitation function with the 4α OCM is formulated after a brief explanation of the 4α OCM framework together with the shell-model framework for ¹⁶O. Results and discussions are given in Sec. III, together with the EWSR of the isoscalar monopole transition. Finally, we present a summary in Sec. IV.

II. FORMULATION

First we formulate the isoscalar monopole strength function within the framework of the 4α OCM. Then, the formulation of the shell-model analysis is presented for ¹⁶O within the model space of 0s, 0p, 0d1s, and 0f 1p shells.

A. Monopole strength function

The strength function S(E) of the monopole excitation from the ¹⁶O ground state 0_1^+ is defined with use of the isoscalar monopole operator $\mathcal{O} = \sum_{i=1}^{16} (\mathbf{r}_i - \mathbf{R}_{c.m.})^2$ as follows:

$$S(E) = \sum_{n} \delta(E - E_n) |\langle 0_n^+| \sum_{i=1}^{16} (\mathbf{r}_i - \mathbf{R}_{\text{c.m.}})^2 |0_1^+\rangle|^2, \quad (1)$$

where \mathbf{r}_i ($i = 1 \sim 16$) are the coordinates of nucleons, $\mathbf{R}_{c.m.} = \frac{1}{16} \sum_{i=1}^{16} \mathbf{r}_i$ is the center-of-mass (c.m.) coordinate of ¹⁶O, and E_n denotes the excitation energy of the 0_n^+ state of ¹⁶O. However, the response function for the transition operator \mathcal{O} is defined as

$$R(E) = \langle 0_1^+ | \frac{\mathcal{O}^{\dagger} \mathcal{O}}{E - H + i\epsilon} | 0_1^+ \rangle, \qquad (2)$$

with ϵ representing an infinitesimal positive number. Then, R(E) is related to S(E) through

$$S(E) = -\frac{1}{\pi} \text{Im}[R(E)] = \sum_{n} \delta(E - E_n) |\langle 0_n^+ | \mathcal{O} | 0_1^+ \rangle|^2.$$
(3)

When the state $|0_n^+\rangle$ is a resonance state with the complex energy $E_n - i\Gamma_n/2$, the strength function is expressed as

$$S(E) = -\frac{1}{\pi} \text{Im}[R(E)]$$

= $\frac{1}{\pi} \sum_{n} \frac{\Gamma_n/2}{(E - E_n)^2 + (\Gamma_n/2)^2} |\mathcal{M}(0_n^+ - 0_1^+)|^2$, (4)

$$\mathcal{M}(0_n^+ - 0_1^+) = \langle 0_n^+ | \sum_{i=1}^{N} (\boldsymbol{r}_i - \boldsymbol{R}_{\text{c.m.}})^2 | 0_1^+ \rangle, \qquad (5)$$

where Γ_n represents the width of the 0_n^+ state. The isoscalar monopole transition matrix element, $\mathcal{M}(0_n^+ - 0_1^+)$, has a relation with the *E*0 transition matrix element $M(E0, 0_n^+ - 0_1^+)$ for the 0_1^+ and 0_n^+ states with the total isospin T = 0 as follows:

$$M(E0, 0_n^+ - 0_1^+) \equiv \langle 0_n^+ | \sum_{i=1}^{10} \frac{1 + \tau_{3i}}{2} (\mathbf{r}_i - \mathbf{R}_{\text{c.m.}})^2 | 0_1^+ \rangle$$
$$= \frac{1}{2} \mathcal{M}(0_n^+ - 0_1^+).$$
(6)

The EWSR of the isoscalar monopole transition [14] reads

$$\sum_{n} (E_n - E_1) |\mathcal{M}(0_n^+ - 0_1^+)|^2 = \frac{2\hbar^2}{m} \times 16 \times R^2, \quad (7)$$

$$R^{2} = \frac{1}{16} \langle 0_{1}^{+} | \sum_{i=1}^{16} (\boldsymbol{r}_{i} - \boldsymbol{R}_{\text{c.m.}})^{2} | 0_{1}^{+} \rangle, \qquad (8)$$

where R and m represent the rms radius of the ground state and nucleon mass, respectively. Here, we assume that the *NN* interaction has no velocity dependence. Employing the experimental charge radius of ¹⁶O ($R_c = 2.70$ fm [13]), the value of *R* in Eq. (8) is estimated to be 2.58 fm, in which the effects of the charge radius of proton ($\langle r^2 \rangle_{\text{proton}} = 0.8791^2$ fm²) and that of neutron ($\langle r^2 \rangle_{\text{neutron}} = -0.1149$ fm²) [13] are subtracted from the charge radius of ¹⁶O (R_c): $R = \sqrt{R_c^2 - \langle r^2 \rangle_{\text{proton}} - \langle r^2 \rangle_{\text{neutron}}} = 2.58$ fm. Then, the total EWSR value, $\frac{2\hbar^2}{m} \times 16 \times R^2$, is 8.83×10^3 fm⁴ MeV.

It is instructive to see a characteristic feature of the isoscalar monopole operator in Eq. (5), which can be decomposed into two parts, internal parts and relative parts, with respect to 4α clusters in ¹⁶O (as well as α and ¹²C clusters in ¹⁶O). Because the operator in Eq. (5) has a quadratic form with respect to the coordinates of nucleons, the following interesting identities are realized [14,16,22]:

$$\sum_{i=1}^{16} (\mathbf{r}_{i} - \mathbf{R}_{c.m.})^{2}$$

$$= \sum_{k=1}^{4} \sum_{i=1}^{4} (\mathbf{r}_{i+4(k-1)} - \mathbf{R}_{\alpha_{k}})^{2} + \sum_{k=1}^{4} 4(\mathbf{R}_{\alpha_{k}} - \mathbf{R}_{c.m.})^{2}, \quad (9)$$

$$= \sum_{k=1}^{4} (\mathbf{r}_{i} - \mathbf{R}_{k})^{2} + \sum_{k=1}^{16} (\mathbf{r}_{i} - \mathbf{R}_{k})^{2} + 3\boldsymbol{\xi}_{2}^{2} \quad (10)$$

$$= \sum_{i=1}^{\infty} (\mathbf{r}_i - \mathbf{R}_{\alpha})^2 + \sum_{i=5}^{\infty} (\mathbf{r}_i - \mathbf{R}_{\rm C})^2 + 3\boldsymbol{\xi}_3^2, \qquad (10)$$

where $\mathbf{R}_{\alpha_k} = (1/4) \sum_{i=1}^{4} \mathbf{r}_{i+4(k-1)}$ is the c.m. coordinate of the *k*th α cluster, and $\boldsymbol{\xi}_j$ (j = 1-3) are Jacobi coordinates with respect to the c.m. coordinates of 4α clusters $(\mathbf{R}_{\alpha_k}, k = 1-4)$: $\boldsymbol{\xi}_1 = \mathbf{R}_2 - \mathbf{R}_1, \boldsymbol{\xi}_2 = \mathbf{R}_3 - (\mathbf{R}_1 + \mathbf{R}_2)/2$, and $\boldsymbol{\xi}_3 = \mathbf{R}_4 - (\mathbf{R}_1 + \mathbf{R}_2 + \mathbf{R}_3)/3$. In Eq. (10), $\mathbf{R}_{\alpha} = (1/4) \sum_{i=1}^{4} \mathbf{r}_i$ and $\mathbf{R}_{\rm C} = (1/12) \sum_{i=5}^{16} \mathbf{r}_i$ stand for the c.m. coordinates of α and ¹²C clusters, respectively. Here we should recall the useful identity of $\sum_{k=1}^{4} 4(\mathbf{R}_{\alpha_k} - \mathbf{R}_{\rm c.m.})^2 = \sum_{j=1}^{3} \mu_k \boldsymbol{\xi}_j^2$ in Eq. (9), where μ_j ($\mu_1 = 2, \mu_2 = 8/3$, and $\mu_3 = 3$) correspond to the reduced masses with respect to the Jacobi coordinates $\boldsymbol{\xi}_j$. The fact that the isoscalar monopole operator consists of the two parts, the internal part and the relative part, plays an important role in the monopole excitation of ¹⁶O (see Sec. III).

B. 4α OCM

The total wave function $\tilde{\Psi}(J^{\pi})$ of the 4 α system with total angular momentum J^{π} in the OCM framework is expressed by the product of the internal wave functions of α clusters $\phi(\alpha)$ and the relative wave function $\Psi(J^{\pi})$ among the 4 α clusters,

$$\tilde{\Psi}(J^{\pi}) = \Psi(J^{\pi})\phi(\alpha_1)\phi(\alpha_2)\phi(\alpha_3)\phi(\alpha_4).$$
(11)

The relative wave function $\Psi(J^{\pi})$ is expanded in terms of Gaussian basis functions as follows:

$$\Psi(J^{\pi}) = \sum_{c,\nu} A_c(\nu) \Phi_c(\nu), \qquad (12)$$

$$\Phi_{c}(\nu) = \widehat{S}[[\varphi_{l_{1}}(\boldsymbol{\xi}_{1},\nu_{1})\varphi_{l_{2}}(\boldsymbol{\xi}_{2},\nu_{2})]_{l_{12}}\varphi_{l_{3}}(\boldsymbol{\xi}_{3},\nu_{3})]_{J}, \quad (13)$$
$$\langle u_{F}|\Psi(J^{\pi})\rangle = 0, \quad (14)$$

where $\boldsymbol{\xi}_1$, $\boldsymbol{\xi}_2$, and $\boldsymbol{\xi}_3$ are the Jacobi coordinates describing internal motions of the 4α system. \widehat{S} stands for the symmetrization operator acting on all α particles obeying

Bose statistics. ν denotes the set of size parameters ν_1 , ν_2 , and ν_3 of the normalized Gaussian function, $\varphi_l(\boldsymbol{\xi}, \nu_i) = N_{l,\nu_i} \boldsymbol{\xi}^l \exp(-\nu_i \boldsymbol{\xi}^2) Y_{lm}(\hat{\boldsymbol{\xi}})$, and *c* the set of relative orbital angular momentum channels $[[l_1, l_2]_{l_12}, l_3]_J$ depending on either of the coordinate type of *K* or *H* [20,25], where l_1 , l_2 , and l_3 are the orbital angular momenta with respect to the corresponding Jacobi coordinates. Equation (14) represents the orthogonality condition that the total wave function (12) should be orthogonal to the Pauli-forbidden states of the 4α system, u_F 's, which are constructed from Pauli forbidden states between two α particles in 0*S*, 0*D*, and 1*S* states [26]. The ground state with the dominant shell-model-like configuration $(0s)^4(0p)^{12}$ can be described properly in the present 4α OCM framework, as discussed below.

The 4α Hamiltonian for $\Psi(J^{\pi})$ is given as follows:

$$\mathcal{H} = \sum_{i}^{4} T_{i} - T_{c.m.} + \sum_{i < j}^{4} \left[V_{2\alpha}^{(N)}(i, j) + V_{2\alpha}^{(C)}(i, j) \right] + \sum_{i < j < k}^{4} V_{3\alpha}(i, j, k) + V_{4\alpha}(1, 2, 3, 4),$$
(15)

where T_i , $V_{2\alpha}^{(N)}(i, j)$, $V_{2\alpha}^{(C)}(i, j)$, $V_{3\alpha}(i, j, k)$, and $V_{4\alpha}(1, 2, 3, 4)$ stand for the operators of kinetic energy for the *i*th α particle, two-body, Coulomb, three-body, and four-body forces between α particles, respectively. The c.m. kinetic energy $T_{c.m.}$ is subtracted from the Hamiltonian. The effective α - α interaction $V_{2\alpha}^{(N)}$ is constructed by the folding procedure from an effective two-nucleon force. Here we take the modified Hasegawa-Nagata (MHN) force [27] as the effective NNforce, which is constructed based on the *G*-matrix theory. It is noted that the folded α - α potential reproduces the α - α scattering phase shifts and energies of the ⁸Be ground state and of the Hoyle state. The three-body force is phenomenologically introduced so as to fit the ground-state energy of the ¹²C with the framework of the 3 α OCM. The same force parameter set as used in Ref. [28] is adopted in the present calculation. For the later discussion (see Sec. III A) the energy spectra of ¹²C



FIG. 3. (Color online) Energy spectra of ¹²C with the 3α OCM [28] together with the experimental one [13]. The experimental energy level of the 2^+_2 state is cited from Ref. [29].

with the 3α OCM calculation is shown in Fig. 3 together with the experimental one. It is known that the 3α OCM reasonably describes the structure of the $0^+_{1,2}$, $2^+_{1,2}$, 4^+_1 , 3^-_1 , and 1^-_1 states in ¹²C [28]. In addition, the phenomenological four-body force [30] is adjusted to the ground-state energy of ¹⁶O. The three-body and four-body forces are short range, and, hence, they only act in compact configurations. The coefficients $A_c(\nu)$ in Eq. (12) are determined according to the Rayleigh-Ritz variational principle.

The isoscalar monopole matrix element is evaluated as follows:

$$\mathcal{M}^{\text{OCM}}(0_{n}^{+}-0_{1}^{+}) = \langle \tilde{\Psi}(0_{n}^{+}) | \sum_{i=1}^{16} (\boldsymbol{r}_{i} - \boldsymbol{R}_{\text{c.m.}})^{2} | \tilde{\Psi}(0_{1}^{+}) \rangle, \qquad (16)$$

$$= \langle \Psi(0_n^+) | \sum_{k=1}^{4} 4 (\mathbf{R}_{\alpha_k} - \mathbf{R}_{\text{c.m.}})^2 | \Psi(0_1^+) \rangle + 16 \times R(\alpha)^2 \delta_{n1},$$
(17)

$$= \langle \Psi(0_n^+) | 2\boldsymbol{\xi}_1^2 + \frac{8}{3}\boldsymbol{\xi}_2^2 + 3\boldsymbol{\xi}_3^2 | \Psi(0_1^+) \rangle + 16 \times R(\alpha)^2 \delta_{n1},$$
(18)

where Ψ is the 4 α OCM wave function in Eq. (12). In Eqs. (16)–(18) we used the relation

$$\langle \phi(\alpha_1)\phi(\alpha_2)\phi(\alpha_3)\phi(\alpha_4)| \sum_{k=1}^{4} \sum_{i=1}^{4} \left(\boldsymbol{r}_{i+4(k-1)} - \boldsymbol{R}_{\alpha_k} \right)^2 \\ \times |\phi(\alpha_1)\phi(\alpha_2)\phi(\alpha_3)\phi(\alpha_4)\rangle = 16 \times R(\alpha)^2,$$
(19)

where $R(\alpha)$ is the rms radius of α particle, $R(\alpha) = \sqrt{\frac{1}{4} \langle \sum_{i=1}^{4} (\mathbf{r}_i - \mathbf{R}_{\alpha})^2 \rangle_{\alpha}}$ [20,28]. It is important to study the EWSR of the isoscalar monopole transition within the framework of the 4 α OCM. We call it the OCM-EWSR, and its definition reads

$$\sum_{n} (E_{n} - E_{1}) |\mathcal{M}^{\text{OCM}}(0_{n}^{+} - 0_{1}^{+})|^{2}$$
$$= \frac{1}{2} \langle \Psi(0_{1}^{+}) | [\mathcal{O}_{\text{OCM}}, [\mathcal{H}, \mathcal{O}_{\text{OCM}}]] |\Psi(0_{1}^{+}) \rangle, \qquad (20)$$

$$=\frac{2\hbar^2}{m}\langle\Psi(0_1^+)|\sum_{k=1}^4 4(\boldsymbol{R}_{\alpha_k}-\boldsymbol{R}_{\rm c.m.})^2|\Psi(0_1^+)\rangle,\quad(21)$$

$$=\frac{2\hbar^2}{m}\times 16\times (R^2-R(\alpha)^2), \qquad (22)$$

where \mathcal{H} is given in Eq. (15) and $\mathcal{O}_{\text{OCM}} = \sum_{k=1}^{4} 4(\mathbf{R}_{\alpha_k} - \mathbf{R}_{\text{c.m.}})^2$ [see Eq. (17)], and R denotes the rms radius of the ¹⁶O ground state given in Eq. (8). It is noted that the 4α OCM can describe the shell-model-like structure of the ¹⁶O ground state, as shown later. Then, the ratio of the OCM-EWSR to the total EWSR in Eq. (7) is

$$\frac{\text{OCM-EWSR}}{\text{total EWSR}} = 1 - \left(\frac{R(\alpha)}{R}\right)^2 = 1 - \left(\frac{1.47}{2.58}\right)^2 = 0.68.$$
(23)

Here we use $R(\alpha) = 1.47$ fm and R = 2.58 fm, which are estimated from the experimental charge radii (1.68 and 2.70 fm, respectively [13]), subtracting the effects of the charge radius of proton and that of neutron from them, the method of which is the same as that shown in previous section. This result means that the 4 α OCM framework shares about 70% of the total EWSR value (the OCM-EWSR is also discussed in the Appendix). This is one of the important reasons that the 4 α OCM works rather well in reproducing the isoscalar monopole transitions in the low-energy region of ¹⁶O, as shown later.

In the present paper, the energies E_n and isoscalar monopole matrix elements \mathcal{M} in Eq. (4) are obtained by the 4α OCM calculation. As for the widths Γ_n , we estimate the α -decay widths with the *R*-matrix theory [31],

$$\Gamma_L = 2P_L(a)\gamma_L^2(a), \qquad (24)$$

$$P_L(a) = \frac{\kappa a}{F_L^2(ka) + G_L^2(ka)},$$
(25)

$$\gamma_L^2(a) = \theta_L^2(a) \gamma_W^2(a), \qquad (26)$$

$$\gamma_{\rm W}^2(a) = \frac{3\hbar^2}{2\mu a^2},$$
 (27)

where k, a, and μ are the wave number of the α -¹²C relative motion, the channel radius, and the reduced mass, respectively, and F_L , G_L , and $P_L(a)$ are the regular and irregular Coulomb wave functions and the corresponding penetration factor, respectively. The reduced width of $\theta_L^2(a)$ is related to the reduced width amplitude or overlap amplitude \mathcal{Y}_L as $\theta_L^2(a) = \frac{a^3}{3} \mathcal{Y}_L^2(a)$, and the definition of \mathcal{Y}_L is presented as

$$\mathcal{Y}_{L}(r) = \sqrt{\frac{4!}{3!1!}} \left\langle \left[\frac{\delta(r'-r)}{r'^{2}} Y_{L}(\hat{r}') \Phi_{L}(^{12}\mathrm{C}) \right]_{0} \middle| \Psi(0_{n}^{+}) \right\rangle.$$
(28)

Here, $\Phi_L({}^{12}C)$ is the wave function of ${}^{12}C$, given by the 3α OCM calculation [28], and *r* is the relative distance between the c.m. of ${}^{12}C$ and the α particle. The spectroscopic factor of the $\alpha + {}^{12}C(L)$ channel S_L^2 in the 0_n^+ state of ${}^{16}O$, defined as

$$S_L^2 = \int_0^\infty dr \left[r \mathcal{Y}_L(r) \right]^2, \qquad (29)$$

is useful to analyze the obtained wave functions.

In the present study, we perform more careful analyses than the previous ones [20], in particular, for identifying the 0^+ states around the 4α threshold. The calculation of the resonant state in the bound-state approximation is usually done by diagonalizing the Hamiltonian with the use of a finite number of square-integrable basis wave functions. The positive-energy eigenstates obtained by the diagonalization are divided into resonant states and continuum states, and many methods for carrying out the division are proposed [32]. In the present study, a pseudopotential method is adopted to divide the resonant states and continuum states, as shown below.

Let us us first consider a repulsive pseudopotential V that is added to the original Hamiltonian H, yielding

$$H'(\delta) = H + \delta \times V, \tag{30}$$

where δ is a constant used to vary the strength of the pseudopotential. As increasing into negative values the constant δ

from the physical value, $\delta = 0$, the eigenenergy of this new Hamiltonian $H'(\delta)$ decreases for any resonance state, which is eventually transformed into a bound state. On the contrary, continuum states show almost no change in their eigenvalues as δ increases into the negative region. In the present 4α OCM framework, it is important to study the eigenenergies with changing the constant δ but with no change in the threshold energies of the $\alpha + {}^{12}C$ and 4α decay channels, even though we introduce the pseudopotential V. Here, we take the four-body potential $V_{4\alpha}$ in Eq. (15) as the pseudopotential V, because the choice is convenient for practical reasons in the present numerical calculation. This pseudopotential method is simple but helpful to identify the resonant states under the bound-state approximation. As a result, we obtained almost the same results as the previous ones [20], as is shown below.

C. Shell-model calculation

The shell-model Hamiltonian of ¹⁶O adopted here is presented as follows:

$$H = \sum_{i=1}^{16} t_i - T_{\text{c.m.}} + \sum_{i< j=1}^{16} [v^{(\text{C})}(i, j) + v^{(\text{LS})}(i, j)], \quad (31)$$

where t_i denotes the kinetic energy of the *i*th nucleon and $v^{(C)}$ ($v^{(LS)}$) represents the central (*LS*) force of the effective *NN* interaction. The c.m. kinetic energy $T_{c.m.}$ is subtracted from the total kinetic energy. The model space adopted here covers all configurations of 1p1h and 2p2h within the 0s, 0p, 0d1s, and 0f1p shells. The spurious states of the c.m. motion are eliminated with the Lawson's method [33].

In the present study, we take the Volkov No. 2 force [34] and G3RS force [35] for $v^{(C)}$ and $v^{(LS)}$, respectively. The Majorana parameter (*M*) in the Volkov No. 2 force, the multiplying factor (ζ_0^{LS}) of the G3RS, and the nucleon size parameter (*b*) are chosen so as to reproduce as well as possible the total binding energies of the ground states of ¹⁶O and ¹⁵O, the *LS* splitting between $3/2_1^-$ and $1/2_1^-$ in ¹⁵O, and the rms radius of the ground state of ¹⁶O. The following two parameter sets are adopted: case A for (M, ζ_0^{LS} , b) = (0.665, 1.7, 1.70 fm) and case B for (0.620, 2.6, 2.0 fm). The present shell-model code is based on the code in Refs. [36,37].

III. RESULTS AND DISCUSSION

A. 4a OCM calculation

The energy levels of 0^+ states in ¹⁶O obtained by the present 4α OCM calculation are shown in Fig. 4 and Table I. One can make the one-to-one correspondence of the six lowest 0^+ states observed up to $E_x \simeq 16$ MeV in the 4α OCM calculation. The reader is reminded that the $\alpha + {}^{12}C$ OCM cluster model [16] can reproduce only the lowest three 0^+ states. We obtained almost the same results as the previous 4α OCM calculation [20]. The six 0^+ states have the following characteristic structures [20]: (1) the ground state (0^+_1) has dominantly a doubly closed-shell structure; (2) the 0^+_2 state at $E_x = 6.05$ MeV and the 0^+_3 state at $E_x = 12.05$ MeV have mainly $\alpha + {}^{12}C$ structures [39], where the α particle



FIG. 4. (Color online) Comparison of energy spectra among experiment, the 4α OCM calculation [20], and the $\alpha + {}^{12}C$ model calculation [16], where the $\alpha + {}^{12}C$ and 4α thresholds are shown. Experimental data are taken from Ref. [13] and from Ref. [38] for the 0_4^+ state.

orbits around the ¹²C(0⁺₁) core in an *S* wave and around the ¹²C(2⁺₁) core in a *D* wave, respectively, the results of which are consistent with the previous studies with α + ¹²C OCM [16] and the α + ¹²C generator coordinate method (GCM) [19]; (3) the 0⁺₄ (E_x = 13.6 MeV) and 0⁺₅ (E_x = 14.1 MeV) states mainly have α + ¹²C(0⁺₁) structure with higher nodal behavior and α + ¹²C(1⁻) structure, respectively, where in the latter the α particle moves around the ¹²C(1⁻) core (corresponding to the first 1⁻ state at E_x = 10.84 MeV having an intermediate structure between the shell-model-like structure and cluster structure [18]) in *P* orbit; and (4) the 0⁺₆ state at 15.1 MeV is a strong candidate of the 4 α condensate, (0*S*)⁴_{α}, with the probability of 61%.

These characteristic features of the structures of the six 0^+ states can be verified from the analysis of the spectoscopic factors S_L^2 defined in Eq. (29). The results are shown in Fig. 5. Because the ground state has a closed shell structure with the dominant component of SU(3)(λ , μ) = (0, 0) [40], the values of the spectroscopic factors S_L^2 for 0_1^+ in Fig. 5 can be explained by the SU(3) nature of the state. This SU(3) character was confirmed by the recent no-core shell model [41]. As mentioned above, the structures of the 0^+_2 and 0_3^+ states are well established as having the $\alpha + {}^{12}C(0_1^+)$ and $\alpha + {}^{12}C(2^+_1)$ cluster structures, respectively. These structures of the 0^+_2 and 0^+_3 states are confirmed by the 4α OCM calculation. In fact, one sees in Fig. 5 that the S^2 factors for the $\alpha + {}^{12}C(0^+_1)$ and $\alpha + {}^{12}C(2^+_1)$ channels are dominant in the 0_2^+ and 0_3^+ states, respectively. In the 0_3^+ state, however, the S² factors of the $\alpha + {}^{12}C(4^+_1)$ channel in the 4 α OCM calculation are rather large as compared with the result of the $\alpha + {}^{12}C$ OCM one [16]. This is attributable to the following facts. (1) The calculated excitation energy of the ${}^{12}C(4_1^+)$ state in the present 4α OCM calculation is underestimated by $E_x \sim 6 \text{ MeV}$ (see Fig. 3), while it is set to the experimental value ($E_x = 14.1 \text{ MeV}$) in the $\alpha + {}^{12}\text{C}$ OCM calculation; and (2) thus, in the 4α OCM calculation a coupled-channel

TABLE I. Excitation energies (E_x) , charge rms radii (R_c) , *E*0 transition matrix elements [M(E0)], and particle decay widths (Γ) of the 0⁺ states in ¹⁶O obtained by the 4 α OCM calculation and α + ¹²C OCM model calculation [16], together with the experimental data [13,38]. They are given in the unit of MeV, fm, fm², and MeV, respectively. The experimental monopole matrix elements are obtained by the ¹⁶O(*e*, *e'*) reaction [13]. *P*^{e.w.} represents the percentage of the energy weight strength to the isoscalar monopole EWSR [see Eq. (7)]. The finite size effects of α particle and ¹²C are taken into account in estimating R_c with the 4 α OCM and α + ¹²C OCM (see Ref. [28] for details).

	4α OCM				$\alpha + {}^{12}C \text{ OCM}$			Experiment				
	E_x	R_c	M(E0)	Г	E_x	R_c	M(E0)	E_x	R_c	M(E0)	P ^{e.w.} (%)	Г
0^{+}_{1}	0.00	2.7			0.00	2.5		0.00	2.70			
0_{2}^{+}	6.37	3.0	3.9		6.57	2.9	3.88	6.05		3.55 ± 0.21	3.5	
0^{+}_{3}	9.96	3.1	2.4		10.77	2.8	3.50	12.05		4.03 ± 0.09	8.9	
0_{4}^{+}	12.56	4.0	2.4	0.60	_	-	_	13.60		No data		0.6
0_{5}^{+}	14.12	3.1	2.6	0.20	_	_	_	14.01		3.3 ± 0.7	6.9	0.185
0_{6}^{+}	16.45	5.6	1.0	0.14	_	_	-	15.10		No data		0.166

effect of the $\alpha + {}^{12}C(4^+_1)$ channel with the $\alpha + {}^{12}C(0^+_1, 2^+_1)$ channel is reinforced and, consequently, the S^2 factor of the $\alpha + {}^{12}C(4^+_1)$ channel becomes larger. We expect that the S^2 factor of the $\alpha + {}^{12}C(4^+_1)$ channel will be smaller when the excitation energy of ${}^{12}C(4^+_1)$ is properly reproduced in the 4α OCM calculation.

Table I lists the E0 transition matrix elements M(E0). The M(E0) value of the 0_2^+ state is reproduced well, while that of the 0_3^+ state underestimates the experimental result, and this trend is similar to the result of the $\alpha + {}^{12}C$ OCM model [16]. However, the 0_4^+ and 0_5^+ states mainly have the $\alpha + {}^{12}C(0_1^+)$ structure with higher nodal behavior and an $\alpha + {}^{12}C(1^-)$ structure, respectively. The E0 transition matrix element of the 0_5^+ state is reproduced nicely within the experimental





error (see Table I). In Table I, the largest rms radius is about 5 fm for the 0_6^+ state, the wave function of which has a large overlap amplitude with the $\alpha + {}^{12}C(0_2^+)$ channel (see Fig. 3 in Ref. [20]). Hence, the S^2 factor of the $\alpha + {}^{12}C(0_2^+)$ channel is dominant in the 0_6^+ state (see Fig. 5), whereas those in the other channels are much suppressed. This dominance of the S^2 factor of the $\alpha + {}^{12}C(0_2^+)$ channel is one of the evidences for the 0_6^+ state being the 4α -condensate state, $(0S)_{\alpha}^4$, because the Hoyle state has the main configuration, $(0S)_{\alpha}^3$ [28,42,43], and the overlap amplitude of $\langle (0S)_{\alpha}^3 | (0S)_{\alpha}^4 \rangle$ becomes large.

As for the decay widths of the 0_4^+ and 0_5^+ states, the results are shown in Table I. The calculated width of the 0_4^+ state is ~600 keV, which is quite a bit larger than that found for the 0_5^+ state ~200 keV. Both are quantitatively consistent with the corresponding experimental data, 600 and 185 keV, respectively. However, the decay width of the 0_6^+ state is very small, 140 keV, in reasonable agreement with the corresponding experimental value of 166 keV, indicating that this state is unusually long lived. We should note that our calculation consistently reproduces the ratio of the widths of the 0_4^+ , 0_5^+ , and 0_6^+ states, that is, about 6 : 3 : 2, respectively (see Table I).

Comparing the energy levels of the six 0^+ states with the experimental monopole response function of ¹⁶O shown in Fig. 1, one notices that the energy positions of the fine structures in the low-energy region ($10 \leq E_x \leq 16$ MeV) of the experimental response function seem to be in good correspondence with the energy levels of 0^+_3 , 0^+_4 , 0^+_5 , and 0^+_6 in Fig. 4 and Table I. It should be noted that the peak corresponding to the 0^+_2 state at $E_x = 6.05$ MeV is not visible in Fig. 1, because of an energy cut in the experimental condition [12]. Thus, it is important to study the isoscalar monopole strength function within the framework of the 4α OCM calculation.

B. Isoscalar monopole excitation function with the 4α OCM calculation

Figure 6 shows the calculated isoscalar monopole strength function of 16 O defined in Eq. (4), where we use the calculated



FIG. 6. (Color online) Calculated isoscalar monopole strength functions of ¹⁶O (bold line) and experimental data (thin line; see Fig. 1 [12]). Here we use the calculated monopole matrix elements and the calculated decay widths for the six 0⁺ states up to $E_x \simeq 16$ MeV obtained by the 4α OCM calculation (see Table I), although the experimental excitation energies for the six 0⁺ states are employed (see text).

monopole matrix elements and the calculated decay widths for the six 0^+ states up to $E_x \simeq 16$ MeV obtained by the 4α OCM calculation (see Table I); also the experimental excitation energies for the six 0^+ states are employed. We take into account the experimental energy resolution of 50 keV [12] for the width Γ_n in Eq. (4) through $\Gamma_n = \sqrt{\Gamma_n (\text{OCM})^2 + 0.050^2}$, where $\Gamma_n(OCM)$ denotes the calculated decay width of the nth 0⁺ state of ¹⁶O given in Table I. The calculated strength function is normalized so as to match the calculated strength of the 12.1-MeV peak to the experimental one. We can see a rather good correspondence with the experimental data. The fine structures in the calculated strength function, that is, one peak at $E_x = 12.1$ MeV (corresponding to the 0^+_3 state), one shoulderlike peak at $E_x = 13.8 \text{ MeV} (0_4^+)$, and two peaks at $E_x = 14.1 \text{ MeV}(0_5^+) \text{ and } 15.1 \text{ MeV}(0_6^+), \text{ are well reproduced.}$ As mentioned above, the fine structures in the energy region of $10 \leq E_x \leq 15$ MeV as well as the sharp peak at $E_x \simeq 6$ MeV (corresponding to the 0^+_2 state) are difficult to reproduce by any mean-field calculations [4,6-10,15], as far as the present authors know. The calculated values of S(E) at each 0^+ state $(0^+_2 - 0^+_6)$ in Fig. 6 are approximately proportional to the squared values of the respective calculated monopole matrix elements.

It is instructive and interesting to discuss the mechanism of why the five α cluster states $(0_2^+, 0_3^+, 0_4^+, 0_5^+, \text{ and } 0_6^+)$ of ¹⁶O are excited relatively strongly from the ground state with the doubly closed-shell-like structure [14]. Their monopole matrix elements shown in Table I are comparable to the single-particle strength (~5.4 fm² [14]) and share about 20% of the total EWSR value. Because the mechanism is closely related to the property of the ground state of ¹⁶O as shown below, we first demonstrate its interesting properties with the use of the microscopic wave function and then discuss the monopole matrix elements in the OCM calculation. The wave function of the ¹⁶O ground state has dominantly the doubly closed-shell-model configuration $(0s)^4(0p)^{12}$ with the nucleon size parameter $v = M\omega/2\hbar$ (*M*, nucleon mass), corresponding to the SU(3) (λ , μ) = (0, 0) wave function [40]. This doubly closed-shell-model wave function is mathematically equivalent to a single-cluster-model wave function of $\alpha + {}^{12}C$ as well as 4α with the total harmonic oscillator quanta Q = 12 [14,16,22],

$$\frac{1}{\sqrt{16!}} \det|(0s)^4 (0p)^{12}| \times [\phi_{\rm c.m.}(\boldsymbol{R}_{\rm c.m.})]^{-1}$$
(32)

$$= N_0 \sqrt{\frac{12!4!}{16!}} \mathcal{A}\{[u_{40}(\boldsymbol{\xi}_3, 3\nu)\phi_{L=0}(^{12}\mathrm{C})]_{J=0}\phi(\alpha)\}, \quad (33)$$

$$= N_2 \sqrt{\frac{12!4!}{16!}} \mathcal{A}\{[u_{42}(\boldsymbol{\xi}_3, 3\nu)\phi_{L=2}(^{12}\mathrm{C})]_{J=0}\phi(\alpha)\}, \quad (34)$$

$$= \hat{N}_{0} \sqrt{\frac{4!4!4!}{16!}} \mathcal{A} \left\{ \left[u_{40}(\boldsymbol{\xi}_{3}, 3\nu) \left[u_{40} \left(\boldsymbol{\xi}_{2}, \frac{8}{3}\nu \right) \right. \right. \right. \\ \left. \times u_{40}(\boldsymbol{\xi}_{1}, 2\nu) \right]_{L=0} \right]_{J=0} \phi(\alpha_{1}) \phi(\alpha_{2}) \phi(\alpha_{3}) \phi(\alpha_{4}) \left. \right\}, \quad (35)$$

$$\phi_{\text{c.m.}}(\boldsymbol{R}_{\text{c.m.}}) = \left(\frac{32\nu}{\pi}\right)^{3/4} \exp(-16\nu \boldsymbol{R}_{\text{c.m.}}^2),$$
 (36)

where $\phi_{c.m.}$ denotes the wave function of the c.m. motion of ¹⁶O, and $N_{0,2}$ (\hat{N}_0) are the normalization constants. $\phi(\alpha)$ and $\phi_L(^{12}C)$ represent, respectively, the internal wave function of the α cluster with the $(0s)^4$ configuration and that of ${}^{12}C$ with the angular momentum of L belonging to the SU(3) irreducible representation $(\lambda, \mu) = (0, 4)$ for the $(0s)^4 (0p)^8$ configuration. The relative wave functions between the α and ¹²C clusters in Eqs. (33) and (34) are described by the harmonic oscillator wave function $u_{QLm}(\boldsymbol{\xi}, \beta) = u_{QL}(\boldsymbol{\xi}, \beta)Y_{Lm}(\boldsymbol{\hat{\xi}})$ with the node number n = (Q - L)/2 and Q = 4. One can prove Eqs. (33) and (34) with help of the Bayman-Bohr theorem [44]. The reader should be reminded that the dominance of the doubly closed-shell structure in the ground state of ¹⁶O is confirmed by the no-core shell model with realistic NN forces [41]. Equations (33) and (34) mean that the doubly closed-shellmodel wave function has an α + ¹²C cluster degree of freedom. In addition, Eq. (35) demonstrates that the doubly closed-shellmodel wave function also possesses a 4α cluster degree of freedom.

In the 4α OCM, the monopole matrix elements from the ¹⁶O ground state to the $\alpha + {}^{12}C$ cluster states are evaluated with use of Eqs. (17) and (18). The validity of using the formulas is based on the following three facts found within the microscopic framework [14,16,22]: (1) The ground state of ¹⁶O is of the SU(3) $(\lambda, \mu) = (0, 0)$ nature with the α -clustering degree of freedom; (2) the monopole matrix elements come dominantly from the relative part of the monopole operator referring to the $\alpha + {}^{12}C$ relative motion, $3\xi_3^2$; (3) the contribution from the other parts of the monopole operator becomes significantly smaller for the $\alpha + {}^{12}C$ cluster states; and (4) the α -clustertype ground-state correlation significantly enhances the monopole strength compared with the case of the ¹⁶O ground state being the pure SU(3) $(\lambda, \mu) = (0, 0)$ wave function. These are certainly the reasons why the estimation of the monopole matrix element using Eqs. (17) and (18) in the 4α OCM calculation gives a reasonable reproduction for the experimental data. In fact, the ground state of ¹⁶O obtained in the 4α OCM has the dominant SU(3) (λ , μ) = (0, 0) component with Q = 12, and only the relative part with respect to the $\alpha + {}^{12}C$ relative motion, $3\xi_3{}^2$, in the monopole matrix element (18) gives the major contribution to the monopole matrix elements $M(E0; 0^+_1 - 0^+_{2,3})$ in the present 4α OCM calculation.

However, the reason why the 0_6^+ state with the 4α -gas-like character has a relatively large monopole strength ($\sim 1 \text{ fm}^2$) can be also understood from the property of the ground state of 16 O. The doubly closed-shell-model wave function in Eq. (32) is mathematically equivalent to the single 4α cluster wave function with Q = 12 in Eq. (35). This equation means that the ground state of ¹⁶O with the $(0s)^4(0p)^{12}$ configuration inherently has a 4α -cluster degree of freedom. The relative part (or second term) of the monopole operator in Eq. (9), $\sum_{k=1}^{4} 4(\mathbf{R}_{\alpha_k} - \mathbf{R}_{c.m.})^2 = \sum_{k=1}^{3} \mu_k \boldsymbol{\xi}_k^2$, can excite the relative motion among the 4α particles. In other words, the monopole operator has an ability to populate democratically 4α particles by $2\hbar\omega$ with respect to the c.m. coordinate of ¹⁶O. The resultant state, thus, has some amount of the overlap with the 4α -gaslike state, that is, 0_6^+ , with the 4 α -condensate-like structure [20,21]. The overlap value corresponds certainly to the monopole matrix element, M(E0). As shown in Ref. [45], this 0_6^+ state can well be described by a 4 α -condensate-type microscopic wave function, called the THSR wave function [42]. In this THSR framework, the monopole matrix element to the 4 α -condensate-like state is estimated to be M(E0) = 1.2fm², similar to that in the 4 α OCM, $M(E0; 0_1^+ - 0_6^+) = 1.0$ fm^2 in Table I, which is calculated with the use of Eqs. (17) and (18). Thus, the evaluation of the monopole matrix elements using Eqs. (17) and (18) in the OCM framework is useful and gives a reasonable estimate for the monopole transition to the α -¹²C cluster states and 4 α -gas-like states. It is noted that the mechanism of the 4α -gas-like state being populated by the monopole transition is similar to that of the Hoyle state with the 3α -gas-like structure being excited by the monopole transition, although the ground state of ¹²C has a shell-model-like compact structure with the main configuration of SU(3) $(\lambda, \mu) = (0, 4)$. The detailed discussion is given in Ref. [14].

As for the 0_5^+ state, its main configuration is $\alpha + {}^{12}C(1_1^-)$ with the *P*-wave orbiting of an α cluster around the ${}^{12}C(1_1^-)$ core, as mentioned above. According to the Bayman-Bohr theorem, the SU(3) $(\lambda, \mu) = (0, 0)$ state of ¹⁶O has no component of the $\alpha + {}^{12}C(1_1^-)$ channel. However, the monopole strength to the 0_5^+ state is as large as 3 fm² (see Table I). This is the reason that the 0_5^+ state has small but important components of the $\alpha + {}^{12}C(0^+_1, 2^+_1)$ and $\alpha + {}^{12}C(0^+_2)$ configurations, as one can see from the spectroscopic factors shown in Fig. 5. It is noted that the $\alpha + {}^{12}C(0^+_2)$ configuration is likely to be an alternative of the 4α -gas-like state with the dominant $(0S)^4_{\alpha}$ configuration. Because these three configurations, $\alpha + {}^{12}C(0^+_1, 2^+_1)$ and $(0S)^4_{\alpha}$, can be excited from the ground state of ¹⁶O by the monopole operator as discussed above, their respective contributions are coherently added to provide the relatively large monopole strength to the 0_5^+ state. However, the situation of the 0_4^+ state, characterized mainly by the higher nodal $\alpha + {}^{12}C(0^+_1)$

state, is similar to the case of the 0_5^+ state. From Fig. 5, the 0_4^+ state has also small but non-negligible components of the $\alpha + {}^{12}C(0_1^+, 2_1^+)$ and $\alpha + {}^{12}C(0_2^+)$ configurations, which contribute to the monopole strength for the 0_4^+ state.

Finally, it is remarked that the calculated strength function in Fig. 6 takes into account only the contributions from the resonant states $(0_2^+, \ldots, 0_6^+)$. The continuum states of the $\alpha + {}^{12}C$ channel ($E_x = 7.16$ MeV) and 4α disintegrated channel ($E_x = 14.44$ MeV) can contribute to the strength. It is interesting to study the effect of the continuum states to the monopole strength function up to $E_x \simeq 16$ MeV, although the contribution from the 4α continuum states may be small in the energy region up to $E_x \simeq 16$ MeV because of their small phase space.

C. Shell-model calculation

In the shell-model calculations using the spherical basis, formulated in Sec. II C, the total energy of the ground state of ¹⁶O is well reproduced in cases A and B: $E^{cal}(^{16}O) = -127.57$ MeV and -126.79 MeV, respectively, for cases A and B vs $E^{exp}(^{16}O) = -127.62$ MeV. Case A gives the rms radius of ¹⁶O, 2.50 fm, the value of which is in good correspondence with the experimental one (2.58 fm), while the case B does the larger radius (2.94 fm). In addition, the energies of the $1/2_1^-$ and $3/2_1^-$ states in ¹⁵O are reasonably reproduced in the cases A and B: $E^{cal}(1/2_1^-) = -109.27$ MeV and $E^{cal}(3/2_1^-) = -103.18$ MeV (-110.05 and -103.77) for the case A (B) vs $E^{exp}(1/2_1^-) = -111.96$ MeV and $E^{exp}(3/2_1^-) = -105.78$ MeV.

We found that the excitation energy of the 0^+_2 state in ${}^{16}O$ is as large as 30 MeV in the present shell-model calculation, while the experimental value is as small as 6.7 MeV, the value of which is reproduced by the cluster model (see Table I). It is noted that the present shell-model space includes only the 1p1h and 2p2h configurations up to the 0f1p shell. Because the 1p1h configurations, in particular, $(1s_{1/2})(0s_{1/2})^{-1}$, $(1p_{3/2})(0p_{3/2})^{-1}$, and $(1p_{1/2})(0p_{1/2})^{-1}$, should give a crucial contribution to the monopole transition strengths, it is instructive to investigate how strongly the components of the three 1p1h configurations, P(1p1h), distribute among the various 0^+ states. We found three energy regions in which the components P(1p1h) become significantly large: (1) $E_x \sim 32$ MeV with $P(1p1h) \sim 81\%$, contributed by the 0_2^+ state; (2) $E_x \sim 48$ MeV with $P(1p1h) \sim 65\%$; and (3) $E_x \sim 60$ MeV with $P(1p1h) \sim 40\%$. This split of the 1p1h component into the three energy regions is consistent with the RRPA result [10,12], in which the monopole strength concentrates mainly into three energy regions (see Fig. 1). This is in line with the RPA and QRPA calculations (Fig. 3 in Ref. [15]), although the excitation energies of the three energy regions are different. The fact that the excitation energies of the three energy regions in the present shell-model calculations are higher than those in the RRPA, RPA, and QRPA calculations is reasonable because our calculation employs the spherical harmonic oscillator basis, and the model space taken in its calculation covers only the 1p1h and 2p2h configurations within the 0s, 0p, 0d1s, and 0f1p shells (see Sec. II C).

The fact that the present spherical shell-model calculation has great difficulty in reproducing the low excitation energies of the 0_2^+ and 0_3^+ states is likely to be a common feature of the no-core shell-model calculations [41], the FMD calculations [7–9], and the coupled-cluster calculations [46]. Exceptions are a few conventional shell-model works, for example, by Brown and Green [47] in 1966 and Arima *et al.* [48] in 1967, as far as the present authors know. Here it is instructive to briefly present their main results.

Brown and Green discussed the low-lying three 0^+ states of ¹⁶O with the deformed-shell model [47]. It is proposed that the three 0^+ states $(0_1^+, 0_2^+, \text{ and } 0_3^+)$ can be described by the mixture among the 0p0h, 2p2h, and 4p4h states. In their calculation, the unperturbed energies of the 0p0h, 2p2h, and 4p4h states are treated as free parameters adjusted to give the observed spectra, although the coupling strengths among the 0p0h, 2p2h, and 4p4h states are estimated with some approximations based on SU(3) algebra [49]. Then they found that the 0_1^+ state has a dominant configuration of 0p0h type, while the main configuration of the $0^+_2(0^+_3)$ state is of the 4p4h type (2p2h type). With use of the Brown-Green wave functions, Bertsch [50] calculated the E0 transition strength between the 0_2^+ and 0_1^+ states, $M(E0; 0_2^+ - 0_1^+)$, defined in Eq. (6). His result is M(E0) = 1.6-2.8 fm², corresponding to the experimental data (3.8 fm^2) , although the monopole transition strength between the 0_3^+ and 0_1^+ states was not discussed in that paper. The important point in the Brown-Green calculation is that the unperturbed energy of the 4p4h configuration is taken to be lower than that of the 2p2h one. It is found that if the 2p2h state lies lower than the 4p4h state, the calculated B(E2)transition rates between the resulting levels for the 0^+ states and 2^+ states are difficult to reconcile with the experimental data, although they could not present the reason why the unperturbed energy of the 4p4h configuration becomes lower than that of the 2p2h one. This schematic model proposed by Brown and Green was confirmed by the large-basis spherical shell-model calculations mixing the $(0 + 2 + 4)\hbar\omega$ excitations [51,52]. Although they succeed in reproducing the low-lying spectrum of ¹⁶O, the single-particle energies are adjusted to fit six low-lying T = 0 states in ¹⁶O, including the 0^+_2 and 0^+_3 states [51]. Thus, the problem of why the excitation energy of the 0^+_2 state is as small as 6.05 MeV remains unclear in those shell-model calculations.

After Brown-Green's work, Arima et al. proposed a weak coupling picture [48] and showed that one can understand the appearance of the low-lying 0^+_2 and 0^+_3 states in ¹⁶O if one assumes weak coupling between four particles (4p) in the sd shell and four holes (4h) in the p shell. They estimated the coupling strength with a shell model by employing the experimental excitation energy ($E_x = 0.86 \text{ MeV}$) of the $1/2_1^$ state in 19 F which is described by the 4p in sd shell and 1h in the p shell. They eventually found that the coupling strength between 4p and 4h in ¹⁶O becomes significantly weak. This weakness of the 4p and 4h interactions is nothing but the basic assumption of an α -cluster model in ¹⁶O; that is, one can easily imagine that the 4p (4h) state corresponds to the α (¹²C) cluster. The 4p in the sd shell can obtain an extraordinarily large binding energy owing to the α -cluster correlation, and thus the states consisting of 4p (in the sd shell) and 4h (in the

p shell) can be expected to lie at lower excitation energies, compared with the 2p2h states.

A structure study of ¹⁶O which explicitly treats the α -cluster degree of freedom was performed by Suzuki in 1976 with the semi-microscopic cluster model, $\alpha + {}^{12}C$ OCM [16]. In this model the relative motion between the α and ${}^{12}C$ clusters is solved, taking into account the coupling between the $\alpha - {}^{12}C$ relative motion and the internal rotational motion of ${}^{12}C(0_1^+, 2_1^+, 4_1^+)$. Almost all levels of ${}^{16}O$ up to about $E_x \simeq 14$ MeV including the ground state with the dominant configuration of the doubly closed-shell structure and the electromagnetic transition rates (E0, E2, and E3) among them are reproduced well, together with the low-energy $\alpha - {}^{12}C$ scattering cross sections [53] (see Table I for the monopole strengths). It is found that a lot of states in ¹⁶O up to $E_x \simeq 14$ MeV have a weak coupling structure of $\alpha + {}^{12}C$, that is, loosely bound $\alpha + {}^{12}C$ cluster structure. In particular, the 0^+_2 [0^+_3] state has a weakly coupling structure of the α and ${}^{12}C(\tilde{0}_1^+)$ clusters [α and ${}^{12}C(2_1^+)$], in which the α cluster moves predominantly around the ¹²C cluster with S wave [D wave] (see also Sec. III A). These successes in the $\alpha + {}^{12}C$ OCM mean that (1) the weak coupling picture [48] is realized in 16 O, and the 4p (4h) state of the 4p4h configuration in shell-model picture can be interpreted as the α (¹²C) cluster; and (2) the reason why the energy of the 4p4h configurations is lower than that of the 2p2h one is considered to be the α -cluster correlation for the 4p state. Although the α -cluster structures for the 0_2^+ and 0^+_3 states are much different from what Brown and Green thought [47], one should mention that their conjecture, the energy of the 4p4h configurations being lower that of the 2p2h one, is remarkable, because it has played an important role in the progress of our understanding the structure of 16 O.

D. Two features of isoscalar monopole transitions

As discussed in Sec. III A, the fine structures observed in the experimental monopole strength function in the low-energy region up to $E_x \simeq 16$ MeV can be rather well reproduced within the 4 α OCM framework. However, these fine structures are difficult to be reproduced by any mean-field theory. This result means that the α clustering degree of freedom is inevitable to reconcile the low-energy behavior of the monopole strength function in ¹⁶O with experiment. On the contrary, the RPA calculations look likely to reproduce approximately the three bump structures in the experimental monopole strength function in the higher-energy region of $16 \leq E_x \leq 40$ MeV, although some normalization and energy-shifting procedures for the calculated strength function is needed to fit the experimental data (see Sec. I).

Here we should remind the reader that the ground state of ¹⁶O is described dominantly by a doubly closed-shell structure, $(0s)^4(0p)^{12}$, in the 4 α OCM calculation, as well as the RPA, QRPA, and RRPA calculations. As discussed in Sec. III A, the doubly closed-shell-model wave function is mathematically equivalent to a single α -cluster wave function. This result means that the ground-state wave function originally has an α -clustering degree of freedom together with the single-particle degree of freedom. In other words, the ground-state wave

function of ¹⁶O has a duality of α -clustering character and mean-field-type character.

From these facts, one can notice that there exist two types of the isoscalar monopole excitation of ¹⁶O; that is, the monopole excitations to cluster states are dominant in the lower-energy part ($E_x \lesssim 16$ MeV) of the monopole strength function, whereas the monopole excitation of the one-particle one-hole (1p1h) type contributes to the higher-energy region $(16 \lesssim E_x \lesssim 40 \text{ MeV})$. This also is in line with the first 0^+ excited state of the α particle which is situated at ~ 20 MeV. Thus, one can expect that the reproduction of the experimental isoscalar monopole strength function of ¹⁶O in the full energy region up to $E_x \sim 40$ MeV will definitely fail, if one does not take into account simultaneously the α -cluster-type four-body correlations as well as the 1p1h- and 2p2h-type correlations in the structure study of ¹⁶O. To tackle the issue, a structure calculation is desirable to be performed in which one uses a huge model space covering fully the α -type correlations together with the 1p1h- and 2p2h-type correlations

We here report on a trial calculation using the $\alpha + {}^{12}C$ cluster basis and collective basis for studying the isoscalar monopole strength of ¹⁶O [22]. In the latter respect the symplectic group Sp(6,R) [23,24] as the collective basis was used. Because the generators of the Sp(6,R) group contain the monopole and quadrupole operators with respect to the nucleon coordinates and conjugate momenta, the group is expected to reproduce the EWSR for the operators. Although there are no effective NN interactions that are suited for cluster-Sp(6,R) mixed-basis calculation, they took a phenomenological treatment, because their main purpose was to investigate the effect of the Sp(6,R) group to the cluster states [54,55]. They obtained the following two results. First comes that the effect of the Sp(6,R) basis to the cluster states in the low-energy region is not very large. It is noted that the 0^+_2 and 0_3^+ states (see Fig. 4) are the cluster states that are reproduced by the $\alpha + {}^{12}C$ model. The second result is that there are three states in the higher-energy region ($20 \leq E_x \leq 40 \text{ MeV}$) that correspond to the three peaks of the monopole excitation function and share about 70% of the EWSR, although the excitation energies of the three states are higher by about 3 MeV. These results support our finding that the isoscalar monopole excitations in light nuclei, as mentioned above, are dominated by two features: α -cluster states at low energies and shell-model states at higher energies. An analysis with a mixed 4α cluster and symplectic group basis may be useful for this study.

IV. SUMMARY

We have investigated the monopole strength function in the low-energy region up to $E_x \simeq 16$ MeV within the framework of 4α OCM, which has succeeded in reproducing all the six 0⁺ states observed up to $E_x \simeq 16$ MeV and in giving good agreement with all of the available data such as the decay widths, monopole transition strengths, and rms radius of the ground state. It was found that the fine structures at the low-energy region up to $E_x \simeq 16$ MeV in the experimental monopole strength function obtained by the ${}^{16}O(\alpha, \alpha')$ experiment is rather satisfactorily reproduced within the 4α

OCM framework. On the contrary, mean-field calculations have encountered difficulties in reproducing the fine structures of the monopole strength function at the low-energy region, as well as the monopole matrix elements for the 0_2^+ ($E_x =$ 6.05 MeV), 0_3^+ ($E_x = 12.05$ MeV), and 0_5^+ ($E_x = 14.01$ MeV) states obtained in the ${}^{16}O(e, e')$ experiment. These results mean that the α clustering degree of freedom is inevitably necessary to reconcile the monopole strength (amounting to about 20%) of the EWSR) in the low-energy region with experiment. However, the 4α cluster model has difficulties in reproducing the gross three bump structures of the monopole strength function at the higher-energy region of $16 \lesssim E_x \lesssim 40$ MeV. The gross bump structures look likely to be qualitatively reproduced by the mean-field-theory calculations such as RRPA, RPA, and QRPA, although the energy positions of the three bumps and the absolute values of the strength functions quantitatively deviate from the experimental data. In general, one can expect the interplay between clustering degrees of freedom and mean-field degrees of freedom. Because the interplay affects the monopole strength function, there is a possibility that the isoscalar monopole strength at the higher-energy region of $16 \leq E_x \leq 40$, in particular at its lower-energy part, contains the influence by clustering degrees of freedom. The fact that no mean-field-theory calculations have satisfactorily reproduced the isoscalar monopole strength function even at the higher energy region ($16 \leq E_x \leq 40$ MeV) demonstrates the need to investigate the interplay between clustering and mean-field degrees of freedom in the isoscalar monopole strength function.

From the above results, one concludes that there exist two features of the isoscalar monopole excitations of ¹⁶O; that is, the monopole excitation to cluster states dominates the low-energy region ($E_x \lesssim 16$ MeV), sharing about 20% of the EWSR, while that to the 1p1h-type states looks likely to be predominant at the higher-energy region. We indicated that the existence of these two types of the monopole excitations stems from the fact that the ground state of ¹⁶O with the dominant doubly closed-shell configuration $(0s)^4(0p)^{12}$ that is the dominant SU(3) $(\lambda \mu) = (00)$ symmetry has, in fact, a dual feature inasmuch as it can equivalently be described by a cluster wave function of the α type, as can be shown with the Bayman-Bohr theorem. When the monopole operator activates α -type degrees of freedom in the ground state, α -cluster states are excited, while in the case of the monopole operator acting on the 1p1h-type degree of freedom in the ground state, collective states of the 1p1h type are populated. Thus, one will fail to reproduce the experimental isoscalar monopole strength function of ¹⁶O up to $E_x \sim 40$ MeV if the α -cluster-type four-body correlations, as well as the 1p1h- and 2p2h-type correlations, are not simultaneously taken into account in the structure study of ¹⁶O.

The existence of two features of isoscalar monopole excitation which originates from the dual nature of the ground state seems to be general in light nuclei, and the case of 16 O discussed in the present paper is typical. This is attributable to the nuclear SU(3) symmetry which is well verified in the ground state of light nuclei. According to the Bayman-Bohr theorem, an SU(3) wave function is mathematically equivalent to a cluster-model wave function. Thus, the ground state which

has a dominant SU(3) symmetry is considered to have the dual nature similar to the case of ¹⁶O, which generates the two features in the isoscalar monopole excitation. However, this feature will be vanishing with increasing mass number, and eventually only the 1p1h-type collective motions are strongly excited, maybe, in the mass region beyond the *fp*-shell nuclei. The reason for this is that the quality of the nuclear SU(3)symmetry in the ground state of light nuclei is gradually disappearing because of the stronger effect from the spin-orbit forces in heavier nuclei. This means that the dual nature of the ground state is also corroding with increasing mass number. It is an intriguing subject to study theoretically and experimentally how these two features are changing with the mass number. Thus, it is strongly hoped that systematic experiments of analyzing the existence of these two features of monopole excitations will be performed in near future.

(Japan Society for the Promotion of Science) Grant-in-Aid for Scientific Research (C) (21540283) and Young Scientists (B) (21740209).

APPENDIX: CLUSTER SUM RULE OF ISOSCALAR MONOPOLE TRANSITION

In this Appendix, we first discuss the EWSR of the isoscalar monopole transition for ^{*A*}Z nucleus. Then we discuss the EWSR of the isoscalar monopole transition within the framework of the OCM, called the OCM-EWSR, in the case of the $n\alpha$ OCM and two-cluster OCM (${}^{A_1}Z_1$ and ${}^{A_2}Z_2$ with $Z_1 = A_1/2$ and $Z_2 = A_2/2$) for a self-conjugated nucleus A = 4n. Finally, a general formula of the OCM-EWSR value is presented in the case of the *k*-cluster OCM of ${}^{Z}A$ nucleus (k = 2, 3, ...), composed of the *k* clusters (${}^{Z_1}A_1, {}^{Z_2}A_2, ..., {}^{Z_k}A_k$).

The EWSR of the isoscalar monopole transition [14] for ${}^{A}Z$ nucleus [see Eq. (7) for ${}^{16}O$] is given as

ACKNOWLEDGMENTS

The authors thank Professor K. Katō for useful discussions and comments. This work was partially supported by JSPS

$$\sum_{n} (E_{n} - E_{1}) \left| \langle 0_{n}^{+} | \sum_{i=1}^{A} (\mathbf{r}_{i} - \mathbf{R}_{\text{c.m.}})^{2} | 0_{1}^{+} \rangle \right|^{2} = \frac{2\hbar^{2}}{m} \times A \times R^{2},$$
(A1)

$$R^{2} = \frac{1}{A} \langle 0_{1}^{+} | \sum_{i=1}^{A} (\boldsymbol{r}_{i} - \boldsymbol{R}_{\text{c.m.}})^{2} | 0_{1}^{+} \rangle,$$
(A2)

where R represents the rms radius of the ground state, and other notations are self-evident. Here, we assume that the NN interaction has no velocity dependence. The isoscalar monopole operator in Eq. (A1) can be decomposed as

$$\sum_{i=1}^{4n} (\mathbf{r}_i - \mathbf{R}_{\text{c.m.}})^2 = \sum_{k=1}^n \sum_{i=1}^4 (\mathbf{r}_{i+4(k-1)} - \mathbf{R}_{\alpha_k})^2 + \sum_{k=1}^n 4 (\mathbf{R}_{\alpha_k} - \mathbf{R}_{\text{c.m.}})^2,$$
(A3)

$$= \sum_{i \in A_1} \left(\boldsymbol{r}_i - \boldsymbol{R}_{A_1} \right)^2 + \sum_{i \in A_2} \left(\boldsymbol{r}_i - \boldsymbol{R}_{A_2} \right)^2 + \frac{A_1 A_2}{A_1 + A_2} \boldsymbol{\xi}^2, \tag{A4}$$

where $\mathbf{R}_{\alpha_k} = (1/4) \sum_{i=1}^{4} \mathbf{r}_{i+4(k-1)}$ is the c.m. coordinate of the *k*th α cluster, \mathbf{R}_{A_1} (\mathbf{R}_{A_2}) stands for the c.m. coordinate of the $^{A_1}Z_1$ ($^{A_2}Z_2$) cluster, and $\boldsymbol{\xi} = \mathbf{R}_{A_2} - \mathbf{R}_{A_1}$ denotes the relative coordinate between the two clusters. As discussed in Sec. II B, the isoscalar monopole operator in the $n\alpha$ OCM gives nonzero contribution only for the second term of Eq. (A3), and that in the two-cluster OCM provides with nonzero contribution only for the third term of Eq. (A4).

In the $n\alpha$ OCM, the total wave function of 0^+ state is presented as

$$|0^{+}\rangle = \Phi^{(\text{OCM})}(0^{+}) \prod_{k=1}^{n} \phi(\alpha_{k}),$$
 (A5)

where $\phi(\alpha_k)$ is the internal wave function of *k*th α cluster and $\Phi^{(\text{OCM})}(0^+)$ stands for the relative wave function among the $n\alpha$ clusters (see Sec. II B). Then the EWSR-OCM is evaluated as

$$\sum_{p} \left(E_{p}^{(\text{OCM})} - E_{1}^{(\text{OCM})} \right) \left| \langle 0_{p}^{+} | \sum_{i=1}^{4n} (\mathbf{r}_{i} - \mathbf{R}_{\text{c.m.}})^{2} | 0_{1}^{+} \rangle \right|^{2},$$

$$= \sum_{p} \left(E_{p}^{(\text{OCM})} - E_{1}^{(\text{OCM})} \right) \left| \langle \Phi^{(\text{OCM})} (0_{p}^{+}) | \sum_{k=1}^{n} 4 \left(\mathbf{R}_{\alpha_{k}} - \mathbf{R}_{\text{c.m.}} \right)^{2} | \Phi^{(\text{OCM})} (0_{1}^{+}) \rangle \right|^{2}$$

$$= \frac{2\hbar^{2}}{m} \left\langle \Phi^{(\text{OCM})} (0_{1}^{+}) | \sum_{k=1}^{n} 4 \left(\mathbf{R}_{\alpha_{k}} - \mathbf{R}_{\text{c.m.}} \right)^{2} | \Phi^{(\text{OCM})} (0_{1}^{+}) \rangle,$$

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(A6)

$$=\frac{2\hbar^2}{m}\times A\times (R^2-R(\alpha)^2),$$

where $\Phi^{(\text{OCM})}(0_p^+)$ and $E_p^{(\text{OCM})}$ are the eigenwave function and eigenvalue of the *p*th 0⁺ state obtained by solving the *n* α OCM equation. Here the α - α interaction is assumed to be velocity-independent, and $R(\alpha)$ stands for the rms radius of the α cluster. The derivation of Eq. (A6) should be referred to Sec. II B. Then the ratio of the OCM-EWSR to the total EWSR in Eq. (A1) is

$$\frac{\text{OCM-EWSR}}{\text{total EWSR}} = 1 - \left(\frac{R(\alpha)}{R}\right)^2.$$
 (A7)

This result is common to all *n* and hence is the same as Eq. (23) for n = 4. As shown in Sec. II B, the ratio for the 4α OCM is 68%.

In the case of the two-cluster OCM, the total wave function of 0^+ state is presented as

$$|0^{+}\rangle = \Phi^{(2-\text{clus.})}(0^{+})\phi({}^{A_{1}}Z_{1})\phi({}^{A_{2}}Z_{2}), \tag{A8}$$

where $\phi({}^{A_1}Z_1) [\phi({}^{A_2}Z_2)]$ is the internal wave function of the ${}^{A_2}Z_2$ cluster $({}^{A_2}Z_2)$ and $\Phi^{(2clus.)}(0^+)$ stands for the relative wave function between the two clusters. Then the EWSR-OCM is presented as

$$\sum_{p} \left(E_{p}^{(2\text{clus.})} - E_{1}^{(2\text{clus.})} \right) \left| \langle 0_{p}^{+} | \sum_{i=1}^{4n} (\mathbf{r}_{i} - \mathbf{R}_{\text{c.m.}})^{2} | 0_{1}^{+} \rangle \right|^{2},$$

$$= \sum_{n} \left(E_{n}^{(2\text{clus.})} - E_{1}^{(2\text{clus.})} \right)$$

$$\times \left| \langle \Phi^{(2\text{clus.})}(0_{n}^{+}) | \frac{A_{1}A_{2}}{A_{1} + A_{2}} \boldsymbol{\xi}^{2} | \Phi^{(2\text{clus.})}(0_{1}^{+}) \rangle \right|^{2},$$

$$= \frac{2\hbar^{2}}{m} \langle \Phi^{(2\text{clus.})}(0_{1}^{+}) | \frac{A_{1}A_{2}}{A_{1} + A_{2}} \boldsymbol{\xi}^{2} | \Phi^{(2\text{clus.})}(0_{1}^{+}) \rangle,$$

$$= \frac{2\hbar^{2}}{m} \times A \times \left[R^{2} - \frac{A_{1}}{A} R(A_{1})^{2} - \frac{A_{2}}{A} R(A_{2})^{2} \right], \quad (A9)$$

where $\Phi^{(2\text{clus.})}(0_n^+)$ and $E_p^{(2\text{clus.})}$ are the eigenwave function and eigenvalue of the *p*th 0⁺ state obtained by solving the two-cluster OCM equation. Here the two-cluster potential is assumed to be velocity independent, and $R(A_1) [R(A_2)]$ stands for the rms radius of the ${}^{A_1}Z_1 ({}^{A_2}Z_2)$ cluster. It is noted that this EWSR-OCM is realized in the case of the coupled-channel

- D. H. Youngblood, C. M. Rozsa, J. M. Moss, D. R. Brown, and J. D. Bronson, Phys. Rev. Lett. **39**, 1188 (1977).
- [2] D. H. Youngblood, H. L. Clark, and Y.-W. Lui, Phys. Rev. Lett. 82, 691 (1999).
- [3] V. R. Pandharipande, Phys. Lett. B 31, 635 (1970).
- [4] J. P. Blaizot, D. Gogny, and B. Grammaticos, Nucl. Phys. A 265, 315 (1976).
- [5] D. Lebrun, M. Buenerd, P. Martin, P. de Saintignon, and G. Perrin, Phys. Lett. B 97, 358 (1980).
- [6] S. Drożdż, S. Nishizaki, J. Speth, and J. Wambach, Phys. Rep. 197, 1 (1990).

OCM, for example, $\alpha + {}^{12}C(0^+_1, 2^+_1, 4^+_1)$, where there is no contribution of the internal monopole transitions in both the α and ${}^{12}C$ clusters. Then the ratio of the OCM-EWSR to the total EWSR in Eq. (A1) is

$$\frac{\text{OCM-EWSR(2-cluster)}}{\text{total EWSR}} = 1 - \frac{A_1}{A} \left(\frac{R(A_1)}{R}\right)^2 - \frac{A_2}{A} \left(\frac{R(A_2)}{R}\right)^2. \quad (A10)$$

This ratio for the $\alpha + {}^{12}C$ OCM is 31%, which is about half of that in the 4 α OCM.

Finally, let us discuss the OCM-EWSR value, in general, in the case of the *k*-cluster OCM of a ^ZA nucleus (k = 2, 3, ...), composed of the *k* clusters (${}^{Z_1}A_1, {}^{Z_2}A_2, ..., {}^{Z_k}A_k$). The total wave function of 0⁺ state is given as

$$|0^{+}\rangle = \Phi^{(k-\text{clus.})}(0^{+}) \prod_{i=1}^{k} \phi^{(A_{i}} Z_{i}),$$
 (A11)

where $\phi({}^{A_i}Z_i)$ is the internal wave function of the ${}^{A_i}Z_i$ cluster and $\Phi^{(k-\text{clus.})}(0^+)$ stands for the relative wave function among the *k* clusters. Then, the EWSR-OCM is presented as

$$\sum_{p} \left(E_{p}^{(k-\text{clus.})} - E_{1}^{(k-\text{clus.})} \right) \left| \langle 0_{p}^{+} | \sum_{i=1}^{4n} (\boldsymbol{r}_{i} - \boldsymbol{R}_{\text{c.m.}})^{2} | 0_{1}^{+} \rangle \right|^{2},$$
$$= \frac{2\hbar^{2}}{m} \times A \times \left[R^{2} - \sum_{i=1}^{k} \frac{A_{i}}{A} R(A_{i})^{2} \right], \qquad (A12)$$

where $R(A_i)$ denotes the rms radius of the ${}^{A_i}Z_i$ cluster, and we assumed no contributions from the internal monopole transitions in the ${}^{Z_i}A_i$ nucleus (i = 1, 2, ..., k). The proof of Eq. (A12) is similar to those of Eqs. (A6) and (A9). Then, the ratio of the OCM-EWSR to the total EWSR in Eq. (A1) is

$$\frac{\text{OCM-EWSR}(k\text{-cluster})}{\text{total EWSR}} = 1 - \sum_{i=1}^{k} \frac{A_i}{A} \left(\frac{R(A_i)}{R}\right)^2.$$
 (A13)

The present results can be applied to the monopole transition in neutron-rich nuclei. For example, the ratio for the $\alpha + \alpha + n + n$ OCM of ¹⁰Be amounts to 68%.

- [7] N. Paar, P. Papakonstantinou, H. Hergert, and R. Roth, Phys. Rev. C 74, 014318 (2006).
- [8] P. Papakonstantinou, R. Roth, and N. Paar, Phys. Rev. C 75, 014310 (2007).
- [9] P. Papakonstantinou and R. Roth, Phys. Lett. B 671, 356 (2009).
- [10] Z. Ma, N. Van Giai, H. Toki, and M. L'Huillier, Phys. Rev. C 55, 2385 (1997).
- [11] T. Furuta, K. H. O. Hasnaoui, F. Gulminelli, C. Leclercq, and A. Ono, Phys. Rev. C 82, 034307 (2010).
- [12] Y.-W. Lui, H. L. Clark, and D. H. Youngblood, Phys. Rev. C 64, 064308 (2001).

- [13] F. Ajzenberg-Selove, Nucl. Phys. A 460, 1 (1986); D. R. Tilley,
 H. R. Weller, and C. M. Cheves, *ibid.* 564, 1 (1993); I. Angeli,
 At. Data Nucl. Data Tables 87, 185 (2004).
- [14] T. Yamada, Y. Funaki, H. Horiuchi, K. Ikeda, and A. Tohsaki, Prog. Theor. Phys. **120**, 1139 (2008).
- [15] D. Gambacurta, M. Grasso, and F. Catara, Phys. Rev. C 81, 054312 (2010).
- [16] Y. Suzuki, Prog. Theor. Phys. 55, 1751 (1976); 56, 111 (1976).
- [17] S. Saito, Prog. Theor. Phys. 40, 893 (1968); 41, 705 (1969);
 Prog. Theor. Phys. Suppl. 62, 11 (1977).
- [18] K. Ikeda, H. Horiuchi, and S. Saito, Prog. Theor. Phys. Suppl. 68, 1 (1980).
- [19] M. Libert-Heinemann, D. Baye, and P.-H. Heenen, Nucl. Phys. A 339, 429 (1980).
- [20] Y. Funaki, T. Yamada, H. Horiuchi, G. Röpke, P. Schuck, and A. Tohsaki, Phys. Rev. Lett. 101, 082502 (2008).
- [21] T. Yamada, Y. Funaki, H. Horiuchi, G. Röpke, P. Schuck, and A. Tohsaki, in *Cluster in Nuclei - Vol. 2*, edited by C. Beck, Lecture Notes in Physics (Springer-Verlag, Berlin, 2011), Vol. 848, Chap. 5.
- [22] Y. Suzuki and S. Hara, Phys. Rev. C 39, 658 (1989).
- [23] G. Rosensteel and D. J. Rowe, Phys. Rev. Lett. 38, 10 (1977); Ann. Phys. 126, 343 (1980).
- [24] D. J. Rowe, Rep. Prog. Phys. 48, 1419 (1985).
- [25] M. Kamimura, Phys. Rev. A 38, 621 (1988); E. Hiyama,
 Y. Kino, and M. Kamimura, Prog. Part. Nucl. Phys. 51, 223 (2003).
- [26] H. Horiuchi, Prog. Theor. Phys. 58, 204 (1977); Prog. Theor. Phys. Suppl. 62, 90 (1977).
- [27] A. Hasegawa and S. Nagata, Prog. Theor. Phys. 45, 1786 (1971);
 F. Tanabe, A. Tohsaki, and R. Tamagaki, *ibid*. 53, 677 (1975).
- [28] T. Yamada and P. Schuck, Euro. Phys. J. A 26, 185 (2005).
- [29] M. Itoh, H. Akimune, M. Fujiwara, U. Garg, N. Hashimoto, T. Kawabata, K. Kawase, S. Kishi, T. Murakami, K. Nakanishi, Y. Nakatsugawa, B. K. Nayak, S. Okumura, H. Sakaguchi, H. Takeda, S. Terashima, M. Uchida, Y. Yasuda, M. Yosoi, and J. Zenihiro, Phys. Rev. C 84, 054308 (2011).
- [30] In the present 4α OCM calculation, the four-body force is taken as $V_{4\alpha}(1, 2, 3, 4) = V_{4\alpha}^{(0)} \exp[-\beta_4 \sum_{i< j=1}^4 (\boldsymbol{R}_{\alpha_i} - \boldsymbol{R}_{\alpha_j})^2]$, where $V_{4\alpha}^{(0)} = 12\,000$ MeV and $\beta_4 = 0.15$ fm⁻². The expectation value of the four-body force as well as the three-body force $\sum_{i< j< k}^4 V_{3\alpha}(i, j, k)$ is less than 7% in the present study, compared with that of the two-body force, $\sum_{i< j=1}^4 [V_{2\alpha}^{(N)}(i, j) + V_{2\alpha}^{(C)}(i, j)]$ [see Eq. (15)].

- PHYSICAL REVIEW C 85, 034315 (2012)
- [31] A. M. Lane and R. G. Thomas, Rev. Mod. Phys. 30, 257 (1958).
- [32] V. I. Kukulin, V. M. Krasnopol'ky, and J. Horáček, *Theory of Resonances* (Akademia, Praha, 1989).
- [33] R. D. Lawson, *Theory of Nuclear Shell Model* (Clarendon Press, Oxford, 1980).
- [34] A. B. Volkov, Nucl. Phys. 74, 33 (1965).
- [35] R. Tamagaki, Prog. Theor. Phys. 39, 91 (1968).
- [36] T. Myo, S. Sugimoto, K. Katō, H. Toki, and K. Ikeda, Prog. Theor. Phys. 117, 257 (2007).
- [37] T. Myo, A. Umeya, H. Toki, and K. Ikeda, Phys. Rev. C 84, 034315 (2011).
- [38] T. Wakasa et al., Phys. Lett. B 653, 173 (2007).
- [39] H. Horiuchi and K. Ikeda, Prog. Theor. Phys. 40, 277 (1968).
- [40] J. P. Elliott, Proc. R. Soc. London A 245, 128 (1958); 245, 562 (1958).
- [41] T. Dytrych, K. D. Sviratcheva, C. Bahri, J. P. Draayer, and J. P. Vary, Phys. Rev. Lett. 98, 162503 (2007).
- [42] A. Tohsaki, H. Horiuchi, P. Schuck, and G. Röpke, Phys. Rev. Lett. 87, 192501 (2001).
- [43] Y. Funaki, A. Tohsaki, H. Horiuchi, P. Schuck, and G. Röpke, Phys. Rev. C 67, 051306 (2003).
- [44] B. F. Bayman and A. Bohr, Nucl. Phys. 9, 596 (1958-1959).
- [45] Y. Funaki, T. Yamada, A. Tohsaki, H. Horiuchi, G. Röpke, and P. Schuck, Phys. Rev. C 82, 024312 (2010).
- [46] M. Wloch, D. J. Dean, J. R. Gour, M. Hjorth-Jensen, K. Kowalski, T. Papenbrock, and P. Piecuch, Phys. Rev. Lett. 94, 212501 (2005).
- [47] G. E. Brown and A. M. Green, Nucl. Phys. 75, 401 (1966).
- [48] A. Arima, H. Horiuchi, and T. Sebe, Phys. Lett. B 24, 129 (1967).
- [49] In Ref. [47], they take the SU(3)(λ , μ) = (0, 0), (6, 2), and (8, 4) states as the 0p0h, 2p2h, and 4p4h states, respectively. The latter two SU(3) states correspond, respectively, to the states having maximal deformation among the $2\hbar\omega$ and $4\hbar\omega$ states described by the nuclear SU(3) model [40].
- [50] G. F. Bertsch, Phys. Lett. 21, 70 (1966).
- [51] W. C. Haxton and C. Johnson, Phys. Rev. Lett. 65, 1325 (1990).
- [52] E. K. Warburton, B. A. Brown, and D. J. Millener, Phys. Lett. B 293, 7 (1992).
- [53] Y. Suzuki, T. Ando, and B. Imanishi, Nucl. Phys. A 295, 365 (1978).
- [54] K. T. Hecht, Phys. Rev. C 16, 2401 (1977).
- [55] K. T. Hecht and D. Braunschweig, Nucl. Phys. A 295, 34 (1978).