

Uncertainties in nuclear transition matrix elements for neutrinoless $\beta\beta$ decay: The heavy Majorana neutrino mass mechanism

P. K. Rath,¹ R. Chandra,^{2,3} P. K. Raina,^{3,4} K. Chaturvedi,⁵ and J. G. Hirsch⁶

¹*Department of Physics, University of Lucknow, Lucknow-226007, India*

²*Department of Applied Physics, Babasaheb Bhimrao Ambedkar University, Lucknow-226025, India*

³*Department of Physics and Meteorology, Indian Institute of Technology, Kharagpur-721302, India*

⁴*Department of Physics, Indian Institute of Technology, Ropar, Rupnagar-140001, Punjab, India*

⁵*Department of Physics, Bundelkhand University, Jhansi-284128, India*

⁶*Instituto de Ciencias Nucleares, Universidad Nacional Autónoma de México, 04510 México, D.F., México*

(Received 8 June 2011; revised manuscript received 5 December 2011; published 11 January 2012)

Employing four different parametrizations of the pairing plus the multipolar type of effective two-body interaction and three different parametrizations of the Jastrow type of short-range correlations, the uncertainties in the nuclear transition matrix elements $M_N^{(0\nu)}$ due to the exchange of heavy Majorana neutrino for the $0^+ \rightarrow 0^+$ transition of neutrinoless double beta decay of ^{94}Zr , ^{96}Zr , ^{98}Mo , ^{100}Mo , ^{104}Ru , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes in the PHFB model are estimated to be around 35%. Excluding the nuclear transition matrix elements calculated with the Miller-Spencer parametrization of Jastrow short-range correlations, the uncertainties are found to be smaller than 20%.

DOI: [10.1103/PhysRevC.85.014308](https://doi.org/10.1103/PhysRevC.85.014308)

PACS number(s): 21.60.-n, 23.40.Hc, 27.60.+j

I. INTRODUCTION

In addition to establishing the Dirac or Majorana nature of neutrinos, the observation of $(\beta\beta)_{0\nu}$ decay is a convenient tool to test the lepton number conservation, possible hierarchies in the neutrino mass spectrum, the origin of neutrino mass, and CP violation in the leptonic sector. Furthermore, it can also ascertain the role of various gauge models associated with all possible mechanisms, namely the exchange of light neutrinos, heavy neutrinos, the right-handed currents in the left-right symmetric model (LRSM), the exchange of sleptons, neutralinos, squarks, and gluinos in the R_p -violating minimal supersymmetric standard model, the exchange of leptoquarks, existence of heavy sterile neutrinos, compositeness, extradimensional scenarios, and Majoron models, allowing the occurrence of $(\beta\beta)_{0\nu}$ decay. Stringent limits on the associated parameters have already been extracted from the observed experimental limits on the half-life of $(\beta^-\beta^-)_{0\nu}$ decay [1] and, presently, all the experimental attempts are directed for its observation. The experimental and theoretical studies devoted to $(\beta\beta)_{0\nu}$ decay over the past decades have been recently reviewed by Avignone *et al.* [2] and references therein.

Presently, there is an increased interest in calculating reliable NTMEs for $(\beta^-\beta^-)_{0\nu}$ decay from the exchange of heavy Majorana neutrinos to ascertain the dominant mechanism contributing to it [3,4]. The lepton number violating $(\beta^-\beta^-)_{0\nu}$ decay was studied by Vergados by taking a Lagrangian consisting of left-handed as well as right-handed leptonic currents [5]. In the QRPA, the $(\beta^-\beta^-)_{0\nu}$ decay from the exchange of heavy Majorana neutrinos was studied by Tomoda [6]. The decay rate of $(\beta^-\beta^-)_{0\nu}$ mode in the LRSM was derived by Doi and Kotani [7]. Hirsch *et al.* [8] have calculated all the required nuclear transition matrix elements (NTMEs) in the QRPA and limits on the effective light neutrino mass $\langle m_\nu \rangle$, heavy neutrino mass $\langle M_N \rangle$, right-handed heavy neutrino $\langle M_R \rangle$, $\langle \lambda \rangle$, $\langle \eta \rangle$, and mixing angle $\tan\xi$ have been obtained.

The heavy neutrino mechanism was also studied in the QRPA without [9] and with pn pairing [10]. In the heavy Majorana neutrino mass mechanism, Simkovic *et al.* [11] have studied the role of induced weak magnetism and pseudoscalar terms and it was found that they are quite important in ^{48}Ca nucleus. The importance of the same induced currents in both the light and heavy Majorana neutrino exchange mechanisms was also studied using the pn RQRPA [12] as well as SRQRPA [3].

In spite of the remarkable success of the large-scale shell model (LSSM) calculations of the Strassbourg-Madrid group [13], there is a necessity for large configuration mixing to reproduce the structural complexity of medium and heavy mass nuclei. On the other hand, the QRPA and its extensions have emerged as successful models by including a large number of basis states and in correlating the single- β GT strengths and half-lives of $(\beta^-\beta^-)_{2\nu}$ decay in addition to explaining the observed suppression of $M_{2\nu}$ [14,15]. In the mass region $90 \leq A \leq 150$, there is a subtle interplay of pairing and quadrupolar correlations and their effects on the NTMEs of $(\beta^-\beta^-)_{0\nu}$ decay have been studied in the interacting shell model (ISM) [16,17], deformed QRPA model [18–21], and projected-Hartree-Fock-Bogoliubov (PHFB) model [22,23].

The possibility of constraining the values of the gauge parameters using the measured lower limits on the $(\beta^-\beta^-)_{0\nu}$ decay half-lives relies heavily on the model-dependent NTMEs. Different predictions are obtained by employing different nuclear models, and within a given model, varying the model space, single-particle energies (SPEs), and effective two-body interaction. In addition, a number of issues regarding the structure of NTMEs, namely the effect of pseudoscalar and weak magnetism terms on the Fermi, Gamow-Teller, and tensorial NTMEs [24,25], the role of finite size of nucleons (FNS) as well as short-range correlations (SRC) vis-a-vis the radial evolution of NTMEs [16,26–28], and the value

of the axial-vector coupling constant g_A are also sources of uncertainty and remain to be investigated.

It was observed by Vogel [29] that in the case of well-studied ^{76}Ge , the calculated decay rates $T_{1/2}^{0\nu}$ differ by a factor of 6–7 and consequently, the uncertainty in the effective neutrino mass (m_ν) is about 2–3. Thus, the spread between the calculated NTMEs can be used as the measure of the theoretical uncertainty. In case the $(\beta\beta)_{0\nu}$ decay of different nuclei is observed, Bilenky and Grifols [30] have suggested that the results of calculations of NTMEs of the $(\beta^-\beta^-)_{0\nu}$ decay can be checked by comparing the calculated ratios of the corresponding NTMEs squared with the experimentally observed values.

Bahcall *et al.* [31] and Avignone *et al.* [32] have calculated averages of all the available NTMEs, and their standard deviation is taken as the measure of theoretical uncertainty. On the other hand, Rodin *et al.* [33] have calculated nine NTMEs with three sets of basis states and three realistic two-body effective interactions of charge-dependent Bonn, Argonne, and Nijmen potentials in the QRPA as well as RQRPA and estimated the theoretical uncertainties by making a statistical analysis. It was noticed that the variances are substantially smaller than the average values and the results of QRPA, albeit slightly larger, are quite close to the RQRPA values. Faessler and coworkers have further studied uncertainties in NTMEs from short-range correlations using the unitary correlation operator method (UCOM) [26] and self-consistent coupled cluster method (CCM) [27].

The PHFB model has the advantage of treating the pairing and deformation degrees of freedom on equal footing and projecting out states with good angular momentum. However, the single β -decay rates and the distribution of GT strength, which require the structure of the intermediate odd Z -odd N nuclei, cannot be studied in the present version of the PHFB model. In spite of this limitation, the PHFB model in conjunction with pairing plus quadrupole-quadrupole (PQQ) [34] was successfully applied to reproduce the lowest yrast states, electromagnetic properties of the parent and daughter nuclei, and the measured $(\beta^-\beta^-)_{2\nu}$ -decay rates [35,36]. In the PHFB formalism, the existence of an inverse correlation between the quadrupole deformation and the size of NTMEs $M_{2\nu}$, $M^{(0\nu)}$, and $M_N^{(0\nu)}$ was observed [22,23]. Further, it was noticed that the NTMEs are usually large for a pair of spherical nuclei, almost constant for small deformation, suppressed depending on the difference in the deformation $\Delta\beta_2$ of parent and daughter nuclei, and having a well defined maximum when $\Delta\beta_2 = 0$ [22,23].

In Ref. [37], a statistical analysis was performed for extracting uncertainties in eight (twelve) NTMEs for $(\beta^-\beta^-)_{0\nu}$ decay from the exchange of light Majorana neutrino, calculated in the PHFB model with four different parametrizations of pairing plus the multipolar type of effective two-body interaction [23] and two (three) different parametrizations of the Jastrow type of SRC [27]. In confirmation with the observation made by Simkovic *et al.* [27], it was noticed that the Miller-Spencer type of parametrization is a major source of uncertainty and its exclusion reduces the uncertainties from 10%–15% to 4%–14%. Presently, the same procedure was adopted to estimate the theoretical uncertainties associated

with the NTMEs $M_N^{(0\nu)}$ for $(\beta^-\beta^-)_{0\nu}$ decay from the exchange of heavy Majorana neutrino. In Sec. II, a brief discussion of the theoretical formalism is presented. The results for different parametrizations of the two-body interaction and SRC vis-a-vis radial evolution of NTMEs are discussed in Sec III. In the same section, the averages as well as standard deviations are calculated for estimating the theoretical uncertainties. Finally, the conclusions are given in Sec. IV.

II. THEORETICAL FORMALISM

In the charged current weak processes, the current-current interaction under the assumption of zero mass neutrinos leads to terms which, except for vector and axial vector parts, are proportional to the lepton mass squared, and hence negligible. However, it was reported by Simkovic *et al.* [24,25] that the contribution of the pseudoscalar term is equivalent to a modification of the axial vector current because of PCAC and greater than the vector current. The contributions of pseudoscalar and weak magnetism terms in the mass mechanism can change $M^{(0\nu)}$ up to 30% and the change in $M_N^{(0\nu)}$ is considerably larger. In the shell model [16,38], IBM [39], and GCM + PNAMP [40], the contributions of these pseudoscalar and weak magnetism terms to $M^{(0\nu)}$ have also been investigated. However, it was reported by Suhonen and Civitarese [41] that these contributions are relatively small and can be safely neglected. Therefore, the investigation of this issue is of definite interest and is reported in the present work.

In the two nucleon mechanism, the half-life $T_{1/2}^{0\nu}$ for the $0^+ \rightarrow 0^+$ transition of $(\beta^-\beta^-)_{0\nu}$ decay from the exchange of heavy Majorana neutrino between nucleons having finite size is given by [6,7]

$$[T_{1/2}^{0\nu}(0^+ \rightarrow 0^+)]^{-1} = \left(\frac{m_p}{\langle M_N \rangle}\right)^2 G_{01} |M_N^{(0\nu)}|^2, \quad (1)$$

where m_p is the proton mass and

$$\langle M_N \rangle^{-1} = \sum_i U_{ei}^2 m_i^{-1}, \quad m_i > 1 \text{ GeV}, \quad (2)$$

and in the closure approximation, the NTMEs $M_N^{(0\nu)}$ is of the form [12,26,27],

$$M_N^{(0\nu)} = -M_{Fh} + M_{GTh} + M_{Th}, \quad (3)$$

where

$$M_\alpha = \sum_{n,m} \langle 0_F^+ || O_{\alpha, nm} \tau_n^+ \tau_m^+ || 0_I^+ \rangle, \quad (4)$$

with

$$O_{Fh} = H_{Fh}(r_{nm}), \quad (5)$$

$$O_{GTh} = \sigma_n \cdot \sigma_m H_{GTh}(r_{nm}), \quad (6)$$

$$O_{Th} = [3(\sigma_n \cdot \hat{\mathbf{r}}_{nm})(\sigma_m \cdot \hat{\mathbf{r}}_{nm}) - \sigma_n \cdot \sigma_m] H_{GTh}(r_{nm}). \quad (7)$$

The exchange of heavy Majorana neutrinos gives rise to short-ranged neutrino potentials, which with the consideration

of FNS are given by

$$H_{\alpha h}(r_{nm}) = \frac{2R}{(m_p m_e)\pi} \int f_{\alpha h}(qr_{nm}) h_{\alpha}(q) q^2 dq, \quad (8)$$

where $f_{\alpha h}(qr_{nm}) = j_0(qr_{nm})$ for $\alpha = F$ as well as GT and $f_{Th}(qr_{nm}) = j_2(qr_{nm})$.

Furthermore, the $h_F(q)$, $h_{GT}(q)$ and $h_T(q)$ are written as

$$h_F(q) = \left(\frac{g_V}{g_A}\right)^2 \left(\frac{\Lambda_V^2}{q^2 + \Lambda_V^2}\right)^4, \quad (9)$$

$$h_{GT}(q) = \frac{g_A^2(q^2)}{g_A^2} \left[1 - \frac{2}{3} \frac{g_P(q^2)q^2}{g_A(q^2)2m_p} + \frac{1}{3} \frac{g_P^2(q^2)q^4}{g_A^2(q^2)4m_p^2} \right] + \frac{2}{3} \frac{g_M^2(q^2)q^2}{g_A^2 4m_p^2} \\ \approx \left(\frac{\Lambda_A^2}{q^2 + \Lambda_A^2}\right)^4 \left[1 - \frac{2}{3} \frac{q^2}{(q^2 + m_\pi^2)} + \frac{1}{3} \frac{q^4}{(q^2 + m_\pi^2)^2} \right] + \left(\frac{g_V}{g_A}\right)^2 \frac{\kappa^2 q^2}{6m_p^2} \left(\frac{\Lambda_V^2}{q^2 + \Lambda_V^2}\right)^4, \quad (10)$$

$$h_T(q) = \frac{g_A^2(q^2)}{g_A^2} \left[\frac{2}{3} \frac{g_P(q^2)q^2}{g_A(q^2)2m_p} - \frac{1}{3} \frac{g_P^2(q^2)q^4}{g_A^2(q^2)4m_p^2} \right] + \frac{1}{3} \frac{g_M^2(q^2)q^2}{g_A^2 4m_p^2} \\ \approx \left(\frac{\Lambda_A^2}{q^2 + \Lambda_A^2}\right)^4 \left[\frac{2}{3} \frac{q^2}{(q^2 + m_\pi^2)} - \frac{1}{3} \frac{q^4}{(q^2 + m_\pi^2)^2} \right] + \left(\frac{g_V}{g_A}\right)^2 \frac{\kappa^2 q^2}{12m_p^2} \left(\frac{\Lambda_V^2}{q^2 + \Lambda_V^2}\right)^4, \quad (11)$$

where the form factors are given by

$$g_A(q^2) = g_A \left(\frac{\Lambda_A^2}{q^2 + \Lambda_A^2}\right)^2, \\ g_M(q^2) = \kappa g_V \left(\frac{\Lambda_V^2}{q^2 + \Lambda_V^2}\right)^2, \quad (12) \\ g_P(q^2) = \frac{2m_p g_A(q^2)}{(q^2 + m_\pi^2)} \left(\frac{\Lambda_A^2 - m_\pi^2}{\Lambda_A^2}\right),$$

with $g_V = 1.0$, $g_A = 1.254$, $\kappa = \mu_p - \mu_n = 3.70$, $\Lambda_V = 0.850$ GeV, $\Lambda_A = 1.086$ GeV, and m_π is the pion mass.

Substituting Eqs. (5)–(11) in Eq. (3), there is one term, associated with h_F [Eq. (9)] contributing to M_{Fh} , while $M_{GT h}$ has four terms, denoted by M_{GT-AA} , M_{GT-AP} , M_{GT-PP} , and M_{GT-MM} , which correspond to the four terms in h_{GT} [Eq. (10)]. The tensor contribution M_{Th} has three terms, denoted by M_{T-AP} , M_{T-PP} , and M_{T-MM} , which correspond to the three terms in h_T [Eq. (11)]. Their contributions to the total nuclear matrix element are discussed in Sec. III.

The short-range correlations (SRC) arise mainly from the repulsive nucleon-nucleon potential from the exchange of ρ and ω mesons and have been incorporated by using the effective transition operator [42], the exchange of ω meson [43], UCOM [26,44], and the self-consistent CCM [27]. The SRC can also be incorporated phenomenologically by the Jastrow type of correlations with Miller-Spencer parametrization [45]. Furthermore, it was shown in the self-consistent CMM [27] that the SRC effects of the Argonne and CD-Bonn two-nucleon potentials are weak and it is possible to parametrize them by the Jastrow type of correlations within a few percent accuracy.

Explicitly,

$$f(r) = 1 - ce^{-ar^2}(1 - br^2), \quad (13)$$

where $a = 1.1, 1.59, \text{ and } 1.52 \text{ fm}^{-2}$, $b = 0.68, 1.45, \text{ and } 1.88 \text{ fm}^{-2}$, and $c = 1.0, 0.92, \text{ and } 0.46$ for Miller-Spencer parametrization, Argonne V18, and CD-Bonn NN potentials, respectively. In this work the NTMEs $M_N^{(0\nu)}$ are calculated in the PHFB model for the above-mentioned three sets of parameters for the SRC, denoted as SRC1, SRC2, and SRC3, respectively.

In Fig. 1, we plot the neutrino potential $H_N(r, \Lambda) = H_{Fh}(r, \Lambda) f(r)$ with the three different parametrizations of SRC. It is noticed that the potentials from FNS and FNS + SRC3 are peaked at the origin whereas the peaks from FNS + SRC1 and FNS + SRC2 are at $r \approx 0.6$ fm and $r \approx 0.5$ fm, respectively. The shapes of these functions have a definite influence on the radial evolution of NTMEs $M_N^{(0\nu)}$ for $(\beta^-\beta^-)_{0\nu}$ decay from the exchange of heavy Majorana neutrino as discussed in Sec. III.

The calculation of $M_N^{(0\nu)}$ in the PHFB model was discussed in our earlier work [22,37] and one obtains the following expression for NTMEs $M_\alpha^{(0\nu)}$ of $(\beta^-\beta^-)_{0\nu}$ decay [37],

$$M_\alpha^{(0\nu)} = [n^{J_i=0} n^{J_f=0}]^{-1/2} \int_0^\pi n_{(Z,N),(Z+2,N-2)}(\theta) \sum_{\alpha\beta\gamma\delta} (\alpha\beta | O_\alpha | \gamma\delta) \\ \times \sum_{\varepsilon\eta} \frac{(f_{Z+2,N-2}^{(\pi)*})_{\varepsilon\beta}}{[(1 + F_{Z,N}^{(\pi)}(\theta)) f_{Z+2,N-2}^{(\pi)*}]_{\varepsilon\alpha}} \\ \times \frac{(F_{Z,N}^{(v)})_{\eta\delta}}{[(1 + F_{Z,N}^{(v)}(\theta)) f_{Z+2,N-2}^{(v)*}]_{\gamma\eta}} \sin\theta d\theta, \quad (14)$$

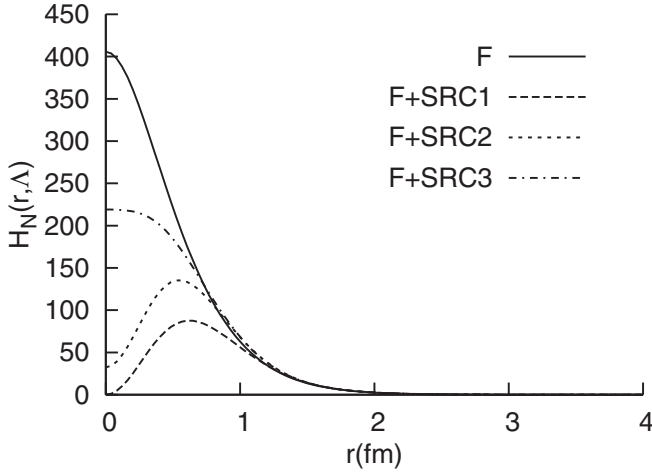


FIG. 1. Radial dependence of $H_N(r, \Lambda) = H_{Fh}(r, \Lambda) f(r)$ for the three different parametrizations of the SRC. In the case of FNS, $f(r) = 1$.

and the expressions for calculating n^J , $n_{(Z,N),(Z+2,N-2)}(\theta)$, $f_{Z,N}$, and $F_{Z,N}(\theta)$ are given in Refs. [22,37].

The calculation of matrices $f_{Z,N}$ and $F_{Z,N}(\theta)$ requires the amplitudes (u_{im} , v_{im}) and expansion coefficients $C_{ij,m}$, which specify the axially symmetric HFB intrinsic state $|\Phi_0\rangle$ with $K = 0$. Presently, they are obtained by carrying out the HFB calculations through the minimization of the expectation value of the effective Hamiltonian given by [23]

$$H = H_{sp} + V(P) + V(QQ) + V(HH), \quad (15)$$

where H_{sp} , $V(P)$, $V(QQ)$, and $V(HH)$ denote the single-particle Hamiltonian, the pairing, quadrupole-quadrupole, and hexadecapole-hexadecapole part of the effective two-body interaction, respectively. The HH part of the effective interaction $V(HH)$ is written as [23]

$$V(HH) = -\left(\frac{\chi_4}{2}\right) \sum_{\alpha\beta\gamma\delta} \sum_{\nu} (-1)^{\nu} \langle \alpha | r^4 Y_{4,\nu}(\theta, \phi) | \gamma \rangle \times \langle \beta | r^4 Y_{4,-\nu}(\theta, \phi) | \delta \rangle a_{\alpha}^{\dagger} a_{\beta}^{\dagger} a_{\delta} a_{\gamma}, \quad (16)$$

with $\chi_4 = 0.2442 \chi_2 A^{-2/3} b^{-4}$ for $T = 1$, and twice of this value for $T = 0$ case, following Bohr and Mottelson [46].

In Refs. [22,35,36], the strengths of the like particle components χ_{pp} and χ_{nn} of the QQ interaction were kept fixed. The strength of proton-neutron (pn) component χ_{pn} was varied so as to reproduce the excitation energy of the 2_2^+ state $E_{2_2^+}$ for the considered nuclei, namely $^{94,96}\text{Zr}$, $^{94,96,98,100}\text{Mo}$, $^{98,100,104}\text{Ru}$, $^{104,110}\text{Pd}$, ^{110}Cd , $^{128,130}\text{Te}$, $^{128,130}\text{Xe}$, ^{150}Nd , and ^{150}Sm as closely as possible to the experimental values. This is denoted as $PQQ1$ parametrization. Alternatively, one can employ a different parametrization of the χ_{2pn} , namely $PQQ2$ by taking $\chi_{2pp} = \chi_{2nn} = \chi_{2pn}/2$ and the excitation energy $E_{2_2^+}$ can be reproduced by varying the χ_{2pp} . Adding the HH part of the two-body interaction to $PQQ1$ and $PQQ2$ and by repeating the calculations, two more parametrizations of

the effective two-body interactions, namely $PQQH1$ and $PQQH2$ were obtained [37].

The four different parametrizations of the effective pairing plus multipolar correlations provide us four different sets of wave functions. With three different parametrizations of the Jastrow type of SRC and four sets of wave functions, sets of 12 NTMEs $M_N^{(0\nu)}$ are calculated for estimating the associated uncertainties in the present work. The uncertainties associated with the NTMEs $M_N^{(0\nu)}$ for $(\beta^-\beta^-)_{0\nu}$ decay are estimated statistically by calculating the mean and the standard deviation

TABLE I. Calculated NTMEs $M_N^{(0\nu)}$ in the PHFB model with four different parametrizations of effective two-body interaction and three different parametrizations of the Jastrow type of SRC for the $(\beta^-\beta^-)_{0\nu}$ decay of $^{94,96}\text{Zr}$, $^{98,100}\text{Mo}$, ^{104}Ru , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes from the exchange of heavy Majorana neutrino exchange. (a), (b), (c), and (d) denote $PQQ1$, $PQQH1$, $PQQ2$, and $PQQH2$ parametrizations, respectively. See the footnote on p. 3 of Ref. [37] for further details.

Nuclei		F	$F + S$		
			SRC1	SRC2	SRC3
^{94}Zr	(a)	236.9498	77.5817	138.2606	191.3897
	(b)	220.3794	72.4285	128.7496	178.0783
	(c)	205.8370	72.9303	124.3248	168.5705
	(d)	211.0437	68.9323	122.9710	170.3572
^{96}Zr	(a)	177.7479	56.4909	102.4434	142.8831
	(b)	185.5251	59.5338	107.2877	149.3117
	(c)	170.8199	54.2382	98.4051	137.2870
	(d)	175.4730	56.0746	101.2963	141.1240
^{98}Mo	(a)	355.1915	117.0804	208.2494	287.5615
	(b)	346.1118	116.4967	204.5667	281.0515
	(c)	358.5109	118.0563	210.1150	290.2080
	(d)	343.4160	115.2077	202.6977	278.7158
^{100}Mo	(a)	365.8004	122.2000	215.8882	296.9869
	(b)	361.9877	122.6611	214.7455	294.4297
	(c)	368.4056	123.2364	217.5391	299.1598
	(d)	328.9795	111.4464	195.1601	267.5869
^{104}Ru	(a)	274.0700	89.7666	160.7925	222.1151
	(b)	264.9015	88.1515	156.2893	215.1076
	(c)	258.2796	84.6746	151.6002	209.3600
	(d)	247.0603	82.3208	145.8435	200.6645
^{110}Pd	(a)	424.6601	140.3359	249.6835	344.3187
	(b)	379.9404	127.4915	224.6563	308.6907
	(c)	407.2163	134.6824	239.4733	330.1888
	(d)	390.3539	130.6314	230.5392	316.9996
^{128}Te	(a)	190.5325	62.4373	111.5143	154.1796
	(b)	231.8024	77.4559	136.7936	188.1893
	(c)	220.7156	73.5158	130.0810	179.0960
	(d)	235.4814	78.6367	138.9052	191.1366
^{130}Te	(a)	236.0701	81.5493	141.3447	192.7610
	(b)	231.5921	79.3844	138.1901	188.8492
	(c)	233.0024	80.4020	139.4400	190.2194
	(d)	230.5282	78.9888	137.5307	187.9675
^{150}Nd	(a)	163.8037	55.8968	97.8169	133.6912
	(b)	130.1364	43.8840	77.3178	105.9993
	(c)	160.2720	54.6713	95.6942	130.8005
	(d)	131.9781	44.6741	78.5433	107.5715

defined by

$$\overline{M}_N^{(0\nu)} = \frac{\sum_{i=1}^k M_N^{(0\nu)}(i)}{N}, \quad (17)$$

and

$$\Delta \overline{M}_N^{(0\nu)} = \frac{1}{\sqrt{N-1}} \left[\sum_{i=1}^N (\overline{M}_N^{(0\nu)} - M_N^{(0\nu)}(i))^2 \right]^{1/2}. \quad (18)$$

III. RESULTS AND DISCUSSIONS

The model space, SPEs, parameters of the PQQ type of effective two-body interactions, and the method to fix them have already been given in Refs. [22,35,36]. It turns out that with $PQQ1$ and $PQQ2$ parametrizations, the experimental excitation energies of the 2^+ state E_{2^+} [47] can be reproduced within about 2% accuracy. The electromagnetic properties, namely reduced $B(E2:0^+ \rightarrow 2^+)$ transition probabilities, deformation parameters β_2 , static quadrupole moments $Q(2^+)$, and gyromagnetic factors $g(2^+)$ are in overall agreement with the experimental data [48,49].

A. Short-range correlations and radial evolutions of NTMEs

In the approximation of the finite size of nucleons with the dipole form factor (F) and finite size plus SRC ($F + S$), the theoretically calculated 12 NTMEs $M_N^{(0\nu)}$ using the four sets of HFB wave functions generated with $PQQ1$, $PQQHH1$, $PQQ2$, and $PQQHH2$ parametrizations of the effective two-body interaction and three different parametrizations of the Jastrow type of SRC for $^{94,96}\text{Zr}$, $^{98,100}\text{Mo}$, ^{104}Ru , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes are presented in Table I.

To analyze the role of different components of NTME $M_N^{(0\nu)}$, the decomposition of the latter into Fermi, different terms of Gamow-Teller, and tensor matrix elements of ^{100}Mo are presented in Table II for $PQQ1$ parametrization. From the inspection of Table II, the following observations emerge.

- (i) The contribution of conventional Fermi matrix elements $M_{Fh} = M_{F-VV}$ is about 20% to the total matrix element.
- (ii) The Gamow-Teller matrix element is noticeably modified by the inclusion of the pseudoscalar and weak magnetism terms in the hadronic currents. While M_{GT-PP} increases the absolute value of M_{GT-AA} , M_{GT-AP} has a significant contribution with the opposite sign in all cases. The term M_{GT-MM} is smaller than others, and the introduction of short-range correlations changes its sign.
- (iii) The tensor matrix elements have a very small contribution, smaller than 2%, to the total transition matrix elements.
- (iv) The inclusion of short-range correlations changes the nuclear matrix elements significantly, whose effects are large for the Gamow-Teller and Fermi matrix elements but small in the case of the tensor ones.
- (v) The Miller-Spencer parametrization of the short-range correlations, SRC1, cancels out a large part of the radial function H_N , as shown in Fig. 1. The same cancellation reduces the calculated matrix element to about one-third of its original value. The other two parametrizations of the short-range correlations, namely SRC2 and SRC3, have a sizable effect, which is in all cases much smaller than SRC1.

With respect to the point nucleon case, the change in $M_N^{(0\nu)}$ is about 30%–34% because of the FNS. With the inclusion of effects from FNS and SRC, the NTMEs change by about 75%–79%, 58%–62%, and 43%–47% for $F + \text{SRC1}$, $F + \text{SRC2}$, and $F + \text{SRC3}$, respectively. It is noteworthy that the SRC3 has practically negligible effect on the finite-size case. Furthermore, the maximum variation in $M_N^{(0\nu)}$ from $PQQHH1$, $PQQ2$, and $PQQHH2$ parametrization with respect to $PQQ1$ interaction is about 24%, 18%, and 26%, respectively.

In the QRPA [26,27], ISM [16], and PHFB [28,37], the radial evolution of $M^{(0\nu)}$ from the exchange of light Majorana neutrino was already studied. In both QRPA and

TABLE II. Decomposition of NTMEs for the $(\beta^- \beta^-)_{0\nu}$ decay of ^{100}Mo including finite size effect (F) and SRC ($F + S$) for the $PQQ1$ parametrization.

NTMEs	F	$F + S$		
		SRC1	SRC2	SRC3
M_F	68.6223	35.8191	54.0101	64.2516
M_{GT-AA}	-370.5960	-144.5650	-242.3340	-316.3250
M_{GT-AP}	174.4640	43.3631	93.9737	137.5100
M_{GT-PP}	-66.3082	-8.3767	-28.5936	-48.0727
M_{GT-MM}	-41.7693	16.3949	7.6213	-13.3421
M_{GT}	-304.2095	-93.1837	-169.3326	-240.2298
M_{T-AP}	9.4369	9.2393	10.0332	10.0610
M_{T-PP}	-3.6622	-3.5528	-3.9226	-3.9386
M_{T-MM}	1.2567	1.1163	1.3438	1.3722
M_T	7.0314	6.8028	7.4544	7.4945
$ M_N^{(0\nu)} $	365.8004	122.2000	215.8882	296.9869

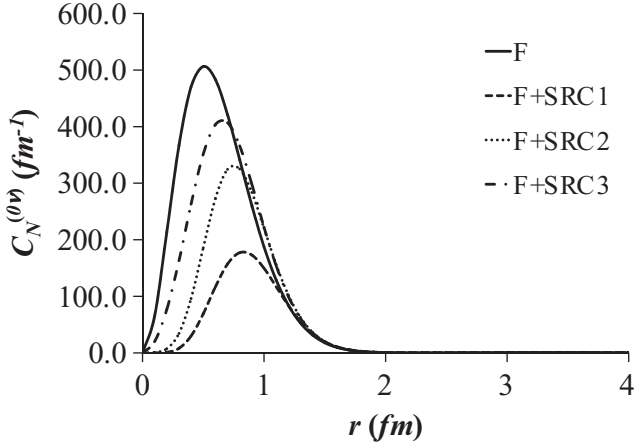


FIG. 2. Radial dependence of $C_N^{(0v)}(r)$ for the $(\beta^-\beta^-)_{0v}$ decay of ^{100}Mo isotope.

ISM calculations, it was established that the contributions of decaying pairs coupled to $J = 0$ and $J > 0$ almost cancel beyond $r \approx 3$ fm and the magnitude of $C^{(0v)}$ for all nuclei undergoing $(\beta^-\beta^-)_{0v}$ decay have their maximum at about the internucleon distance $r \approx 1$ fm. These observations were also made in the PHFB model [28,37]. Similarly, the radial evolution of $M_N^{(0v)}$ can be studied by defining

$$M_N^{(0v)} = \int C_N^{(0v)}(r) dr. \quad (19)$$

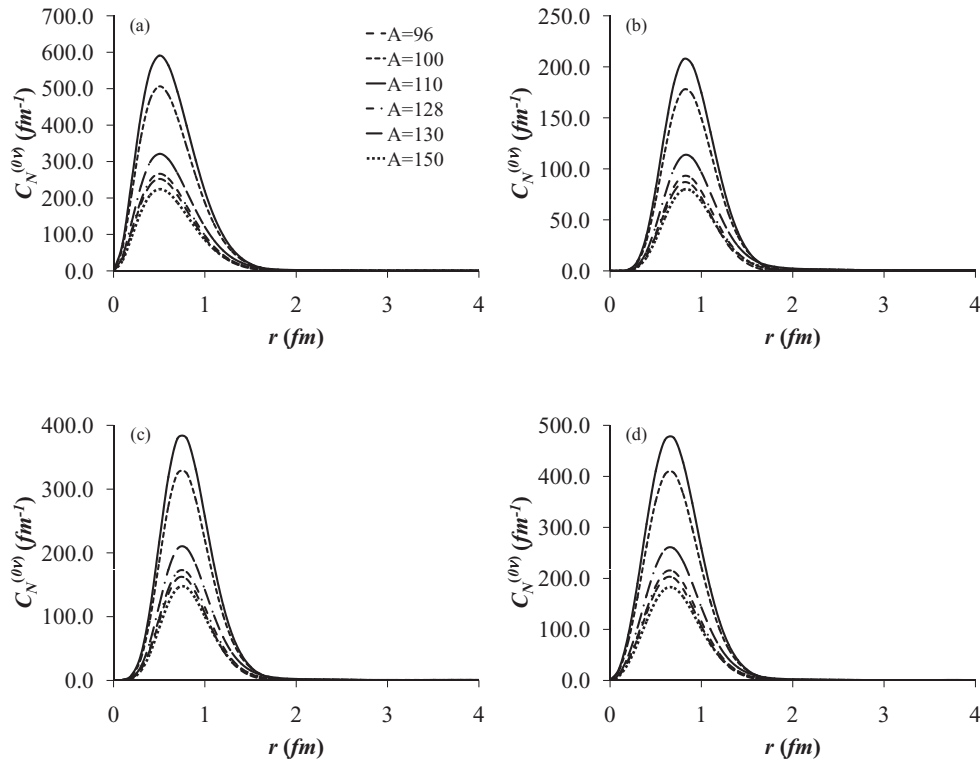


FIG. 3. Radial dependence of $C_N^{(0v)}(r)$ for the $(\beta^-\beta^-)_{0v}$ decay of ^{96}Zr , ^{100}Mo , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes. In this figure, (a), (b), (c), and (d) correspond to F , $F + \text{SRC1}$, $F + \text{SRC2}$, and $F + \text{SRC3}$, respectively.

The radial evolution of $M_N^{(0v)}$ was studied for four cases, namely F , $F + \text{SRC1}$, $F + \text{SRC2}$, and $F + \text{SRC3}$. To make the effects of finite size and SRC more transparent, we plot them for ^{100}Mo in Fig. 2. In the case of finite-sized nucleons, the $C_N^{(0v)}$ are peaked at $r \approx 0.5$ fm and with the addition of SRC1 and SRC2, the peak shifts to about 0.8 fm. However, the position of peak is shifted to 0.7 fm for SRC3. In Fig. 3, we plot the radial dependence of $C_N^{(0v)}$ for six nuclei, namely ^{96}Zr , ^{100}Mo , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd and the same observations remain valid. In addition, the same features in the radial distribution of $C_N^{(0v)}$ are noticed in the cases of $PQQ2$, $PQQHH1$, and $PQQHH2$ parametrizations.

B. Uncertainties in NTMEs

The uncertainties associated with the NTMEs $M_N^{(0v)}$ for $(\beta^-\beta^-)_{0v}$ decay are estimated by performing a statistical analysis by using Eqs. (17) and (18). In Table I, sets of 12 NTMEs $M_N^{(0v)}$ of $^{94,96}\text{Zr}$, $^{98,100}\text{Mo}$, ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes are displayed, which are employed to calculate the average values $\overline{M}_N^{(0v)}$ as well as uncertainties $\Delta\overline{M}_N^{(0v)}$ tabulated in Table III for the bare axial vector coupling constant $g_A = 1.254$ and quenched value of $g_A = 1.0$.

It turns out that in all cases, the uncertainties $\Delta\overline{M}_N^{(0v)}$ are about 35% for $g_A = 1.254$ and $g_A = 1.0$. Furthermore, we estimate the uncertainties for eight NTMEs $\overline{M}_N^{(0v)}$ calculated using the SRC2, and SRC3 parametrizations and the uncertainties in NTMEs reduce to about 16%–20% with

TABLE III. Average NTMEs $\overline{M}_N^{(0\nu)}$ and uncertainties $\Delta\overline{M}_N^{(0\nu)}$ for the $(\beta^-\beta^-)_{0\nu}$ decay of $^{94,96}\text{Zr}$, $^{98,100}\text{Mo}$, ^{104}Ru , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes. Both bare and quenched values of g_A are considered. Case I and Case II denote calculations with and without SRC1, respectively.

$\beta^-\beta^-$ emitters	g_A	Case I		Case II	
		$\overline{M}_N^{(0\nu)}$	$\Delta\overline{M}_N^{(0\nu)}$	$\overline{M}_N^{(0\nu)}$	$\Delta\overline{M}_N^{(0\nu)}$
^{94}Zr	1.254	126.2146	44.9489	152.8378	27.1912
	1.0	142.9381	49.1752	172.1620	29.3965
^{96}Zr	1.254	100.5313	36.8858	122.5048	21.9209
	1.0	114.4851	40.3246	138.6328	23.5263
^{98}Mo	1.254	202.5006	71.6345	245.3957	41.8882
	1.0	230.1520	78.3244	277.2795	44.9878
^{100}Mo	1.254	206.7533	73.0792	250.1870	43.7119
	1.0	235.0606	79.9885	282.7964	47.1334
^{104}Ru	1.254	150.5572	53.9389	182.7216	31.9382
	1.0	171.8075	59.0467	207.1750	34.3939
^{110}Pd	1.254	231.4743	82.4924	280.5688	49.1588
	1.0	263.4339	90.3033	317.3947	53.0150
^{128}Te	1.254	126.8285	46.3381	153.7370	29.4676
	1.0	143.9772	50.6942	173.5263	31.8554
^{130}Te	1.254	136.3856	46.9164	164.5378	27.2226
	1.0	154.3797	51.2511	185.2849	29.1907
^{150}Nd	1.254	85.5467	31.4473	103.4294	20.9802
	1.0	97.3640	34.5024	117.0160	22.8729

the exclusion of the Miller-Spencer type of parametrization. In Table IV, average NTMEs for case II along with NTMEs calculated in other models have been presented. It is noteworthy that in the models employed in Refs. [6,8,9],

TABLE IV. Average NTMEs $\overline{M}_N^{(0\nu)'} (= (g_A/1.254)^2 \overline{M}_N^{(0\nu)})$ for the $(\beta^-\beta^-)_{0\nu}$ decay of $^{94,96}\text{Zr}$, $^{98,100}\text{Mo}$, ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes. Both bare and quenched values of g_A are considered. The superscripts a and b denote the Argonne and CD-Bonn potentials.

$\beta^-\beta^-$ emitters	g_A	$\overline{M}_N^{(0\nu)'}$	QRPA [6]	QRPA [8]	QRPA [9]	QRPA [10]	SRQRPA ^a [3]	SRQRPA ^b [3]	$T_{1/2}^{0\nu}$ (yr)	Ref.	$\langle m_N \rangle$ (GeV)
^{94}Zr	1.254	152.84 ± 27.19							1.9×10^{19}	[50]	$2.57^{+0.46}_{-0.46} \times 10^4$
	1.0	109.48 ± 18.69									$1.84^{+0.31}_{-0.31} \times 10^4$
^{96}Zr	1.254	122.50 ± 21.92				99.062			9.2×10^{21}	[51]	$2.68^{+0.48}_{-0.48} \times 10^6$
	1.0	88.16 ± 14.96									$1.93^{+0.33}_{-0.33} \times 10^6$
^{98}Mo	1.254	245.40 ± 41.89							1.0×10^{14}	[52]	$9.70^{+1.70}_{-1.70}$
	1.0	176.33 ± 28.61									$6.97^{+1.13}_{-1.13}$
^{100}Mo	1.254	250.19 ± 43.71	155.960	333.0	56.914	76.752	259.8	404.3	4.6×10^{23}	[53]	$3.43^{+0.60}_{-0.60} \times 10^7$
	1.0	179.84 ± 29.97					191.8	310.5			$2.47^{+0.41}_{-0.41} \times 10^7$
^{110}Pd	1.254	280.57 ± 49.16							6.0×10^{17}	[54]	$2.43^{+0.43}_{-0.43} \times 10^4$
	1.0	201.84 ± 33.71									$1.75^{+0.29}_{-0.29} \times 10^4$
^{128}Te	1.254	153.74 ± 29.47	122.669	303.0		101.233			1.1×10^{23}	[55]	$2.06^{+0.39}_{-0.39} \times 10^6$
	1.0	110.35 ± 20.26									$1.48^{+0.27}_{-0.27} \times 10^6$
^{130}Te	1.254	164.54 ± 27.22	108.158	267.0		92.661	239.7	384.5	3.0×10^{24}	[56]	$5.67^{+0.94}_{-0.94} \times 10^7$
	1.0	117.83 ± 18.56					176.5	293.8			$4.06^{+0.64}_{-0.64} \times 10^7$
^{150}Nd	1.254	103.43 ± 20.98	153.085	422.0					1.8×10^{22}	[57]	$5.99^{+1.21}_{-1.21} \times 10^6$
	1.0	74.41 ± 14.55									$4.31^{+0.84}_{-0.84} \times 10^6$

effects from higher order currents have not been included. We also extract lower limits on the effective mass of heavy Majorana neutrino $\langle M_N \rangle$ from the largest observed limits on half-lives $T_{1/2}^{0\nu}$ of $(\beta^-\beta^-)_{0\nu}$ decay. The extracted limits are $\langle M_N \rangle > 5.67^{+0.94}_{-0.94} \times 10^7$ GeV and $> 4.06^{+0.64}_{-0.64} \times 10^7$ GeV, from the limit on half-life $T_{1/2}^{0\nu} > 3.0 \times 10^{24}$ yr of ^{130}Te [56] for $g_A = 1.254$ and $g_A = 1.0$, respectively.

IV. CONCLUSIONS

We have employed the PHFB model, with four different parametrizations of pairing plus multipole effective two-body interaction, to generate sets of four HFB intrinsic wave functions, which reasonably reproduced the observed spectroscopic properties, namely the yrast spectra, reduced $B(E2:0^+ \rightarrow 2^+)$ transition probabilities, static quadrupole moments $Q(2^+)$, and g factors $g(2^+)$ of participating nuclei in $(\beta^-\beta^-)_{2\nu}$ decay, as well as their $M_{2\nu}$ [35,36]. Considering three different parametrizations of the Jastrow type of SRC, sets of 12 NTMEs $M_N^{(0\nu)}$ for the study $(\beta^-\beta^-)_{0\nu}$ decay of $^{94,96}\text{Zr}$, $^{98,100}\text{Mo}$, ^{104}Ru , ^{110}Pd , $^{128,130}\text{Te}$, and ^{150}Nd isotopes in the heavy Majorana neutrino mass mechanism have been calculated.

The study of effects because of finite size of nucleons and SRC reveal that in the case of heavy Majorana neutrino exchange, the NTMEs change by about 30%–34% from finite size of nucleons and the SRC1, SRC2, and SRC3 change them by 75%–79%, 58%–62%, and 43%–47%, respectively. Furthermore, it was noticed through the study of radial evolution of NTMEs that the FNS and SRC play a more crucial role in the heavy than in the light Majorana neutrino exchange mechanism.

Finally, a statistical analysis was performed by employing the sets of 12 NTMEs $M_N^{(0\nu)}$ to estimate the uncertainties for $g_A = 1.254$ and $g_A = 1.0$. It turns out that the uncertainties are about 35% for all the considered nuclei. Exclusion of the Miller-Spencer parametrization of the Jastrow type of SRC, reduces the maximum uncertainties to a value smaller than 20%. The best extracted limit on the effective heavy Majorana neutrino mass $\langle M_N \rangle$ from the available limits on experimental half-lives $T_{1/2}^{0\nu}$ using average NTMEs $\overline{M}_N^{(0\nu)}$ calculated in the PHFB model is $>5.67_{-0.94}^{+0.94} \times 10^7$ GeV and $>4.06_{-0.64}^{+0.64} \times 10^7$ GeV for the ^{130}Te isotope.

ACKNOWLEDGMENTS

This work is partially supported by the Department of Science and Technology (DST), India vide sanction No. SR/S2/HEP-13/2006, DST-RFBR Collaboration via Grant No. RUSP-935, Consejo Nacional de Ciencia y Tecnología (Conacyt)-México, European Union-Mexico Science and Technology International Cooperation Fund (FONCICYT) Project No. 94142, and Dirección General de Asuntos del Personal Académico, Universidad Nacional Autónoma de México (DGAPA-UNAM) Project No. IN102109-3.

-
- [1] H. V. Klapdor-Kleingrothaus, I. V. Krivosheina, and I. V. Titkova, *Int. J. Mod. Phys. A* **21**, 1159 (2006).
- [2] F. T. Avignone, S. R. Elliott, and J. Engel, *Rev. Mod. Phys.* **80**, 481 (2008).
- [3] F. Simkovic, J. D. Vergados, and A. Faessler, *Phys. Rev. D* **82**, 113015 (2010); A. Faessler, A. Meroni, S. T. Petcov, F. Simkovic, and J. Vergados, *ibid.* **83**, 113003 (2011).
- [4] V. Tello, M. Nemevsek, F. Nesti, G. Senjanovic, and F. Vissani, *Phys. Rev. Lett.* **106**, 151801 (2011).
- [5] J. D. Vergados, *Phys. Rep.* **133**, 1 (1986).
- [6] T. Tomoda, *Rep. Prog. Phys.* **54**, 53 (1991).
- [7] M. Doi and T. Kotani, *Prog. Theor. Phys.* **89**, 139 (1993).
- [8] M. Hirsch, H. V. Klapdor-Kleingrothaus, and O. Panella, *Phys. Lett. B* **374**, 7 (1996).
- [9] G. Pantis and J. D. Vergados, *Phys. Rep.* **242**, 285 (1994).
- [10] G. Pantis, F. Simkovic, J. D. Vergados, and A. Faessler, *Phys. Rev. C* **53**, 695 (1996).
- [11] F. Simkovic, G. V. Efimov, M. A. Ivanov, and V. E. Lyubovitskij, *Z. Phys. A* **341**, 193 (1992).
- [12] F. Simkovic, M. Nowak, W. A. Kaminski, A. A. Raduta, and A. Faessler, *Phys. Rev. C* **64**, 035501 (2001).
- [13] E. Caurier, A. Poves, and A. P. Zuker, *Phys. Lett. B* **252**, 13 (1990); E. Caurier, F. Nowacki, A. Poves, and J. Retamosa, *Phys. Rev. Lett.* **77**, 1954 (1996); *Nucl. Phys. A* **654**, 973c (1999).
- [14] P. Vogel and M. R. Zirnbauer, *Phys. Rev. Lett.* **57**, 3148 (1986).
- [15] O. Civitarese, A. Faessler, and T. Tomoda, *Phys. Lett. B* **194**, 11 (1987).
- [16] E. Caurier, J. Menéndez, F. Nowacki, and A. Poves, *Phys. Rev. Lett.* **100**, 052503 (2008).
- [17] J. Menéndez, A. Poves, E. Caurier, and F. Nowacki, [arXiv:0809.2183](https://arxiv.org/abs/0809.2183) [nucl-th].
- [18] L. Paceaescu, A. Faessler, and F. Simkovic, *Phys. At. Nucl.* **67**, 1210 (2004).
- [19] R. Álvarez-Rodríguez, P. Sarriguren, E. Moya de Guerra, L. Paceaescu, A. Faessler, and F. Simkovic, *Phys. Rev. C* **70**, 064309 (2004).
- [20] M. S. Yousef, V. Rodin, A. Faessler, and F. Simkovic, *Phys. Rev. C* **79**, 014314 (2009).
- [21] D. L. Fang, A. Faessler, V. Rodin, and F. Simkovic, *Phys. Rev. C* **82**, 051301(R) (2010).
- [22] K. Chaturvedi, R. Chandra, P. K. Rath, P. K. Raina, and J. G. Hirsch, *Phys. Rev. C* **78**, 054302 (2008).
- [23] R. Chandra, K. Chaturvedi, P. K. Rath, P. K. Raina, and J. G. Hirsch, *Europhys. Lett.* **86**, 32001 (2009).
- [24] F. Simkovic, G. Pantis, J. D. Vergados, and A. Faessler, *Phys. Rev. C* **60**, 055502 (1999).
- [25] J. D. Vergados, *Phys. Rep.* **361**, 1 (2002).
- [26] F. Simkovic, A. Faessler, V. Rodin, P. Vogel, and J. Engel, *Phys. Rev. C* **77**, 045503 (2008).
- [27] F. Simkovic, A. Faessler, H. Mütter, V. Rodin, and M. Stauf, *Phys. Rev. C* **79**, 055501 (2009).
- [28] P. K. Rath, R. Chandra, K. Chaturvedi, P. K. Raina, and J. G. Hirsch, *Phys. Rev. C* **80**, 044303 (2009).
- [29] P. Vogel, in *Current Aspects of Neutrino Physics*, edited by D. O. Caldwell (Springer, New York, 2001), Chap. 8, p. 177.
- [30] S. M. Bilenky and J. A. Grifols, *Phys. Lett. B* **550**, 154 (2002).
- [31] J. N. Bahcall, H. Murayama, and C. Peña-Garay, *Phys. Rev. D* **70**, 033012 (2004).
- [32] F. T. Avignone III, G. S. King III, and Yu. G. Zdesenko, *New J. Phys.* **7**, 6 (2005).
- [33] V. A. Rodin, A. Faessler, F. Simkovic, and P. Vogel, *Phys. Rev. C* **68**, 044302 (2003).
- [34] M. Baranger and K. Kumar, *Nucl. Phys. A* **110**, 490 (1968).
- [35] R. Chandra, J. Singh, P. K. Rath, P. K. Raina, and J. G. Hirsch, *Eur. Phys. J. A* **23**, 223 (2005).
- [36] S. Singh, R. Chandra, P. K. Rath, P. K. Raina, and J. G. Hirsch, *Eur. Phys. J. A* **33**, 375 (2007).
- [37] P. K. Rath, R. Chandra, K. Chaturvedi, P. K. Raina, and J. G. Hirsch, *Phys. Rev. C* **82**, 064310 (2010).
- [38] M. Horoi and S. Stoica, *Phys. Rev. C* **81**, 024321 (2010).
- [39] J. Barea and F. Iachello, *Phys. Rev. C* **79**, 044301 (2009).
- [40] T. R. Rodríguez and G. Martínez-Pinedo, *Phys. Rev. Lett.* **105**, 252503 (2010).
- [41] J. Suhonen and O. Civitarese, *Phys. Lett. B* **668**, 277 (2008).
- [42] H. F. Wu, H. Q. Song, T. T. S. Kuo, W. K. Cheng, and D. Strottman, *Phys. Lett. B* **162**, 227 (1985).
- [43] J. G. Hirsch, O. Castaños, and P. O. Hess, *Nucl. Phys. A* **582**, 124 (1995).
- [44] M. Kortelainen and J. Suhonen, *Phys. Rev. C* **76**, 024315 (2007); M. Kortelainen, O. Civitarese, J. Suhonen, and J. Toivanen, *Phys. Lett. B* **647**, 128 (2007).
- [45] G. A. Miller and J. E. Spencer, *Ann. Phys. (NY)* **100**, 562 (1976).
- [46] A. Bohr and B. R. Mottelson, *Nuclear Structure*, Vol. I (World Scientific, Singapore, 1998).
- [47] M. Sakai, *At. Data Nucl. Data Tables* **31**, 399 (1984).
- [48] P. Raghavan, *At. Data Nucl. Data Tables* **42**, 189 (1989).
- [49] S. Raman, C. W. Nestor Jr., and P. Tikkanen, *At. Data Nucl. Data Tables* **78**, 1 (2001).

- [50] R. Arnold *et al.*, *Nucl. Phys. A* **658**, 299 (1999).
[51] J. Argyriades *et al.*, *Nucl. Phys. A* **847**, 168 (2010).
[52] J. H. Fremlin and M. C. Walters, *Proc. Phys. Soc. London, Sect. A* **65**, 911 (1952).
[53] R. Arnold *et al.*, *Phys. Rev. Lett.* **95**, 182302 (2005).
[54] R. G. Winter, *Phys. Rev.* **85**, 687 (1952).
[55] C. Arnaboldi *et al.*, *Phys. Lett. B* **557**, 167 (2003).
[56] C. Arnaboldi *et al.*, *Phys. Rev. C* **78**, 035502 (2008).
[57] J. Argyriades *et al.*, *Phys. Rev. C* **80**, 032501(R) (2009).