Direct photon calculations in heavy-ion collisions at $\sqrt{s_{NN}} = 62.4A$ to 200A GeV in a (3 + 1)-dimensional hybrid approach

Bjørn Bäuchle^{*} and Marcus Bleicher

Frankfurt Institute for Advanced Studies, Frankfurt am Main, Germany and Institut für Theoretische Physik, Goethe-Universität, Frankfurt am Main, Germany (Received 6 September 2010; revised manuscript received 17 November 2010; published 7 December 2010)

Direct photon spectra from central Au + Au and Cu + Cu collisions at $\sqrt{s_{NN}} = 62.4, 130$, and 200A GeV are

calculated within the ultrarelativistic quantum molecular-dynamics microscopic transport model and a micro + macro hybrid model. In the latter approach, the high-density part of the transport evolution is replaced by an ideal (3 + 1)-dimensional hydrodynamic calculation. We study the impact of viscosity and full local thermalization and compare the calculations to measurements obtained by the PHENIX Collaboration. We find a reasonable agreement with the experimental data for calculations involving a quark-gluon-plasma phase.

DOI: 10.1103/PhysRevC.82.064901

PACS number(s): 25.75.Cj

I. INTRODUCTION

Heavy-ion physics is widely used as a tool for the exploration of the phase diagram of strongly interacting matter. In the collision of heavy nuclei, the nucleons may be compressed and heated sufficiently to create a new state of matter that consists of partonic degrees of freedom, the quark-gluon-plasma (QGP) phase [1,2]. Indeed, proposed signatures for the QGP, like strong jet quenching and large elliptic flow, have been found by experiments at the Brookhaven National Laboratory Relativistic Heavy Ion Collider (BNL-RHIC) [3–8].

Inferring knowledge about the central regions of a heavyion collision is very difficult, since even if a plasma is created, its lifetime and size are beyond the experimental reach for direct observation, so we are limited to the study of particles that are emitted from the reaction zone. Unfortunately, firstprinciples calculations of quantum chromodynamics (QCD) processes are only possible if all involved scales are much larger than the QCD scale, $\Lambda_{QCD} \approx 0.2$ GeV. However, in a heavy-ion collision, most particles have momenta comparable to Λ_{QCD} . Therefore, more phenomenological approaches are necessary to explore the bulk of the matter.

While the abundance of hadronic particles that are produced in a heavy-ion collision are emitted at the end of the reaction and carry only indirect information from the early stages, electromagnetic probes allow for an undisturbed view into all stages of the reaction. Photons and leptons escape the reaction zone without rescattering due to their very small cross section, but for the same reason, their abundances are rather low compared to hadronic species [9].

Three different electromagnetic particle species are currently being measured in heavy-ion experiments: single- and dielectrons, single- and dimuons, and photons. Direct photons have the advantage that they are created in scatterings of the partonic or hadronic medium and are therefore directly coupled to the region of interaction. The leptons, however, are usually created in pairs, either in the (initial-state) Drell-Yan process or by the decay of hadrons. In addition, one of the leptons might be a neutrino, which escapes observation. Since this process is governed by the weak interaction, the decay usually happens outside the fireball. Single leptons are used, therefore, to reconstruct weakly decaying heavy quarks, while the invariant mass distribution of dileptons can be used to extract spectral functions of vector mesons.

Previous calculations of direct photons from transport theory include work with the ultrarelativistic quantum moleculardynamics (UrQMD) model by Dumitru *et al.* [10] and Bäuchle *et al.* [11] and with hadron-string dynamics (HSD) by Bratkovskaya *et al.* [12]. Hydrodynamics has been used in many direct photon calculations; see, for example [13–20].

The extraction of the yield of photons from the fireball (direct photons) is hindered by a huge background of photons from hadronic decays outside the fireball, which is dominated by the π^0 and η decays. However, experimental techniques for the extraction of direct photon yields are well developed and allow us to disentangle these late-stage contributions from the scattering contribution. The experimental methods include a direct estimation of the background via invariant mass analysis of the photons [21,22], the analysis of interference patterns (using a Hanbury-Brown–Twiss analysis) [23], and the extrapolation of the spectra of low-mass dileptons to the photon point [24].

In this paper, we apply a previously established model for direct photon emission from hadronic and partonic sources [11] and apply it to collision systems measured by the STAR and PHENIX Collaborations at BNL-RHIC. In Sec. II, we briefly introduce the model and the parameters used for the present calculations, and in Sec. III we show the direct photon spectra obtained with our calculations as well as comparisons to the available data from the PHENIX Collaboration [25,26].

II. THE MODEL

In the present work, direct photon spectra are calculated in the framework of the microscopic UrQMD transport model [27–29] using the hybrid option introduced in version 3.3 [30–33]. While UrQMD itself is a hadronic transport model that includes only hadronic and string degrees of freedom and employs PYTHIA [34] for scatterings at high momentum

^{*}baeuchle@th.physik.uni-frankfurt.de

TABLE I. The critical energy densities for the mapping from hydrodynamics to transport theory for the various equations of state (EOS). The nuclear ground-state energy density is $\epsilon_0 = 146 \text{ MeV/fm}^3$.

| EOS | $\epsilon_{ m crit}$ |
|--------|----------------------|
| HG-EOS | $5\epsilon_0$ |
| χ-EOS | $7\epsilon_0$ |
| BM-EOS | $5\epsilon_0$ |

transfer, the hybrid option allows us to substitute the highdensity part of the evolution by a (3 + 1)-dimensional ideal hydrodynamic [31] description. In this part, other-than-hadronic degrees of freedom and phase transitions may be included.

The inclusion of an intermediate phase into the model raises the need for two interfaces to go from the particlebased description of the transport model to the density-based description of the hydrodynamic model and back again.

The mapping from transport simulation to hydrodynamics is performed at $t_{\text{start}} = 0.6$ fm. Here, the energy-density, baryon number density, and momentum densities are calculated from all particles at midrapidity. Particles with a rapidity |y| > 2 are propagated in the cascade and do not interact with the bulk medium.

The transition from hydrodynamics back to the cascade proceeds gradually, mapping the temperatures and chemical potentials to particles via the Cooper-Frye formula [35] when all cells in the same transverse slice (i.e., at the same position



FIG. 1. (Color online) Comparison of the data from the PHENIX Collaboration [25] (black squares) with cascade calculations (red solid lines) for central to peripheral collisions. The green dash-dotted lines show the sum of pQCD calculations [25,41] and the cascade contribution. For the most central collisions, 00%–10% and 10%–20%, the spectra from hybrid calculations with the BM-EOS plus pQCD contribution are shown (violet dotted lines).

TABLE II. Fit results for the low- p_{\perp} part ($p_{\perp} < 2.5$ GeV) of the cascade calculations of Au + Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV (see Fig. 1). The fit function is $f(p_{\perp}) = A \exp(-\frac{p_{\perp}}{T_{\text{slope}}})$, and d.o.f. denotes degrees of freedom.

| Centrality | T _{slope} (MeV) | $A (\text{GeV}^{-2})$ | χ^{2} (d.o.f.) |
|------------|--------------------------|----------------------------------|---------------------|
| 00%-10% | 231.9 ± 9.4 | 2.39 ± 0.67 | 0.038 |
| 00%-92% | 231.4 ± 8.5 | 0.41 ± 0.11 | 0.032 |
| 10%-20% | 234.0 ± 10.0 | 1.26 ± 0.37 | 0.041 |
| 20%-30% | 239.0 ± 11.4 | 0.56 ± 0.18 | 0.049 |
| 30%-40% | 239.0 ± 13.1 | 0.27 ± 0.10 | 0.065 |
| 40%-50% | 243.0 ± 13.4 | 0.12 ± 0.04 | 0.064 |
| 50%-60% | 235.4 ± 8.8 | $(5.64 \pm 1.43) \times 10^{-2}$ | 0.032 |
| 60%-92% | 250.5 ± 11.8 | $(6.91 \pm 2.08) \times 10^{-3}$ | 0.044 |

along the beam direction) have diluted below a critical energy density (see Table I). After the transition to the cascade, rescatterings and decays are calculated in the well-known UrQMD model. For more detailed information on the hybrid model, the reader is referred to [32,36].

A. Equations of state

Three different equations of state (EOS) are compared in this work. The effects of thermalization at the transition from the initial-stage cascade to hydrodynamics can be explored with the hadron gas EOS (HG-EOS) [37], which has the same degrees of freedom as the transport phase. To investigate the effects of partonic matter and a phase transition, we use two different models for the EOS: The chiral equation of state χ -EOS [38] has a crossover phase transition to chirally restored and deconfined matter, while the bag model equation



FIG. 2. (Color online) Comparison of data from the PHENIX Collaboration [26] (black squares) to cascade calculations (red crosses) and hybrid-model calculations with HG-EOS (blue solid lines), χ -EOS (orange dashed lines), and BM-EOS (violet dotted lines) for central (0%–20%) and midcentral (20%–40%) collisions. The contribution from initial pQCD scatterings [26,41] have been added to all spectra. The spectra from central collisions have been scaled by a factor of 10³ to enhance readability.



FIG. 3. (Color online) Direct photon spectra calculated with the hybrid model and HG-EOS (solid blue lines), χ -EOS (dashed orange lines), and BM-EOS (dotted violet lines) without prompt photon contribution. The left panels show calculations for $\sqrt{s_{\text{NN}}} = 62.4$ GeV, the middle panels show calculations for $\sqrt{s_{\text{NN}}} = 130$ GeV, and the right panel shows calculations for $\sqrt{s_{\text{NN}}} = 200$ GeV. The upper panels show calculations for Au + Au collisions, while the lower panels show calculations for Cu + Cu collisions. In each panel, the upper curves are central collisions (00%–20%) and the lower curves are midcentral collisions (20%–40%).

of state BM-EOS [31] has a first-order phase transition to a QGP. In both equations of state, the transition happens at around $T_{\rm C} \approx 170$ MeV. A comparison between the chiral equation of state and lattice QCD results can be found in Ref. [39].

B. Photon emission sources

Due to the small creation probability of direct photons, their emission is calculated perturbatively. That is, the evolution of the underlying event remains unaltered by the emission of direct photons.

The set of channels for direct photon production differs in the transport and hydrodynamic parts of the model. The most important channels, though, are common to both parts, namely $\pi\pi \rightarrow \gamma\rho$ and $\pi\rho \rightarrow \gamma\pi$. Besides photon emission from the QGP, channels with strangeness are included in the hydrodynamic part. The corresponding rates for hadronic photon emission from each hydrodynamic cell are taken from Turbide *et al.* [15]. For the partonic emission, the parametrizations are from Arnold *et al.* [40]. In the transport part, additional processes including an η meson are included. The corresponding cross sections have been calculated by Kapusta *et al.* [13]. Although Kapusta and Turbide use different Lagrangians to derive their cross sections and rates, earlier investigations (see [11]) have shown that the thermal rates that can be extracted from Kapusta's cross sections using this model agree very well with those parametrized by Turbide *et al.* The same investigations have shown that the contributions of the hadronic processes that are not common to both models contribute about equally, but not significantly to the final spectra. The numerical implementation for direct photon emission is explained in detail in [11].

At high transverse momenta, another source becomes important, namely the prompt contribution from hard scatterings of partons in the initial nuclei. The spectra predicted by next-to-leading-order processes in QCD (NLO-pQCD) calculations from Gordon and Vogelsang [41] fit the experimental data from the PHENIX Collaboration [25] rather well at high p_{\perp} . Therefore, the pQCD contributions from [41], scaled by the number of binary collisions $\langle N_{\rm coll} \rangle$, are added to the soft photons calculated here.

It has been pointed out by Fries *et al.* [42] and Qin *et al.* [43] that at intermediate transverse momentum, jet-quenching and jet-medium interactions might increase the direct photon yield from hard pQCD processes. The effects of these processes are neglected in the current work.

TABLE III. Fit results for the low- p_{\perp} part ($p_{\perp} < 2.5 \text{ GeV}$) of the spectra from central (0%–20%) and midcentral (20%–40%) Au + Au collisions. The fit function is $f(p_{\perp}) = A \exp(-\frac{p_{\perp}}{T_{\text{slope}}})$. The data are shown in Fig. 2 (for $\sqrt{s_{\text{NN}}} = 200 \text{ GeV}$) and Fig. 3 ($\sqrt{s_{\text{NN}}} = 62.4 \text{ GeV}$, upper left panel and $\sqrt{s_{\text{NN}}} = 130 \text{ GeV}$, upper central panel).

| $\sqrt{s_{ m NN}}$ | EOS | Centrality | $T_{\rm slope}$ (MeV) | $A (\text{GeV}^{-2})$ | $\frac{\chi^2}{\text{d.o.f.}}$ |
|--------------------|-----------|------------|-----------------------|-----------------------|--------------------------------|
| 200 | Transport | 0%-20% | 232.5 ± 9.8 | 1.65 ± 0.48 | 0.041 |
| 200 | HG-EOS | 0%-20% | 246.7 ± 8.6 | 3.63 ± 0.83 | 0.025 |
| 200 | χ-EOS | 0%-20% | 261.9 ± 8.7 | 10.13 ± 2.05 | 0.020 |
| 200 | BM-EOS | 0%-20% | 251.4 ± 9.7 | 16.37 ± 4.03 | 0.029 |
| 200 | Transport | 20%-40% | 237.3 ± 12.1 | 0.38 ± 0.13 | 0.057 |
| 200 | HG-EOS | 20%-40% | 243.4 ± 8.3 | 1.32 ± 0.30 | 0.025 |
| 200 | χ-EOS | 20%-40% | 253.0 ± 8.0 | 4.11 ± 0.82 | 0.020 |
| 200 | BM-EOS | 20%-40% | 240.6 ± 9.0 | 7.61 ± 1.90 | 0.030 |
| 130 | Transport | 0%-20% | 232.5 ± 9.1 | $(9.87 \pm 2.67)^{a}$ | 0.035 |
| 130 | HG-EOS | 0%-20% | 246.3 ± 8.5 | 3.42 ± 0.66 | 0.024 |
| 130 | χ-EOS | 0%-20% | 261.2 ± 8.5 | 9.67 ± 1.93 | 0.019 |
| 130 | BM-EOS | 0%-20% | 250.2 ± 9.6 | 15.84 ± 3.88 | 0.039 |
| 130 | Transport | 20%-40% | 257.2 ± 11.3 | $(5.48 \pm 1.50)^{b}$ | 0.036 |
| 130 | HG-EOS | 20%-40% | 242.4 ± 7.6 | 1.26 ± 0.26 | 0.021 |
| 130 | χ-EOS | 20%-40% | 252.7 ± 7.9 | 4.01 ± 0.80 | 0.019 |
| 130 | BM-EOS | 20%-40% | 240.6 ± 8.8 | 7.46 ± 1.82 | 0.029 |
| 62.4 | Transport | 0%-20% | 242.1 ± 13.5 | $(5.29 \pm 1.95)^{a}$ | 0.066 |
| 62.4 | HG-EOS | 0%-20% | 247.3 ± 8.1 | 3.19 ± 0.67 | 0.022 |
| 62.4 | χ-EOS | 0%-20% | 261.8 ± 8.2 | 9.24 ± 1.78 | 0.018 |
| 62.4 | BM-EOS | 0%-20% | 250.3 ± 9.5 | 15.13 ± 3.65 | 0.028 |
| 62.4 | Transport | 20%-40% | 232.8 ± 9.4 | $(4.18 \pm 1.16)^{b}$ | 0.038 |
| 62.4 | HG-EOS | 20%-40% | 245.8 ± 8.0 | 1.21 ± 0.26 | 0.022 |
| 62.4 | χ-EOS | 20%-40% | 253.9 ± 7.7 | 3.82 ± 0.73 | 0.018 |
| 62.4 | BM-EOS | 20%-40% | 240.8 ± 8.6 | 7.33 ± 1.74 | 0.028 |

 $^{a} \times 10^{-2}$. $^{b} \times 10^{-3}$.

III. RESULTS

The comparison between direct photon spectra at low and intermediate transverse momentum p_{\perp} from cascade calculations and data from the PHENIX Collaboration [25] for Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV is shown in Fig. 1. One clearly observes that the hadronic transport model (full lines) does not saturate the upper limits of the experimental data. In all centrality bins, the prompt photon yield is significantly larger than predicted by the hadronic cascade. The ratio between pQCD and hadronic contributions is fairly constant among the centrality bins. For comparison, Fig. 1 also shows the spectra obtained with the hybrid model using the BM-EOS for the two most central bins, 00%–10% and 10%–20%, which agree nicely with the data. Thermal fits to the low- p_{\perp} parts of the cascade spectra show inverse slope parameters of $T_{slope} \approx 235$ MeV throughout the centrality bins; see Table II.

A more detailed exploration of the low- p_{\perp} part of the direct photon calculation is shown in Fig. 2. Here, the low- p_{\perp} data obtained by extrapolating the dilepton yield to zero invariant mass [26] for central (00%–20%) and midcentral (20%–40%) collisions are shown in comparison to cascade calculations (red crosses) and hybrid calculations with HG-EOS (solid blue lines), χ -EOS (dashed orange lines), and BM-EoS (dotted violet lines) and prompt (pQCD) photon calculations. All calculated spectra include the $\langle N_{coll} \rangle$ -scaled prompt photon contribution. In both centrality bins, the direct photon spectra obtained with the BM-EOS and χ -EOS, which include a phase transition to a deconfined state of matter, are significantly higher than the hadronic HG-EOS calculations and agree with the measured data.

A similar picture presents itself in Au + Au collisions at lower incident energy $\sqrt{s_{\text{NN}}} = 62.4$ and 130 GeV, shown in the upper panels of Fig. 3. The cascade calculations have been omitted from the figure for clarity.

Thermal fits to the spectra (see Table III) show inverse slope parameters in the range $233 < T_{slope} < 262$ MeV, with the cascade calculations showing the smallest and the χ -EOS hybrid calculations showing the largest values of T_{slope} . HG-EOS and BM-EOS calculations show similar inverse slope parameters. The integrated yield *A* is highest in BM-EOS hybrid calculations. The spectra from the hybrid calculations are rather similar for the different beam energies.

Hybrid-model calculations for central (0%-20%) and midcentral (20%-40%) Cu + Cu collisions are shown in the lower panels of Fig. 3 for all EOS. The thermal fits (see Table IV) again show no significant energy dependence of inverse slope parameter T_{slope} or yield A. We observe a clear ordering of the total yield between the EOS, with yield from the BM-EOS calculations being higher than that of the χ -EOS calculations, and both yields exceeding that of HG-EOS calculations. However, the inverse slope parameters are similar in HG-EOS _._.

| TABLE IV. Fit results for the low- p_{\perp} part ($p_{\perp} < 2.5 \text{ GeV}$) of the spectra from central (0%–20%) and midcentral (20%–40%) Cu + 0 | Cu |
|--|----|
| collisions. The fit function is $f(p_{\perp}) = A \exp(-\frac{p_{\perp}}{T_{\text{slope}}})$. The data are shown in Fig. 3, lower panels. | |

- - - --

| $\sqrt{s_{ m NN}}$ | EOS | Centrality | T_{slope} (MeV) | $A (\text{GeV}^{-2})$ | $\frac{\chi^2}{\text{d.o.f.}}$ |
|--------------------|--------|------------|--------------------------|-----------------------|--------------------------------|
| 200 | HG-EOS | 0%-20% | 252.0 ± 9.6 | $(4.84 \pm 1.38)^{a}$ | 0.057 |
| 200 | χ-EOS | 0%-20% | 251.5 ± 7.3 | 1.77 ± 0.39 | 0.033 |
| 200 | BM-EOS | 0%-20% | 237.7 ± 7.8 | 3.61 ± 0.94 | 0.047 |
| 200 | HG-EOS | 20%-40% | 254.6 ± 13.2 | $(1.61 \pm 0.62)^{a}$ | 0.103 |
| 200 | χ-EOS | 20%-40% | 242.9 ± 7.0 | $(7.25 \pm 1.63)^{a}$ | 0.036 |
| 200 | BM-EOS | 20%-40% | 229.2 ± 7.5 | 1.60 ± 0.43 | 0.051 |
| 130 | HG-EOS | 0%-20% | 250.0 ± 9.3 | $(4.78 \pm 1.35)^{a}$ | 0.056 |
| 130 | χ-EOS | 0%-20% | 250.9 ± 7.1 | 1.76 ± 0.37 | 0.031 |
| 130 | BM-EOS | 0%-20% | 238.1 ± 7.9 | 3.56 ± 0.93 | 0.048 |
| 130 | HG-EOS | 20%-40% | 240.4 ± 7.7 | $(1.99 \pm 0.50)^{a}$ | 0.044 |
| 130 | χ-EOS | 20%-40% | 242.8 ± 7.1 | $(6.99 \pm 1.59)^{a}$ | 0.036 |
| 130 | BM-EOS | 20%-40% | 228.5 ± 7.7 | 1.58 ± 0.44 | 0.054 |
| 62.4 | HG-EOS | 0%-20% | 248.2 ± 7.7 | $(4.71 \pm 1.11)^{a}$ | 0.039 |
| 62.4 | χ-EOS | 0%-20% | 250.2 ± 6.8 | 1.71 ± 0.35 | 0.029 |
| 62.4 | BM-EOS | 0%-20% | 236.8 ± 7.4 | 3.52 ± 0.88 | 0.044 |
| 62.4 | HG-EOS | 20%-40% | 242.8 ± 6.9 | $(1.87 \pm 0.41)^{a}$ | 0.034 |
| 62.4 | χ-EOS | 20%-40% | 241.7 ± 6.3 | $(6.71 \pm 1.37)^{a}$ | 0.029 |
| 62.4 | BM-EOS | 20%-40% | 227.0 ± 6.7 | 1.62 ± 0.40 | 0.042 |
| . 10 1 | | | | | |

 $^{a} \times 10^{-1}$.

and χ -EOS calculations but significantly lower in BM-EOS calculations.

is assumed, and are dominant throughout all p_{\perp} if a purely hadronic scenario is assumed.

IV. SUMMARY

We examined the direct photon spectra obtained with a transport and a transport + hydrodynamics hybrid model for collisions of Au + Au and Cu + Cu at energies of $\sqrt{s_{\rm NN}} = 62.4, 130$, and 200 GeV. We find that the hadronic models (transport model and hybrid model with HG-EoS) underpredict the data, while calculations with a deconfined state of matter (hybrid model with χ -EOS or BM-EOS) fit the data much better.

Thermal fits to the data show no significant beam energy dependence on the spectra. The inverse slope parameters obtained by fitting the low-transverse momentum part of the spectra are in the range of $227 < T_{slope} < 262$ MeV, which is significantly above the expected transition temperature to deconfined matter.

Prompt photons from the initial early hard proton-proton scatterings are found to be a significant source of direct photon emission above $p_{\perp} = 3.5$ GeV if an EOS with phase transition

V. OUTLOOK

Future work with this model will include the extraction of radial and elliptic flow parameters v_1 and v_2 for more differential analyses. Also, the influence of changing the criteria for the transition between the transport and hydrodynamic phases in the hybrid model will be examined in the future.

ACKNOWLEDGMENTS

This work has been supported by the Frankfurt Center for Scientific Computing (CSC), the GSI, and the BMBF. B.B. gratefully acknowledges support from the Deutsche Telekom Stiftung, the Helmholtz Research School on Quark Matter Studies, and the Helmholtz Graduate School for Hadron and Ion Research. This work was supported by the Hessian LOEWE initiative through the Helmholtz International Center for FAIR. The authors thank Elvira Santini for valuable discussions and Henner Büsching for experimental clarifications.

- [1] J. W. Harris and B. Muller, Annu. Rev. Nucl. Part. Sci. 46, 71 (1996).
- [2] S. A. Bass, M. Gyulassy, H. Stoecker, and W. Greiner, J. Phys. G 25, R1 (1999).
- [3] S. S. Adler et al. (PHENI Collaboration), Phys. Rev. Lett. 91, 072301 (2003).
- [4] S. S. Adler et al. (PHENI Collaboration), Phys. Rev. C 69, 034910 (2004).
- [5] J. Adams et al. (STAR Collaboration), Nucl. Phys. A 757, 102 (2005).
- [6] B. B. Back et al., Nucl. Phys. A 757, 28 (2005).
- [7] I. Arsene et al. (BRAHMS Collaboration), Nucl. Phys. A 757, 1 (2005).
- [8] K. Adcox et al. (PHENIX Collaboration), Nucl. Phys. A 757, 184 (2005).
- [9] C. Gale, arXiv:0904.2184 [hep-ph].

BJØRN BÄUCHLE AND MARCUS BLEICHER

- [10] A. Dumitru, M. Bleicher, S. A. Bass, C. Spieles, L. Neise, H. Stocker, and W. Greiner, Phys. Rev. C 57, 3271 (1998).
- [11] B. Bauchle and M. Bleicher, Phys. Rev. C 81, 044904 (2010).
- [12] E. L. Bratkovskaya, S. M. Kiselev, and G. B. Sharkov, Phys. Rev. C 78, 034905 (2008).
- [13] J. I. Kapusta, P. Lichard, and D. Seibert, Phys. Rev. D 44, 2774 (1991); 47, 4171(E) (1993).
- [14] D. K. Srivastava and B. Sinha, Phys. Rev. C 64, 034902 (2001).
- [15] S. Turbide, R. Rapp, and C. Gale, Phys. Rev. C 69, 014903 (2004).
- [16] S. Turbide, C. Gale, S. Jeon, and G. D. Moore, Phys. Rev. C 72, 014906 (2005).
- [17] F. M. Liu, T. Hirano, K. Werner, and Y. Zhu, J. Phys. G 36, 064072 (2009).
- [18] F. M. Liu, T. Hirano, K. Werner, and Y. Zhu, Phys. Rev. C 80, 034905 (2009).
- [19] K. Dusling, Nucl. Phys. A 839, 70 (2010).
- [20] K. Dusling and I. Zahed, arXiv:0911.2426 [nucl-th].
- [21] M. M. Aggarwal *et al.* (WA98 Collaboration), arXiv:nucl-ex/0006007.
- [22] M. M. Aggarwal *et al.* (WA98 Collaboration), Phys. Rev. Lett. 85, 3595 (2000).
- [23] M. M. Aggarwal *et al.* (WA98 Collaboration), Phys. Rev. Lett. 93, 022301 (2004).
- [24] A. Adare *et al.* (PHENIX Collaboration), Phys. Rev. C 81, 034911 (2010).
- [25] S. S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. Lett. 94, 232301 (2005).
- [26] A. Adare *et al.* (PHENIX Collaboration), Phys. Rev. Lett. 104, 132301 (2010).

PHYSICAL REVIEW C 82, 064901 (2010)

- [27] S. A. Bass et al., Prog. Part. Nucl. Phys. 41, 255 (1998).
- [28] M. Bleicher et al., J. Phys. G 25, 1859 (1999).
- [29] H. Petersen, M. Bleicher, S. A. Bass, and H. Stocker, arXiv:0805.0567 [hep-ph].
- [30] D. H. Rischke, S. Bernard, and J. A. Maruhn, Nucl. Phys. A 595, 346 (1995).
- [31] D. H. Rischke, Y. Pursun, and J. A. Maruhn, Nucl. Phys. A 595, 383 (1995); 596, 717(E) (1996).
- [32] H. Petersen, J. Steinheimer, G. Burau, M. Bleicher, and H. Stocker, Phys. Rev. C 78, 044901 (2008).
- [33] UrQMD v3.3 is available at [http://urqmd.org/].
- [34] T. Sjostrand, S. Mrenna, and P. Z. Skands, J. High Energy Phys. 05 (2006) 026.
- [35] F. Cooper and G. Frye, Phys. Rev. D 10, 186 (1974).
- [36] J. Steinheimer, V. Dexheimer, M. Bleicher, H. Petersen, S. Schramm, and H. Stocker, Phys. Rev. C 81, 044913 (2010).
- [37] D. Zschiesche, S. Schramm, J. Schaffner-Bielich, H. Stoecker, and W. Greiner, Phys. Lett. B 547, 7 (2002).
- [38] J. Steinheimer, S. Schramm, and H. Stocker, arXiv:0909.4421 [hep-ph].
- [39] J. Steinheimer, S. Schramm, and H. Stocker, arXiv:1009.5239 [hep-ph].
- [40] P. B. Arnold, G. D. Moore, and L. G. Yaffe, J. High Energy Phys. 12 (2001) 009.
- [41] L. E. Gordon and W. Vogelsang, Phys. Rev. D 48, 3136 (1993).
- [42] R. J. Fries, B. Muller, and D. K. Srivastava, Phys. Rev. Lett. 90, 132301 (2003).
- [43] G. Y. Qin, J. Ruppert, C. Gale, S. Jeon, and G. D. Moore, Phys. Rev. C 80, 054909 (2009).