⁷Be decay scheme and the solar neutrino problem

Eric B. Norman, Timothy E. Chupp, Kevin T. Lesko, and John L. Osborne Nuclear Physics Laboratory, University of Washington, Seattle, Washington 98195

Patrick J. Grant* and Gene L. Woodruff

Nuclear Engineering Department, University of Washington, Seattle, Washington 98195 (Received 29 November 1982)

The decay scheme of ⁷Be has been reinvestigated. Known numbers of ⁷Be nuclei were produced in targets via the ⁷Li(p, n) reaction. Following activation, the yields of 478-keV γ rays from the targets were measured. From three such measurements, performed at different bombarding energies, the ⁷Be decay branching ratio to the first excited state of ⁷Li has been determined to be 10.8±0.4%. The implications of this result for the solar neutrino problem are discussed.

RADIOACTIVITY ⁷Be: measured I_{γ} ; deduced branching ratio to first excited state in ⁷Li. Neutron detector. Ge(Li) detector.

The long-standing discrepancy between the experimental and theoretical values for the solar neutrino flux is one of the outstanding problems in physics. The combined results of nearly ten years of observations by Davis and collaborators¹ have established that the mean capture rate for the ³⁷Cl detector is 1.8 ± 0.9 SNU (3σ uncertainty), where 1 SNU $=10^{-36}$ neutrino captures per target atom per second. Over an even longer period of time, numerous theoretical estimates of the solar neutrino flux have been performed. The most recent calculations by Bahcall *et al.*² and Filippone and Schramm³ yield values of 7.6 ± 3.3 SNU and 7.0 ± 3.0 SNU, respectively, for the ³⁷Cl detector. The cause for this disagreement between experiment and theory continues to be a subject of great interest.

Because of its relatively high threshold energy (0.81 MeV), the 37 Cl detector is mainly sensitive to the high energy neutrinos produced by the beta decay of 8 B. The production of 8 B in the sun is the result of a very small branch of the proton-proton chain, the last four reactions of which are

$$^{3}\text{He} + ^{4}\text{He} \rightarrow ^{7}\text{Be} + \gamma$$
, (1)

$$^{7}\mathrm{Be} + p \rightarrow ^{8}\mathrm{B} + \gamma \quad , \tag{2}$$

$$^{8}B \rightarrow ^{8}Be + e^{+} + \nu \quad . \tag{3}$$

$$^{8}\text{Be} \rightarrow ^{4}\text{He} + ^{4}\text{He} \quad . \tag{4}$$

Hence, the ³⁷Cl solar neutrino capture rate is nearly linearly dependent upon the rates for the ³He(⁴He, γ)⁷Be and ⁷Be(p,γ)⁸B reactions.⁴ The ⁷Be decay scheme plays a role in some experimental determinations of the cross sections for both of these reactions.

The ³He(⁴He, γ)⁷Be reaction was studied experimentally in the 1960's by Parker and Kavanagh⁵ and by Nagatani, Dwarakanath, and Ashery⁶ and was reinvestigated recently by Kräwinkel *et al.*⁷ and by Osborne *et al.*⁸ In all of these studies, the yields of the capture γ rays were measured in beam. The results of the two early measurements agree reasonably well with each other. However, the recent work of Kräwinkel *et al.*⁷ suggests that the cross section factor, $S_{34}(0)$, may be 30–50% lower than the previous values. If this result were correct, it would substantially reduce the discrepancy between the ex-



FIG. 1. Decay scheme of ⁷Be. B_0 and B_{γ} are the ⁷Be electron-capture decay percentages to the ⁷Li ground state and to the 477.6 keV first excited state, respectively.

27

1728

© 1983 The American Physical Society

perimental and theoretical values of the solar neutrino flux. The results of the subsequent measurement by Osborne *et al.*⁸ again agree with the early results.

Another technique to determine the ${}^{3}\text{He}({}^{4}\text{He},\gamma){}^{7}\text{Be}$ cross section is to measure the 'Be activity following the bombardment. As can be seen in Fig. 1, ⁷Be electron capture decays with a 53.3 d half-life to the ground state and first excited state of ⁷Li. Thus by measuring the yield of the 478-keV γ rays from an activated target cell, the ⁷Be production cross section can be determined. This method was used by Osborne et al.8 and recently by Robertson et al.,⁹ and both experiments yielded results consistent with those obtained from the early inbeam measurements. Obviously, a parameter which enters the calculation of the ${}^{3}\text{He}({}^{4}\text{He},\gamma)$ cross section from the 478-keV γ -ray yield is B_{γ} , the ⁷Be decay branching ratio to the $J^{\pi} = \frac{1}{2}^{-}$ first excited state of ⁷Li.

This same branching ratio also plays a role in some measurements of the ${}^{7}\text{Be}(p,\gamma){}^{8}\text{B}$ cross sections. One of the methods that has been used to determine the thicknesses of ${}^{7}\text{Be}$ targets used in such experiments is to measure the yield of 478-keV γ rays from the targets.^{10,11} Again, in order to calculate the number of ${}^{7}\text{Be}$ nuclei contained in the target from the γ -ray yield, B_{γ} must be known.

Over the years, B_{γ} has been measured a number of times by both direct and indirect techniques. $^{12-22}$ In the direct B_{γ} determinations, electron-capture decays to both the ground state and 478-keV excited state of ⁷Li were observed in measurements of the Auger electron and γ -ray yields. The indirect methods rely on a measurement of the number of ⁷Be nuclei produced in a target and a separate measurement of the γ -ray yield. The results of these experiments agree remarkably well with one another and yield a mean value of 10.4%.²³ However, Rolfs et al.²⁴ recently reported a value of 15.4±0.8% based upon an indirect measurement. If this result were correct, it would cause the ${}^{3}\text{He}({}^{4}\text{He},\gamma)$ cross sections obtained by Osborne et al.8 and by Robertson et al.⁹ in their activation experiments to be reduced by approximately 33%, thus bringing them into closer agreement with the result of Kräwinkel et al.⁷ On the other hand, if B_{γ} were 15.4%, then the thicknesses of ⁷Be targets inferred from γ -ray measurements would have to be reduced, and hence the ${}^{7}\text{Be}(p,\gamma)$ reaction rates determined from such measurements would have to be increased by approximately 50%.

Because of its importance in both of these aspects of the solar neutrino problem, we have reinvestigated the decay scheme of ⁷Be using an indirect activation method. The number of ⁷Be nuclei produced in a target via the ⁷Li(p,n) reaction was determined by measuring the neutron yield. Following activation, the target was counted and the yield of the 478-keV γ ray was determined. From three such measurements performed at different proton bombarding energies B_{γ} was determined.

Targets consisting of approximately 700 μ g/cm² ⁷LiF evaporated onto thick Au backings were bombarded with proton beams of a few nanoamperes from the University of Washington's FN tandem accelerator. At the bombarding energies used in the present experiment, 1.90, 2.10, and 3.50 MeV, the only neutron-producing reaction possible is the ⁷Li(p,n)⁷Be reaction. The neutron yield (and hence the number of ⁷Be nuclei produced in each target) was measured with a large neutron detection system, the details of which will be described elsewhere.²⁵ Briefly, the system consists of ten ³He-filled proportional counters that are imbedded in a 1.5-m diameter graphite moderator which surrounds the target area. The detectors are oriented perpendicular to the beam direction and are positioned around the surface of a half-cylinder. The beam is stopped in the target assembly and the outgoing neutrons are moderated in a 15-cm diameter iron sphere immediately surrounding the target and in the large graphite pile.

Four different techniques were used to measure the detection efficiency of this system. Three of these methods measured the absolute efficiency averaged over three different energy intervals while the fourth method determined the relative efficiency as a function of neutron energy. Two of the measurements involved the use of calibrated neutron sources, a 252 Cf fission source calibrated to $\pm 3\%$ and a 238 Pu- 13 C source calibrated to $\pm 10\%$. We emphasize that both of these sources had been calibrated by techniques independent of the ⁷Be decay scheme. Each source was placed at the normal target position and counting was done for a measured period of time. The third absolute efficiency determination was obtained by placing a thick PbF₂ target in the target position and bombarding it with a beam of 2 nA of 10.0-MeV alpha particles. At this bombarding energy, the only significant neutron

 TABLE I. Results of neutron detector efficiency measurements.

Neutron		Efficiency
source		(%)
²³⁸ Pu- ¹³ C		5.69±0.57
$^{19}\mathrm{F}(\alpha,n)^{22}\mathrm{Na}$		5.32 ± 0.29
²⁵² Cf		5.49 ± 0.27
	Weighted mean	5.45 ± 0.18

TABLE II. Results of the present measurements of B_{γ} using the ⁷Li(*p*,*n*) reaction.

Ep	Mean E_n	Bγ
(MeV)	(MeV)	(%)
1.90	0.05	11.2±0.5
2.10	0.20	10.8 ± 0.4
3.50	1.29	<u>10.6±0.4</u>
	Weighted mean ^a	10.8±0.4

^aUncertainty in mean value includes systematic 3.3% uncertainty in ⁷Be reference source calibration.

production mechanism is the ${}^{19}F(\alpha,n)^{22}Na$ reaction. The neutron yield from a Pb target was measured and was found to be negligible at 10 MeV. The neutron yield from the PbF₂ was measured and following the bombardment, the ${}^{22}Na$ activity in the target was determined by measuring the yield of 1275-keV γ rays with a well-shielded Ge(Li) detector. The efficiency of the Ge(Li) detector at this energy was determined with two calibrated ($\pm 5\%$) ${}^{22}Na$ sources obtained from New England Nuclear. Finally, by dividing the observed neutron yield by the number of ${}^{22}Na$ nuclei produced in the target, the neutron detection efficiency was obtained. The results of these three measurements are summarized in Table I.

Each of the above-mentioned efficiency measurements involved neutrons with continuous energy spectra, from a few keV up to several MeV. To measure the dependence of the efficiency on the neutron energy, we bombarded a tritium target with 2-6 MeV protons. At low bombarding energies, the neutrons produced by the ${}^{3}\text{H}(p,n)$ reaction are emitted in a fairly narrow energy range. Thus, as the proton energy was varied, the mean neutron energy was varied from approximately 0.7 to 3 MeV. The observed neutron yields were divided by the known ${}^{3}\text{H}(p,n)$ cross sections²⁶ to obtain the detector efficiency. The results of these measurements showed that the efficiency varies by less than 15% over this energy range. Furthermore, a Monte Carlo calculation of the efficiency was performed utilizing a three-dimensional model of the detector. These calculations suggest that the efficiency varies by no more than 10% over the range of neutron energy 1-3 MeV.

Following the ${}^{7}Li(p,n)$ activations, each target was counted with a well-shielded 79 cm³ Ge(Li) detector and the yield of the 478-keV ⁷Be γ ray was measured. The efficiency of the Ge(Li) detector at this energy was determined with a ⁷Be source from Isotope Products Laboratories whose γ -ray emission rate had been calibrated to $\pm 3.3\%$. Additional checks on the γ -ray detection efficiency were made with a number of standard calibrated sources. Dividing the observed γ -ray yield by the inferred number of ⁷Be nuclei in each target thus yields the ⁷Be decay branching percentage to the 477.6-keV level in ⁷Li. The results of this procedure are summarized in Table II. The mean value obtained from the three measurements of B_{γ} is 10.8±0.4%. The uncertainty in the mean value includes both statistical and systematic errors.

In Table III, the results of the present experiment are compared with previous measurements of B_{γ} . As can be seen, the present results are in good agreement with the early measurements, but definitely do not agree with the recent measurement of Rolfs *et al.*²⁴ We do not have an explanation for this discrepancy. However, we feel that the combined

Author	Year	Method ^a	B_{γ} (%)
Rumbough et al. (Ref. 12)	1938	A	10^{+20}_{-7}
Williamson and Richards (Ref. 13)	1949	В	10.7 ± 2.0
Turner (Ref. 14)	1949	В	11.8 ± 1.2
Dickson and Randle (Ref. 15)	1951		12.3 ± 0.6
(reevaluation of Turner)			
Taylor and Merritt (Ref. 16)	1962	С	10.32 ± 0.16
Poenitz (Refs. 17 and 18)	1966	В	10.5 ±0.2
Mutterer (Ref. 19)	1970	С	10.47 ± 0.20
Szabo et al. (Ref. 20)	1972	D	10.4 ±0.3
Poenitz and DeVolpi (Ref. 21)	1973	В	10.42 ± 0.18
Goodier et al. (Ref. 22)	1974	С	10.35 ± 0.08
Rolfs et al. (Ref. 24)	1982	D	15.4 ±0.8
Present work	1982	В	10.8 ±0.4

TABLE III. Results of present and previous measurements of B_{γ} .

^aIndirect: (A) Measured neutron yield from ⁶Li(d, n) and γ yield; (B) measured neutron yield from ⁷Li(p, n) and γ yield; (D) measured α yield [and ⁷Be yield (Ref. 24)] from ¹⁰B(p, α) and γ yield. Direct: (C) Measured Auger-electron yield and γ yield.

results of ten independent experiments, using several different techniques, support a value of B_{γ} of approximately 10.4%. This was the value used by Osborne *et al.*⁸ and by Robertson *et al.*⁹ in the analysis of their activation experiments. Thus the ³He(⁴He, γ) cross section factors determined in their two experiments are in good agreement with each other and with the results of Parker and Kavanagh⁵

*Present address: The Boeing Company, Seattle, WA 98031.

and Nagatani et al.⁶ This same value of B_{γ} was

used in the 7Be target thickness measurements of

- ¹R. Davis, Jr., B. T. Cleveland, and J. K. Rowley, in Proceedings of the Workshop on Science Underground, Los Alamos, 1982 AIP Conf. Proc. (AIP, New York, to be published).
- ²J. N. Bahcall et al., Rev. Mod. Phys. <u>54</u>, 767 (1982).
- ³B. W. Filippone and D. N. Schramm, Astrophys. J. <u>253</u>, 393 (1982).
- ⁴J. N. Bahcall, N. A. Bahcall, and R. K. Ulrich, Astrophys. J. <u>156</u>, 559 (1969).
- ⁵P. D. Parker and R. W. Kavanagh, Phys. Rev. <u>131</u>, 2578 (1963).
- ⁶K. Nagatani, M. R. Dwarakanath, and D. Ashery, Nucl. Phys. <u>A128</u>, 325 (1969).
- ⁷H. Kräwinkel et al., Z. Phys. A <u>304</u>, 307 (1982).
- ⁸J. L. Osborne et al., Phys. Rev. Lett. <u>48</u>, 1664 (1982).
- ⁹R. G. H. Robertson et al., Phys. Rev. C <u>27</u>, 11 (1983).
- ¹⁰C. Wiezorek et al., Z. Phys. A <u>282</u>, 121 (1977).
- ¹¹B. W. Filippone et al., Phys. Rev. Lett. <u>50</u>, 412 (1983).
- ¹²L. H. Rumbough, R. B. Roberts, and L. R. Hafstad, Phys. Rev. <u>54</u>, 657 (1938).
- ¹³R. M. Williamson and H. T. Richards, Phys. Rev. <u>76</u>,

Wiezorek *et al.*¹⁰ and Filippone *et al.*¹¹ Thus, no changes are required in the results they obtained for the ⁷Be(p, γ) cross sections. The discrepancy between the experimental and theoretical values for the solar neutrino flux, therefore, remains a tantalizing puzzle.

This work was supported in part by the U. S. Department of Energy.

- 614 (1949).
- ¹⁴C. M. Turner, Phys. Rev. <u>76</u>, 148 (1949).
- ¹⁵J. M. Dickson and T. C. Randle, Proc. Phys. Soc. (London) <u>A64</u>, 902 (1951).
- ¹⁶J. G. V. Taylor and J. S. Merritt, Can. J. Phys. <u>40</u>, 926 (1962).
- ¹⁷W. P. Poenitz, J. Nucl. Energy <u>20</u>, 825 (1966).
- ¹⁸W. P. Poenitz, Nuclear Data for Reactors (IAEA, Vienna, 1966), Vol. 1, p. 277.
- ¹⁹M. Mutterer, Neutron Standards and Flux Normalization (AEC, Argonne, 1970), Symposium Series, Vol. 23, p. 452.
- ²⁰J. Szabo, J. Csiakai, and M. Varnagy, Nucl. Phys. <u>A195</u>, 527 (1972).
- ²¹W. P. Poenitz and A. DeVolpi, Int. J. Appl. Rad. Isot. <u>24</u>, 471 (1973).
- ²²I. W. Goodier, J. L. Makepeace, and A. Williams, Int. J. Appl. Radia. Isot. <u>25</u>, 373 (1974).
- ²³F. Ajzenberg-Selove, Nucl. Phys. <u>A320</u>, 66 (1979).
- ²⁴C. Rolfs et al. (unpublished).
- ²⁵P. J. Grant et al. (unpublished).
- ²⁶H. Liskien and A. Paulsen, Nucl. Data Tables <u>11</u>, 569 (1973).