# **Tensor part of the Skyrme energy density functional. II. Deformation properties of magic and semi-magic nuclei**

M. Bender,<sup>1,2</sup> K. Bennaceur,<sup>3</sup> T. Duguet,<sup>4,5</sup> P.-H. Heenen,<sup>6</sup> T. Lesinski,<sup>3,7,8</sup> and J. Meyer<sup>3</sup>

<sup>1</sup>*Universite Bordeaux, Centre d'Etudes Nucl ´ eaires de Bordeaux Gradignan, UMR5797, F-33175 Gradignan, France ´*

<sup>2</sup>*CNRS/IN2P3, Centre d'Etudes Nucleaires de Bordeaux Gradignan, UMR5797, F-33175 Gradignan, France ´*

<sup>3</sup>*Universite de Lyon, F-69003 Lyon, France; Institut de Physique Nucl ´ eaire de Lyon, CNRS/IN2P3, Universit ´ e Lyon 1, ´*

*F-69622 Villeurbanne Cedex, France*

<sup>4</sup>*CEA, Irfu, SPhN, Centre de Saclay, F-911191 Gif-sur-Yvette, France*

<sup>5</sup>*National Superconducting Cyclotron Laboratory and Department of Physics and Astronomy, Michigan State University,*

*East Lansing, Michigan 48824, USA*

<sup>6</sup>*PNTPM, CP229, Universite Libre de Bruxelles, B-1050 Bruxelles, Belgium ´*

<sup>7</sup>*Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996, USA*

<sup>8</sup>*Physics Division, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA*

(Received 20 September 2009; published 3 December 2009)

We study systematically the impact of the time-even tensor terms of the Skyrme energy density functional, i.e., terms bilinear in the spin-current tensor density, on deformation properties of closed-shell nuclei corresponding to 20, 28, 40, 50, 82, and 126 neutron or proton shell closures. We compare results obtained with three different families of Skyrme parametrizations whose tensor terms have been adjusted on properties of spherical nuclei: (i) T*IJ* interactions proposed in the first article of this series [T. Lesinski *et al.*, Phys. Rev. C **76**, 014312 (2007)] that were constructed through a complete readjustment of the rest of the functional and (ii) parametrizations whose tensor terms have been added perturbatively to existing Skyrme interactions, with or without readjusting the spin-orbit coupling constant. We analyze in detail the mechanisms at play behind the impact of tensor terms on deformation properties and how studying the latter can help screen out unrealistic parametrizations. It is expected that findings of the present article are to a large extent independent of remaining deficiencies of the central and spin-orbit interactions and will be of great value for the construction of future, improved energy functionals.

DOI: [10.1103/PhysRevC.80.064302](http://dx.doi.org/10.1103/PhysRevC.80.064302) PACS number(s): 21*.*30*.*Fe, 21*.*10*.*Dr, 21*.*10*.*Pc, 21*.*60*.*Jz

## **I. INTRODUCTION**

Our experimental knowledge of the evolution of shell structure in atomic nuclei as a function of proton and neutron numbers has largely increased over the past few years. The difficulty to describe these new results has triggered the search for mechanisms that could alter nuclear shell structure when going toward unstable nuclides and approaching nucleon drip lines. One such mechanism that has an impact on the shell structure of stable and unstable nuclei is provided by the tensor force between nucleons [\[1\]](#page-27-0).

The tensor force is a key ingredient of all vacuum nucleonnucleon interactions. It is also explicitly included in the shell-model interaction for Hamiltonians constructed from first principles. By contrast, it was absent from methods based on the introduction of a self-consistent mean field [\[2\]](#page-27-0), until recent studies based on the Skyrme or Gogny modeling of the in-medium strong interaction [\[3–8\]](#page-27-0).

The renewed interest in the residual tensor interaction is due to its very specific effect on nuclear spectra. It brings a correction to binding energies and to spin-orbit splittings that fluctuates with the filling of shells. Its introduction seems, therefore, necessary to improve the predictive power of meanfield-based methods. The Skyrme and Gogny parametrizations are viewed today as nuclear energy density functionals (EDF). Their derivation from first principles is still lacking, although significant advances have been made recently for the pairing part [\[9\]](#page-27-0). Instead, one still has to resort to the phenomenological

construction of an EDF and adjust its free parameters, including those associated with tensor terms, to data. How to perform this adjustment in an optimal manner is still an unsettled question. The main problem is to find experimental observables that can unambiguously be related to a mean-field result and are primarily sensitive to a specific part of the EDF, e.g., the tensor part. Such a link between experiment and theory is very often obscured by collective fluctuations around the mean-field states [\[10,11\]](#page-27-0).

The tensor contribution to single-particle energies depends on the filling of shells. It (nearly) vanishes in spin-saturated nuclei, whereas it might be significant when only one level out of two spin-orbit partners is filled. The breaking of spherical shells in deformed nuclei leads to a strong modification of the net spin saturation; hence, the contribution of the tensor force to mean fields and the total energy evolves with deformation. To the best of our knowledge, this was never studied in the published literature. The authors of Ref. [\[8\]](#page-27-0) considered deformation but did not study the impact of tensor terms on deformation properties as such.

In local energy density functionals of the Skyrme type, the tensor force manifests itself through terms that are bilinear in the local pseudotensor spin-current density. In Ref. [\[6\]](#page-27-0), referred to as Article I in what follows, the impact of these terms on the mean-field ground states of spherical nuclei was studied. In the most general case, three bilinear combinations of the pseudotensor spin-current density can be <span id="page-1-0"></span>constructed from a zero-range tensor force, one of them being also generated by a momentum-dependent central zero-range two-body force. All of these terms extrapolate differently when going from spherical to deformed shapes. For the sake of simplicity, we will denote all of the three terms that are bilinear in the spin-current tensor density as *tensor terms* throughout this article, although one of them is already present as soon as a zero-range central velocity-dependent interaction is used as a starting point to derive the EDF.

A word of caution: We cannot expect to significantly improve the agreement with all experimental data, as we learned in our study of spherical nuclei that most deficiencies of the single-particle spectra predicted by standard Skyrme interactions without tensor terms persist in our fits including these terms. Indeed, one of the major results of Article I is that the current form of the *central* Skyrme interaction is not flexible enough to allow for a satisfactory description of single-particle spectra. Adding tensor terms and adjusting them tightly to very specific spectroscopic quantifies often amplifies the deficiencies of the central and spin-orbit parts as currently used, which seem to establish a compromise that averages over the details of shell structure. This is consistent with the recent study of Kortelainen *et al.* [\[12\]](#page-27-0), who point out that it seems impossible to satisfactorily describe single-particle levels of doubly magic nuclei with a standard Skyrme energy density functional including tensor terms, even when relaxing all constraints from bulk properties.

The aim of the present study is to investigate the generic influence of the tensor terms on deformation properties. The questions to be addressed here are as follows:

- (i) How do the tensor terms influence the topography of deformation energy curves of even-even nuclei, given their spherical shell-structure, in particular regarding the number of spin-unsaturated levels at sphericity?
- (ii) How much of these changes is caused by the tensor terms themselves and how much is caused by the rearrangement of all other terms during the fit of the parametrizations?
- (iii) How do the three different tensor terms behave in deformed nuclei depending on the symmetries chosen for the nuclear shapes? For spherically symmetric systems, two of them reduce to the same functional form, whereas the third one vanishes.

The answers to these questions will to a large extent remain independent of remaining deficiencies of the central and spinorbit interactions and will be of great value for the construction of future improved energy functionals. We will address the question of how the surface and surface symmetry energy coefficients change in dependence of the coupling constants of the tensor terms and how this influences the deformation energy at large deformation in a future work.

## **II. THE SKYRME ENERGY FUNCTIONAL WITH TIME-REVERSAL SYMMETRY**

The total energy of a nucleus can be modeled by an energy density functional  $[2,13-15]$  that is the sum of five terms: the uncorrelated kinetic energy, a Skyrme potential energy functional that models the strong interaction in the particle-hole channel, a pairing energy functional, a Coulomb energy functional whose exchange term is treated using the Slater approximation and correction terms that approximately remove the excitation energy due to spurious motions caused by broken symmetries

$$
\mathcal{E} = \mathcal{E}_{\text{kin}} + \mathcal{E}_{\text{Sk}} + \mathcal{E}_{\text{pairing}} + \mathcal{E}_{\text{Coulomb}} + \mathcal{E}_{\text{corr}}.\tag{1}
$$

The Skyrme energy density functional is local and can be decomposed into isoscalar  $t = 0$  and isovector  $t = 1$ contributions of central, spin-orbit and tensor terms

$$
\mathcal{E}_{\text{Sk}} = \int d^3 r \sum_{t=0,1} \mathcal{H}_t^{\text{Sk}}(\mathbf{r})
$$
 (2a)

$$
= \int d^3r \sum_{t=0,1} \left[ \mathcal{H}_t^c(\mathbf{r}) + \mathcal{H}_t^{ls}(\mathbf{r}) + \mathcal{H}_t^l(\mathbf{r}) \right]. \tag{2b}
$$

The physics contained in the Skyrme functional has been discussed in great detail in the literature [\[2,6,15–18\]](#page-27-0). We will repeat here only those aspects that are relevant for the present study.

#### **A. Local densities and currents**

Each of the terms in the Skyrme energy density functional (2a) can be further decomposed into one part that depends on time-even densities only and another part that is bilinear in time-odd densities and currents [\[6,15\]](#page-27-0). We follow the common practice to call them "time-even" and "time-odd" parts of the energy density functional, respectively, although the energy density functional  $\mathcal E$  itself is time-even by construction. Throughout this article, we will assume that time-reversal symmetry is not broken; hence, the densities and currents entering the time-odd part of the energy density functional vanish exactly. This allows one to represent the Skyrme part of the energy density functional through six independent local densities:

$$
\rho_q(\mathbf{r}) = \rho_q(\mathbf{r}, \mathbf{r}')|_{\mathbf{r} = \mathbf{r}'},\tag{3a}
$$

$$
\tau_q(\mathbf{r}) = \nabla \cdot \nabla' \, \rho_q(\mathbf{r}, \mathbf{r}')|_{\mathbf{r} = \mathbf{r}'},\tag{3b}
$$

$$
J_{q,\mu\nu}(\mathbf{r}) = -\frac{i}{2} (\nabla_{\mu} - \nabla'_{\mu}) s_{q,\nu}(\mathbf{r}, \mathbf{r}')|_{\mathbf{r} = \mathbf{r}'}, \qquad (3c)
$$

which are the scalar density  $\rho_q(\mathbf{r})$ , the scalar kinetic density  $\tau_q(\mathbf{r})$ , and the spin-current pseudotensor density  $J_{q,\mu\nu}(\mathbf{r})$  for protons and neutrons  $q = p, n$ . They can be constructed from neutron and proton density matrices expressed in the position basis [\[15,18\]](#page-27-0)

$$
\rho_q(\mathbf{r}\sigma, \mathbf{r}'\sigma') = \langle \hat{a}^{\dagger}_{r'\sigma'q} \hat{a}_{r\sigma q} \rangle
$$
  
=  $\frac{1}{2} \rho_q(\mathbf{r}, \mathbf{r}') \delta_{\sigma\sigma'} + \frac{1}{2} \mathbf{s}_q(\mathbf{r}, \mathbf{r}') \cdot \langle \sigma' | \hat{\sigma} | \sigma \rangle$ , (4)

where

$$
\rho_q(\mathbf{r}, \mathbf{r}') = \sum_{\sigma} \rho_q(\mathbf{r}\sigma, \mathbf{r}'\sigma),
$$
  

$$
\mathbf{s}_q(\mathbf{r}, \mathbf{r}') = \sum_{\sigma\sigma'} \rho_q(\mathbf{r}\sigma, \mathbf{r}'\sigma') \langle \sigma' | \hat{\sigma} | \sigma \rangle.
$$
 (5)

<span id="page-2-0"></span>Proton and neutron densities can be recoupled to isoscalar  $t =$ 0 and isovector  $t = 1$  densities, for example,  $\rho_0(\mathbf{r}) = \rho_n(\mathbf{r}) + \rho_n(\mathbf{r})$  $\rho_p(\mathbf{r})$  and  $\rho_1(\mathbf{r}) = \rho_n(\mathbf{r}) - \rho_p(\mathbf{r})$ , and similarly for  $\tau_t(\mathbf{r})$  and  $J_{t,\mu\nu}(\mathbf{r}).$ 

The Cartesian spin-current pseudotensor density  $J_{\mu\nu}(\mathbf{r})$  can be separated into its pseudoscalar, antisymmetric vector, and symmetric and traceless symmetric pseudotensor parts,

$$
J_{\mu\nu}(\mathbf{r}) = \frac{1}{3}\delta_{\mu\nu}J^{(0)}(\mathbf{r}) + \frac{1}{2}\sum_{\kappa=x}^{z} \epsilon_{\mu\nu\kappa}J_{\kappa}^{(1)}(\mathbf{r}) + J_{\mu\nu}^{(2)}(\mathbf{r}), \qquad (6)
$$

where  $\delta_{\mu\nu}$  is the Kronecker symbol and  $\epsilon_{\mu\nu\kappa}$  the Levi-Civita tensor. The Cartesian components of the pseudoscalar, vector, and traceless pseudotensor parts, expressed in terms of the Cartesian pseudotensor density, are given by

$$
J^{(0)}(\mathbf{r}) = \sum_{\mu=x}^{z} J_{\mu\mu}(\mathbf{r}),
$$
  
\n
$$
J_{\kappa}^{(1)}(\mathbf{r}) = \sum_{\mu,\nu=x}^{z} \epsilon_{\kappa\mu\nu} J_{\mu\nu}(\mathbf{r}),
$$
  
\n
$$
J_{\mu\nu}^{(2)}(\mathbf{r}) = \frac{1}{2} [J_{\mu\nu}(\mathbf{r}) + J_{\nu\mu}(\mathbf{r})] - \frac{1}{3} \delta_{\mu\nu} \sum_{\kappa=x}^{z} J_{\kappa\kappa}(\mathbf{r}).
$$
\n(7)

The radial component of  $\mathbf{J} = \sum_{\kappa} J_{\kappa}^{(1)} \mathbf{e}_{\kappa}$  is the only nonzero contribution when spherical symmetry is imposed. The pseudoscalar  $J^{(0)}(\mathbf{r})$  term still vanishes when rotational symmetry is broken, but parity remains conserved.

### **B. Skyrme's tensor force**

The Skyrme energy functional representing the central, tensor, and spin-orbit interactions can be written in different ways. The most traditional one  $[16,17]$  is to consider the functional as generated by a zero-range two-body effective interaction including a density-dependent term. In his seminal articles [\[19,20\]](#page-27-0), Skyrme introduced two tensor interactions that have not been considered in standard parametrizations so far. An "even" tensor force with the coupling constant *te* mixes relative *S* and *D* waves, whereas an "odd" tensor force with the coupling constant  $t_0$  mixes relative *P* and *F* waves

$$
v^{\mathrm{t}}(\mathbf{r}) = \frac{1}{2}t_e\{[3(\sigma_1 \cdot \mathbf{k'})(\sigma_2 \cdot \mathbf{k'}) - (\sigma_1 \cdot \sigma_2)\mathbf{k'}^2]\delta(\mathbf{r})
$$
  
+  $\delta(\mathbf{r})[3(\sigma_1 \cdot \mathbf{k})(\sigma_2 \cdot \mathbf{k}) - (\sigma_1 \cdot \sigma_2)\mathbf{k}^2]\}$   
+  $t_o[3(\sigma_1 \cdot \mathbf{k'})\delta(\mathbf{r})(\sigma_2 \cdot \mathbf{k}) - (\sigma_1 \cdot \sigma_2)\mathbf{k'} \cdot \delta(\mathbf{r})\mathbf{k}],$  (8)

where we use the shorthand notation  $\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2$  for the relative position vector between the two particles, whereas  $\mathbf{k} = -\frac{i}{2}(\nabla_1 - \nabla_2)$  is the operator for relative momenta acting to the right and **k** its complex conjugate acting to the left. The vectors formed by the Pauli spin matrices are denoted by  $\sigma_1$ and  $\sigma_2$ .

With the symmetry restrictions that we have imposed, only the time-even part of the energy density corresponding to the tensor force (8) differs from zero and is given by

$$
\mathcal{H}_t^{\text{t}}(v^{\text{t}}) = -B_t^T \sum_{\mu,\nu=x}^z J_{t,\mu\nu} J_{t,\mu\nu}
$$

$$
-B_t^F \left[ \frac{1}{2} \left( \sum_{\mu=x}^z J_{t,\mu\mu} \right)^2 + \frac{1}{2} \sum_{\mu,\nu=x}^z J_{t,\mu\nu} J_{t,\nu\mu} \right]. \quad (9)
$$

The labels of the coupling constants  $B_t^T$  and  $B_t^F$  refer to the time-odd terms they multiply in the energy functional when time-reversal invariance is broken, ensuring Galilean invariance [\[6\]](#page-27-0). The notation  $\mathcal{H}_t^t(v^t)$  stresses that Eq. (9) provides the contribution to tensor *terms* coming from the Skyrme zero-range tensor effective *interaction v*<sup>t</sup> given by Eq.  $(8)$ . When starting from the tensor force  $(8)$ , the four coefficients  $B_t^T$  and  $B_t^F$  in Eq. (9) are determined by  $t_e$  and  $t_o$ . As discussed in the next section, the central part of Skyrme's interaction also gives rise to terms proportional to  $J_{t,\mu\nu}J_{t,\mu\nu}$  in the energy density [\[6,15\]](#page-27-0).

The three terms present in Eq. (9) couple the derivatives of the single-particle wave functions and the spin matrices in different ways. In the first term, the derivatives contained in both  $J_{t,\mu\nu}$  are taken along the same direction, as are the two Pauli spin matrices. The two other terms have a structure more typical of what is expected for a tensor interaction that couples a vector in space with a Pauli spin matrix. In the second term they are coupled within a given  $J_{t,\mu\mu}$ , whereas in the third therm they are coupled between the  $J_{t,\mu\nu}$ s. It is the simultaneous presence of these three terms that is the signature of an actual tensor interaction.

#### **C. The Skyrme energy functional**

The complete time-even part of the Skyrme energy density functional is obtained by combining the central, spin-orbit, and tensor contributions

$$
\mathcal{H}_t^{\text{Sk}} = C_t^{\rho} [\rho_0] \rho_t^2 + C_t^{\Delta \rho} \rho_t \Delta \rho_t + C_t^{\tau} \rho_t \tau_t + C_t^{\nabla \cdot J} \rho_t \nabla \cdot \mathbf{J}_t
$$
  

$$
-C_t^T \sum_{\mu,\nu=x}^{\tilde{z}} J_{t,\mu\nu} J_{t,\mu\nu}
$$
  

$$
-C_t^F \left[ \frac{1}{2} \left( \sum_{\mu=x}^{\tilde{z}} J_{t,\mu\mu} \right)^2 + \frac{1}{2} \sum_{\mu,\nu=x}^{\tilde{z}} J_{t,\mu\nu} J_{t,\nu\mu} \right]. \quad (10)
$$

When the energy functional  $(10)$  is generated from a Skyrme interaction, the coupling constants  $C_t$  are the sum of the coupling constants  $A_t$  coming from the central and spin-orbit forces and those of the tensor force  $B_t$  and are defined in the Appendix A of Article I.<sup>1</sup>

One can alternatively consider  $\mathcal{H}^{Sk}_t$  as a functional of local densities in a more general sense and abandon the link to effective interactions. The coefficients  $C_t$  are then fixed independently, except for constraints that must be imposed

<sup>&</sup>lt;sup>1</sup>The expressions for  $B_1^T$ ,  $B_0^F$ ,  $B_0^{\Delta s}$ , and  $B_0^{\nabla s}$ , Eqs. (A3)–(A6) given in the published version of Article I have to be multiplied with  $-1$ , whereas the expressions given in the preprint are correct.

<span id="page-3-0"></span>to fulfill Galilean invariance, cf. Article I. In principle, the twelve constants  $C_t$  can furthermore depend on densities, but in all standard Skyrme parametrizations extensively tested up to now, only  $C_t^{\rho}$  does depend on the isoscalar local density  $\rho_0$ .

### **D. Choice of independent coupling constants in the energy density functional**

The tensor terms are given in Eq.  $(10)$  as a function of the Cartesian representation of the spin-current tensor density. Using the pseudoscalar, vector, and pseudotensor components  $J^{(0)}$ ,  $J^{(1)}$ , and  $J^{(2)}$  introduced in Eq. [\(7\)](#page-2-0), which is more appropriate when spherical symmetry is imposed, one obtains an alternative form

$$
\mathcal{H}_{t}^{t} = -C_{t}^{T} \sum_{\mu,\nu=x}^{z} J_{t,\mu\nu} J_{t,\mu\nu}
$$

$$
-C_{t}^{F} \left[ \frac{1}{2} \left( \sum_{\mu=x}^{z} J_{t,\mu\mu} \right)^{2} + \frac{1}{2} \sum_{\mu,\nu=x}^{z} J_{t,\mu\nu} J_{t,\nu\mu} \right]
$$

$$
= C_{t}^{J0} (J_{t}^{(0)})^{2} + C_{t}^{J1} \mathbf{J}_{t}^{2} + C_{t}^{J2} \sum_{\mu,\nu=x}^{z} J_{t,\mu\nu}^{(2)} J_{t,\mu\nu}^{(2)}.
$$
 (11)

In the last line of Eq.  $(11)$  we have introduced new coupling constants for the terms bilinear in the pseudoscalar, vector, and pseudotensor parts of the spin-current pseudotensor density. Their relation to the coupling constants defined in Eq. [\(10\)](#page-2-0) is given by

$$
C_t^{J0} = -\frac{1}{3}C_t^T - \frac{2}{3}C_t^F,
$$
 (12a)

$$
C_t^{J1} = -\frac{1}{2}C_t^T + \frac{1}{4}C_t^F,\tag{12b}
$$

$$
C_t^{J2} = -C_t^T - \frac{1}{2}C_t^F.
$$
 (12c)

In general, the tensor part of the energy density  $(10)$  depends on four independent parameters. This is most obvious in a Cartesian representation of the tensor functional, where these four parameters are provided by the  $C_t^T$  and  $C_t^F$ . In Article I, however, we have used two different coupling constants to characterize the tensor terms that fulfill

$$
C_0^J = 2C_0^{J1},\tag{13a}
$$

$$
C_1^J = 2C_1^{J1}.
$$
 (13b)

These two coupling constants are sufficient to describe the strength of the isoscalar and isovector tensor terms in static<sup>2</sup> calculations in spherical symmetry as the pseudoscalar and pseudotensor parts of the spin-current tensor density are zero by construction.

When starting from a central and a tensor force, the ratios between the isospin components of the different terms will not be proportional, i.e.,  $C_0^{J_1}/C_1^{J_1} C_0^{J_2}/C_1^{J_2}$ ,  $C_0^{T}/C_1^{T}$ , and  $C_0^F/C_1^F$  will not be equal. The same property is lost also when

one separates the tensor interaction strength between particles of the same and different isospins.

The tensor terms of existing Skyrme parametrizations have been adjusted on spherical nuclei, for which one has timereversal invariance and  $J<sub>t</sub><sup>2</sup>$  is the only nonzero term in Eq. (11). Hence, the values of  $C_t^{J0}$  and  $C_t^{J2}$  have not been fixed by these fits and one has to make additional choices when going beyond sphericity. In the present work, parity is still conserved as a good quantum number such that the only problem is to fix the values of the two constants  $C_t^{j2}$ . The solution to this problem is not unique and a set of reasonable choices is given by:

- (i) One can consider that the tensor terms of the energy functional are generated by the central and tensor parts of a Skyrme force. There is then an univocal relation between  $C_t^{J2}$  and  $t_e$  and  $t_o$  and the balance between the various terms in Eq.  $(11)$  is automatically fixed. Unless otherwise noted, we will use this choice throughout the present article for the parameter sets T*IJ* introduced in Article I. This choice does not permit to set  $C_t^{J1}$  and  $C_t^{J2}$  simultaneously to zero without imposing unrealistic constraints on the central Skyrme interaction. In particular, the parametrization T22 has been constructed in such a way that  $C_t^{J1} = 0$  and that the tensor terms vanish for spherical shapes. The values of  $C_t^{j2}$  are then nonzero and the contribution of the pseudotensor terms does not vanish for deformed shapes.
- (ii) Many authors set to zero the pseudotensor part of the tensor terms (11) for axially deformed Skyrme EDF calculations [\[21,22\]](#page-27-0) although it is *a priori* nonzero. They keep only the vector part **J** that also appears in the spin-orbit part of the EDF. The main motivation for this choice is that the spin-current tensor density in cylindrical coordinates has a complicated form [\[23\]](#page-27-0). This can either be viewed as an approximation or as specific choice of the tensor terms such that  $C_t^{J2} = 0$ .
- (iii) Another possible choice is to set  $C_t^F$  to zero. Together with suitable choices for the time-odd part of the EDF, this allows us to keep the functional form of the standard central Skyrme EDF but with coupling constants of the symmetric tensor terms that are independent of those of a central Skyrme interaction. This choice has been made by the authors of Ref. [\[8\]](#page-27-0).
- (iv) A choice similar to the previous one is to take  $C_t^T$  equal to zero, keeping only the antisymmetric combination of the spin-current tensor density.
- (v) Finally, one can take any ratio of  $C_t^F / C_t^T$  leading to a given  $C_t^{J1}$  value, which interpolates between the two previous choices.

When choosing  $C_t^{J1}$  and  $C_t^{J2}$  to be independent, we have the following interrelations between coupling constants

$$
C_t^{J0} = -C_t^{J1} + \frac{5}{6}C_t^{J2},\tag{14a}
$$

$$
C_t^T = -C_t^{J1} - \frac{1}{2}C_t^{J2},\tag{14b}
$$

$$
C_t^F = 2C_t^{J1} - C_t^{J2}.
$$
 (14c)

The multitude of possible choices for the tensor parametrization opens the risk of an inconsistent use of the coupling

<sup>2</sup>We recall that in the quasi-particle random-phase approximation (QRPA) and other dynamical methods all components of  $J_{\mu\nu}$  might be nonzero in the response transition densities also in spherical symmetry, as do the time-odd densities not addressed here.

<span id="page-4-0"></span>constants of the tensor part of a given Skyrme parametrization. In particular, each choice leads to very different coupling constants in the "time-odd" part of the EDF, which can lead to significant differences.

Eventually, the bilinear part of the functional can be recoupled into terms that contain only densities of the same isospin on the one hand and terms that couple proton and neutron densities on the other hand. Such a representation is often used to characterize the interaction strength in the vector part of the tensor terms through coupling constants  $\alpha$  of the like-particle  $J_t^2$  terms and  $\beta$  of the proton-neutron  $J_t^2$  term

$$
\alpha = C_0^J + C_1^J = 2(C_0^{J1} + C_1^{J1}), \tag{15a}
$$

$$
\beta = C_0^J - C_1^J = 2(C_0^{J1} - C_1^{J1}).
$$
 (15b)

The relation of  $\alpha$  and  $\beta$  to the coupling constants of Skyrme's central and tensor forces can be found in Article I. All other coupling constants of the energy density [\(10\)](#page-2-0) can be recoupled in the same manner, of course.

#### **E. The single-particle Hamiltonian**

The isospin representation of the EDF is very convenient for a discussion of its physical content. The codes that we have developed, however, use a different representation [\[17\]](#page-27-0) that is better suited to construct the mean-fields with the symmetries chosen here. The central and spin-orbit parts of the Skyrme EDF have been described in Ref. [\[17\]](#page-27-0). The additional tensor terms that were not addressed are given by

$$
\mathcal{H}^{t} = b_{14} \sum_{\mu,\nu=x}^{z} J_{0,\mu\nu} J_{0,\mu\nu}
$$
  
+  $b_{16} \left[ \left( \sum_{\mu=x}^{z} J_{0,\mu\mu} \right)^{2} + \sum_{\mu,\nu=x}^{z} J_{0,\mu\nu} J_{0,\nu\mu} \right]$   
+  $\sum_{q=n,p} \left\{ b_{15} \sum_{\mu,\nu=x}^{z} J_{q,\mu\nu} J_{q,\mu\nu}$   
+  $b_{17} \left[ \left( \sum_{\mu=x}^{z} J_{q,\mu\mu} \right)^{2} + \sum_{\mu,\nu=x}^{z} J_{q,\mu\nu} J_{q,\nu\mu} \right] \right\}.$  (16)

The coupling constants of  $(16)$  are related to those of  $(11)$ through

$$
b_{14} = -C_0^T + C_1^T = C_0^{J1} - C_1^{J1} + \frac{1}{2}C_0^{J2} - \frac{1}{2}C_1^{J2}
$$
  
\n
$$
b_{15} = -2C_1^T = 2C_1^{J1} + C_1^{J2}
$$
  
\n
$$
b_{16} = -\frac{1}{2}C_0^F + \frac{1}{2}C_1^F = -C_0^{J1} + C_1^{J1} + \frac{1}{2}C_0^{J2} - \frac{1}{2}C_1^{J2}
$$
  
\n
$$
b_{17} = -C_1^F = -2C_1^{J1} + C_1^{J2}.
$$
\n(17)

The mean-field equations for protons and neutrons, obtained by functional derivative techniques  $[2,15]$  from the energy functional (16), read

$$
\hat{h}_q(\mathbf{r})\psi_i(\mathbf{r}) = \epsilon_i \psi_i(\mathbf{r}),\tag{18}
$$

with the one-body Hamiltonian corresponding to the energy functional  $(10)$  given by<sup>3</sup>

$$
\hat{h}_q(\mathbf{r}) = U_q(\mathbf{r}) - \nabla \cdot B_q(\mathbf{r}) \nabla \n- \frac{i}{2} \sum_{\mu,\nu=x}^z [W_{q,\mu\nu}(\mathbf{r}) \nabla_\mu + \nabla_\mu W_{q,\mu\nu}(\mathbf{r})] \hat{\sigma}_\nu, \qquad (19)
$$

where the  $\hat{\sigma}_{\nu}$  denote the Pauli matrices. The expressions for the single-particle potential  $U(\mathbf{r})$  and the inverse effective mass  $B(r)$  are the same as those given in Ref. [\[17\]](#page-27-0). When the energy functional depends on the vector part of the spin-current tensor only, the second line of Eq. (19) boils down to  $-i\mathbf{W}_q(\mathbf{r}) \cdot \nabla \times$ *σ* with  $W_q(r) = \sum_{\mu \nu \kappa} \epsilon_{\mu \nu \kappa} W_{q, \mu \nu}(r) e_{\kappa}$ , where  $e_{\kappa}$  is the unit vector in  $\kappa$  direction. With the full spin-current tensor, one has to consider

$$
W_{q,\mu\nu}(\mathbf{r}) = \frac{\delta E}{\delta J_{q,\mu\nu}(\mathbf{r})} = -b_9 \sum_{\kappa=x}^z \epsilon_{\kappa\mu\nu} (\nabla_{\kappa} \rho + \nabla_{\kappa} \rho_q) + 2b_{14} J_{\mu\nu} + 2b_{15} J_{q,\mu\nu} + 2b_{16} J_{\nu\mu} + 2b_{17} J_{q,\nu\mu} + 2b_{16} \left[ \sum_{\kappa=x}^z J_{\kappa\kappa} \right] \delta_{\mu\nu} + 2b_{17} \left[ \sum_{\kappa=x}^z J_{q,\kappa\kappa} \right] \delta_{\mu\nu}
$$
(20)

instead. The terms in the first line of Eq. (20) originate from the spin-orbit part of the functional, the other two lines from the tensor part. The terms in the last line of Eq.  $(20)$  might be nonzero only when parity is broken.

We recall that for constrained calculations, as discussed below, the constraints do not contribute to the observable total energy, which is still obtained from  $\mathcal{E}$ , Eq. [\(1\)](#page-1-0), whereas the eigenvalues  $\epsilon_i$  of the mean-field Hamiltonian used to construct the Nilsson diagrams contain a contribution from the constraint [\[24\]](#page-27-0).

### **III. PARAMETERIZATIONS**

For an overview of earlier choices made for the coupling constants of the tensor terms in standard parametrizations of the Skyrme energy functional, we refer to Article I. We will limit ourselves here to recent parametrizations that explore the impact of tensor terms on single-particle spectra.

#### **A. T***IJ* **parametrizations of Lesinski** *et al.*

The main aim of the present article is to test the deformation properties of magic and semimagic nuclei obtained with the parametrizations T*IJ* introduced in Article I. The fit of these parametrizations is based on the same protocol as the one of the SLy*x* parametrizations [\[13,14\]](#page-27-0), with a few minor changes explained in Article I. We have found in Article I that to add a tensor term to a standard Skyrme EDF does not globally

<sup>3</sup>For the standard Skyrme functional [\(11\)](#page-3-0) with non-densitydependent coupling constants of the spin-orbit and tensor terms, the second line can be simplified into  $-i \sum_{\mu\nu} W_{q,\mu\nu}(\mathbf{r}) \nabla_{\mu} \hat{\sigma}_{\nu}$  for the symmetries chosen here.

<span id="page-5-0"></span>

FIG. 1. Coupling constants  $C_0^J$  and  $C_1^J$  of the tensor terms for the parametrizations discussed in the article.

correct its deficiencies for the prediction of masses, radii, or single-particle properties of semimagic nuclei. In fact, very different values of the  $C_t^J$  constants are required for each of these observables to be accurately reproduced, and even these values should vary in different mass regions. Instead of trying to construct a single "best" EDF in Article I, we have studied the impact of the tensor terms on different observables by constructing a set of 36 parametrizations, each corresponding to given values of  $C_0^J$  and  $C_1^J$ , and all other Skyrme parameters being determined by the same fitting procedure. In this way, a wide range of the effective coupling constants  $C_0^J$  and  $C_1^J$  is systematically covered.

In the present study, we limit ourselves to a small subset of these 36 parametrizations, i.e., T22, T26, T44, and T62 in most cases. Figure 1 shows their location in the parameter space of  $C_0^J$  and  $C_1^J$ . The parametrization T22 has by construction vanishing tensor terms at sphericity. It should have properties close to those of SLy4 in which tensor terms have been neglected. The parametrizations T26, T44, and T62 share the same value of the isoscalar tensor coupling constant  $C_0^J = 120 \text{ MeV fm}^5$  and differ by the isovector one  $C_1^J$ , which takes the values 120, 0, and  $-120$  MeV fm<sup>5</sup> respectively.

Parameterizations having the same proton-neutron coupling constant  $\beta$  are aligned along the first diagonal in Fig. 1, those with the same value of  $\alpha$  are aligned along the antidiagonal. For parametrizations T*IJ*, the coefficient of the proton-neutron tensor term increases with the first index *I* for the fixed like-particle tensor term, whereas that of the like-particle tensor term increases with the second index *J* for fixed proton-neutron tensor coupling. Let us recall that the integers *I* and *J* are related to the constants *α* and *β* by

$$
\alpha = 60(J - 2) \text{ MeV fm}^5, \qquad (21a)
$$

$$
\beta = 60(I - 2) \text{ MeV fm}^5. \tag{21b}
$$

#### **B. The parametrization of Colo`** *et al.*

The Skyrme parametrization SLy5 introduced in Ref. [\[14\]](#page-27-0) is one of the two SLy*x* functionals that include the tensor terms generated from the central part of the Skyrme force. Colò *et al.* [\[5\]](#page-27-0) have added a tensor force to it, keeping all the other coupling constants of the parametrization at their original values. We will call this interaction  $SLy5 + T$  ("SLy5 plus tensor") in what follows. The parameters of the tensor force were adjusted in spherical symmetry to single-particle energies along the chains of  $N = 82$  isotones and  $Z = 50$  isotopes. The empirical values for these energies were obtained as separation energies of the last particle in states of an odd-*A* nucleus, assumed to be dominated by one single-particle configuration. They were compared to the eigenvalues of the one-body Hamiltonian, Eq. [\(19\)](#page-4-0), in the neighboring even-even nucleus. The resulting coupling constants in MeV fm<sup>5</sup> are  $C_0^J = -19.333$ and  $C_1^J = -70.466$  or, equivalently,  $\alpha = -89.8$  and  $\beta = 51.9$ . This parametrization has been used in studies of spherical shell structure in Ref. [\[7\]](#page-27-0) and of the Gamov-Teller strength distribution in  $90Zr$  and  $208Pb$  through RPA calculations in Refs.  $[25,26]$ . As can be seen in Fig. 1, SLy5 + T explores a different region of the parameter space than the T*IJ* parametrizations;  $\alpha$  being negative and its modulus larger than *β*.

#### **C. The parametrization of Zalewski** *et al.*

In Ref. [\[8\]](#page-27-0), Zalewski *et al.* did refit the spin-orbit and tensor coupling constants of some standard Skyrme interactions. We will consider here two of their fits that are based on the SLy4 functional [\[14\]](#page-27-0). In a first step, Zalewski *et al.* readjusted  $C_t^{\nabla J}$ and  $C_t^J$  to carefully selected spin-orbit splittings in <sup>40</sup>Ca, 48Ca, and 56Ni, keeping all other coupling constants of the energy functional at their original values. The single-particle energies of a spherical mean-field calculation of doubly magic nuclei have been identified with the separation energy of a nucleon with the same quantum numbers to or from the odd neighboring nuclei. They were compared to experimental separation energies corrected through a macroscopic model taking into account the influence of the coupling of the single-particle state to collective vibrations of the surface. The values of the parameters resulting from this procedure are  $C_0^J = -45$ ,  $C_1^J = -60$ ,  $C_0^{\nabla J} = -60$ , and  $C_1^{\nabla J} = -20$  (all in  $MeV$  fm<sup>5</sup>) and define an EDF called SLy4T. The strength of the standard zero-range spin-orbit force of SLy4T is equal to  $W_0 = 80 \,\text{MeV fm}^5$  and is much lower than in the original SLy4 parametrization, for which  $W_0 = 123$  MeV fm<sup>5</sup>.

Such a modification of  $C_t^J$  and  $C_t^{\nabla J}$  from their original values without changing the other parameters of the EDF degrades prohibitively the masses calculated with SLy4T with respect to those obtained with SLy4. For this reason, Zalewski *et al.* refitted all parameters of SLy4T except  $C_t^J$  and  $C_t^{\nabla J}$ in a second step to restore a reasonable description of bulk properties, leading to the parametrization  $SLy4T_{min}$ .

The parametrizations SLy4T and SLy4T $_{min}$  are explicitly constructed as energy density functionals without making reference to any underlying central, spin-orbit, or tensor force. In particular, the authors chose to set the two coupling constants  $C_t^F$  of the asymmetric Cartesian tensor term to zero and to vary only  $C_t^T$ . This automatically fixes the value of  $C_t^{J2}$  to be equal to two times that of  $C_t^{J1}$  for  $t = 0$  and  $t = 1$ . Although the coupling constants  $C_t^{\nabla J}$  of the spin-orbit term were readjusted, the ratio between the isoscalar and isovector

<span id="page-6-0"></span>TABLE I. Skyrme parametrizations discussed in this work. Procedure: variational (V) corresponds to Skyrme parametrizations where, for given tensor coefficients, all other Skyrme parameters are fitted following the procedure of Ref. [\[14\]](#page-27-0), slightly modified in Article I for the T*IJ* interactions, while perturbative (P) labels parametrizations for which tensor terms are added without refit. Type: the parametrization is treated as an interaction (I) or a functional (F). Central: the contribution to the tensor terms coming from the central part of the interaction is included or not. Tensor: a tensor interaction is included or not.



coupling constants  $C_0^{\nabla J} / C_1^{\nabla J}$  was kept at the value of the original fit.

To analyze the consequences of the choices made in the fitting strategy of Zalewski *et al.* and those of Article I, we performed two additional fits using the same protocol as in Ref. [\[6\]](#page-27-0) but exploring a different region around (−45, −60) MeV fm<sup>5</sup> corresponding to SLy4T in the  $C_0^J$ ,  $C_1^J$  plane of Fig. [1.](#page-5-0) For the first one, called SLy4T<sub>self</sub> hereafter, we fixed  $C_t^J$  and  $C_t^{\nabla J}$ at their SLy4T values but readjusted all the other constants of the functional to obtain the "best" EDF corresponding to our protocol. This parametrization differs from  $SLy4T_{min}$  by the fit protocol and by our choice to keep the interrelations between the coupling constants of the tensor terms as obtained from a two-body central and tensor forces. To study also the impact of the readjustment of the spin-orbit interaction, we constructed a second parametrization, called TZA hereafter, where we additionally vary  $W_0$ , resulting to a value  $W_0 =$ 111*.*934 MeVfm5. This parametrization is thus fitted exactly as the T*IJ* ones, except that it is outside of the rectangular parameter space for  $C_0^J$  and  $C_1^J$  considered in Article I. The coupling constants for SLy4Tself and TZA can be found in the *Physical Review* archive [\[27\]](#page-27-0). The properties of all the interactions that we have used are summarized in Table I.

### **IV. RESULTS**

#### **A. Technical details**

The wave functions are constructed with the code EV8 [\[28,29\]](#page-27-0) that has been modified to include the tensor terms in the energy density functional and the single-particle Hamiltonian. The energies have been recalculated after convergence with a code that uses a more accurate algorithm for the derivatives.

Pairing correlations are treated with the Lipkin-Nogami (LN) method to avoid the breakdown of Bardeen-Cooper-Schrieffer pairing and the resulting discontinuities in the

deformation energy curves. We use an effective densitydependent zero-range pairing interaction with two soft cutoffs at 5 MeV above and below the Fermi energy as described in Ref. [\[30\]](#page-27-0). For consistency with our recent calculations [\[31,32\]](#page-27-0), we chose a strength of  $-1000 \,\text{MeV fm}^{-3}$  for light and medium-heavy nuclei and  $-1250 \,\text{MeV}$  fm<sup>-3</sup> for <sup>186</sup>Pb and 208Pb.

### **B. General comments**

The magic numbers close to stability can be divided into two categories: up to 20, they correspond to a spin-saturated closure of major oscillator shells, whereas above 20, they are created by the spin-orbit interaction that pushes down the level with largest *j* value into the gap between the oscillator shells for 28 or even into the oscillator shell below the gap for 50, 82, and 126. One usually labels a spherical nucleus spin saturated when all pairs of spin-orbit partners are either occupied or empty. In realistic mean-field calculations this will not result in an exact cancellation of the spin-current tensor density, as it should be the case for an exact spin saturation. This has two origins. First, the radial wave functions of the spin-orbit partners are not identical, and, second, pairing correlations will smear out the distribution of occupation numbers. The effect of these two factors will be discussed in a forthcoming publication [\[33\]](#page-27-0). Tensor terms in the time-even part of the energy functional fluctuate with the spin-current tensor density  $J_{\mu\nu}$ , which is small in spin-saturated systems and large whenever only the lower level of a pair of spin-orbit partners is filled, while its partner level remains empty. For an illustration we refer to Article I.

Deformation breaks this simple picture. As soon as it sets in, the spin saturation or nonsaturation disappears and the energy due to tensor terms varies in a way related to the sign of the coupling constants. Close to sphericity, this can be determined by looking at the  $C_t^J$  coefficients. For  $N = Z$  spin saturated

<span id="page-7-0"></span>

FIG. 2. (Color online) Nilsson diagram of the neutrons (top), change of the total contribution from the tensor terms to the total energy relative to the values at the spherical shape (middle), and deformation energy relative to the spherical shape (bottom) for 56Ni obtained with the parametrizations T22, T24, and T26. The energy scale is the same for the two lower panels.

nuclei and parametrizations with  $C_0<sup>J</sup>$  positive, the contribution from the tensor interaction is zero at sphericity and becomes repulsive as soon as deformation sets in. For  $N = Z$  spinunsaturated nuclei, the tensor contribution will be largest at sphericity and decrease with deformation.

## **C. 56Ni**

Let us start our study by looking in detail into  $56Ni$ , the lightest doubly magic nucleus with major proton and neutron shell closures due to the spin-orbit interaction. The values and systematics of static [\[34,35\]](#page-27-0) and transition moments [\[36\]](#page-27-0) of low-lying states around <sup>56</sup>Ni suggest that it is not as good an inert magic core as other doubly magic nuclei. This feature is also observed in shell-model calculations [\[37\]](#page-27-0) and is at the origin of substantial corrections found between "empirical" and "bare" single-particle energies in Ref. [\[38\]](#page-27-0). Several welldeformed rotational bands coexisting with the spherical shellmodel-type states have been observed, one of them down to a  $2^{+}$  level at 5.351 MeV [\[39\]](#page-27-0).

#### *1. Key quantities*

Figure 2 provides three key quantities for the analysis of the deformation properties of  $56$ Ni: the neutron Nilsson diagram,

the contribution of tensor terms to the deformation energy and the total energy, all as a function of axial quadrupole deformation. Three parametrizations have been used, T22, T24, and T26, which differ in the strength of the tensor terms. The proton Nilsson diagram is very similar to the one for neutrons, except for an overall shift due to the Coulomb interaction. The dependence of these quantities on the axial quadrupole deformation is shown as a function of the dimensionless deformation *β*<sup>2</sup> of the mass density distribution defined as

$$
\beta_2 = \sqrt{\frac{5}{16\pi}} \frac{4\pi}{3R^2A} \langle 2z^2 - y^2 - x^2 \rangle, \tag{22}
$$

where  $R = 1.2A^{1/3}$  fm.

The presence of tensor terms in the energy functional has an obvious impact on the single-particle levels. An increase of the tensor interaction results in a reduction of the spin-orbit fields and in a smaller spherical gap at  $N = 28$ . The net result is a sizable decrease of the splitting of the 1*f* levels from T22 to T62. At the same time, tensor terms also modify the slope of the Nilsson levels at small deformations, whereas, at large deformation, the levels predicted by the three parametrizations nearly lie on top of each other.

At sphericity, the tensor contribution for T22 is zero by construction. In practice, one can see that the tensor energy remains close to zero for all deformations. As soon as the nucleus is deformed, the decrease of spin nonsaturation strongly affects the tensor terms. Parameterizations like T24 and T26 have a positive like-particle coupling constant *α* and give repulsive tensor energies at sphericity. This repulsion is decreased by deformation, which reduces the tensor terms by several MeV. For T26, the total energy curve obtained as a function of deformation is softer than without the tensor interaction. In particular, the shoulder at prolate deformations becomes lower in energy and more pronounced. This structure is associated with the rotational band observed down to spin  $2^+$  [\[39\]](#page-27-0). The gain in total deformation energy, however, is much smaller than the gain in deformation energy from the tensor terms.

In the following subsections, we will analyze the origin of these differences and their dependence on the fit strategy.

#### *2. Contributions to the total energy*

Let us first recall that the interactions constructed in Article I differ by their choice of the strengths of the tensor interaction; in addition, all the other terms of the energy functional are also different because each interaction is refitted on the same set of data. It is, therefore, interesting to examine how the different terms of the functional vary from one set to another and how the changes induced by the tensor interaction are, in fact, largely attenuated by a readjustment of the entire functional.

Figure [3](#page-8-0) presents the decomposition of the total binding energy into the contributions from the various terms in the energy functional, Eq. [\(1\)](#page-1-0), for parametrizations T22, T24, and T26. Those panels showing contributions to the Skyrme energy functional are labeled by their content in densities, Eq. [\(10\)](#page-2-0), whereas the other panels provide the kinetic energy (plus the

<span id="page-8-0"></span>

FIG. 3. (Color online) Decomposition of the total energy of <sup>56</sup>Ni obtained with the parametrizations T22, T24, and T26 into the various contributions to the EDF, Eqs.  $(1)$  and  $(10)$ , as a function of the quadrupole deformation  $\beta_2$  (see text).

one-body center-of-mass correction), the Coulomb energy, and the pairing energy (including the Lipkin-Nogami correction).

Unlike in Fig. [2,](#page-7-0) Fig. 3 shows here the absolute values for the total binding energy and the tensor contributions. The binding energy of 56Ni is included in the data the T*IJ* parametrizations are adjusted to and, indeed, the total energy (lower left panel) differs by only a few 100 keV at sphericity. Also, as already pointed out, the deformation energy curves obtained with the three parametrizations differ by less than 2 MeV. These similarities result from a complicated compensation between the various components of the energy. All of them, with the exception of the Coulomb energy, differ on a much larger scale, both in absolute values and deformation dependence.

It can be seen also that each term of the EDF has a very different dependence on deformation and that the total deformation energy also always results from subtle compensations. Both the kinetic energy and the part of the Skyrme functional that contributes to  $E/A$  in infinite homogeneous nuclear matter,  $C^{\rho}[\rho]\rho^2 + C^{\tau}\rho\tau$ , vary by about 70 MeV as a function of deformation. The part of the EDF that does not depend on gradient terms is obtained by summing these two terms and the Coulomb energy and is represented in the panel in the fourth row on the left. It varies with deformation by about 15 MeV. For all parametrizations, the latter curves exhibit pronounced prolate and oblate minima. The gradient term  $C^{\Delta \rho} \rho \Delta \rho$  even amplifies the preference for deformed minima. The combined spin-orbit and tensor terms, shown individually and summed up in the three lower right panels, are the necessary ingredient to obtain a spherical ground state in 56Ni. This underlines the fact that the spin-orbit and tensor terms are not only important for single-particle spectra but also might play a crucial role for the total binding energy, in particular for its deformation dependence. Interestingly, for this nucleus and the parametrizations shown, the contributions of the spin-orbit and tensor terms are of opposite sign and sum up such that their sum is much less dependent on the parametrization than the individual terms.

Figure [4](#page-9-0) provides the same decomposition for various variants of SLy4 adjusted with different strategies, i.e., SLy4, SLy5, SLy4T<sub>min</sub>, and SLy5 + T.

The results obtained with the two functionals SLy4 and SLy5 adjusted in Ref. [\[14\]](#page-27-0) are quite close for all terms, except of course for the tensor contribution, excluded in the case of SLy4 and restricted to the contribution from the central interaction for SLy5. All components of the energy differ slightly because the coupling constants are completely refitted in both cases. The total energies obtained with these two parametrizations, however, are quite close.

In particular, although SLy4 and SLy5 correspond to slightly different coupling constants, they lead to results that differ on a much smaller scale than  $SLy5$  and  $SLy5 + T$ , which differ only by the tensor terms. This underlines the role of self-consistency: the difference in the tensor terms is responsible for changes in the single-particle properties, which ultimately induce changes in each individual contribution to the energy functional.

The similarity between the curves obtained with SLy4 and T22 (Figs. 3 and [4\)](#page-9-0) shows that the slight differences between the interactions have no significant effect.

#### *3. Decomposition of the tensor terms*

The energy contribution from the tensor terms can be decomposed in several ways. We first compare the contributions from the central and tensor parts of the parametrizations T22 and T26 in Fig. [5.](#page-9-0) As explained in Sec. IID, such a decomposition has a meaning only when assuming an underlying force but not for genuine functionals. The central contribution is very similar for both parametrizations (and all others from the T*IJ* family), which is a consequence of its correlation with effective masses and surface tensions

<span id="page-9-0"></span>

FIG. 4. (Color online) Same as Fig. [3,](#page-8-0) but for the parametrizations SLy4, SLy4 $T_{min}$ , SLy5, and SLy5 + T.

through  $t_1$  and  $t_2$  terms of the two-body Skyrme force; see the discussion of Fig. [3](#page-8-0) in Article I. The contribution from central and tensor parts cancel nearly exactly for T22 for all deformations. As exemplified by T26, the contributions from the central and tensor forces to the tensor terms have the same sign for 56Ni for all other T*IJ* parametrizations that have zero or positive values for  $C_0^J$  and  $C_1^J$ .

In Fig. 6, we decompose the tensor energy in the Cartesian and in the angular-momentum coupled representations. Isovector contributions for this  $N = Z$  nucleus are smaller than 20 keV for all deformations. In the left panels, the contributions corresponding to the symmetric and asymmetric terms in the Cartesian representation are plotted. Both are of the order of a few MeV for T22 and of opposite sign to ensure a total contribution close to zero. They are also



FIG. 5. (Color online) Decomposition of the isoscalar tensor energy obtained in 56Ni for the parametrizations T22 and T26 into the contributions from the central and tensor parts of the forces.

of similar magnitude for T26 but repulsive in both cases. Results from T24 are intermediate between those of T22 and T26. The right panels show the contributions to the total tensor energy from the vector and pseudotensor terms in the angular-momentum coupled representation. The contribution from the pseudoscalar term is zero with the symmetries assumed here. Except for T22, where the vector contribution is zero by construction, the pseudotensor contribution is two orders of magnitude smaller than the vector one. We found similar results for all other parametrizations with nonzero  $C_0^{J1}$ and for all nuclei studied here. This justifies the common practice of neglecting the pseudotensor terms for the purpose of calculating binding energies in situations where Galilean invariance is not an issue.

All decompositions of the tensor energy exhibit the same trend: this energy decreases with deformation, without



FIG. 6. (Color online) Decomposition of the isoscalar tensor energy obtained in <sup>56</sup>Ni for the parametrizations T22, T24, and T26 into the contributions from symmetric and asymmetric terms in the Cartesian representation (left), or the contributions from vector and pseudotensor contributions (right). The same energy scale is used for all panels except the one for the pseudotensor contribution  $J^{(2)}J^{(2)}$ .

<span id="page-10-0"></span>TABLE II. Eigenvalues corresponding to the neutron  $pf$  and the  $g_{9/2+}$  orbitals obtained for the single-particle Hamiltonian at spherical shape in  $56$ Ni (see text).

Parametrization	$\epsilon_{f7/2}$	$\epsilon_{f_{5/2}}$	$\Delta \epsilon_f$	$\epsilon_f^{\text{cent}}$	$\epsilon_{p_{3/2}}$	$\epsilon_{p_{1/2}}$	$\Delta \epsilon_p$	$\epsilon_p^{\text{cent}}$	$\Delta \epsilon_f / \Delta \epsilon_p$	$\epsilon_f^{\text{cent}} - \epsilon_p^{\text{cent}}$	$\epsilon_{g_{9/2}}$
Experiment $(\pm S_{1n})$	$-16.64$	$-9.48$	1.02	$-13.57$	$-10.25$	$-9.14$	0.37	$-9.88$	2.75	3.69	
Empirical [38]	$-16.93$	$-9.53$	1.06	$-13.76$	$-10.36$	$-8.48$	0.63	$-9.73$	1.68	4.03	
T <sub>22</sub>	$-16.18$	$-7.70$	1.21	$-12.54$	$-11.21$	$-9.19$	0.67	$-10.54$	1.80	2.00	$-4.26$
T <sub>26</sub>	$-15.47$	$-8.97$	0.93	$-12.68$	$-10.98$	$-9.19$	0.60	$-10.38$	1.56	2.30	$-3.56$
T44	$-15.56$	$-8.69$	0.98	$-12.61$	$-11.09$	$-9.20$	0.63	$-10.46$	1.56	2.15	$-3.75$
T62	$-15.61$	$-8.52$	1.01	$-12.57$	$-11.22$	$-9.25$	0.66	$-10.56$	1.54	2.00	$-3.89$
SL <sub>v</sub> 5	$-16.01$	$-8.03$	1.14	$-12.59$	$-11.11$	$-9.17$	0.65	$-10.47$	1.76	2.12	$-4.05$
$SLy5 + T$	$-16.66$	$-7.09$	1.37	$-12.56$	$-11.15$	$-9.01$	0.72	$-10.44$	1.91	2.12	$-4.67$
SL <sub>v4</sub>	$-16.17$	$-7.80$	1.20	$-12.58$	$-11.13$	$-9.14$	0.66	$-10.47$	1.81	2.11	$-4.20$
SLy4T	$-15.49$	$-8.71$	0.97	$-12.58$	$-11.20$	$-9.47$	0.57	$-10.62$	1.69	1.96	$-3.54$
$SLy4T_{min}$	$-15.63$	$-8.72$	0.99	$-12.67$	$-11.26$	$-9.50$	0.59	$-10.67$	1.68	1.98	$-3.57$
$SLy4T_{self}$	$-15.73$	$-8.59$	1.02	$-12.67$	$-11.29$	$-9.50$	0.60	$-10.70$	1.70	1.97	$-3.71$
TZA	$-16.57$	$-7.08$	1.36	$-12.50$	$-11.33$	$-9.18$	0.72	$-10.61$	1.90	1.89	$-4.65$

exhibiting much structure. This behavior can be understood rather easily in this  $N = Z$  nucleus where the  $f_{7/2}$ - orbitals are filled at sphericity while the *f*5*/*2<sup>−</sup> ones are empty. This situation makes the tensor interaction maximal. As soon as deformation sets in, this simple picture is destroyed: the single-particle levels loose their purity and cross, cf. the upper panel of Fig. [2.](#page-7-0)

#### *4. Single-particle spectra at sphericity*

We give in Table  $II$  the eigenvalues of the single-particle Hamiltonian  $\hat{h}_q$ , Eq. [\(18\)](#page-4-0) for neutron orbitals in the *pf* shell obtained in calculations of 56Ni imposing spherical shape. The position of the  $1g_{9/2^+}$  level is also given, although it is far above the Fermi energy at spherical shape, as it determines the size of the deformed gap at 28 in the Nilsson diagram through its downsloping  $j_z = 1/2$  levels, cf. Fig. [2.](#page-7-0)

We also give the renormalized spin-orbit splittings

$$
\Delta \epsilon_{\ell} = \frac{1}{2\ell + 1} (\epsilon_{j = \ell - 1/2} - \epsilon_{j = \ell + 1/2}), \tag{23}
$$

which for a standard modified oscillator potential would be independent on the quantum numbers of the single-particle states, the centroids

$$
\epsilon_{\ell}^{\text{cent}} = \frac{\ell+1}{2\ell+1} \epsilon_{j=\ell+1/2} + \frac{\ell}{2\ell+1} \epsilon_{j=\ell-1/2},\tag{24}
$$

of spin-orbit partners for the 2*p* and 1*f* levels, the ratio  $\Delta \epsilon_f / \Delta \epsilon_p$  of the spin-orbit splittings of the 1*f* and 2*p* levels, and the distance of their centroids  $\epsilon_f^{\text{cent}} - \epsilon_p^{\text{cent}}$ .

Experimental separation energies from or into low-lying levels in the odd-*A* neighbors of 56Ni that have the characteristics of a single-particle configuration are given in the first line of Table II. Empirical values for the single-particle energies are given in the second line. These quantities are usually compared to the eigenvalues  $\epsilon_i$  of the mean-field Hamiltonian, although many factors make this comparison questionable; see for instance Ref. [\[8\]](#page-27-0) and Sec. [IV B](#page-6-0) of Article I. One source of ambiguity is the coupling of the particle or hole

outside the closed shell to the vibrations of the core. Using the schematic extended unified model, the authors of Ref. [\[38\]](#page-27-0) attempted to remove this effect to determine "bare" values of the single-particle energies by reverse engineering from the low-lying excitation spectra of 56Ni and its odd-*A* neighbors. Although model dependent, we include these values here in the second line of Table  $II$  to have a rough estimate of the order of magnitude of the corresponding corrections. In any event, for an  $N = Z$  nucleus such as <sup>56</sup>Ni, there is also a contribution from the Wigner energy to the separation energies [\[40\]](#page-27-0), which is not considered in Ref. [\[38\]](#page-27-0). Its main effect for a magic nucleus is to render the gap in the separation energies much larger than the gap in the spectrum of eigenvalues of the mean field. Using a schematic model, Chasman [\[40\]](#page-27-0) estimates the correction from the Wigner energy to the size of the  $N = 28$ gap in the empirical spectrum from separation energies to be larger than 2 MeV.

The results for the parametrizations T22, T26, T44, and T62 are given in the next four lines. By construction, there is no tensor contribution at sphericity for T22. The other three interactions share the same isoscalar coupling constant  $C_0^J$  = 120 MeV fm<sup>5</sup> but differ in their isovector one  $C_1^J$ . The presence of a tensor term has a small effect on the absolute position of the 2*p* levels, which move at most by 200 keV, much less than the  $1f$  levels for which the changes go up to  $1.2 \text{ MeV}$ . The tensor term is mainly responsible for a reduction of the spin-orbit splittings, whereas the centroids of the 2*p* and 1*f* levels are affected to a much smaller extent. A change in the centroid position cannot be directly related to the tensor terms because they do not contribute directly to the part of the mean field that governs it. The modification of the centroids is a nonlinear effect induced by the tensor. Although small, the net effect is visible, in particular for the distance between the 1*f* and 2*p* centroids that are pulled into opposite directions.

The shift of the centroids is correlated to the isoscalar tensor coupling constants (cf. T22 and T44), but unexpectedly for a  $N = Z$  nucleus, also slightly to the isovector ones (cf. T26, T44, and T62). The isovector densities and currents induced by the isospin breaking Coulomb interaction are small and do not <span id="page-11-0"></span>significantly contribute to mean fields and energies. The differences between the centroids obtained with T26, T44, and T62 are predominantly a consequence of the readjustment of the entire energy functional for each strength of the tensor terms.

The larger impact of the tensor terms on the 1*f* levels than on the 2*p* ones is still more apparent when a trivial angularmomentum factor in the spin-orbit splitting  $\Delta \epsilon_{\ell}$  is taken out. This result has a geometrical origin discussed in Fig. [16](#page-17-0) of Article I for a different example: a zero-range tensor interaction has the largest impact on spin-orbit splittings for those levels that have the same nodal structure as the ones that dominate the spin-current **J**.

As discussed in Article I and exemplified in Fig. [4,](#page-9-0) the isoscalar tensor term has the tendency to reduce the spin-orbit splitting in spin-unsaturated nuclei for the T*IJ* parametrizations studied here. To maintain a given splitting, the spin-orbit coupling constant has to be increased. Thus, the reduction of the spin-orbit splittings obtained with T44, as compared to those from T22, results from the partial compensation of the change in tensor and spin-orbit contributions. By contrast, the spin-orbit splittings obtained with T26, T44, and T62 are fairly independent on the value of the isovector tensor coupling constant  $C_1^J$ . The reason is twofold: on the one hand, the isovector spin-current tensor density is negligibly small in an  $N = Z$  nucleus and all direct isovector contributions to the spin-orbit field are suppressed. On the other hand, changing  $\overline{C}_1^J$  in the fit does not induce a significant change of the strength of the spin-orbit interaction within the protocol of the T*IJ* interactions.

The values obtained with  $SLy5$  and  $SLy5 + T$  are listed in the next two lines of Table  $II$ . The negative value chosen for  $C_0<sup>J</sup>$  in SLy5 + T is not compensated by a readjustment of the spin-orbit strength and leads to a substantial increase of all spin-orbit splittings. The negligible changes in the position of the centroids gives an indication of the order of magnitude of rearrangement effects from self-consistency.

The results obtained with SLy4, SLy4T, SLy4T<sub>min</sub>,  $SLy4T<sub>self</sub>$ , and TZA can be found in the last five lines of Table [II.](#page-10-0) Although the tensor coupling constants of SLy4T and  $SLy5 + T$  are similar, cf. Fig. [1,](#page-5-0) their behavior with respect to the parametrizations from which they have been constructed is quite different. The spin-orbit splitting of the 1*f* levels obtained with  $SLy5 + T$  is much larger than with  $SLy5$ , while that of SLy4T is smaller than that of SLy4. The good agreement with experiment obtained with SLy4T is not surprising because 56Ni is one of the data that has been used to adjust the spin-orbit and tensor strengths. The origin of the differences between these interactions is the additional reduction of the spin-orbit force in SLy4T to about 2*/*3 of its original value. For 56Ni, the reduced spin-orbit interaction of SLy4T overcompensates the effect of the tensor interaction. The single-particle spectra obtained with SLy4T, SLy4T<sub>min</sub>, and SLy4T<sub>self</sub> differ slightly, which results from self-consistency in the calculations and the readjustment of the other coupling constants of the functional. Because TZA has the same tensor coupling constants as SLy4T, but an increased spin-orbit interaction, it predicts too large spin-orbit splittings.

The single-particle spectra from SLy4 and T22, obtained from almost the same fit protocol, are very close as should

be expected. Also, the single-particle spectrum obtained with SLy4 lies in between those from SLy5 and  $SLy5 + T$ , as could be expected from their tensor coupling constants, Fig. [1,](#page-5-0) and the similarities of the respective fits.

Comparing calculated single-particle energies to empirical ones from Ref. [\[38\]](#page-27-0) and to experimental separation energies, the fine-tuning of the spin-orbit splittings that constitutes the main difference between the interactions studied here does not significantly improve the overall agreement with data. The main deficiency shared by all parametrizations is that the relative distance between the centroids of the 1*f* and 2*p* levels is nearly 2 MeV too small, leading to different sequences of the 1*f*5*/*2<sup>−</sup> and 2*p*1*/*2<sup>−</sup> levels in calculations and data. This result is consistent with the suspicion raised in Ref. [\[41\]](#page-27-0) that a substantial increase of the distance between the centroids given by SLy4 might be needed to reproduce the shape coexistence phenomena around  $74$ Kr. However, a more careful analysis of the physics that connects the single-particle spectra and the observable separation energies is needed before a final conclusion can be drawn.

#### *5. Nilsson diagrams*

At sphericity, the single-particle spectra obtained with interactions adjusted using the same protocol exhibit minor differences, with a splitting of the 1*f* levels varying by about 250 keV. Variations are slightly larger when the tensor term is added perturbatively. The effects of these differences on the dependence of the single-particle levels on deformation can be found in Fig. 7. Two T*IJ* parametrizations only are plotted, as



FIG. 7. (Color online) Neutron Nilsson single-particle diagrams for 56Ni obtained using various families of parametrizations (see text).

<span id="page-12-0"></span>the results do not depend on the isovector coupling constant  $C_1^J$ . Nilsson diagrams for protons differ mainly by a constant shift due to the Coulomb interaction.

The most striking insight from Fig. [7](#page-11-0) is that changing the strength of the isoscalar tensor coupling modifies the slope of the level dependence on deformation, especially close to sphericity. The impact on deformed shell gaps depends, however, on how a tensor term has been introduced.

For refitted parametrizations such as as T22 and T26, the difference between the position of the 1*f*7*/*2<sup>−</sup> levels at sphericity is compensated by the change of the slope of the single-particle levels in such a way that the gap at  $\beta_2 = 0.5$ has about the same size. A similar result is obtained for all T*IJ* and SLy*x* parametrizations.

In contrast, for interactions with perturbatively added or rescaled terms, the size of the deformed gap is strongly modified: it becomes smaller for  $SLy5 + T$ , and larger for  $SLy4T<sub>min</sub>$ . The origin of the latter difference is the reduced spin-orbit interaction for  $SLy4T_{min}$ . With the same tensor coupling constants as  $SLy4T_{min}$ , TZA leads to results similar to those of the T*IJ* parametrizations.

Similar results and similar observations can be made are obtained for all other nuclei discussed hereafter.

#### *6. Deformation energy and its tensor term contribution*

Deformation energy curves are plotted in Fig. 8. To facilitate their correlation with shell structure, the corresponding neutron and proton single-particle spectra at sphericity are given in the upper panels. The total deformation energy is given on the left-hand side in the lower panels, whereas the difference between the tensor energy contribution at a given deformation and at sphericity is provided on the right-hand side. This tensor energy contribution decreases with deformation for TIJ parametrizations and for SLy5, for which  $C_0<sup>J</sup>$  is positive. It increases for the other parametrizations that have a negative  $C_0^J$ .

The differences between the energy curves are not directly linked to the evolution of the tensor energy. They can, in fact, be related to the shell effects that are seen in Fig. [7](#page-11-0) and, in particular, to the relative size of the spherical and prolate deformed  $N = Z = 28$  gaps. A well-defined spherical minimum and a pronounced shoulder at a deformation around  $\beta_2 = 0.5$ are obtained for SLy4, SLy5, and the T*IJ* parametrizations. This structure in the energy curve is at a position qualitatively in agreement with the properties of a superdeformed rotational band observed in  $56$ Ni [\[39\]](#page-27-0).

The combined effect of the reduced spherical  $N = Z = 28$ gap and the large deformed gap obtained with SLy4T has the unphysical consequence that the ground state corresponds to a superdeformed minimum. This result is related to a deficiency of SLy4 on which the tensor interaction has no effect. This interaction, indeed, predicts too small a distance between the centroids of the  $1f$ ,  $2p$  orbitals (see Table [II\)](#page-10-0). In such a case, reproducing the empirical data for the splitting of the 1*f* levels does not guarantee a realistic shell structure. In practice, the gap between the 1*f*7*/*2<sup>−</sup> and 2*p*3*/*2<sup>−</sup> levels becomes too small,



FIG. 8. (Color online) The single-particle spectra at spherical shape are shown in the upper panel for neutrons (left) and protons (right) The deformation energy (left) and the variation of the total tensor energy (right) are plotted on the lower panels for  $56$ Ni and different Skyrme parametrizations, as indicated.

whereas the distance between the 1*f*7*/*2<sup>−</sup> and 1*g*9*/*2<sup>+</sup> levels is now too large.

The parametrization  $SLy5 + T$  has the inverse drawback: the gap at 28 is too large at sphericity, preventing the formation of a secondary gap at large deformation.

The deformation energy and the relative change of the tensor terms obtained with  $SLy4T$  (shown) and with  $SLy4T_{min}$ (not shown) cannot be distinguished within the resolution of Fig. 8. This is less obvious than one might think. The refit that leads from  $SLy4T$  to  $SLy4T_{min}$  changes the absolute binding energy of 56Ni by nearly 6 MeV from −469*.*522 (SLy4T) to −475.480 MeV (SLy4T<sub>min</sub>).

The results obtained with SLy4Tself and TZA confirm the crucial role of the spin-orbit strength. Both interactions have been adjusted with the same protocol and the SLy4T values for the tensor coefficients, but SLy4T<sub>self</sub> shares the same spin-orbit strength as SLy4T while it has been freely varied for TZA. This variation of the spin-orbit leads to results quite close to those of the T*IJ* parametrizations for TZA, in contrast to what is obtained with  $SLy4T_{self}$ .

TABLE III. Tensor coupling constants of T44 and three extensions of T44 beyond the spherical symmetry constructed using the freedom of choice given by an energy functional (see text). All values are in MeV fm<sup>5</sup>. Coupling constants are given for all three representations of the tensor part of the energy functional, Eqs. [\(11\)](#page-3-0) and [\(16\)](#page-4-0). All coupling constants not shown are identical.

	T44	T44 II	T44 III	<b>T44 IV</b>
$C_0^{J0}$	$-20.994$	$-60$	40	$-160$
$C_1^{J0}$	50.027	$\mathbf{0}$	$\overline{0}$	0
$C_0^{J1}$	60	60	60	60
$C_1^{J1}$	$\theta$	$\mathbf{0}$	$\Omega$	$\overline{0}$
$C_0^{J2}$	46.806	$\mathbf{0}$	120	$-120$
$C_1^{J2}$	62.433	$\mathbf{0}$	$\overline{0}$	$\boldsymbol{0}$
$\boldsymbol{C}_{0}^{T}$	$-83.403$	$-60$	$-120$	$\overline{0}$
$\boldsymbol{C}_{1}^{T}$	$-31.216$	$\theta$	0	$\overline{0}$
$\mathcal{C}^F_0$	73.194	120	$\Omega$	240
$\mathcal{C}^F_1$	$-62.433$	$\theta$	0	$\overline{0}$
$b_{14}$	52.187	60	120	0
$b_{15}$	62.433	$\mathbf{0}$	$\overline{0}$	$\Omega$
$b_{16}$	$-67.813$	$-60$	$\Omega$	$-120$
$b_{17}$	62.433	$\boldsymbol{0}$	0	0

### *7. The freedom of using an energy density functional*

As discussed in Sec. [III,](#page-4-0) the fits of tensor couplings have all been performed assuming spherical symmetry. Two coupling constants have been fixed in this way, either *te* and *to* or  $C_t^{J1}$ ,  $t = 0$ , 1. There remains the freedom to choose the  $C_t^{J2}$ coefficients. It has been assumed, for the interactions T*IJ*, that there are underlying two-body central, spin-orbit, and tensor forces, cf. Sec. [II D.](#page-3-0) In this case, there is a one-to-one correspondence between the  $C_t^{J_1}$  coefficients and parameters  $t_1$ ,  $x_1$ ,  $t_2$ ,  $x_2$  of the central Skyrme force and parameters  $t_e$  and  $t<sub>o</sub>$  of the tensor force. With this choice, all coupling constants of the energy functional,  $C_t^{J2}$ ,  $t = 0$ , 1, or, alternatively, the  $C_t^T$  and  $C_t^F$  coupling are also fixed. As explained in Sec. [II D,](#page-3-0) other choices can be made. To explore the impact of doing so, we have constructed three variants of the parametrization T44, listed in Table III. These consist of setting  $C_0^{J2}$  (case II) or  $C_0^F$  (case III) or  $C_0^T$  to zero (case IV). Each choice leads to different values for the coupling constants of the pseudoscalar and pseudotensor part of the tensor terms. We concentrate here on the isoscalar part of the functional, because for the purpose of our study isovector effects are negligible in a  $N = Z$  nucleus.

The total deformation energy and the variation of tensor contributions with respect to their value at sphericity are plotted in Fig. 9 as a function of deformation. The total tensor energies (bottom left), and the decomposition into vector and pseudotensor contributions on the one hand (top), and into symmetric and asymmetric Cartesian components on the other hand (middle), are presented.

As expected from the smallness of the pseudotensor contribution in Fig. [6,](#page-9-0) the choice II where the coefficient  $C_0^{J2}$ is set to zero does not lead to any sizable difference with T44. The results for the two other choices are less obvious. As can be



FIG. 9. (Color online) Deformation energy curve (lower right), tensor energy (lower left), and its decomposition into symmetric, asymmetric, vector, and pseudotensor parts for 56Ni as obtained with the variants of T44 defined in Table III.

seen on the middle panels of Fig. 9, the Cartesian components of the tensor energy are both sizable and one could expect that setting one of the two coefficients of these terms to zero will have a large effect. One can see in the middle panels of Fig. 9 that both terms show significant variation with respect to deformation. The decomposition is even significantly different for T44 and choice II, although leading in both cases to the same total energy. Self-consistency effects are such that for the choices III (where  $J_{q,\mu\nu}J_{q,\nu\mu}$  has a coefficient equal to zero) and IV (where  $J_{q,\mu\nu}J_{q,\mu\nu}$  has a coefficient equal to zero), the remaining contribution has a variation very similar to the variation of the total tensor energy of T44.

The way a parametrization is extended beyond sphericity has a small, but visible, effect on the variation of the total energy with deformation. Setting the coefficient of *Jq,µνJq,µν* equal to zero (option IV) significantly softens the energy curve with an oblate shoulder nearly degenerate with the spherical configuration. The extension of a parametrization beyond sphericity is not a trivial choice and the procedure followed for such extension should always be made transparent. The small differences seen in a simple nucleus such as <sup>56</sup>Ni could become more dramatic in other nuclei.

#### **D. Doubly-magic nuclei**

$$
1. \ ^{40}Ca
$$

The  $N = Z = 20$  nucleus <sup>40</sup>Ca is the heaviest known doubly magic nucleus that exhibits oscillator shell closures. The variation of the total energy, the change of the contribution to the energy of the tensor terms, and the single-particle spectra at spherical shape are plotted in Fig. [10.](#page-14-0) The configuration of  $40$ Ca is spin-saturated at sphericity, and the corresponding tensor energy is due only to small effects as pairing and

<span id="page-14-0"></span>

FIG. 10. (Color online) Same caption as Fig.  $8$  but for  $40$ Ca. Note that the scale of the total deformation energy differs from the one of the tensor terms by a factor two.

nonidentity of the single-particle wave functions of spin-orbit partners. Spin saturation disappears as soon as the nucleus is deformed, but the filling of single-particle states remains identical for protons and neutrons. Thus, the contribution of the tensor terms to the total binding energy induced by deformation is almost purely isoscalar and has the same sign as the isoscalar coupling constant  $C_0^J$ . The tensor term is close to zero at sphericity and increases with deformation. As the coupling constant  $C_0^{J_1}$  of the isoscalar vector part of the tensor terms is the same for T26, T44, and T62,  $C_0^J = 120 \text{ MeV fm}^5$ , these interactions give nearly identical tensor contributions to the total energy.

Despite this feature, the energy curves corresponding to T26, T44, and T62 are not identical. One can see that the softness of the energy curves increases with the strength of the isovector part of the tensor interaction. Once again, these changes result from the readjustment of the coupling constants of all terms of the energy functional when varying  $C_0<sup>J</sup>$  and  $C_1<sup>J</sup>$ to a sample of data that includes both  $N = Z$  and  $N \neq Z$ nuclei. In fact, as for 56Ni discussed in Figs. [3](#page-8-0) and [4,](#page-9-0) most components of the energy show larger variations between the parametrizations than the tensor energy. This example illustrates particularly well the fact that the impact of the tensor



FIG. 11. (Color online) Same as Fig.  $7$  but for  ${}^{40}Ca$ .

terms on deformation energies cannot be foreseen from the sole knowledge of the tensor coupling constants and of the variation of the degree of spin saturation with deformation. The energy of the shoulder that can be seen at a  $\beta_2$  value around 0.5 is significantly lowered for the tensor interactions T44 and T62. This shoulder can be related to the existence of a superdeformed band in  ${}^{40}Ca$  [\[42,43\]](#page-27-0).

The isoscalar tensor coupling is attractive for  $SLy5 + T$ and SLy4T. The energy curve obtained with  $SLy5 + T$  is only slightly different from the one calculated with SLy5 around sphericity but presents a significant lowering of the energy of the shoulder. The situation is quite different for SLy4T. The energy curve varies in opposite direction from the tensor energy and is stiffer than that obtained with SLy4. In particular, the shoulder is obtained at a much larger excitation energy. This clearly is a consequence of the reduced spin-orbit interaction, as this feature of SLy4T is shared by  $SLy4T_{self}$  but not TZA. The energy of the tensor term is larger for the parametrization TZA than for SLy4T, although the tensor coefficients have the same values in both cases. The increase (in absolute value) of the tensor energy is compensated by changes in other terms of the energy functional, in particular a slight reduction of the spin-orbit term, in such a way that the energy curves obtained with SLy4 and TZA are nearly undistinguishable.

The qualitatively different deformation dependence of the tensor energy found for <sup>56</sup>Ni and <sup>40</sup>Ca is accompanied by systematic differences in the single-particle levels given in Fig. 11. As <sup>40</sup>Ca is spin-saturated, the spectrum at sphericity is nearly the same for all parametrizations, except SLy4T. The contributions of the tensor terms are indeed small, although not exactly zero. The largest differences between the results

<span id="page-15-0"></span>obtained with the T*IJ* parametrizations are those for T22 and T26. They are related to the slight difference between their coupling constants and, in particular, to the larger strength of the spin-orbit of T26 compared to T22, cf. Article I. This effect of the spin-orbit interaction on the single-particle levels is more drastic when its strength is explicitly adjusted to spin-orbit splittings in this mass region, as illustrated by the comparison between SLy4T and SLy4.

As for <sup>56</sup>Ni, Fig. [7,](#page-11-0) the tensor interaction affects the slope of the single-particle levels shown in Fig. [11](#page-14-0) as a function of deformation. However, the changes in slopes for a given parametrization are opposite for both nuclei. This is related to the difference in the tensor contribution to the spin-orbit field: it increases with deformation for  ${}^{40}Ca$ , while it decreases in  ${}^{56}Ni$ . The change in slope at least partly compensates the differences found at spherical shape when going to deformed ones. As a consequence, the spectra around the Fermi energy at large deformation are close for refitted parametrizations such as T22, T26, and TZA. On the contrary, the differences between the spectra for a perturbative interaction such as  $SLy5 + T$  and the original one increases with deformation. Differences are larger at all deformations for SLy4T compared to SLy4 and TZA, as a consequence of a perturbative modification of the tensor and the spin-orbit parametrizations.



FIG. 12. (Color online) Same caption as Fig. [8](#page-12-0) but for <sup>48</sup>Ca. Note that the scale of the total deformation energy differs from the one of the tensor terms by a factor two.

## *2.* **<sup>48</sup>***Ca*

The  $Z = 20$ ,  $N = 28$  nucleus <sup>48</sup>Ca is spin saturated in protons and unsaturated in neutrons at sphericity. The variations of the tensor energy and of the total energy with deformation are given in Fig. 12. The upper panels show that the four T*IJ* parametrizations behave very differently in contrast to the case of  $40$ Ca. The tensor energy is nearly independent of deformation for T44, whereas it decreases with deformation for T26 and increases for T62. As confirmed by the behaviors of the  $T = 0$  and  $T = 1$  components of the tensor energy that are plotted in Fig. 13, the isoscalar contribution to the tensor energy does not vary much for all T*IJ* parametrizations, whereas the isovector contribution presents an extremum at sphericity and goes rapidly to zero with deformation. One can relate the behavior of the tensor energy to the fact that the spin-current density is the largest for neutrons at sphericity but nearly zero for protons. With increasing deformation the proton spin-current density grows, whereas the neutron one is reduced, as can be deduced from the decomposition of the tensor terms into their *nn*, *pp*, and *np* contributions provided in Fig. 13. Note for T62 that the *nn* and *pp* contributions are close to zero at all deformations by construction, but this does not result in the proportionality of  $T = 0$  and  $T = 1$  components with respect to each other and to the *np* contribution. In general, the *nn* and *pp* contributions to both  $T = 0$  and  $T = 1$  part are nonzero and deformation dependent. Similarly, from the (by construction) nearly vanishing *np* contribution found at all deformations for T26 it also cannot be concluded that the  $T = 0$  and  $T = 1$  components are proportional to each other and the sum of the *nn* and *pp* contributions.

For 48Ca, the differences between the deformation energy curves in Fig. 12 are clearly correlated with the variations in



FIG. 13. (Color online) Decomposition of the total contribution of the tensor terms to the total energy (lower left) into its isoscalar  $(T = 0)$  and isovector  $(T = 1)$  parts (upper and middle left) and into its *nn*, *pp*, and *np* contributions (right panels) for 48Ca and for the parametrizations T26, T44, and T62.

the isovector tensor energy. The curve is softer for a repulsive isovector contribution to the tensor energy, as it is for T26, and stiffer for an attractive isovector contribution, as it is for T62. This counterintuitive outcome is the consequence of the rapid decrease of the isovector tensor energy from large values at sphericity to very small values with deformation, cf. Fig. [13.](#page-15-0) The results obtained with the interactions  $SLy4T$  and  $SLy5 + T$ are plotted in the lower panels of Fig. [12.](#page-15-0) The isovector tensor term has the same sign for these two interactions as for T62 leading also to stiffer energy curves. However, the magnitude of the effect is smaller, the isovector coupling constants being smaller.

Compared to all other interactions but T62, the SLy4T, SLy4T<sub>min</sub>, and SLy4T<sub>self</sub> functionals give a larger  $Z = 20$ gap at the expense of a reduced  $Z = 28$  one. This is the consequence of the tightly adjusted spin-orbit splittings of the  $1f$  levels in <sup>40</sup>Ca, <sup>48</sup>Ca, and <sup>56</sup>Ni through an attractive tensor interaction in conjunction with a reduced spin-orbit force. The reduced spin-orbit interaction also switches the ordering of the  $1d_{3/2}$ + and  $2s_{1/2}$ + levels below the  $Z = 20$  gap, at variance with empirical data. Keeping the negative tensor coupling constants of SLy4T, but allowing for the readjustment of the spin-orbit strength in TZA, brings the level spacings back to values close to the ones of the original SLy4 interaction.

## *3.* **<sup>68</sup>***Ni*

It is usually assumed from its spectrum that  $^{68}$ Ni is doubly magic [\[44–46\]](#page-27-0). The excitation energy of its first  $2^+$  state is, indeed, large and the  $B(E2; 0^+_{gs} \rightarrow 2^+_1)$  value small [\[47\]](#page-27-0). However, while the  $Z = 28$  proton shell closure is clearly visible in the mass systematics along the  $N = 40$  isotonic chain [\[48\]](#page-28-0), there is no pronounced discontinuity in the masses of Ni isotopes when crossing  $N = 40$  [\[48,49\]](#page-28-0), which hints at a more complex situation. An alternative explanation of the properties of the first  $2^+_1$  state in <sup>68</sup>Ni is based on the impossibility to construct the first  $2^+$  state as a simple neutron 1p-1h excitation. Indeed, the odd-parity *pf* shell is completely filled and the  $1g_{9/2^+}$  orbital empty  $[50,51]$  and at least a 2p-2h excitation is needed to construct a positive-parity state. The interpretation of <sup>68</sup>Ni as a doubly magic nucleus has also been questioned by shell model and QRPA calculations [\[52\]](#page-28-0). Let us mention, finally, that the first excited state of  $^{68}$ Ni is a 0<sup>+</sup> level with a small  $B(E0)$  value to the ground state [\[53\]](#page-28-0), pointing to a possible shape coexistence with weak mixing.

The deformation energy curves and the variation of the tensor energy with deformation are plotted in Fig. 14 for <sup>68</sup>Ni. With  $N = 40$  and  $Z = 28$ , this nucleus is spin-saturated for neutrons and spin-unsaturated for protons.

It is instructive to compare  ${}^{68}$ Ni and  ${}^{48}$ Ca (Fig. [12\)](#page-15-0). Both nuclei have, indeed, similar single-particle configurations. The 1*f*7*/*2<sup>−</sup> subshell is completely filled for protons in 68Ni and for neutrons in 48Ca, and both nuclei are spin saturated for the other type of nucleons. However, the comparison of both nuclei in fact indicates large differences. The variance can be related to two factors. First, the degeneracy of the shells that makes the  $N = 40$  gap for neutrons is much larger than those that make the  $Z = 20$  gap for protons. Second, the



FIG. 14. (Color online) Same caption as Fig. [8](#page-12-0) but for 68Ni.

 $Z = 20$  and  $N = 40$  gaps have different sizes. The latter is not large enough to suppress pairing correlations, such that the spin-saturation is broken and the neutron spin-current density is non-negligible at spherical shape, in particular for the T26 parametrization predicting the smallest  $N = 40$ gap. The main consequence is that for 68Ni all contributions to the tensor energy at spherical shape are nonzero, unless suppressed by their coupling constant, see Fig. [15.](#page-17-0) Also, the *np* component is not just increasing with deformation but fluctuating, most obviously for T62, where it is the only sizable nonzero contribution. By contrast, the rapidly varying *nn* and *pp* contributions for T26 fortuitously add up such that this parametrization presents the smallest variation of the tensor energy among the parametrizations with the same isoscalar coupling constant. The behavior of the tensor energy is reflected in the total energy curves in Fig. 14: it creates an inflexion of the energy curve at  $\beta_2 \approx 0.4$ , which is sufficiently large for T62 to create a secondary minimum.

The tensor energy varies less with deformation for both the SLy5 and  $SLy5 + T$  interactions; therefore, the energy curves obtained with both of these parametrizations and with SLy4 are very similar. The gap at  $N = 40$  obtained with SLy4T is larger, having the size of a major shell closure. It results in a very stiff energy curve. Such a behavior is not directly related to the variation of the tensor energy with deformation, but

<span id="page-17-0"></span>

FIG. 15. (Color online) Same caption as Fig. [13](#page-15-0) but for 68Ni.

to the small spin-orbit strength, as shown by comparing the results of  $SLy4T_{self}$  and TZA, just like for  ${}^{40}Ca$ .

## *4.* **<sup>78</sup>***Ni*

Although it has been observed for the first time more than a decade ago [\[54\]](#page-28-0), not much is yet known about the neutron-rich <sup>78</sup>Ni besides its existence and its  $\beta$ -decay half-life [\[55\]](#page-28-0). The relatively long half-life and the systematics of the separation energies to and from  $N = 50$  isotones down to  $Z = 30$ suggests that the  $N = 50$  shell closure persists for <sup>78</sup>Ni [\[56\]](#page-28-0). In such a case,  $^{78}$ Ni is spin nonsaturated both in protons at  $Z = 28$  and neutrons at  $N = 50$ . The spin-currents from the unsaturated proton 1*g*9*/*2<sup>+</sup> and neutron 1*f*7*/*2<sup>−</sup> levels point into the same direction. Their degeneracies differ by only two, and their nodeless radial wave functions are sufficiently similar that their contributions to the isovector spin-current density nearly cancel. This suggests that the isovector tensor terms cannot play a decisive role in 78Ni, in spite of this nucleus' large isospin asymmetry. The isovector decomposition of the tensor terms for T26, T44, and T62 shown in Fig. 16 indeed confirms that the  $T = 1$  contribution is small for all deformations. The  $T = 0$  contribution dominates and is nearly the same for all three parametrizations such that the different relative weight of the *nn*, *np*, and *pp* contributions does not play a significant role. One can, therefore, expect that its energy curves depend on the tensor parametrization in a way similar to that of  $56$ Ni. Figure [17](#page-18-0) indicates that this is indeed the case for all the interactions that we have studied. In particular, all interactions predict the behavior of a doubly magic nucleus. The TIJ parametrizations with nonzero  $C_0<sup>J</sup>$  values lead to softer energy curves than T22, SLy4, and SLy5, with an inflexion point around  $\beta_2 = 0.3$ , whereas SLy5 + T gives a much stiffer deformation energy curve. The difference between the total deformation energy curves from T26, T44, and T62 is even smaller for  ${}^{78}$ Ni than what was found for  ${}^{56}$ Ni. This suggests



FIG. 16. (Color online) Same caption as Fig. [13](#page-15-0) but for 78Ni.

that the readjustment and self-consistency effects at its origin for  $56\text{Ni}$  are compensated by the asymmetry in  $78\text{Ni}$  just in such a way that the net isospin dependence vanishes for this nucleus.

The parametrizations  $SLy4T$  and  $SLy4T_{self}$  give again results that are qualitatively different from those of the others: they lead to a very pronounced deformed minimum at an excitation energy around 1 MeV, although the spherical gaps at  $N = 50$  and  $Z = 28$  are not smaller than those from T26, for example. This is again a consequence of the reduced contribution of the spin-orbit interaction to the deformation energy for these interactions.

## *5.* **<sup>100</sup>***Sn*

The proton-rich and probably heaviest bound  $N = Z$ doubly magic nucleus  $100Sn$  has been observed more than a decade ago [\[57,58\]](#page-28-0). Up to now, the only spectroscopic information in the direct vicinity of the nucleus is a 172-keV  $\gamma$  ray observed in <sup>101</sup>Sn [\[59\]](#page-28-0). It has been tentatively interpreted as corresponding to the transition between an excited 7*/*2<sup>+</sup> level to the  $5/2^+$  ground state although the order of these two levels is not firmly established. In any case, the distance between these two levels is much lower than the energy difference between the spherical  $1g_{7/2^+}$  and  $2d_{5/2^+}$  orbitals predicted by all Skyrme parametrizations plotted in Fig. [18.](#page-18-0) The tensor interaction has some effect on this spacing: it decreases from more than 1 MeV for T22 down to 600 keV for T26. Of course, the comparison between the energy levels in 101Sn and the single-particle energies supposes that both are pure single-particle configurations that is far from being established. Similar discrepancies with other parametrizations of the self-consistent mean field were reported in Ref. [\[59\]](#page-28-0). From this, however, one cannot safely draw the conclusion that T26 is the most realistic among the parametrizations studied here. The distance between the  $1g_{7/2+}$  and  $2d_{5/2+}$ 

<span id="page-18-0"></span>

FIG. 17. (Color online) Same caption as Fig. [8](#page-12-0) but for 78Ni.

levels depends on the balance between the distance of the centroids of the 1*g* and 2*d* levels as well as on their respective spin-orbit splittings, none of which can be expected to be described well throughout the chart of nuclei for any of the current parametrizations of the Skyrme EDF, see Article I and Ref. [\[12\]](#page-27-0).

The variation of the deformation energy and of the energy contribution of the tensor terms with quadrupole deformation for  $100$ Sn are presented in Fig. 18. For all parametrizations, the results are very similar to those obtained for  $56Ni$ , Fig. [8,](#page-12-0) and 78Ni, Fig. 17. The main difference is that the structure appearing at moderate deformation in the total deformation energy surface is less pronounced and located at higher excitation energies.

## *6.* **<sup>132</sup>***Sn and* **<sup>208</sup>***Pb*

The results obtained for the two heavy doubly magic nuclei <sup>132</sup>Sn and <sup>208</sup>Pb are presented in Figs. [19](#page-19-0) and [20.](#page-19-0) All neutron single-particle spectra exhibit the usual problem of all mean-field interactions that the  $1h_{11/2}$  level in  $132$ Sn is not intruding the *gds* shell [\[2\]](#page-27-0), as suggested by empirical data. The overall behavior of the energy curves below 8 MeV is very



FIG. 1[8](#page-12-0). (Color online) Same caption as Fig.  $8$  but for  $100$  Sn.

similar for most interactions. The stiffness of the deformation energy is marginally modified by the tensor interaction and much less than one might have expected from the variation of the single-particle spectra. The TIJ interactions with  $C_0^J$ coefficients different from zero give slightly softer deformation energy curves than T22 or SLy4. However, the dependence of the relative tensor energy on the value of the isovector coupling constant  $C_1^{J1}$  is very small for all deformations for the same reason as for 78Ni, in spite of the large asymmetry of both nuclei. The reduction of the spin-orbit strength for SLy4T and  $SLy4T_{self}$  leads to a prolate shoulder at about 10 MeV excitation energy. As already found for lighter nuclei, the variation of the tensor energy as a function of deformation can be large, up to  $8 \text{ MeV in }^{132}\text{Sn}$  and  $4 \text{ MeV in }^{208}\text{Pb}$  for the rather small range of deformations covered in Figs. [19](#page-19-0) and [20.](#page-19-0) This significant variation is, to a large extent, absorbed by the rearrangement of the other terms of the Skyrme functional, and it does not affect significantly the total energy curves. The same mechanism that suppresses the isovector tensor terms for <sup>78</sup>Ni is also at play in  $132$  Sn and  $208$  Pb; hence, the variation of the tensor and total energy with deformation of both nuclei is mainly correlated with the isoscalar tensor coupling constant  $C_0^J$ .

<span id="page-19-0"></span>

FIG. 19. (Color online) Same caption as Fig. [8](#page-12-0) but for <sup>132</sup>Sn.

## **E. Selected Zr isotopes**

The  $Zr$ ,  $Z = 40$ , isotopic chain exhibits a rich spectroscopy, the neutron-deficient and neutron-rich isotopes being very deformed and the stable ones being spherical [\[60\]](#page-28-0). Selfconsistent mean-field methods experience large difficulties to reproduce these very rapid variations of shapes in detail [\[21](#page-27-0)[,61,62\]](#page-28-0).

## *1.* **<sup>90</sup>***Zr*

The deformation energy and variation of the tensor contribution with quadrupole deformation for  $90Zr$  are given in Fig. [21.](#page-20-0) There is a very close similarity between the results obtained for  ${}^{90}Zr$  and those for  ${}^{48}Ca$ ; see Fig. [12.](#page-15-0) In both cases, protons are spin-saturated at sphericity, whereas neutrons are nonsaturated as they fully occupy the lowest shell of a pair of spin-orbit partners.

## *2.* **<sup>80</sup>***Zr*

The situation is different for the  $N = Z = 40$  isotope <sup>80</sup>Zr. In spite of its double subshell closure, the sparse available spectroscopic data suggest that  ${}^{80}Zr$  has a large quadrupole



FIG. 20. (Color online) Same caption as Fig. [8](#page-12-0) but for 208Pb.

deformation with a  $\beta_2$  value around 0.4. A rotational band built on the ground state has been observed up to a spin of  $10 \hbar$   $[63–65]$ , although it appears to be slightly distorted at low spin. The large deformation of states in  $80Zr$  is also supported by the observation of strongly coupled rotational bands built on several Nilsson states in adjacent  $^{79}Y$  [\[66\]](#page-28-0) and  $^{81}Zr$  [\[67\]](#page-28-0). In the absence of information on the transition matrix elements at the bottom of the band in  ${}^{80}Zr$ ; however, it is not ruled out that spherical and deformed configurations might coexist in this nucleus and are strongly mixed in the ground state.

The deformation energy curves can be seen in Fig. [22.](#page-20-0) Protons and neutrons are spin-saturated at sphericity. As a consequence, the predicted properties of this nucleus present similarities with those of  ${}^{40}$ Ca: the energy of the tensor terms obtained using the T26, T44, and T62 interactions is very similar as only the isoscalar part of the tensor terms gives a sizable contribution. The total deformation energy, however, does exhibit a weak dependence on the value of  $C_1^{J_1}$ . Also, comparing T22 and the other T*IJ* interactions, the variation of the tensor terms with deformation is in opposite direction to that of the total energy. Both results illustrate the importance of the changes induced in all terms of the functional by the fitting procedure.

<span id="page-20-0"></span>

FIG. 21. (Color online) Same caption as Fig. [8](#page-12-0) but for <sup>90</sup>Zr.

The situation is different for interactions obtained by a perturbative procedure. In this case, the addition of a tensor term to an existing parametrization leads to more drastic changes. This is illustrated by the comparison between the energy curves obtained with  $SLy5$  and  $SLy5 + T$ . For the latter, the deformed minimum is pulled down and becomes degenerate with the spherical configuration. The situation is opposite for SLy4 and SLy4T. The large gap obtained with SLy4T for  $Z = N = 40$  has a dramatic effect on the energy curve, which shows a sharp spherical minimum. The large reduction of the spin-orbit strength is making this nucleus doubly magic and pushes the deformed minimum to a very high energy, although the tensor interaction for SLy4T is more attractive for deformed configurations. This effect is corrected for by TZA. However, none of the parametrizations gives a deformed ground state, a deficiency shared by many modern Skyrme interactions [\[21\]](#page-27-0).

The decomposition of the energy into its central and spin-orbit + tensor components is given in Fig.  $23$  for the interactions T22, T24, and T26, and Fig. [24](#page-21-0) for SLy4,  $SLy4T_{min}$ ,  $SLy5$ , and  $SLy5 + T$ . They confirm the result found for  $56$ Ni, the topography of the energy curves result from subtle cancellations between the bulk contributions and the terms containing gradients. Again, the comparison of





FIG. 22. (Color online) Same caption as Fig. [8](#page-12-0) but for  ${}^{80}Zr$ .

different parametrizations indicates that the readjustment of the parameters counteracts the self-consistency effects. The results obtained using variational interactions are qualitatively very similar. On the contrary, all components of the energy calculated with perturbative interactions differ significantly from those of the original interaction. A major qualitative difference with  $56$ Ni is that the bulk terms give coexisting near-degenerate spherical and deformed minima in <sup>80</sup>Zr and that the compensation between the gradient, spin-orbit, and tensor terms can tip the balance in one or the other direction. A weak spin-orbit strength, such as for the SLy4T interaction, now has the effect of favoring the spherical minimum much too strongly, pushing the deformed minimum very high in energy.

### *3.* **<sup>96</sup>***Zr*

The 96Zr isotope combines spherical subshell closures at  $N = 56$  and  $Z = 40$ . Its low-energy spectrum exhibits several unusual features. The systematics of masses in its immediate vicinity, a first  $2^+$  level with a large excitation energy and one of the smallest  $B(E2)$  values known in heavy nuclei are all consistent with the expectation that  $96Zr$  is a rigid spherical

<span id="page-21-0"></span>

FIG. 2[3](#page-8-0). (Color online) Same caption as Fig. 3 but for <sup>80</sup>Zr.

nucleus. Other observables indicate the fragility of both shells. The *B*(*E*3) value of the  $3^-_1 \rightarrow 0^+_1$  transition is among the strongest known for a vibrational nucleus [\[68,69\]](#page-28-0); the charge radius is enhanced compared to the droplet-model trend [\[70\]](#page-28-0), thereby pointing to substantial ground-state correlations; and the *g* factors of the  $2^+_1$  and  $3^-_1$  hint at a complex superposition of several neutron and proton excitations across subshell closures [\[71\]](#page-28-0). It also shares with  $90Zr$  and  $98Zr$  the rare feature to have a low-lying  $0^+$  state as a first excited state [\[72\]](#page-28-0). As for many of the light doubly magic nuclei studied above, this  $0^+$  state is interpreted as a deformed state resulting from the simultaneous 2p-2h excitation of protons and neutrons across the respective gaps [\[73\]](#page-28-0).

The single-particle spectra at spherical shape and the energy curves of  $96Zr$  are presented in Fig. [25.](#page-22-0) As in  $90Zr$ , the neutrons



FIG. 2[4](#page-9-0). (Color online) Same caption as Fig. 4 but for  ${}^{80}Zr$ ,  $Z =$  $N = 40.$ 

are spin unsaturated and the protons spin saturated. However, in this case, two levels contribute to the neutron spin-current density at sphericity,  $1g_{9/2+}$  and  $2d_{5/2+}$ ; hence, this density is larger than in <sup>90</sup>Zr. Thanks to that, the differences between the interactions are amplified. In particular, the contribution from the isovector tensor terms might become very large, as can be seen in Fig. [26,](#page-22-0) and drastically change the distance and even ordering of the single-particle levels; see Fig. [25.](#page-22-0) Compared to T22, T26, and T44, the parametrization T62 gives much larger  $N = 56$  and  $Z = 40$  gaps and also pushes up the  $2d_{3/2+}$  and 1*g*7*/*2<sup>+</sup> neutron levels, both located above the Fermi energy. The  $SLy5 + T$  interaction has the same tendency but in a less pronounced way. At least one of the tensor coupling constants is negative for both interactions.

<span id="page-22-0"></span>

FIG. 25. (Color online) Same caption as Fig. [8](#page-12-0) but for <sup>96</sup>Zr.



FIG. 26. (Color online) Same caption as Fig. [13](#page-15-0) but for  $96Zr$ .

The deformed minima obtained with the T26 and SLy5 interactions are at variance with data, as is the very stiff energy surface obtained with T62.

## *4.* **<sup>100</sup>***Zr*

A large set of experimental data (charge radii [\[70\]](#page-28-0), rotational bands [\[74\]](#page-28-0), and  $B(E2)$  values [\[75\]](#page-28-0)) demonstrate that 100Zr is located in a region of deformed nuclei. An excited band built on a  $0^+$  state coexisting with the ground-state band [\[74\]](#page-28-0) and the large  $E0$  transition strength between the  $0^+$  states [\[76\]](#page-28-0) indicate the coexistence of shapes with different deformations, the state with the largest deformation being the ground state.

The single-particle spectra at spherical shape, the deformation energy curve and the variation of the tensor energy are plotted against quadrupole deformation in Fig. 27. The overall behavior of the tensor energy shows many similarities with <sup>96</sup>Zr. The results obtained with the T44 and T62 parametrizations indicate, however, a larger contribution from the isovector tensor terms. The four additional neutrons shift



FIG. 27. (Color online) Same caption as Fig. [8](#page-12-0) but for  $100Zr$ ,  $Z = 40, N = 60.$ 

the neutron Fermi energy into a region of large level density above the  $N = 56$  subshell closure. The positions of the  $2d_{3/2^+}$ and 2*g*7*/*2<sup>+</sup> are very much dependent on the sign and size of the isovector coupling constant  $C_1^J$ . For a positive value as in T26, the  $2d_{3/2^+}$  level is close to the Fermi level and is occupied in such a way that it partially cancels the contribution from the  $2d_{5/2+}$  orbital. The isovector tensor terms are in this case strongly reduced. In contrast, for a negative  $C_1^J$ coefficient as in T62, the  $2d_{3/2+}$  level is pushed up and crosses the  $1h_{11/2}$ - level, increasing the neutron spin-current density. Results obtained with the  $SLy5 + T$  interaction, for which  $C_1^J$ is also negative, are similar, although less pronounced. For even larger negative values of the tensor coupling constants, this feedback mechanism will ultimately generate an abnormal level ordering for certain midshell nuclei, cf. Appendix B of Article I. For SLy4T and SLy4T<sub>self</sub>, this feedback mechanism is suppressed by the reduced spin-orbit interaction, whereas for TZA, it is present.

Most total deformation energy curves in Fig. [27](#page-22-0) exhibit spherical, prolate, and oblate minima. The inclusion of beyond mean-field correlations should favor the prolate minima and create a  $0^+$  excitation exhibiting some amount of configuration mixing. Such results are consistent with experiment. The spherical minimum is too much below the deformed one to expect that additional correlations from the projection of  $J = 0$ states will make  $100Zr$  deformed in its ground state. For T62, the deformation energy curve looks like that of a doubly magic nucleus. For SLy4T and SLy4T<sub>self</sub>, it is the reduced spin-orbit interaction that reinforces the proton  $Z = 40$  shell closure. The prolate minimum becomes a shoulder around 5 MeV, leading to the coexistence of spherical and oblate minima.

### *5.* **<sup>110</sup>***Zr*

The only experimental information available about the very neutron-rich  $110Zr$  is that it is a bound nucleus [\[77\]](#page-28-0). It presents the particularity to combine two spin-saturated oscillator shells,  $Z = 40$  and  $N = 70$ . The corresponding gaps are still large for parametrizations like  $SLy4T$  and  $SLy4T<sub>self</sub>$  with a reduced spin-orbit strength and <sup>110</sup>Zr behaves like a doubly magic nucleus, c.f. Fig. 28. For all other parametrizations, but T62, weak subshell closures remain at both these nucleon numbers. The gaps are too small to enforce a rigid spherical shape but sufficient to prevent the existence of a clear-cut unique deformed minimum to describe the ground state. Instead, all interactions but  $SLy4T$ ,  $SLy4T_{self}$ , and T62 predict a complicated pattern of three coexisting spherical, prolate, and oblate structures. For T62, there is no prolate minimum, and the spherical configuration is favored because of the semimagic character of this nucleus with a large  $Z = 40$  shell closure. For  $SLy4T$  and  $SLy4T<sub>self</sub>$ , there is a single, very sharp, spherical minimum typical of a doubly magic character.

#### **F. Heavy semimagic nuclei**

Let us conclude our survey with two selected heavy semimagic nuclei.



FIG. 2[8](#page-12-0). (Color online) Same caption as Fig. 8 but for  $110Zr$ ,  $Z = 40, N = 70.$ 

## *1.* **<sup>120</sup>***Sn*

The stable semimagic  $120$ Sn is the lightest of heavy tin isotopes for which no coexisting deformed rotational band at low excitation energy has been observed [\[78\]](#page-29-0). The neutron number  $N = 70$  of <sup>120</sup>Sn corresponds to a magic number for neutrons in a pure harmonic oscillator picture. This simple picture is destroyed by the spin-orbit interaction that pushes the  $1h_{11/2}$ - across the *N* = 70 gap, creating a shell closure at  $N = 82$ . In fact, data suggest that this oscillator shell does not survive even as a subshell closure, as the empirical 11*/*2<sup>−</sup> intruder level is below the 3*/*2<sup>+</sup> level and degenerate with the  $1/2^+$  state in <sup>132</sup>Sn. As mentioned, it is a well-known problem of virtually all energy functionals that the 1*h*11*/*2<sup>−</sup> intruder level is predicted to lie slightly above the *gds* shell [\[2\]](#page-27-0). This deficiency was related in Fig. [17](#page-18-0) of Article I to a too-high position of the centroid of the 1*h* levels.

The energy surfaces obtained with the T*IJ* interactions are presented on the left-hand side of Fig. [29,](#page-24-0) and the Nilsson diagrams for four selected parametrizations are presented in Fig. [30.](#page-24-0) The neutron contribution to the tensor energy is

<span id="page-24-0"></span>

FIG. 29. (Color online) Same caption as Fig. [8](#page-12-0) but for <sup>120</sup>Sn,  $Z = 50$ ,  $N = 70$ . All panels share the same energy scale.

small at sphericity, as the neutrons are predicted to be spin saturated (see Fig. 30) at variance with experiment. As soon as deformation sets in, the tensor energy increases for the four T*IJ* parametrizations. However, the total energy curves are much closer than one would expect from the difference between the tensor energies. The most significant difference is obtained for deformations between the spherical minimum and the prolate shoulder.

The energy curves calculated with the SLy4, SLy5, and  $SLy5 + T$  interactions are nearly identical, except for a small lowering of the prolate shoulder for the latter.

The situation is quite different for SLy4T. As a consequence of its weak spin-orbit strength, the neutron intruder level is halfway in the gap between the major shells. The energy gap at  $N = 70$  remains very large and significantly reduces the neutron level density around the Fermi energy for deformations up to  $\beta_2$  values around 0.1. The tensor energy decreases with deformation but not sufficiently to compensate the effect of the decrease of the spin-orbit strength. The net effect on the energy curve is that it is much stiffer than with the original SLy4 parameters, artificially making 120Sn a doubly magic nucleus similar to  $132$ Sn.



FIG. 30. (Color online) Proton and neutron Nilsson diagrams for 120Sn obtained with the parametrizations as indicated. Solid lines denote levels of positive parity, dotted lines levels of negative parity, and the red dashed lines denote the Fermi energy.

## *2.* **<sup>186</sup>***Pb*

The heavy, neutron-deficient  $N = 104$ ,  $Z = 82$  Pb isotope, <sup>186</sup>Pb, exhibits a triple shape coexistence of spherical, prolate, and oblate shapes, with the unique feature that its two lowest excited levels are  $0^+$  states [\[79,80\]](#page-29-0). The deformation energy curves obtained with all interactions tested here are plotted in Fig. [31.](#page-25-0) They are compatible with the experimental data and present a spherical minimum and excited oblate, prolate, and often also superdeformed minima, in most cases all separated by small barriers.

The tensor energy and the impact of the tensor terms on the total energy are similar to those found for  $120$ Sn. The differences between the T*IJ* interactions are the largest between the minima or shoulders;  $SLy5 + T$  slightly moves the excitation energies of excited minima compared to SLy5, and the deformation energy curves from  $SLy4T$  and  $SLy4T_{self}$ are stiffer than the others, at least for prolate deformations. The reduced spin-orbit strength for  $SLy4T$  and  $SLy4T_{self}$  pulls the neutron  $1i_{13/2^+}$  intruder back toward the  $N = 126$  gap. This is inconsistent with the existence of a very low-lying isomeric  $13/2$ <sup>+</sup> states located at a few tens of keV excitation energy in surrounding odd-*A* Pb isotopes. *α*-decay hindrance factors

<span id="page-25-0"></span>

FIG. 31. (Color online) Same caption as Fig. [10](#page-14-0) but for 186Pb,  $Z = 82, N = 104.$ 

suggest indeed that this state is well described by a neutron in the  $1i_{13/2^+}$  level coupled to a spherical core [\[81\]](#page-29-0).

Nuclei in this mass region are less affected by the tensor terms, as hinted already in Article I by the analysis of the spincurrent density at spherical shape in the Pb isotopic chain. Still, the tensor terms modify the balance between the excitation energy of the coexisting minima. As the relative position of the minima is sensitive to all terms of the EDF, this quantity cannot be used to safely validate the tensor coupling constants.

#### **V. SUMMARY AND CONCLUSIONS**

We have studied the impact of tensor terms in the Skyrme energy density functional on deformation properties of magic and semimagic nuclei. This work is a continuation of a previous study limited to spherical symmetry, as published in Article I [\[6\]](#page-27-0). The study has been focused on a representative sample of parametrizations introduced in Article I, which covers a wide range of values for the isoscalar and isovector tensor coupling constants and allows us to disentangle their respective role. These parametrizations are adjusted with a fit protocol very similar to that of the successful SLy*x* parametrizations

[\[13,14\]](#page-27-0). We also considered two other recent families of energy functionals also based on the SLy*x* ones but constructed following very different strategies. For the parametrization  $SLy5 + T [5]$  $SLy5 + T [5]$  a tensor force was perturbatively added to  $SLy5$ without any readjustment of the other parameters. For SLy4T [\[8\]](#page-27-0), the tensor and spin-orbit coupling constants were fixed without any readjustment of the other parameters of SLy4. The related parametrizations  $SLy4T_{min}$  [\[8\]](#page-27-0) and  $SLy4T_{self}$  and TZA introduced here allow us to disentangle the origin of the different results obtained with SLy4T and the T*IJ*s as being due to the perturbative fit, the change in the spin-orbit strength, or the choice of tensor coupling constants.

A first result that we have obtained concerns the order of magnitude of the different components of the tensor term. In spherical coordinates, it can be decomposed in vector and pseudotensor contributions. For all studied parametrizations for which the coefficients of both terms have the same order of magnitude, the pseudotensor contribution is at least one order of magnitude lower than the vector one. This justifies the common practice of neglecting the pseudovector contribution to the energy.

The shell effects induced by the tensor interaction fluctuate as a function of the filling of single-particle orbits. This effect has motivated the introduction of the tensor force to explain the evolution of the shell structure of spherical nuclei along isotopic lines. Inevitably, it leads also to a pattern for the size and deformation dependence of the contribution of the tensor terms to the total energy that depends on the fillings of orbitals:

- (i) For doubly spin-saturated nuclei at sphericity, such as  $^{40}$ Ca and  $^{80}Zr$ , the tensor energy is close to zero at spherical shape and increases in absolute value with deformation.
- (ii) For doubly spin-unsaturated doubly magic nuclei such as  ${}^{56}Ni$ ,  ${}^{78}Ni$ ,  ${}^{100}Sn$ ,  ${}^{132}Sn$ , and  ${}^{208}Pb$ , the absolute value of the tensor energy is the largest at sphericity and decreases with deformation.
- (iii) For doubly spin-unsaturated doubly magic  $N = Z$  nuclei like  $56$ Ni and  $100$ Sn, the tensor energy is obviously dominated by the isoscalar part of the tensor interaction. The same conclusion holds, however, also for the  $N \neq$ *Z* nuclei  $^{78}$ Ni,  $^{132}$ Sn, and  $^{208}$ Pb, in spite of their large asymmetry  $N - Z$ . The reason for that is that the proton and neutron spin-currents densities are very similar in size, sign, and spatial distribution in these nuclei; hence, they nearly cancel each others' contribution to the isovector spin-current at all deformations.
- (iv) The isovector tensor contribution to the energy plays a significant role only for doubly magic nuclei that combine a spin-saturated configuration for one nucleon species with a spin-unsaturated configuration for the other, such as in  $48Ca$ ,  $68Ni$ , and  $90Zr$ .
- The behavior of nuclei without large shell or subshell closures for at least one nucleon species does not follow simple rules. These nuclei are most sensitive to the values of the tensor coupling constants, at least within the sample of nuclei studied here. In nuclei with a large density of single-particle levels around the Fermi surface, there are highly nonlinear feedback

effects at play. For large absolute values of their coupling constants, the tensor terms reduce or amplify themselves through the reordering of levels around the Fermi energy, as exemplified by  $96Zr$  and  $100Zr$ .

Self-consistency is implemented at two different levels in the method that we have used: in the fitting procedure of the interaction and in the solution of the mean-field equations.

- (i) The perturbative addition of only a tensor term, like for  $SLy5 + T$ , to an existing parametrization will modify all contributions to the mean fields and the energy.
- (ii) The self-consistency of the mean-field induces a rearrangement of the single-particle wave functions and consequently of all densities affecting at the end all observables. This effect is exemplified by the comparison between the results obtained with SLy5 and  $SLy5 + T$ , which share all coupling constants except those of the tensor terms.
- (iii) Using a protocol mainly based on infinite nuclear matter properties, binding energies and charge radii, as the Saclay-Lyon protocol [\[13,14\]](#page-27-0), the changes in the coupling constants due to the self-consistency of fits tend to counteract the self-consistency in the mean field. This is exemplified by the comparison between the results obtained with SLy4, SLy5, and  $SLy5 + T$ . For most nuclei studied here and for most quantities not directly affected by the tensor terms, the differences between the predictions of the first two are on a much smaller scale than the differences between the latter two. A perturbative modification of a well-adjusted parametrization might spoil its predictive power in unexpected ways. Our results confirm the suspicion of the authors of the perturbatively constructed  $SLy5 + T$ [\[5\]](#page-27-0), who indeed intend their interaction as a tool for explorative studies only and state that "an ambitious refitting program [...] should be [...] undertaken" for more detailed studies.
- (iv) Self-consistency of the fits and/or the calculations has the consequence that the total deformation energy obtained with different interactions varies in most cases on a much smaller scale than the tensor contributions. In some cases such as  ${}^{80}Zr$ , they even might go into opposite directions.
- (v) The tensor and spin-orbit contributions to the total energy and to the spin-orbit fields are tightly interwoven. Constraining both too tightly in a small region of the nuclear chart might be misleading when aiming at a universal functional. This is exemplified by SLy4T with its spin-orbit and tensor coupling constants fitted very carefully to suitably chosen spin-orbit splittings in  ${}^{40}Ca$ ,  $^{48}Ca$ , and  $^{56}Ni$ . The failure of SLy4T to extrapolate well clearly points to missing physics, either in the form of missing terms in the functional or missing correlations.
- (vi) The strong reduction of the spin-orbit strength for SLy4T improves the description of spin-orbit splittings in light nuclei but amplifies the problems from the

wrong positioning of centroids. Also, the spin-orbit splittings in heavy nuclei are much too small. The strong reduction of the spin-orbit strength to about 2*/*3its original value is specific to the SLy4-based interaction constructed in Ref. [\[8\]](#page-27-0). For their SkP and SkO based fits, the reduction is much more moderate.

The size and deformation dependence of the tensor energy is correlated with the impact of the tensor terms on single-particle spectra. In Article I, we analyzed how the tensor terms affect the position and relative distance of single-particle energies for spherical shapes. Concerning the dependence of Nilsson diagrams on the tensor force, the following can be stated:

- (i) Tensor terms modify the slope of the levels in the Nilsson diagram. For the magic nuclei studied here this happens in particular around sphericity, where the tensor contribution to the spin-orbit field  $W_{q,\mu\nu}$  often changes rapidly with deformation.
- (ii) When comparing interactions with different values of the tensor coupling constants that are otherwise completely refitted, the change in slope compensates at large deformations to a large extent the differences between single-particle spectra found at spherical shape. For those interactions, the deformed single-particle spectra around the Fermi energy are often nearly identical in spite of the different tensor interactions. In such fit protocol, the coupling constants of the tensor terms control the balance between spherical and deformed shell gaps.
- (iii) In perturbative fits, in particular those where more than one term is rescaled, deformed shell structure is affected as well.

It can be expected that these finding are to a large extent independent of remaining deficiencies of the central and spinorbit interactions and will be of great value for the construction of future, improved energy functionals. We will address the question of how the surface and surface symmetry energy coefficients of the interactions change as a function of the coupling constants of the tensor terms and how this correlates with energy at large deformation in future work. A study of the "time-odd terms" in the energy functional that originate from a tensor force is underway as well. A point of special interest will be the analysis of potential finite-size instabilities using the technique of Refs. [\[82,83\]](#page-29-0).

## **ACKNOWLEDGMENTS**

This work was supported by the "Interuniversity Attraction Pole" (IUAP) of the Belgian Scientific Policy Office under project P6/23; by the US National Science Foundation under Grant No. PHY-0456903, and by the US Department of Energy under Contract Nos. DE-FG02-96ER40963, DE-FG02-07ER41529 (University of Tennessee), and DE-AC05- 00OR22725 with UT-Battelle, LLC (Oak Ridge National Laboratory).

- <span id="page-27-0"></span>[1] T. Otsuka, T. Suzuki, R. Fujimoto, H. Grawe, and Y. Akaishi, Phys. Rev. Lett. **95**, 232502 (2005).
- [2] M. Bender, P.-H. Heenen, and P.-G. Reinhard, Rev. Mod. Phys. **75**, 121 (2003).
- [3] T. Otsuka, T. Matsuo, and D. Abe, Phys. Rev. Lett. **97**, 162501 (2006).
- [4] B. A. Brown, T. Duguet, T. Otsuka, D. Abe, and T. Suzuki, Phys. Rev. C **74**, 061303(R) (2006).
- [5] G. Colò, H. Sagawa, S. Fracasso, and P. F. Bortignon, Phys. Lett. **B646**, 227 (2007); [Erratum-*ibid.* **B668**, 457 (2008)].
- [6] T. Lesinski, M. Bender, K. Bennaceur, T. Duguet, and J. Meyer, Phys. Rev. C **76**, 014312 (2007).
- [7] W. Zou, G. Colò, Z. Ma, H. Sagawa, and P. F. Bortignon, Phys. Rev. C **77**, 014314 (2008).
- [8] M. Zalewski, J. Dobaczewski, W. Satuła, and T. R. Werner, Phys. Rev. C **77**, 024316 (2008).
- [9] T. Duguet and T. Lesinski, Eur. Phys. J. Special Topics **156**, 207 (2008).
- [10] M. Bender, G. F. Bertsch, and P.-H. Heenen, Phys. Rev. C **73**, 034322 (2006).
- [11] M. Bender, G. F. Bertsch, and P.-H. Heenen, Phys. Rev. C **78**, 054312 (2008).
- [12] M. Kortelainen, J. Dobaczewski, K. Mizuyama, and J. Toivanen, Phys. Rev. C **77**, 064307 (2008).
- [13] E. Chabanat, P. Bonche, P. Haensel, J. Meyer, and R. Schaeffer, Nucl. Phys. **A627**, 710 (1997).
- [14] E. Chabanat, P. Bonche, P. Haensel, J. Meyer, and R. Schaeffer, Nucl. Phys. **A635**, 231 (1998); [Erratum-*ibid.* **A643**, 441 (1998)].
- [15] E. Perlińska, S. G. Rohoziński, J. Dobaczewski, and W. Nazarewicz, Phys. Rev. C **69**, 014316 (2004).
- [16] D. Vautherin and D. M. Brink, Phys. Rev. C **5**, 626 (1972).
- [17] P. Bonche, H. Flocard, and P.-H. Heenen, Nucl. Phys. **A467**, 115 (1987).
- [18] J. Dobaczewski, J. Dudek, S. G. Rohoziński, and T. R. Werner, Phys. Rev. C **62**, 014310 (2000).
- [19] T. H. R. Skyrme, Phil. Mag. **1**, 1043 (1956).
- [20] T. H. R. Skyrme, Nucl. Phys. **9**, 615 (1958).
- [21] P.-G. Reinhard, D. J. Dean, W. Nazarewicz, J. Dobaczewski, J. A. Maruhn, and M. R. Strayer, Phys. Rev. C **60**, 014316 (1999).
- [22] N. Chamel, S. Goriely, and J. M. Pearson, Nucl. Phys. **A812**, 72 (2008).
- [23] M. V. Stoitsov, J. Dobaczewski, W. Nazarewicz, and P. Ring, Comput. Phys. Commun. **167**, 43 (2005).
- [24] S. Åberg, H. Flocard, and W. Nazarewicz, Annu. Rev. Part. Sci. **40**, 439 (1990).
- [25] C. L. Bai, H. Sagawa, H. Q. Zhang, X. Z. Zhang, G. Colò, and F. R. Xu, Phys. Lett. **B675**, 28 (2009).
- [26] C. L. Bai, H. Q. Zhang, X. Z. Zhang, F. R. Xu, H. Sagawa, and G. Colo, Phys. Rev. C ` **79**, 041301(R) (2009).
- [27] See EPAPS document No. E-PRVCAN-80-071911 for the coupling constants of the  $SLy4T<sub>self</sub>$  and TZA parametrizations. For more information on EPAPS, see http://www.aip.org/pubservs/ epaps.html.
- [28] P. Bonche, H. Flocard, P.-H. Heenen, S. J. Krieger, and M. S. Weiss, Nucl. Phys. **A443**, 39 (1985).
- [29] P. Bonche, H. Flocard, and P.-H. Heenen, Comput. Phys. Commun. **171**, 49 (2005).
- [30] C. Rigollet, P. Bonche, H. Flocard, and P.-H. Heenen, Phys. Rev. C **59**, 3120 (1999).
- [31] M. Bender, H. Flocard, and P. H. Heenen, Phys. Rev. C **68**, 044321 (2003).
- [32] M. Bender, P. Bonche, T. Duguet, and P.-H. Heenen, Phys. Rev. C **69**, 064303 (2004).
- [33] M. Bender and P.-H. Heenen (2009) (unpublished).
- [34] J. S. Berryman, K. Minamisono, W. F. Rogers, B. A. Brown, H. L. Crawford, G. F. Grinyer, P. F. Mantica, J. B. Stoker, and I. S. Towner, Phys. Rev. C **79**, 064305 (2009).
- [35] T. E. Cocolios, A. N. Andreyev, B. Bastin, N. Bree, J. Büscher, J. Elseviers, J. Gentens, M. Huyse, Yu. Kudryavtsev, D. Pauwels, T. Sonoda, P. Van den Bergh, and P. Van Duppen, Phys. Rev. Lett. **103**, 102501 (2009).
- [36] J. N. Orce, B. Crider, S. Mukhopadhyay, E. Peters, E. Elhami, M. Scheck, B. Singh, M. T. McEllistrem, and S. W. Yates, Phys. Rev. C **77**, 064301 (2008).
- [37] T. Otsuka, M. Honma, and T. Mizusaki, Phys. Rev. Lett. **81**, 1588 (1998).
- [38] L. Trache, A. Kolomiets, S. Shlomo, K. Heyde, H. Dejbakhsh, C. A. Gagliardi, R. E. Tribble, X. G. Zhou, V. E. Iacob, and A. M. Oros, Phys. Rev. C **54**, 2361 (1996).
- [39] D. Rudolph, C. Baktash, M. J. Brinkman, E. Caurier, D. J. Dean, M. Devlin, J. Dobaczewski, P.-H. Heenen, H.-Q. Jin, D. R. LaFosse, W. Nazarewicz, F. Nowacki, A. Poves, L. L. Riedinger, D. G. Sarantites, W. Satuła, and C.-H. Yu, Phys. Rev. Lett. **82**, 3763 (1999).
- [40] R. R. Chasman, Phys. Rev. Lett. **99**, 082501 (2007).
- [41] M. Bender, P. Bonche, and P.-H. Heenen, Phys. Rev. C **74**, 024312 (2006).
- [42] E. Ideguchi, D. G. Sarantites, W. Reviol, A. V. Afanasjev, M. Devlin, C. Baktash, R. V. F. Janssens, D. Rudolph, A. Axelsson, M. P. Carpenter, A. Galindo Uribarri, D. R. LaFosse, T. Lauritsen, F. Lerma, C. J. Lister, P. Reiter, D. Seweryniak, M. Weiszflog, and J. N. Wilson, Phys. Rev. Lett. **87**, 222501 (2001).
- [43] C. J. Chiara, E. Ideguchi, M. Devlin, D. R. LaFosse, F. Lerma, W. Reviol, S. K. Ryu, D. G. Sarantites, C. Baktash, A. Galindo Uribarri, M. P. Carpenter, R. V. F. Janssens, T. Lauritsen, C. J. Lister, P. Reiter, D. Seweryniak, P. Fallon, A. Görgen, A. O. Macchiavelli, and D. Rudolph, Phys. Rev. C**67**, 041303(R) (2003).
- [44] R. Broda, B. Fornal, W. Królas, T. Pawłat, D. Bazzacco, S. Lunardi, C. Rossi-Alvarez, R. Menegazzo, G. de Angelis, P. Bednarczyk, J. Rico, D. De Acuña, P. J. Daly, R. H. Mayer, M. Sferrazza, H. Grawe, K. H. Maier, and R. Schubart, Phys. Rev. Lett. **74**, 868 (1995).
- [45] T. Ishii, M. Asai, A. Makishima, I. Hossain, M. Ogawa, J. Hasegawa, M. Matsuda, and S. Ichikawa, Phys. Rev. Lett. **84**, 39 (2000).
- [46] O. Sorlin, S. Leenhardt, C. Donzaud, J. Duprat, F. Azaiez, F. Nowacki, H. Grawe, Z. Dombradi, F. Amorini, A. Astier, ´ D. Baiborodin, M. Belleguic, C. Borcea, C. Bourgeois, D. M. Cullen, Z. Dlouhy, E. Dragulescu, M. Górska, S. Grévy, D. Guillemaud-Mueller, G. Hagemann, B. Herskind, J. Kiener, R. Lemmon, M. Lewitowicz, S. M. Lukyanov, P. Mayet, F. de Oliveira Santos, D. Pantalica, Yu.-E. Penionzhkevich, F. Pougheon, A. Poves, N. Redon, M. G. Saint-Laurent, J. A. Scarpaci, G. Sletten, M. Stanoiu, O. Tarasov, and Ch. Theisen, Phys. Rev. Lett. **88**, 092501 (2002).
- [47] N. Bree, I. Stefanescu, P. A. Butler, J. Cederkäll, T. Davinson, P. Delahaye, J. Eberth, D. Fedorov, V. N. Fedosseev, L. M. Fraile, S. Franchoo, G. Georgiev, K. Gladnishki, M. Huyse,

<span id="page-28-0"></span>O. Ivanov, J. Iwanicki, J. Jolie, U. Köster, Th. Kröll, R. Krücken, B. A. Marsh, O. Niedermaier, P. Reiter, H. Scheit, D. Schwalm, T. Sieber, J. Van de Walle, P. Van Duppen, N. Warr, D. Weisshaar, F.Wenander, and S. Zemlyanoy, Phys. Rev. C**78**, 047301 (2008).

- [48] S. Rahaman, J. Hakala, V. Elomaa, T. Eronen, U. Hager, A. Jokinen, A. Kankainen, I. Moore, H. Penttila, S. Rinta-Antila, ¨ J. Rissanen, A. Saastamoinen, C. Weber, and J. Äystö, Eur. Phys. J. A **34**, 5 (2007).
- [49] C. Guénaut, G. Audi, D. Beck, K. Blaum, G. Bollen, P. Delahaye, F. Herfurth, A. Kellerbauer, H.-J. Kluge, J. Libert, D. Lunney, S. Schwarz, L. Schweikhard, and C. Yazidjian, Phys. Rev. C **75**, 044303 (2007).
- [50] H. Grawe, M. Górska, J. Döring, C. Fahlander, M. Palacz, F. Nowacki, E. Caurier, J. M. Daugas, M. Lewitowicz, M. Sawicka, M. Pfützner, R. Grzywacz, K. Rykaczewski, O. Sorlin, S. Leenhardt, F. Azaiez, M. Rejmund, K. Hauschild, and J. Uusitalo, in *Tours Symposium on Nuclear Physics IV*, edited by M. Lewitowicz, Y. T. Oganessian, H. Akimune, M. Ohta, H. Utsunomiya, T. Yamagata, M. Arnould, and T. Wada (AIP, 2001), Vol. 561 of *AIP Conf. Proc.*, pp. 287–300.
- [51] H. Grawe and M. Lewitowicz, Nucl. Phys. **A693**, 116 (2001).
- [52] K. Langanke, J. Terasaki, F. Nowacki, D. J. Dean, and W. Nazarewicz, Phys. Rev. C **67**, 044314 (2003).
- [53] T. Kibédi and R. H. Spear, At. Data Nucl. Data Tables 89, 77 (2005).
- [54] C. Engelmann, F. Ameil, P. Armbruster, M. Bernas, S. Czajkowski, P. Dessagne, C. Donzaud, H. Geissel, A. Heinz, Z. Janas, C. Kozhuharov, Ch. Miehé, G. Münzenberg, M. Pfützner, C. Röhl, A. Schwab, C. Stéphan, K. Sümmerer, L. Tassan-Got, and B. Voss, Z. Phys. A **352**, 351 (1995).
- [55] P. T. Hosmer, H. Schatz, A. Aprahamian, O. Arndt, R. R. C. Clement, A. Estrade, K.-L. Kratz, S. N. Liddick, P. F. Mantica, W. F. Mueller, F. Montes, A. C. Morton, M. Ouellette, E. Pellegrini, B. Pfeiffer, P. Reeder, P. Santi, M. Steiner, A. Stolz, B. E. Tomlin, W. B. Walters, and A. Wöhr, Phys. Rev. Lett. 94, 112501 (2005).
- [56] J. Hakala, S. Rahaman, V.-V. Elomaa, T. Eronen, U. Hager, A. Jokinen, A. Kankainen, I. D. Moore, H. Penttila, S. Rinta- ¨ Antila, J. Rissanen, A. Saastamoinen, T. Sonoda, C. Weber, and J. Äystö, Phys. Rev. Lett. **101**, 052502 (2008).
- [57] R. Schneider, J. Friese, J. Reinhold, K. Zeitelhack, T. Faestermann, R. Gernhäuser, H. Gilg, F. Heine, J. Homolka, P. Kienle, H. J. Körner, H. Geissel, G. Münzenberg, and K. Sümmerer, Z. Phys. A **348**, 241 (1994).
- [58] M. Lewitowicz, R. Anne, G. Auger, D. Bazin, C. Borcea, V. Borrel, J. M. Corre, T. Dörfler, A. Fomichov, R. Grzywacz, D. Guillemaud-Mueller, R. Hue, M. Huyse, Z. Janas, H. Keller, S. Lukyanov, A. C. Mueller, Yu. Penionzhkevich, M. Pfützner, F. Pougheon, K. Rykaczewski, M. G. Saint-Laurent, K. Schmidt, W. D. Schmidt-Ott, O. Sorlin, J. Szerypo, O. Tarasov, J. Wauters, and J. Zylicz, Phys. Lett. **B332**, 20 (1994).
- [59] D. Seweryniak, M. P. Carpenter, S. Gros, A. A. Hecht, N. Hoteling, R. V. F. Janssens, T. L. Khoo, T. Lauritsen, C. J. Lister, G. Lotay, D. Peterson, A. P. Robinson, W. B. Walters, X. Wang, P. J. Woods, and S. Zhu, Phys. Rev. Lett. **99**, 022504 (2007).
- [60] J. Eberth, R. A. Meyer, and K. Sistemich, eds., in *Proceedings of the International Workshop on Nuclear Structure of the Zirconium Region, Bad Honnef, 1988* (Springer, Berlin/Heidelberg, 1988).
- [61] F. Tondeur, Nucl. Phys. **A359**, 278 (1981).
- [62] J. Skalski, P. H. Heenen, and P. Bonche, Nucl. Phys. **A559**, 221 (1993).
- [63] C. J. Lister, M. Campbell, A. A. Chishti, W. Gelletly, L. Goettig, R. Moscrop, B. J. Varley, A. N. James, T. Morrison, H. G. Price, J. Simpson, K. Connel, and O. Skeppstedt, Phys. Rev. Lett. **59**, 1270 (1987).
- [64] C. J. Lister, P. J. Ennis, A. A. Chishti, B. J. Varley, W. Gelletly, H. G. Price, and A. N. James, Phys. Rev. C **42**, R1191 (1990).
- [65] S. M. Fischer, C. J. Lister, D. P. Balamuth, R. Bauer, J. A. Becker, L. A. Bernstein, M. P. Carpenter, J. Durell, N. Fotiades, S. J. Freeman, P. E. Garrett, P. A. Hausladen, R. V. F. Janssens, D. Jenkins, M. Leddy, J. Ressler, J. Schwartz, D. Svelnys, D. G. Sarantites, D. Seweryniak, B. J. Varley, and R. Wyss, Phys. Rev. Lett. **87**, 132501 (2001).
- [66] S. D. Paul, C. Baktash, W. Satuła, C. J. Gross, I. Birriel, R. M. Clark, R. A. Cunningham, M. Devlin, P. Fallon, A. Galindo Uribarri, T. Ginter, D. R. Lafosse, J. Kay, F. Lerma, I. Y. Lee, C. Leyland, A. O. Macchiavelli, B. D. MacDonald, S. J. Metcalfe, A. Piechaczek, D. C. Radford, W. Reviol, L. L. Riedinger, D. Rudolph, K. Rykaczewski, D. G. Sarantites, J. X. Saladin, D. Shapira, G. N. Sylvan, S. L. Tabor, K. S. Toth, W. Weintraub, D. F. Winchell, V. Q. Wood, R. Wyss, and C. H. Yu, Phys. Rev. C **58**, R3037 (1998).
- [67] N. Mărginean, D. Bucurescu, C. Rossi Alvarez, C. A. Ur, Y. Sun, D. Bazzacco, S. Lunardi, G. de Angelis, M. Axiotis, E. Farnea, A. Gadea, M. Ionescu-Bujor, A. Iordăchescu, W. Krolas, Th. Kröll, S. M. Lenzi, T. Martinez, R. Menegazzo, D. R. Napoli, P. Pavan, Zs. Podolyak, M. De Poli, B. Quintana, and P. Spolaore, Phys. Rev. C **69**, 054301 (2004).
- [68] H. Mach, S. Ćwiok, W. Nazarewicz, B. Fogelberg, M. Moszyński, J. Winger, and R. L. Gill, Phys. Rev. C 42, R811 (1990).
- [69] D. J. Horen, R. L. Auble, G. R. Satchler, J. R. Beene, I. Y. Lee, C. Y. Wu, D. Cline, M. Devlin, R. Ibbotson, and M. W. Simon, Phys. Rev. C **48**, R2131 (1993).
- [70] P. Campbell, H. L. Thayer, J. Billowes, P. Dendooven, K. T. Flanagan, D. H. Forest, J. A. R. Griffith, J. Huikari, A. Jokinen, R. Moore, A. Nieminen, G. Tungate, S. Zemlyanoi, and J. Äystö, Phys. Rev. Lett. **89**, 082501 (2002).
- [71] G. Kumbartzki, N. Benczer-Koller, J. Holden, G. Jakob, T. J. Mertzimekis, M. J. Taylor, K. H. Speidel, R. Ernst, A. E. Stuchbery, C. W. Beausang, and R. Krücken, Phys. Lett. **B562**, 193 (2003).
- [72] H. Mach, G. Molnár, S. W. Yates, R. L. Gill, A. Aprahamian, and R. A. Meyer, Phys. Rev. C **37**, 254 (1988).
- [73] K. Heyde, E. D. Kirchuk, and P. Federman, Phys. Rev. C **38**, 984 (1988).
- [74] J. K. Hwang, A. V. Ramayya, J. H. Hamilton, J. O. Rasmussen, Y. X. Luo, D. Fong, K. Li, C. Goodin, S. J. Zhu, S. C. Wu, M. A. Stoyer, R. Donangelo, X.-R. Zhu, and H. Sagawa, Phys. Rev. C **74**, 017303 (2006).
- [75] J. K. Hwang, A. V. Ramayya, J. H. Hamilton, Y. X. Luo, A. V. Daniel, G. M. Ter-Akopian, J. D. Cole, and S. J. Zhu, Phys. Rev. C **73**, 044316 (2006).
- [76] H. Mach, M. Moszyński, R. L. Gill, G. Molnár, F. K. Wohn, J. A. Winger, and J. C. Hill, Phys. Rev. C **41**, 350 (1990).
- [77] M. Bernas, C. Engelmann, P. Armbruster, S. Czajkowski, F. Ameil, C. Böckstiegel, P. Dessagne, C. Donzaud, H. Geissel,

<span id="page-29-0"></span>A. Heinz, Z. Janas, C. Kozhuharov, Ch. Miehé, G. Münzenberg, M. Pfützner, W. Schwab, C. Stéphan, K. Sümmerer, L. Tassan-Got, and B. Voss, Phys. Lett. **B415**, 111 (1997).

- [78] R. Julin, Physica Scripta **T56**, 151 (1995).
- [79] A. N. Andreyev, M. Huyse, P. Van Duppen, L. Weissman, D. Ackermann, J. Gerl, F. P. Heßberger, S. Hofmann, A. Kleinbohl, G. Münzenberg, S. Reshitko, C. Schlegel, H. Schaffner, P. Cagarda, M. Matos, S. Saro, A. Keenan, C. Moore, C. D. O'Leary, R. D. Page, M. Taylor, H. Kettunen, M. Leino, A. Lavrentiev, R. Wyss, and K. Heyde, Nature **405**, 430 (2000).
- [80] A. N. Andreyev, M. Huyse, P. V. Duppen, L. Weissman, D. Ackermann, J. Gerl, F. P. Heßberger, S. Hofmann,

A. Kleinbohl, G. Münzenberg, S. Reschitko, C. Schlegel, H. Schaffner, P. Cagarda, M. Matos, S. Saro, A. Keenan, C. J. Moore, C. D. O'Leary, R. D. Page, M. J. Taylor, H. Kettunen, M. Leino, A. Lavrentiev, R. Wyss, and K. Heyde, Nucl. Phys. **A682**, 482c (2001).

- [81] A. N. Andreyev, M. Huyse, P. Van Duppen, J. F. C. Cocks, K. Helariutta, H. Kettunen, P. Kuusiniemi, M. Leino, W. H. Trzaska, K. Eskola, and R. Wyss, Phys. Rev. Lett. **82**, 1819 (1999).
- [82] T. Lesinski, K. Bennaceur, T. Duguet, and J. Meyer, Phys. Rev. C **74**, 044315 (2006).
- [83] D. Davesne, M. Martini, K. Bennaceur, and J. Meyer, Phys. Rev. C **80**, 024314 (2009).