# **Cluster decay of hot**  $56$ Ni\* **formed in the**  $32$ S $+24$ Mg reaction

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The decay of <sup>56</sup>Ni<sup>\*</sup>, formed in <sup>32</sup>S+<sup>24</sup>Mg reaction at the incident energies  $E_{c.m.}$ =51.6 and 60.5 MeV (where c.m. is the center of mass), is calculated as a cluster decay process within the preformed cluster-decay model of Gupta et al. [Phys. Rev. C 65, 024601 (2002)] reformulated for hot compound systems. Interestingly enough, the cluster decay process is shown to contain the complete structure of both the measured fragment cross sections and total kinetic energies (TKEs). The observed deformed shapes of the exit channel fragments are simulated by introducing the neck-length parameter at the scission configuration, which nearly coincides with the <sup>56</sup>Ni saddle configuration. This is the only parameter of the model, which, though, is also defined in terms of the binding energy of the hot compound system and the ground-state binding energies of the various emitted fragments. For the temperature effects included in shell corrections only, the normalized  $\alpha$ -nucleus *s*-wave cross sections calculated for nuclear shapes with outgoing fragments separated within nuclear proximity limit (here  $\sim$  0.3 fm) can be compared with the experimental data, and the TKEs are found to be in reasonably good agreement with experiments for the angular momentum effects added in the sticking limit for the moment of inertia. The incident energy effects are also shown in predicting different separation distances and angular momentum values for the best fit. Also, some light particle production (other than the evaporation residue, not treated here) is predicted at these energies and, interestingly,  ${}^{4}$ He, which belongs to evaporation residue, is found missing as a dynamical cluster-decay fragment. Similar results are obtained for temperature effects included in all the terms of the potential energy. The non- $\alpha$  fragments are now equally important, and hence present a more realistic situation with respect to experiments.

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# **I. INTRODUCTION**

Experimentally, <sup>56</sup>Ni is an extensively studied compound system by using different entrance channels, namely,  $^{16}$ O  $1+^{40}Ca$ ,  $^{28}Si+^{28}Si$ , and  $^{32}S+^{24}Mg$ , and at various incident energies ranging from 1.5 to 2.2 times the Coulomb barrier (see, e.g., Ref.  $[1]$  and the other direct and more recent papers  $[2-8]$ . At such incident energies, the incident flux is found to get trapped by the formation of a compound nucleus (CN), which is in addition to a significant large-angle elastic scattering cross section. For lighter masses  $(A_{CN}$ <44), such a compound nucleus decays subsequently by the emission of mainly light particles  $(n, p, \alpha)$  and  $\gamma$  rays; i.e., with a very small component of heavy fragment  $(A>4)$  emission. An experimental measure of this so-called particle evaporation residue yield is the CN fusion cross section. For somewhat heavier systems, such as  ${}^{48}Cr$  and  ${}^{56}Ni$ , a significant decay strength to  $A > 4$  fragments (the mass-asymmetric channels) is also observed, which could apparently not arise from a direct reaction mechanism because of the large massasymmetry differences between the entrance and exit channels. The measured angular distributions and energy spectra are consistent with fissionlike decays of the respective compound systems.

For the  ${}^{32}S+{}^{24}Mg\rightarrow{}^{56}Ni*$  reaction, in one of the experiments, the mass spectra for  $A=12-28$  fragments and the total kinetic energy (TKE) for only the most favored (enhanced yields)  $\alpha$ -nucleus fragments are measured at the energies  $E_{lab}$ =121.1 and 141.8 MeV, or equivalently, at  $E_{c.m.}$  $=$  51.6 and 60.5 MeV, respectively [2,3] (where c.m. is the center of mass). Note that  $56Ni$  is a negative *Q*-value system (negative  $Q_{\text{out}}$ , different for different exit channels) and hence would decay only if it were produced in heavy ion reactions with sufficient compound nucleus excitation energy  $E_{CN}^*$  (= $E_{c.m.}$ + $Q_{in}$ ), to compensate for the negative  $Q_{out}$ , the deformation energy of fragments  $(E_d)$ , their total kinetic energy  $(TKE)$ , and the total excitation energy  $(TXE)$ , in the exit channel, as

$$
E_{CN}^* = |Q_{out}(T)| + E_d(T) + TKE(T) + TXE(T); \qquad (1)
$$

see Fig. 1, where  $E_d$  is neglected because fragments are considered to be spherical. Here  $Q_{\text{in}}$  is positive (=16.68 MeV for  $32S+24Mg$  entrance channel) and hence adds to the entrance channel kinetic energy  $E_{\text{c.m.}}$  of the two incoming nuclei in their ground states. In another experiment  $[4]$  for  $32S+24Mg$  reaction at  $E_{cm} = 51.0$  and 54.5 MeV, the excitation-energy spectra for only the symmetric  $^{28}Si + ^{28}Si$ and near-symmetric  $^{24}Mg+^{32}S$  channels are measured, whose analysis indicate that a specific set of states in  $^{28}$ Si correspond to highly deformed bands. In other words, the expected shapes of some of the observed fragments in the exit channel could be relatively deformed. It is interesting to note that this result is supported by a very recent study of the <sup>28</sup>Si+<sup>28</sup>Si reaction at  $E_{\text{c.m.}}$  = 55 MeV, where the population of highly excited states in the  $^{24}Mg$ ,  $^{28}Si$ , and  $^{32}S$  nuclei indicated a selective and enhanced population of deformed bands  $[6]$ . In a still other recent experiment  $[7]$ , the incident energy used in the same  ${}^{32}S+{}^{24}Mg\rightarrow{}^{56}Ni*$  reaction is  $E_{lab}$  $=130$  MeV and an enhanced emission yield by a factor



FIG. 1. The *s*-wave  $(\ell=0)$  scattering potential for <sup>56</sup>Ni<sup>\*</sup>  $\rightarrow$ <sup>12</sup>C+<sup>44</sup>Ti, calculated for no temperature effects in  $E_c$  and  $V_p$ , i.e.,  $V(R) = E_c + V_p$ . The *Q* values are calculated from *T*-dependent binding energy  $B(T) = V_{LDM} + \delta U(T)$ . The actually calculated decay path for  $V(R_a) = Q_{\text{eff}}(\Delta R) = V[C_t(T) + \Delta R]$  is shown, where  $\overline{\Delta R}$  is an average of the separation distances for different fragmentations (different  $\eta$  values).

of 1.5–1.8 is observed for <sup>8</sup>Be over the two  $\alpha$  particles. The aim of our present work is to understand some of the results of these experiments.

The above stated light particle  $(A \leq 4)$  production, the evaporation residue, is very satisfactorily understood as the equilibrated compound nucleus emission in the statistical Hauser-Feshbach analysis  $[3,9-12]$ , using the LILITA or CAS-CADE codes. The Hauser Feshbach calculations are also extended to include the complex fragments, such as the ones observed in the experiments mentioned above. These are considered in the, so-called, BUSCO code  $[10]$  or the Extended Hauser-Feshbach scission-point model  $[12]$ . Within the framework of the Extended Hauser-Feshbach method [12], the above noted observed enhanced emission of  ${}^{8}$ Be over the evaporation of two  $\alpha$  particles in the <sup>32</sup>S+<sup>24</sup>Mg reaction is shown related to an increased deformation of the heavier fragment  $^{48}Cr$  [7]. The emission of complex fragments  $[A>4]$ , also called the intermediate mass fragments (IMFs) or "clusters"] is alternatively treated as the binary fission of a compound nucleus in the statistical fission models  $[13,14]$ , using the GEMINI code  $[9]$  or the saddle-point "transition-state" model  $[3,5,11]$ . The transition-state model, treating the complex fragment emission as a compoundnucleus fission process (the fusion-fission), seems to explain the observed mass spectra and excitation-energy spectra rather well for the  $32\text{S}+24\text{Mg}$  reaction at the two energies used in respective experiments  $[3,4]$ . Also, the measured TKE for the symmetric fission is comparable to the saddlepoint potential energy at  $\ell = 36\hbar$  [3]. Then, there are other processes, such as the deep-inelastic orbiting or scattering, which have also been studied for this reaction but do not seem to explain the observed data  $[3]$ .

In the statistical fission models  $[13,14]$ , the fission decay of a compound nucleus is determined by the phase space (level density) available at the "transition" configuration, which is saddle or scission in these models. For light systems, this choice can lead to a significant population of many energetically allowed mass channels, though there is no structure information of the compound system in these fission models. However, the structure effects of the compound system seem to influence the observed yields strongly since strong resonance behavior is observed in the measured excitation functions of large-angle elastic and inelastic scattering yields in several light systems (see, e.g., Ref.  $[6]$ ). One possibility to account for such structure effects is via the process of fragment (or clusters) preformation in a compound nucleus and its subsequent decay as a cluster decay process, proposed recently by some of us  $[15,16]$ . The structure information enters the process via the preformation probability (also known as the spectroscopic factors) of the fragments. We follow this approach of preformed cluster decay  $[15,16]$ in this paper.

The cluster-decay process was recently studied  $[15]$  for the compound system  $56Ni^*$ , using the preformed clusterdecay model (PCM) of Gupta and collaborators  $[17–21]$ . It was shown that for the decay of  $56Ni^*$ , the two processes of binary fission (the dynamical collective mass transfer calculated, by some of us  $[22-25]$ , in the quantum mechanical fragmentation theory  $[26–28]$  and cluster decay are almost indistinguishable, particularly at higher angular momenta. However, this work was a simple model study where the role of the TKE was analyzed and found to be significant for  $\alpha$ -nucleus structure in the measured yields. This model is more recently reformulated  $[16]$  for the IMFs emitted from an excited  $116Ba*$  compound nucleus produced in the low energy  $58$ Ni +  $58$ Ni reaction. The IMFs in  $116$ Ba\* are shown to be produced as multiple "clusters" of masses  $A < 20$  and only at  $E_{lab} > 200$  MeV, in agreement with experiments. Both of these works  $[15,16]$  show that the IMFs in the decay of excited 116Ba\* or the complete mass spectra in decay of excited <sup>56</sup>Ni<sup>\*</sup> have their origin in the macroscopic liquid drop energy (the shell effects are almost zero at the excitation energies involved). For  $116Ba*$  decay, the light particle  $(Z \le 2)$  emission, other than the promptly emitted via the statistical evaporation process (not treated in this model), is also shown to be given, but at higher energies where only the pure liquid drop model (LDM) energies enter the calculations. Thus, the macroscopic liquid drop energy  $V_{\text{LDM}}$  is shown playing the most important role in the cluster decay calculations. Apparently, the compound nucleus being hot at the energies involved, the  $V_{\text{LDM}}$  should also depend on the temperature *T*. This is done here in this paper for the decay of  $56Ni*$  formed in the  $32S+24Mg$  reaction at the two energies,  $E_{\text{c.m.}}$  = 51.6 and 60.5 MeV [2,3]. Also, the other terms of the potential, which constitute the scattering potential *V*(*R*), are considered *T* dependent.

The *T*-dependent liquid drop model used is that of Davidson *et al.* [29], which is based on the semiempirical mass formula of Seeger [30]. The model parameters of Seeger's formula at  $T=0$  are refitted in view of the present availability of a larger dataset for binding energies [31]. For the *T* dependence in  $V(R)$ , we follow Davidson *et al.* [29] and some other authors  $[32]$ , as discussed below. The deformation effects of the fragments (and the neck formation between them) are included here within the extended model of Gupta and collaborators  $|33-35|$ , via a neck-length parameter at the scission configuration which simulates the twocenter nuclear shape parametrization, used for both the light and heavy nuclear systems. A similar method has been used earlier by other authors  $[3,11,12]$ , discussed below.

The dynamical cluster decay model for hot compound systems, a reformulation of the PCM of Gupta and coworkers  $[17-21]$  for ground-state decays, is presented in Sec. II and its application to the hot  $56Ni*$  nucleus data from Refs.  $[2,3]$  in Sec. III. The (statistical) evaporation of light particles that occur promptly before the beginning of the binary decay process of cluster emission studied here, is not included in this paper. Hence, any discussion of light particle emission is that of one which is in addition to the ones emitted promptly. Finally, a summary of our results is presented in Sec. IV.

## **II. THE DYNAMICAL CLUSTER-DECAY MODEL FOR HOT COMPOUND SYSTEMS**

The cluster-decay model developed here is the PCM of Gupta *et al.*  $[17-21]$  for the ground-state decays, reformulated for hot and excited compound systems. In this model, we treat the complex fragments (the IMFs or clusters) as dynamical collective mass motion of preformed fragments through the barrier. It is based on the well-known dynamical (or quantum mechanical) fragmentation theory  $[26-28]$  developed for fission and heavy ion reactions, and used later for predicting the exotic cluster radioactivity  $[36-38]$  also. This theory is worked out in terms of the collective coordinates of mass asymmetry  $\eta = (A_1 - A_2)/(A_1 + A_2)$  and relative separation *R*, which in a PCM allows to define the decay half-life  $T_{1/2}$ , or the decay constant  $\lambda$ , as

$$
\lambda = \frac{\ln 2}{T_{1/2}} = P_0 \nu_0 P,\tag{2}
$$

where the preformation probability  $P_0$  refers to  $\eta$  motion and the penetrability *P* to *R* motion. Apparently, the two motions are taken as decoupled, an assumption justified in our earlier works [26,27,39]. The  $\nu_0$  is the barrier assault frequency. In terms of the partial waves, the decay cross section

$$
\sigma = \frac{\pi}{k^2} \sum_{\ell=0}^{\ell_c} (2\ell+1) P_0 P; \quad k = \sqrt{\frac{2\mu E_{\text{cm}}}{\hbar^2}}, \tag{3}
$$

with  $\mu = [A_1A_2/(A_1+A_2)]m = \frac{1}{4}Am(1-\eta^2)$  as the reduced mass and  $\ell_c$ , the critical (maximum) angular momentum, defined later; *m* is the nucleon mass. This means that  $\lambda$  in Eq. (2) gives the *s*-wave cross section, with a normalization constant  $\nu_0$ , instead of the  $\pi/k^2$  in Eq. (3). However, in the present calculations, made for  $\ell=0$  case, the normalization constant is obtained empirically from the experimental data.

For  $\eta$ -motion, we solve the stationary Schrödinger equation in  $\eta$ , at a fixed *R*,

$$
\left\{-\frac{\hbar^2}{2\sqrt{B}_{\eta\eta}}\frac{\partial}{\partial\eta}\frac{1}{\sqrt{B}_{\eta\eta}}\frac{\partial}{\partial\eta}+V_R(\eta,T)\right\}\psi^{\nu}(\eta)=E^{\nu}\psi^{\nu}(\eta),\tag{4}
$$

with  $\nu=0,1,2,3,...$  and  $R=R_a=C_t(=C_1+C_2)$ , the first turning point, fixed empirically for the ground-state  $(T=0)$ decay since this value of  $R$  (instead of the compound nucleus radius  $R_0$ ) assimilates to a good extent the effects of both the deformations  $\beta_i$  of two fragments and neck formation between them [35]. In other words, the deformation effects of the two fragments are included here in the scattering potential  $V(R, T=0)$  for each  $\eta$  by raising the first turning point  $R_a$  from  $R_a = R_0$  to  $R_a = C_t$  or  $C_t + \Sigma \delta R(\beta_i)$ , which is equivalent of lowering of the barrier, as is found to be the case for deformed fragments [35]. This method of inclusion of fragment deformation and the parametrization of the neck zone via a neck-length parameter  $\delta R$  in the present calculations is quite similar to what has been achieved in both the transition-state model of Sanders  $(3,11)$  (in saddle point configuration) and the Extended Hauser-Feshbach method of Matsuse and collaborators  $[12]$  (in scission point configuration). It is also shown in Ref.  $|35|$  that the alternative of calculating the fragmentation potential  $V(\eta)$  and scattering potential  $V(R)$  for deformed nuclei is not practical since the experimental deformation parameters for all the possible fragments  $(A_1, A_2)$ , required for calculating  $V(\eta)$ , are generally not available. The deformation effects of nuclei in our calculations are further included via the Süssmann central radii  $C_i = R_i - (b^2/R_i)$ , with the radii  $R_i = 1.28A_i^{1/3} - 0.76$  $+0.8A_i^{-1/3}$  fm and surface thickness parameter  $b=0.99$  fm. Note that the  $C_t$  are different for different  $\eta$  values, and hence  $C_t$  is  $C_t(\eta)$ .

The eigensolutions of Eq.  $(4)$  give the preformation probability

$$
P_0 = \sqrt{B_{\eta\eta}} |\psi[\eta(A_i)]|^2 (2/A), \qquad (5)
$$

 $(i=1 \text{ or } 2)$ , where  $\psi(\eta)$  is  $\psi^{\nu=0}(\eta)$  if the ground-state solution is chosen. However, the decay of  $56$ Ni in the ground state  $(T=0, R_a=C_t)$  is not allowed since  $Q_{out}(T=0)$  is negative.

For the decay of a hot compound nucleus, we use an ansatz  $\lceil 16 \rceil$  for the first turning point,

$$
R_a = C_t(\eta, T) + \Delta R(\eta, T),\tag{6}
$$

which depends on the total kinetic energy  $TKE(T)$ . The corresponding potential  $V(R_a)$  acts like an effective, positive  $Q$ value,  $Q_{\text{eff}}$ , for the decay of the hot compound system at temperature *T* to two fragments in the exit channel observed in the ground states  $(T=0)$ . Thus, in terms of the respective binding energies  $B$ ,  $Q_{\text{eff}}$  is defined as

$$
Q_{eff}(T) = B(T) - [B_1(T=0) + B_2(T=0)]
$$
  
= TKE(T) = V(R<sub>a</sub>). (7)

Since,  $R_a = C_t(\eta)$  for  $T=0$ ,  $\Delta R(\eta)$  corresponds to the change in TKE at *T* with respect to its value at  $T=0$ , and hence can be estimated exactly for the temperature effects included in the scattering potential  $V(R)$ . Note that in Eq.  $(6)$   $C_t$  is also taken to depend on temperature, as is defined in the following. Also,  $\Delta R$  depends on  $\eta$ . In the following, however, based on our earlier work  $[16]$ , instead, we use a constant average value  $\Delta R$  (independent of  $\eta$ ) which also takes care of the additional  $\Sigma \delta R(\beta_i)$  effects of the deformations of fragments and neck formation between them. Note that  $\Delta R$  is the only parameter of the model, though it is shown that the structure of the calculated mass spectrum is nearly independent of the exact choice of this parameter value. The corresponding  $Q_{\text{eff}}$  is denoted as  $Q_{\text{eff}}(\Delta R)$ .

In the above definition of  $Q_{\text{eff}}$ , apparently the two fragments would come out of the barrier and go to ground state  $(T\rightarrow 0)$  only by emitting some light particle(s) and/or  $\gamma$  rays of energy, defined as (see Fig.  $8$ )

$$
E_x = B(T) - B(0) = Q_{\text{out}}(T) - Q_{\text{out}}(T=0) + \Delta B
$$
  
=  $Q_{\text{eff}}(T) - Q_{\text{out}}(T=0) = \text{TKE}(T) - \text{TKE}(T=0).$  (8)

Equation  $(8)$  means that one can also write

$$
Q_{\text{eff}}(T) = \text{TKE}(T) = Q_{\text{out}}(T=0) + E_x = \text{TKE}(T=0) + E_x, \tag{9}
$$

which is what one observes experimentally, i.e., the fragments in the ground state with  $Q_{\text{out}}(T=0)$  [=TKE(*T*=0)] and light particle(s) and  $\gamma$  rays of energy  $E_r$ . The remaining excitation energy of the decaying system is then

$$
E_{\text{CN}}^* - E_x = |Q_{\text{out}}(T)| + \text{TKE}(T = 0) + \text{TXE}(T), \quad (10)
$$

which again shows that the exit channel fragments are obtained with their TKE in the ground state, i.e., with TKE  $(T=0)$ . The excitation energy TXE  $(T)$  in Eq.  $(10)$  is used in the secondary emission of light particles from the fragments, which are not treated here. Instead, we compare our calculations with the primary pre-secondary-evaporation fragment emission data.

We notice from Eq.  $(7)$  that for the ground-state  $(T=0)$ decay,

$$
Q_{\text{eff}}(T=0) = Q_{\text{out}}(T=0) = \text{TKE}(T=0),
$$
 (11)

as is the case for exotic cluster radioactivity  $\lfloor 21,38 \rfloor$ . In fact, one can write Eq.  $(7)$  as

$$
Q_{\text{eff}}(T) = Q_{\text{out}}(T) + \Delta B,\tag{12}
$$

where

$$
\Delta B = [B_1(T) + B_2(T)] - [B_1(T=0) + B_2(T=0)], \quad (13)
$$

the difference of binding energies at temperature *T* and the ground-state binding energies of the two fragments. Also, for the ground-state  $(T=0)$  decays, according to Eq. (8),  $E<sub>x</sub>$  $=0$  (no particle or  $\gamma$ -ray emission), as is known to be true for exotic cluster radioactivity  $[21,38]$ .



FIG. 2. The fragmentation potential for <sup>56</sup>Ni at  $T=0, R=C_t$ , using the experimental binding energies (solid squares)  $[31]$  and the empirically fitted Seeger's binding energies (solid circles) with the new constants of Table I. Here, MS14 means the shell corrections from the empirical method of Myers and Swiatecki [40] with *Z* and  $N=14$  as the magic numbers.

Thus, at temperature *T*, the preformation factor  $P_0$  in Eq. (5) is calculated at  $R_a = C_t(\eta) + \Delta R$ , with the temperature effects also included in  $\psi(\eta)$  through a Boltzmann-like function,

$$
|\psi|^2 = \sum_{\nu=0}^{\infty} |\psi^{\nu}|^2 \exp(-E^{\nu}/T), \qquad (14)
$$

with the compound nucleus temperature  $T$  (in MeV) related as

$$
E_{\text{CN}}^* = (A/9)T^2 - T; \tag{15}
$$

and for the penetrability *P*, Eqs. (6) and (7) for each  $\eta$  and *T* values mean that

$$
V(R_a) = V(C_t + \overline{\Delta R}) = V(R_b) = Q_{\text{eff}}(\overline{\Delta R}) = \text{TKE}(T),
$$
\n(16)

with  $R<sub>b</sub>$  as the second turning point, and penetrability *P* calculated as the WKB tunneling probability for the path shown in Fig. 1 (or Fig. 8), as

$$
P = \exp\bigg[-\frac{2}{\hbar} \int_{R_a}^{R_b} \{2\,\mu \big[ \, V(R) - Q_{\text{eff}} \big] \}^{1/2} dR \bigg],\tag{17}
$$

solved analytically  $[18]$ .

The fragmentation potential  $V_R(\eta, T)$  at any temperature  $T$ , in Eq.  $(4)$ , is calculated within the Strutinsky renormalization procedure, as

$$
V_R(\eta, T) = \sum_{i=1}^{2} \left[ V_{\text{LDM}}(A_i, Z_i, T) \right] + \sum_{i=1}^{2} \left[ \delta U_i \right] \exp \left( -\frac{T^2}{T_0^2} \right) + E_c(T) + V_P(T) + V_{\ell}(T), \tag{18}
$$

where the *T*-dependent liquid drop energy  $V_{LDM}(T)$  is that of Ref. [29], with the (Seeger's) constants at  $T=0$  refitted to give the experimental binding energies  $B$  [31], defined as  $B$  $=V_{LDM}(T=0)+\delta U$ . The shell corrections  $\delta U$  are calculated in the ''empirical method'' of Myers and Swiatecki [40]. Some of these details are given in Appendix I. Figure 2 illustrates the kind of comparisons obtained for  $V(\eta)$  calculated at  $R = C_1 + C_2 = C_t$  and  $T=0$  for the experimental and newly fitted binding energies. Apparently, the binding energies fit within  $1-1.5$  MeV.

The  $V_p$  is an additional attraction due to the nuclear proximity potential  $[41]$ , which is also considered temperature dependent here,

$$
V_P(R,T) = 4\,\pi\overline{R}(T)\,\gamma b(T)\Phi(s,T),\tag{19}
$$

where  $\overline{R}(T)$  and  $\Phi(s,T)$  are, respectively, the inverse of the root mean square radius of the Gaussian curvature and the universal function, which is independent of the geometry of the system, given by

$$
\Phi(s,T) = \begin{cases}\n-\frac{1}{2}(s-2.54)^2 - 0.0852(s-2.54)^3 & \text{for } s \le 1.2511 \\
-3.437 \exp\left(-\frac{s}{0.75}\right) & \text{for } s \ge 1.2511,\n\end{cases}
$$
\n(20)

$$
\bar{R}(T) = \frac{C_1(T)C_2(T)}{C_t(T)},
$$
\n(21)

and  $\gamma$  is the specific nuclear surface tension given by

$$
\gamma = 0.9517 \left[ 1 - 1.7826 \left( \frac{N - Z}{A} \right)^2 \right] \text{ MeV fm}^{-2}. \quad (22)
$$

In Eq. (20),  $s(T)$  (=[ $R - C<sub>t</sub>(T)$ ]/ $b(T)$  is the overlap distance, in units of b, between the colliding surfaces. The temperature dependence in radii  $R_i$  is given as [29,32]

$$
R_i(T) = r_0(T)A_i^{1/3} = 1.07(1 + 0.01T)A_i^{1/3},
$$
 (23)

with the surface width

$$
b(T) = 0.99(1 + 0.009T^2). \tag{24}
$$

The same temperature dependence of  $R(T)$  is also used for Coulomb potential  $E_c(T) = Z_1 Z_2 e^2 / R(T)$ , where the charges  $Z_i$  are fixed by minimizing the potential  $V_R(\eta, T)$  in the charge asymmetry coordinate  $\eta_Z = (Z_1 - Z_2)/(Z_1 + Z_2)$ . The shell corrections  $\delta U$  in Eq. (18) are considered to vanish exponentially for  $T_0$ =1.5 MeV [42].

Also, for the angular momentum effects (so far included here for the calculation of total kinetic energy only)

$$
V_{\ell}(T) = \frac{\hbar^2 \ell(\ell+1)}{2I(T)}.
$$
 (25)

In the nonsticking limit, where  $R_a = C_1(T) + C_2(T) + \Delta R$  $= C_t(T) + \Delta R$ , the moment of inertia in Eq. (25) is given by

$$
I(T) = I_{NS}(T) = \mu R_a^2.
$$
 (26)

In this case, the separation distance  $\Delta R$  is assumed to be beyond the range of nuclear proximity forces, which is about 2 fm. However, when  $\Delta R$  is within the range of nuclear proximity  $(< 2$  fm), we get in the complete sticking limit

$$
I(T) = I_S(T) = \mu R_a^2 + \frac{2}{5} A_1 m C_1^2 + \frac{2}{5} A_2 m C_2^2.
$$
 (27)

For the  $\ell$  value, in terms of the bombarding energy  $E_{\text{c.m.}}$  of the entrance channel  $\eta_{\rm in}$ , we have

$$
\ell = \ell_c = R_a \sqrt{2\,\mu \left[E_{\text{c.m.}} - V(R_a, \eta_{\text{in}}, \ell = 0)\right]}/\hbar, \qquad (28)
$$

or, alternatively, it could be fixed for the vanishing of fusion barrier. In this work, however, we use  $\ell=0$  for the IMF cross sections and take  $\ell_c$  as a variable parameter for TKE calculations (see Fig.  $8$ ).

The mass parameters  $B_{\eta\eta}(\eta)$ , representing the kinetic energy part in Eq. (4), are the smooth classical hydrodynamical masses  $[43]$ , since we are dealing here with a situation where the shell effects are almost completely washed out.

The assault frequency  $v_0$ , in Eq. (2), is given simply as

$$
\nu_0 = \frac{(2E_2/\mu)^{1/2}}{R_0},\tag{29}
$$

with the kinetic energy of the lighter fragment  $E_2$  $=(A_1 / A)Q_{\text{eff}}$ , for the  $Q_{\text{eff}}$  shared between the two fragments as inverse of their masses. However, for the calculations of s-wave cross sections, instead of  $v_0$ , we use an empirically determined normalization constant.

Finally, the temperature-dependent scattering potential  $V(R,T)$ , normalized to the exit channel binding energy, is

$$
V(R,T) = Z_1 Z_2 e^2 / R(T) + V_P(T) + V_{\ell}(T). \tag{30}
$$



FIG. 3. The fragmentation potentials  $V(\eta, R, T)$  for <sup>56</sup>Ni<sup>\*</sup> compound system, calculated at the ground state  $(T=0, R_a=C_t)$  and at various temperatures with  $R_a = C_1(T) + C_2(T) + \Delta R$  values as shown. The *T* dependence is included only in the shell corrections.

This means that all energies are measured with respect to  $B_1(T) + B_2(T)$ , and the fragments go to ground state (*T*  $\rightarrow$ 0) via the emission of light particle(s) and/or  $\gamma$  rays of energy  $E_x$ .

#### **III. CALCULATIONS**

The calculations are made in two steps:  $(i)$  with temperature effects included only in the shell corrections, i.e., using  $\delta U(T)$ , but *T*-independent  $V_{LDM}$  and  $V(R)$ ; and (ii) with temperature effects included also in both the liquid drop energy and scattering potential, i.e., using  $V_{LDM}(T)$ ,  $\delta U(T)$ , and  $V(R,T)$ . This allows us to study explicitly the role of temperature in different terms of the potential. In both sets of the calculations, we first take  $\ell=0$ , i.e., use  $V_{\ell}=0$  throughout, but then study the effect of adding this term to the potential  $V(R)$  for calculating the TKE alone.

#### **A. Temperature effects only in shell corrections**

Figure 3 gives our calculated fragmentation potentials *V*( $\eta$ ,*T*) for <sup>56</sup>Ni<sup>\*</sup> at *T*=0, as well as at other two temperatures referring to the compound nucleus excitation energies  $E_{CN}^*$  of the experiments of Refs. [2,3]. The *R* values chosen are  $R = R_a = C_t$  at  $T=0$ , and, as before [16],  $R = R_a + \Delta R$ with  $\Delta R$  = 0.30 and 0.31 fm, arbitrarily, for *T* = 3.39 and 3.60 MeV, which correspond to the experimental energies  $E_{c.m.}$  $=$  51.6 and 60.5 MeV, respectively. The near independence of the structure in  $V(\eta)$  on *R* value was studied in our earlier



FIG. 4. The fragment preformation probability  $P_0$  for <sup>56</sup>Ni<sup>\*</sup>, calculated by using the fragmentation potentials in Fig. 3 for the two experimental *T* values only.

works [21,28]. The  $\delta U$  at these temperatures reduce almost to zero. However, we notice that the  $N=Z$ , $A=4n$   $\alpha$ -nucleus structure is obtained at all temperatures, which has its origin apparently in the macroscopic liquid drop energy and is due to the ''Wigner term'' in it, as was also shown earlier in Refs. [15,16]. Note that here the  $V_{\text{LDM}}$  and other *R*-dependent terms  $(E_c$  and  $V_p$ ) are not yet *T* dependent (see the following section). This means that for use of only  $\delta U$  as *T* dependent, the  $N=Z$   $\alpha$ -nuclei fragments should be produced preferentially in the decay of  $56Ni*$  at all temperatures.

The preformation probability  $P_0$  of the fragments, calculated for the potentials in Fig. 3, is given in Fig. 4. The case of  $T=0$  is not shown here since cold <sup>56</sup>Ni (in the ground state) cannot decay because of its negative  $Q$  value. Interestingly enough, for both the temperatures (the two temperatures are nearly the same), the yields are large for only a small window of  $A \le 16$  fragments, including the light particles ( $A \le 3$ ). Also, the  $\alpha$ -nucleus fragments <sup>4</sup>He, <sup>8</sup>Be,  $12C$ , and  $16O$ , and the light particle  $1H$  (in addition to the evaporation residues, not included here) are preferentially preformed. This means that, out of all the fragments observed in the decay of  $56Ni^*$ , the ones with  $A \le 16$  are strongly preformed. The other ones with  $A > 16$ , if observed, must have larger penetrability *P*, since the decay constant is a combined effect of both the preformation factor  $P_0$  and penetrability  $P(v_0)$  is nearly constant).

Figure 5 gives the results of our calculation for the normalized decay constants, equivalently, the *s*-wave production cross sections for *only* the most favored (largest yields or cross sections)  $\alpha$ -nucleus fragments, compared with the experimental data at two energies, taken from Fig. 9 of Ref. [3]. In the lower panel, the calculation at  $E_{\text{c.m.}}$  = 51.6 MeV, using  $\Delta R$ =0.3 fm, is fully normalized to the experimental data for the favored  $\alpha$ -nucleus fragments only. Then, in the upper panel, for the higher energy  $E_{\text{c.m.}}$  = 60.5 MeV, we find that for the use of the same normalization as obtained in lower panel and for a further normalization of the  $A=12$ fragment yield, the best fit to the  $\alpha$ -nuclei fragment data is obtained for  $\Delta R$ =0.29 fm, a value lower than that used for





FIG. 5. The calculated *s*-wave cross sections for the  $\alpha$ -nucleus fragments compared with the measured ones produced in the reaction  ${}^{32}S+{}^{24}Mg \rightarrow {}^{56}Ni*$  at  $E_{c.m.}=51.6$  and 60.5 MeV. The data are from Fig. 9 of Ref. [3]. The calculations for  $E_{\text{c.m.}} = 51.6 \text{ MeV}$  in the lower panel are made for  $\Delta R$ =0.30 fm, and are normalized completely to the experimental data. Using the same normalization, the calculations for  $E_{\text{c.m.}}$ =60.5 MeV in the upper panel are made for  $\Delta R$ =0.29, 0.30, and 0.31 fm and compared with the experimental data, for a further normalization of the data at fragment mass *A* = 12. Only the  $\alpha$ -nucleus fragments are studied, since they have the largest cross sections. The dotted lines are drawn only for guiding the eyes.

the lower incident energy  $E_{\text{c.m.}}$  = 51.6 MeV. This is contrary to the expected behavior of increased *R* at higher temperatures, but, as we shall see below in Fig. 7, this is a result of our having not included here the contribution of angular momentum term in the fragmentation potential  $[V_\ell=0$  in  $V(\eta, T)$ ] and hence in the cross sections. Also, the inclusion of temperature effects in other terms (the  $V_{LDM}$ ,  $E_c$ , and  $V_p$ ) are important, as is shown below in Sec. III B. Hence, Fig. 5 (and Fig. 7 below) shows that the dynamical clusterdecay model contains the required structure of the measured yields (and TKEs) in this experiment  $[2,3]$ .

Figure 6 shows the complete mass spectra for decay of 56Ni\* calculated at both the energies and compared with the measured yields [3]. The calculated yields are for the energetically favored, most probable, mass fragments (see Figs. 3) and 4). Note that the experimental data in Refs.  $[2,3]$  are available only for fragments heavier than mass 11, and in steps of mass 1 for  $E_{\text{c.m.}}$ =60.5 MeV, but in steps of only mass 2 for  $E_{\text{c.m}}$  = 51.6 MeV due to a deteriorated mass resolution at the lower bombarding energy. For comparisons, the calculations are normalized to the experimental data for onefragment mass  $(A=20)$  only. The role of the penetrability *P* is evident in this figure, since some of the strongly preformed fragments, such as  ${}^{4}$ He and  ${}^{1}$ H in Fig. 4, are now shown as less favored decays (smaller cross sections, not shown in Fig. 6 since they lie below the chosen scale). The same is true for weakly preformed fragments (in Fig. 4), with  $A > 16$ . Specifically, amongst the light particles, mass-3 fragment  $(^{3}He)$ is shown to be produced with a large cross section, and for lighter fragments ( $A$ <12), instead of  $A=8$  ( ${}^{8}$ Be), the fragments with  $A=6$  and 10 are shown to be produced with larger cross sections. This means that of all the residue products  $(A \leq 4$ , not studied here), only the mass-3 fragment  $(3He)$  is produced and that the mass-4 ( $^{4}He$ ) fragment is not at all produced as a dynamical cluster-decay fragment. This



FIG. 6. Same as for Fig. 5, but studied for all the fragments at  $E_{c.m.}=51.6 \text{ MeV}, \Delta R=0.30 \text{ fm}$ (upper panel), and  $E_{c.m.}=60.5 \text{ MeV}$  and  $\Delta R$  $=0.29$  fm (lower panel). The calculations are normalized to the experimental data for one fragment mass  $(A=20)$  only. The calculated (*s*-wave) cross sections are for the energetically most favored fragments in  $\eta$  coordinate, i.e., fragments lying at the minimum in the fragmentation potential  $V(\eta)$ , minimized in  $\eta$ <sub>Z</sub> coordinate.-



FIG. 7. The measured and calculated total kinetic energy (TKE) for average  $\overline{\Delta R}$  for the reaction  ${}^{32}S+{}^{24}Mg\rightarrow{}^{56}Ni^* \rightarrow A_1+A_2$ , at the two incident energies. The calculations for  $\ell \neq 0$  are made for both the cases of sticking and nonsticking limits (see text). The data are from Fig.  $5$  (summed over all the angles) of Ref.  $[3]$ . The same data are also given in Fig. 10 of Ref. [11], where it should be noted that Fig. 10(a) refers to  $E_{\text{c.m.}}$  = 60.5 MeV and Fig. 10(b) to  $E_{\text{c.m.}}$  $= 51.6$  MeV.

nonoccurrence of <sup>4</sup>He as a dynamical cluster-decay product is an interesting result, giving a strong support to the credential of the model. For mass-8 (Be) decay, perhaps the contribution of higher  $\ell$  values is important. For the heavier fragments  $(A > 20)$ , the calculated cross sections are rather small due to the fact that here the contribution of only  $\ell=0$  term is considered. Also, in experiments it is difficult to separate the contributions of direct (such as  $\alpha$ -transfer and orbiting processes) and compound-nucleus yields for the heavy mass fragments  $(A>20)$  (see Ref. [1] and references therein). Thus, in view of the fact that we are dealing here with only the  $\ell=0$  case and that the temperature effects are not included in full in the potential, the comparisons in Fig. 6 between the theory and experiments could be said to be at least reasonable.

Figure 7 shows the results of our calculation for total kinetic energy (TKE), with angular momentum  $\ell$  effects included only in the scattering potential  $V(R)$ . We notice that the calculated TKEs for the sticking limit (using  $I<sub>S</sub>$ ) compare reasonably good with the experimental data. This means that, even though  $\Delta R$  is nonzero (=0.29 and 0.3 fm), the sticking limit for the moment of inertia is preferred. Also, unlike the  $\Delta R$  values, the  $\ell$  values required for the case of higher energy data is now of a larger value  $(\ell = 25\hbar$  for  $E_{\text{c.m.}}$ = 60.5 MeV as compared to 24 $\hbar$  for  $E_{\text{c.m.}}$ = 51.6 MeV), as expected. The measured TKEs are taken from Ref.  $[3]$ .

#### **B. Fully temperature-dependent potential**

Figure 8 shows the scattering potential  $V(R,T,\ell)$  for tem-



FIG. 8. Same as for Fig. 1, but with  $\ell$ , and *T* dependences included in  $E_c$  and  $V_p$  also, i.e., the scattering potential is  $V(R,T,\ell) = E_c(T) + V_p(T) + V_{\ell}(T)$  with *Q* value now calculated from  $B(T) = V_{LDM}(T) + \delta U(T)$ . Only the sticking limit of moment of inertia is used in  $V_{\ell}(T)$ . The  $T=0$  potential is shown for comparisons. For all  $\ell$  values, the decay path (dotted line), shown for  $\Delta R$ , begins at  $R = R_a$  (marked explicitly). The distribution of energies and definitions of other quantities such as  $\Delta B$  and  $E_x$  are indicated for the calculated  $\Delta R$  value.

perature effects included in all the terms of the potential (compare this figure for  $\ell=0$  with Fig. 1, where temperature effects are included in  $\delta U$  only). Notice that as  $\ell$  value increases, the TKE( $\Delta R$ ) value increases, since the decay path for all the  $\ell$  values begins at  $R=R_a$ . Figure 9 gives our



FIG. 9. Same as for Fig. 3, but for *T* dependence in all the terms of the fragmentation potential, and at  $\overline{\Delta R}$  values as shown.



FIG. 10. Same as for Fig. 4, but for the fragmentation potential of Fig. 9.

calculated fragmentation potentials  $V(\eta, T)$ . The *T* values chosen are the same as in Fig. 3, where temperature effects were included only in the shell corrections. The *R* values here are  $R(T) = C_1(T) + C_2(T) + \Delta R = C_1(T) + \Delta R$ , with  $\Delta R$  values as shown in the figure. We notice in Fig. 9 that, due to the inclusion of temperature effects in all terms, the minima in the potential, which were earlier only for  $\alpha$  nuclei, are now obtained for both the  $\alpha$  and non- $\alpha$  fragments. This happens, possibly, due to the pairing energy term  $\delta(T)$  in formula (A2) of Davidson *et al.* [29], which goes to zero for *T*>2 MeV. Thus, with the addition of temperature, not only the shell structure effects go to zero but also the explicitly preferred  $\alpha$ -nucleus structure washes out. Also, we notice that the light particle  $(A \leq 4)$  structure changes; in particular, the minimum at  $4$ He disappears and a shallow minimum at  ${}^{2}$ H appears.

Figure 10 gives the preformation factors  $P_0$  for the two experimentally chosen temperatures only, since the groundstate  $(T=0)$  decay is not possible. We notice that the formation yields are large only for light fragments  $(A \leq 16)$  and are of the same orders as in Fig. 4, except that now the non- $\alpha$  fragments are also preformed equally strongly. However, the calculated decay constants, equivalently, the fragment  $(s$ -wave) production cross sections, in Fig. 11 do not show much improvement in their comparisons with experiments. The comparisons are now somewhat better for the heavier fragments, but the yields for fragments lighter than  $A=9$  are very low, lying below the chosen scale. On the other hand, the calculated TKEs in Fig. 12 compare nicely (even better than in Fig. 7) with the experimental data. Only the case of sticking limit is shown since the  $\Delta R$  values are still within the proximity limits. Note that the  $\ell$ -dependent contribution is so far added here only in the scattering potential  $V(R,T)$ , and not yet in the fragmentation potential  $V(\eta, T)$ , which is needed for both the preformation factor and penetrability. This extension is being carried out.

## **IV. SUMMARY**

In summary, we have reformulated for hot nuclear systems the preformed cluster-decay model (PCM) of Gupta and collaborators for ground-state decays and applied it for the



FIG. 11. Same as for Fig. 6, but for *T* dependence in all the terms of the fragmentation potential, and at  $\Delta R$  values as shown. For lighter fragments, the calculated yields are not shown as they lie below the chosen scale.



FIG. 12. Same as for Fig. 7, but for *T* dependence in all the terms of the fragmentation potential, and at  $\Delta R$  values as shown.

first time to the decay of a light compound nucleus such as <sup>56</sup>Ni<sup>\*</sup> formed in the reaction <sup>32</sup>S+<sup>24</sup>Mg carried out at two incident energies  $E_{\text{c.m.}}$ =51.6 and 60.5 MeV [2,3]. In this experiment, the mass spectra for fragments heavier than mass 12 and the total kinetic energies (TKEs) for only the favored  $\alpha$ -nucleus fragments are measured. Also, at another energy, in between the two above, an enhanced yield is observed for  ${}^{8}$ Be over the two  $\alpha$ -particle emissions [7]. Our calculations are made first for the temperature effects included only in shell corrections and then in all terms of the potential, and in each case for  $\ell=0$  only. The contribution due to  $\ell$  is added only for estimating the TKEs. Similar to the saddle-point model  $[11]$  and/or the scission-point model  $[12]$ , the deformations of the fragments are taken into account by the parametrization of the neck-in zone, proposed by Gupta and collaborators [33–35]. This quantity is  $\eta$  dependent and could be calculated but is taken as a parameter here, which is the only parameter of the model.

For the temperature effects included in shell corrections only, we find that the  $\alpha$ -nucleus fragments are favorably preformed and are due to the macroscopic liquid drop energy alone since the shell effects are almost zero at the energies under consideration. The calculated decay constants or the normalized *s*-wave cross sections, in particular for the  $\alpha$ -nucleus fragments are found to contain the complete structure of the experiments for a nuclear shape with fragments separated by about 0.3 fm which is within the limits of nuclear proximity effects. Some of the light particles (other than the ones constituting the evaporation residue, not included here) are also predicted to be there in the mass spectra, but  $4$ He is shown to be absent. With angular momentum effects included, the calculated TKEs are found to compare rather nicely with experimental data for the moment of inertia calculated for a sticking limit.

For the full temperature effects in the potential, the non- $\alpha$ fragments are also preformed equally strongly as the  $\alpha$ -nucleus fragments. The cluster decay process now occurs at a somewhat larger separation distance, which is also temperature dependent. Hence, the TKEs for a sticking moment of inertia are now in somewhat better agreement with the experiments. However, the comparison between the calculated (*s*-wave) and measured mass spectra is not improved much, which calls for the inclusion of  $\ell$ -dependent potential in the calculations of yields also, which is underway.

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## **APPENDIX: TEMPERATURE-DEPENDENT BINDING ENERGIES**

In Eq.  $(18)$  we have defined, within the Strutinsky renormalization procedure, the binding energy *B* of a nucleus at temperature *T* as the sum of liquid drop energy  $V_{LDM}(T)$  and shell correction  $\delta U(T)$ ,

$$
B(T) = V_{\text{LDM}}(T) + \delta U \exp\left(-\frac{T^2}{T_0^2}\right). \tag{A1}
$$

The *T*-dependent liquid drop part of the binding energy  $V_{LDM}(T)$  used here is that of Davidson *et al.* [29], based on the semiempirical mass formula of Seeger [30], as

$$
V_{\text{LDM}}(T) = \alpha(T)A + \beta(T)A^{2/3} + \left(\gamma(T) - \frac{\eta(T)}{A^{1/3}}\right)
$$

$$
\times \left(\frac{I^2 + 2|I|}{A}\right) + \frac{Z^2}{r_0(T)A^{1/3}} \left(1 - \frac{0.7636}{Z^{2/3}}\right)
$$

$$
- \frac{2.29}{\left[r_0(T)A^{1/3}\right]^2} + \delta(T)\frac{f(Z,A)}{A^{3/4}}, \tag{A2}
$$

where

$$
I = a_a (Z - N), \quad a_a = 1,
$$

and, respectively, for even-even, even-odd, and odd-odd nuclei,

$$
f(Z,A) = (-1,0,1).
$$

For  $T=0$ , Seeger [30] obtained the constants, by fitting all even-even nuclei and 488 odd-*A* nuclei available at that time, as

$$
\alpha(0) = -16.11
$$
 MeV,  $\beta(0) = 20.21$  MeV,

$$
\gamma(0) = 20.65
$$
 MeV,  $\eta(0) = 48.00$  MeV,

with the pairing energy term

$$
\delta(0) = 33.0 \text{ MeV},
$$

Ζ	$\boldsymbol{N}$	$\alpha(0)$	$a_a$	Ζ	$\boldsymbol{N}$	$\alpha(0)$	$\boldsymbol{a}_a$	Ζ	$\boldsymbol{N}$	$\alpha(0)$	$\boldsymbol{a}_a$
$\mathbf{1}$	$\overline{c}$	$-15.85$ 0.10		6	$\overline{9}$	$-15.70$ 0.10		10	$\boldsymbol{7}$	$-15.70$	0.50
	3	$-16.95$ 0.12			10	$-15.10$ 0.10			$\,8\,$	$-15.90$	0.90
	$\overline{4}$	$-13.00$ 0.05			$11\,$	$-14.80$ 0.10			13	$-15.95$	0.50
	5	$-13.70$ 0.12			12, 13, 15, 16	$-15.00$ 0.80			14	$-15.70$ 0.50	
2	$\mathbf{1}$	$-15.50$ 0.10			14	$-14.85$ 0.80			$9 - 12, 15 - 22$	$-16.16$	0.88
	$\sqrt{2}$	$-16.00$ 0.10		7	$\ensuremath{\mathfrak{Z}}$	$-14.30$ 0.20		11	$\boldsymbol{7}$	$-15.55$	0.50
	3	$-16.80$ 0.30			$\overline{4}$	$-15.20$ 0.50			$\,8\,$	$-15.80$	0.50
	4,5	$-14.20$ 0.30			$\sqrt{5}$	$-16.20$ 0.80			14	$-15.95$	0.50
	6	$-13.50$ 0.10			$\sqrt{6}$	$-16.55$ 0.80			$9 - 13,15 - 24$	$-16.20$	0.86
	7,8	$-13.00$ 0.10			$\boldsymbol{7}$	$-16.80$ 0.80		12	$8 - 10$	$-16.11$	0.90
3	1,2,4,5	$-16.60$ 0.10			$\,8\,$	$-16.30$	0.80		$11 - 25$	$-16.20$	0.86
	3	$-16.98$ 0.98			$\overline{9}$	$-16.20$ 0.80		13	$8 - 10$	$-16.11$	0.90
	6	$-13.80$ 0.98			10,11	$-15.90$ 0.94			$11 - 26$	$-16.22$	0.84
	$\tau$	$-14.30$ 0.40			12	$-15.75$ 0.94		14	$8 - 12$	$-16.11$	0.90
	8,9	$-13.20$ 0.10			13	$-15.80$ 0.94			$13 - 20,27,28$	$-16.28$	0.84
4	1	$-13.00$ 0.01			14	$-15.65$ 0.94			$21 - 26$	$-16.22$ 0.84	
	$\overline{c}$	$-14.50$ 0.10			15	$-15.90$ 0.94		15	$9 - 13,20 - 31$	$-16.30$ 0.82	
	$\mathfrak{Z}$	$-16.20$ 0.80			16	$-16.00$ 0.94			$14 - 19$	$-16.36$ 0.78	
	$\overline{4}$	$-16.98$ 0.98			$17\,$	$-16.10$ 0.93		16	$10-14,21-28$	$-16.30$ 0.82	
	5	$-16.70$ 0.60		8	$\overline{4}$	$-14.00$ 0.94			$15 - 20$	$-16.40$ 0.78	
	6	$-15.50$ 0.80			$\sqrt{5}$	$-15.25$ 0.94			$29 - 33$	$-16.32$ 0.80	
	$\tau$	$-15.30$ 0.50			$\sqrt{6}$	$-15.90$ 0.94		17	$11-14,20,21,29-34$	$-16.36$ 0.78	
	8	$-14.30$ 0.10			$\boldsymbol{7}$	$-16.35$ 0.94			$15 - 19$	$-16.45$ 0.78	
	9	$-14.00$ 0.10			$8\,$	$-16.20$ 0.94			$22 - 28$	$-16.32$ 0.82	
	10	$-13.30$ 0.01			$\overline{9}$	$-16.18$ 0.94		18	$12 - 14, 21, 22, 31 - 35$	$-16.36$ 0.78	
5	$\overline{2}$	$-14.60$ 0.10			$10\,$	$-15.95$ 0.94			$15 - 20$	$-16.45$	0.78
	3	$-16.50$ 0.10			$11\,$	$-15.93$ 0.94			$23 - 30$	$-16.32$ 0.78	
	$\overline{4}$	$-16.60$ 0.60			12,14	$-15.85$ 0.94		19	13, 14, 22, 23, 30 - 36	$-16.38$ 0.78	
	5	$-16.99$ 0.10			13	$-15.90$ 0.94			$15 - 21$	$-16.44$ 0.78	
	$\epsilon$	$-16.60$ 0.60			15	$-16.10$ 0.94			$24 - 29$	$-16.36$ 0.80	
	$\tau$	$-16.30$ 0.10			16	$-16.15$ 0.90		20	$14, 15, 22 - 37$	$-16.38$	0.78
	$8\,$	$-15.35$ 0.10			17	$-16.30$ 0.92			$16 - 21$	$-16.48$	0.78
	9	$-15.10$ 0.10			18	$-16.11$ 0.92		21	$15 - 23,31 - 38$	$-16.42$ 0.77	
	10	$-14.45$ 0.10		9	$\sqrt{5}$	$-15.25$ 0.90			$24 - 30$	$-16.38$ 0.78	
	11	$-14.10$ 0.10			$\sqrt{6}$	$-15.90$ 0.90		22	$16 - 39$	$-16.42$ 0.77	
	12	$-13.45$ 0.10			$\tau$	$-16.28$ 0.90		23	$17 - 40$	$-16.42$ 0.77	
	13	$-13.10$ 0.10			9	$-16.30$ 0.90		24	$18 - 25$	$-16.45$ 0.77	
	14	$-13.00$ 0.40			10	$-16.15$ 0.90			$26 - 41$	$-16.42$ 0.77	
6	$\mathbf{2}$	$-13.00$ 0.10			8,11,17,19,20	$-16.20$ 0.90		25	$19 - 26$	$-16.46$ 0.77	
	3	$-13.85$ 0.80			12	$-16.01$	0.90		$27 - 42$	$-16.42$ 0.77	
	$\overline{4}$	$-15.70$ 0.10			13	$-16.05$ 0.90		26	$19 - 43$	$-16.46$ 0.77	
	5,7	$-16.50$ 0.10			14	$-15.95$ 0.90		27	$21 - 28$	$-16.48$ 0.77	
	6	$-16.65$ 0.10			15,16,18	$-16.11$ 0.90			$29 - 45$	$-16.46$ 0.77	
	8	$-15.90$ 0.10		10	6	$-15.25$ 0.50		28	$22 - 48$	$-16.48$ 0.77	

TABLE I. Refitted bulk and asymmetry constants for Seeger's mass formula.

from Ref. [44]. Evidently, these constants need be refitted since a large amount of data has become available  $[31]$ , particularly for neutron-rich nuclei. We found that the measured binding energies could be fitted within  $1-1.5$  MeV by changing the bulk constant  $\alpha(0)$  and introducing a proton, neutron asymmetry constant  $a_a$ . The  $\alpha(0)$  works as an overall scaling factor and  $a_a$  controls the curvature of the experimental parabola (and hence helps to fit the binding energies for neutron-rich nuclei), as expected. Table I gives the new  $\alpha(0)$ and  $a_a$  constants for all the known nuclei with  $1 \le Z \le 28$ , relevant to the present problem. The kind of comparisons obtained between the experimental and calculated binding energies is already illustrated in Fig. 2.

The *T*-dependent constants in Eq. (A2) were obtained numerically by Davidson *et al.* [29] for the available experimental information on excited states of 313 nuclei in the mass region  $22 \leq A \leq 250$  by determining the partition function  $Z(T)$  of each nucleus in the canonical ensemble and making a least squares fit of the excitation energy,

$$
E_{ex}(T) = V_{\text{LDM}}(T) - V_{\text{LDM}}(T=0)
$$

to the ensemble average

$$
E_{ex}(T) = T^2 \frac{\partial}{\partial T} \ln \mathcal{Z}(T).
$$

The  $\alpha(T)$ ,  $\beta(T)$ ,  $\gamma(T)$ ,  $\eta(T)$ , and  $\delta(T)$  thus obtained are given in Fig. 1 of Ref. [29] for  $T \le 4$  MeV, extrapolated linearly for higher temperatures. For the bulk constant  $\alpha(T)$ , instead, an empirically fitted expression to a Fermi gas model is used, as

$$
\alpha(T) = \alpha(0) + \frac{T^2}{15}.
$$

Also, the  $\delta(T)$  is constrained to be positive definite at all temperatures, with  $\delta(T>2 \text{ MeV})=0$ . Finally, the analytical form for  $r_0(T)$ , taken from Ref. [45], is

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$$
r_0(T) = 1.07(1 + 0.01T).
$$

For the shell corrections  $\delta U$  in Eq. (A1), since there is no microscopic shell model known that gives the shell corrections for light nuclei, we use the empirical formula of Myers and Swiatecki  $[40]$ . For spherical shapes,

$$
\delta U = C \bigg[ \frac{F(N) + F(Z)}{(A/2)^{2/3}} - c A^{1/3} \bigg],
$$
 (A3)

where

$$
F(X) = \frac{3}{5} \left( \frac{M_i^{5/3} - M_{i-1}^{5/3}}{M_i - M_{i-1}} \right) (X - M_{i-1}) - \frac{3}{5} (X^{5/3} - M_{i-1}^{5/3}),
$$
\n(A4)

with  $X = N$  or *Z*,  $M_{i-1} < X < M_i$  and  $M_i$  as the magic numbers 2, 8, 14 (or 20), 28, 50, 82, 126, and 184 for both neutrons and protons. The constants are  $C = 5.8$  MeV and  $c$  $=0.26$ . In this paper, we refer to the use of magic numbers 14 or 20 as MS14 or MS20 parametrization.

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