

## Production of spectator hypermatter in relativistic heavy-ion collisions

A. S. Botvina,<sup>1,2</sup> K. K. Gudima,<sup>1,3</sup> J. Steinheimer,<sup>1</sup> M. Bleicher,<sup>1</sup> and I. N. Mishustin<sup>1,4</sup>

<sup>1</sup>*Frankfurt Institute for Advanced Studies, J.W. Goethe University, D-60438 Frankfurt am Main, Germany*

<sup>2</sup>*Institute for Nuclear Research, Russian Academy of Sciences, 117312 Moscow, Russia*

<sup>3</sup>*Institute of Applied Physics, Academy of Sciences of Moldova, MD-2028 Kishinev, Moldova*

<sup>4</sup>*Kurchatov Institute, Russian Research Center, 123182 Moscow, Russia*

(Received 6 May 2011; revised manuscript received 17 October 2011; published 5 December 2011)

Possible formation of large hyperfragments in relativistic heavy-ion collisions is studied within two transport models, the Dubna cascade model and UrQMD model. Our goal is to explore a new mechanism for the formation of strange nuclear systems via capture of hyperons by relatively cold spectator matter produced in semiperipheral collisions. We investigate basic characteristics of the produced hyperspectators and estimate the production probabilities of multistrange systems. Advantages of the proposed mechanisms over an alternative coalescence process are analyzed. We also discuss how such hyperfragments can be detected by taking into account the background of free hyperons. This investigation is important for the development of new experimental methods for producing hypernuclei in peripheral relativistic nucleus-nucleus collisions, which are now underway at GSI and are planned for the future FAIR and NICA facilities.

DOI: [10.1103/PhysRevC.84.064904](https://doi.org/10.1103/PhysRevC.84.064904)

PACS number(s): 25.75.-q, 21.80.+a, 25.70.Mn, 13.75.Ev

### I. INTRODUCTION

Strange baryons (hyperons) were discovered in the 1950s in reactions induced by cosmic rays. In nuclear reactions at high energies strange particles (baryons and mesons) are produced abundantly, and they are strongly involved in the reaction dynamics. The specifics of hypernuclear physics is that there is no direct experimental way to study hyperon-nucleon ( $YN$ ) and hyperon-hyperon ( $YY$ ) interactions (where  $Y = \Lambda, \Sigma, \Xi, \text{ or } \Omega$ ). When hyperons are captured by nuclei, hypernuclei are produced; these can live long enough in comparison with nuclear reaction times. Therefore, a nucleus may serve as a laboratory offering a unique opportunity to study basic properties of hyperons and their interactions. Double- and multistrange nuclei are especially interesting, because they are more suitable for extracting information about the hyperon-hyperon interaction and strange matter properties.

The investigation of hypernuclei allows us to answer many fundamental questions. Here we mention only some of them. First of all, studying the structure of hypernuclei helps to understand the structure of conventional nuclei too [1]. These studies lead to an extension of the nuclear chart into the strangeness sector [2,3]. Second, hypernuclei provide a bridge between traditional nuclear physics (dealing with protons and neutrons) and hadron physics. Strangeness is an important degree of freedom for the construction of QCD-motivated models of strong interactions [4]. And last but not least, strange particles are abundantly produced in nuclear matter at high densities, which are realized in the core of neutron stars [5]. The only way to describe realistically these physical conditions is to study the hyperon interactions in the laboratory and to select theoretical models which pass the careful comparison with experimental data.

Traditionally, information about hypernuclear interactions is obtained from spectroscopy of hypernuclei combined with theoretical analyses. Production of kaons is often used for

triggering hypernuclei, and by using kaon beams one can produce double hypernuclei. The theoretical studies are mainly concentrated on calculating the structure of nearly cold hypernuclei with baryon density around the nuclear saturation density,  $\rho_0 \approx 0.15 \text{ fm}^{-3}$ . In this work we consider relativistic nucleus-nucleus collisions leading to copious production of hyperons and analyze new opportunities for hypernuclear physics realized in this case. As seen in experiments in the GeV domain of bombarding energies [6],  $\Lambda$  hyperons are produced mainly in the participant zone; however, they have a broad rapidity distribution, so that a certain fraction of them can even be found in the spectator kinematical region. Some of these  $\Lambda$  hyperons may be captured by nuclear spectator fragments produced in peripheral collisions, as was already discussed by using simplified coalescence-like approaches [7,8]. Recently, this possibility was clearly demonstrated within realistic dynamical calculations [9]. Indeed, experiments with light-ion beams at the Lawrence Berkeley National Laboratory (LBL) [10] and the Joint Institute for Nuclear Research (JINR) [11] have demonstrated that hypernuclei can be formed in such reactions. The production of large excited spectator residues is well established in relativistic heavy-ion collisions also (see, e.g., Ref. [12]). At a later stage, these excited residues undergo deexcitation via evaporation, fission, or multifragmentation [13]. Therefore, we expect that the capture of hyperons by spectators may lead to the formation of a big lump of excited matter containing a strangeness admixture. In the following, these excited spectators will break up into conventional and hyperfragments [14].

It should be noted that light nuclei and hypernuclei can be produced in central nucleus-nucleus collisions, as has been recently demonstrated in Relativistic Heavy Ion Collider (RHIC) experiments [15], as well as at lower energies (see, also Ref. [16]). However, because of the very high excitation energy released in the overlapping zone, it will be only possible to produce very light hypernuclei ( $A \lesssim 4$ ) in this way.

## II. MODELING RELATIVISTIC NUCLEUS-NUCLEUS COLLISIONS

### A. The Dubna cascade model with quark-gluon string approach

One of the first models designed to describe the dynamics of energetic heavy-ion collisions was the intranuclear cascade model developed in Dubna [17]. In the following we refer to it as the Dubna cascade model (DCM). The DCM is based on the Monte Carlo solution of a set of the Boltzmann-Uehling-Uhlenbeck relativistic kinetic equations with the collision terms, including cascade-cascade interactions. For particle energies below 1 GeV it is sufficient to consider only nucleons, pions, and deltas. The model includes a proper description of pion and baryon dynamics for particle production and absorption processes. In the original version the nuclear potential is treated dynamically, i.e., for the initial state it is determined using the Thomas-Fermi approximation, but later on its depth is changed according to the number of knocked-out nucleons. This allows one to account for nuclear binding. The Pauli principle is implemented by introducing a Fermi distribution of nucleon momenta as well as Pauli blocking factors for scattered nucleons.

At energies higher than about 10 GeV, the quark-gluon string model (QGSM) is used to describe elementary hadron collisions [18,19]. This model is based on the  $1/N_c$  expansion of the amplitude for binary processes, where  $N_c$  is the number of quark colors. Different terms of the  $1/N_c$  expansion correspond to different diagrams, which are classified according to their topological properties. Every diagram defines how many strings are created in a hadronic collision and which quark-antiquark or quark-diquark pairs form these strings. The relative contributions of different diagrams can be estimated within Regge theory, and all QGSM parameters for hadron-hadron collisions were fixed from the analysis of experimental data. The breakup of strings via creation of quark-antiquark and diquark-antidiquark pairs is described by the Field-Feynman method [20], using phenomenological functions for the fragmentation of quarks, antiquarks, and diquarks into hadrons. The modified non-Markovian relativistic kinetic equation, having a structure close to the Boltzmann-Uehling-Uhlenbeck kinetic equation, but accounting for the finite formation time of newly created hadrons, is used for simulations of relativistic nuclear collisions. One should note that QGSM considers the two lowest SU(3) multiplets in mesonic, baryonic, and antibaryonic sectors, so interactions between almost 70 hadron species are treated on the same footing. This is a great advantage of this approach which is important for the proper evaluation of the hadron abundances and characteristics of the excited residual nuclei. The above noted two energy extremes were bridged by the QGSM extension downward in the beam energy [21].

In the course of a nucleus-nucleus collision, strange particles are produced in both primary and secondary baryon and meson interactions ( $B + B \rightarrow BYK$ ,  $M + B \rightarrow YK$ ). The produced hyperons can propagate and rescatter on other particles, and, occasionally, they may be located inside the projectile or target spectators. These hyperons can be absorbed by the spectators if their kinetic energy in the rest frame of the

residual nucleus is lower than the attractive potential energy, i.e., the hyperon potential. In this case an excited residual system with nonzero strangeness will be formed. As known from previous studies the  $\Lambda$ -hyperon potential at the normal nuclear density  $V_\Lambda(\rho_0)$  is around  $-30$  MeV. In our simulations we calculate the local nucleon density  $\rho$  at the hyperon's position by taking into account only the nucleons in the vicinity of this hyperon, within a radius of 2 fm. This local density is then used to calculate the effective potential  $V_\Lambda(\rho)$ . Usually this potential is softer than the potential in normal nuclear matter, since spectators are quite dilute after primary interactions. The density dependence of this potential is parametrized following Ref. [22]:

$$V_\Lambda(\rho) = -\alpha \frac{\rho}{\rho_0} \left[ 1 - \beta \left( \frac{\rho}{\rho_0} \right)^{2/3} \right], \quad (1)$$

where  $\alpha = 57.5$  MeV and  $\beta = 0.522$ . In our calculations we follow the propagation of each  $\Lambda$  hyperon during the whole reaction time, up to about 100 fm/c. One possibility for treatment of the absorption is to fix the position and time when this absorption first happened and to remove the hyperon from simulations at the later times. We call this the first-capture criterion. However, the rescattering processes proceed further in time, and after this absorption some surrounding nucleons may escape from the spectator, as a result of their interactions with other hadrons. Then  $V_\Lambda(\rho)$  will decrease and the absorbed hyperon may become free, or, after subsequent interactions, it may be captured again in another part of the spectator. To account for these processes we check the absorption criterion regularly during the whole reaction time. In the following we show DCM results obtained with this more sophisticated criterion unless otherwise stated.

### B. The ultrarelativistic quantum molecular dynamics approach

For the investigation of spectator hypermatter formation in high-energy heavy-ion collisions we have also employed the ultrarelativistic quantum molecular dynamics (UrQMD v2.3) model [23,24]. This nonequilibrium transport approach constitutes an effective solution of the relativistic Boltzmann equation. The underlying degrees of freedom are hadrons and strings that are excited in energetic binary collisions. Mean fields can in principle be taken into account in this framework too, but for the present investigation we have run the model in the so-called cascade mode without external potentials.

The nucleon's coordinates are initialized according to a Woods-Saxon profile in coordinate space and their momenta are assigned randomly according to the Fermi distribution in the rest frame of the corresponding nucleus. The hadrons are propagated on straight lines until the collision criterion is fulfilled: The collision takes place if the covariant relative distance  $d_{\text{trans}}$  between two particles gets smaller than the distance  $d_0$  corresponding to the total cross section  $\sigma_{\text{tot}}$  (which depends on the energy  $s$  and the kind of hadrons  $h$ ),

$$d_{\text{trans}} \leq d_0 = \sqrt{\frac{\sigma_{\text{tot}}(\sqrt{s}, h)}{\pi}}. \quad (2)$$

The reference frame that is used for the time ordering of the collisions is the c.m. system of the nucleus-nucleus collision. However, each individual collision process is calculated in the rest frame of the binary collision.

In the UrQMD model 55 baryon and 32 meson species, ground state particles, and all resonances with masses up to 2.25 GeV are implemented with their specific properties and interaction cross sections. In addition, full particle-antiparticle symmetry is applied. Until now, isospin symmetry is assumed and only flavor-SU(3) states are taken into account. The elementary cross sections are calculated by detailed balance, or the additive quark model, or are fitted and parametrized according to the available experimental data. For resonance excitations and decays the Breit-Wigner formalism is employed assuming vacuum parameters.

Toward higher energies, the treatment of subhadronic degrees of freedom is of major importance. In the present model, these degrees of freedom enter via the introduction of a formation time for hadrons produced in the fragmentation of strings [25–27]. String formation and fragmentation is treated according to the Lund model. For hard collisions with large momentum transfer ( $>1.5$  GeV/ $c$ ) the Pythia model is used for the simulation of final states. The UrQMD transport model is successful in describing the yields and the  $p_T$  spectra of various secondary particles in  $pp$  and  $pA$  collisions [28]. A compilation of results of the recent version UrQMD-2.3 compared to experimental data can be found in [29].

For our present purposes we have applied the following procedure: During the UrQMD calculational run, we check for all  $\Lambda$ s and  $\Xi$ s in time steps of  $\Delta t = 0.5$  fm/ $c$ , first selecting only particles which are in the rapidity interval  $y_b \pm \Delta y$ , where  $y_b$  is the beam rapidity and  $\Delta y = 0.267$ , the rapidity spread associated with the Fermi momenta of nucleons in the initial nucleus. For such particles the net baryon density in the local rest frame of the strange particle is then estimated. This is done by calculating the zero component of the net baryon current of all baryons within the same rapidity interval. For this purpose we assume that each baryon is represented by a Gaussian wave package of width  $\sigma = 1$  fm, which is Lorentz contracted in the direction of motion.

For the hyperon absorption by spectators the same potential criterion as in the DCM was adopted: We calculate the kinetic energy of the strange particles, both  $\Lambda$  and  $\Xi$  hyperons, with respect to the rest frame of the projectile nucleus. Whenever this energy is less than the corresponding potential, calculated with Eq. (1), the particle is absorbed and removed from the cascade simulation. This corresponds exactly to the first-capture criterion used in the DCM. We have checked that the results do only weakly depend on the depth of the hyperon potential, when it is varied in a reasonable interval of  $\pm 20\%$ .

### III. FORMATION OF HYPERON-RICH SPECTATOR MATTER

A detailed picture of peripheral relativistic heavy-ion collisions has been established in many experimental and theoretical studies. Nucleons from the overlapping parts of the projectile and target (participant zone) interact

intensively between themselves and with other hadrons produced in primary and secondary collisions. Nucleons from the nonoverlapping parts do not interact strongly, and they form the residual nuclear systems, which we call spectators. In all transport models the production of hyperons is associated with nucleon-nucleon collisions, e.g.,  $p + n \rightarrow n + \Lambda + K^+$ , or collisions of secondary mesons with nucleons, e.g.,  $\pi^+ + n \rightarrow \Lambda + K^+$ . Strange particles are mainly produced in the participant zone; however, they can rescatter and populate the whole momentum space around the colliding nuclei.

In the present paper we are interested mostly in hyperons which propagate with velocities close to the initial velocities of the nuclei, i.e., in the vicinity of nuclear spectators. We assume that such hyperons can be absorbed by the spectators if their kinetic energy (in the rest frame of the spectator) is lower than the potential generated by neighboring spectator nucleons. Due to the secondary interactions these spectators are excited and have a dilute (subnuclear) density. As was previously investigated in detail [12,30,31], the further evolution of the spectators depends on their excitation energy. At lower excitation energies (below 2–3 MeV per nucleon) they deexcite via evaporation and/or fission-like processes, while at higher excitation energies they undergo multifragmentation [13]. We expect the same behavior for the residual nuclei containing hyperons.

As was also found in previous studies of spectator fragmentation [30], there exists a strong correlation between the excitation energy transferred to the spectator matter and the number of high-energy particles produced in the participant zone. On the other hand, the multiplicity of strange particles is also proportional to the number of participants. Therefore, we expect that the hyperspectators will be very excited, and finally they will break up into many fragments, with some of them containing the captured hyperons. This process was already analyzed in Ref. [14].

The DCM gives a very good description of yields and spectra of free hyperons and kaons produced in relativistic heavy-ion collisions. For example, in Fig. 1 the calculated inclusive rapidity distributions of  $\Lambda$  and  $\Sigma^0$  hyperons are compared with experimental data of the E896 experiment at the Alternating Gradient Synchrotron (AGS) [32]. As one can clearly see, the primary nucleon-nucleon interactions, which are noted as “without rescattering,” are responsible only for a small fraction of the strange particles. The secondary interactions of baryons and mesons provide the dominant contribution to the production of hyperons at all rapidities, and they are crucially important for description of the data. Similarly, the UrQMD model proved to be very successful in describing the strangeness production in high-energy nuclear collisions (see, e.g., [29]).

Since both the DCM and UrQMD model can follow the evolution of all interacting particles in space and time, we can predict where and when the produced  $\Lambda$  is absorbed in a projectile or target spectator. As a first example, in Fig. 2 we demonstrate the spatial distribution of  $\Lambda$  hyperons, which satisfy the capture criterion described above. This figure shows the coordinates of absorption points projected into the transverse plane ( $X$  and  $Y$  coordinates), perpendicular to the beam axis ( $Z$  coordinate). The spectator residues are formed mostly

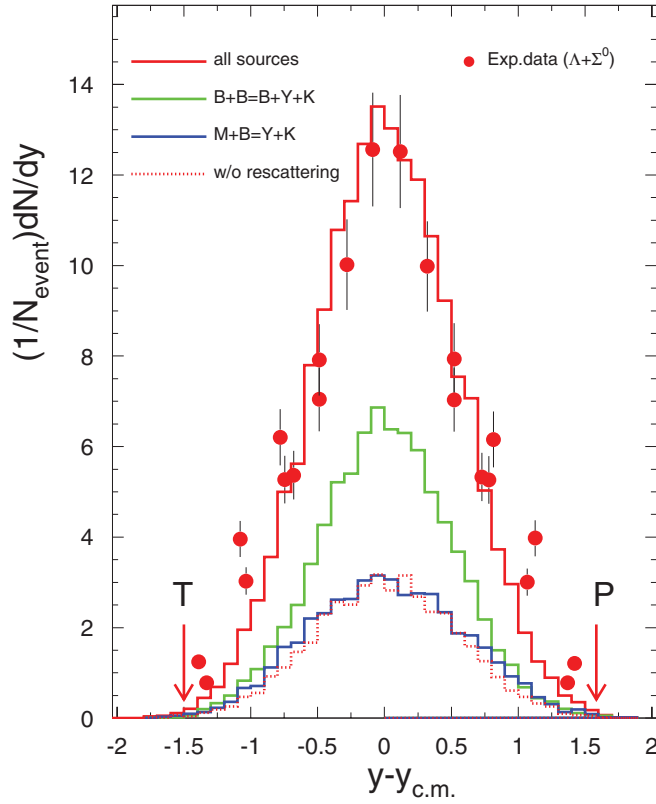


FIG. 1. (Color online) Rapidity distribution of  $\Lambda$  and  $\Sigma^0$  hyperons produced in collisions of Au projectiles with momentum 11 GeV/c per nucleon on a Au target. Red circles are experimental data [32]. Histograms show contributions of different processes as predicted by the DCM. Rapidities of the projectile (P) and target (T) are shown by arrows.

from the nonoverlapping parts of target and projectile. The results are presented for Au + Au collisions at a beam energy of 20 GeV per nucleon and an impact parameter of 8.5 fm. Such collisions will be studied in future experiments at the Facility for Antiproton and Ion Research (FAIR) (Darmstadt, Germany) and at the Nuclotron-based Ion Collider Facility (NICA) (JINR, Dubna, Russia). The figure accumulates the simulation results after  $2 \times 10^5$  collision events and integrated over the reaction time. For a better resolution of the scatter plot we show only about 10% of the absorption points taken randomly. Only DCM results are presented, since the UrQMD gives the same qualitative predictions.

In this plot and the following figures we show the elementary hadron interactions which are responsible for the production of the absorbed  $\Lambda$  hyperons. This information is important to understand the physics of the absorption process. One can see that the contribution of direct nucleon-nucleon and pion-nucleon interactions is rather small, since the kinetic energy of produced hyperons is quite high in the rest frame of the nuclear residues. Only few hyperons emerging from these interactions may be captured by spectators. The largest number of absorbed hyperons is produced in the secondary interactions of strange particles with nucleons, e.g., in strangeness exchange reactions induced by antikaons and in the rescattering of fast hyperons on slow spectator nucleons.

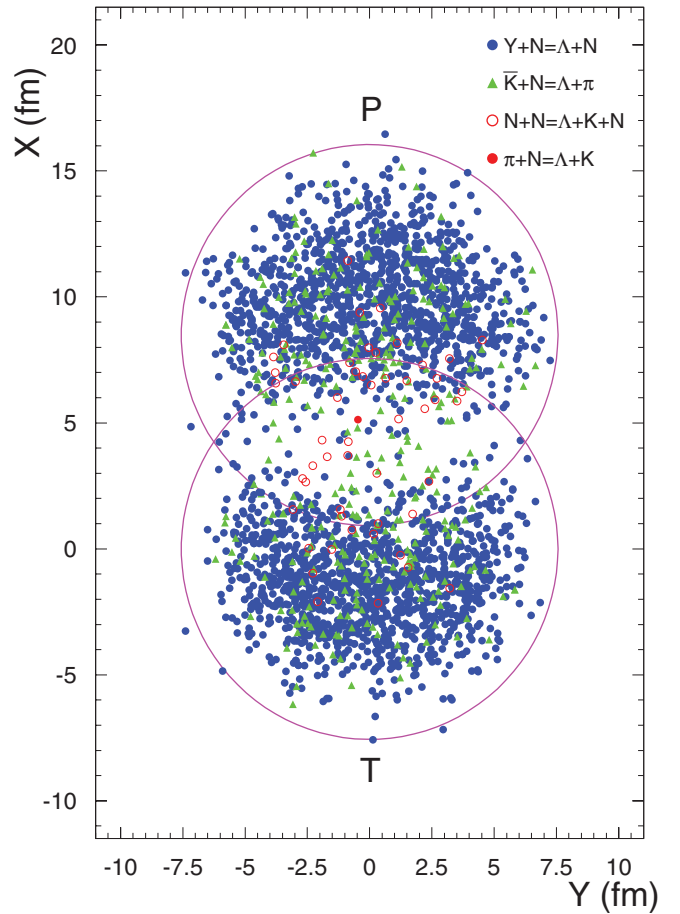


FIG. 2. (Color online) Coordinates of  $\Lambda$  absorption points projected into the transverse plane perpendicular to the beam axis. Results of DCM calculations are shown for Au (20 A GeV) + Au collisions with an impact parameter of 8.5 fm. The circles are outer contours of projectile (P) and target (T) nuclei at this impact parameter. Processes creating the absorbed hyperon, such as interactions of secondary hyperons, antikaons, nucleons, and pions with nucleons, are indicated by different symbols.

It is instructive to investigate the absorption rate as a function of time to extract the coordinates of absorption points along the beam axis. The correlation between the time and X-Z coordinates of the absorption points is illustrated in Fig. 3. One can see that some absorption events happen at a very early time ( $\lesssim 10$  fm/c), when hyperons produced in the hot participant are captured by a few nucleons located in this zone. However, the contribution of this mechanism to the total capture yield is rather small. Moreover, due to intensive interactions taking place in the participant zone, many of the surrounding nucleons will be kicked out at later stages and the hyperons can become free again. This is why in the DCM calculations we have introduced a multiple check of the absorption criterion at later time steps.

After the first 10 fm/c, when the projectile and target residues are completely separated in coordinate space, the capture process is entirely associated with the interactions of strange particles in the nonoverlapping zones of colliding nuclei. Reactions induced by antikaons and hyperons with



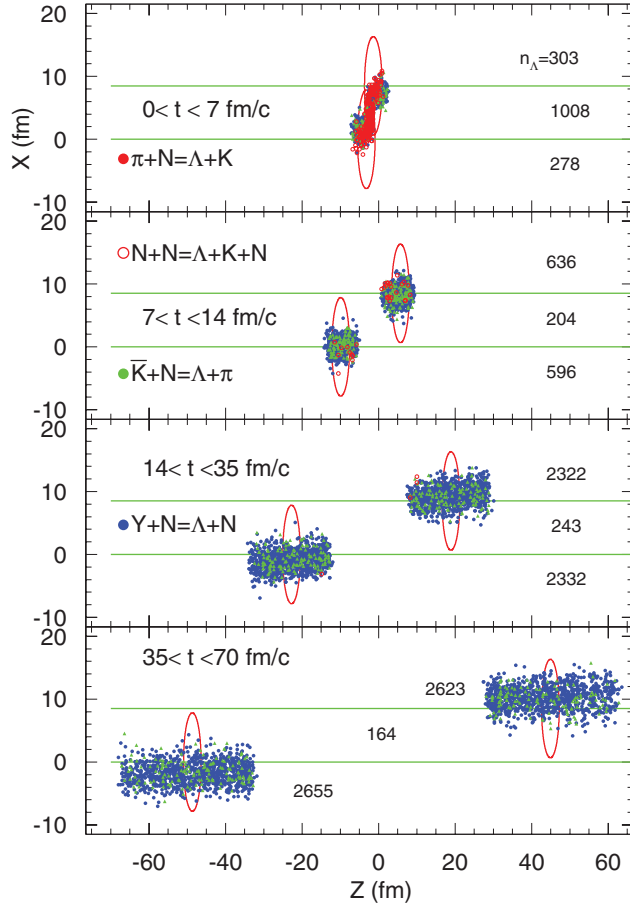


FIG. 3. (Color online) The same DCM calculations with symbol notation as in Fig. 2, but in the  $X$ - $Z$  plane. The symbols show the coordinates of  $\Lambda$  absorption points in projectile and target spectators. Ellipses show the average positions of projectile and target nuclei during time intervals indicated in the figure (from top to bottom) in the system of equal velocities. The numbers of hyperons,  $n_\Lambda$  (per  $2 \times 10^5$  events), captured in the participant and spectator zones during these intervals are noted on the right side.

velocities close to projectile and target velocities provide the dominant contribution. One can see from Fig. 3 that this process continues for tens of fm/c, and it is quite possible to produce a hypernuclear residue even at the time of around 50 fm/c. By looking at the  $X$  coordinates of the absorption points one can clearly see that the region of the  $\Lambda$  absorption moves from the overlapping zone at early time to the nonoverlapping spectator parts for later times, as a result of many secondary interactions.

The space-integrated time evolution of the  $\Lambda$  absorption is presented in Fig. 4 for the same Au + Au collisions at 20A GeV projectile energy. For convenience, the capture rate is normalized to the total number of collision events. This figure allows for a quantitative estimate of the absorption rates associated with different interaction channels. From the DCM results one can see that about 90% of  $\Lambda$ s absorbed in spectators come from secondary reactions when previously produced hyperons rescatter on spectator nucleons. After the first 50–70 fm/c the absorption rate becomes low and it drops

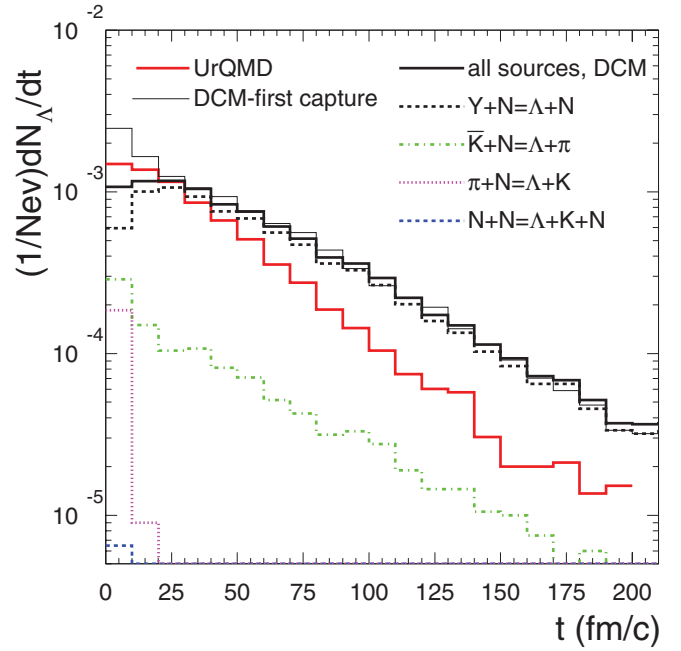


FIG. 4. (Color online) The rate of the  $\Lambda$  absorption by spectators (vs time) as calculated in the DCM and the UrQMD model for the Au (20 A GeV) + Au reaction. Thick black, green, blue, and violet lines give the total rate and contributions of specific reactions to production of these  $\Lambda$  hyperons as predicted by the DCM. The red line is the total rate predicted by UrQMD. The thin black line is the DCM calculation using the first-capture criterion as in UrQMD.

exponentially, signalling that the dynamical process is coming to an end, and the equilibrium stage of the spectator's evolution should be considered.

We show two types of DCM calculations: The thin black line corresponds to the “first-capture” absorption criterion (as discussed in Sec. II A), i.e., the  $\Lambda$  hyperons are considered as finally absorbed after the capture criterion is first satisfied. No further propagation of these hyperons is simulated. The same criterion is used in all UrQMD calculations. The thick black line represents the second DCM criterion which takes into account a time evolution of the nucleon density in the vicinity of the captured hyperon. As was mentioned, some nucleons can leave this neighborhood at a later time, leading to a decrease of the local density so that the hyperon may escape. As one can see in Fig. 2 this happens mostly at the early stage ( $\lesssim 10$ – $20$  fm/c) with hyperons captured in the overlap zone. On the other hand, at later times some of these hyperons may be recaptured again in the spectator zone. Therefore, the time scale for the absorption process will be longer in this case. By comparing these two calculations one can see that the total number of captured hyperons decreases only slightly in this case (not more than by 30%).

In Fig. 4 one can see a noticeable difference between the two considered models: the absorption rate decreases faster in the UrQMD. As our analysis shows, one of the reasons is that the two models use somewhat different parametrizations for the  $\Lambda N$  elastic cross section  $\sigma_{\Lambda N}$ . In the UrQMD this cross section is taken to be constant,  $\sigma_{\Lambda N} \approx 30$  mb, at all relative momenta. While in the DCM a behavior consistent with a

$\sim 1/v$  law is assumed at relative momenta below 1 GeV/ $c$ . For this reason, according to UrQMD, hyperons undergo fewer interactions with spectator nucleons at late times (when relative  $\Lambda N$  velocities are small), can leave the spectators without rescattering, and, therefore, have lower absorption probability. However, because of lack of direct experimental information on the  $\Lambda$ -nucleon scattering cross section we do not give a preference to a particular parametrization and consider both results as meaningful. Another reason is that the number of spectator nucleons participating in absorption of  $\Lambda$  is larger for the DCM case (see the next section). It is especially important at later times, since both the production of hyperons in secondary reactions and their rescattering depend on the size of spectators. We emphasize that in the end these two dynamical models give rather similar integrated yields of produced hyperspectators.

#### IV. CHARACTERISTICS OF SPECTATOR HYPERMATTER

As one can see in Fig. 4, 80–90% of the  $\Lambda$ s are absorbed in spectators by the time of 80 fm/ $c$ . At this stage all secondary interactions are practically over and we can determine the parameters of the spectator residues. Figure 5 shows the time evolution of the average mass number of spectators as predicted by DCM and UrQMD calculations. The decrease of masses with time is obviously caused by secondary interactions of fast particles with spectator nucleons. In the UrQMD model a nucleon is defined as spectator if it has not participated in elementary interactions. On the other hand, the DCM definition of spectators includes nucleons which underwent interactions and have been recaptured by the

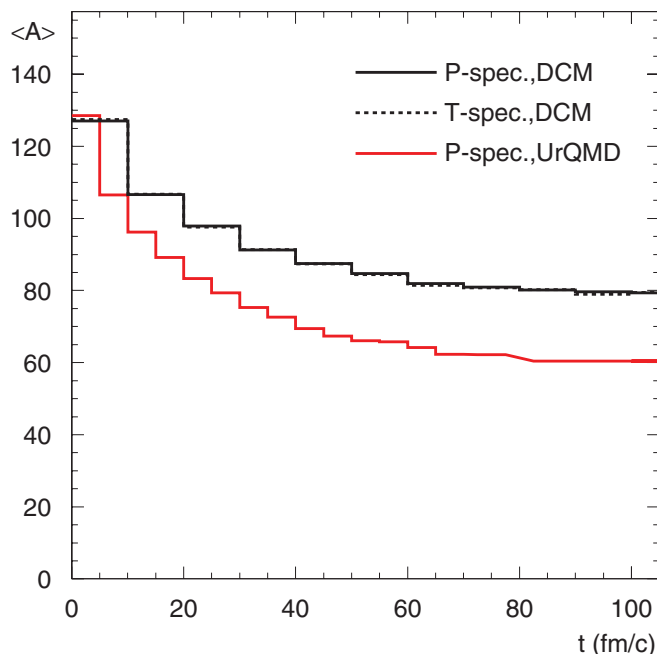


FIG. 5. (Color online) Time evolution of the mean number of nucleons in the spectator (P: projectile, T: target) residual nuclei after  $\Lambda$  absorption as predicted by the DCM and UrQMD calculations. The reaction is as in Fig. 2.

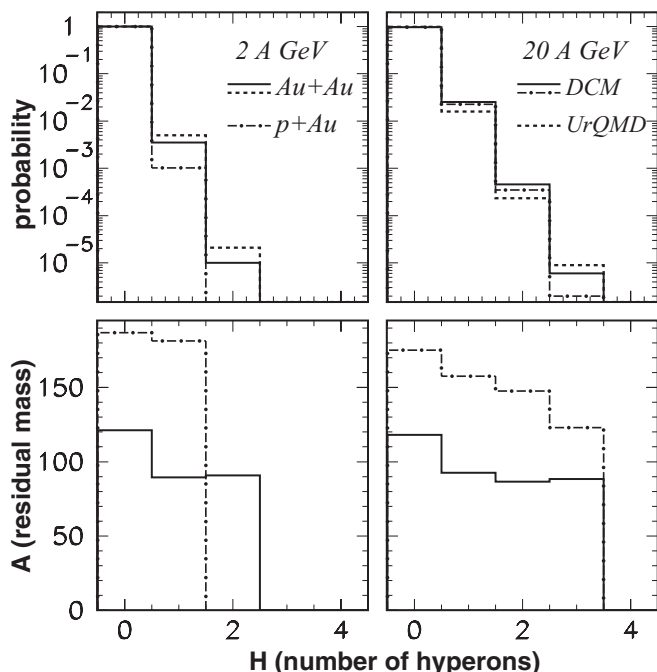


FIG. 6. Probability for formation of conventional and strange spectator residuals (top panels) and their mean mass numbers (bottom panels) vs the number of captured  $\Lambda$  hyperons ( $H$ ), calculated with the DCM and the UrQMD model for  $p + Au$  and  $Au + Au$  collisions with energy of 2 GeV per nucleon (left panels) and 20 GeV per nucleon (right panels). The reactions and energies are noted in the figure by different histograms.

nuclear potential. For this reason the DCM spectator residues are bigger. It is seen that the total number of nucleons in the spectators is nearly saturated after 70 fm/ $c$ . During the dynamical stage the high-energy neutrons and protons interact similarly; therefore, the average neutron-to-proton ratio in the spectators is almost the same as in colliding nuclei.

These spectator residues acquire some excitation energy, which can be calculated within both models too. However, it is well known from experiments with conventional nuclear matter that in relativistic collisions the average temperature of such residues is saturated at the level of 5–6 MeV, despite large event-by-event fluctuations of the excitation energy [13,33]. We expect similar average temperatures in hyperresidues too, especially when the number of captured hyperons is much smaller than the number of nucleons.

One important practical result of this study is the estimated probability for producing spectator residues with different numbers of captured  $\Lambda$ s. These probabilities are shown in Fig. 6 for residues containing up to three  $\Lambda$  hyperons, together with their mean masses, for  $Au + Au$  and  $p + Au$  collisions with energies of 2 and 20 A GeV. The chosen beam energies correspond to the present GSI experiments and the future FAIR experiments. The calculations were performed with minimum bias, i.e., integrated over all impact parameters. For each reaction  $10^6$  Monte Carlo events were generated. According to the DCM calculation, at the highest energy the probability for capturing one hyperon in the spectator amounts to a few percent, and two hyperons may be captured with a

significant probability of about  $\sim 5 \times 10^{-4}$ . Furthermore, even three  $\Lambda$  hyperons may be captured by spectators with a low but measurable probability of  $\sim 10^{-5}$ . This estimate is based only on few events, therefore, its statistical significance is limited. However, we believe, it is sufficient for a qualitative conclusion that such events can be obtained experimentally. The absorption of a higher number of hyperons is also feasible. This new mechanism opens a unique opportunity to produce and study multistrange systems, which are not conceivable in other nuclear reactions. Since the predicted masses of the hyperspectators are quite large, one can even speak about the formation of excited hypermatter. We have found that variation of the hyperon potential within a reasonable range does only weakly influence the capture rates at the highest energy, because these hyperons undergo multiple rescattering and decelerate considerably in the spectator matter.

Compared to the DCM results, the probabilities for producing hyperspectators predicted by the UrQMD model can differ by a factor of 2. This difference shows the uncertainty in the calculations based on the best theoretical models available at present. Still the qualitative agreement between the two transport models gives us confidence that this new method can be successfully used for producing new hypernuclei.

From calculations with the UrQMD model we have also estimated the probability of a  $\Xi$  being absorbed in the spectator fragment at  $E_{\text{lab}} = 20 A \text{ GeV}$  to be on the order of  $10^{-5}$  per event. Such an absorption would eventually lead to the formation of  $\Xi$  hypernuclei in future experiments. This opens the interesting possibility to study a new multistrange system which is even less understood than the conventional  $\Lambda$  hypernuclei.

For completeness, in Fig. 7 we show the DCM predictions for the mass distributions of residues (representing both

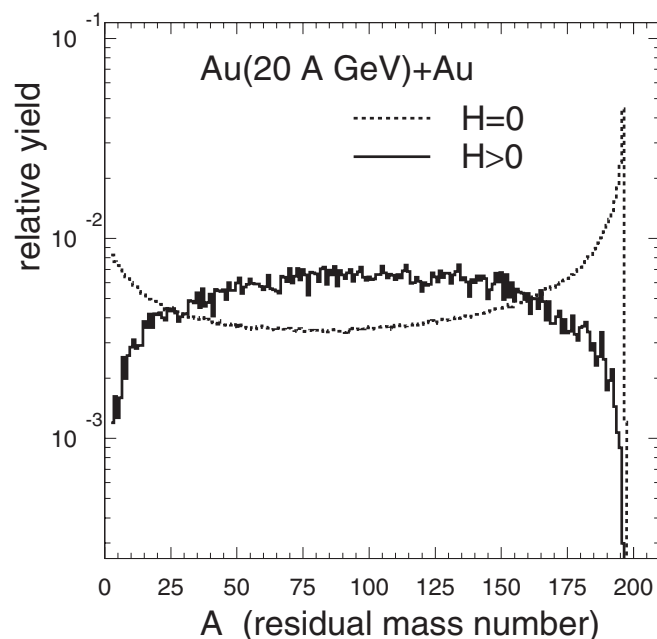


FIG. 7. Relative mass distribution of conventional ( $H = 0$ ) and hypernuclear ( $H > 0$ ) spectators after DCM calculations for Au + Au collisions with projectile energy of 20 GeV per nucleon.

conventional matter and hypermatter), integrated over all impact parameters. One can notice a significant difference in the shape: for conventional residues the distribution is very broad, and it has a sharp peak around the projectile mass, corresponding to very peripheral collisions. This result is consistent with information on the spectator mass distribution extracted from analyses of experimental data [12,13,30,31]. The heavy residuals are weakly excited since only a few nucleons participated in the reaction. When hyperons are produced in peripheral heavy-ion collisions, the spectator mass distribution is different: In this case the most probable way to produce a captured  $\Lambda$  is to consider events when many nucleons of the projectile and target are involved in the interaction and, therefore, can leave the nuclei. As a result, the spectator residues are smaller and the shape of the mass distribution is changed. For this reason, it is unlikely to produce hyperresidues with masses around the initial mass in heavy-ion collisions. However, the distribution remains very broad and practically all intermediate-mass residues can be produced with considerable probability.

We should note that the formation of hyperresidues was earlier discussed in Ref. [34] for the reactions initiated by protons. As one can see from Fig. 6 the probabilities for production of multistrange hyperresidues in such reactions are lower than in nucleus-nucleus collisions. This is easy to understand, since more hyperons in the spectator region can be produced when more particles participate in the interactions. However, the channels with absorption of one  $\Lambda$ , and even with two  $\Lambda$ s at the highest energy, can be still very probable in  $p + \text{Au}$  collisions. In addition, the masses of hyperresidues will be larger, since in this case fewer target nucleons are involved in the reaction and the emission of nucleons will be lower. The main problem expected for proton beams is that experimentally it is very difficult to identify decay products of a slow hypernucleus in a background of free  $\Lambda$  decays. Relativistic heavy-ion collisions have essential advantages: Because of the Lorentz factor their lifetime becomes longer and projectile hyperfragments can travel a longer distance. This makes it possible to use sophisticated vertex detectors and fragment separation techniques for their identification.

## V. RAPIDITY DISTRIBUTIONS OF HYPERRESIDUES AND FREE $\Lambda$ HYPERONS

For the experimental identification of hypernuclei it is of crucial importance to know the background associated with free hyperons produced in a reaction. For example, let us consider the reaction  ${}^6\text{Li} + {}^{12}\text{C}$  at a beam energy of 1.9 GeV per nucleon, which was employed on the first stage of the HypHi experiment at GSI [35]. Figure 8 shows the rapidity distributions of the free and captured  $\Lambda$ s produced in this reaction. One can see that for these light nuclei the yield of free  $\Lambda$ s in the spectator region is 1–2 orders of magnitude larger than the yield of hypernuclei originating from the projectile spectator. For this reason, the fraction of pions coming from the decay of hypernuclei is very small and careful correlation measurements are required to identify these hypernuclei. It is interesting that this fraction is higher for the larger target

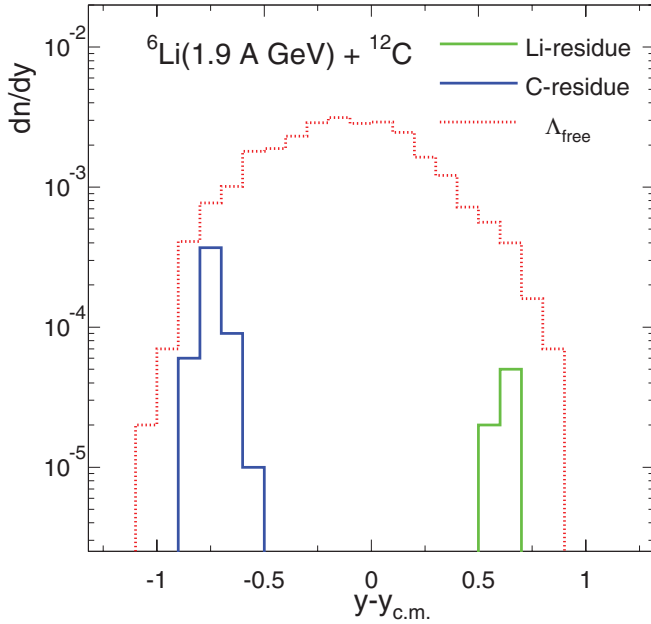


FIG. 8. (Color online) DCM calculations of rapidity distribution for free  $\Lambda$  and for hyperresidues with one  $\Lambda$  coming from target (C) and projectile (Li) spectators in the reaction  ${}^6\text{Li}$  (1.9 A GeV) +  ${}^{12}\text{C}$ .

spectator; however, in the experiment only projectile fragments can be detected.

A higher ratio of spectator hypernuclei to free  $\Lambda$ s is obtained in reactions involving heavy nuclei. Figure 9 presents the result of the DCM calculation for Au + Au collisions at a beam energy of 20 A GeV, as planned in future FAIR experiments. One can see clearly that the excess of free  $\Lambda$ s over hypernuclei in spectator kinematic zones is minimal in this case, i.e., less

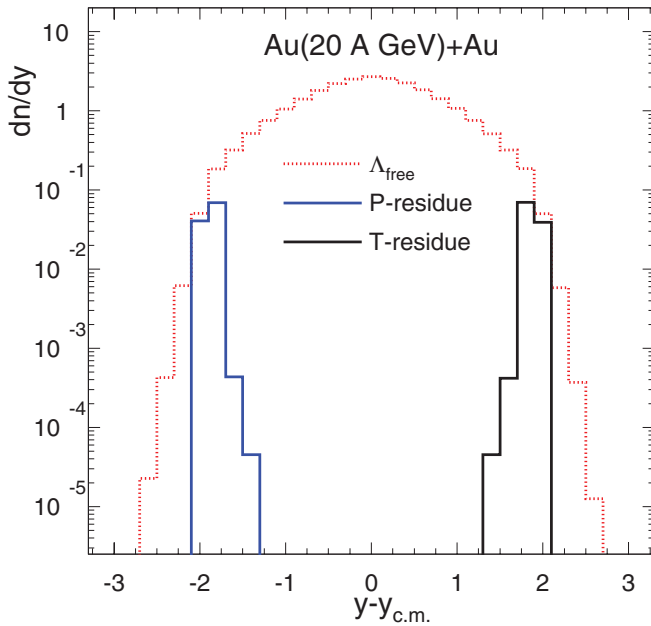


FIG. 9. (Color online) DCM calculations of rapidity distribution for free  $\Lambda$  and for hyperresidues coming from target (T) and projectile (P) spectators in the reaction  ${}^{197}\text{Au}$  (20 A GeV) +  ${}^{197}\text{Au}$ .

than a factor of 2–3. This significant improvement has a simple explanation: More secondary hyperons interact with spectator nucleons at a later time of the reaction ( $t > 20$  fm/c). These hyperons have velocities close to the velocities of spectators; therefore, it is more probable for them to have other spectator nucleons nearby.

## VI. MOMENTUM DISTRIBUTION OF ABSORBED HYPERONS

As follows from the capture criterion (see Sec. II) the absorbed  $\Lambda$  hyperons have small momenta in the rest frame of the spectators. Direct DCM and UrQMD calculations of the invariant momentum distribution of the absorbed  $\Lambda$ s for Au (20 A GeV) + Au collisions are shown in Fig. 10. One can see that the distribution is rather narrow with a smooth cutoff for hyperon momenta greater than 250–300 MeV/c. However, this distribution is not steplike, and we can even approximate the slope of the spectrum at  $P > 100$  MeV/c by a Boltzmann distribution with apparent temperature  $T \approx 10$  MeV. This low temperature indicates that these hyperons are rather slow in the considered reference frame, and this is a typical temperature for particles coming from the spectator matter as was demonstrated long ago by the analysis of ALADIN data [30].

Sometimes the coalescence criterion for cluster formation is used, when a hyperon is considered to be absorbed by a nuclear cluster if they are close in momentum and coordinate space. Our results shown in Fig. 10 justify using the coalescence criterion in a simple consideration, when secondary interactions in spectators are neglected. Moreover, our potential criterion works similar to the coalescence: if a particle propagates

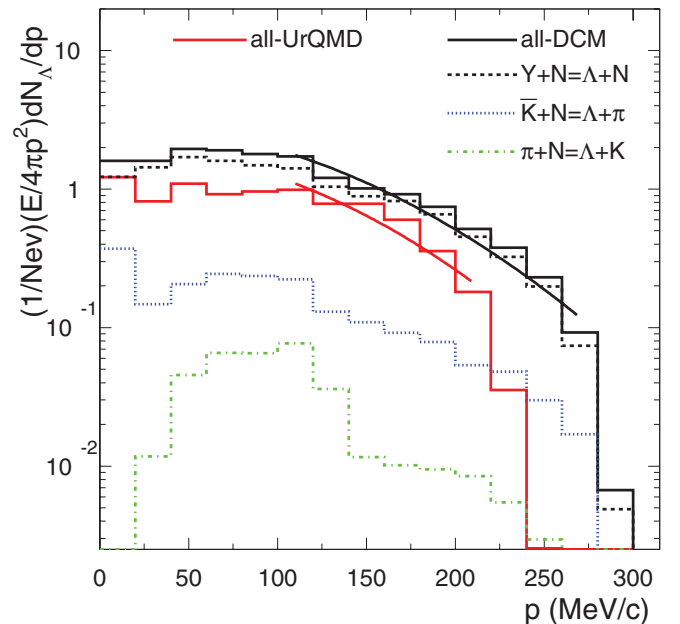


FIG. 10. (Color online) Momentum distribution of  $\Lambda$  hyperons captured in the spectators. The notation is the same as in Fig. 4. Solid curves show the Boltzmann fits of the distributions with temperature  $\sim 10$  MeV.



at late times only those nucleons will coalesce which have small relative momenta and close coordinates. Recently the Giessen Boltzmann-Uehling-Uhlenbeck (GiBUU) model [9] was generalized to describe hypernuclei production in similar relativistic nucleus-nucleus collisions. The authors have determined excited residues after the dynamical stage. Then the statistical multifragmentation model (SMM) [13] was applied to simulate the deexcitation and multifragmentation of these residues into conventional fragments. Afterward, a coalescence prescription was used to form hyperfragments from these conventional fragments and free  $\Lambda$ s. However, an important difference of our approach is the assumption that the absorption is taking place rather fast (during first few tens of fm/c), i.e., before the breakup of excited residues. Namely, their secondary decay is responsible for production of final cold hypernuclei. The deexcitation itself takes a considerable time, about  $10^2$ – $10^4$  fm/c depending on excitation energy. Therefore, the final yields of specific hypernuclei in our approach will be different from those obtained from the approach of Ref. [9].

On the other hand, the standard coalescence model can be successfully used to take into account the final state interaction between free nucleons leading to the formation of the lightest clusters, as was shown already in Ref. [17]. We expect that this mechanism may work in the case of light nuclei and free hyperons too. Similar results concerning the production of light hyperfragments can be obtained within the thermal models [36–38]. A connection between parameters of coalescence and thermal models describing production of the lightest clusters at high excitation energies has been under discussion for a long time (see, e.g., Refs. [39,40]). In principle, the coalescence prescription can be applied for free nucleons and hyperons to investigate the formation of both conventional light nuclei (such as  $d$ ,  $t$ , and He) and light hypernuclear fragments produced in the spectator region. However, our preliminary calculations for Au + Au collisions have shown that the probability for the production of hyperclusters by this mechanism is by several orders of magnitude smaller than that for the absorption of hyperons in the spectators, as considered in this paper. On the other hand, we have found that the coalescence mechanism is more efficient for the formation of hyperclusters in the midrapidity zone [41].

## VII. DISINTEGRATION OF HOT SPECTATOR MATTER INTO HYPERFRAGMENTS

In the final stage of the reaction the excited hyperresidues undergo deexcitation. At low excitation energy, this should be an evaporation (and may be a fission) process, similar to what is well known for the conventional nuclei. However, as we see from our calculations and also from analyses of experimental data [13,30,31], the thermalized residual nuclei have rather high excitation energies in the nuclear scale, and, consequently, they should undergo multifragmentation with a characteristic time of about 100 fm/c. The generalization of the SMM [13] into the strangeness sector by including  $\Lambda$  hyperons has been done in Ref. [14]. It was demonstrated

that the fragment mass distributions are quite different for fragments with different strangeness content. This means that the multifragmentation of excited hypernuclear systems proceeds in a different way compared with conventional nuclei. The reason is the additional binding energy of hyperons in nuclear matter. It was also shown that the yields of fragments with two  $\Lambda$ s depend essentially on the mass formulas (i.e., on details of  $\Lambda N$  and  $\Lambda\Lambda$  interactions) used for the calculations [14,42]. Therefore, an analysis of double hypernuclei can help to improve these mass formulas and reveal information about the hyperon-hyperon interaction. In Ref. [43] the decay of light, excited hypersystems was considered within the framework of the Fermi breakup model. It was also concluded that the production rate of single and double hypernuclei is directly related to their binding energy.

We would like to mention another canonical statistical model [44], which was developed in line with [14]. It gives similar results concerning hypernuclei produced after the decay of excited hypersystems. A detailed analysis of the deexcitation process taking into account the primary dynamical stage of the reaction will be presented in a subsequent publication.

## VIII. CONCLUSION

Within the DCM and UrQMD model we have investigated the production of hyperons in peripheral relativistic heavy-ion collisions and their capture by the attractive potential of nuclear spectator residues. In contrast to the coalescence mechanism, which may be responsible for the formation of light hyperfragments, this capture process can also lead to the production of medium and heavy hypernuclei. We predict significant yields of such hypernuclear spectators with a very broad mass distribution. An important advantage of this method over the traditional reactions induced by kaons is that it provides a natural way to produce multihyperon systems. As our calculations show, the relative probabilities to produce residual nuclei with two and three  $\Lambda$ s in Au + Au reactions at 20 A GeV are about  $5 \times 10^{-4}$  and  $10^{-5}$ , respectively. They should be sufficient for systematic studies of such systems in the future FAIR and NICA experiments.

Conventional spectator matter in relativistic nuclear collisions was intensively investigated during the last 20 years [13,33]. Similarly, we expect the production of excited hypermatter at a baryon density  $\rho \sim 0.1$ – $0.8\rho_0$  and temperatures around  $T \sim 3$ – $7$  MeV, which are typical for the coexistence region of the nuclear liquid-gas phase transition. This opens the possibility to study this phase transition in nuclear matter with a strangeness admixture, which is important for the physics of neutron stars. An analysis of the disintegration of this matter into cold nonstrange fragments and hyperfragments will reveal information about the properties of hypernuclei, their binding energies, and, finally,  $YN$  and  $YY$  interactions. Hypernuclei of all sorts are expected to be produced in this way, including nuclei around drip lines, which is not possible with traditional methods.

It would be very helpful to perform further detailed theoretical calculations, which include all stages of the reaction: (1) production of hyperon-rich spectator nuclear systems,

(2) formation of hyperfragments via decay of these systems, and (3) decay of the hypernuclei via weak interaction with production of nonstrange particles in the final state. The DCM and UrQMD approaches can be used for the description of the dynamical stage and they can be followed by statistical models, e.g., the SMM. In this case all products of the reaction can be determined on an event-by-event basis. Advanced correlation measurements, which are necessary to identify hypernuclei, can be tested with these models too.

The detection of spectator hypernuclei may suffer from a background problem, which may require sophisticated experimental methods to solve. However, it has already been demonstrated experimentally that in peripheral collisions of light projectiles and targets the cross section for the capture of produced  $\Lambda$  hyperons by projectile residues may be of the order of a few microbarns [11], and this cross section allows for an experimental identification of the produced hypernuclei

[35]. As we have demonstrated, involving heavy projectiles and targets can indeed improve the signal-to-background ratio significantly. We believe that hypernuclear physics will benefit strongly from exploring new production mechanisms and the novel detection technique associated with spectator fragmentation reactions.

#### ACKNOWLEDGMENTS

This work was supported by BMBF, HGS-HIRE, and the Hessian LOEWE initiative through the Helmholtz International Center for FAIR (HIC for FAIR). We also acknowledge partial support from DFG Grant No. 436 RUS 113/957/0-1, as well as Grants No. RFBR 09-02-91331 and No. NSH-7235.2010.2 (Russia). K.K.G. thanks FIAS for its hospitality.

- 
- [1] O. Hashimoto and H. Tamura, *Prog. Part. Nucl. Phys.* **57**, 564 (2006).
- [2] J. Schaffner, C. B. Dover, A. Gal, C. Greiner, and H. Stoecker, *Phys. Rev. Lett.* **71**, 1328 (1993).
- [3] W. Greiner, *J. Mod. Phys. E* **5**, 1 (1995).
- [4] P. Papazoglou, S. Schramm, J. Schaffner-Bielich, H. Stoecker, and W. Greiner, *Phys. Rev. C* **57**, 2576 (1998).
- [5] J. Schaffner and I. N. Mishustin, *Phys. Rev. C* **53**, 1416 (1996).
- [6] X. Lopez *et al.*, *Prog. Part. Nucl. Phys.* **53**, 149 (2004).
- [7] M. Wakai, H. Bando, and M. Sano, *Phys. Rev. C* **38**, 748 (1988).
- [8] J. Aichelin and K. Werner, *Phys. Lett. B* **274**, 260 (1992).
- [9] Th. Gaitanos, H. Lenske, and U. Mosel, *Phys. Lett. B* **663**, 197 (2008); **675**, 297 (2009).
- [10] K. J. Nield *et al.*, *Phys. Rev. C* **13**, 1263 (1976).
- [11] S. Avramenko *et al.*, *JETP Lett.* **48**, 516 (1988); *Nucl. Phys. A* **547**, 95c (1992).
- [12] A. S. Botvina *et al.*, *Nucl. Phys. A* **584**, 737 (1995).
- [13] J. P. Bondorf, A. S. Botvina, A. S. Iljinov, I. N. Mishustin, and K. Snepken, *Phys. Rep.* **257**, 133 (1995).
- [14] A. S. Botvina and J. Pochodzalla, *Phys. Rev. C* **76**, 024909 (2007).
- [15] The STAR Collaboration, *Science* **328**, 58 (2010).
- [16] L. M. Barbier *et al.*, *Phys. Rev. Lett.* **60**, 405 (1988).
- [17] V. D. Toneev and K. K. Gudima, *Nucl. Phys. A* **400**, 173c (1983).
- [18] V. D. Toneev, N. S. Amelin, K. K. Gudima, and S. Yu. Sivoklov, *Nucl. Phys. A* **519**, 463c (1990).
- [19] N. S. Amelin, E. F. Staubo, L. S. Csernai, V. D. Toneev, and K. K. Gudima, *Phys. Rev. C* **44**, 1541 (1991).
- [20] R. D. Field and R. P. Feynman, *Nucl. Phys. B* **136**, 1 (1978).
- [21] N. S. Amelin, K. K. Gudima, S. Yu. Sivoklov, and V. D. Toneev, *Sov. J. Nucl. Phys.* **52**, 272 (1990).
- [22] I. Ahmad, M. Mian and M. Z. Rahman Khan, *Phys. Rev. C* **31**, 1590 (1985).
- [23] M. Bleicher *et al.*, *J. Phys. G* **25**, 1859 (1999).
- [24] S. A. Bass *et al.*, *Prog. Part. Nucl. Phys.* **41**, 255 (1998).
- [25] B. Andersson, G. Gustafson, and B. Nilsson-Almqvist, *Nucl. Phys. B* **281**, 289 (1987).
- [26] B. Nilsson-Almqvist and E. Stenlund, *Comput. Phys. Commun.* **43**, 387 (1987).
- [27] T. Sjostrand, *Comput. Phys. Commun.* **82**, 74 (1994).
- [28] E. L. Bratkovskaya *et al.*, *Phys. Rev. C* **69**, 054907 (2004).
- [29] H. Petersen, M. Bleicher, S. A. Bass, and H. Stoecker, [arXiv:0805.0567v1](https://arxiv.org/abs/0805.0567v1) [hep-ph].
- [30] H. Xi *et al.*, *Z. Phys. A* **359**, 397 (1997).
- [31] R. P. Scharenberg *et al.*, *Phys. Rev. C* **64**, 054602 (2001).
- [32] S. Albergo *et al.*, *Phys. Rev. Lett.* **88**, 062301 (2002).
- [33] J. Pochodzalla, *Prog. Part. Nucl. Phys.* **39**, 443 (1997).
- [34] Z. Rudy, W. Cassing *et al.*, *Z. Phys. A* **351**, 217 (1995).
- [35] The HypHI Collaboration, Letter of Intent (2005), [<http://www.gsi.de/documents/DOC-2005-Feb-432-2.pdf>]; Progress Report (2006), [<http://www.gsi.de/documents/DOC-2006-Jan-2-1.pdf>].
- [36] P. Braun-Munzinger, J. Stachel, *J. Phys. G* **21**, L17 (1995).
- [37] J. Steinheimer, M. Mitrovski, T. Schuster, H. Petersen, M. Bleicher, and H. Stoecker, *Phys. Lett. B* **676**, 126 (2009).
- [38] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stoecker, *Phys. Lett. B* **697**, 203 (2011).
- [39] A. Mekjian, *Phys. Rev. Lett.* **38**, 640 (1977).
- [40] W. Neubert and A. S. Botvina, *Eur. Phys. J. A* **17**, 559 (2003).
- [41] K. K. Gudima *et al.* (in preparation).
- [42] C. Samanta, A. S. Botvina, I. N. Mishustin, and W. Greiner (in preparation).
- [43] A. S. Lorente, A. S. Botvina, and J. Pochodzalla, *Phys. Lett. B* **697**, 222 (2011).
- [44] S. Das Gupta, *Nucl. Phys. A* **822**, 41 (2009); V. Topor Pop, S. Das Gupta, *Phys. Rev. C* **81**, 054911 (2010).