Determination of the emission probability of the 21.54 keV γ ray in the decay of ¹⁵¹Sm

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The emission probability of the 21.54 keV γ ray in the decay of ¹⁵¹Sm was experimentally determined by means of the measurements of the 21.54 keV γ -ray emission rate and the activity of ¹⁵¹Sm with high-resolution Ge, Si(Li) γ -ray spectrometry and liquid scintillation counting (LSC), respectively. The resulting emission probability of the 21.54 keV γ ray is 0.0324% with a standard uncertainty of 0.0013% based on a quadratic summation of the uncertainty components from the measurements of the γ -ray emission rate and the activity of ¹⁵¹Sm. The result agrees with the latest literature value (partially theoretically based), 0.0318 \pm 0.0022%, with reduced uncertainty.

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I. INTRODUCTION

¹⁵¹Sm is a long-lived fission product with a fission yield of ~0.4% and can be used as a tracer in the fields of environment and life sciences. ¹⁵¹Sm also is a important nuclide in astrophysics for characterizing nucleosynthesis in the slow neutron capture process, and an excellent thermometer for the *s* process [1]. The present investigation on the emission probability of 21.54 keV γ rays in the decay of ¹⁵¹Sm was undertaken to resolve some of the discrepancies in the literature and to reduce the uncertainties in the existing decay schemes of ¹⁵¹Sm. Furthermore, an accurate value of 21.54 keV γ -ray intensity is needed to accurately assay ¹⁵¹Sm by γ -ray spectrometry for various applications, such as nuclear physics, medicine, and nuclear chemistry.

A decay scheme constructed in Ref. [2] is displayed in Fig. 1 with the new half-life measured by our laboratory [3]. A large majority of β^- transition events go directly to the ground state of ¹⁵¹Eu, only a weak β branch (0.91%) feeds the first excited state at 21.54 keV followed only by a 21.54 keV γ transition to the ground state of ¹⁵¹Eu via the emission of either γ rays or internal conversion electrons. The emission probability of the γ rays depopulating the 21.54 keV state has not been accurately measured up to now principally because of the very weak β branch and the very large internal conversion coefficient (ICC). In the Nuclear Data Sheets, the latest BR value of the 21.54 keV γ ray was calculated using the equation $I_{\beta}/(1+\alpha)$ to be 0.0318 \pm 0.0022% [4], where the β branch I_{β} is 0.91 \pm 0.06% [5,6], and the theoretical ICC α is 27.6 [7] with 1.4% uncertainty coming from Raman-band calculation. The large uncertainties in this value hindered their use in the above-mentioned fields. So, the emission probability of 21.54 keV γ rays in the decay of ¹⁵¹Sm should be experimentally determined in an effort to examine the calculated value and reduce its uncertainty.

In present work, the nuclide ¹⁵¹Sm was produced by the neutron irradiation of a Sm₂O₃ target enriched in ¹⁵⁰Sm to 87.34% in the heavy water research reactor at China Institute of Atomic Energy for ~15 days. The neutron flux was 4.5×10^{13} n/cm²s at the sample irradiation site. The ratio of ¹⁵¹Sm/¹⁵⁰Sm in the activated sample was measured by a thermal ionization mass spectrometer (TIMS) to be ~3.36 × 10^{-3} [3]. The emission probability of the γ ray was obtained by the equation $BR = \gamma PS(21.54)/A$, where the gamma emission rate, $\gamma PS(21.54)$, of the ¹⁵¹Sm sample was measured with both a high-resolution Ge and a Si(Li) spectrometer, and the radioactivity, A, with liquid scintillation counting (LSC).

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II. EXPERIMENTAL

A. Measurement of emission rate of 21.54 keV γ ray in ¹⁵¹Sm decay

After a cooling time for more than 5 yr, an irradiated Sm₂O₃ sample (enriched in ¹⁵⁰Sm) was dissolved with 68% HNO₃. The solution was heated on a hot plate and evaporated to near dryness, then 20 mL 0.01M HNO₃ was added to produce an essentially salt-free ¹⁵¹Sm solution to eliminate the influence of the self-absorption effect on the γ emissivity. The concentration of Sm in sample solution, measured with ICP-MS, is (2.044 ± 0.031) × 10⁻⁴ g Sm/mL. In order to improve the accuracy, two point sources (the diameter of the active zone is ~0.5 mm with an accurately known activity ratio were made of the salt-free ¹⁵¹Sm solution.

The sources were measured by an ORTEC Si(Li) $\gamma(x)$ ray spectrometer system composed of a Si(Li) detector of SLP-10180-P, a DSPEC-PLUS 92X unit, and a computer [8]. The system has an energy resolution of 0.18 keV for a 5.9 keV γ ray. Efficiency functions were established for the energy range of interest (properly covering 21.54 keV) by using 21.99, 22.16, and 23.17 keV x rays of ^{110m}Ag; 20.78 and 21.34 keV x rays and 26.34 keV γ ray of ²⁴¹Am; and the 32.19 keV x ray of ¹³⁷Cs for measurement positions of 0, 1.0, and 2.0 cm from the detector surface. The emission rates of the above $x(\gamma)$ rays were obtained from the accurately known activities of the standard

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FIG. 1. The decay scheme of 151 Sm.

^{110m}Ag, ²⁴¹Am, and ¹³⁷Cs sources and their respective branching ratios from the literature [2]. As shown in Fig. 1, there are no cascade or crossover transitions related to 21.54 keV γ ray, thus no coincidence summing effect on its emission rate measurement. The resulting peak efficiency curves at three counting positions are expressed by Eqs. (1)–(3):

$$\ln (\varepsilon_P) = -9.964590 + 4.83383 \ln (E) - 0.93256 \ln (E)^2 \times 0 \text{ cm } 20 \text{ keV} \leqslant E \leqslant 32 \text{ keV},$$
(1)

$$\ln(\varepsilon_p) = -14.4430 - 7.4269 \ln(E) - 1.36746 \ln(E)^2 \times 1.0 \text{ cm } 20 \text{ keV} \le E \le 32 \text{ keV},$$
(2)

$$\ln(\varepsilon_p) = -12.0512 + 5.43801 \quad \ln(E) - 1.02625 \quad \ln(E)^2 \times 2.0 \text{ cm } 20 \text{ keV} \leqslant E \leqslant 32 \text{ keV}.$$
(3)

The intrinsic full peak efficiency of 21.54 keV, $\varepsilon_p(21.54)$, was calculated using formulas (1), (2), and (3) to be 1.997%, 1.078%, and 0.655% for 0, 1.0, and 2.0 cm counting positions,

respectively, with a relative standard uncertainty of 3.0% derived from the uncertainties of the peak efficiency curves, counting statistics of 21.54 keV peak, and the emission rates of the x rays in the standard sources used for efficiency calibration. A background γ spectrum showed no interference with the measurement of the 21.54 keV peak. A Si(Li) γ (x) spectrum is shown in Fig. 2.

In addition, an *N*-type HPGe-detector system (resolution: 1.8 keV, relative efficiency: 47%) was also used to measure the same samples. The two samples were measured at the position 7 cm from the surface of the HPGe detector for more than 20 h. A γ spectrum is shown in Fig. 3. The detector was previously calibrated for efficiency at the above counting geometry. All results of 21.54 keV γ ray emission rate are listed in Table I. The final value of $\gamma PS = 0.2029 \pm 0.0069$ Bq/mg was obtained by Eq. (4):

$$\overline{\gamma PS} = \frac{\sum_{i}^{N} \gamma PS_{i}/u_{i}^{2}}{\sum_{i}^{N} 1/u_{i}^{2}},$$
(4)

where $\gamma P S_i$ and μ_i stand for the γ emissivities of 21.54 keV γ ray and uncertainties measured in different positions of different detectors. The average 21.54 keV γ emission rate per mg Sm is $\gamma PS = 994.5$, with a relative standard deviation of 3.0%.

B. Activity measurement of ¹⁵¹Sm

¹⁵¹Sm is essentially a pure β^- decay nuclide with a very weak β branch (0.91%) feeding the first excited state of ¹⁵¹Eu at 21.54 keV. The activity of ¹⁵¹Sm was measured by liquid scintillation counting (LSC) which was already mentioned in our former paper [3]. The main difficulty in LSC is the discrimination between the nuclide of interest, ¹⁵¹Sm, and the radioactive contaminants present in the sample solution which can induce complex spectra overlapping and quenching effects. As can be seen from the γ -ray spectrum of the ¹⁵¹Sm sample shown in Fig. 4, ¹⁵²Eu and ¹⁵⁴Eu are the main



FIG. 2. A Si(Li) γ (x)-ray spectrum of a ¹⁵¹Sm sample.



FIG. 3. A HPGe γ -ray spectrum of a ¹⁵¹Sm sample.



FIG. 4. A HPGe γ -ray spectrum of a ¹⁵¹Sm sample.

TABLE I. Results of 21.54 keV γ emission rate per mg Sm [using the Si(Li) and HPGe].

Instrument, sample No., counting site	Peak count rate (s/mg Sm)	Efficiency	Emission rate $(\gamma/s/mg Sm)$	Uncertainty. %
Si(Li), 1, 0 cm	18.89	0.01938	975	3.1
Si(Li), 1, 1 cm	10.72	0.01078	994	3.1
Si(Li), 1, 2 cm	6.738	0.00669	1007	3.1
Si(Li), 2, 0 cm	18.14	0.01938	935	3.2
Si(Li), 2, 1 cm	10.83	0.01078	1004	3.1
Si(Li), 2, 2 cm	6.957	0.00669	1040	3.5
HPGe, 1, 7 cm	17.33	0.01726	1004	3.1
HPGe, 2, 7 cm	17.19	0.01726	995	3.2
Average 21.54 keV γ emission rate per mg Sm \pm RSD			994.5	3.0



FIG. 5. A LSC entire spectrum of a ¹⁵¹Sm sample (log-log coordinate).

 β^{-} - γ -emitting contaminants. The activity of ¹⁵²Eu in a 1-mg Sm sample, $A(^{152}$ Eu), was calculated using formula (5) to be 3.74 \times 10³ Bq, with a relative standard uncertainty of 1.3% mainly derived from the uncertainties of the peak efficiency curve.

$$A(^{152}\text{Eu}) = \text{Cps}(344.3) / [\varepsilon(344.3) B(344.3)],$$
(5)

where Cps(344.3), ε (344.3), and B(344.3) stand for peak counts per second, peak efficiency, and branching ratio of 344.3 keV γ ray in the decay of ¹⁵²Eu, respectively. The activity of ¹⁵⁴Eu was calculated using the same method to be 4.18×10^3 Bq/mg Sm sample, with a relative standard uncertainty of 1.3%. Several other γ rays of these two nuclides (such as 244.7 keV, 1085.8 keV, 1112.1 keV γ ray of ¹⁵²Eu and 591.7 keV, 723.3 keV, 1004.8 keV γ ray of ¹⁵⁴Eu) were also used in calculating the activities of ¹⁵²Eu and ¹⁵⁴Eu. and the results are consistent with the previous results within uncertainty limits, which confirms that the total activity of 152 Eu and 154 Eu was $< 8 \times 10^3$ Bq/mg Sm sample and can be neglected compared to the activity of 151 Sm ($\sim 3 \times 10^6$ Bq/mg Sm sample). Furthermore, the end-point energy for 151 Sm $\beta^$ decay is 76.7 keV, while those for ¹⁵²Eu and ¹⁵⁴Eu β^- decay are \sim 500 keV; therefore the contribution from ¹⁵²⁺¹⁵⁴Eu to the energy region of ¹⁵¹Sm in the LSC spectrum will be even smaller.

A model 3170 TR/SC Tri-Carb LSC spectrometer with low-background double photomultiplier tubes was used to measure the activity of ¹⁵¹Sm. The detection efficiency is a function of the β^- spectrum of the radionuclide and the

light yield of the scintillator, which depends on the LSC cocktail itself but also on the chemical composition of the source. The LSC was calibrated by the efficiency tracing method using a ¹⁴C standard source ($A = 1896 \pm 19$ DPS, certified by China National Institute of Metrology) with the same chemical and physical parameters (i.e., volume, source-to-cocktail volume ratio, acidity concentration) and quenching agent (nitromethane) as for the ¹⁵¹Sm sample. Blank contribution was corrected for by using the same LS cocktail containing natural Sm. A ¹⁵¹Sm source containing 26.1 mL sample solution (containing 5.335 mg Sm) was measured, and a LSC spectrum is shown in Fig. 5. Taking into account the detection efficiency, the specific activity of ¹⁵¹Sm is therefore (3.07 ± 0.08) × 10⁶ Bq/mg Sm.

III. RESULTS AND DISCUSSION

The emission probability of the 21.54 keV γ ray of ¹⁵¹Sm, *BR*, was calculated from the absolute γ ray emission rates $\gamma PS(21.54)$ and the activity (*A*) of ¹⁵¹Sm in samples with the same mass of Sm by Eq. (6):

$$BR = \gamma P S(21.54)/A,\tag{6}$$

where $\gamma PS(21.54)$ (=994.5 ± 29.9 Bq/mg) is the specific γ emissivity of 21.54 keV γ ray, A (=3.07 ± 0.08 × 10⁶ Bq/mg) is the specific activity of ¹⁵¹Sm. The emission probability of 21.54 keV γ ray of ¹⁵¹Sm is thus 0.0324 ± 0.0013%. The uncertainty budget is given in Table II.

TABLE II. Uncertainty budget in measurement of emission probability of 21.54 keV γ ray in the decay of ¹⁵¹Sm.

	Uncertainty component	Relative standard uncertainty (%)
γ ray emission rates	Counting statistics	0.5
, ,	Full energy peak efficiency	3.0
Activity of ¹⁵¹ Sm	Impurities	2.3
	Detection efficiency for ¹⁵¹ Sm	1.0
	Counting statistics	0.2
	Sample weighing	0.1
	Total	3.9

The result of the emission probability of 21.54 keV γ ray in the decay of ¹⁵¹Sm from this work, $0.0324 \pm 0.0013\%$, agrees with that of the latest literature value, $0.0318 \pm 0.0022\%$ [4], within uncertainty limits. The relative standard uncertainty of our result is reduced by ~50%, relative. More importantly, as mentioned above, since the literature value was derived from a weak β^- branching ratio (with a large uncertainty) and a theoretical ICC, the cited uncertainty may be much smaller

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than it should be, while the result from the present work was obtained completely by experimental measurements.

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