Poincaré invariant three-body scattering at intermediate energies

T. Lin,[1] Ch. Elster,[1,2] W. N. Polyzou,[3] H. Witała,[4] and W. Glöckle[5]
[1]*Institute of Nuclear and Particle Physics, and Department of Physics and Astronomy, Ohio University, Athens, Ohio 45701, USA*
[2]*Physics Division, Argonne National Laboratory, Argonne, Illinois 60439, USA*
[3]*Department of Physics and Astronomy, University of Iowa, Iowa City, Iowa 52242, USA*
[4]*Institute of Physics, Jagiellonian University, PL-30059 Cracow, Poland*
[5]*Institute for Theoretical Physics II, Ruhr-University Bochum, D-44780 Bochum, Germany*
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The Faddeev equation for three-nucleon scattering, based on an exactly Poincaré invariant formulation of quantum mechanics, is solved for projectile energies up to 2 GeV. As in the nonrelativistic three-body problem, the three-body dynamics is determined, up to three-body interactions, by the two-body dynamics and cluster properties. The two-body interactions are determined, up to a unitary scattering equivalence, by two-body scattering data, which in our application are generated by a nonrelativistic Malfliet-Tjon interaction. The Faddeev equation is directly solved in a kinematic momentum representation without employing a partial-wave decomposition. The solution of the Faddeev equation is generated using Padé summation, and the numerical feasibility and stability of the solution is demonstrated. Scattering observables for elastic and breakup scattering are calculated for projectile energies in the intermediate energy range up to 2 GeV, and compared with their nonrelativistic counterparts. The convergence of the multiple scattering series is investigated as a function of the projectile energy in different scattering observables and configurations. The complementary roles of kinematic and dynamical contributions to our Poincaré invariant model are investigated. Approximations to the two-body interaction embedded in the three-particle space are compared with the exact treatment.

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I. INTRODUCTION

The lightest nuclei can be accurately modeled as systems of nucleons interacting via effective two- and three-body forces motivated, e.g., by meson exchange. This picture is expected to break down at a higher energy scale, where the physics is more efficiently described in terms of subnuclear degrees of freedom. Few-body methods are an essential tool for determining model Hamiltonians that describe low-energy nuclear physics. Few-body methods have the potential to be equally as useful for modeling light nuclei as few-degree of freedom systems at intermediate energies. To successfully do this, a number of challenges need to be addressed. These include replacing the nonrelativistic theory with a relativistic one and overcoming limitations imposed by interactions fit to elastic nucleon-nucleon (NN) scattering data, including new degrees of freedom that appear above the pion production threshold, as well as solving numerical problems related to the proliferation of partial waves characteristic of scattering calculations at higher energies. Thus the intermediate energy regime is a new territory for few-body calculations that waits to be explored.

In this paper we address two of these challenges. We demonstrate that it is now possible to perform converged three-body scattering calculations at energies up to 2 GeV laboratory kinetic energy. Key elements are a consistent implementation of a Poincaré symmetric quantum theory [1] and the use of direct integration methods that avoid the partial-wave decomposition, successfully applied below the pion-production threshold [2]. In a series of publications [3–5], the technique for solving the nonrelativistic momentum-space Faddeev equation without partial waves has been mastered for bound as well as scattering states. For projectile energies of a GeV or more, the partial-wave expansions, which are efficient at low energies, become increasingly more complicated as the energy is increased. Direct integration methods have clear advantages at intermediate energies.

The framework used in this paper is one of two approaches that have been successfully applied to model realistic relativistic few-body systems. Field theory motivated approaches are based on few-body extensions of the Bethe-Salpeter equation. The structure of the equations is dictated by quantum field theory, while the input uses a combination phenomenology and theoretically motivated assumptions. Quasipotential reductions of these equations simplify the numerical problem without making any compromises on the physics input. The first exact realistic treatment of the relativistic three-nucleon bound state problem by Stadler and Gross [6] was based on the Gross equation, which is a quasipotential equation that puts two of the three nucleons on their mass shells. The alternative treatment of relativistic few-body dynamics used in this paper builds on the success of nonrelativistic few-nucleon physics. The theoretical foundation is few-body quantum mechanics of particles where the Galilean symmetry of nonrelativistic quantum mechanics is replaced by a Poincaré symmetry. The relation between the two- and three-body dynamics is fixed, up to three-body interactions by cluster properties, while the two-body input can be constructed from the same interactions that fit experimental data in the nonrelativistic two-body problem.

In 1939, Wigner [1] showed that Poincaré invariance of a quantum theory is equivalent to the existence of a unitary representation of the Poincaré group on the Hilbert space of the quantum theory. Dirac [7] demonstrated the nonlinear

difficulties in constructing dynamical representations of the Poincaré group. Bakamjian and Thomas [8] solved this problem for $N = 2$, Coester [9] extended this to $N = 3$ with S-matrix cluster properties, and Sokolov [10] and Coester and Polyzou [11] constructed a complete solution for any fixed number of particles consistent with cluster properties and a spectral condition. These constructions provide an example of a class of theories consistent with the axioms of quantum field theory [12], except for microscopic locality, which requires a theory with an infinite number of degrees of freedom and cannot be tested experimentally. The model solved in this paper is based directly on the three-body construction given by Coester in 1965 [9]. This formalism has no direct connection to quantum field theory or quasipotential equations; it is similar in spirit to the nonrelativistic three-body problem, where two-body interactions fit to scattering data and cluster properties fix the three-body Hamiltonian up to three-body forces. The eventual goal is to have a mathematical model that can provide a quantitative description of few-nucleon systems at energies of a few GeV. We use our model to explore the role of Poincaré invariance in strongly interacting few-body systems and the validity of various approximation methods that are employed at these energies. The advantage of this approach is that Poincaré invariance, cluster properties, and the spectral condition are realized exactly.

The dynamical model in this paper is given by a dynamical unitary representation of the Poincaré group on the three-nucleon Hilbert space. The mass Casimir operator of this representation replaces the nonrelativistic center-of-mass Hamiltonian. Time-dependent scattering theory can be used to derive expressions for on-shell S-matrix elements in terms of "transition operators" constructed directly in terms of the resolvent of the mass Casimir operator and the interactions appearing in that operator. The Faddeev equation can be formulated and the solution can be used to construct these transition operators. As in the nonrelativistic case, the Faddeev equation has a compact iterated kernel, which implies that it can be uniformly approximated by a finite matrix.

The Faddeev equation in the relativistic theory has the same operator structure as the Faddeev equation in the nonrelativistic theory; however, there are several new complications that must be overcome. The first is that the three-body mass Casimir operator is a nonlinear function of the subsystem mass operators, so the two-body interactions appear in the three-body mass operator in a nonlinear manner. This requires developing techniques for treating square roots of noncommuting operators [13]. A second complication is that the permutation operators in the relativistic case involve coupling irreducible representations of the Poincaré group in different orders using Clebsch-Gordan coefficients for the Poincaré group; the permutation operators are exactly Racah coefficient for the Poincaré group, which have a more complicated spin and momentum dependence than the nonrelativistic permutation operators. These operators play an important role in constructing the interactions that appear in the mass Casimir operator. A third complication is that high precision two-body interactions [14–16] are normally associated with nonrelativistic two-body problems; there is no comparable fit in the relativistic case. This problem is addressed by a method employed to treat the relativistic many-body problem [17]. Finally, the group theoretic method used for solving the nonlinear problem of adding interactions in a Poincaré invariant manner does not satisfy cluster properties; this has to be repaired with an interaction-dependent unitary transformation [11]. The resolution of all of these issues in the context of our three-nucleon model is discussed in detail in Ref. [18], where we constructed the Faddeev kernel and performed a first-order calculation in the multiple scattering series. In this paper, we complete what we started in Ref. [18] by providing a fully converged solution to the Faddeev equation for laboratory kinetic energies up to 2 GeV.

While the method that we use addresses the problem of how to incorporate high-precision interactions into our framework in a manner that reproduces the experimental two-body scattering matrix elements, in our calculations we actually use the Malfliet-Tjon [19] interaction, which is linear combinations of Yukawa interactions that commute with the spin. In the context of our general framework, this is equivalent to assuming that the measured scattering data are generated by the associated nonrelativistic two-body model. The Malfliet-Tjon interaction is often used in benchmark calculations because it leads to a two-body bound state wave function with the same symmetry as the spatial part of the wave function in realistic models of the deuteron. In addition, its core is sufficiently hard that it provides a robust test of numerical methods for computing the momentum-dependent part of the three-body wave functions. It simplifies our calculations because it is a spin-independent interaction. We want to point out that the relativistic Faddeev equation with high-precision spin-dependent interactions has been solved below the pion-production threshold in a partial-wave basis [20–22]. This method becomes more inefficient when energy is increased above 250 MeV.

This article is organized as follows. In Sec. II, the formulation of the Poincaré invariant Faddeev equation is given, and numerical aspects for computing the Faddeev kernel are discussed. In Secs. III and IV, we present calculations for elastic and breakup processes in the intermediate energy regime from 0.2 to 1.5 GeV. Our focus here is the investigation of the convergence of the multiple scattering series as a function of projectile kinetic energy. We compare our calculations with selected breakup observables and investigate a simple approximation of the embedding of the two-body interaction into the three-body problem.

II. SOLVING THE RELATIVISTIC FADDEEV EQUATION

The dynamics in Poincaré invariant quantum mechanics is given by the mass Casimir operator of a dynamical representation of the Poincaré group. This mass operator plays the same role in the relativistic theory as the center-of-mass Hamiltonian plays in the nonrelativistic theory. In this framework, the Faddeev equation is a reformulation of the scattering eigenvalue problem (including asymptotic conditions) as an integral equation with a compact iterated kernel, which can be uniformly approximated by a finite dimensional matrix. Unlike other relativistic integral equation

methods, it has no relation to Bethe-Salpeter or quasipotential equations.

A detailed formulation of three-body scattering in Poincarè invariant quantum mechanics has been given in Ref. [18], where the driving term and kernel in the relativistic Faddeev equation was used to estimate cross sections for elastic as well as breakup scattering to leading order in the multiple scattering series. This is now being complemented by a complete solution of the relativistic Faddeev equation based on the numerical techniques previously used to solve the nonrelativistic Faddeev equation [3]. For the convenience of the reader, essential equations are repeated; but for the detailed derivation of the expressions, we refer the reader to Ref. [18].

Abstractly, the Faddeev equation in our Poincaré invariant model has the same form as in the Galilean invariant case. The symmetrized transition operators $U(z)$ for elastic scattering and $U_0(z)$ for breakup reactions can be expressed in terms of the solution $T(z)$ of the symmetrized Faddeev equation

$$\begin{aligned} U(z) &= P(z - M_0) + PT(z), \\ U_0(z) &= (1 + P)\, T(z), \end{aligned} \tag{2.1}$$

where M_0 is the invariant mass operator for three noninteracting particles, and the permutation operator P is given by $P = P_{12}P_{23} + P_{13}P_{23}$. The operator $T(z)$ is the solution to the symmetrized Faddeev equation

$$T(z) = T_1(z)P + T_1(z)P(z - M_0)^{-1}T(z), \tag{2.2}$$

where the operator $T_1(z)$ is the solution of

$$T_1(z) = V_1 + V_1(z - M_0)^{-1}T_1(z). \tag{2.3}$$

In Eq. (2.3), $V_1 = V_{23} = M_{23} - M_0$ is the two-body interaction embedded in the three-body Hilbert space, and M_{23} is the invariant mass operator for two-interacting particles and a spectator. The interaction V_1 is related to a two-body NN interaction, v_{23}, fit to scattering data by

$$V_1 = \sqrt{4\left(\mathbf{k}_1^2 + m_n^2\right) + 4m_n v_{23} + \mathbf{q}_1^2} - \sqrt{4\left(\mathbf{k}_1^2 + m_n^2\right) + \mathbf{q}_1^2}, \tag{2.4}$$

where $\mathbf{k}_i$ and $\mathbf{q}_i$ are the Poincaré-Jacobi momenta defined below. The computation of the kernel of Eq. (2.3) is nontrivial; we compute it exactly using techniques based on formal scattering theory, which are discussed in detail in Refs. [13,18].

Dynamical representations of the Poincaré group are constructed by diagonalizing the interacting mass Casimir operator

$$M = M_0 + V_1 + V_2 + V_3, \tag{2.5}$$

which is rotationally and translationally invariant, in a noninteracting irreducible representation of the Poincaré group [18]. The resulting eigenstates transform irreducibly with respect to a representation in which the mass Casimir is the interacting mass operator (2.5). In our direct integration approach, the SU(2) little group is uncoupled to obtain vectors that Wigner rotate together under kinematic Lorentz transformation and permutations, allowing us to realize the dynamical symmetry without directly using a partial-wave analysis. Having established the connection of the mass operator to the dynamical representation of the Poincaré group, three-body scattering observables can be calculated by solving the Faddeev equation (2.2).

The vector variables obtained by this uncoupling process are also the natural variables to use in the Faddeev equations. These variables are similar to the standard nonrelativistic Jacobi momenta, which are defined by transforming the single-particle momenta $\mathbf{p}_i$ to the rest frame of the three-body system using a Galilean boost. The corresponding Poincaré-Jacobi momenta $(\mathbf{q}_i, \mathbf{k}_i)$ are defined by replacing the Galilean boosts by rotationless Lorentz boosts $B^{-1}(\mathbf{P}/M_0)$ and $\mathbf{p}_i$ by the single-particle (on-shell) four-momenta p_i:

$$q_i = B^{-1}(\mathbf{P}/M_0)p_i = \left(\sqrt{\mathbf{q}_i^2 + m_n^2}, \mathbf{q}_i\right), \tag{2.6}$$

and

$$k_i = B^{-1}[(\mathbf{q}_j + \mathbf{q}_k)/m_{0jk}]\tfrac{1}{2}(q_j - q_k) = (0, \mathbf{k}_i), \tag{2.7}$$

where M_0 is the invariant mass of the noninteracting three-particle system, and m_{0jk} is the invariant mass of the pair of noninteracting particles with momenta q_j and q_k. Because the arguments of the boost in Eqs. (2.6) and (2.7) are operators, rather than fixed parameters, the Poincaré-Jacobi momenta are not four vectors: instead, they all undergo identical Wigner rotations under Lorentz transformations. Explicit forms of the relations (2.6) and (2.7) are

$$\mathbf{q}_i = \mathbf{p}_i + \frac{\mathbf{P}}{M_0}\left(\frac{\mathbf{P}\cdot\mathbf{p}_i}{M_0 + \sqrt{M_0 + \mathbf{P}^2}} - \sqrt{m^2 + \mathbf{p}_i^2}\right), \tag{2.8}$$

$$\begin{aligned} \mathbf{q} \equiv \mathbf{q_i} &= -(\mathbf{q_j} + \mathbf{q_k}), \\ \mathbf{k} \equiv \mathbf{k_i} = \mathbf{k_{jk}} &= \frac{1}{2}(\mathbf{q_j} - \mathbf{q_k}) \\ &\quad - \frac{1}{2}(\mathbf{q_j} + \mathbf{q_k})\left(\frac{E_j - E_k}{E_j + E_k + \sqrt{(E_j + E_k)^2 - (\mathbf{q_j} + \mathbf{q_k})^2}}\right), \end{aligned} \tag{2.9}$$

where $E_i \equiv E(\mathbf{q_i}) = \sqrt{m^2 + \mathbf{q_i}^2}$. In addition, the transformation from eigenstates labeled by single-particle momenta $\mathbf{p}_i$ to eigenstates labeled by Poincaré-Jacobi momenta is

$$|\mathbf{p_1}, \mathbf{p_2}, \mathbf{p_3}\rangle = \left|\frac{\partial(\mathbf{P}, \mathbf{k}, \mathbf{q})}{\partial(\mathbf{p_1}, \mathbf{p_2}, \mathbf{p_3})}\right|^{1/2} |\mathbf{P}, \mathbf{k}, \mathbf{q}\rangle, \tag{2.10}$$

where for $\mathbf{P} = \mathbf{0}$ the Jacobian becomes

$$\begin{aligned} &\left|\frac{\partial(\mathbf{P}, \mathbf{k}, \mathbf{q})}{\partial(\mathbf{p_1}, \mathbf{p_2}, \mathbf{p_3})}\right|^{1/2}_{\mathbf{P}=\mathbf{0}} \\ &= \left(\frac{\sqrt{[E(\mathbf{q_2}) + E(\mathbf{q_3})]^2 - \mathbf{q}^2}\ [E(\mathbf{q_2}) + E(\mathbf{q_3})]}{4E(\mathbf{q_2})E(\mathbf{q_3})}\right)^{1/2}. \end{aligned} \tag{2.11}$$

This Jacobian appears in the expressions for the permutation operators.

In the above expression, we chose, without loss of generality, particle 1 as the spectator. The matrix elements of the permutation operator are then explicitly calculated as

$$\begin{aligned} \langle \mathbf{k}', \mathbf{q}'|P|\mathbf{k}, \mathbf{q}\rangle = N(\mathbf{q}', \mathbf{q})\big[&\delta\left(\mathbf{k}' - \mathbf{q} - \tfrac{1}{2}\mathbf{q}'\, C(\mathbf{q}, \mathbf{q}')\right) \\ &\times \delta\left(\mathbf{k} + \mathbf{q}' + \tfrac{1}{2}\mathbf{q}\, C(\mathbf{q}', \mathbf{q})\right) \end{aligned}$$

$$+\delta\left(\mathbf{k}' + \mathbf{q} + \tfrac{1}{2}\mathbf{q}'\, C(\mathbf{q},\mathbf{q}')\right) \times \delta\left(\mathbf{p} - \mathbf{q}' - \tfrac{1}{2}\mathbf{q}\, C(\mathbf{q}',\mathbf{q})\right)\Big], \qquad (2.12)$$

where the function $N(\mathbf{q}',\mathbf{q})$ contains the product of two Jacobians and reads

$$N(\mathbf{q},\mathbf{q}') \equiv N(q,q',x') = \frac{\sqrt{E(\mathbf{q})+E(\mathbf{q}+\mathbf{q}')}\,\sqrt{E(\mathbf{q}')+E(\mathbf{q}+\mathbf{q}')}}{4E(\mathbf{q}+\mathbf{q}')} \times \frac{\sqrt[4]{[E(\mathbf{q})+E(\mathbf{q}+\mathbf{q}')]^2-\mathbf{q}'^2}\;\sqrt[4]{[E(\mathbf{q}')+E(\mathbf{q}+\mathbf{q}')]^2-\mathbf{q}^2}}{\sqrt{E(\mathbf{q})E(\mathbf{q}')}}, \qquad (2.13)$$

with $x' = \hat{\mathbf{q}}\cdot\hat{\mathbf{q}}'$. The function $C(\mathbf{q},\mathbf{q}')$ is calculated as

$$C(\mathbf{q}',\mathbf{q}) \equiv C(q',q,x') = 1 + \frac{E(\mathbf{q}') - E(\mathbf{q}'+\mathbf{q})}{E(\mathbf{q}')+E(\mathbf{q}'+\mathbf{q})+\sqrt{[E(\mathbf{q}')+E(\mathbf{q}'+\mathbf{q})]^2-\mathbf{q}^2}}. \qquad (2.14)$$

These permutation operators, which change the order of coupling, are essentially Racah coefficients for the Poincaré group. In the nonrelativistic case, the functions $N(\mathbf{q}',\mathbf{q})$ and $C(\mathbf{q}',\mathbf{q})$ both reduce to the constant 1 and have the relatively compact form of the matrix elements of P given in, e.g., Refs. [3,5].

In this paper, we call the modifications of the Faddeev equation associated with use of Poincaré Jacobi momenta and the associated permutation operators "kinematic corrections."

In matrix form, the Faddeev equation, Eq. (2.2), reads

$$\langle \mathbf{k},\mathbf{q}\|T\|\varphi_d,\mathbf{q}_0\rangle = \langle \mathbf{k},\mathbf{q}\|T_1 P\|\varphi_d,\mathbf{q}_0\rangle + \langle \mathbf{k},\mathbf{q}\|T_1 P(z-M_0)^{-1}T\|\varphi_d,\mathbf{q}_0\rangle, \qquad (2.15)$$

where we have factored out a delta function in the total momentum and set $\mathbf{P}=\mathbf{0}$. The "dynamical modifications" that appear in this relativistic form of the Faddeev equation are contained in the the operator $\langle \mathbf{k},\mathbf{q}|T_1(z)|\mathbf{k}',\mathbf{q}\rangle$, which is constructed by solving Eq. (2.3) using the embedded interaction (2.4). The solution of this equation is needed fully off-shell, and the methods of computation are discussed in Ref. [18]. Inserting complete sets of states in Eq. (2.2) and explicitly evaluating the permutation operator leads to

$$\langle \mathbf{k},\mathbf{q}\|T(\mathsf{W})\|\varphi_d,\mathbf{q}_0\rangle = N(\mathbf{q},\mathbf{q}_0)\, T_s\left(\mathbf{k},\mathbf{q}_0+\frac{1}{2}\mathbf{q}\, C(\mathbf{q}_0,\mathbf{q}),\mathbf{q};\varepsilon\right)\varphi_d\left(\mathbf{q}+\frac{1}{2}\mathbf{q}_0\, C(\mathbf{q},\mathbf{q}_0)\right) + \int d^3q'\, N(\mathbf{q},\mathbf{q}')\frac{T_s\left(\mathbf{k},\mathbf{q}'+\frac{1}{2}\mathbf{q}C(\mathbf{q}',\mathbf{q}),\mathbf{q};\varepsilon\right)\left\langle \mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),\mathbf{q}'\|T(\mathsf{W})\|\varphi_d\mathbf{q}_0\right\rangle}{\mathsf{W}-(\sqrt{m^2+\mathbf{q}^2}+\sqrt{m^2+\mathbf{q}'^2}+\sqrt{m^2+(\mathbf{q}+\mathbf{q}')^2})+i\epsilon}. \qquad (2.16)$$

The quantities W and $\mathbf{q}_0$ are determined by the laboratory kinetic energy $E_{\rm lab}$ of the incident nucleon,

$$\mathsf{W}^2 = (m+m_d)^2 + 2m_d E_{\rm lab}. \qquad (2.17)$$

The nucleon rest mass is given by m, the rest mass of the deuteron is $m_d = 2m - \varepsilon_d$, where ε_d is the deuteron binding energy. The Poincaré-Jacobi momentum between projectile and target, $\mathbf{q_0}$, is related to $E_{\rm lab}$ by

$$\mathbf{q}_0^2 = \frac{m_d^2 E_{\rm lab}}{\mathsf{W}^2}(E_{\rm lab}+2m). \qquad (2.18)$$

The invariant parametric energy ε which enters the two-body t matrix is given by $\varepsilon = W - \sqrt{m^2+\mathbf{q}^2}$. Since we consider bosons, we introduce the symmetrized two-body transition

matrix T_s

$$T_s(\mathbf{k},\mathbf{k}',\mathbf{q};\varepsilon) = T_1(\mathbf{k},\mathbf{k}',\mathbf{q};\varepsilon) + T_1(-\mathbf{k},\mathbf{k}',\mathbf{q};\varepsilon) = T_1(\mathbf{k},\mathbf{k}',\mathbf{q};\varepsilon) + T_1(\mathbf{k},-\mathbf{k}',\mathbf{q};\varepsilon). \quad (2.19)$$

This two-body t matrix has a simple pole at $\mathsf{W} = \sqrt{m^2+\mathbf{q}^2} + \sqrt{m_d^2+\mathbf{q}^2}$. Thus, for the practical calculation, we need to take this pole explicitly into consideration by defining

$$\hat{T}_s = \left(\mathsf{W} - \sqrt{m^2+\mathbf{q}^2} - \sqrt{m_d^2+\mathbf{q}^2}\right)T_s, \qquad \hat{T} = \left(\mathsf{W} - \sqrt{m^2+\mathbf{q}^2} - \sqrt{m_d^2+\mathbf{q}^2}\right)T, \quad (2.20)$$

and solving Eq. (2.16) for $\hat{T}$,

$$\begin{aligned}\langle \mathbf{k},\mathbf{q}\|\hat{T}(\mathsf{W})\|\varphi_d,\mathbf{q}_0\rangle = {} & N(\mathbf{q},\mathbf{q}_0)\hat{T}_s\left(\mathbf{k},\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q}),\mathbf{q};\varepsilon\right)\varphi_d\left(\mathbf{q}+\frac{1}{2}\mathbf{q}_0\,C(\mathbf{q},\mathbf{q}_0)\right) \\ & + \int d^3q' N(\mathbf{q},\mathbf{q}')\frac{\hat{T}_s\left(\mathbf{k},\mathbf{q}'+\frac{1}{2}\mathbf{q}C(\mathbf{q}',\mathbf{q}),\mathbf{q};\varepsilon\right)}{\mathsf{W}-\left(\sqrt{m^2+\mathbf{q}'^2}+\sqrt{m_d^2+\mathbf{q}'^2}\right)+i\epsilon} \\ & \times \frac{\left\langle \mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),\mathbf{q}'\|\hat{T}\|\varphi_d\mathbf{q}_0\right\rangle}{\mathsf{W}-\left(\sqrt{m^2+\mathbf{q}^2}+\sqrt{m^2+\mathbf{q}'^2}+\sqrt{m^2+(\mathbf{q}+\mathbf{q}')^2}\right)+i\epsilon}. \end{aligned} \quad (2.21)$$

For the explicit calculation, we introduce the independent variables [3]

$$k = |\mathbf{k}|, \quad q = |\mathbf{q}|, \quad x_k = \hat{\mathbf{k}}\cdot\hat{\mathbf{q}}_0, \quad x_q = \hat{\mathbf{q}}\cdot\hat{\mathbf{q}}_0, \quad x_{kq}^{q_0} = (\widehat{\mathbf{q}_0\times\mathbf{q}})\cdot(\widehat{\mathbf{q}_0\times\mathbf{k}}), \quad (2.22)$$

so that $\langle \mathbf{k},\mathbf{q}|\hat{T}|\varphi_d,\mathbf{q}_0\rangle = \hat{T}(k,x_k,x_{kq}^{q_0},x_q,q)$, is a function of five variables. In the variables of Eq. (2.22) and defining $y_{kq} = x_k x_q + \sqrt{1-x_k^2}\sqrt{1-x_q^2}x_{kq}^{q_0}$, the final expression for Eq. (2.16) reads

$$\begin{aligned}\hat{T}(k,x_k,x_{kq}^{q_0},x_q,q) = {} & N(q,q_0,x_q)\varphi_d\left(\left|\mathbf{q}+\frac{1}{2}\mathbf{q}_0\,C(\mathbf{q},\mathbf{q}_0)\right|\right) \\ & \times \hat{T}_s\left(k,\left|\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q})\right|,y_{k,\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q})};\mathsf{W}-\sqrt{m^2+\mathbf{q}^2}\right) \\ & + \int d^3q' N(q,q',x')\frac{\hat{T}_s\left(k,|\mathbf{q}'+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}',\mathbf{q})|,y_{k,\mathbf{q}'+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}',\mathbf{q})};\mathsf{W}-\sqrt{m^2+\mathbf{q}^2}\right)}{\mathsf{W}-\left(\sqrt{m^2+\mathbf{q}'^2}+\sqrt{m_d^2+\mathbf{q}'^2}\right)+i\epsilon} \\ & \times \frac{\hat{T}\left(|\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}')|,y_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q_0},x^{q_0}_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q'},y_{q'q_0},q'\right)}{\mathsf{W}-(\sqrt{m^2+\mathbf{q}^2}+\sqrt{m^2+\mathbf{q}'^2}+\sqrt{m^2+q^2+q'^2+2qq'x'})+i\epsilon}. \end{aligned} \quad (2.23)$$

While the deuteron pole can be numerically taken care of with a single subtraction in the q' integration, the free three-nucleon propagator in the second term under the integral of Eq. (2.23) contains singularities depending on q' as well as x' leading to a singular region in the $q-q'$ plane. To simplify the calculation, we carry out the integration of the kernel in a frame in which the z axis is along the direction of $\mathbf{q}$. In this frame, $x' = \hat{\mathbf{q}}'\cdot\hat{\mathbf{q}}$ and ϕ' is the azimuthal angle of $\mathbf{q}'$. With these definitions, one has

$$y_{q'q_0} = x_q x' + \sqrt{1-x_q^2}\sqrt{1-x'^2}\,\cos(\phi_{q_0}-\phi'), \qquad y_{kq'} = x_p x' + \sqrt{1-x_k^2}\sqrt{1-x'^2}\,\cos(\phi_k-\phi'), \quad (2.24)$$

where ϕ_k and ϕ_{q_0} are the azimuthal angles of $\mathbf{k}$ and $\mathbf{q}_0$ in the frame described above. Since there is a freedom in choosing the x axis, we may place $\mathbf{q}_0$ in the xz plane, this gives $\phi_{q_0} = 0$. With this choice, ϕ_k is evaluated as

$$\cos\phi_k = \frac{x_k - y_{kq}x_q}{\sqrt{1-y_{kq}^2}\sqrt{1-x_q^2}}. \quad (25)$$

Explicit definitions of the remaining variables appearing in Eq. (2.23) are given in the Appendix.

For the integration of the $3N$ propagator, each singularity in the x' integration (for fixed q') is explicitly taken into account by a subtraction. However, this leads to logarithmic singularities in q' at the boundaries $x' = \pm 1$. These we integrate

in the semianalytic fashion introduced in Ref. [3] by using cubic Hermite splines. While using cubic Hermite splines is advantageous in dealing with the logarithmic singularities, this method is not as effective as Gauss-Legendre quadrature when integrating over large, nonsingular regions. Thus, to make the most efficient use of both methods, we divide the interval of the q' integration into several integration regions, and we use Gauss-Legendre quadrature in the nonsingular integrals while keeping the cubic Hermite splines in the small regions around the singularities. With this procedure, we are able to successfully integrate over the Faddeev kernel with sufficient accuracy. For the final solution of Eq. (2.23), the kernel is successively applied, and the resulting terms are summed up as Padé sums. At the higher energies, we will also carry out the Neumann sum.

Our explicit calculations are based on a simple interaction of Malfliet-Tjon type consisting of a superposition of an attractive and repulsive Yukawa interaction that supports a bound state with the deuteron binding energy. The parameters of this nonrelativistic interaction are given in Ref. [3]. To obtain a relativistic interaction that is phase shift equivalent with the nonrelativistic one, we employ a scheme in which $4m$ multiplied with the interaction is added to the square of the noninteracting two-body mass operator. This procedure was introduced by Coester, Pieper, and Serduke [17] and used here in the form given in Ref. [13]. It guarantees that differences in the relativistic and nonrelativistic calculations first appear in the three-body calculations.

Before entering a detailed study on relativistic effects, we want to present further details on the numerical quality of our solution of the relativistic Faddeev equation. One internal consistency check of the solution is provided by the optical theorem, which states that the total cross section, being the sum of the total elastic cross section, $\sigma_{\rm el}$, and the total breakup cross section, $\sigma_{\rm br}$, must be equal to the imaginary part of the transition operator for elastic scattering U in the forward direction. In the center-of-momentum (c.m.) frame, this relation reads

$$\sigma_{\rm el} + \sigma_{\rm br} = \sigma_{\rm tot} = -16\pi^3 \frac{E_n(q_0)E_d(q_0)}{q_0 W} \mathrm{Im}(U(q_0, x = 1)). \tag{2.26}$$

Listed in Table I are our fully relativistic calculations of the total cross sections for elastic scattering and breakup reaction for projectile energies from 0.1 to 2.0 GeV, together with the total cross sections. The total cross sections are calculated as the sum of the elastic and breakup cross sections, $\sigma_{\rm tot}$, and via the optical theorem, $\sigma_{\rm op}$, from the imaginary part of the operator U in the forward direction, $x = 1$. A comparison of those two numbers for the total cross section shows that our calculations fulfill the optical theorem to about 1% or better up to 1 GeV. This error increases to about 3% at 2 GeV. Here we did not push the calculations any farther, since our model potential is too simple to take it to much higher energies anyway. For the sake of showing the numerical quality of our calculations, we included 2 GeV in Table I, but we do not show any further observables at this energy.

The transition amplitude of Eq. (2.23) is a function of five variables and is the solution of an integral equation in three dimensions. Thus, in the calculation the dependence of the result on the various choices of grids has to be considered. As far as the momentum grids are concerned, the accuracy of the calculation is most sensitive to the q grid, as already found in Ref. [3]. In Fig. 1, we show the dependence of the relative error $\Delta_q = \frac{\sigma_{\rm op}-\sigma_{\rm tot}}{\sigma_{\rm op}} \times 100$ in the optical theorem as function of the size of the q grid, N_q, for a calculation at 1 GeV projectile laboratory kinetic energy. The slope of Δ_q shows that indeed the accuracy of the calculation is strongly influenced by the size of this grid. For our calculation, $N_q = 50$ is sufficient at 1 GeV. Next, we consider the sensitivity of the calculation to the size of different angle grids. In Table II, we give the cross sections for elastic scattering and breakup reactions together with the total cross section $\sigma_{\rm op}$ extracted from the optical theorem when varying the size of the different angle grids. We can see that the results are most sensitive with respect to the grids in x_q and x'. It is common wisdom in calculations using an angular momentum basis that as the energy of the projectile increases, the number of partial waves needed to obtain a converged result increases rather quickly. In our three-dimensional calculations, all partial waves are included. The increase in energy manifests itself in a two-body t matrix acquiring a more pronounced peak structure in the forward and backward directions with respect to the angle between the two momentum vectors [23]. This peak structure at $x_q = \pm 1$ must be adequately covered in calculations at higher energies to ensure converged results. In Fig. 2, we show the relative error

TABLE I. Total elastic and breakup cross sections together with the total cross section extracted via the optical theorem calculated from a Malfliet-Tjon type potential as function of the projectile laboratory kinetic energy.

$E_{\rm lab}$ (GeV)	$\sigma_{\rm op}$ (mb)	$\sigma_{\rm tot}$ (mb)	$\sigma_{\rm el}$ (mb)	$\sigma_{\rm br}$(mb)
0.1	349.4	350.6	273.4	77.2
0.2	195.1	194.6	158.6	36.0
0.5	106.2	106.8	72.2	34.6
0.8	74.2	74.5	46.6	27.9
1.0	62.3	61.8	37.7	24.1
1.2	54.6	55.3	33.0	22.3
1.5	43.7	44.9	26.0	18.9
2.0	33.0	34.1	18.9	15.2

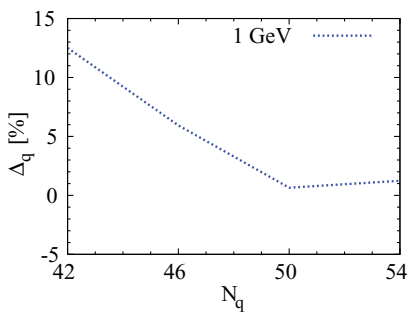

FIG. 1. (Color online) Percent error $\Delta_q = \frac{\sigma_{\rm op}-\sigma_{\rm tot}}{\sigma_{\rm op}} \times 100$ in the optical theorem as a function of the grid points in the momentum q for a calculation at 1.0 GeV.

TABLE II. Relativistic total elastic cross section, total breakup cross section, and total cross section extracted via the optical theorem calculated from a Malfliet-Tjon type potential at 1 GeV as function of the grid points. The double prime quantities are the integration variables.

$E_{\rm lab}$ (GeV)	q	x_q	x_{pq}^{q0}	x_p	p	q'	x'	ϕ''	$\sigma_{\rm op}$ (mb)	$\sigma_{\rm tot}$ (mb)	$\sigma_{\rm el}$ (mb)	$\sigma_{\rm br}$ (mb)
1.0	50	28	12	20	50	50	20	20	0.6154×10^2	0.6069×10^2	0.3678×10^2	0.2391×10^2
1.0	50	32	12	20	50	50	20	20	0.6225×10^2	0.6184×10^2	0.3774×10^2	0.2411×10^2
1.0	50	36	12	20	50	50	20	20	0.6257×10^2	0.6233×10^2	0.3812×10^2	0.2421×10^2
1.0	50	40	12	20	50	50	20	20	0.6250×10^2	0.6229×10^2	0.3809×10^2	0.2420×10^2
1.0	50	32	12	20	50	50	20	20	0.6225×10^2	0.6184×10^2	0.3774×10^2	0.2411×10^2
1.0	50	32	12	20	50	50	24	20	0.6194×10^2	0.6153×10^2	0.3753×10^2	0.2400×10^2
1.0	50	32	12	20	50	50	28	20	0.6199×10^2	0.6133×10^2	0.3744×10^2	0.2389×10^2
1.0	50	32	12	20	50	50	32	20	0.6187×10^2	0.6136×10^2	0.3746×10^2	0.2390×10^2
1.0	50	32	12	20	50	50	20	20	0.6225×10^2	0.6184×10^2	0.3774×10^2	0.2411×10^2
1.0	50	32	16	20	50	50	20	20	0.6225×10^2	0.6184×10^2	0.3773×10^2	0.2411×10^2
1.0	50	32	12	24	50	50	20	20	0.6225×10^2	0.6177×10^2	0.3777×10^2	0.2400×10^2
1.0	50	32	12	20	50	50	20	24	0.6221×10^2	0.6180×10^2	0.3773×10^2	0.2408×10^2

$\Delta_x = \frac{\sigma_{\rm op}-\sigma_{\rm tot}}{\sigma_{\rm op}} \times 100$ in the optical theorem as a function of the size of the x_q grid for three different projectile laboratory kinetic energies. The necessity of increasing the x_q grid with increasing projectile energy is clearly seen. Whereas for 0.2 GeV $N_{x_q} = 24$ is clearly sufficient, at 0.5 GeV one needs already at least 28 points, whereas at 1 GeV a minimum of 36 points is required. This conclusion is also reached in our Table III, which shows the relativistic differential cross section at selected angles while varying the x_q grid. Note that the angle x_q is related to the angular momentum of the relative motion between the spectator and the interacting pair. The angle x_p, which is related to the angular momentum of the interacting pair, is not nearly as sensitive as x_q. In Table II, we vary the x_p grid from 20 to 24 points and see hardly any difference.

It is illustrative to contrast the computational algorithm for direct integration with the experience gained when using a partial-wave basis in the $3N$ system. Our experience tells us that at $E_{\rm lab} = 200$ MeV, the total angular momentum of the $2N$ subsystem j needs to be $j_{\rm max} = 5$ to reach convergence. Furthermore, the maximum total angular momentum J of the $3N$ system required to reach convergence is $J_{\rm max} = 25/2$. Let us assume that $J_{\rm max} = j_{\rm max} + I_{\rm max}$, where $I_{\rm max} = s_i + \lambda$ is the maximal angular momentum of the projectile nucleon with respect to the target pair, s_1 is the spin of the projectile,

TABLE III. Relativistic elastic differential cross sections for selected scattering angles calculated at 1 GeV for a Malfliet-Tjon type potential as function of the size of the x_q grid. The last column indicates the percent difference with respect to the calculations for the corresponding angle in the rows above.

$E_{\rm lab}$ (GeV)	q	x_q	x_{pq}^{q0}	x_p	p	q'	x'	ϕ''	θ(deg)	$\frac{d\sigma}{d\Omega}$ (mb/sr)	Δ (%)
1.0	50	28	12	20	50	50	20	20	0.0	0.6123×10^3	
									21.8	0.1142×10^1	
									62.1	0.1159×10^{-2}	
									102.3	0.4193×10^{-2}	
									151.5	0.1233×10^{-2}	
1.0	50	32	12	20	50	50	20	20	0.0	0.6266×10^3	2.3
									21.8	0.1149×10^1	0.6
									62.1	0.1268×10^{-2}	8.5
									102.3	0.4117×10^{-2}	1.9
									151.5	0.1233×10^{-2}	0.001
1.0	50	36	12	20	50	50	20	20	0.0	0.6318×10^3	0.8
									21.8	0.1170×10^1	1.8
									62.1	0.1234×10^{-2}	2.8
									102.3	0.4127×10^{-2}	0.2
									151.5	0.1230×10^{-2}	0.2
1.0	50	40	12	20	50	50	20	20	0.0	0.6319×10^3	0.02
									21.8	0.1169×10^1	0.1
									62.1	0.1234×10^{-2}	0.01
									102.3	0.4201×10^{-2}	1.8
									151.5	0.1234×10^{-2}	0.03

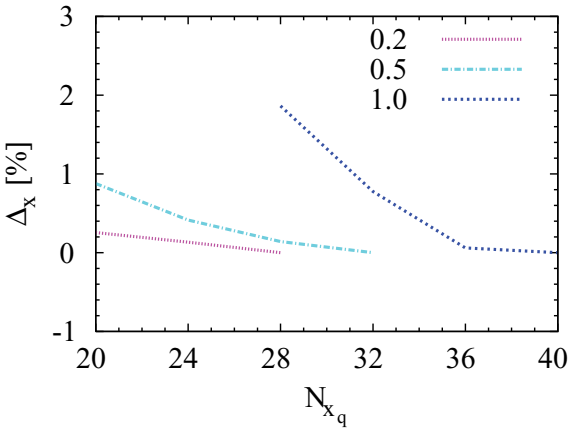

FIG. 2. (Color online) Percent error $\Delta_x = \frac{\sigma_{op}-\sigma_{tot}}{\sigma_{op}} \times 100$ in the optical theorem as a function of the grid points in the angle grid x_q, when this grid is increased successively by four Gauss-Legendre points. The different curves correspond to the three different laboratory projectile energies in GeV, indicated in the legend.

and λ is the relative orbital angular momentum between the projectile and target pair. This leads to $I_{max} = 15/2$ for a $3N$ scattering calculation at $E_{lab} = 200$ MeV. Disregarding the spin degree of freedom for the three nucleons, leading to the three-boson model under consideration here, we find that $J_{max} = 12$ with $l_{max} = 5$ and $\lambda_{max} = 7$ are necessary for a converged calculation at $E_{lab} = 200$ MeV. In the three-boson case, l and λ take the roles of j and I.

To estimate the corresponding maximal number of angular momenta needed for $E_{lab} = 1$ GeV, one needs the effective deuteron radius r_0, which leads to $\lambda_{max} = 7$ at $E_{lab} = 200$ MeV. Nonrelativistically, the $3N$ c.m. energy is given as $3/4q_0^2 = 2/3E_{lab}$, leading to $q_0 \simeq 400$ MeV/c at $E_{lab} = 200$ MeV and $q_0 \simeq 900$ MeV/c at $E_{lab} = 1$ GeV. If we roughly set $\lambda_{max} = q_0 \times r_0$, then we find at $E_{lab} = 200$ MeV a value $r_0 \simeq 3.5$ fm, which appears reasonable. Applying the same value at $E_{lab} = 1$ GeV then leads to $\lambda_{max} = 15$. Using our experience in calculating the NN system in the GeV regime [24], where one needs for converged NN observables at least $j_{max} = 14$ at 1 GeV, we estimate that a converged partial-wave 3N calculation of the three-boson system would need $J_{max} = l_{max} + \lambda_{max} = 14 + 15 = 29$.

Let us now regard the two cases: (a) $E_{lab} = 200$ MeV, $l_{max} = 5$, $\lambda_{max} = 7$, $J_{max} = 12$, and (b) $E_{lab} = 1$ GeV, $l_{max} = 14$, $\lambda_{max} = 15$, $J_{max} = 29$ in a partial-wave decomposition. To illustrate the tremendous number of partial waves needed in case (b) compared to the feasible case (a) it is sufficient to consider simple algebra for different values of J. Take for example $J = 5$. Then simple counting yields 30 different l-λ combinations in case (a) and 125 in case (b). For $J = 10$, this number increases in case (b) to 160. Moreover, since the number of total J's at 1 GeV is more than twice the number of J's at 200 MeV, it appears quite unreasonable to enforce a partial-wave decomposition at energies far above 200 MeV in the three-boson (nucleon) system. In addition, it would also be numerically very demanding to evaluate the various ingredients in the Faddeev equation reliably for the very high angular momenta.

III. RESULTS AND DISCUSSION

In the following, we present our results for elastic and breakup scattering in the energy regime from about 200 to 1500 MeV laboratory projectile kinetic energy. We start with a comparison of our model calculation to calculations based on a realistic NN force at lower energies to show that even though our model is very simple, we see similar features in the cross sections. Then we study relativistic effects at higher energies. There are several questions we want to address. First, we want to identify scattering observables that are sensitive to the difference between the relativistic and nonrelativistic formulations of the three-body problems and to study the size of those relativistic effects as a function of increasing energy. This can at present only be done with our model interaction. Second, we want to study the convergence properties of the Faddeev multiple scattering series as a function of the projectile kinetic energy. Here, the question of interest is whether, once the energy is high enough, it is sufficient to only consider the first few terms in the multiple scattering series. In addition, we also want to study some approximations to our relativistic scheme.

A. Comparison with calculations based on a realistic NN interaction at 200 MeV

The laboratory kinetic energy of 200 MeV is a perfect energy to study if the features of the $3N$ system we find based on our model interaction are also present in calculations based on a realistic model of the NN interaction, which describes the NN observables with high accuracy. The so-called high-precision interactions are fitted up to about 350 MeV, but strictly speaking, they are only valid below the pion-production threshold. We also know [20–22] that relativistic effects are already visible at 200 MeV.

We choose the CD-Bonn interaction [15] for this comparison. In Fig. 3, we show the np total cross section extracted from the SAID database [25] together with the total

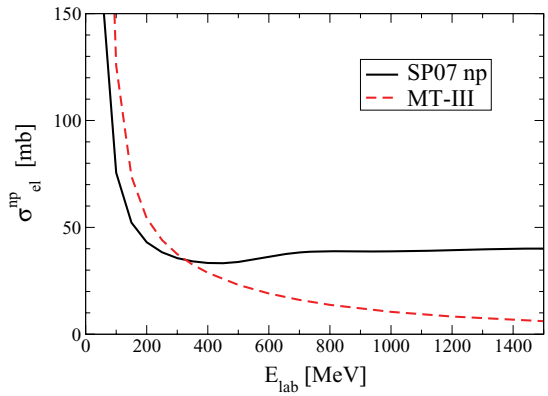

FIG. 3. (Color online) Neutron-proton differential cross section as a function of the projectile laboratory kinetic energy. The solid line represents the "experimental" cross section obtained from the SAID data base [25], and the dashed line shows the two-body cross section obtained from the Malfliet-Tjon-III potential [23] used in our calculations.

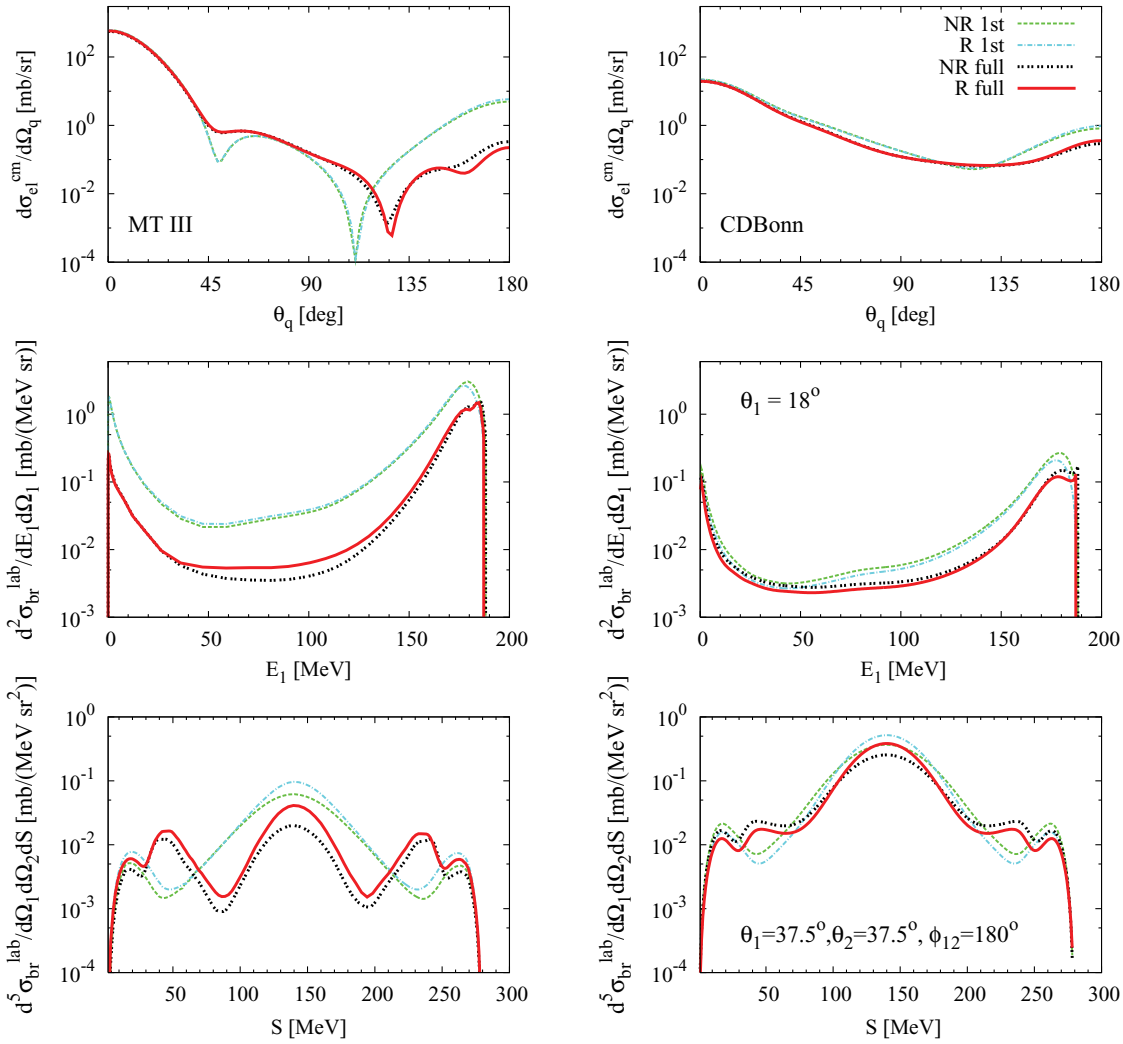

FIG. 4. (Color online) Three-nucleon scattering at $E_{\rm lab} = 200$ MeV. The left column shows results obtained from the Malfliet-Tjon-III potential assuming boson symmetry and no partial-wave decomposition, the right column shows the corresponding realistic calculations obtained with the CD-Bonn [15] potential where partial-wave decomposition is applied. The top row displays the differential cross section for elastic scattering, the middle row shows the breakup cross section for inclusive scattering for the laboratory angle $\theta_1 = 18^\circ$ of the outgoing particle. The bottom row shows the five-fold differential for exclusive breakup reaction as function of the arc-length S. The laboratory angles of the outgoing particles are $\theta_1 = \theta_2 = 37^\circ$, and $\phi_{12} = 180^\circ$. The fully relativistic converged Faddeev calculations are given by the solid lines (R), the corresponding nonrelativistic calculations by the long-dashed lines (NR). In addition the relativistic (dash-dot) and nonrelativistic (dotted) first-order calculations are shown.

cross section obtained from the Malfliet-Tjon-III (MT-III) interaction assuming bosonic symmetry. The parameters of the MT-III interaction [3] are adjusted such that a two-body bound state at $E_d = 2.23$ MeV is supported. Figure 3 shows that the experimental np total cross section falls slightly below the two-body cross section predicted by our model at energies smaller than $E_{\rm lab} \simeq 300$ MeV, is about equal between 300 and 400 MeV, and then reaches a constant value from about 600 MeV on, while our model prediction continues to decrease. The slight rise of the experimental value around 600 MeV is a manifestation of the influence of the $\Delta(1232)$ resonance in the NN system. The CD-Bonn interaction is fitted to NN observables to about 350 MeV laboratory projectile energy and thus coincides with the SAID result up to that energy.

In Fig. 4, we show a comparison of elastic and breakup cross sections at 200 MeV projectile laboratory energy for the three-dimensional calculations based on our MT-III model interaction and calculations based on a partial-wave decomposition employing the CD-Bonn potential. The top row displays the differential cross section. We see that in both cases the difference between the fully relativistic calculation and the nonrelativistic one is overall quite small and mostly visible at the backward angles, an observation already made in Ref. [20]. The differential cross section in the forward direction is much larger for our model interaction, which is consistent with the larger two-body total cross section. In addition, there are more diffraction minima in the bosonic case than in the fermionic case; however, the minimum at around 130° is present in both calculations. In the middle row, we display

the cross section for inclusive breakup scattering as a function of the laboratory kinetic energy of the ejected particle at fixed laboratory angle of 18°. Both cross sections are qualitatively similar, the fully converged Faddeev calculation gives a lower cross section than the first-order calculation, indicating the importance of rescattering contributions at this low energy. The difference between the relativistic and nonrelativistic calculations is quite small in both cases. In the calculation based on the CD-Bonn interaction, the final-state interaction (FSI) peak is more pronounced due to the virtual bound state in the $^1S_0$ state. The latter is absent in the MT-III model. The bottom row shows the five-fold differential cross section as a function of the arc-length S for configuration in which the laboratory angles $\theta_1 = \theta_2 = 37^\circ$ are measured in the scattering plane ($\phi_{12} = 180^\circ$). The position of the peaks is identical for both calculations, which is a manifestation of the fact that peak structures are given by the kinematics of the problem. In both cases, the relativistic calculation gives a significantly larger cross section for the central peak at $S \approx 140$ MeV than the nonrelativistic result, an increase by a factor of ∼1.5 for the full partial-wave calculation and by a factor of ∼2 for the full three-dimensional calculation. This increase is already present in both first-order calculations. For the MT-III model, this trend is the same for all other peaks, whereas for the CD-Bonn model, the nonrelativistic calculations gives a slightly larger cross section than the relativistic one in the peaks at small and large values of arc-length S.

Summarizing, the comparison of the cross sections obtained from our model interaction MT-III with those given by a realistic NN interaction such as CD-Bonn at 200 MeV indicates that despite our model being quite simple, the qualitative features of especially the breakup cross sections are very similar. The differences between the fully relativistic calculations and their nonrelativistic counterparts are still quite small at this low energy for elastic scattering and inclusive breakup. For the exclusive breakup, however, even at this energy complete configurations with large changes of the nonrelativistic cross section due to relativity can be found. This sensitivity of the complete breakup to relativistic effects has already been observed in Refs. [21,22].

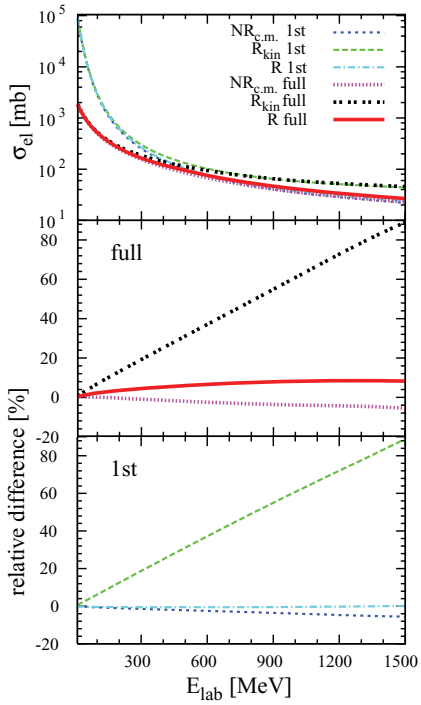

FIG. 5. (Color online) Total cross section for elastic scattering as function of the projectile kinetic energy (top panel). The fully relativistic Faddeev calculation is shown as solid line, the corresponding first-order term by the short dashed line. Calculations which only use relativistic kinematics, i.e., the Lorentz transformation between laboratory and c.m. frame together with the relativistic phase-space factor (labeled R_{kin}) are given as dotted line for a full Faddeev calculation and as short-dashed line for the first-order term. Calculations that only take into account the Lorentz transformations between the laboratory and c.m. frame (labeled $NR_{c.m.}$) are shown as a dotted line for the full Faddeev calculation and as a dash-dotted line for the first-order one. The middle panel shows the relative difference between the fully relativistic Faddeev calculation and the nonrelativistic one (solid line) together with the difference to the nonrelativistic calculation if only relativistic kinematics is considered. The bottom panel shows the corresponding relative differences when only the first-order term is taken into account.

B. Elastic scattering at intermediate energies

Starting from our model interaction, we now consider three-body scattering in the energy regime up to 1.5 GeV. The total cross section for elastic scattering is related to the symmetrized transition operator U of Eq. (2.1) via

$$\sigma_{el} = (2\pi)^4 \int d\Omega \frac{E_n^2(q_0) E_d^2(q_0)}{W^2} |\langle \varphi_d, \hat{\mathbf{q}} q_0 | U | \varphi_d, \mathbf{q_0} \rangle|^2. \quad (3.1)$$

In Fig. 5, we display the total cross section for elastic scattering as a function of the projectile kinetic energy up to 1.5 GeV obtained from our fully converged relativistic Faddeev calculation as well as the one obtained from the first-order term. It is obvious that, especially for energies below 300 MeV, the contribution of the rescattering terms is huge. Since the logarithmic scale from the top panel is unsuited to extracting detailed information about the size of relativistic effects, we show in the lower two panels the relative difference of the relativistic calculations with respect to their nonrelativistic counterparts. The bottom panel displays the relative difference between the relativistic first-order term and its nonrelativistic counterpart as a dotted line. Essentially, the first-order calculation does not show any effect. This is theoretically consistent when having in mind that in first order ($T^{1st} = tP$) only the two-body t matrix enters into the cross section. The relativistic two-body t matrix is constructed to be phase-shift equivalent to the nonrelativistic one via the Coester-Pieper-Serduke (CPS) method [13,17]. Thus, seeing no difference between the fully relativistic and the corresponding nonrelativistic calculations indicates that relativistic effects are taken into account consistently at the two-body level. Making the same comparison with fully converged Faddeev calculations (solid line in the middle panel) indicates that relativistic effects in the three-body problem increase the

elastic scattering total cross section with increasing energy. At our highest energy, 1.5 GeV, this increase is about 8.3%.

Often only effects due to relativistic kinematics are taken into account. Here we have the opportunity to study the consequences of such a simple approximation. For the calculations labeled $R_{\rm kin}$ we only consider the Lorentz transformations from laboratory to c.m. frame and the relativistic phase-space factor of Eq. (3.1), whereas the matrix elements of the operator U are calculated from the solution of the nonrelativistic Faddeev equation. The relative difference between this calculation and a completely nonrelativistic calculation is indicated by the short-dashed line in the bottom panel of Fig. 5, where only the first-order term is considered. The triple-dotted curve in the middle panel is the same comparison, but now between fully converged Faddeev calculations. For both, full and first-order calculations, the effect is huge. To understand better which piece of the kinematics included is responsible for this large enhancement of the cross section, we also plot in Fig. 5 calculations (labeled $\rm NR_{c.m.}$) that only contain the Lorentz transformation between the laboratory and c.m. frame, but carry the nonrelativistic phase-space factor in Eq. (3.1). The dash-dotted line in the lower panel shows the first-order calculation, and the dotted line in the middle panel the full Faddeev calculation. Using a Lorentz transformation in the change of frames has the effect that the two-body t matrix is calculated at a slightly different c.m. momentum $\mathbf{q_0}$, and thus there is a small effect, about a 5% underestimation of the total cross section. The huge effect is entirely due to the relativistic phase-space factor, and the relativistic dynamics then has an equally large effect of the opposite sign. This interplay of increasing effects due to relativistic phase-space factor and decreasing effects due to relativistic dynamics has been already observed in the partial-wave-based Faddeev calculations with realistic interactions [20]. It led for the elastic scattering cross section at 250 MeV to relativistic effects which were relatively small and restricted to backward angles. Recent measurements of the neutron-deuteron (nd) differential cross section at 248 MeV [26] indicate that for discrepancies of theoretical prediction in this observable, short-range components of a three-nucleon force are equally important.

The problem with approximating relativistic effects only through kinematics and phase-space factors can be easily understood in the 2+1 body problem, where the phase equivalence is achieved by choosing the invariant mass as a function of the nonrelativistic two-body Hamiltonian, $M = f(h)$. The eigenvalues equation for the scattering problems

$$|\psi\rangle = \frac{1}{f(w) - f(h_0) + i0^+}(f(h) - f(h_0))|\psi\rangle,$$
$$|\psi\rangle = \frac{1}{w - h_0 + i0^+}(h - h_0)|\psi\rangle, \tag{3.2}$$

are equivalent, but the replacement of $f(w)$ by w must be compensated by replacing the interaction $f(h) - f(h_0)$ by $h - h_0$. Including only kinematic relativistic effects is equivalent to making the replacement $h_0 \to f(h_0)$ without making the compensating replacement $v = h - h_0 \to f(h) - f(h_0)$.

We also study a more sophisticated approximation to the relativistic dynamics. In Ref. [18] we described in detail how we obtain the transition amplitude of the $2N$ subsystem, $T_1(\mathbf{k}, \mathbf{k}', \mathbf{q}; z) \equiv \langle \mathbf{k}|T_1(\mathbf{q}; z)|\mathbf{k}'\rangle$, embedded in the three-particle Hilbert space which enters the Faddeev equation, Eq. (2.15). The fully off-shell amplitude is the solution of a first resolvent type equation [13] given by

$$\langle \mathbf{k}|T_1(\mathbf{q}; z)|\mathbf{k}'\rangle = \langle \mathbf{k}|T_1(\mathbf{q}; z'|\mathbf{k}'\rangle - \int d\mathbf{k}''\langle \mathbf{k}|T_1(\mathbf{q}; z)|\mathbf{k}''\rangle \left(\frac{1}{z - \sqrt{4(m^2 + \mathbf{k}''^2) + \mathbf{q}^2}} - \frac{1}{z' - \sqrt{4(m^2 + \mathbf{k}''^2) + \mathbf{q}^2}}\right) \langle \mathbf{k}''|T_1(\mathbf{q}; z')|\mathbf{k}'\rangle. \tag{3.3}$$

Here $T_1(z')$ is taken to be the right half-shell with $z' = \sqrt{4(m^2 + \mathbf{k}'^2) + \mathbf{q}^2} + i\epsilon$. Note that in this equation the unknown matrix element is to the *left* of the kernel. It was suggested in Ref. [13] that a reasonable approximation to this embedded $2N$ transition amplitude might by the Born term of the above integral equation, which is

$$\langle \mathbf{k}|T_1^H(\mathbf{q}; z)|\mathbf{k}'\rangle \simeq \langle \mathbf{k}|T_1(\mathbf{q}; z'|\mathbf{k}'\rangle. \tag{3.4}$$

In this approximation, the fully off-shell $2N$ transition amplitude is replaced by a half-shell amplitude. The effect of this approximation is not large in elastic scattering, as shown in Fig. 6, where we plot the differential cross section in the forward direction for the fully relativistic calculations and the ones containing the approximation of Eq. (3.4) to the boost (curves labeled H). Consistently and independent of projectile energy, approximating the embedded two-body t matrix by the half-shell t matrix leads to an underprediction of the differential cross section in the forward direction. Though not plotted, this also leads to smaller total cross section for elastic scattering.

Finally, we want to investigate the convergence of the multiple scattering series as a function of the projectile laboratory kinetic energy. One might expect that with increasing energy only a few terms in the multiple scattering series are sufficient for a converged result. Our converged relativistic Faddeev calculations now allow a detailed study. This is of particular interest, since relativistic calculations in the energy regime around 1 GeV have recently been published [27–29]; they are carried out in a multiple scattering expansion of the Faddeev equations up to second order and use the off-shell continuation of the experimental NN amplitudes as two-body input.

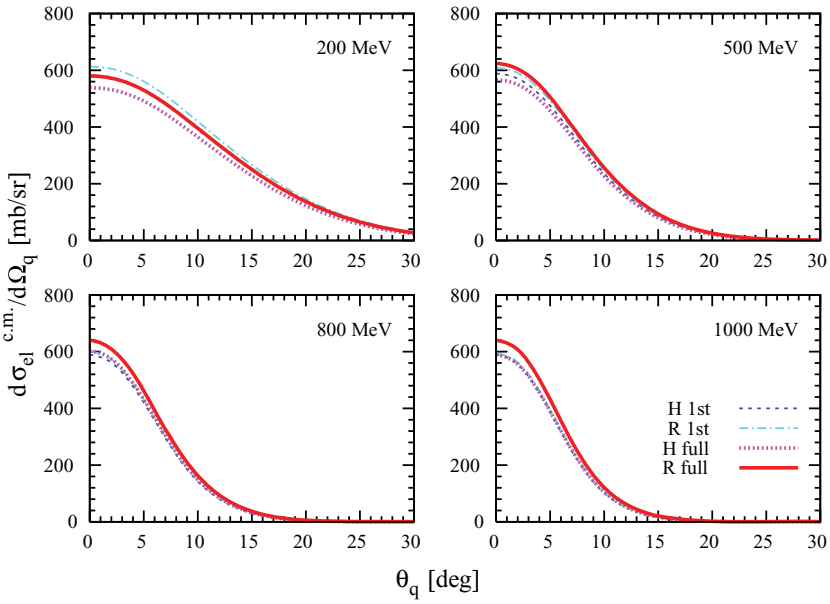

FIG. 6. (Color online) Differential cross section for elastic scattering as function of c.m. angle θ_q for selected laboratory kinetic energies. The converged solution of the relativistic Faddeev equation is given as a solid line. The dotted line shows the converged solution of the relativistic Faddeev equation in which the fully off-shell $2N$ t matrix is replaced by the half-shell t matrix. The corresponding first-order calculations are given by the short-dashed and the dash-dotted lines.

First we want to consider the convergence of the Faddeev multiple scattering series in the total cross sections for elastic scattering as well as breakup reactions as a function of projectile kinetic energy. In the bottom row of Fig. 7, the different orders (successively summed up as Neumann sum to the order indicated in the legend) are shown as functions of the projectile laboratory energy. We see a distinct difference in the behavior of the elastic total cross section from that of the breakup total cross section. While the elastic total cross section converges very rapidly, the total breakup cross section does not. The left upper panel of Fig. 7 shows the elastic total cross section as a function of the order in the multiple scattering series (the orders are successively summed up the order indicated on the x axis). Even at 200 MeV, there is very little change due to contributions from the second- or higher order rescattering terms. For the higher energies, the first-order term already captures the essential physics. This is very different for the total breakup cross section, where for 200 MeV projectile energy the full solution of the Faddeev equation is clearly necessary. For energies of 1 GeV and higher, at least one rescattering contribution (second order in the multiple scattering series) is necessary to come close to the full solution.

Since the total cross section for elastic scattering might be insensitive to higher orders in the Faddeev multiple scattering series, we plot in Fig. 8 the differential cross section at forward and backward angles as a function of the order in the multiple scattering series for the same laboratory projectile energies. Here we see that at the lowest energy, 0.2 GeV, the convergence is not as fast as the total cross section suggests. In fact, at least five orders are necessary, which is consistent with the experience from nonrelativistic calculations at low

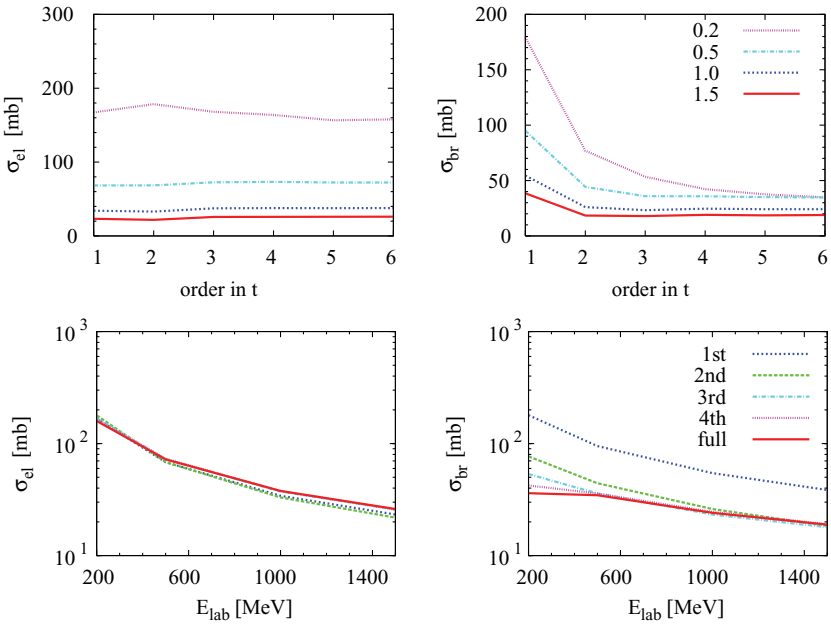

FIG. 7. (Color online) Total cross section for elastic scattering (left column) and for breakup reaction (right column). The bottom row shows both relativistic cross sections as function of the projectile laboratory kinetic energy, when starting from the first order in the Faddeev calculation successively the next three orders are added, together with the fully converged calculation (solid line). The top row displays the change in both total cross sections as a function of the order in the multiple scattering series for selected $E_{\rm lab}$ values in GeV (see legend in top right panel).

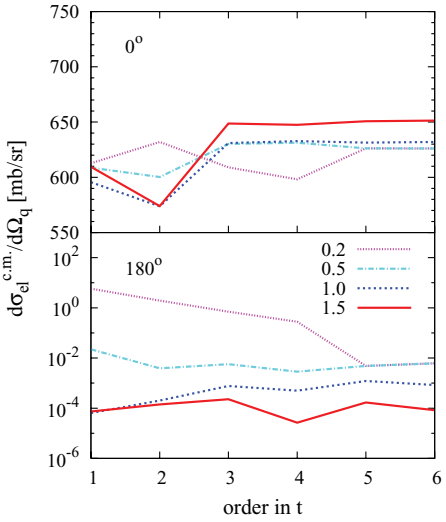

FIG. 8. (Color online) Differential cross section for elastic scattering in forward (0°) and backward (180°) direction as function of the order in the multiple scattering series for selected $E_{\rm lab}$ values indicated in the legend in units of GeV.

energies [2]. For energies of 1 GeV and higher, the forward direction is converged at the third order in t, whereas the backward angle is not as sensitive (it should be pointed out that the cross section in the backward direction is about five orders of magnitude smaller than the one in the forward direction). It seems accidental that the multiple scattering series converges faster at 0.5 GeV than at 1 GeV. However, a similar finding was presented in Ref. [27], where it was observed that polarization observables for elastic proton-deuteron (pd) scattering at 395 MeV were described better than those at 1.2 GeV, when calculating the Faddeev multiple scattering series up to the second order.

C. Breakup scattering at intermediate energies

The calculation of breakup cross sections requires the knowledge of the matrix element $\langle \mathbf{k}, \mathbf{q}|U_0|\varphi_d, \mathbf{q_0}\rangle$ in Eq. (2.1). For details of the derivation, we refer the reader to Ref. [18] and only give the final expressions here. The five-fold differential cross section for exclusive breakup is given in the laboratory frame as [30]

$$\frac{d^5\sigma_{\rm br}^{\rm lab}}{d\Omega_1 d\Omega_2 dE_1} = (2\pi)^4 \frac{E(q_0)E_d(q_0)E(q)}{2k_{\rm lab}m_d} \times \frac{p_1 p_2^2}{p_2(\mathsf{E}-E(p_1)) - E(p_2)(\mathbf{P}-\mathbf{p}_1)\cdot\hat{\mathbf{p}}_2} \times E(k)\sqrt{4E^2(k)+\mathbf{q}^2}\ |\langle \mathbf{k}, \mathbf{q}\|U_0\|\varphi_d, \mathbf{q}_0\rangle|^2 . \quad (3.5)$$

Here E is the total energy of the system and $\mathbf{P}$ its total momentum. The subscripts 1 and 2 indicate the two outgoing particles. In inclusive breakup, only one of the particles is detected, and thus one of the angles in Eq. (3.5) is integrated out. This leads to the inclusive breakup cross section in the laboratory frame[1]

$$\frac{d^3\sigma_{\rm br}^{\rm lab}}{d\Omega_1 dE_1} = (2\pi)^4 \frac{\mathsf{E}}{\mathsf{W}} \frac{E(q_0)E_d(q_0)}{4k_{\rm lab}m_d} \frac{p_1 k_a E(q)[4E^2(k_a)+\mathbf{q}^2]}{\sqrt{4E^2(k_a)+(\mathbf{P}-\mathbf{p}_1)^2}} \times \int d\Omega_k |\langle \mathbf{k}_a\mathbf{q}|U_0|\varphi_d\mathbf{q}_0\rangle|^2, \quad (3.6)$$

where k_a is determined by the condition $\mathsf{W} = \sqrt{4(m^2+\mathbf{k}_a^2)+\mathbf{q}^2} + \sqrt{m^2+\mathbf{q}^2}$.

In Refs. [18,30,31] we already pointed out and demonstrated that relativistic kinematics is essential to obtain the correct position of, e.g., the peak for quasifree scattering (QFS), especially at higher energies. The difference between a nonrelativistic calculation of the breakup cross section and a relativistic one is quite large at higher energies. However, one may argue that this difference is artificially large, since it is natural to use relativistic kinematics at higher energies. Therefore, here we will *not* compare relativistic calculations with entirely nonrelativistic calculations, but rather with calculations in which the three-body transition amplitude has been obtained from the solution of a nonrelativistic Faddeev equation but the transformations between the laboratory frame and the c.m. frame are Lorentz transformations. This is equivalent to comparing the relativistic and nonrelativistic calculations in the center-of-momentum frame.

In addition, we use the relativistic phase-space factor for the cross sections. In Fig. 9, we show the inclusive breakup cross section as a function of the laboratory kinetic energy of the ejected particle at fixed angle $\theta_1 = 24^\circ$ for different projectile kinetic energies calculated from the full solution of the relativistic Faddeev equations together with "nonrelativistic" calculations using the above-defined relativistic kinematics. There is still a shift of the position of the QFS peak toward lower ejectile energies, which increases with increasing projectile energy. There is also a very visible effect of the relativistic phase-space factor used together with the nonrelativistic three-body transition amplitude. At 1000 MeV, the size of the QFS peak is a factor of 2 larger than the exact relativistic calculation. For the lower energies, the first-order calculation yields a significantly higher QFS peak than the full calculation; whereas for the higher energies, the peak height is almost the same for the first-order and the full calculation.

Next we investigate in detail the convergence of the Faddeev multiple scattering series in the region of the QFS peak as a function of the projectile energy. In Fig. 10, we display calculations at selected energies from 200 to 1000 MeV. The solid line represents the solution of the relativistic Faddeev equation, whereas the other curves show the Neumann sum of the multiple scattering series containing the sum up to the order in the two-body t matrix as indicated in the legend. For the lowest energy, 200 MeV, it is obvious that the multiple scattering series does not converge fast. This changes considerably as the projectile kinetic energy grows. Though the variation of the different orders is not as large anymore at 500 MeV, the multiple scattering series must still be summed

[1]Note that in Ref. [18], the five-fold differential breakup cross sections contain an erroneous factor of 1/3.

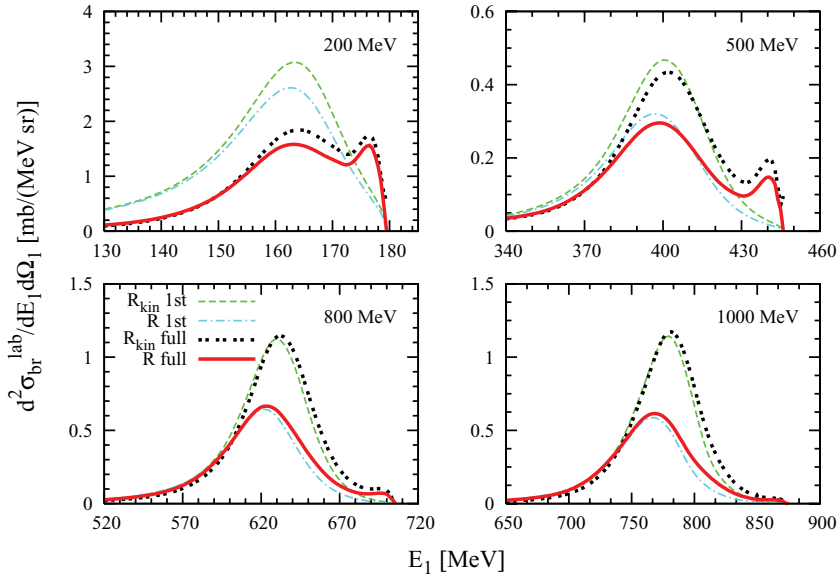

FIG. 9. (Color online) Inclusive breakup cross section as function of the laboratory kinetic energy E_1 of the emitted particle at an emission angle $\theta_1 = 24^\circ$. The incident laboratory kinetic energy for each cross section is indicated in each panel. The solid lines (R full) represent the converged relativistic Faddeev calculation, and the dotted line the corresponding first-order calculations (R 1st). The lines labeled R_{kin} correspond to calculations in which only relativistic kinematics is taken into account.

up to fourth order in the QFS peak to coincide with the full result, whereas at 800 MeV already the second order is almost identical with the full result, and even a first-order calculation can be considered quite good. This trend continues as the energy grows. Of course, first-order calculations are never able to capture the FSI peak at the maximum energy of the ejectile, nor do they describe the high-energy shoulder of the QFS peak. However, our study indicates that for energies in the GeV regime, it is very likely sufficient to consider only one rescattering term when studying inclusive breakup reactions in the vicinity of the QFS peak.

Finally, we also want to study the approximation suggested in Eq. (3.4), namely, replacing the off-shell two-body transition amplitude embedded in the three-body Hilbert space by the half-shell one. The calculations based on the approximation of Eq. (3.4) and labeled (H) are plotted in Fig. 11 together with the exact solution. Considering only the first-order calculation, we observe a similar trend as in the differential cross section for elastic scattering, the approximation slightly underpredicts the exact result, independent of the energy under consideration. However, when this approximate two-body transition amplitude is iterated to all orders in the Faddeev equation, the deviations from the exact calculations become larger. At 800 and 1000 MeV, the iteration of the exact amplitude increases the cross section in the QFS peak, whereas it decreases for the approximation with respect to the first-order term. At 200 and 500 MeV, the approximation not only gives a smaller cross section in the QFS peak but also fails to develop a FSI peak toward the maximum allowed ejectile energies. From this we conclude that Eq. (3.4) does not provide a good approximation for inclusive breakup cross sections. Our calculations indicate that at energies of 1 GeV or higher, it is important to carry out the Poincaré invariant aspects of the calculation exactly. They also indicate that it is sufficient to consider only one rescattering term to capture most features of the cross section. Although these conclusions are based on the

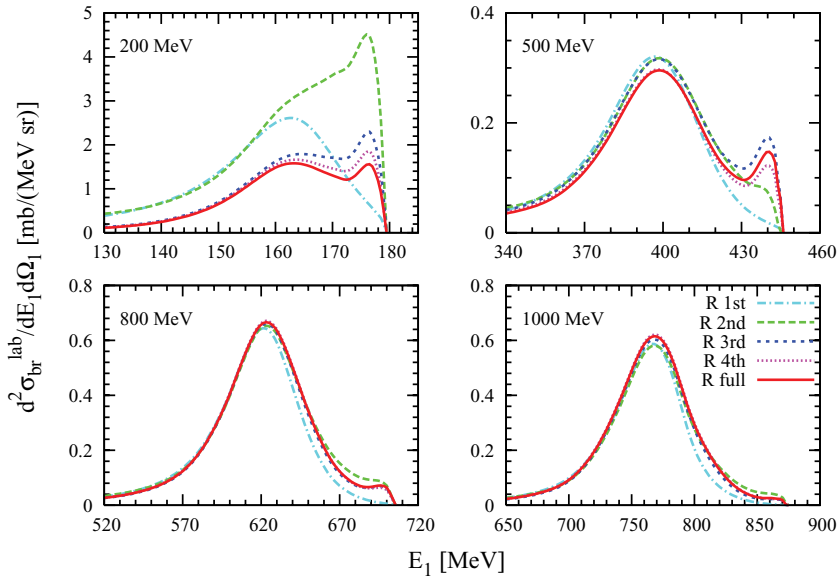

FIG. 10. (Color online) Inclusive breakup cross section as function of the laboratory kinetic energy E_1 of the emitted particle at an emission angle $\theta_1 = 24^\circ$. The incident laboratory kinetic energy for each cross section is indicated in each panel. The solid lines (R) represent the converged relativistic Faddeev calculation. The triple-dotted line shows the first-order calculation, for the short dashed line the second-order contribution is added to the previous, for the dash-dotted line the third order is added, and for the dotted line the fourth order.

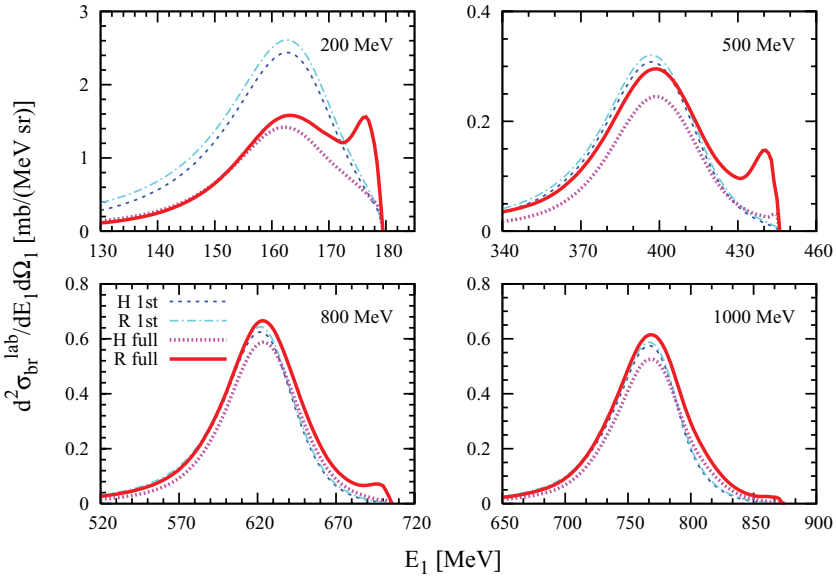

FIG. 11. (Color online) Same as Fig. 10, but comparing the converged relativistic Faddeev calculation (R) with the calculation in which the fully off-shell two-body t matrix is replaced by the half-shell one (H). The calculations labeled "1st" stand for the corresponding first-order calculations.

use of a simple model two-body interaction, we conjecture that calculations based on realistic interactions will have similar characteristics.

For our study of exclusive breakup scattering in the intermediate energy regime, we choose two different experimental situations for which data are available. First we consider the ^{2}H$(p, 2p)n$ reaction at 508 MeV, where the two outgoing protons are measured for a given angle pair θ_1-θ_2 in the scattering plane [32]. Since the convergence of the multiple scattering series is already discussed in Ref. [30], we only want to investigate the effect of the approximations previously given in this reaction. In Fig. 12, selected angle configurations are shown. The left column of the figure shows the first-order calculations; the right column the full solution of the Faddeev equation. The exact first-order calculation is given by the dotted line in the left column, and the exact full solution by the solid line in the right column. The angle combination $\theta_1 = 41.5^\circ, \theta_2 = 41.4^\circ$ is a QFS configuration. First, we see that in a QFS configuration, the first-order calculation is already almost identical to the full Faddeev calculation [30], whereas

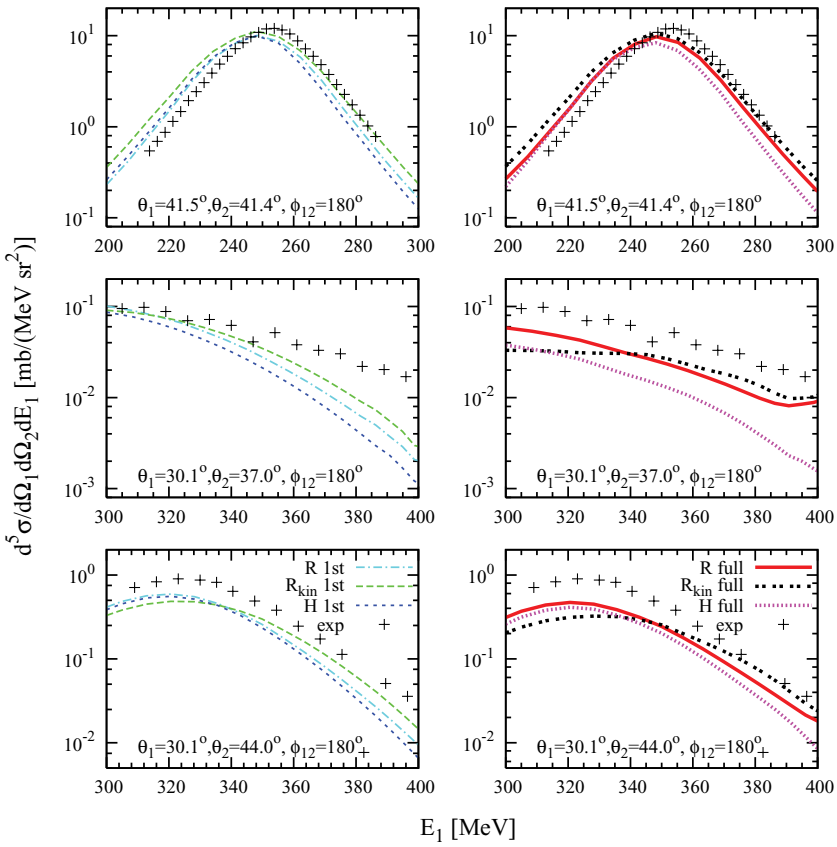

FIG. 12. (Color online) Exclusive differential cross section for the ^{2}H$(p, 2p)n$ reaction at 508 MeV laboratory projectile energy for different proton angle pairs θ_1-θ_2 with respect to the beam axis as a function of the laboratory kinetic energy of the first detected proton. The left column represents first-order calculation; right column gives full solution of the Faddeev equation. The curves labeled R represent the full relativistic calculations; for curves labeled R_{kin} only relativistic kinematics is taken into account (see text); and for curves labeled H the fully off-shell t matrix is replaced by the half-shell one. The data are taken from Ref. [32].

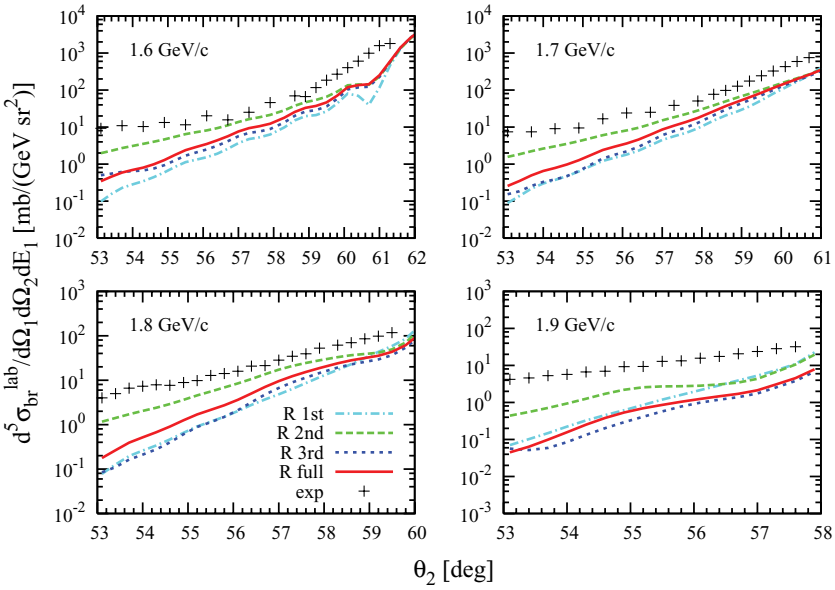

FIG. 13. (Color online) Exclusive differential cross section for the reaction ^{1}H$(d, 2p)n$ at 2 GeV deuteron energy as function of the angle θ_2 of the second of the outgoing protons for a fixed first proton momentum indicated in the figure. Curves represents the solution of the full relativistic Faddeev equation, the result of the first-order calculation, and the results of the successive addition of second- and third-order terms. The data are taken from Ref. [33].

this is not the case for the other configurations shown. If only relativistic kinematics is considered, namely, the Lorentz transformations between laboratory and c.m. frame together with the relativistic phase-space factor, and a nonrelativistic three-body transition amplitude is employed, we obtain the double-dotted curve for the first-order order calculations and the dashed line for the full solution of the Faddeev calculation. Again, the QFS configuration is quite insensitive to this approximation. However, the deviation from the exact calculation is quite visible in the other two configurations shown. Finally, we also consider the approximation suggested by Eq. (3.4), which is indicated by the dash-dotted line, labeled H in the left column (first-order calculation) and the dotted line in the right column (full solution of the Faddeev equation). Here we see that even in the QFS configuration, there are already deviations of this approximation for the high-energy shoulder. The approximation underpredicts the full solution. This tendency becomes stronger for the other two configurations. The interesting property of this approximation is that while it appears to be a reasonable approximation to the Faddeev kernel, the errors in the approximation increase when the equation is iterated. Thus we conclude that this approximation, though simplifying the calculation of the two-body t matrix embedded in the three-body Hilbert space, does not seem to capture essential structures of the two-body t matrix. The failure of this approximation, which approximates the of-shell two-body transition operator in the Faddeev equation with the half-shell transition operator, suggests that some care is necessary in modeling the off-shell behavior of the transition operators in more phenomenological schemes.

For the breakup reaction at a slightly higher energy, we consider the ^{1}H$(d, 2p)n$ reaction at 2 GeV deuteron kinetic energy [33]. Here the two outgoing protons are measured. Energetically, this reaction would correspond to pd scattering at roughly 1 GeV and thus is within the range of the calculations presented here. In Fig. 13, we show the five-fold differential cross section as function of the angle of the second detected proton for four different momenta of the first detected proton. The full relativistic Faddeev calculation is represented by the solid line. To investigate the convergence of the multiple scattering series, we show the first-order calculation as a dotted line, then successively add one (second order) and two (third order) rescattering terms to the leading order. In this reaction, the first two rescattering terms are about the same size but have opposite sign, so the third-order calculations are very close to the first-order one. We also observe that the third-order calculation is already so close to the full Faddeev calculation that the Neumann series can be considered converged with three terms.

IV. SUMMARY AND CONCLUSION

In this work, we demonstrated the feasibility of applying Poincaré invariant quantum mechanics to model three-nucleon reactions at energies up to 2 GeV. This is an important first step for studying dynamical models of strongly interacting particles in the energy range where subnuclear degrees of freedom are thought to be relevant. At these energies, the Poincaré invariance of the theory is an essential symmetry. At lower energies, nonrelativistic quantum mechanical models are powerful tools for understanding the dynamics of strongly interacting nucleons. At higher energies, the physics is more complicated, but one can expect that it is still dominated by a manageable number of degrees of freedom. Poincaré invariant quantum mechanics is the only alternative to quantum field theory where it is possible to realize the essential requirements of Poincaré invariance, spectral condition, and cluster properties [12]. It has the advantage that the Faddeev equation provides a mathematically well-defined method for exactly solving the strong interaction dynamics. The Faddeev equation in this framework is more complicated than the corresponding nonrelativistic equation because of the nonlinear relation between the mass and energy in relativistic theories, but these

difficulties can be overcome [13,18,34]. An important advance that allows these calculations to be extended to energies in the GeV range is the use of numerical methods based on direct integrations, rather than partial-wave expansions [3,4]. These have been successfully applied to the nonrelativistic three-nucleon problem. This paper demonstrates that they can also be successfully applied to the relativistic problem, even with its additional complications.

The model presented here involves three nucleons interacting with a spin-independent Malfliet-Tjon [19] type of interaction. It differs from more realistic interactions [14–16] in that it is spin independent and it does not give a high-precision fit to the two-body scattering data. In addition, the model is for fixed numbers of particles, not allowing pion production, which is an open channel at these energies. While these limitations must be addressed in realistic applications, the three-body Faddeev calculations presented in this paper provide a powerful framework for both testing approximations and examining the sensitivity of scattering observables to relativistic effects.

To investigate relativistic effects, we treat the interaction as if it were determined by fitting the cross section obtained by solving the nonrelativistic Lippmann-Schwinger equation to scattering data. When this is done with a realistic interaction, the experimental differential cross section is properly transformed from the laboratory frame to the center-of-momentum frame before the fit is done. The result of this process is that the computed differential cross section agrees with the fully relativistic experimental differential cross section in the c.m. frame as a function of the relative momentum. Thus, even though the two-body scattering observables are computed with a nonrelativistic equation, there is nothing nonrelativistic about the result. At the two-body level, the corresponding relativistic Lippmann-Schwinger equation must be designed to give the same scattering observables. This can be achieved by expressing the relativistic mass operator as a simple function of the nonrelativistic c.m. Hamiltonian [17,35]. The important consequence of this is that it does not make sense to relate the relativistic and nonrelativistic two-body models using p/m expansions; the prediction of the relativistic and nonrelativistic two-body models are identical. Real differences in the dynamics appear when the two-body dynamical operators are used to formulate the three-body dynamics. How this must be done in two- and three-body cases is dictated, up to three-body interactions, by cluster properties. The Faddeev equations for the relativistic and nonrelativistic systems have identical operator forms. The permutation operators, two-body transition operators, and free resolvents that are input to the Faddeev equation have different forms in the relativistic and nonrelativistic equations. These differences are responsible for differences in the relativistic and nonrelativistic three-body calculations.

The calculations presented in this paper have a number of consequences. The most important result was a demonstration that direct integration methods can be successfully applied to extend the energy range for converged solutions to Faddeev equations to intermediate energies. Our estimates of the number of partial waves needed for calculations at different energies suggest that it is not currently practical to extend existing partial-wave calculations beyond a few hundred MeV, while in this paper we have demonstrated convergence of the direct integration methods for laboratory energies up to 2 GeV.

While our model interaction is not realistic, when we compared the results of our calculations to relativistic calculations at 200 MeV that have been performed with realistic interactions [20–22] in a partial-wave basis, we found that the qualitative features of the realistic model are reproduced in our simple model, suggesting that some of the conclusions derived from our model should be applicable to models with realistic interactions.

Having a model with which it is possible to perform numerically exact solutions of scattering observables in the intermediate energy range provides us with a tool to test approximations that have been used in other calculations as well as to look for observables that are sensitive to the differences between relativistic and nonrelativistic models.

One common approximation that we tested is the replacement of nonrelativistic kinematic factors by the corresponding relativistic kinematic factors in a nonrelativistic model. Our tests clearly illustrated a big effect, but most of it is canceled by the associated dynamical corrections. This suggests that including only kinematic corrections can actually provide large relativistic effects. Such an approach should never be used in the absence of a complete theory with which relativistic effects can be rigorously estimated.

A second important set of approximations are multiple scattering approximations. These are expected to improve at higher energies, but it is important to understand in the context of models based on realistic interactions how high these energies have to be for convergence.

Our conclusions are that the convergence of the multiple scattering series is nonuniform. Even at 200 MeV our calculations show that the first-order term reproduces the total elastic cross section; for the total breakup cross section, at least one more iteration is needed up to about 600 MeV. Both of these observations turn out to be misleading when one investigates the differential cross sections.

While the total elastic cross section is reproduced at 200 MeV by the first-order term, the correct angular distribution requires at least five orders in the multiple scattering series. Even at 1 GeV the first-order approximation is not accurate enough at forward angles.

For inclusive breakup reactions, our computations show that the first-order calculation does not give the right size of the quasifree peak even at 1 GeV; however, for 800 MeV and above the second-order term is a good approximation. For exclusive breakup, the convergence of the multiple scattering series even at 1 GeV energy depends on a specific configuration.

Another type of approximation that is employed is the use of on-shell transition operators with a phenomenological representation of the off-shell dependence. In our formulation of the three-body problem, that off-shell behavior needs to be computed by solving a singular integral equation. It was suggested in Ref. [13] that simply replacing the off-shell two-body T by its on-shell value might be a good approximation. This was based on the observation that the difference between the on- and of-shell Faddeev kernels was small. Our calculations show that while this does not lead

to a large effect in the elastic cross section, the off-shell effects lead to nontrivial modifications when one considers the breakup cross sections. This shows that such approximation should not be used and also suggests that phenomenological parametrizations of the off-shell behavior of the two-body amplitudes need to be carefully tested, especially for the breakup reaction.

While a number of calculations have shown small relativistic effects for the three-body binding energy, nontrivial effects have already been observed in scattering observables at 200 MeV [20–22]. Our model confirms these previously observed effects and indicates that they continue into the intermediate energy region. Our calculations exhibited a number of sensitivities to relativistic effects in the breakup observables. Both the shape and size of the quasielastic peak differ from the nonrelativistic quantities.

This paper demonstrates the need for a relativistic description of few-nucleon dynamics in the intermediate energy range and shows that the problem is amenable to a numerically exact solution, using direct integration, for laboratory energies up to 2 GeV. In the future, relativistic few-body calculations will be important tools for testing the validity of approximations, such as the eikonal approximation. Obviously, extensions to include spin-dependent interactions, meson channels, and interactions that are fit to higher energy data will be needed for realistic applications. The success of the calculations in this paper provide a strong motivation for continuing this program.

ACKNOWLEDGMENTS

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APPENDIX

This Appendix contains explicit definitions of the quantities that appear in Eq. (2.23). The remaining variables in Eq. (2.23) are explicitly evaluated as

$$\begin{aligned}
&\left|\mathbf{q}+\frac{1}{2}\mathbf{q}'\,C(\mathbf{q},\mathbf{q}')\right| \\
&\quad=\sqrt{q^2+\frac{1}{4}q'^2\,C^2(q,q',x')+qq'x'\,C(q,q',x')}, \\
&\left|\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q})\right| \\
&\quad=\sqrt{q_0^2+\frac{1}{4}q^2\,C^2(q_0,q,x_q)+qq_0x_q\,C(q_0,q,x_q)}, \\
&\left|\mathbf{q}'+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}',\mathbf{q})\right| \\
&\quad=\sqrt{q'^2+\frac{1}{4}q^2\,C^2(q',q,x')+qq'x'\,C(q',q,x')}, \qquad \text{(A1)}
\end{aligned}$$

and

$$\begin{aligned}
y_{k,\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q})} &= \frac{\mathbf{k}\cdot\left(\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q})\right)}{k\left|\mathbf{q}_0+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}_0,\mathbf{q})\right|} \\
&= \frac{kq_0x_p+\frac{1}{2}kqy_{kq}\,C(q_0,q,x_q)}{k\sqrt{q_0^2+\frac{1}{4}q^2\,C^2(q_0,q,x_q)+qq_0x_q\,C(q_0,q,x_q)}}, \\
y_{k,\mathbf{q}'+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}',\mathbf{q})} &= \frac{\mathbf{k}\cdot\left(\mathbf{q}'+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}',\mathbf{q})\right)}{k\left|\mathbf{q}'+\frac{1}{2}\mathbf{q}\,C(\mathbf{q}',\mathbf{q})\right|} \\
&= \frac{kq'y_{kq'}+\frac{1}{2}kqy_{kq}\,C(q',q,x')}{k\sqrt{q'^2+\frac{1}{4}q^2\,C^2(q',q,x')+qq'x'\,C(q',q,x')}}, \\
y_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q_0} &= \frac{\left(\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}')\right)\cdot\mathbf{q}_0}{q_0\left|\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}')\right|} \\
&= \frac{qq_0x_q+\frac{1}{2}q'q_0y_{q_0q'}\,C(q,q',x')}{q_0\sqrt{q^2+\frac{1}{4}q'^2\,C^2(q,q',x')+qq'x'\,C(q,q',x')}},
\end{aligned} \qquad \text{(A2)}$$

and

$$x^{q_0}_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q'} = \frac{y_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q'}-y_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q_0}y_{q_0q'}}{\sqrt{1-y^2_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q_0}}\sqrt{1-y^2_{q_0q'}}}, \qquad \text{(A3)}$$

with

$$\begin{aligned}
y_{\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}'),q'} &= \frac{\left(\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}')\right)\cdot\mathbf{q}'}{q'\left|\mathbf{q}+\frac{1}{2}\mathbf{q}'C(\mathbf{q},\mathbf{q}')\right|} \\
&= \frac{qq'x'+\frac{1}{2}q'^2\,C(q,q',x')}{q'\sqrt{q^2+\frac{1}{4}q'^2\,C^2(q,q',x')+qq'x'\,C(q,q',x')}}. \qquad \text{(A4)}
\end{aligned}$$

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