High-spin states and the first signs of band termination in 156Ho

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High-spin states have been populated in 156 Ho with the $^{14}N+^{148}Nd$ reaction at 96 MeV. The known level scheme has been considerably extended using the Gammasphere spectrometer. Configuration assignments for these bands have been deduced from their aligned angular momenta, signature splittings and from a systematic survey of the neighboring nuclei and these are compared with theoretical $B(M1)/B(E2)$ predictions from a semiclassical model. Evidence is also presented for the onset of band termination, or the beginning of a prolate collective to oblate noncollective shape change near the yrast line, at the highest spin states in two of the observed strongly coupled bands. $[$0556-2813(98)01905-0]$

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I. INTRODUCTION

Historically odd-odd nuclei have always proven difficult to study. This is because of their high single-particle level density, at low excitation energy, which results in a multitude of near-yrast bands or configurations. As an example of these complexities, the ground-state spin of the odd-odd nucleus ¹⁵⁶Ho has been difficult to establish and remains uncertain. To date arguments have been presented for ground-state spin assignments of $I=1\hbar$ [1] and $I=4\hbar$ [2] from hyperfine interaction measurements and $I = 5\hbar$ from the electron-capture decay of ¹⁵⁶Er, [3] and also $I=9\hbar$ [4]. However, it is precisely the fact the odd-odd nuclei have been somewhat neglected in the past which makes their study, with the new generation of powerful γ -ray spectrometers, particularly attractive.

In the mass 150 region the study of odd-odd nuclei has been mainly restricted to low spins. These odd-odd studies have played an important role in demonstrating the shapedriving effects of high-*j* intruder orbits and that these nuclei are "soft" with respect to triaxial deformations [5]. In addition, odd-odd nuclei have served as useful references in assigning configurations in the neighboring nuclei.

This paper describes the results of an experiment which populated high-spin states in the odd-odd nucleus ${}^{156}_{67}$ Ho₈₉. The known level scheme has been considerably extended and evidence is presented for a change in structure, signified by the onset of band termination, at the highest spin states in two of the rotational bands. The transition to band termination is accompanied by a change in nuclear shape in the yrast states from a well-deformed prolate to a weakly-deformed oblate shape $[6,7]$. Such behavior might be expected for ¹⁵⁶Ho because some of the rotational bands in the neighboring isotopes ${}^{155}_{67}$ Ho [7] and ${}^{157}_{67}$ Ho [8] and isotones ${}^{155}_{66}$ Dy [9] and $^{157}_{68}$ Er [10], are observed to terminate in fully aligned configurations at large angular momentum $I \approx 30-40\hbar$.

II. EXPERIMENT

High-spin states were populated in ¹⁵⁶Ho with the 148 Nd(14 N,6*n*) reaction at 96 MeV. Stacked thin (≈ 800 μ g/cm²) targets of ¹⁴⁸Nd, enriched to 94.1%, were used. Triple- and higher-fold coincidence events were collected at a rate of \approx 3000 per second. A total of \approx 10⁹ unfolded triplecoincident events were collected with the thirty-six escapesuppressed germanium detectors of the Gammasphere array [11] at the Lawrence Berkeley Laboratory. The 36 detectors were arranged in three rings of five at each of the forward and backward angles of 17.3°, 31.7°, and 37.4° and the remaining six detectors were in a ring at 90° with respect to the beam direction.

The data were sorted into a three-dimensional matrix or "cube" and a triple-coincidence analysis was performed with the LEVIT8R code $[12]$. Double-gates were placed in this cube and the γ -ray intensities and coincidence relationships were used to determine the order of the γ rays in the level

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FIG. 1. (a) The level scheme for $156H\text{O}$ showing the new bands (bands 2a, 3a, 4, 5, 8a, and 8b) and the extensions to the known bands (bands 1a, 1b, 2b, 3b, 6a, 6b, 7a, and 7b). Note that the ground-state spin is uncertain (see text) and in the figure all spins are drawn relative to a ground-state spin which is assumed to be $4\hbar$ at zero excitation energy. (b) A partial level scheme for ¹⁵⁶Ho which shows the lower-spin region of bands 1a and 1b expanded for clarity.

scheme. In addition, a two-dimensional matrix was constructed in which all events in coincidence with one of the six 90 ° detectors were incremented on one of the axes. The other axis contained those events in coincidence with any other non-90 ° detector. The angular correlation ratios, *R* $= I(for)/I(90)$ extracted from these matrices, were used to determine the multipolarity of the γ rays using the directional correlation from orientated states (DCO) method $[13]$.

III. RESULTS

The new level scheme for 156 Ho, extracted from these data, is shown in Figs. $1(a)$ and $1(b)$. In the present study the known $\left[4,7,14-16\right]$ yrast rotational bands in 156 Ho (bands 1a and 1b) have been extended and six new bands (bands 2a, $3b, 4, 5, 8a$ and $8b)$ have been established. The average intensity of these bands, relative to the yrast band 1b, are given in Table I. In the introduction it was pointed out that the ground-state spin of 156Ho has not been definitely established. In this paper the $I=4\hbar$ assignment [2] will be used throughout; however, all spins and parities are uncertain and are shown in parentheses in Figs. $1(a)$ and $1(b)$. The DCO ratios for some of the most intense γ -ray transitions are given in Table II. With these data it was not always possible to obtain DCO ratios for the highest-spin transitions in any of these bands due to their small intensities. In this case those in-band γ -ray transitions were generally assumed to be $\Delta I = 2$ transitions if their energies were reasonably smooth extensions of the lower-spin collective rotational states.

The yrast band 1a has been extended $|7|$ from spin (24) h to spin (36) h and its signature-partner band 1b has been extended from spin $(29)\hbar$ [7] to spin $(35)\hbar$; see Fig. 1(a). The DCO ratios for the transitions in bands 1a and 1b are consistent with those of other known $\Delta I=2$ quadrupole transitions in 156Ho; see Table II. Seven new interband transitions between bands 1a and 1b were also established; see

TABLE I. The average intensities of the bands observed in ¹⁵⁶Ho normalized to band 1b at low rotational frequencies.

Band	% Intensity signature (0) a	% Intensity signature $(+1)$ b
1	51.2 ± 1.5	100 ± 1.9
2	8.5 ± 0.3	8.5 ± 0.4
3	19.0 ± 1.1	21.9 ± 1.1
4	4.8 ± 0.4	.
5	2.8 ± 0.3	.
6	11.6 ± 0.5	14.5 ± 0.6
7	14.5 ± 0.7	14.7 ± 0.7
8	6.4 ± 0.3	4.1 ± 0.3

TABLE II. The DCO ratios for bands 1, 2, 3, and 7 in 156 Ho. The symbol <1.0 indicates that the transition was more intense in the 90° relative to the forward-spectrum. DCO ratios were not calculated for some transitions since they were too weak and/or heavily contaminated by other γ rays with similar energies. (The weakest intensity transitions were generally only assigned in the level scheme after a series of double-gated spectra were summed.)

E_{γ}	DCO ratio (Band 1a)	Assignment	E_{γ}	DCO ratio (Band 1a)	Assignment
91	1.02 ± 0.19	$E2\,$	710	0.52 ± 0.04	Dipole $(E1)$
130	0.57 ± 0.05	$\Delta I = 1, M1/E2$	861	0.45 ± 0.03	Dipole $(E1)$
264	0.60 ± 0.04	$\Delta I = 1, M1/E2$	952	< 1.0	Dipole
300	0.55 ± 0.01	$\Delta I = 1$, $M1/E2$	E_{γ}	DCO ratio (Band 3a)	Assignment
327	0.61 ± 0.04	$\Delta I = 1$, $M1/E2$	195	0.66 ± 0.10	$\Delta I = 1, M1/E2$
394	0.93 ± 0.08	E2	239	0.57 ± 0.05	$\Delta I = 1$, $M1/E2$
512	0.95 ± 0.06	$E2\,$	382/383	$0.98 + 0.08$	E2
610	1.04 ± 0.09	$E2\,$	462	1.03 ± 0.07	$E2\,$
693	0.92 ± 0.11	$E2$	532	1.00 ± 0.04	$E2\,$
758	1.08 ± 0.12	$E2$	602	1.01 ± 0.05	$E2\,$
802	0.90 ± 0.12	$E2$	676	0.87 ± 0.06	$E2\,$
E_{γ}	DCO ratio (Band 1b)	Assignment	752	0.97 ± 0.07	$E2\,$
134	0.67 ± 0.03	$\Delta I = 1, M1/E2$	856	0.83 ± 0.09	$E2\,$
211/212	0.58 ± 0.01	$\Delta I = 1$, $M1/E2$	931	0.97 ± 0.14	$E2\,$
284	$0.59 + 0.06$	$\Delta I = 1, M1/E2$	956	1.20 ± 0.28	$E2\,$
340	$0.96\!\pm\!0.06$	E2	130	0.45 ± 0.05	Dipole
476	0.98 ± 0.01	$E2\,$	340	0.59 ± 0.06	$\Delta I = 1$ (<i>M</i> 1)
583	1.03 ± 0.07	$E2\,$	582	0.90 ± 0.10	$\Delta I = 0$ (<i>M</i> 1)
671	1.00 ± 0.13	$E2\,$	844	0.62 ± 0.04	$\Delta I = 1$ (<i>M</i> 1)
744	0.97 ± 0.07	$E2\,$	724	0.69 ± 0.06	$\Delta I = 1$ (<i>M</i> 1)
797	0.98 ± 0.19	$E2\,$	E_{γ}	DCO ratio (Band 3b)	Assignment
820	1.06 ± 0.25	$E2$	223	0.66 ± 0.07	$\Delta I = 1, M1/E2$
E_{γ}	DCO ratio (Band 2a)	Assignment	263	0.51 ± 0.03	$\Delta I = 1, M1/E2$
290	0.59 ± 0.10	$\Delta I = 1$, $M1/E2$	418	1.04 ± 0.13	E2
316	0.47 ± 0.07	$\Delta I = 1, M1/E2$	502	0.93 ± 0.11	$E2\,$
356	0.48 ± 0.17	$\Delta I = 1, M1/E2$	568	1.12 ± 0.08	E2
512	0.92 ± 0.11	E2	631	0.89 ± 0.03	E2
568	0.94 ± 0.19	$E2\,$	699	0.91 ± 0.07	$E2\,$
649	1.45 ± 0.28	$E2\,$	772	0.96 ± 0.07	$E2\,$
726	0.87 ± 0.10	$E2\,$	529	1.19 ± 0.23	$\Delta I = 0$
800	0.61 ± 0.07	$E2\,$			(E2/M1)
879	1.00 ± 0.17	$E2\,$	606	0.94 ± 0.11	$\Delta I = 0$
927	1.07 ± 0.27	$E2\,$			(E2/M1)
807	0.49 ± 0.08	Dipole $(E1)$	737	1.15 ± 0.10	M1
966	0.55 ± 0.19	Dipole $(E1)$	847	0.84 ± 0.08	M1
E_γ	DCO ratio (Band 2b)	Assignment	901	0.36 ± 0.04	$\cal M$ 1
252	0.54 ± 0.07	$\Delta I = 1$, $M1/E2$	E_{γ}	DCO ratio (Band 7a)	Assignment
293	0.52 ± 0.08	$\Delta I = 1, M1/E2$	228	0.40 ± 0.02	M1
541	0.91 ± 0.16	E2	572	0.91 ± 0.07	E2
608	0.95 ± 0.10	$E2\,$	619	1.16 ± 0.10	$E2\,$
692	0.91 ± 0.11	$E2\,$	E_{γ}	DCO ratio (Band 7b)	Assignment
776	0.97 ± 0.14	$E2\,$	606	1.09 ± 0.05	E ₂
868	0.87 ± 0.12	$E2\,$	676	1.21 ± 0.09	E ₂
935	1.02 ± 0.13	$E2\,$	800	1.31 ± 0.18	$E2\,$

Figs. $1(a)$ and $2(a)$. These transitions have DCO ratios which are consistent with those of other known $\Delta I = 1$ mixed $M1/E2$ transitions in ¹⁵⁶Ho.

Four members of band 2b were previously reported in Ref. [7]. However, in that study band 2b was proposed to directly decay into the yrast states. The present data disagree with the previous decay pattern in that band 2b is connected to the yrast band 1a by four new transitions $(561, 710, 861)$ and 952 keV); see Fig. $1(a)$. Band 2b has also been extended by two transitions $(999 \text{ and } 958 \text{ keV})$ and nine members of its signature-partner rotational band, band 2a, have been established; see Fig. $2(b)$. Band $2a$ is also connected to the yrast bands 1a and 1b at spins (17) and (19) by two new transitions (at 966 and 807 keV). Nine interband transitions

FIG. 2. (a) A sum of 6 double-gates in band 1b (the 797 keV with the 134- to 744-keV transitions). (b) A sum of 36 double-gated spectra for band 2a (the sum includes all the in-band transitions double gated with the 512-keV transition). (c) A sum of 21 double-gated spectra for band 3b from all of the combinations of the 502- to 913-keV transitions). (d) A sum of all of the doubled-gated combinations of the 624- to 1007-keV transitions in band 4. (e) A sum of all of the doubled-gated combinations of the in-band transitions in band 5. (f) A double-gated spectrum (the 162- with the 163-keV transition) on band 6b. (g) A double-gated spectrum (the 534- with the 534-keV transition) on band 7b.

were also observed to link these signature-partner bands. These interband transitions have DCO ratios which are consistent with those of other $\Delta I = 1$ mixed *M*1/*E*2 transitions and are shown in Table II.

Five members of band 3a were also previously reported in Ref. $|7|$ and this band has been extended by five new stretched $E2$ transitions. In addition the lowest spin γ -ray transition in band 3b (502.5 keV) was previously reported in Ref. [15]; however, the present data disagree with the two higher in-band transitions which were reported at 607.6 and 725.0 keV and also with its decay into the yrast band 1a (via a 381.9-keV transition). These two proposed in-band transitions in Ref. $[15]$ are in fact established, in this work, as transitions which link this band into the yrast states. Band 3b is connected to the yrast bands 1a and 1b by seven new transitions and band 3a by eleven new transitions; see Figs. $1(a)$ and $2(c)$. In addition the eight new members of band 3b are shown in Fig. $1(a)$. Bands 3a and 3b are linked by thirteen new interband transitions. These interband transitions have DCO ratios which are consistent with those of other $\Delta I = 1$ mixed *M*1/*E*2 transitions; see Table II.

The newly established bands 4 and 5 display different behavior to the other rotational bands in ¹⁵⁶Ho in that only one signature is observed; see Figs. $1(a)$, $2(d)$ and $2(e)$. The decoupled band 5 decays to the yrast bands 1a and 1b around spin (21) to $(22)\hbar$; see Fig. 2(e). Band 5 is observed to be in coincidence with the yrast band up to the 744- and 758-keV transitions and therefore, the lowest observed spin in band 5 must be $\geq (21-22)\hbar$. However, the discrete γ -ray linking transitions could not be established. The decoupled band 4 is composed of a new sequence of seven $\Delta I = 2$ transitions; see Fig. $1(a)$. Band 4 is observed to be in coincidence with the yrast band up to the 758- and 802-keV transitions at spin $(22-24)\hbar$ and the lowest observed spin in band 4 must be $\geq (24)\hbar$. However, like band 5, the discrete linking transitions could not be established.

The strongly coupled bands 6a and 6b were first reported in Ref. 7. In the present work bands 6a and 6b have been extended with one new transition in band 6a (679 keV) and two new transitions in band 6b $(754 \text{ and } 789 \text{ keV})$ and by one dipole connecting transition (345 keV) . It was not possible to link this band into the yrast states because its intensity is weak and fragmented. However, the transitions in the band are observed in coincidence with the 81- and 96-keV transitions of the low-spin states; see Fig. $2(f)$.

Band 7b was previously established by Waddington *et al.* $[14]$ where four states in one signature of band 7b and one in the other signature band 7a were observed. The 228.2-keV transition which was reported to link band 7b to the yrast states is in conflict with our data which show that the (12^-) to (11^{-}) 211-keV transition is not in coincidence with the 228-keV transition; see Fig. $2(g)$. The transitions which link bands 7a and 7b into the yrast states were not observed in these data probably due to their low intensity and/or fragmented decay path. After the third in-band transition bands 7a and 7b backbend and two new sequences have been established to high spin; see Fig. $1(a)$. In addition, at the same rotational frequency another two signatures, bands 8a and 8b, are observed to decay into bands 7a and 7b, respectively. The DCO ratios for some of these transitions are given in Table II.

IV. DISCUSSION

In order to establish configurations for the newly observed rotational bands and to confirm the configurations of the existing bands in 156 Ho a systematic survey of the neighboring nuclei has been performed. Cranked-shell model (CSM) and total Routhian surface (TRS) calculations, based on those described in $[17]$, predict that the yrast configuration has a deformation which is characterized by a prolate shape $(\beta_2=0.21, \beta_4=0.02, \text{ and } \gamma \approx 0^0)$ at low rotational frequencies ($\hbar \omega$ < 0.4 MeV) which becomes slightly γ soft at higher rotational frequencies. Table III summarizes the labeling convention used in this work for the lowest orbitals around

TABLE III. The convention for labeling the lowest orbitals around the neutron $N=89$ and the proton $Z=67$ Fermi surfaces. The integer *n* refers to the *n*th orbit of a particular parity π and signature α .

	Asymptotic Nilsson neutron state	
Label	$(\hbar \omega = 0)$	$(\pi,\alpha)_n$
A	$i_{13/2}$ 660 1/2	$(+, +1/2)$ ₁
B	$i_{13/2}$ [660]1/2	$(+,-1/2)$ ₁
C	$i_{13/2}$ [651] 3/2	$(+, +1/2)$,
D	$i_{13/2}$ [651] 3/2	$(+,-1/2)$,
E	$h_{9/2}$ [521] 3/2	$(-, +1/2)$
F	$h_{9/2}$ [521] 3/2	$(-,-1/2)$
X	$h_{11/2}$ [505]11/2	$(-, +1/2)$
Y	$h_{11/2}$ [505]11/2	$(-,-1/2)$
Label	Asymptotic Nilsson proton state	(π,α)
	$(\hbar \omega = 0)$	
A_p	$h_{11/2}$ [523]7/2	$(-,-1/2)$
B_p	$h_{11/2}$ [523]7/2	$(-, +1/2)$
E_p	$g_{7/2}$ 404 7/2	$(+,-1/2)$ ₁
F_p	$g_{7/2}$ [404]7/2	$(+, +1/2)$ ₁
X_{n}	$h_{9/2}$ 541 1/2	$(+,-1/2)$,
Y_p	$h_{9/2}$ 541 1/2	$(+, +1/2)$,

the $N=89$ and $Z=67$ Fermi surfaces.

In order to check the validity of the proposed configurations, from the aligned angular momentum and the systematics of the region, the experimental $B(M1)/B(E2)$ ratios have been compared with those predicted from theoretical calculations. The intensity branching ratios from the data, λ , were used to calculate the experimental $B(M1)/B(E2)$ ratios using the standard relations

$$
\frac{B(M1;I \to I - 1)}{B(E2;I \to I - 2)} = 0.693 \frac{E_{\gamma}^{5}(I \to I - 2)}{E_{\gamma}^{3}(I \to I - 1)} \frac{1}{\lambda(1 + \delta^{2})}
$$
(1)

in units of (μ_N/e) b)².

The theoretical calculations were based on the geometric model of Dönau and Frauendorf $[18]$. In this model the ratio of reduced transition probabilities for the $\Delta I=1$ to $\Delta I=2$ transitions is calculated with

$$
\frac{B(M1:I \to I-1)}{B(E2:I \to I-2)} = \frac{12}{5Q_0^2 \cos^2(\gamma + 30^\circ)} \left[1 - \frac{K^2}{1 - (\frac{1}{2})^2} \right]^{-2}
$$

$$
\times \left\{ \left(1 - \frac{K^2}{I^2} \right)^{1/2} \left[K_1(g_1 - g_R) \times \left(1 \pm \frac{\Delta e'}{\hbar \omega} \right) + \sum_n K_n(g_n - g_R) \right] - \frac{K}{I} \left[(g_1 - g_R) i_i + \sum_n (g_n - g_R) i_n \right] \right\}^2,
$$
\n(2)

where

$$
K = K_1 + K_2 + K_3 + \cdots. \tag{3}
$$

TABLE IV. Parameters used in the theoretical $B(M1)/B(E2)$ ratio calculation from the geometric model; see text for details. Additionally the quadrupole moment was taken to be 5.5 e b, γ $=1^\circ$, and $g_R = Z/A$.

Neutron state	Shell	g factor	alignment, i	Κ
A,B $\lceil 660 \rceil 1/2$	$i_{13/2}$	-0.177	5.5	1/2
E,F [521]3/2	$f_{7/2}$	-0.328	2.5	3/2
X, Y [505] 11/2	$h_{11/2}$	-0.209	1.0	11/2
Proton orbit	Shell	g factor	alignment, i	K
A_p , B_p [523]7/2	$h_{11/2}$	$+1.214$	2.5	7/2
E_p , F_p [404]7/2	87/2	$+0.739$	1.0	7/2

The subscripts *n* refer to the quasiparticles, or aligned pairs of quasiparticles, that couple to form the band. Δe^{\prime} is the signature splitting of the level energies in the rotating frame. Quasiparticle 1 is deemed to be the particle responsible for the signature splitting. In the calculation the signaturesplitting term was set to zero so that just the average theoretical $B(M1)/B(E2)$ ratio was determined. In addition the *B*(*E*2) value was calculated assuming a quadrupole moment of 5.5 *e* b from a rough average of the measured values in 157 Ho [19]. Table IV shows the values of the parameters that were used in the calculation.

The actual experimental signature splitting between the states of the strongly coupled bands in 156Ho, observed in these data, was estimated according to

$$
\Delta E_{split} = E(I) - E(I - 1)
$$

$$
- \frac{1}{2} [E(I + 1) - E(I) + E(I - 1) - E(I - 2)], \quad (4)
$$

where $E(I)$ is the excitation energy (in keV) of the state with spin $I\hbar$.

A. Configurations of the bands

The high-spin systematic behavior of the nuclei around 156Ho is dominated by the presence of the neutron intruder $i_{13/2}$ and proton $h_{11/2}$ orbitals. These high-*j* orbitals lie close to the Fermi surface at zero rotational frequency and decrease rapidly in energy as the rotational frequency increases. Bands 1a and 1b have the lowest excitation energy at low spin and are expected to be built upon the lowest neutron configuration, A, $i_{13/2}$ [660] $1/2$ ⁺ coupled with both signatures, A_p and B_p , of the the lowest proton configuration, $h_{11/2}$ [523]7/2⁻. Figure 3 shows the experimental aligned angular momentum, *i*, versus rotational frequency extracted from the data, using the prescription of Ref. $[20]$, for all of the bands observed in 156 Ho and the yrast band in 156 Dy [21] for comparison. A reference term, representing the contribution from the collective rotation of the nucleus, has been subtracted according to Harris [22] with parameters, \mathcal{I}_0 =32.1 MeV⁻¹ \hbar^2 and \mathcal{I}_1 =34.0 MeV⁻³ \hbar^4 from Refs. [21,23]. The 156 Dy data show the back-bending effect of the first (AB) alignment ($\Delta i = 10\hbar$) of a pair of $i_{13/2}$ neutrons near $\hbar \omega \approx 0.28$ MeV. In the neighboring nuclei the AB

FIG. 3. (a) The experimental aligned angular momentum as a function of rotational frequency for bands 1, 2 and 3 in 156Ho and the $i_{13/2}$ band in ¹⁵⁶Dy for comparison. (b) The experimental aligned angular momentum for bands 4, 5, 6, 7, and 8 in 156 Ho. The Harris parameters \mathcal{I}_0 =32.1 \hbar^2 /MeV and \mathcal{I}_1 =34 \hbar^4 /MeV³ were used as a reference band.

band crossing occurs at $\hbar \omega \approx 0.27$ MeV in ¹⁵⁷Ho [8], at $\hbar \omega \approx 0.3$ MeV in ¹⁵⁵Ho [7,24] and at $\hbar \omega \approx 0.28$ MeV in 158 Er [25].

1. Bands 1a and 1b

Figure 3(a) shows that neither signature of the $156Ho$ yrast bands 1a or 1b undergo the first (AB) neutron band crossing. This nonobservation of the AB crossing in $156Ho$ is consistent with the configurations of bands 1a and 1b involving either the A or the B orbital since the AB crossing would then be blocked. The value of the aligned angular momentum ($\approx 5\hbar$) of these bands at low frequency, $\hbar \omega \approx 0.1$ –0.2 MeV, rules out the possibility that both A and B are occupied $(i_A + i_B = 10\hbar)$. Figure 3(a) also shows that bands 1a and 1b undergo a crossing at $\hbar \omega \approx 0.40$ MeV with an aligned angular momentum gain $\Delta i = 5.0\hbar$. The behavior of this crossing is consistent with that expected for a BC crossing [26]. In this case the configuration for the yrast bands can be built upon neither the B nor the C orbitals. $\text{(In }^{157}\text{Ho}\ [8]$ the BC crossing occurs at a rotational frequency of $\hbar \omega \approx 0.37$ MeV with an aligned angular momentum gain of Δi \approx 5.9 \hbar .) Therefore, we assign the initial low-spin neutron configuration for bands 1a and 1b as the A orbit.

Employing similar blocking arguments for the proton configurations then the nonobservation of the first proton band crossing, $A_p B_p$, which takes place at ≈ 0.43 MeV in 157 Ho and the neighboring nuclei [21,23,27], can be understood if the configuration of bands 1a and 1b are built upon the lowest A_p and B_p $h_{11/2}$ proton [523]7/2 orbitals. In conclusion all of these observations are consistent with bands 1a

FIG. 4. Signature splitting as a function of spin (see text) for (a) bands 1 (filled squares) and 2 (open diamonds) based on the $[523]5/2$ orbit in ¹⁵⁶Ho, (b) $[523]7/2$ bands in ¹⁵⁸Ho (open squares) and 160 Ho, (filled squares) (c) bands 3 (filled triangles) and 6 (open diamonds) 156 Ho, (d) [404]7/2 bands (filled squares) in 157 Ho, (e) bands 7 (filled triangles) based on $\lceil 505 \rceil$ 11/2 configuration in ¹⁵⁶Ho, (f) bands based on $[505]11/2$ orbital on 155 Dy (filled squares) and 156 Dy (open squares). Note the ordinate scale change between the first, second and third panels.

and 1b having the AA_p and AB_p configurations, respectively.

The experimental signature splitting for bands 1a and 1b, which is governed by the lowest A_p and B_p $h_{11/2}$ [523]7/2 proton orbitals, is shown by the filled squares in Fig. $4(a)$. Also shown in Fig. $4(b)$ is the signature splitting of the corresponding bands, built upon similar configurations, in the neighboring nuclei, 158 Ho (open squares) [28,29] and in 160 Ho (filled squares) [30]. From Figs. 4(a) and 4(b) it can be observed that the signature splitting between bands 1a and 1b and that in the $h_{11/2}$ [523]7/2 bands in ¹⁵⁸Ho and ¹⁶⁰Ho are similar above spin $20\hbar$. However, one difference should be pointed out: the signature splitting at low spin in 156Ho is larger than that of 158 Ho and at high spin this signature splitting is comparable. The open diamonds in Fig. $4(a)$ will be discussed in Sec. IV A 2.

The experimental $B(M1)/B(E2)$ ratios for bands 1a and 1b, Fig. $5(a)$, exhibit a regular staggering between the states of odd and even spin and have an experimental $B(M1)/B(E2)$ ratio of about 1.1μ ^N e b² at spin (14) which reduces to about $0.5\mu_N / e b^2$ at spin $(24)\hbar$. The theoretical values (solid lines) based on the AA _p proposed configurations (assuming zero signature splitting) qualitatively reproduce this trend. In summary bands 1a and 1b are well described by the proposed AA_p and AB_p configurations assignments. These configuration assignments and those for the other bands in 156Ho are summarized in Table V.

2. Bands 2a and 2b

Figure $3(a)$ shows that bands 2a and 2b have a larger experimental aligned angular momentum at low rotational

FIG. 5. Experimental and theoretical $B(M1)/B(E2)$ ratios as a function of spin for bands 1a, 1b, 2a, 2b, 3a, 3b, 6a, 6b, 7a, and 7b in 156Ho. The solid line represents the theoretical calculated from the semiclassical model of Dönau and Frauendorf based on the proposed configurations; see text for details. Filled symbols refer to signature 0 states and open symbols refer to signature 1 states. In the figure, the uncertainties result from the uncertainty in the gamma-ray intensities and do not include those which result from the choice of parameters employed in the calculation.

frequencies than that of the yrast bands 1a and 1b. We argue that this is because bands 2a and 2b are based upon a configuration involving the aligned pair of $i_{13/2}$ neutrons, AB. Notice how the curve for bands 2a and 2b seem to follow on from the 156 Dy points after the back-bending point at 0.25 MeV $[21]$. Further, if bands 2a and 2b involve both the A and the B orbitals, then neither the BC nor the AD crossings would be expected to occur. This is consistent with the observation that no crossing takes place around $\hbar \omega \approx 0.40$ MeV [8]; see Fig. 3(a). From a study of the neighboring nuclei and the CSM predictions the next lowest neutron orbital that this band could be built upon is the negative parity E, $h_{9/2}$ [521] $3/2$ orbit.

TABLE V. Summary of the configuration assignments for the rotational bands observed in 156Ho.

Band	Configuration
1a	$AA_n \rightarrow ABCA_n$
1b	$AB_p \rightarrow ABCB_p$
2a	$EABA_p \rightarrow Band$ termination
2 _b	$EABB_p \rightarrow Band$ termination
3a	$AE_p \rightarrow$ Band termination
3b	$AF_p \rightarrow$ Band termination
4	$X_pA \rightarrow X_pABC$
4	(alternative configuration $X_pA \rightarrow X_pAA_pB_p$)
5	X_p EAB
бa	BF_p
6b	BE_n
7a	$XA_p \rightarrow XABA_p$
7b	$YA_p \rightarrow YABA_p$
8a	$XA_p \rightarrow XBCA_p$
8b	$YA_p \rightarrow YBCA_p$

The second gain in aligned angular momentum which occurs in this band at $\hbar \omega \approx 0.45$ MeV in Fig. 3(a) cannot be fully reconciled with that expected from the first pair of protons $A_p B_p$. This is because, in this region, this alignment usually occurs over a large range in rotational frequency as a consequence of its large interaction strength $[21,23,27]$ as illustrated by the ¹⁵⁶Dy yrast band in Fig. 3(a) from $\hbar \omega$ \approx 0.35-0.55 MeV. In addition, Fig. 3(a) also shows that the gain in aligned angular momentum, $\Delta i \approx 5-6\hbar$, continues to rise even at the highest observed rotational frequencies. The nature of this anomalous aligned angular momentum gain will be discussed in the Sec. IV B. In summary, bands 2a and 2b are interpreted as being built upon a configuration which has the EAB orbitals coupled to the proton A_p and B_p orbitals. These bands are, therefore, based upon the ABEA *^p* and ABEB*^p* configurations.

The signature splitting in bands 2a and 2b is governed by that of the proton $h_{11/2}$ [523]7/2 orbital which is expected to be similar in magnitude to that shown by bands 1a and 1b. Figure $4(a)$ shows this similarity and Fig. $4(b)$ the similarity with bands based on the π [523]7/2 orbital in the neighboring nuclei 158Ho and 160Ho. Interestingly, signature inversion occurs in both bands 1a and 1b and bands 2a and 2b but at a different spin. The theoretical $B(M1)/B(E2)$ ratios for bands 2a and 2b based on the EBAA *^p* configuration are also in qualitative agreement with these data; see Fig. $5(b)$. From this comparison with the neighboring nuclei and bands 1a and 1b it can be deduced that there is good agreement with the proposed configuration for bands 2a and 2b.

3. Bands 3a and 3b

Bands 3a and 3b do not show any evidence for an AB neutron band crossing around $\hbar \omega \approx 0.25-0.30$ MeV; see Fig. $3(a)$. This implies that the AB configuration is blocked for these bands and that they are built upon either the A or the B orbital. The possibility for both A and B being occupied is ruled out since the low frequency aligned angular momentum is small ($\approx 5\hbar$). It can also be observed that bands 3a and 3b appear to undergo a smooth increase in aligned angular momentum from 0.25 to 0.4 MeV; see Fig. $3(a)$. This increase could be associated with the BC crossing which would imply that bands 3a and 3b are built upon the neutron A $i_{13/2}$ [660] $1/2$ configuration. Alternatively this increase could arise from the AD crossing which would imply that bands 3a and 3b are built upon the neutron B $i_{13/2}$ [660]1/2 configuration. However, in this mass region the $A_p B_p$ proton crossing also takes place around these rotational frequencies [see the smooth aligned angular momentum increase in the 156 Dy yrast band in Fig. 3(a)] which is similar to the behavior of bands 3a and 3b over the rotational frequency range $(0.25 \text{ to } 0.4 \text{ MeV})$. Furthermore, the nonobservation of this interaction in bands 1a and 1b and bands 2a and 2b is consistent because both of their configuration assignments contain either the A_p or B_p quasiparticle. These aligned angular momentum properties are consistent with the bands having E_p and F_p proton $g_{7/2}$ [404]7/2 configurations.

Another gain in aligned angular momentum occurs at 0.45 MeV for bands 3a and 3b; see Fig. $3(a)$. This gain is at least \approx 5–6 \hbar and continues to rise even at the highest observed rotational frequencies. This alignment gain is similar to that observed in bands 2 and will be discussed in Sec. IV B.

Figure $4(c)$ shows the signature splitting for bands 3a and 3b; notice the scale change between Figs. $4(a)$ and $4(c)$. Also shown in the figure is the signature splitting observed in a band based upon a similar proton $g_{7/2}$ [404]7/2 configuration in the neighboring nucleus 157 Ho [8], Fig. 4(d). Reasonable agreement in these nuclei with the absolute magnitude of the splitting is observed. The experimental $B(M1)/B(E2)$ ratios for bands 3a and 3b have an average ratio of about 0.5μ _N/*e* b² for all spins. The theoretical ratios based on the AE_p (F_p) configuration qualitatively reproduce this trend as a function of spin; see Fig. $5(c)$. In summary band 3a is interpreted as being built upon the AE*^p* and band 3b upon the AF_p configuration.

4. Band 4

The configuration assignment for band 4 has been deduced from its aligned angular momentum shown in Fig. $3(b)$ and the fact that it is a decoupled band which means it must be based upon an orbit which has large signature splitting. The lowest-lying proton orbital with a large signature splitting, from the systematics of the neighboring nuclei, is the X_p $h_{9/2}$ [541]1/2 orbit. A decoupled band with similar alignment properties has been observed in 157 Ho [8]. Since band 4 shows no evidence for the AB band crossing, its configuration must, therefore, be built upon either the A or the B orbital. However, at a rotational frequency of $\hbar \omega \approx 0.43$ MeV a smooth increase in aligned angular momentum is observed; see Fig. $3(b)$. This aligned angular momentum gain is most probably caused by the alignment of a pair of BC neutrons. Comparison with the BC crossing observed in the yrast bands 1a and 1b reveals that the total aligned angular momentum gain, $5\hbar$, is very similar. Therefore, band 4 is assigned an AX_p configuration at low rotational frequencies and an $ABCX$ _p configuration at higher rotational frequencies. However, there are alternative configuration assignments which could also explain the behavior of band 4. For example, the X_p $h_{9/2}$ [541]1/2 orbit could increase the deformation which may in turn reduce the frequency at which the BC crossing takes place. In this case the crossing observed at $\hbar \omega \approx 0.43$ MeV could be the A_pB_p crossing and band 4 might be based on the $AX_{p}A_{p}B_{p}$ configuration.

5. Band 5

Band 5 is also a decoupled band and exhibits very similar behavior to the bands 2a and 2b above the AB band crossing in that its initial aligned angular momentum at its lowest observed rotational frequency is large. Indeed, Fig. $3(b)$ indicates that band 5 shows the start of a decrease in aligned angular momentum at low rotational frequencies where the AB band crossing is expected and therefore, can neither be built upon the A nor the B orbital below this rotational frequency. The next lowest neutron orbit is the E, $f_{7/2}$ [521] $3/2$ orbital. At frequencies around $\hbar \omega = 0.46$ MeV band 5 gains aligned angular momentum which is interpreted as the start of the proton $A_p B_p$ band crossing. In this case band 5 cannot be built upon either the A_p or B_p configuration and must be built upon a proton configuration which has a large signature splitting since it is a decoupled band. The systematics of the neighboring nuclei suggest that this is most likely the proton $h_{11/2}$ [541] $1/2$ X_p orbital. Therefore, band 5 is assigned a four-quasiparticle $E(AB)X_p$ configuration above the AB band crossing. A band with similar properties has been observed in 157 Ho [8].

6. Bands 6a and 6b

Bands 6a and 6b do not show any evidence for the AB band crossing; see Fig. $3(b)$, which implies that the configuration of these bands involve either the A or B orbits being occupied. Since the low frequency aligned angular momentum is about $7\hbar$, bands 6a and 6b cannot have both A and B occupied. Figure $3(b)$ shows that a gain in aligned angular momentum occurs around 0.35 MeV which is consistent with the BC crossing. In this case bands 6a and 6b would be built upon the A orbital coupled with one of the lower proton orbitals, for example, the AE _{*p*} and AF _{*p*} configurations. However, since the more intensely populated bands 3a and 3b are already associated with the AE _{*p*} and AF _{*p*} configurations, the crossing at 0.35 MeV, in bands 6a and 6b, is more likely the AD crossing. In this case the configurations for bands 6a and 6b could be based on the BE_p and BF_p orbits and would be less intensely populated than bands 3a and 3b, which is consistent with these data; see Table I. Unfortunately bands 6a and 6b were not observed to sufficiently high rotational frequencies (≥ 0.45 MeV) to determine whether the first proton band crossing occurs.

The signature splitting between bands 6a and 6b is shown in Fig. $4(c)$ and is very similar in magnitude to that of bands 3a and 3b and to that of the bands based upon similar configurations in the neighboring nuclei $157Ho$ [8]. The experimental $B(M1)/B(E2)$ ratios for bands 6a and 6b are about 0.5μ _N/*e* b² for the entire spin range over which they are observed. The theoretical ratios (solid lines) based on the BE*^p* configuration are in good agreement with the experimental ratios extracted from these data; see Fig. $5(d)$.

7. Bands 7a, 7b and 8a, 8b

At low rotational frequencies the signature splitting between bands 7a and 7b is very small. This suggests that these strongly coupled bands are most likely built upon configurations which have large Ω quantum numbers such as the neutron X and Y $h_{11/2}$ [505]11/2 orbital coupled with the lowest proton orbit, A_p . Bands with similar properties have been established in the neighboring holmium isotopes, for example, in 157 Ho [8] and in 155 Ho [9]. At a rotational frequency of $\hbar \omega \approx 0.3$ MeV bands 7a and 7b undergo a backbend; see Fig. $3(b)$. This band crossing in interpreted as an AB crossing since the observed $10\hbar$ gain in alignment is only consistent with this crossing. Notice how the curves for bands 7a and 7b resemble the AB crossing observed in the 156 Dy yrast band; see Fig. 3(a) and also the behavior of bands 2a and 2b above $\hbar \omega \approx 0.25$ MeV. Above this rotational frequency ($\hbar \omega = 0.3$ MeV) bands 7a and 7b do not show evidence for any other neutron or proton band crossings, presumably because their underlying single-particle orbit configuration blocks all possible band crossings; see Fig. $3(b)$. At these high rotational frequencies bands 7a and 7b are interpreted as being based on XABA *^p* and YABA*^p* configurations.

At rotational frequencies around $\hbar \omega \approx 0.3$ two new sequences, bands 8a and 8b, are observed to feed into bands 7a and 7b; see Fig. $1(a)$. These bands are interpreted as the continuation of bands 7a and 7b above the AB crossing and are based on the XA _{*p*} and YA _{*p*} configurations. These bands undergo an alignment gain at $\hbar \omega = 0.33$ MeV which is interpreted as a BC crossing.

The signature splitting between bands 7a and 7b below the AB crossing is extremely small; see Fig. $4(e)$. This splitting is in good agreement with that of the bands involving the $[505]11/2$ orbit configuration in the neighboring nuclei 156 Dy [21] and 155 Dy [9] and is smaller than that observed in bands 1a and 1b $([523]7/2)$ and bands 3a and 3b $([404]7/2)$ since they are based upon the signature splitting of the strongly-coupled $[505]11/2$ orbit. Below the AB crossing bands 7a and 7b have $B(M1)/B(E2)$ ratios of about 0.6 μ_N/e b² at spin (14) \hbar which decreases to 0.1 μ_N/e b² at spin (18) \hbar . This trend is reasonably reproduced in the theoretical calculations based on the XA ^{*p*} configuration; see Fig. $5(e).$

Above the AB crossing the experimental signature splitting between bands 7a and 7b; see Fig. $4(e)$, are in good agreement with those of the bands based upon similar configurations in the neighboring 156 Dy [21] and 155 Dy [9], Fig. $4(f)$. These signature splittings are smaller than those observed in bands 1a and 1b and bands 3a and 3b since they are based upon the signature splitting of the strongly coupled $h_{11/2}$ [505]11/2 orbit. The experimental *B*(*M*1)/*B*(*E*2) ratios for bands 7a and 7b, above the AB crossing, have large uncertainties due to the low intensity and contamination of these bands. The theoretical ratios based on the XABA *^p* configuration are also presented in Fig. $5(f)$ for comparison.

B. Band termination

The excitation energies of the states in ¹⁵⁶Ho are plotted in Fig. 6 versus spin relative to a rigid-rotor reference. In the figure it can be observed that the highest-spin states in bands 2a and 2b and 3a and 3b begin to show the start of a downsloping trend compared with the excitation energy minus rigid-rotor energy above spin $30\hbar$. This trend has also been observed in the neighboring nuclei and was suggested to be characteristic of a nucleus which is undergoing a transition from collective to single-particle behavior with a corresponding nuclear shape change from weakly deformed prolate to a weakly deformed oblate [31]. When a nucleus in this mass region approaches band termination its angular momentum or spin is gained rather efficiently only requiring a relatively small amount of energy. Therefore, if the excitation energy of a rigid rotor reference is subtracted, the levels associated with the band terminating sequence should be down-sloping with respect to spin $[31]$.

The band termination scenario was first suggested for the rotational bands in 155 Ho and 156 Ho in Ref. [7]. The spins at which these bands terminate is determined by the number of valence particles outside the $^{146}_{64}$ Gd closed shell. For example, the dashed lines in Fig. 6 represent two of the rotational bands in $^{155}_{66}$ Dy which show behavior consistent with the approach to band termination $[9,32]$. The theoretical calculations for this nucleus predict that there should be a low-lying $\pi[(h_{11/2})^2] \nu[(i_{13/2})^2 (f_{7/2})^3 (h_{9/2})^2]$ configuration which terminates at spin $75/2 - \hbar$ [32]. In agreement with this scenario, Emling *et al.* [32] presented evidence that the

FIG. 6. Excitation energy of the states in bands 1, 2, and 3 relative to a rigid rotor. The EAB configuration which terminates in ¹⁵⁵Dy is shown with a dashed line for comparison. The solid line represents the ABC configuration in 155Dy which shows no evidence for band termination.

high-spin states in this band at spins $71/2$ ⁻ and $75/2$ ⁻ were long-lived with lifetimes greater than 1.5 ps.

In an analogous manner two terminating states with especially low excitation energies have been observed in ¹⁵⁷Ho [8] at spins $75/2$ ⁻ and $87/2$ ⁻ \hbar . The $87/2$ ⁻ state was described as a closed 146Gd core with eleven valence particles in a fully aligned

$$
\pi[(h_{11/2})^3]_{(27^-/2)} \otimes \nu[(i_{13/2})^2 (f_{7/2})^3 (h_{9/2})^3]_{30+}
$$

configuration. The $75/2$ ⁻ state has one $f_{7/2}$ neutron antialigned with a

$$
\pi[(h_{11/2})^3]_{(27^-/2)} \otimes \nu[(i_{13/2})^2 (f_{7/2})_4^4 + (h_{9/2})^2]_{24+}
$$

configuration. In 157 Ho the lifetime measurements described in Ref. $[19]$ were consistent with the interpretation of band termination because the high-spin states $(272/2)$ were observed to be long lived with lifetimes greater than 1.1 ps. Indeed an 8 ps lifetime was deduced for the $75/2$ ⁻ state, corresponding to a transition quadrupole moment of ≤ 1 *e* b which is consistent with the extremely weak collectivity expected for these band terminating states. The heavier Er nuclei show similar behavior, for example, in 157 Er [23] and 158 Er [33], many single-particle and fully aligned band terminating states are observed at spins around 40^+ and 46^+ \hbar .

Thus in 156Ho, which has 10 valence particles, low-lying $\pi(h_{11/2})^3 \otimes \nu[(i_{13/2})^2 (f_{7/2})^3 (h_{9/2})^2]$ configurations which terminate at spins 40^+ and 41^+ are expected. In $^{155}_{66}$ Dy [9] it was observed that the bands based upon the EAB and FAB configurations showed a decrease in excitation energy minus rigid-rotor value above spins $\approx 30\hbar$ where they are crossed by the terminating sequences; see the dashed lines in Fig. 6. In $^{156}_{67}$ Ho an analogous situation seems to occur when the

 $EABA_p(B_p)$ configuration of bands 2a and 2b start to downslope near spin 35 \hbar where they are presumably crossed by the $\pi(h_{11/2})^3 \otimes \nu[(i_{13/2})^2(f_{7/2})^3(h_{9/2})^2]$ configuration; see Fig. 6. If these down-sloping terminating sequences cut through the yrast line in this manner, then based on an extrapolation of the 155Dy data, it is expected that the higher-lying nonyrast $(+,0)$, $(+,1)$ bands are crossed at slightly lower spins; see Fig. 6. Such a scenario is consistent with the anomalous alignment behavior of bands 3a and 3b at high spin. Note also that the configurations involving the ABC (bands 1a and 1b) and $ABX(Y)$ (bands 8a and 8b) orbitals in 156Ho do not show any evidence for downsloping at the highest spins consistent with the behavior of bands based on similar configurations in 155 Dy; see the solid lines in Fig. 6.

Therefore, it appears that the most likely explanation for this anomalous gain in aligned angular momentum in bands 2a and 2b and 3a and 3b in 156 Ho is due to the fact that they are starting to undergo a change from collective to a singleparticle nature with a corresponding change in shape which affects the aligned angular momentum. However, with the statistics of the present data it was not possible to observe these bands up to sufficiently high spin to observe the actual band terminating or favored high-spin states in ¹⁵⁶Ho. Interestingly the observation of these favored states, at their specific angular momentum values, could be used to make strong arguments for the low-spin assignments of the states in this nucleus. Indeed, further data are required to probe the bands in 156Ho at higher spin and ultimately lifetime information on the higher-spin states should confirm the weak collectivity of these states and the existence of band termination.

V. SUMMARY

In summary new rotational bands have been established and the previous level scheme considerable extended in 156Ho with the Gammasphere spectrometer. Configurations for these bands have been deduced from their aligned angular momenta, signature splittings and from a systematic survey of the neighboring nuclei and these are compared with theoretical $B(M1)/B(E2)$ predictions from a semiclassical model. Good agreement is found with the proposed configurations. Evidence is presented for the approach to band termination, at the highest spin states in two of these strongly coupled bands which is consistent with the expectations from neighboring nuclei where this phenomenon has been established. However, detailed lifetime measurements are ultimately required to measure the reduction in the collectivity of these states.

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- [1] R. Neugart, E. Arnold, W. Borchers, W. Neu, G. Ulm, and K. Wendt, in *Proceedings of the 5th International Conference on Nuclei Far from Stability*, Rosseau Lake, Canada 1987, edited by I. S. Towner (American Institute of Physics, New York, 1988), p. 126.
- [2] G. D. Alkhazov, A. E. Barzakh, I. Ya. Chubukov, V. P. Denisov, V. S. Ivanov, N. B. Buyanov, V. N. Fedoseyev, V. S. Letokhov, V. I. Mishin, and S. K. Sekatsky, Nucl. Phys. **A504**, 549 (1989).
- [3] K. D. Schilling, L. Kaubler, W. Andrejtscheff, T. M. Muminov, V. G. Kalinnikov, N. Z. Marupov, F. R. May, and W. Seidel, Nucl. Phys. **A299**, 189 (1978).
- [4] R. G. Helmer, Nucl. Data Sheets **65**, 65 (1992); **49**, 383 $(1986).$
- [5] R. Bengtsson *et al.*, Nucl. Phys. **A415**, 189 (1984).
- [6] T. Bentgsson and I. Ragnarsson, Phys. Scr. **T5**, 165 (1983).
- [7] J. Simpson *et al.*, Proceedings of XXIII International Winter Meeting on Nuclear Physics, Bormio, Italy, 1985; R. V. F. Janssens (private communication).
- [8] D. C. Radaford *et al.*, Nucl. Phys. **A545**, 665 (1992).
- [9] R. Vlastou et al., Nucl. Phys. **A580**, 133 (1994).
- [10] S. J. Gale *et al.*, J. Phys. G 21, 193 (1995).
- $[11]$ I. Y. Lee, Nucl. Phys. $A520$, 641c (1990) .
- [12] D. C. Radford, Nucl. Instrum. Methods Phys. Res. A 361, 297 $(1995).$
- [13] K. S. Krane, R. M. Steffen, and R. M. Wheeler, Nucl. Data Tables 11, 351 (1973).
- [14] J. Waddington et al., Scientific/Technical Report No. 147, University of Bergen, Norway, 1984.
- [15] S. H. Bhatti *et al.*, Z. Phys. A **353**, 119 (1995).
- [16] G. Løvhøiden, Phys. Scr. 25, 459 (1982).
- [17] W. Nazarewicz, R. Wyss, and A. Johnson, Nucl. Phys. A503, 285 (1989).
- [18] F. Dönau and S. Frauendorf, in *Proceedings of the Conference*

on High Angular Momentum Properties of Nuclei, Oak Ridge, Tennessee, 1982, edited by N. R. Johnson (Harvard Academic, New York, 1983); F. Dönau, Nucl. Phys. A471, 469 (1987).

- [19] H. Gascon *et al.*, Nucl. Phys. **A513**, 344 (1990).
- [20] R. Bengtsson and S. Frauendorf, Nucl. Phys. A314, 27 (1979); A327, 139 (1979).
- [21] M. A. Riley *et al.*, Nucl. Phys. **A486**, 456 (1988).
- [22] S. Harris, Phys. Rev. B 138, 509 (1965).
- [23] J. Simpson *et al.*, J. Phys. G 12, L67 (1986).
- [24] G. B. Hagemann, B. Herskind, J. Kownacki, B. M. Nyako, P. J. Nolan, J. F. Sharpey-Schafer, and P. O. Tjom, Nucl. Phys. A424, 365 (1984).
- [25] J. Simpson, P. A. Butler, P. D. Forsyth, J. F. Sharpey-Schafer, J. D. Garrett, G. B. Hagemann, B. Herskind, and L. P. Ekström, J. Phys. G 10, 383 (1984).
- [26] J. D. Morrison, J. Simpson, M. A. Riley, H. W. Cranmer-Gordon, P. D. Forsyth, D. Howe, and J. F. Sharpey-Schafer, J. Phys. G 15, 1871 (1989).
- [27] M. A. Riley, J. Simpson, M. A. Bentley, P. Fallon, P. D. Forsyth, J. C. Lisle, J. D. Morrison, E. S. Paul, J. F. Sharpey-Schafer, and P. M. Walker, Z. Phys. A 345, 121 (1993).
- [28] N. Rizk and J. Boutet, J. Phys. (France) Lett. **37**, L197 (1976).
- [29] C.-H. Yu, D. M. Cullen, D. Cline, M. Simon, D. C. Radford, I. Y. Lee, A. O. Macchiavelli, in *Proceedings of the workshop on Gammasphere Physics*, Berkeley, California, 1995, edited by M. A. Deleplanque, I. Y. Lee, and A. O. Macchiavelli (World Scientific, Singapore, 1996), p. 254.
- [30] J. A. Pinston, S. Andre, D. Barneoud, C. Foin, J. Genevey, and H. Frisk, Phys. Lett. **137B**, 47 (1984).
- [31] I. Ragnarsson, Z. Xing, T. Bengtsson, and M. A. Riley, Phys. Scr. 34, 651 (1986).
- [32] H. Emling et al., Phys. Lett. B 217, 33 (1989).
- [33] J. Simpson *et al.*, Phys. Lett. B **327**, 187 (1994).