Decay studies of neutron deficient nuclei near the Z = 64 subshell: ¹⁴²Dy, ^{140, 142}Tb, ^{140, 142}Gd, ^{140, 142}Eu, ¹⁴²Sm, and ¹⁴²Pm

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The electron-capture and β^+ -decay branchings (EC/ β^+) and delayed proton decays of A = 142isotopes with $61 \le Z \le 66$ and A = 140 isotopes with $63 \le Z \le 65$ were investigated with the OASIS facility on-line at the Lawrence Berkeley Laboratory SuperHILAC. Electron capture and positron-decay emission probabilities have been determined for ¹⁴²Pm and ¹⁴²Sm decays, and extensive decay schemes have been constructed for ¹⁴²Eu^g(2.34\pm0.12 s), ¹⁴²Gd(70.2\pm0.6 s), ¹⁴⁰Eu(1.51\pm0.02 s), and ¹⁴⁰Gd(15.8\pm0.4 s). Decay schemes for the new isotopes ¹⁴²Tb^g(597±17 ms), ¹⁴²Tb^m(303±17 ms), ¹⁴²Dy(2.3\pm0.3 s), ¹⁴⁰Eu^m(125±2 ms), and ¹⁴⁰Tb(2.4\pm0.2 s) are also presented. We have assigned γ rays to these isotopes on the basis of $\gamma\gamma$ and $x\gamma$ coincidences, and from half-life determinations. Electron-capture and β^+ -decay branchings were measured for each decay, and β delayed proton branchings were determined for ¹⁴²Dy, ¹⁴²Tb, and ¹⁴⁰Tb decays. Q_{EC} values, derived from the measured EC/β^+ branchings and the level schemes are compared with those from the Wapstra and Audi mass evaluation and the Liran and Zeldes mass calculation. The systematics of the N = 77 isomer decays are discussed, and the intense $0^+ \rightarrow 1^+$ and $1^+ \rightarrow 0^+$ ground-state beta decays are compared with shell-model predictions for simple spin-flip transitions.

I. INTRODUCTION

This paper continues our studies of light rare-earth nuclei with $N \leq 82$. Here we report the results of our investigation of the radioactive decays of the neutron deficient tigation of the radioactive decays of the neutron dencient nuclei 142 Dy, 142 Tb^{g+m}, 140 Tb, 142 Gd, 142 Eu^g, 140 Eu^{g+m}, 142 Sm, and 142 Pm. We have previously reported prelimi-nary results for 142 Dy, 142 Tb^{g+m}, 140 Tb, 140 Gd, and 140 Eu^m decays.^{1,2} The decay schemes for 142 Sm and ¹⁴²Pm were already well studied and are summarized by Peker.³ We remeasured the β -emission probabilities for these decays and determined their total electron capture and β^+ branchings. The decay of ¹⁴²Eu^g was investigated initially by Habs et $al.^4$ and by Kennedy et $al.^5$ We have placed additional transitions in the ¹⁴²Eu^g decay scheme and determined new γ -ray and β -emission probabilities. The decay of ¹⁴²Gd was first investigated by Habs et al.⁶ and later a partial decay scheme was constructed by Turcotte et al.⁷ In this paper, we present a more extensive decay scheme for ¹⁴²Gd with absolute γ -ray and β emission probabilities. The decay of ¹⁴⁰Eu^g was first reported by Westgaard,⁸ and a decay scheme was constructed by Beraud et al.⁹ We present a more comprehensive picture of the decay of this isotope and a decay scheme for the 125 ms isomer 140 Eu^m. The isotope ¹⁴⁰Gd was first reported by Redon *et al.*¹⁰ A 20 sec activity identified by Habs et al.⁶ in 1972 as an isomer of ¹⁴⁰Eu was probably due to ¹⁴⁰Gd. Partial ¹⁴⁰Gd decay schemes were prepared by Beraud et al.⁹ and by Turcotte et al.¹¹ We present a more extensive decay scheme with absolute γ -ray and β -emission probabilities. No data other than our previously reported preliminary results^{1,2} could be found in the literature regarding the decay of ^{140,142}Tb and of 142 Dy.

We have measured total electron-capture and β^+ (EC/ β^+) branchings for all the aforementioned decays. From the decay schemes and the measured EC/ β^+ branchings, we have derived $Q_{\rm EC}$ values which agree well with both the evaluated values of Wapstra *et al.*¹² and the calculated values of Liran and Zeldes.¹³

II. EXPERIMENTAL METHODS

Sources of above isotopes were produced by ${}^{92}Mo(HI, xpyn)$ reactions with 261 MeV ${}^{54}Fe$ and 224 MeV ⁵²Cr (for A = 142) and 312 MeV ⁵⁴Fe and 244 MeV ⁵²Cr (for A = 140) ions accelerated at the Lawrence Berkeley Laboratory SuperHILAC. Products recoiling from the ⁹²Mo metal foil target (97% enriched) were mass separated with the OASIS on-line facility,¹⁵ collected on a programmable moving tape, and transported to a shielded detector array consisting of a Si particle $\Delta E \cdot E$ telescope and a hyperpure Ge detector facing the radioactive source, and of a 1-mm-thick plastic scintillator and an *n*-type Ge detector on the other side of the collector tape. An additional *n*-type Ge detector oriented at 90° with respect to the two other Ge detectors was located ~ 4.5 cm from the source. A schematic view of the detector arrangement is shown in Fig. 1. In the experiments with ⁵⁴Fe beams, 8 and 400 s (for A = 142) and 8 and 80 s (for A = 140) tape cycles were used and the close-in *n*-type Ge had 24% efficiency while the 90° detector had 52% efficiency. For the ⁵²Cr experiments, the tape cycles were 2.4 and 1.6 s, and the two *n*-type detectors were interchanged. Coincidences between particles, γ rays, x rays, and positrons were recorded in an eventby-event mode; all events were tagged with a time signal for half-life information. Singles spectra were taken with



FIG. 1. Arrangement of detectors surrounding the mass-separated products collected with the fast-cycling tape system at OASIS.

the close-in detector in a multispectrum mode in which the tape cycle time was divided into eight equal-time counting intervals. During the 2.4-s cycle for A = 142and the 1.6-s cycle for A = 140, singles spectra were also taken in the 90° detector which was less subject to summing than the close-in 52% detector. The detectors were calibrated for absolute efficiency with commercial standard sources. Relative intensities and energies were calibrated with sources of ⁶⁶Ga, ¹⁵²Eu, ²²⁶Ra, and ²⁴¹Am. Analysis of the singles data was performed with the computer code SAMPO (Ref. 16) and coincidence data were analyzed using software described in Refs. 17–19. Further experimental details may be found in Refs. 14 and 15.

III. EXPERIMENTAL RESULTS AND ANALYSIS

A. Gamma-ray data

Transitions were assigned to the individual A = 142and A = 140 activities on the basis of $\gamma \gamma$ and $x \gamma$ coincidences, and by half-life analysis. The coincidence results are summarized in Tables I and II. Multipolarities were adopted from published values except for transitions in the decays of 142 Dy, 142 Tb^{g+m}, and 142 Eu^m, which were analyzed on the basis of coincidence information and decay scheme intensity balances. For some transitions, which were only observed in coincidence data, approximate γ -ray intensities were derived from the data. Coincidence relationships were used to place most γ rays in the level schemes. Transitions which were not observed in coincidence with other γ rays were assumed to feed the ground state if their intensities were comparable to other transitions observed in coincidence. Level energies were calculated by a weighted, least-squares fit of the γ -ray energies to the level scheme.

B. Decay scheme normalization

An absolute normalization of the transition intensities is required to calculate the β -decay intensity populating the ground state and isomeric states in the daughter. Determination of electron-capture and β^+ branching ratios can also help distinguish between ground-state feedings (β^+ dominated) and unobserved decay to high-lying levels (electron-capture dominated) in the daughter.

We determined the electron-capture decay branching for each A = 140 and A = 142 isotope and for ¹⁴⁰Eu^g and ¹⁴⁰Gd decays from K x-ray intensities measured in the hyperpure Ge detector. The group of four K x-ray peaks for each element was analyzed using SAMPO,¹⁶ and good agreement was obtained with relative x-ray intensities in the Table of Isotopes.²⁰ These results are summarized in Table III. The total electron-capture intensities were calculated from the K x-ray intensities after corrections for detector efficiency, fluorescence yield,²¹ internal conversion, and theoretical EC(K)/EC(total) branchings.²² The internal-conversion contribution was calculated from the measured γ -ray intensities assuming that unmeasured multipolarities were either M1 or consistent with the spin change in the proposed level scheme. A 50% uncertainty was assumed for the unmeasured conversion coefficients. For the A = 142 isotopes discussed here, corrections to the K x-ray intensity from internal conversion were generally small.

The relative β^+ -decay intensity was calculated from the 511.0-keV annihilation radiation intensity. This intensity was corrected for annihilation in flight²³ and apportioned to the various isotopes by a multilinear analysis of the multispectrum γ -ray data. For this analysis, the annihilation intensity associated with the *i*th isotope is assumed to be proportional to the intensity $I_i(\gamma)$ of a given γ ray from that decay. The total annihilation intensity thus becomes

Coincidences Gate Sm K x rays 683,768,864,890,1287,1405, 1658,1671,1754,2056,2264, 2354,2374,2420 683 768 768 683,890,1287,1405,1671 1754,2264,2420 890 768 1287 768 1405 768 1754 768 2264 768 Eu K x rays 179,212,223,242,280,284, 307,348,407,503,526,551, 586,591,614,620,632,651, 661,705,732,750,821,824, 853,862,911,936,1000,1133, 1154,1187,1204,1234,1260, 1275,1412,1438,1495,1600, 1779,1847,1949,1957,1982 179 101,223,348,(375),407,482, 572,821,1204,1234,1260, (1275),1302,1307,1600, 1847,1982 212 136,(239),284 216 280 280 223,(228),264,335,472, 910,1133,(1275) 284 212,242,307,330,336,448, 651,1154,1496 526 105 615 824 620 1158 Gd K x rays (389),465,515,693,853, (934),980 (389),465,515 465 515 (389),465,693,853,1399,1587, 1799 853 515 980 389,465 1399 515 Tb K x rays 30,69,182,212 182 30,69,99 212 69

TABLE I. A = 142 coincidence results.

Gate	Coincidences
Sm K x rays	460,531,609,715,1068,
	1098,1294,1402,1420,
	1491,1753,1952
460	531,609
531	460,1068,1098,1402,1491,
	1753,(1759)
609	460,531
685	531
715	531
1068	531
1098	531
1491	531
1753	531
2065	531
Eu K x rays	175,185,191,237,262,272,
	278,297,304,314,344,379,
	418,428,436,447,453,496,
	546,575,708,722,750,902,
	918,1041,1131
175	185,272,278,297,314,372,
	436,575,708,902,918,1041
191	237,262,344,496,558
262	191,269
272	175
297	175
453	269
750	292
Gd K x rays	328,628

TABLE II. A = 140 coincidence results.

$$I(\gamma^{\pm}) = \sum_{i=1}^{n} x_i I_i(\gamma) , \qquad (1)$$

where x_i are the proportionality constants for the *n* contributing β^+ -emitting isotopes. The γ -ray intensity data from the various tape-cycle experiments were analyzed simultaneously using the IMSL library subroutine RLMUL (Ref. 24) to obtain the best fit to the values of the parameter set x_i . This method is superior to a standard multicomponent half-life analysis because data from different experiments can be combined, in one analysis, without concern for the complex genetic relationships between

TABLE III. Comparison of experimental and theoretical K x-ray intensities for A = 142 isotopes.

	$K\alpha_1$		$K\alpha_2$		$K\beta'_1$	$K\beta'_1$		$K\beta_2'$	
	Expt	Th	Expt	Th	Expt	Th	Expt	Th	
Nd	100(2)	100	50(4)	54.9	24(8)	30.0		8.3	
Pm	100.0(12)	100	53.6(8)	55.1	30.6(12)	30.1		8.4	
Sm	100.0(13)	100	54.6(7)	55.2	31.0(3)	30.2	8.8(3)	8.6	
Eu	100.0(10)	100	55.7(8)	55.4	30.9(6)	30.5	8.5(2)	8.7	
Gd	100(8)	100	43(9)	55.6		30.8		8.9	
Tb	100(3)	100	50(2)	55.8	34(4)	31.0	10(3)	8.9	

$t_{1/2}$ Branching intensity							
Isotope	Experiment ^a	Theory ^b	β^+	EC	E_{γ}	I_{γ}^{c}	I_{γ}^{d}
¹⁴² Pm	40.5(5) s	540 s	0.771(27)	0.229(27)	1576.1	0.0196(11)	
¹⁴² Sm	72.49(5) min	22 m	< 0.05	> 0.95			
$^{142}\mathrm{Eu}^{g}$	2.34(12) s	54 s	0.899(16)	0.101(16)	768.0	0.102(7)	≡ 0.102(3)
¹⁴² Gd	70.2(6) s	80 s	0.48(5)	0.52(5)	178.9	0.112(12)	0.113(5)
142 Tb ^g	597(17) ms	3.5 s	0.968(4)	0.032(4)	515.3	0.249(17)	$\equiv 0.249(13)$
¹⁴² Dy	2.3(3) s	4.3 s	0.90(4)	0.10(4)	181.3	0.043(8)	0.051(5)
¹⁴⁰ Eu ^g	1.51(2) s	15 s	0.951(7)	0.049(7)	531.0	0.29(5)	0.30(3)
¹⁴⁰ Gd	15.8(4) s	41 s	0.67(8)	0.33(8)	749.9	0.110(15)	$\equiv 0.110$
¹⁴⁰ Tb	2.4(2) s	1.5 s	> 0.97	< 0.03	328.4	0.96	

TABLE IV. Summary of experimental and theoretical half-lives and decay branching intensities for A = 142 and A = 140.

^aValues for ¹⁴²Pm and ¹⁴²Sm are from Ref. 3. Other values from this work.

^bFrom gross theory, Refs. 36 and 37.

°Normalized to measured EC+ β^+ intensity.

^dEquilibrium intensity normalized to the indicated equilibrium partner.

the members of the mass chain. A significant source of uncertainty in this method was that not all positrons were constrained to annihilate near the source. This resulted in an energy-dependent loss of detection efficiency which we measured as 20(10)% based on the positron intensity observed in the plastic detector. The uncertainty in this correction includes the estimated energy dependence.

The total electron-capture and β^+ intensities for the A = 142 and A = 140 isotopes are shown in Table IV. The half-life relationships in the A = 142 isobaric mass chain are fortuitous because $^{142}\text{Dy} \rightarrow ^{142}\text{Tb}$, $^{142}\text{Gd} \rightarrow ^{142}\text{Eu}^g$, and $^{142}\text{Sm} \rightarrow ^{142}\text{Pm}$ each form equilibrium pairs which can be followed in appropriate tape-cycle times. $^{140}\text{Gd} \rightarrow ^{140}\text{Eu}^g$ likewise form an equilibrium pair. At equilibrium, the ratio of γ -ray intensities for transitions from each isotope must be consistent with the normalizations determined from the x-ray and 511-keV data. The absolute transition intensities obtained from the two methods are also shown and compared in Table IV.

C. Determination of $Q_{\rm EC}$

Values of the total β -decay energies Q_{EC} were derived from the decay schemes and the measured total electron-

capture and β^+ intensities. Assuming a $Q_{\rm EC}$ value and allowed β decay, theoretical EC and β^+ intensities were calculated²² for each β branch in a given decay scheme. The sum of these intensities was then compared with the experimental values and $Q_{\rm EC}$ was varied until the best agreement was obtained. With this method, any decay to unobserved high-lying daughter levels will lead to a systematic underestimate of $Q_{\rm EC}$. Strictly speaking, the values given below should therefore be construed as lower limits. Table V lists the $Q_{\rm EC}$ values derived from our data and compares them with the systematic evaluation of Wapstra and Audi¹² and with calculated values from Liran and Zeldes.¹³ The error margins shown for the experimental values represent statistical errors in the measurement of γ -ray intensities and do not include decay scheme uncertainties and other systematic effects. The overall agreement is very good and lends support to both the reliability of the above method for determining $Q_{\rm FC}$ and the completeness of the decay schemes. As a check of unobserved transitions, we used the decay schemes to calculate the relative intensities of K x rays expected to be in coincidence with the 191.2-keV transition in ¹⁴⁰Eu and with the 531.0-keV γ ray in ¹⁴⁰Sm, and compared them with measured values. The agreement is quite gratifying; the $K/(EC+\beta)$ ratios are 0.052 (decay

TABLE V. Comparison of experimental and theoretical $Q_{\rm EC}$ values.

Isotope	Experiment	Wapstra and Audi ^a	Liran and Zeldes ^b	
¹⁴² Dy	7.1(2)		7.1	
$^{142}\mathrm{Tb}^{g}$	10.4(7)		9.9	
142 Gd	4.2(3)	4.4(4)	4.6	
142 Eu ^g	7.0(3)	7.40(10)	7.5	
¹⁴² Sm	< 2.1	2.09(5)	2.2	
¹⁴² Pm	4.88(16)	4.89(5)	5.1	
¹⁴⁰ Tb	> 11.3	10.7(11)	10.9	
140 Gd	4.8(4)	4.5(7)	5.5	
140 Eu ^g	8.4(4)	8.4(4)	8.3	

^aReference 12.

^bReference 13.

Isotope	$t_{1/2}(\text{expt})$	$t_{1/2}(\mathbf{QRPA})^{\mathbf{a}}$	$t_{1/2}(\mathrm{GT})^{\mathrm{b}}$	$P_p(\text{expt})$	$P_p(\mathbf{QRPA})^{\mathbf{a}}$	$P_p(\mathbf{GT})^{\mathbf{b}}$
¹⁴² Dy	2.3(3) s	2.7 s	4.2 s	$6(3) \times 10^{-4}$	1×10^{-3}	1×10^{-3}
$^{142}\mathrm{Tb}^{g}$	597(17) ms	3.5 s	4.9 s	$2.2(11) \times 10^{-5}$	3×10^{-4}	3×10^{-4}
¹⁴⁰ Tb	2.4(2) s	2.0 s	2.2 s	$2.6(13) \times 10^{-3}$	3×10^{-3}	1×10^{-3}

TABLE VI. Comparison of experimental and calculated half-lives and delayed proton branching ratios.

^aCalculated using QRPA model; Refs. 26 and 25.

^bCalculated using the gross theory of beta decay, Refs. 36 and 37.

scheme) vs 0.049(5) (experiment) and 0.38 (decay scheme) vs 0.39(4) (experiment) for two γ rays, respectively. This indicates that the fraction of missing activity is quite low, at least for ¹⁴⁰Gd and ¹⁴⁰Eu decays. The magnitude of the missing transitions effect can also be estimated by repeating the procedure with a uniformly distributed beta strength with an average $\log ft = 5$ per MeV of excitation energy in the daughter (compatible with allowed beta strength systematics in this region) added to the region above the experimental decay scheme. The difference between the decay scheme value and the added beta strength value depends on $Q_{\rm EC}$. Around 4-5 MeV, the difference is only 100-200 keV, well within the margins of error; around 7 MeV it increases to about 200-500 keV, and becomes as large as 1–2 MeV for $Q_{\rm EC}$ values of 9–10 MeV.

D. Determination of β -delayed proton branchings

In A = 142, β -delayed protons were observed in the Si particle ΔE -E telescope in coincidence with both Tb and Gd K x rays. The proton-decay data were also consistent with two half-life components of about 0.6 and 2 s and are therefore presumed to follow the decay of the ¹⁴²Tb^g and ¹⁴²Dy precursors.²⁵ Protons from the 2.4-s tape-cycle data (146 events) and the 8-s data (115 events) were divided among the two precursors with the aid of multicomponent decay analysis using the precursor half-lives determined from the $\beta\gamma$ data, and coincident x-ray intensities. From the delayed proton intensity and the measured EC and β^+ intensities, we determined a 2.2(11)×10⁻⁵ β -delayed proton branch for ¹⁴²Tb^g decay and a $6(3) \times 10^{-4}$ branch for ¹⁴²Dy decay. In A = 140, β -delayed proton emission in coincidence with Gd K α and $K\beta$ x rays was observed. The proton half-life of 2.0(5) s is consistent with the decay of 140 Tb. Comparing the proton intensity to the intensity of the 328.4-keV γ transition (assuming all decay intensity deexcites through the $2^+ \rightarrow 0^+$ transition), we determined a 2.6(1.3)×10⁻³ β -delayed proton branch in ¹⁴⁰Tb decay.

In Table VI we compare our experimental delayed proton branches with calculated values obtained from a statistical model using beta strength functions based on a QRPA model of Krumlinde and Moller.²⁶ Calculated half-lives based on the same beta strength functions are also shown. Details of the calculations can be found in Ref. 25, and will not be elaborated here. For ¹⁴²Dy and ¹⁴⁰Tb precursors the agreement between the experimental and the calculated values of both half-lives and branches is very good. For ¹⁴²Tb a substantial discrepancy is observed in both parameters (for a discussion of this effect see Sec. V, Conclusions).

IV. DECAY SCHEMES

A. ¹⁴²Sm and ¹⁴²Pm

The half-lives of 142 Sm and 142 Pm were too long to confirm in these experiments. No γ rays could be attributed to the decay of 142 Sm, and the transitions reported by Dewanjee *et al.*²⁷ could not be confirmed. We produced 142 Sm in equilibrium with 142 Pm in the 400-s dwell time experiment. From the equilibrium Pm and Sm K xray intensities and the total 511-keV annihilation intensity at equilibrium, we determined that 142 Sm decays with less than 5% positron emission. Analysis of the equilibrium 142 Pm data yielded a relative positron-decay branching intensity of 0.771 ± 0.027 , which is in good agreement with the value of 0.779 derived by Peker.³ We also determined an 0.0196 ± 0.0011 branching intensity for the 1576.1-keV γ ray from 142 Pm, which differs appreciably from the 0.033 decay branch derived by Peker.³

B. $^{142}Eu^{g}$

With the target-projectile combinations used in our experiments, ¹⁴²Eu could only be produced by the decay of ¹⁴²Gd, and not directly in the reaction. This allowed us to investigate the ¹⁴²Eu^g decay without interference from the decay of the 1.22-min isomer $^{142}\text{Eu}^m$ (Ref. 5). Due to its lower ionization potential, the ¹⁴²Eu^g is partially separated in the mass separator ion source from its parent ¹⁴²Gd, thus allowing the determination of its halflife. The 768.1-keV $2^+ \rightarrow 0^+$ transition in ¹⁴²Sm exhibits a two-component decay curve, corresponding to equilibrium with ¹⁴²Gd and to excess ¹⁴²Eu emitted directly from the ion source. Assuming the adopted ¹⁴²Gd halflife of 70.2(6) s (Sec. IV C), we determined, from a twocomponent fit, that $t_{1/2} = 2.34(12)$ s for ¹⁴²Eu^g. This value is in good agreement with 2.4(2) s measured by Kennedy et al.⁵ We have assigned 14 γ -ray transitions to the decay of ¹⁴²Eu^g. The γ -ray energies and relative intensities are given in Table VII, where they are compared with the intensities of Kennedy *et al.*⁵ Our intensities agree marginally with those of Ref. 5 and we have added six new γ rays. In addition, we propose a tentative E0 transition deexciting the 1451.1-keV level. This transition has not been observed directly but is inferred from excess Sm K x-ray intensity in coincidence with the Sm Kx-rays from electron capture. Internal conversion is very weak in ¹⁴²Eu^g decay and the coincident x-ray intensity is

TABLE VII. Transition energies, level assignments, and relative intensities for ¹⁴²Eu^g decay.

		$I_{\gamma}(\mathrm{rel})$			
E_{γ} (keV)	Level energy	This work ^a	Kennedy et al. ^b		
683.0(1)	1451.1	4.1(4)	5.5(8)		
768.1(1)	768.1	100	100		
864.4(2)	2522.2	0.84(18)			
889.8(1)	1657.9	10.4(7)	13.3(12)		
1287.4(1)	2055.6	11.4(8)	13.7(12)		
1405.4(1)	2173.6	4.0(4)	7.1(8)		
1451.1°	1451.1	~2			
1657.9(1)	1657.9	17.2(13)	13.0(30)		
1754.1(1)	2522.2	14.6(12)	13.0(10)		
2055.8(2)	2055.6	4.4(4)	4.9(6)		
2263.7(4)	3031.9	4.9(11)			
2353.7(3)	2353.7	5.2(10)			
2373.9(3)	2373.9	7.5(9)			
2419.7(2)	3187.9	2.4(9)			

^aFor absolute intensity per 100 decays of $^{142}Eu^g$, multiply by 0.102(7).

^bReference 5.

^cE0 transition intensity inferred from excess Sm K x-ray coincidence intensity.



FIG. 2. Decay scheme for ¹⁴²Eu decay.

about four times as strong as expected. This is consistent with the behavior of the first-excited 0^+ levels in ¹⁴²Nd and 144 Sm (Ref. 19) where E0 decay branchings have been reported. It is also possible that some of the excess coincident x-ray intensity is due to additional, higherenergy E0 transitions. The decay scheme for $^{142}Eu^{g}$ is shown in Fig. 2. Firm (no parentheses) spin and parity assignments are based on (p,t) reaction measurements.²⁸ We determined $I(\beta^+)/I_{\nu}(768) = 8.8(6)$ in agreement with 8.4(11) from Ref. 5. The excited states of unknown spin and parity fall into two categories. Those levels which deexcite to the ground state are most likely 2^+ since low-lying 1⁺ levels are not normally observed in rare-earth even-even nuclei. Levels which deexcite only through the 768.1-keV 2^+ state are probably 0^+ states because ground-state γ -ray feeding from 0⁺ levels would be forbidden; no transitions to the ground state could be associated with the proposed 0^+ states.

C. ¹⁴²Gd

We have assigned 69 γ -ray transitions to the decay of $^{142}\text{Gd.}$ The $\gamma\text{-ray}$ energies and intensities are summarized in Table VIII where they are compared with values of Turcotte *et al.*⁷ The energy agreements are excellent; however, many of our intensities differ significantly. We believe that this is partly because our sources were mass separated while those of Turcotte et al.⁷ were not and contained many contaminants. Figure 3 shows the spectrum of γ rays coincident with Eu K x rays. From the decay of the most intense γ rays observed in the 400-s dwell time experiment, we determined $t_{1/2} = 70.2(6)$ s which agrees, within experimental uncertainty, with 69.1(10) s measured by Turcotte $et al.^7$ The decay scheme for ¹⁴²Gd decay is given in Fig. 4. It is similar to that in Ref. 7 except that we have identified 12 additional levels and placed 28 additional transitions. The multipolarities in Fig. 4 are from the conversion-electron data of Turcotte et al.⁷ Assuming no significant, unobserved decay to excited states, the ground-state β branch is 52(2)%. This agrees with the experimental value of 61(33)% by Turcotte *et al.*,⁷ but not with their adopted value of 94%. The latter value was deduced from the apparent first-forbidden unique feeding of the 178.9-keV level. We observe no net feeding to that level, indicating that our decay scheme is more complete and that this β transition cannot be used for intensity normalization.

D. 142 **Tb**^g

We have assigned 18 γ rays to the decay of ¹⁴²Tb^g. The γ -ray energies and intensities are summarized in Table IX. Several transitions, observed only in the 2.4-s dwell time, were assigned to ¹⁴²Tb^g decay and placed in the level scheme by energy sums and differences. Figure 5 shows a spectrum of γ rays in coincidence with Gd K x rays. The decay scheme for ¹⁴²Tb^g is shown in Fig. 6. The 465-keV transition was observed "in coincidence with itself" and placed twice in the decay scheme. Its intensity was divided on the basis of the coincidence data. The intense 515.3-keV transition (resolved from the 511-

keV annihilation peak using ¹⁶SAMPO) was analyzed as a two-component parent-daughter decay to extract a 597(17) ms half-life for ¹⁴²Tb^g decay. The ¹⁴²Tb^g parent spin is uniquely determined as 1⁺ by the low log*ft* value for decay to the 0⁺ ground state of ¹⁴²Gd. The lowest 0⁺, 2⁺, and 4⁺ states in ¹⁴²Gd are known from in-beam spectroscopy.^{29,30} The 693.7-keV 4⁺ \rightarrow 2⁺ is clearly seen, although no transitions were observed populating the 4⁺ level, and we presume that this level is fed by weak *E*2 transitions from higher-lying 2⁺ levels.

E. ¹⁴²Tb^m isomeric decay

At A = 142, the Tb K x rays, 181.9- and 211.6-keV γ rays were observed to decay with a 303(7) ms half-life.

The spectra of γ rays coincident with the 181.9- and 211.6-keV γ rays are shown in Fig. 7. From the ratio of K x rays to 68.5-keV γ rays in coincidence with the 211.6-keV γ ray, we determined $\alpha_K(68.5)=61\pm5$. This agrees with the theoretical³¹ $\alpha_K(M2)=64$ and gives a $B(M2)=2.5\times10^{-5}$. From the relative intensities of the 29.7- and 68.5-keV γ rays in the 181.9-keV gate we determined that $\alpha_{tot}(29.7)=44(15)$ and the 29.7-keV transition is M1+6(3)% E2. An alternative assignment of E1+4(1)% M2 is ruled out on the basis of single-particle Weisskopf estimate of 44 μ s for this transition. The total K x-ray intensity is consistent with M1, E1, or E2 assignments for the 181.9- and 211.6-keV γ rays. From the systematics of heavier odd-odd Tb we expect a negative-

$I_{\gamma}(\mathrm{rel})$				$I_{\gamma}(\mathrm{rel})$			
	Level	1	Turcotte		Level	,	Turcotte
E_{γ} (keV)	energy	This work ^a	et al. ^b	E_{γ} (keV)	energy	This work ^a	et al. ^b
101.4(1)	280.2	0.92(23)	0.88(15)	631.7(1)	631.7	9.5(8)	
105(1) ^c	631.7	~1		651.3(1)	935.6	2.9(4)	2.0(9)
136(1) ^c	631.7	~1		660.9(1)	660.9	4.3(5)	
178.9(1)	178.9	100.0(15)	100	704.9(1)	704.9	7.9(22)	$7.1(15)^{d}$
203(1) ^c	935.6	~ 0.8	0.85(17) ^d	732.4(1)	732.5	5.1(4)	
212.2(1)	496.6	1.53(15)		750.2(1)	750.2	7.2(7)	
216(1) ^c	496.6	~ 0.5		821(1) ^c	1000.2	~2	
222.8(1)	503.0	14.4(6)	20.5(6) ^d	823.9(1)	1438.4	10.8(25)	
228(1) ^c	732.5	~0.9	$2.1(18)^{d}$	853(1) ^c	1438.4	~1.1	1.3(3) ^d
238.8(1)		1.3(2)		862(1) ^c	1412.9	~ 3	0.5(4) ^d
241.7(2)	526.2	1.5(4)	4.5(3)	886.3(2)			2.9(13) ^d
247.2(1)	750.2	1.60(23)	2.6(6)	910.0(1)	1412.9	2.4(5)	$4.2(11)^{d}$
264.2(1)	544.4	4.1(3)		912.0(2)			$2.9(6)^{d}$
274.3(4)			b	935.6(1)	935.6	4.4(5)	9.6(21) ^d
280.3(1)	280.2	35.9(8)	38.4(4)	1000.2(1)	1000.2	14(2)	
284.4(1)	284.4	55.0(15)	53.9(5)	1073.6(4)			b
306.9(1)	591.3	7.2(5)	$2.6(8)^{d}$	1133(1) ^c	1412.9	~1.3	1.8(4)
330.4(1)	614.6	2.9(5)	6.5(28) ^d	1153.8(1)	1438.4	2.1(5)	5.5(5)
335(1) ^c	614.6	~ 2	$1.5(12)^{d}$	1158(1) ^c	1779.1	~2	
336(1) ^c	619.7	~0.5		1187(1) ^c	1779.1	~6	1.7(8) ^d
347.6(1)	526.2	4.0(6)	$2.6(5)^{d}$	1204.4(1)	1383.3	4(2)	
375.4(1)		2.5(3)		1233.9(1)	1412.9	15.0(8)	$14.6(29)^d$
407.0(1)	585.7	4.8(4)	5.7(4)	1259.6(1)	1438.4	38.2(15)	29.9(6) ^d
448.2(1)	732.5	1.8(4)	2.6(5) ^d	1275(1) ^c	1779.1	~2	7.0(7) ^d
466(1) ^c	750.2		b	1302(1) ^c	1480.9	~ 3	
472(1) ^c	750.2	~1	b	1307(1) ^c	1485.9	~1	
503.0(1)	503.0	6.4(15)	b	1412.4(2)	1412.9	6.8(15)	
526.2(1)	526.2	52.7(15)	38.9(15)	1438.4(2)	1438.4	11(4)	
550.6(1)	550.6	6.4(7)	5.6(12)	1495.0(2)	1779.1	5.9(15)	4.9(6)
553(1) ^c	732.5	~5		1599.7(2)	1779.1	18(3)	8.3(7) ^d
572(1) ^c	750.2	~5	$3.5(20)^{d}$	1779.1(1)	1779.1	22.1(23)	
585.7(2)	585.7	4.7(6)		1846.7(2)	2025.6	6(3)	
591.3(1)	591.3	9.8(7)		1948.6(3)	1948.6	9(4)	
595.9(3)			$2.1(15)^{d}$	1956.6(3)	1956.6	7(3)	
614.5(1)	614.6	13.0(8)	10.8(4)	1982(1) ^c	2160.9	~2	2.8(14) ^d
619.7(1)	619.7	18.3(8)	13.3(5)				

TABLE VIII. Transition energies, level assignments, and relative intensities for ¹⁴²Gd decay.

^aFor absolute intensity per 100 decays of ¹⁴²Gd, multiply by 0.112(11).

^bReference 7.

^cObserved only in coincidence.

^dIntensity was obtained after subtraction of known contaminants.

parity isomer (see the discussion below). Assuming that the 98.3-keV transition has E3 multipolarity, consistent with a reasonable B(E3)=0.026 value, we can assign a 5^- spin to the isomer, 2^+ to the 181.9-keV level, and 3^+ to the 211.6-keV level. No β decay was observed from ¹⁴²Tb^m. Indirect population of the 4^+ and 6^+ levels in ¹⁴²Gd would characterize this decay. We observe weak feeding of the 4^+ level by ¹⁴²Tb^g decay; however, the intensity of this branch remains constant in equilibrium with ¹⁴²Dy and is therefore not due to ¹⁴²Tb^m. We can set an 0.5% upper limit on the β -decay branch assuming the decay populates only levels which deexcite through the 4⁺ level. This lack of β decay confirms our assignment of the β -delayed proton activity to ¹⁴²Tb^g. The isomeric decay scheme of ¹⁴²Tb^m is shown in Fig. 8. A second isomer in ¹⁴²Tb with a half-life of 15 μ s which presumably deexcites to the 5⁻ level has also been reported,³² but was not observed in these experiments.



FIG. 3. Spectrum of $A = 142 \gamma$ rays in the 54% Ge detector coincident with Eu K α x rays in the HPGe detector. Background transitions from ¹⁴²Sm populated by the intense ¹⁴²Eu source produced in this experiment are indicated.

TABLE IX.	Transition	energies,	level	assignments,	and	rel	a-
tive intensities f	for ¹⁴² Tb ^g de	ecay.					

E_{γ} (keV)	Level energy	$I_{\gamma}(\mathrm{rel})^{\mathrm{a}}$
388.8(1)	1368.7	0.90(9)
465.0(1) ^b	980.0	11(2)
	1445.1	2(1)
515.3(1)	515.4	100(7)
693.7(1)	1209.1	4.0(5)
853.1(1)	1368.7	9.7(8)
898.4(3)	2343.9	1.1(3)
934(1) ^c	1914.7	
980.1(1)	980.0	1.8(4)
1299.6(2)	2279.6	1.6(2)
1364.1(4)	2343.9	0.8(2)
1399.2(1)	1914.7	9.6(8)
1587.4(3)	2102.8	1.4(4)
1764.1(2)	2279.6	0.9(2)
1799(1) ^c	2314	~2
1828.7(2)	2343.9	2.7(3)
1915.0(2)	1914.7	2.3(3)
2343.6(3)	2343.9	1.0(2)

^aFor absolute intensity per 100 decays of ${}^{142}Tb^{g}$, multiply by 0.249(7).

^bUnresolved transition, intensity divided using coincidence data. The total 465-keV transition intensity is 13.3(12).

^cTransition observed only in coincidence.

F. ¹⁴²Dy

A single 181.9-keV γ -ray transition, deexciting the level of that energy in ¹⁴²Tb, was assigned to ¹⁴²Dy decay on the basis of half-life and coincidence with Tb K x rays and 511-keV annihilation radiation. This transition, also observed in the 303 ms ¹⁴²Tb^m decay (Sec. IV E), decays with a longer 2.3(3)-s half-life component, consistent with the half-life measured for the ¹⁴²Dy delayed proton branch. The total intensity of the 181.9-keV transition is 6(2)% indicating that the 2⁺ 181.9-keV level is populated by γ transitions from higher levels in ¹⁴²Tb fed by allowed β transitions from the even-even ¹⁴²Dy. Assuming no strong γ -ray transitions were missed, the ¹⁴²Dy decay predominantly populates the ground state of ¹⁴²Tb with a logft=4.1. This confirms the 1⁺ assignment to the ¹⁴²Tb ground state (see also Fig. 8).

G. ¹⁴⁰Eu^g

We have assigned 20 γ rays to the decay of ¹⁴⁰Eu. The γ -ray energies and intensities are summarized in Table X. We have adopted a 1.51(2)-s half-life for ¹⁴⁰Eu^g decay assuming two half-life components for the 531.0-keV γ ray corresponding to feeding in equilibrium with ¹⁴⁰Gd decay and direct production. This value agrees within uncertainties with 1.3(2) s from Redon *et al.*¹⁰ and 1.54(13) s



FIG. 4. Decay scheme for ¹⁴²Gd decay.

from Deslauriers *et al.*³³ The spectrum of γ rays in coincidence with Sm K x rays is shown in Fig. 9 and the ¹⁴⁰Eu^g decay scheme is shown in Fig. 10. The firm daughter spin assignments are from ¹⁴⁰Sm in-beam reaction studies.³⁰ Strong β feeding is observed to the ground (0⁺) and first excited (2⁺) states of ¹⁴⁰Sm constraining the ¹⁴⁰Eu ground state to 1⁺. This assignment conflicts with the 2⁺ assignment by Turcotte *et al.*,¹¹ but is consistent with heavier even-even Eu isotope ground states. The decay of ¹⁴⁰Gd, discussed below, also confirms the 1⁺ assignment. Weak population of the 4⁺ level in ¹⁴⁰Sm has been observed and is most probably due to the deexcitation of higher-lying 2⁺ levels.

H. ${}^{140}Eu^{m}$

A 185.3-keV γ ray was observed to decay with a single-component half-life of 125(2) ms in the 1.6-s dwell

time data. The decay curve for this transition is shown in Fig. 11. This γ ray was not found in coincidence with any transitions. Similar short half-life components were observed for the 174.8-keV γ ray deexciting the firstexcited state in ¹⁴⁰Eu (see below) and for Eu K x rays. We have thus assigned the 125(2)-ms half-life to an isomer in ¹⁴⁰Eu. The intensities of the x rays and γ rays assigned to 140 Eu^m decay are also given in Table X. The multipolarity of the 174.8-keV transition is M1 from Ref. 11, and, if we assume that all of the observed Eu K x rays for 140 Eu^m decay are from the internal conversion of the 174.8- and 185.3-keV transitions, we derive $\alpha_K(expt)=0.19(4)$ for the 185.3-keV transition. This is consistent with the theoretical³¹ value α_K (theory)=0.19 for an E2 transition. There is insufficient Eu K x-ray intensity for the 185.3-keV transition to have M2 or higher multipolarity so it is doubtful that the 185.3-keV level is the isomeric state. We have assigned the 174.8- and 185.3-keV levels



FIG. 4. (Continued).



FIG. 5. Spectrum of $A = 142 \gamma$ rays in the 54% Ge detector coincident with Gd K α x rays in the HPGe detector.



FIG. 6. Decay scheme for 142 Tb^g decay.



FIG. 7. Spectrum of $A = 142 \gamma$ rays in the HPGe detector coincident with the 182- and 212-keV γ rays in the 54% Ge detector.





FIG. 8. Decay schemes for ${}^{140}\text{Eu}^m$, ${}^{142}\text{Tb}^m$, and ${}^{142}\text{Dy}$ decays.

as 2^+ and 3^+ , respectively. This is consistent with significant indirect feeding of the 174.6-keV level by ¹⁴⁰Gd decay and negligible indirect feeding to the 185.3keV level. No Eu K x-ray intensity can be associated with the isomeric transition, establishing its energy at less than the Eu K binding energy of 48.5 keV and consequently setting an upper limit of 234 keV for the isomer energy in ¹⁴⁰Eu (see Fig. 8).

From the recommended upper limits for γ -ray transition probabilities,³⁴ the isomeric transition must be M2or E3. This constrains the isomer to a spin of 5^{-1} if both lower levels are directly populated. Based on the intensity of the 185.3-keV γ ray, we can derive B(M2) values ranging from 6×10^{-5} to 2×10^{-4} , corresponding to isomeric transition energies of 10 to 48 keV, respectively. This is consistent with the degree of hindrance observed for M2 transitions in this region. A similar estimate of B(E3) for the transition feeding the 174.6-keV level yields values substantially exceeding 1 W.u. (W.u. represents Weisskopf unit), contrary to expectations based on systematics of E3 transitions. This indicates that most of the observed intensity of the 174.6 γ ray is due to a 10.7-keV M1 transition (which cannot be observed directly in our experiments) between the 185.3and 174.6-keV levels.

No EC/ β^+ decay was observed with ¹⁴⁰Eu^{*m*} decay. This decay mode would presumably populate the 4⁺ level in ¹⁴⁰Sm, yet no short-lived component was observed for the 715.4-keV level depopulating that state. From the relative intensity of the 174.8-, 185.3-, and 715.4-keV transitions, we can set an upper limit of 1% on the EC/ β^+ branching intensity.

I. ¹⁴⁰Gd

We have assigned 32 γ rays to the decay of ¹⁴⁰Gd. The γ -ray energies and intensities are summarized in Table XI where they are compared with those of Turcotte et al.¹¹ Our energies agree with Ref. 11; however, there is again poor agreement with the intensities, especially with regard to the intense 750-keV transition. It is not apparent why the agreement is so poor except for the fact that no chemical or mass separation was performed in Ref. 11. From the decay of the strongest γ rays, we determined a 15.8(4)-s half-life for ¹⁴⁰Gd. This agrees with 16.2(8) s from Turcotte et al.¹¹ but not with 11(2) s by Redon et al.¹⁰ Habs et al.⁶ reported a 20-s isomer of ¹⁴⁰Eu based on the observed decay of a 531-keV γ ray with that half-life. It is now apparent that they produced that transition from the decay of ¹⁴⁰Eu^g in equilibrium with ¹⁴⁰Gd.

The spectrum of γ rays in coincidence with Eu K x

TABLE X. Transition energies, level assignments, and relative intensities for 140 Eu^{g + m} and 140 Tb decay.

Parent	E_{γ} (keV)	Level energy	$I_{\gamma}(\mathrm{rel})$
¹⁴⁰ Eu ^{g (a)}	39.5	Sm $K\alpha_2$	4.23(16)
	40.1	Sm $K\alpha_1$	7.45(30)
	45.4	Sm $K\beta_{1'}$	1.9(4)
	46.6	Sm $K\beta_{2'}$	0.63(3)
	352.4(2) ^b	1599.1	0.4(2)
	459.9(1)	990.7	11.0(8)
	531.0(1)	531.0	100(9)
	608.6(1)	1599.1	1.9(2)
	685.1(2)	2284.1	0.9(3)
	715.4(2)	1246.5	0.6(1)
	882.7(3)	2482.4	0.2(1)
	1068.0(1)	1599.1	11.0(11)
	1097.7(2)	1628.6	2.0(3)
	1293.6(1)	2284.1	1.2(2)
	1299.4(2)	2102.8	0.3(1)
	1402.2(2)	1933.2	0.9(2)
	1420.3(2)	1420.3	1.2(2)
	1491.3(2)	2482.4	2.1(3)
	1752.8(2)	2284.1	1.9(3)
	1758.7(4)	2289.9	0.4(2)
	1952.0(2)	2482.4	1.4(2)
	2064.9(3)	2595.8	3.2(6)
	2283.9(3 ^b)	2284.1	0.5(2)
	2289.1(5)	2289.9	0.2(1)
$^{140}\mathrm{Eu}^{m(\mathrm{c})}$		Eu $K\alpha,\beta$	50(5)
	174.8(1)	174.6	100(4)
	185.3(1)	185.3	92(4)
¹⁴⁰ Tb ^d		Gd $K\alpha,\beta$	< 6.7
	328.4(2)	328.4	100(18)
	(508) ^e	837	~ 100
	627.8(2)	1456	52(9)

^aFor absolute intensity per 100 decays multiply by 0.29(5).

^bPlaced in the decay scheme by energy sums.

^cFor absolute intensity per 100 decays multiply by 0.39(2).

^dFor absolute intensity per 100 decays multiply by 0.96.

^eNot measured in this experiment. From Ref. 35.

rays is shown in Fig. 12, and the decay scheme for ¹⁴⁰Gd decay is shown in Fig. 13. The ¹⁴⁰Eu level scheme below 750 keV is similar to that of Turcotte *et al.*¹¹ Significant discrepancies exist only for higher levels. We observe substantial β decay to the ¹⁴⁰Eu ground state which is consistent with our 1⁺ assignment and not the 2⁺ assigned in Ref. 11. The 174.6- and 191.2-keV levels are not significantly populated by β decay; this is inconsistent with their previous 1⁺ assignments.¹¹ Higher levels are too weakly populated to confirm 1⁺ assignments with the

exception of the 749.9-keV level where the $\log ft = 4.4$ is indicative of a definite 1⁺ spin and parity.

J. 140Tb

Two γ -rays with energies of 328.4 and 627.8 keV have been assigned to ¹⁴⁰Tb decay on the basis of coincidence with Gd K x rays. The relevant coincidence spectrum is shown in Fig. 14. The 328.4-keV γ ray decays with a single-component 2.4(2)-s half-life. Both γ rays were ob-



FIG. 9. Spectrum of $A = 140 \gamma$ rays in the 54% Ge detector coincident with Sm K α x rays in the HPGe detector.

served in in-beam reaction studies of ¹⁴⁰Gd (Ref. 35) and assigned to the $2^+ \rightarrow 0^+$ and $6^+ \rightarrow 4^+$ transitions. The intensities of these γ rays are given in Table X. The $4^+ \rightarrow 2^+$ transition is known³⁵ to be 508 keV and could not be resolved from the intense 511-keV annihilation radiation. The 627.8-keV transition is about one-half of the intensity of the 328.4-keV transition which is consistent with significant β feeding and a spin of 5, 6, or 7 for the ¹⁴⁰Tb parent. No evidence for the 675-keV 8⁺ $\rightarrow 6^+$ transition was observed. Figure 15 shows the decay scheme for ¹⁴⁰Tb. Here we have assumed that nearly all of the decay populates levels which deexcite through the 4⁺ or 6^+ levels in ¹⁴⁰Gd. From the relative intensities of the 328.4- and 627.8-keV γ rays we conclude that about half of the beta decay directly or indirectly populates each of these two levels. Very little Gd K x-ray intensity is observed with this decay so feeding to high-lying levels must be minimal. Based on these assumptions, we derive $Q_{\rm EC} > 11.3$ MeV, in agreement with the value of 10.7(11) MeV from Wapstra *et al.*¹² From the apparent low log*ft* or intense γ -ray feeding to the 4⁺ and 6⁺ levels, we conclude the ¹⁴⁰Tb spin is probably 5. We cannot determine whether the 2.4-s higher-spin Tb is the ground state or an isomer. However, no Tb K x rays were observed indicat-

			$I_{\gamma}(\text{rel})$
E_{γ} (keV)	Level energy	This work ^a	Turcotte et al. ^b
40.9	Eu $K\alpha_2$	21.9(10)	
41.5	Eu $K\alpha_1$	38.6(16)	
47.0	Eu $K\beta_{1'}$	11.5(6)	
48.3	Eu $K\beta_2$	2.9(2)	
174.8(2)	174.6	108(19)	100
186.7(3)	361.3	3.8(13)	5.9(11)
191.2(1)	191.2	49(9)	44.8(34)
236.7(1)	427.9	8.9(13)	13.8(7)
253.3(2)	427.9	8.9(19)	
261.8(2)	453.3	7.6(13)	10.3(16)
269.0(2)	722.3	1.9(6)	
272.4(1)	447.0	0.9	10.8(13)
278.4(5)	453.3	10.1(13)	4.6(7)
296.6(2)	749.9	12.7(13)	16.2(10)
304.5(2)	722.3	1.3(6)	
313.5(3)	488.1	12.0(25)	
344.5(4)	535.7	1.9(13)	
372.0(2)	546.6	3.8(13)	
379.0(1)	284.4	54(8)	5(3)
417.7(1)	417.7	39(4)	с
427.9(2)	427.9	11.4(13)	
436.4(2)	611.0	7.0(13)	9.8(19)
446.9(3)	447.0	6.3(19)	
453.4(2)	453.3	19(6)	с
495.8(2)	687.0	9.5(13)	с
532.0(4)			9(5)
546.5(2)	546.6	20.3(19)	
558.7(3)	749.9	33(3)	
575.4(1)	749.9	39(4)	26.0(43)
686.2(4)			6(4)
708.1(2)	882.7	9.5(13)	с
722.3(1)	722.3	27(4)	46(6)
749.9(1)	749.9	100(6)	38(12)
774.6(3)			с
903.2(3)	1077.8	3.2(13)	
918(1)	1092.6	9(4) ^c	
982.9(4)			с
1041.4(2)	1216.0	13(6)	
1131.1(3)	1131.1	5.1(13)	5.3(23)

TABLE XI. Transition energies, level assignments, and relative intensities for ¹⁴⁰Gd decay.

^aFor absolute intensity per 100 decays of ¹⁴⁰Gd, multiply by 0.70(8).

^bReference 11.

^cObserved only in coincidence.



FIG. 10. Decay scheme for ¹⁴⁰Eu^g decay.



FIG. 11. Decay curve for the 185.3-keV γ ray from ¹⁴⁰Eu^m decay.

ing that there is no isomeric transition from this state (or that the isomeric transition energy is less than 50 keV). From the systematics of the heavier odd-odd Tb isotopes,³⁸ one should expect a 1^+ ground state with a half-life considerably shorter than 1 s. This state would be only weakly populated in this reaction and decay mostly to the ¹⁴⁰Gd ground state; thus it would probably have escaped detection in these experiments.

V. CONCLUSIONS

The $Q_{\rm EC}$ values deduced from the radioactive decay schemes described in this paper are in good agreement with both the evaluated values of Wapstra and Audi,¹² and the calculated values of Liran and Zeldes.¹³ This agreement suggests that statistical feeding to high-lying daughter levels in the A = 142 and A = 140 mass chains is low or even negligible. The principal reason for this result is the preponderance of $1^+ \rightarrow 0^+$ and $0^+ \rightarrow 1^+$ ground-state to ground-state β transitions with low log ft values in these even mass decay chains. The β strength is strongly dominated by these fast transitions and leaves little remaining intensity to be apportioned statistically at higher excitation energies. This point is emphasized when we compare (Table IV) the measured half-lives with those calculated from the gross β -decay theory.^{36,37} Agreement for the even-even parents is typically within a factor of 2-3 while odd-odd decays are calculated to be about an order of magnitude longer lived than is observed. Inspection of the gross theory beta strength model reveals that it underestimates the strength of decay to levels below the pairing gap in the even-even daughter nuclei. The odd-odd parent beta strength to the eveneven daughter ground state is therefore neglected, while even-even parent decays are assumed to have considerable beta strength to levels near the daughter's ground state. If we were to include the measured ground-state beta strength in the calculation, the agreement between the gross theory and the experimental values would become much better. This same effect can also explain the discrepancy noted in Sec. III D between the experimental β -delayed proton decay branch for the odd-odd low-spin precursor ¹⁴²Tb^g and the value calculated with a quasirandom-phase approximation (QRPA) model.^{25,26} However, good agreement was obtained for both the even-even precursor ¹⁴²Dy and the odd-odd, J^{π} =5 precursor ¹⁴⁰Tb (see Table VI). Like the gross theory, the QRPA beta strength model tends to underestimate the strength of ground-state β transitions from low-spin oddodd nuclei and therefore overestimates calculated halflives and delayed proton branches.

The log ft values for the $0^+ \rightarrow 1^+$ and $1^+ \rightarrow 0^+$ ground-state beta transitions are all less than 5. Below N=82 and above Z=58, the $\pi d_{5/2} \rightarrow v d_{3/2}$ spin-flip transition is expected to be fast, and above Z=64 the $\pi h_{11/2} \rightarrow v h_{9/2}$ transition is important. In the even-even nuclei either a $(\pi d_{5/2})^2$ or a $(\pi h_{11/2})^2$ pair are assumed to decay, and in odd-odd parents the odd proton and neutron are expected to recouple. The systematics³⁸ of the log ft values for the even-even transitions are plotted in Fig. 16(a). The shell model predicts³⁹

$$ft = \frac{6160}{g_a^2 B(\text{GT})}$$
, (2)

where $g_a = 1.263$, the Gamow-Teller matrix element B(GT) = n(4l/2l+1), and *n* is the number of valence protons. These predictions are shown in Fig. 16(b) and are nearly an order of magnitude lower than experiment. This phenomenon has been commented on previously by Nolte *et al.*⁴⁰ for the N = 82 region and by Barden *et al.*⁴¹ for the Z = 50 region ($\pi g_{9/2} \rightarrow v g_{7/2}$ transitions). Towner⁴² has argued that these discrepancies are due to pairing correlations, core polarization, and higher-order phenomena.

The n dependence predicted by the shell model [Eq. (2)] appears to persist in the spin-flip transitions where

the average $\log ft$ value changes between n=2 and n=6by 0.54(12) for N=76-78, Z=60-64, and by 0.40(6) for N=82-86, Z=66-70. These values agree with the expected value of 0.48 from Eq. (2). The N=80 transitions were excluded from the first average because the shell model predicts that the $vd_{3/2}$ orbital is filled, blocking the spin-flip transition. This is demonstrated by comparing the N=76 and $N=78 \log ft$ values. They differ by 0.22(12), consistent with an 0.3 predicted difference if the $vd_{3/2}$ orbital were half filled at N=78. Experimentally, a nearly constant $\log ft = 5.0$ is observed for all decays with N=80. The $vs_{1/2}$ and $vd_{3/2}$ orbital are nearly degenerate in this region, perhaps explaining the residual β strength at N=80.

The systematics³⁸ of the $1^+ \rightarrow 0^+$ decays are shown in Fig. 17. As in the neighboring $0^+ \rightarrow 1^+$ transitions, the



FIG. 12. Spectrum of $A = 140 \gamma$ rays in the 54% Ge detector coincident with Eu K α x rays in the HPGe detector.



FIG. 13. Decay scheme for ¹⁴⁰Gd decay.



FIG. 14. Spectrum of $A = 140 \gamma$ rays in the 54% Ge detector coincident with Gd $K\alpha$ x rays in the HPGe detector.

log ft values for N=79 and Z=59-63 decrease by 0.50(11) from n=1 to n=5, which is slightly less than the 0.70 value predicted by the shell model. This trend is not observed for N=81 where the spin-flip transition is blocked. It is remarkable that these shell-model trends are sustained despite the fact that the log ft values are hindered. However, the calculations by Towner⁴² suggest that the hindrance should be constant for the l=5 orbitals. Similar shell-model trends have also been reported for isomeric E2 transitions in N=82 nuclei.^{40,43-46}

Another interesting feature is the low $\log ft$ values for 142,144 Dy. The valence protons in these isotopes should be $(\pi h_{11/2})^2$ yet the $vh_{9/2}$ neutron orbital lies above the N=82 shell gap. We thus expect the ground-state transition to be dominated by the $\pi d_{5/2} \rightarrow vd_{3/2}$ transition. The $\pi d_{5/2}$ orbital is filled at Z=64 so we expect the log ft values in the light Dy isotopes to be similar to the adjacent Gd decays. The log ft values reported here are much lower than expected and seem consistent with the Dy decay for N > 82. It is likely that the 142,144 Dy log values



FIG. 15. Decay scheme for ¹⁴⁰Tb decay.

are low due to missed β strength to levels near 2 MeV from the $\pi h_{11/2} \rightarrow v h_{9/2}$ transition. This transition was observed in ¹⁴⁶Dy decay⁴⁷ with a log*ft*=3.8(2) that is consistent with the heavier Dy decays. Assuming that this transition occurs with the same log*ft* in the lighter Dy isotopes, about 35% of the decay would populate that resonance in ¹⁴²Tb raising the log*ft* for the ground-state transition to a value comparable to ¹⁴⁰Gd decay. For ¹⁴⁴Dy decay the $\pi h_{11/2} \rightarrow v h_{9/2}$ branch would be only ~7% which is, however, insufficient to raise the groundstate log*ft* to the ¹⁴²Gd value.

As can be seen from Fig. 8, the ¹⁴⁰Eu^m and ¹⁴²Tb^m decays are remarkably similar. The similarity of these decays, despite their straddling the semimagic Z = 64 shell, may be indicative of a common structure and of the disappearance of that shell below N = 78 (Refs. 14 and 48). From the well characterized systematics of the N = 77 isotones it has been established that the odd neutron should be in either a $\frac{1}{2}^+$ or an isomeric $\frac{11}{2}^-$ orbital, presumably related to the $vs_{1/2}$ and $vh_{11/2}$ shell-model states. In our investigations of ¹⁴¹Tb decay, ¹⁴ we assigned $J^{\pi} = \frac{5}{2}^-$ to the ground state and suggested a $\frac{5}{2}[532]$ configuration. ¹⁴³Tb has been assigned $J^{\pi} = \frac{5}{2}^+$ (Ref. 10) and may be the $\frac{5}{2}[402]$ configuration, although other neighboring configurations are possible. The $\frac{5}{2}[402]$ configuration cannot be coupled with the expected neutron configuration to explain the 5⁻ for both odd-odd



FIG. 16. Ground-state $0^+ \rightarrow 1^+ \log ft$ values for even-even nuclei with Z = 60-70 and N = 74-86. (a) Experimental $\log ft$ values (Ref. 38). (b) Predicted $\log ft$ values from shell-model calculations (Ref. 39).



FIG. 17. Experimental $1^+ \rightarrow 0^+$ ground-state log ft values.

N=77 isomers. However, the expected shell-model state at Z=65 is $\pi s_{1/2}$, which can be coupled with the $vh_{11/2}$ neutron to explain the measured spin. At Z=63, the $\pi d_{5/2}$ shell-model state is expected which also cannot couple with the odd neutron to give 5⁻. Alternate proton configurations for Z=63 and/or Z=65 include $\frac{1}{2}[550]$ and $\frac{1}{2}[411]$, both of which lie near the proposed

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configurations for ^{141,143}Tb and could give rise to the correct spin.

For ${}^{142}\text{Tb}^m$ decay we determine that the single-particle $B(M2) \sim 2.5 \times 10^{-5}$ and $B(E3) \sim 0.026$ transition probabilities are very hindered. This is likely if the neutron Nilsson configurations for the 2⁺ and 3⁺ all have low K quantum numbers. The M2 transition in ${}^{140}\text{Eu}^m$ is also very hindered, but the $B(M2) \sim (2-6) \times 10^{-4}$ is almost an order of magnitude larger than that for the corresponding transition in ${}^{142}\text{Tb}$.

The A = 140 and A = 142 isotopes in this study are close to the Z = 64 shell closure which is predicted to disappear near N = 78. The A = 140 isotopes cross Z = 64 at N = 76 and are expected to show signs of significant deformation.^{14,48} Since our experiments sample predominantly 0⁺ and 1⁺ states, no conclusive and unambiguous evidence for such deformation or for the disappearance of the shell gap near ¹⁴²Gd could be found in the data.

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