Electron capture decay of 81 Kr^m

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The 81 Kr^{*m*} electron capture branching ratio to 81 Br has been remeasured and found to be $(2.26\pm0.32)\times10^{-5}$ with a log *ft* value of 4.93 ± 0.06 . This result is 2.5 times smaller than a previous measurement and implies a similar reduction in the rate for solar neutrino capture in an 81 Br-based detector.

This paper describes an improved measurement of the electron capture branching ratio for the decay of the 190 keV, 13-sec isomeric state of ⁸¹Kr to the ground state of ⁸¹Br (see Fig. 1). This Gamow-Teller transition dominates the capture cross section for low energy solar neutrinos on ⁸¹Br. Recent proposals for large-scale solar neutrino experiments using ⁸¹Br have emphasized the need for accurate ($\sim 15\%$) determinations of this transition strength. We therefore undertook the present measurement, using techniques similar to those applied in a previous measurement¹ in our laboratory, but with substantially improved counting rates. With these improved rates, we were able to identify a large background contamination which invalidated the previous result. The 81 Kr^m production technique used in the present experiment reduces this background to a negligible level and enables an improved measurement with an accuracy $\sim 15\%$.

The branching ratio for electron capture of ⁸¹Kr^m is measured by comparing the number of Br $K\alpha$ x rays observed from a pure sample of ⁸¹Kr^m with the far more abundant Kr $K\alpha$ x rays produced in the internal conversion of the isomeric state. As in the previous measurement, the Princeton Isotope Separator was used to implant ⁸¹Kr^m in a Mylar tape which was then moved to a detection region for counting. To reduce the background from the tail of the Kr $K\alpha$ x-ray peak, a plastic scintillator detector with 98% efficiency for internal conversion electrons was used to veto such events. X-ray spectra were accumulated for five successive time intervals following the tape movement to confirm that the Br K x rays exhibited the 13-sec half-life characteristic of ⁸¹Kr^m.

In the earlier experiment, 22 MeV protons were used to produce 81 Kr^m via the (p,n) reaction on a solid potassium bromide target. In our initial attempt to improve this measurement, we used a boiling liquid target of bromoform. This target provided a larger count rate and much longer target lifetime than the original experiment. With this improved target system, we were able to examine a broad mass region near mass 81 and thus discovered a defocused, long-lived activity which produced a background of Br x rays. This contaminant is probably due to 79 Kr (34.9 h) which is also prolifically produced in (p,n) reactions on the natural bromine targets used. The contaminant accounted for a large fraction of the Br x rays observed at mass 81 and was undoubtedly also present in the original measurement. This background could be reduced by improving the mass resolution of the isotope separator, but it was decided to also change the method of producing 81 Kr^m to reduce the background to a negligible level.

⁸¹Kr^m was obtained from the decay of ⁸¹Rb (4.58 h) which selectively populates the 190 keV isomeric state. Millicurie sources of ⁸¹Rb were obtained from Medi-Physics Inc., South Plainfield, NJ. The ⁸¹Rb was extracted with HCl solution from the ion exchange resin on which it was supplied, dried, and placed in the ion source region of the Princeton Isotope Separator. The emanating gas containing ⁸¹Kr^m was ionized, accelerated to 60 keV, mass separated for 81, and implanted in an aluminized Mylar tape (25 μ m thickness). A collimator allowed only the central portion of the isotope separator beam to reach the tape. After accumulating for 30 sec, the tape containing the ⁸¹Kr^m was pulled into the detection region shown in Fig. 2 and a new section of tape began accumulating.

In the detection region the tape-borne source was viewed by a lithium drifted silicon detector [Si(Li)] and two vanes of 0.5 mm Pilot B scintillator coupled to an RCA 8575 photomultiplier. A major improvement was obtained in this experiment by mounting the scintillator within a slot in the light pipe, producing much better light collection and improving the detection efficiency from 79% to 98%. As shown in Fig. 2, a reentrant cavity for the Si(Li) was provided. For x rays in the region of the Br $K\alpha$ peak at 11.94 keV, the Si(Li) detector had a resolution of approximately 240 eV and a maximum total rate of about 1500 counts/sec. The scintillator was primarily sensitive to internal conversion electrons and operated at a maximum rate of about 150 000 counts/sec. A time-topulse-height converter (TPHC) operating on a 100 nsec scale was started by the Si(Li) and stopped either by an event in the scintillator or by the start signal delayed by 60 nsec. The coincidence peak width was 14 nsec FWHM. A valid TPHC pulse was required for every event so that dead time in the TPHC could not distort the Br to Kr ratio. Windows on the coincidence peak and the anticoincidence peak were used to sort the x-ray events from the Si(Li) detector into two spectra. Spectra from five successive 6 sec intervals following tap movement were separately accumulated in order to allow a measurement of the half-life for the Br events. The data taking sequencing was provided by a crystal-controlled microprocessor.

Ex(MeV)

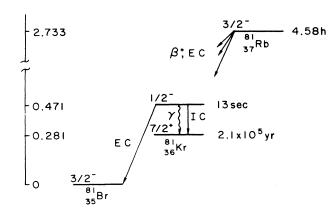


FIG. 1. Decay scheme of ⁸¹Kr^m, showing population of the isomer by ⁸¹Rb as well as subsequent gamma ray, internal conversion, and electron capture decay branches.

A single ⁸¹Rb generator provided an adequate count rate for about 15 h of data taking. Ten generators were used during two running periods at a rate of one per day. Long-term observation of a ⁸¹Rb generator with a lithium drifted germanium gamma ray detector revealed a small ⁸³Rb ($t_{1/2} = 86.2$ d) contamination. ⁸³Rb decays to ⁸³Kr^m ($t_{1/2} = 1.83$ h). The isomer decays exclusively to the stable ⁸³K ground state, therefore producing no Br Ka x rays. Internal conversion from this state could, in principle, contribute to the number of Kr Ka x rays, but the low activity, the isotope separation, and the short residence time in front of the counters make this contribution negligible.

Figure 3(a) shows the sum of all our data for the x rays in anticoincidence with internal conversion (IC) electrons in the scintillator. Those in the coincidence are shown in Fig. 3(b). The much greater statistical accuracy of the krypton K x-ray peaks in Fig. 3(b) allows that spectrum to be scaled and subtracted from the anticoincidence one to produce Fig. 3(c), in which 80% of the Kr K α x-ray peak has been subtracted away. With this subtraction, a signi-

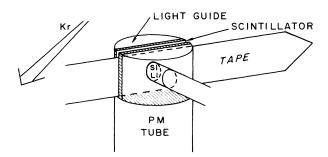


FIG. 2. Schematic overview of the source collection, tape transport, and detection regions of the experiment, showing the high geometric solid angle of the scintillator, improved light coupling to the photomultiplier (PM) and the reentrant cavity for the lithium-drifted silicon detector [Si(Li)].

ficant peak at the energy of the Br $K\alpha$ x ray can be seen in Fig. 3(c). In order to check the consistency and stability of this peak, spectra for individual days were subtracted in this way and analyzed as follows. The data were grouped in two different ways and areas determined with the peak-fitting program FITEK (Ref. 2) to produce the results in Table I. The upper part of the table shows the result of summing over the five (6 sec) decay intervals and grouping into three roughly equivalent data collection periods. The lower part of the table gives the result for each of the decay intervals summed over all the data collected.

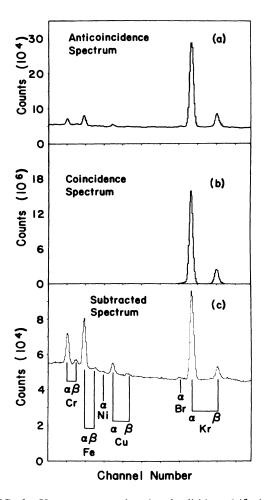


FIG. 3. X-ray spectra taken by the lithium-drifted silicon detector summed over all runs and all decay intervals. (a) X rays not in coincidence with internal conversion electrons detected in the scintillator. The flat background arises from Compton scattering of the 190 keV gamma branch; the krypton x rays from incomplete vetoing by the scintillator; and the Cr, Fe, Ni, and Cu x rays from fluorescing of the detector components by the 190 keV gamma rays. (b) X rays in coincidence with interconversion electrons. Fluorescence of detector constituents and the silicon escape peak are not visible on this scale. (c) Anticoincidence spectrum minus a scaled version of the coincidence one so as to remove 80% of the Kr K α x rays. The small bromine $K\alpha$ x-ray peak is clearly visible on the low energy side of the krypton $K\alpha$ peak.

Chronological intervals						
	Kr $K\alpha$	Kr $K\alpha$		Observed ratio		
	coincidences	anticoincidences	Br $K\alpha$	$(\times 10^{-5})^{a}$		
April 1986	37 434 352	616 558	2760 ± 505	7.25 (18.3%)		
July 1986-1	33 030 192	454 03 1	2393 ± 424	7.15 (17.7%)		
July 1986-2	<u>23 999 040</u>	<u>357 734</u>	1635 ± 402	6.71 (24.6%)		
-	94 463 584	1 428 323	6788 ± 772	7.08 (11.4%)		
		Decay intervals				
Time after		-				
source arrival	Kr $K\alpha$	Kr $K\alpha$		Observed ratio		
(sec)	coincidences	anticoincidences	Br $K\alpha$	$(\times 10^{-5})^{a}$		
0-6	30 952 480	519 609	2355 ± 403	7.48 (17.1%)		
6-12	23 790 928	351 740	1956 ± 480	8.10 (24.5%)		
12-18	17 996 304	251 826	1200 ± 378	6.57 (31.5%)		
18-24	13 521 351	183 782	614 ± 305	4.48 (49.7%)		
24—30	<u>10091307</u>	<u>134 251</u>	566 ± 265	<u>5.53 (46.8%)</u>		
	96 352 370	1 441 208	6690 ± 836	6.83 (12.5%)		
	from fit Br	$K\alpha = 6660.5 \pm 806.2 \Longrightarrow 6.81 \times 10^{-3}$	(12.1%).			

TABLE I. Peak areas and ratios^a for the x-ray spectra grouped by run and by decay interval.

^aRatio=(Br $K\alpha$)/($\sum Kr K\alpha$).

As can be seen, the groupings all give a ratio which is consistent with our adopted value of (6.9×10^{-5}) (±13%) for the ratio of Br K α x rays to Kr K α x rays. The Br x-ray peak areas in the lower part of the table can be fitted to the exponential decay of ⁸¹Kr^m ($t_{1/2}$ =13.10±0.02 sec) (Ref. 3) with no background contribution. The oneparameter fit gives a reduced χ^2 of 0.3 and implies a total of 6661±806 Br K α x rays in the five intervals, as given at the bottom of Table I. (See Fig. 4.) A two-parameter

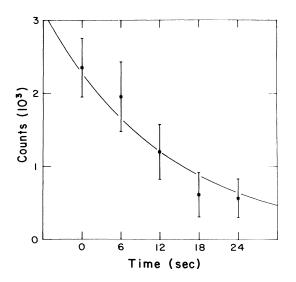


FIG. 4. Decay of the bromine $K\alpha$ peak. Peak areas for sequential 6 sec time bins are plotted at the time after source arrival that accumulation began in that bin. The curve is a one-parameter fit to the data with the half-life fixed at 13.1 sec and with a reduced χ^2 of 0.3.

fit, including a possible flat background such as from long-lived activities, still results in zero background with an upper limit for the background (90% confidence) of 1000 counts or 15% of the observed total.

More stringent limits can be placed by considering the possible Br x-ray producing contaminants. Most can be ruled out by the requirements that they still be present at least 4 h after production and survive being dissolved in HCl and then dried, a process which removes any gaseous components. The only remaining possibilities are ⁷⁹Rb (23 min) and ${}^{80}\text{Br}^m$ (4.42 h). Data were recorded at these two mass settings of the Isotope Separator for approximately an hour each. Small amounts of Kr x rays were seen and upper limits were established on possible Br x rays. Assuming the Kr x rays arise from incomplete mass separation of the ⁸¹Kr^m gives mass separation efficiencies of 3.2×10^{-4} and 8×10^{-5} for 1 and 2 mass number spacings. Applying these to the observed Br x-ray upper limits implies upper limits (90% confidence) for mass 79 and 80 contributions to the total number of mass 81 Br $K\alpha$ x rays of 0.1 and 0.6 counts, respectively. These values are negligibly small, and therefore no additional uncertainty arises from the possibility of contaminants in the Br x ravs.

The high separation efficiencies observed and lack of any large amount of other Kr x-ray producing activities means that the only source for contamination of that peak arises from the possibility of ⁸¹Rb being vaporized, included in the isotope separator beam, and deposited on the tape. The Rb electron capture decay branch would lead to an x ray with no electron and so would be present in the anticoincidence spectrum only. Since this component would have a nearly constant rate over the 30 sec observation period, the number of anticoincident Kr K α x rays in bin 5 sets an upper limit of <0.8% for the fraction of Kr $K\alpha$ x rays from ⁸¹Rb decay itself. This is a negligible

Correctio	Value		Refs.	
(1) Relative efficiency	$\frac{1/\epsilon(11.92 \text{ keV})}{1/\epsilon(12.65 \text{ keV})}$	=	1.018 (1%)	4
(2) Total K x ray to $K\alpha$	$\frac{K/K\alpha \text{ (Br)}}{K/K\alpha \text{ (Kr)}}$	=	0.988 (3%)	5
(3) Fluorescence yield	$\frac{1/\omega_K (\mathbf{Br})}{1/\omega_K (\mathbf{Kr})}$	=	1.039 (3%)	5
(4) Total to K captures or conversions	$\frac{1 + \mathrm{EC}(L + M + \cdots)}{1 + \mathrm{IC}(L + M + \cdots)}$	$\frac{1+0}{1+1/2} = \frac{1+0}{1+1/2}$	$\frac{118}{4.53} = 0.916$ (3%)	5 6
(5) Total decays to conversions	$\frac{1}{1+\gamma/IC}$	1	$\frac{1}{0.49} = 0.329 \ (1.5\%)$	5
		-	Total=0.315 (5.5%)	

TABLE II. Corrections to observed $K\alpha$ x-ray ratio to obtain total branching ratio.

contribution to the uncertainty.

The ratio of observed $K\alpha$ x rays must be multiplied by a number of corrections in order to obtain the total ⁸¹Kr^m electron capture (EC) branching ratio. These corrections and their assumed values are listed in Table II and give a net correction factor of 0.315 ($\pm 5.5\%$). A measured correction of 1.042 ($\pm 1\%$) must be applied to account for Br x rays lost due to random coincidences with the scintillator. This yields a value of (2.26 ± 0.32)×10⁻⁵ for the branching ratio. Combining this with a lifetime of 13.1 sec and a Q value of 0.471 MeV,⁷ which implies a logf of -0.83,⁸ gives a logft of 4.93±0.06 for the Kr ($\frac{1}{2}^{-}$) to Br ($\frac{3}{2}^{-}$) decay. This implies a logft for the inverse Br ($\frac{3}{2}^{-}$) to Kr ($\frac{1}{2}^{-}$) neutrino capture of 5.23±0.06. It is important to note that because of the strongly suspected contamination in the original measurement, this redetermination should replace the previous result.

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Another recent measurement⁹ also gives a result significantly smaller than the original measurement and not inconsistent with our new determination.

This reduction in the expected yield of a bromine-based neutrino detector by a factor of approximately 2.5 is significant in terms of the detailed plans for such a detector, but does not affect the basic conclusion that such a detector is practical, provided a method of single atom counting of krypton can be developed. Moreover, the improved precision of the present measurement will allow the results from such a detector to be clearly interpreted in terms of a solar neutrino flux.

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