

## Lattice gas model with isospin-dependent interactions

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In this paper we continue the investigation of the lattice gas model. The main improvement is that we use two strengths for bonds: one between like particles and another between unlike particles to implement the isospin dependence of nuclear force. The main effect is the elimination of unphysical clusters, such as the dineutron or diproton. It is therefore a better description of a nuclear system. The equation of state in mean field theory is obtained for nuclear matter as well as for  $N \neq Z$  systems. Through numerical and analytical calculation we show that the new model maintains all the important features of the older model. We study the effect of the Coulomb interaction on multifragmentation of a compound system of  $A = 86$ ,  $Z = 40$ , and also for  $A = 197$ ,  $Z = 79$ . For the first case the Coulomb interaction has small effect. For the latter case the effect is much more pronounced but typical signatures of the lattice gas model such as a minimum (maximum) in the value of  $\tau$  ( $S_2$ ) are still obtained but at a much lower temperature. [S0556-2813(98)04204-6]

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### I. INTRODUCTION

Over the past few years we have been developing a lattice gas model for the study of nuclear multifragmentation [1–5]. In this model  $n$  nucleons are placed in  $N$  cubes and they interact with nearest neighbor interactions. In most of this work the interaction between neutron-neutron, proton-proton, and neutron-proton was taken to be identical although we have sometimes used [5] a more complicated model in which interactions between like particles are different from those between unlike particles. The purpose of this paper is to more fully expose this improved model and to examine its relationship with our earlier and simpler model. Among other things we will also show that the important conclusions reached with our simpler model go through in our improved version.

The lattice gas model has the attractive feature that it can provide an equation of state but it can also provide the cluster distribution. Most of the observables are calculated using Monte Carlo simulations. We used the Metropolis algorithm in our simulations. The  $n$  particles are distributed in  $N$  boxes according to the lattice gas Hamiltonian and their momenta are generated from a Maxwell-Boltzmann distribution at a prescribed temperature. The calculation of clusters is straightforward. Two nucleons in neighboring cells are part of the same cluster if the relative kinetic energy is less than the strength of the attractive bond:  $p_r^2/2\mu + \epsilon < 0$ . Here  $p_r$  is the relative momentum,  $\mu$  the reduced mass and  $\epsilon$  is the negative (attractive) interaction. This prescription is sufficient to calculate the cluster distribution. The motivation for introducing two kinds of bond is now obvious. If the neutron-neutron bond or the proton-proton bond is attractive then each numerical simulation can generate dineutrons and diprotons which in reality do not exist in nature as composites. One can get rid of these unphysical clusters by simply making the neutron-neutron and proton-proton bonds zero or repulsive.

For completeness we mention some more features of the

lattice gas model that were studied in [1,2]. One finds that at a certain temperature the distribution of composites is a power law:  $Y(Z) \propto Z^{-\tau}$ , where  $Y(Z)$  is the number (averaged over many simulations) of composites with  $Z$  protons. As is common practice, we will extract a value of  $\tau$  even when the distribution has significantly deviated from a power law [6]. This value is extracted by using

$$\frac{\sum_2^{10} ZY(Z)}{\sum_2^{10} Y(Z)} = \frac{\sum_2^{10} ZZ^{-\tau}}{\sum_2^{10} Z^{-\tau}}. \quad (1.1)$$

We also calculate the second moment  $S_2 = \sum' A^2 Y(A)/n$  where in the sum the largest cluster is excluded. The usefulness of  $S_2$  was emphasized by Campi [7]. Since the lattice gas model has a Hamiltonian, its average energy at any temperature can be calculated. The excitation energy and specific heat  $C_v$  per particle can be calculated. We will find this useful too.

### II. THE EQUATION OF STATE IN SIMPLE MEAN FIELD APPROXIMATION

We have  $N$  boxes in the lattice in which we have to put  $n_n$  neutrons and  $n_p$  protons,  $n_p + n_n = n < N$ ,  $n/N = V_0/V = \rho/\rho_0$ , where  $V_0$ ,  $\rho_0$  are normal nuclear volume and density, respectively. In principle, we can have three kinds of bonds:  $\epsilon_{nn}$ ,  $\epsilon_{pp}$ , and  $\epsilon_{pn}$ . In the most general case where these interactions can take any arbitrary value a rich assortment of phenomena is predicted. The grand canonical partition function of this general lattice gas model can be mapped on to a spin-1 Ising-type model in the presence of a magnetic and quadrupole field. These have been studied in detail in the past [8]. For the nuclear case the values of the interactions are quite restricted and the richness of phenomena disappears. First of all we have to set  $\epsilon_{nn}$  and  $\epsilon_{pp}$  to be either zero

or repulsive so that one avoids producing unphysical dineutron or diproton bound clusters. Charge independence of nuclear forces suggests that we put  $\epsilon_{nn} = \epsilon_{pp}$ . From now on we will write  $\epsilon_{pp}$  for both  $\epsilon_{pp}$  and  $\epsilon_{nn}$ . In our past work [5] and also in other modelings [9] of nuclear collisions using classical mechanics a slightly repulsive  $\epsilon_{pp}$  was used. To avoid proliferation of parameters we will set this bond to zero in this work. The binding energy of nuclear matter fixes the value of  $\epsilon_{pn}$  at  $-5.33$  MeV.

Throughout this work  $\gamma$  stands for the number of nearest neighbors. In three dimensions one has  $\gamma=6$ . We use the Bragg-Williams mean field theory using the canonical ensemble. There are  $N$  boxes and  $n_p$  protons and  $n_n$  neutrons. Let one of the boxes be occupied by a proton. Then, in the Bragg-Williams approximation, among its nearest neighbors, on the average  $\gamma n_p/N$  will be occupied by protons and  $\gamma n_n/N$  by neutrons. The number of  $n$ - $p$  bonds will be  $\gamma n_p n_n/N$ , the number of  $p$ - $p$  bonds will be  $(1/2)\gamma n_p n_p/N$ , where the factor of  $1/2$  remedies the double counting for proton-proton bonds. Similarly starting with a box occupied by a neutron we come up with the same number of neutron-proton bonds and the number of neutron-neutron bonds is determined to be  $(1/2)\gamma n_n n_n/N$ . Thus the interaction energy when there are  $n_p$  protons,  $n_n$  neutrons placed in  $N$  boxes is  $E = \gamma[\epsilon_{pn}n_p n_n + \epsilon_{nn}(n_n^2 + n_p^2)/2]/N$  and the partition function is

$$Z(N, n_p, n_n) = \frac{N!}{(N - n_p - n_n)! n_p! n_n!} \exp(-\beta E). \quad (2.1)$$

We now find pressure  $P$  from the equation  $P = kT[\partial \ln Z / \partial V]_T$  and  $V = a^3 N$ , where  $a^3 = 1/\rho_0$  is the volume of each box. Using Stirling's formula one arrives at

$$P = \rho_0 kT \ln \frac{N}{N - n} + \rho_0 \gamma \epsilon_{pn} (n_p/N)(n_n/N) + \rho_0 \gamma \epsilon_{nn} [(n_n/N)^2 + (n_p/N)^2]/2. \quad (2.2)$$

We introduce an asymmetry parameter  $\eta = (n_n - n_p)/(n_n + n_p)$  which takes value 1 for neutron matter, 0 for nuclear matter, and  $-1$  for proton matter. We can then write

$$P = \rho_0 kT \ln \frac{V}{V - V_0} + \frac{1}{2} \rho_0 \gamma \frac{V_0^2}{V^2} \left[ \frac{\epsilon_{pn} + \epsilon_{nn}}{2} + \frac{1}{2} \eta^2 (\epsilon_{nn} - \epsilon_{pn}) \right]. \quad (2.3)$$

We determine the critical point from  $\partial P / \partial \rho = \partial^2 P / \partial \rho^2 = 0$ . This gives the critical density  $\rho_c = 0.5\rho_0$  and the critical temperature  $-(\gamma/4)[(\epsilon_{pn} + \epsilon_{pp})/2 + (1/2)\eta^2(\epsilon_{nn} - \epsilon_{pn})]$ . In this approximation with  $\epsilon_{pn}$  attractive and  $\epsilon_{nn} = 0$  the critical temperature for nuclear matter (which has  $\eta=0$ ) would be highest at  $-(\gamma/4)(\epsilon_{pn}/2)$  and would fall off quadratically with  $\eta$  to 0 at neutron or proton matter.

The Bragg-Williams approximation is the simplest mean-field approximation. Unfortunately it is rather inaccurate for two kinds of bonds. For nuclear matter at normal density ( $n_p = n_n = N/2$ ) it predicts a huge difference in binding energy for the case  $\epsilon_{pn} = \epsilon_{pp} = \epsilon_{nn} = -5.33$  MeV and the case

$\epsilon_{pp} = \epsilon_{nn} = 0; \epsilon_{pn} = -5.33$  MeV. In the first case it predicts a binding energy of 16 MeV (which is the correct answer) but in the second case it predicts 8 MeV which is a gross underestimation. The correct answer in the second case is also 16 MeV; it is just that sites will be alternately populated by neutrons and protons so that all nearest neighbor bonds are of neutron-proton type. In a similar fashion the Bragg-Williams method predicts analytically that the critical temperature in the second case is half the value of the first case but essentially "exact," although numerical results in the next section will show that the difference is much less, only about 10%. An improved treatment using the Bethe-Peierls approximation is worked out in the Appendix. The mean field calculation shown in this section and the appendix is merely to form a rough idea about the nature of the phase transition. In practical calculations we need to obtain the yields of the composites at a given temperature. Mean field theories do not provide these and we need to do event by event calculation which can be obtained through Monte Carlo samplings.

### III. MONTE CARLO RESULTS

As far as we know exact results with two kinds of bonds are not available. For one kind of bond one can often interpret essentially exact although numerical results from the well-studied spin-1/2 Ising model for use in the lattice gas model. In the absence of such exact results our only recourse is to compare numerical results obtained with two kinds of bonds, i.e.,  $\epsilon_{pn} = -5.33$  MeV,  $\epsilon_{nn} = 0$  with those obtained with one kind of bond that we have used before, i.e.,  $\epsilon_{pn} = \epsilon_{nn} = -5.33$  MeV. We use here  $N = 7^3$ , which is a number appropriate for finite systems that we will investigate. As mentioned in the Introduction the calculation proceeds by first putting the required number of protons and neutrons in the  $N$  boxes using a standard Metropolis algorithm. Nucleons are then assigned momenta from a Monte Carlo sampling of a Boltzmann distribution at a given temperature. The energy of the event can now be calculated. Clusters are then determined as explained in the Introduction. The results shown in Figs. 1 and 2 are obtained by averaging over 1000 events for selected temperatures. For two assumed freeze-out densities we calculate the specific heat, the second moment  $S_2$ , and the deduced values of  $\tau$ . The unit chosen in the graph for temperature is  $T_c = 1.1275|\epsilon_{pn}|$  which is the  $T_c$  for an infinitely large lattice with one kind of bond. We find that the peaking of  $C_v, S_2$ , and the minimum of  $\tau$  happen at a slightly lower temperature (about 10%) with two kinds of bond as compared to when the same bond is used for all the particles. For example, for the minimum of  $\tau$  to appear at the same temperature the value of  $\epsilon_{pp} = \epsilon_{pn}$  has to be set at about 10% lower value than the value of  $\epsilon_{pn}$  when  $\epsilon_{pp}$  is set to zero. Qualitatively and even semiquantitatively the results in the two models look similar when this renormalization of the strength is done. However, the peaking of  $C_v$  is more pronounced in the two bonds model.

All models which employ freeze-out densities assume that the freeze-out density is less than  $0.5\rho_0$ . If the freeze-out density is less than  $0.5\rho_0$  then in the lattice gas model a peak in  $C_v$  will signify the crossing of the coexistence curve and a first order phase transition. The value of specific heat can be

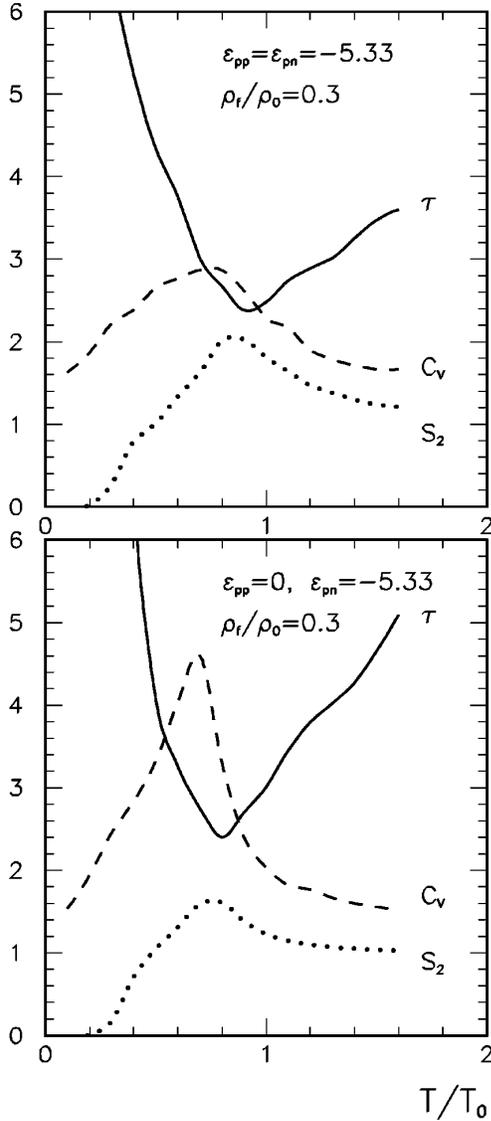


FIG. 1. A comparison of the calculated values of  $\tau$ ,  $C_v$ , and the second moment  $S_2$  in the one type of bond model (top panel) and two types of bonds model (bottom panel). Here as elsewhere  $T_0 \equiv 1.1275|\epsilon_{pn}| \approx 6$  MeV. While  $T_0$  here is merely the scale of energy it also happens to be the critical temperature for an infinite system with one kind of bond. The compound system has  $A = 103$ ,  $Z = 45$ . Notice the maxima and the minimum shift to lower temperature when  $\epsilon_{pp}$  is set to zero. Also  $C_v$  is more sharply peaked.

deduced from the caloric curve [10] but locating the peak is very difficult in experiment. In a recent paper we suggested [11] that since the peaking of  $C_v$  is accompanied by a minimum in  $\tau$  and a maximum in  $S_2$ , the appearance of the last two could be taken as a signal of the phase transition. The appearance of the maxima in  $S_2$  and of the minimum in  $\tau$  in close vicinity of the maximum in  $C_v$  happens in both the versions of the lattice gas model.

#### IV. A STUDY ON THE EFFECTS OF THE COULOMB FORCE

Here we follow the methods employed in Ref. [3]. At that time we studied the influence of the Coulomb force on fragmentation of a system which had 85 nucleons and found the

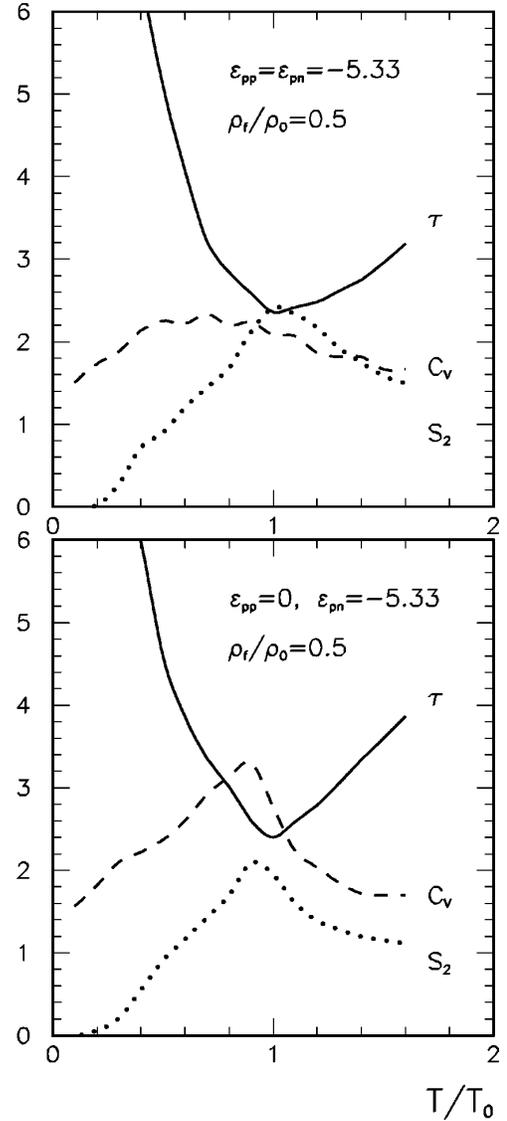


FIG. 2. The same as above except that a higher freeze-out density is used. The number of lattice sites is still  $7^3$ . Here  $A = 171$ ,  $Z = 70$ .

effects to be small. For a much larger system (Au on Au: central collision so that the compound system has  $A \approx 394$ ) we found the effect to be very large. One of the rather unavoidable features of the lattice gas model is the appearance of a minimum in the extracted value of  $\tau$  as a function of temperature. This feature disappeared for the very large system of  $A = 394$  because of the Coulomb interaction.

For completeness a short description of a similar calculation done for two kinds of bonds will be given here. In addition to lattice gas calculations, we do molecular dynamics calculations whose purpose is twofold. One purpose is to check if the predictions of a lattice gas model can resemble those of a molecular dynamics calculation provided the initial conditions are the same and the forces are chosen to be such that they resemble implied forces of the lattice gas model. For this we place the  $n_p$  protons and  $n_n$  neutrons in the  $N$  boxes using, as usual, the Metropolis algorithm. Next we assign the momenta from Monte Carlo sampling of a Maxwell-Boltzmann distribution. Once this is done the lattice gas model immediately gives the cluster distribution us-

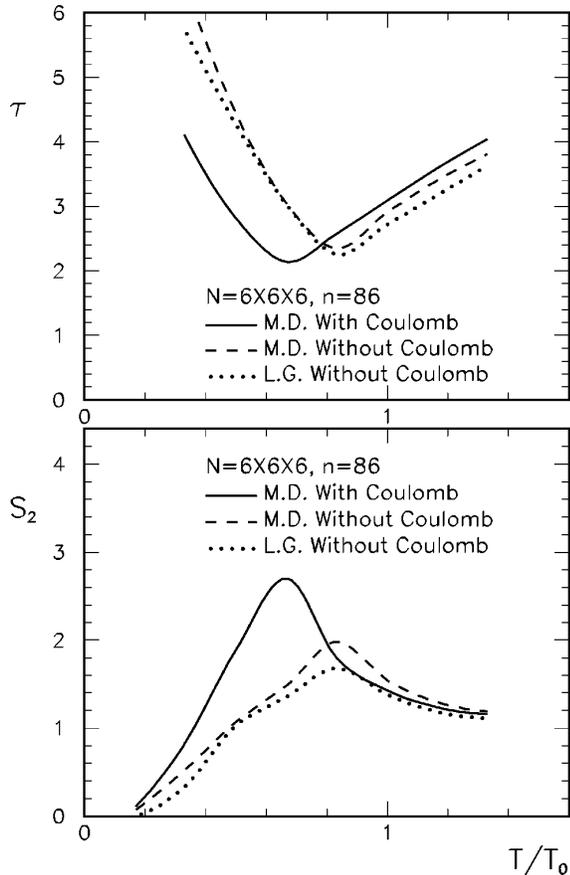


FIG. 3. The top part compares the  $\tau$  values extracted from a lattice gas calculation (dotted curve) with those extracted from a molecular dynamics calculation which had no Coulomb (dashed curve) and one which had Coulomb included in the molecular dynamics calculation (solid curve). Molecular dynamics without Coulomb gives results very similar to those of the lattice gas model. Here the effect of the Coulomb is small. The number of protons was 45. The lower part compares the second moments.

ing the rule that two nucleons are part of the same cluster if  $p_r^2/2\mu + \epsilon < 0$ . To calculate clusters using molecular dynamics we propagate the particles from this initial configuration for a long time under the influence of the chosen force (we will give the force parameters shortly). At asymptotic times the clusters are easily recognized (a detailed discussion of cluster recognition which requires shorter computer times can be found in Ref [12]). The cluster distribution in the two models can now be compared. Figure 3 shows the two prescriptions give nearly the same answer.

We now come to the second and more important purpose of the molecular dynamics calculation. We now add the Coulomb interaction to the nuclear part. The initialization of putting the nucleons in  $N$  boxes is done again but now with the inclusion of Coulomb forces. We then do a molecular dynamics propagation including the Coulomb force. The clusters can again be calculated and compared with the cases where the Coulomb force was ignored. The effects of the Coulomb interaction on fragment distribution was studied in Ref. [13] for number of nucleons 100, half of which were charged.

We now give the force parameters for molecular dynam-

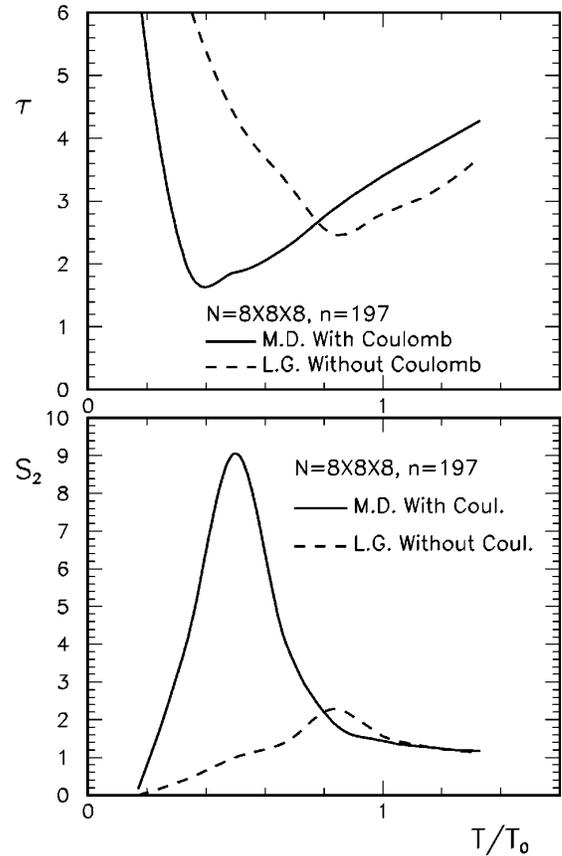


FIG. 4. Effect of Coulomb on  $\tau$  and the second moment for a much larger system;  $A=197$ ,  $Z=79$ . The minimum in  $\tau$  and the maximum in  $S_2$  shift from about 5 to 2.4 MeV because of the Coulomb effect. At some larger Coulomb field the minimum in  $\tau$  will finally disappear.

ics propagation. The neutron-proton potential was taken to be  $v_{pn}(r) = A[B(r_0/r)^p - (r_0/r)^q] \exp\{[1/(r/r_0 - a)]\}$  for  $r/r_0 < a$  and  $v_{pn}(r) = 0$  for  $r/r_0 > a$ . Here  $r_0 = 1.842$  fm is the distance between the centers of two adjacent cubes. We have chosen  $p=2$ ,  $q=1$ ,  $a=1.3$ ,  $B=0.924$ , and  $A=1966$  MeV. With these parameters the potential is minimum at  $r_0$  with the value  $-5.33$  MeV, is zero when the nucleons are more than  $1.3r_0$  apart and becomes strongly repulsive when  $r$  is significantly less than  $r_0$ . We now turn to the nuclear part of like particle interactions. Although we take  $\epsilon_{pp} = 0$  in lattice gas calculations the fact that we do not put two like particles in the same cube would suggest that there is short range repulsion between them. We have taken the nuclear force between two like particles to be the same expression as above plus  $5.33$  MeV up to  $r=1.842$  and zero afterwards:  $v_{pp}(r) = v_{pn}(r) - v_{pn}(r_0)$  for  $r < r_0$  and 0 afterwards. This means there is a repulsive core which goes to zero at  $r_0$  and is zero afterwards.

The results shown in Figs. 3 and 4 can be summarized as follows. Figure 3, first of all, shows that if there is no Coulomb interaction then lattice gas model results are quite close to that of molecular dynamics simulation provided in the latter one starts from the same initial condition and uses a force suitably chosen. Figure 3 also shows that in the case of  $A=85$ ,  $Z=40$  the Coulomb force does not have a large ef-

fect. The minimum in  $\tau$  and the maximum in  $S_2$  are shifted to slightly lower temperature. The effect for  $A=197$ ,  $Z=79$  is much bigger. The minimum in  $\tau$  and the maximum in  $S_2$  are shifted from 4.8 MeV (lattice gas without Coulomb) to about 2.4 MeV. Our previous calculation showed that there is no minimum in  $\tau$  for  $A=394$ ,  $Z=158$ . So somewhere between these two limits the minimum will vanish.

## V. DISCUSSION

The two bond model is a natural progression of the simpler lattice gas model. In this paper we have done calculations with the two bond model. Although in detail the two models differ the major characteristics of the well studied simple model remain unchanged. The lattice gas model remains a quick tool to calculate experimental data.

When the Coulomb force is very strong the lattice gas model can not be relied upon. Figures 3 and 4 give some indication of the reliability of the model in the presence of a Coulomb force. Calculations above indicate that a viable (although much more time consuming) prescription might be to obtain the initial conditions as in a lattice gas model. Put  $n$  nucleons in  $N$  boxes by Metropolis sampling where one includes, in addition to the lattice gas Hamiltonian, the Coulomb force. Obtain momenta of each nucleon from a Monte Carlo sampling of a Maxwell-Boltzmann distribution. Then propagate by molecular dynamics to obtain cluster distributions. Techniques developed in Ref. [12] might be used in order to avoid running molecular dynamics till asymptotic times. One attractive feature of this hybrid model is that the Coulomb force is operative even during the formation of clusters as opposed to other models where the Coulomb force only adds repulsion between the composites already formed.

## ACKNOWLEDGMENTS

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## APPENDIX

We follow the method of Ref. [1]. We break up the lattices into  $N/(\gamma+1)$  blocks, each of which contains 1 central box and  $\gamma$  nearest neighbors to it. We refer to Fig. 5 where for simplicity a two-dimensional lattice is shown. The interactions within each block are taken into account exactly while the interactions between different blocks are treated in an approximate fashion. The grand partition function can be written as the product of the grand partition functions of the  $N/(\gamma+1)$  blocks:

$$Z_{\text{gr}} = z_{\text{gr}}(\text{block } 1) z_{\text{gr}}(\text{block } 2) \cdots z_{\text{gr}}\left(\text{block } \frac{N}{\gamma+1}\right). \quad (\text{A1})$$

We want to write down the grand partition function of the block denoted by 1, 2, 3,  $\gamma$ , and 5. In the general case there

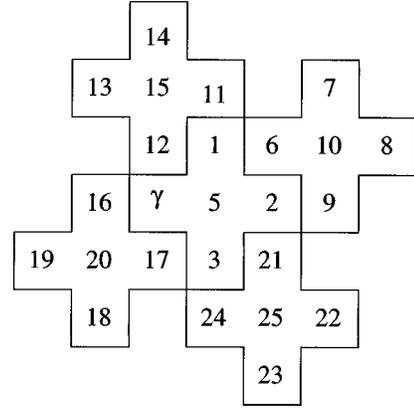


FIG. 5. A square lattice is divided into blocks to illustrate the Bethe-Peierls approximation. See text for details.

will be two absolute fugacities  $\lambda_p$  and  $\lambda_n$ , the first referring to protons and the second to neutrons. The grand partition function for a block can be written as

$$z_{\text{gr}} = A + B + C, \quad (\text{A2})$$

where

$$A = (1 + e^{\lambda_p - \beta \bar{\epsilon}_p} + e^{\lambda_n - \beta \bar{\epsilon}_n}) \gamma, \quad (\text{A3})$$

$$B = e^{\lambda_p} (1 + e^{\lambda_p - \beta \bar{\epsilon}_p - \beta \epsilon_{pp}} + e^{\lambda_n - \beta \bar{\epsilon}_n - \beta \epsilon_{pn}}) \gamma, \quad (\text{A4})$$

$$C = e^{\lambda_n} (1 + e^{\lambda_p - \beta \bar{\epsilon}_p - \beta \epsilon_{pn}} + e^{\lambda_n - \beta \bar{\epsilon}_n - \beta \epsilon_{pp}}) \gamma. \quad (\text{A5})$$

Here  $A$  = contribution to the partition function when the central (innermost) site is empty,  $B$  = contribution to the partition function when the central site has a proton, and  $C$  = contribution to the partition function when the central site has a neutron. Equation (A2) takes the place of Eq. (3.9) in Ref. [1]. The two constants  $\bar{\epsilon}_p$  and  $\bar{\epsilon}_n$  are the average interaction energy with the adjacent block when a proton (neutron) occupies a peripheral site.

The probability that the central site is occupied by a proton is  $n_p/N$ . Thus we have

$$\frac{n_p}{N} = \frac{B}{z_{\text{gr}}}. \quad (\text{A6})$$

But since no particular site is favored over another one, the average occupation of one of the peripheral sites must also be  $n_p/N$ . This gives

$$\frac{n_p}{N} = \frac{E + F + G}{z_{\text{gr}}}, \quad (\text{A7})$$

where

$$E = e^{\lambda_p - \beta \bar{\epsilon}_p} (1 + e^{\lambda_p - \beta \bar{\epsilon}_p} + e^{\lambda_n - \beta \bar{\epsilon}_n}) \gamma^{-1}, \quad (\text{A8})$$

$$F = e^{\lambda_p} e^{\lambda_p - \beta \bar{\epsilon}_p - \beta \epsilon_{pp}} (1 + e^{\lambda_p - \beta \bar{\epsilon}_p - \beta \epsilon_{pp}} + e^{\lambda_n - \beta \bar{\epsilon}_n - \beta \epsilon_{pn}}) \gamma^{-1}, \quad (\text{A9})$$

$$G = e^{\lambda_n} e^{\lambda_p - \beta \bar{\epsilon}_p - \beta \epsilon_{pn}} (1 + e^{\lambda_p - \beta \bar{\epsilon}_p - \beta \epsilon_{pn}} + e^{\lambda_n - \beta \bar{\epsilon}_n - \beta \epsilon_{pp}})^{\gamma-1}. \quad (\text{A10})$$

Similarly two equations can be written for  $n_n/N$ :

$$\frac{n_n}{N} = \frac{C}{z_{\text{gr}}}, \quad (\text{A11})$$

$$\frac{n_n}{N} = \frac{I + J + K}{z_{\text{gr}}}, \quad (\text{A12})$$

where  $I, J, K$  can be written down from expressions  $E, F, G$  by interchanging protons with neutrons. Equations (A6), (A7), (A11), and (A12) determine the four constants  $\lambda_p, \lambda_n, \bar{\epsilon}_p, \bar{\epsilon}_n$ .

For  $n_p = n_n$ , the calculations simplify. Now we have  $\lambda = \lambda_n = \lambda_p$  and  $\bar{\epsilon} = \bar{\epsilon}_n = \bar{\epsilon}_p$ . Then  $B = E + F + G$  [Eqs. (A6) and (A7)] leads to

$$e^{\lambda} (1 + Q e^{\lambda - \beta \bar{\epsilon}})^{\gamma} = e^{\lambda - \beta \bar{\epsilon}} (1 + 2e^{\lambda - \beta \bar{\epsilon}})^{\gamma-1} + e^{2\lambda} (1 + Q e^{\lambda - \beta \bar{\epsilon}})^{\gamma-1} e^{\beta \bar{\epsilon}} Q, \quad (\text{A13})$$

where we have defined  $Q = e^{-\beta_{pp}} + e^{-\beta_{pn}}$ . Dividing both sides of Eq. (A13) by  $e^{\lambda} (1 + Q e^{\lambda - \beta \bar{\epsilon}})^{\gamma-1}$  we obtain

$$1 = e^{-\beta \bar{\epsilon}} \left( \frac{1 + 2e^{\lambda - \beta \bar{\epsilon}}}{1 + Q e^{\lambda - \beta \bar{\epsilon}}} \right)^{\gamma-1}. \quad (\text{A14})$$

We rewrite Eq. (A6) as  $2N/n = z_{\text{gr}}/B$ , where  $n = n_p + n_n$ , to obtain

$$\frac{2N}{n} = 2 + e^{-\lambda} \left( \frac{1 + 2e^{\lambda - \beta \bar{\epsilon}}}{1 + Q e^{\lambda - \beta \bar{\epsilon}}} \right)^{\gamma}. \quad (\text{A15})$$

Using Eq. (A14) the above relation leads to

$$e^{\lambda} = \frac{n}{2(N-n)} e^{\beta \bar{\epsilon} \gamma / (\gamma-1)}. \quad (\text{A16})$$

We define  $x = e^{-\beta \bar{\epsilon} / (\gamma-1)}$ . Going back to Eq. (A15) one can now derive a simple solution for  $x$ :

$$x = \frac{1}{2} \left[ \frac{N-2n}{N-n} + \sqrt{\left( \frac{N-2n}{N-n} \right)^2 + 2Q \frac{n}{N-n}} \right]. \quad (\text{A17})$$

Values of  $\bar{\epsilon}$  and  $e^{\lambda}$  can be obtained from the definition of  $x$  and Eq. (A16).

Let us now go back to the partition function for the lattice as given by Eq. (A1). where it is written as a product of the partition functions of the  $N/(\gamma+1)$  blocks. If we simply  $z_{\text{gr}}$  for each little block as calculated above we will count twice the interaction between neighboring sites in different blocks. For example, the binding energy between 1 and 6 (Fig. 5) is included in  $z_{\text{gr}}$ (block 1) and included again in  $z_{\text{gr}}$ (block 2). We note that on the average there are  $n/N$  particles at each site and each block has  $\gamma$  peripheral sites. Thus when we

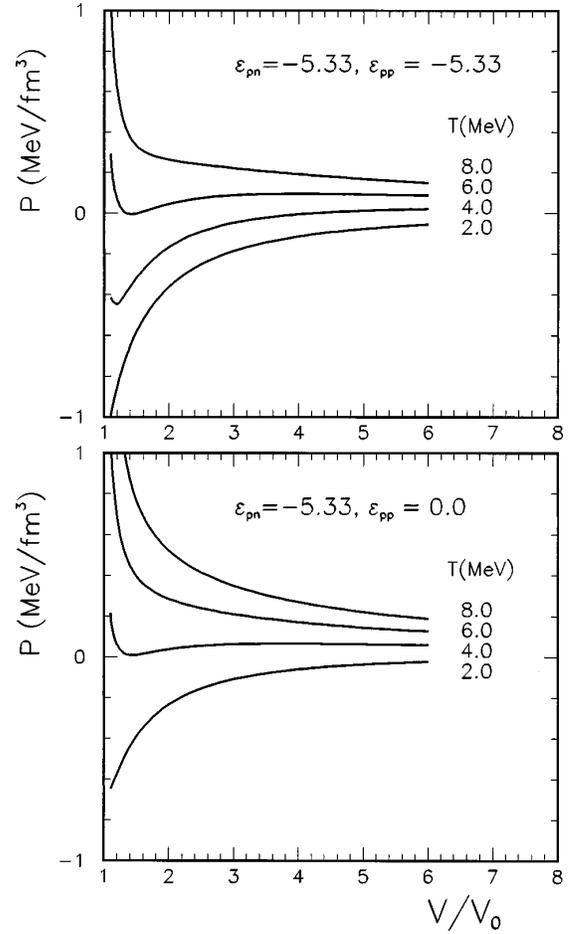


FIG. 6.  $p-V$  diagram in the Bethe-peierls approximation with same bond strength  $-5.33$  MeV (top panel) and for the case where we distinguish between like particle interaction and unlike particle interaction (bottom panel).

evaluate the partition function for the lattice, the partition function for each block should be corrected by the multiplicative factor

$$\text{correction} = e^{(1/2)\beta \bar{\epsilon} \gamma n/N}. \quad (\text{A18})$$

We can now use  $PV = kT \ln Z_{\text{gr}}$ ,  $V = N/\rho_0$ , and  $\ln Z_{\text{gr}} = N/(\gamma+1) \ln z_{\text{gr}}$  to obtain

$$P = \rho_0 kT \frac{1}{\gamma+1} \ln z_{\text{gr}}, \quad (\text{A19})$$

where  $z_{\text{gr}}$  includes the correction factor.

In Fig. 6 we have drawn  $P-V$  diagrams for nuclear matter for two cases: (1)  $\epsilon_{pn} = \epsilon_{pp} = -5.33$  MeV and (2)  $\epsilon_{pn} = -5.33$  MeV and  $\epsilon_{pp} = 0$  MeV. In the Bethe-Peierls approximation the  $T_c$  in the former case appears to be between 6 and 8 MeV and in the second case between 4 and 6 MeV.

- [1] J. Pan and S. Das Gupta, Phys. Lett. B **344**, 29 (1995).
- [2] J. Pan and S. Das Gupta, Phys. Rev. C **51**, 1384 (1995).
- [3] S. Das Gupta and J. Pan, Phys. Rev. C **53**, 1319 (1996).
- [4] L. Beaulieu *et al.*, Phys. Rev. C **54**, R973 (1996).
- [5] S. Das Gupta, J. Pan, and M. B. Tsang, Phys. Rev. C **54**, R2820 (1996).
- [6] S. Pratt, C. Montoya, and F. Ronning, Phys. Lett. B **349**, 261 (1995).
- [7] X. Campi, Phys. Lett. B **208**, 351 (1988).
- [8] J. Sivardiere and J. Lajzerowicz, Phys. Rev. A **11**, 2090 (1975).
- [9] R. J. Lenk, T. J. Schlagel, and V. R. Pandharipande, Phys. Rev. C **52**, 271 (1990).
- [10] J. Pochodzalla *et al.*, Phys. Rev. Lett. **75**, 1040 (1995).
- [11] J. Pan, S. Das Gupta, and M. Grant, Phys. Rev. Lett. **80**, 1182 (1998).
- [12] A. Strachan and C. O. Dorso, Phys. Rev. C **56**, 995 (1997).
- [13] M. Belkacem, V. Latora, and A. Bonasera, Phys. Rev. C **52**, 271 (1995).