Microscopic study of the Δ -nucleus potential from a many-body Hamiltonian for π , N, and Δ

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The strength of the Δ -nucleus potential is calculated from the many-body Hamiltonian for π , N, and Δ constructed by Betz and Lee. The results are compared to the empirical values of the Δ -hole doorway model. It is found that the calculated strength of the central potential reproduces the empirical one to a large extent. However, the calculated spin-orbit potential is much smaller than that of of the Δ -hole model. It is also shown that the pion absorption through the $N\Delta S$ wave only accounts for about half of the total absorption. The sums of other $N\Delta$ partial waves are found to be equally important.

NUCLEAR REACTIONS A model of interactions between π , N, and Δ in nuclear matter. Calculation of the Δ -nucleus potential.

The intermediate pion-nucleus interaction is dominated by Δ -isobar excitation inside the nucleus. This motivated the successful isobar-doorway model¹⁻⁵ which describes the π -nucleus dynamics in terms of Δ propagation in nuclear medium. The Δ -hole propagator contains a phenomenological complex spreading potential $W_{sp}(\vec{r})$ which reflects coupling of the doorway state to more complicated states (e.g., absorption). By a suitable adjustment 3,4 of the strength of $W_{sp}(\vec{r})$, the model yielded a satisfactory description of not only the differential cross sections of π -nucleus elastic scattering but also the gross structure of π -nucleus reaction cross sections. However, without a microscopic understanding of the spreading potential $W_{sp}(\vec{r})$, the model could not be an internally consistent one for analyzing extensive data of pion-absorption by nuclei. The purpose of this paper is to investigate a microscopic origin of $W_{sp}(\vec{r})$ based on the many-body Hamiltonian which has recently been constructed by Betz and Lee⁶ (BL).

It is necessary here to summarize briefly the content of the BL model. This model can be considered a straightforward extension of the conventional nuclear many-body theory to a higher energy region where real pion production can occur during nuclear collisions. The many-body Hamiltonian of BL is written in terms of three elementary degrees of freedom π , N, and Δ ,

$$H = H_0 + H_I , \qquad (1)$$

where

$$H_{0} = \sum_{\tau = \frac{1}{2}, \frac{3}{2}} \int d^{3}p \, c^{\dagger}(\vec{p}, \tau) [E_{\tau}(\vec{p}) - m] c(\vec{p}, \tau) + \int d^{3}k \, \alpha^{\dagger}(\vec{k}) \omega(\vec{k}) \alpha(\vec{k}) , \qquad (2)$$

$$H_{I} = \frac{1}{4} \sum_{\tau_{1}' \tau_{2}' \tau_{1} \tau_{2}} \int d^{3}p'_{1} d^{3}p'_{2} d^{3}p_{1} \, d^{3}p_{2} \, c^{\dagger}(\vec{p}_{1}', \tau_{1}') c^{\dagger}(\vec{p}_{2}', \tau_{2}') c(\vec{p}_{2}, \tau_{2}) c(\vec{p}_{1}, \tau_{1}) \langle \vec{p}_{1}' \tau_{1}', \vec{p}_{2}' \tau_{2}' | V_{0} | \vec{p}_{1} \tau_{1}, \vec{p}_{2} \tau_{2} \rangle$$

$$+ \int d^{3}k \, d^{3}p' d^{3}p [c^{\dagger}(\vec{p}', \frac{3}{2}) c(\vec{p}, \frac{1}{2}) \alpha(\vec{k}) \langle \vec{p}' \frac{3}{2} | h | \vec{p} \frac{1}{2}, \vec{k} \rangle + \text{H.c.}]$$

$$+ \int d^{3}k' \, d^{3}p' \, d^{3}k \, d^{3}p \, c^{\dagger}(\vec{p}, \frac{1}{2}) \alpha^{\dagger}(\vec{k}') c(\vec{p}, \frac{1}{2}) \alpha(\vec{k}) \langle \vec{p}' \frac{1}{2}, \vec{k}' | v | \vec{p} \frac{1}{2}, \vec{k} \rangle . \qquad (3)$$

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FIG. 1. Basic mechanism of the many-body Hamiltonian of Betz and Lee (Ref. 6).

We dropped all z components of spin and isospin to abbreviate the presentation. The creation operators for π , N, and Δ are $\alpha^{\dagger}(\vec{k})$, $c^{\dagger}(\vec{p}, \frac{1}{2})$, and $c^{\dagger}(\vec{p}, \frac{3}{2})$, respectively. The masses are μ , m, and m_{Δ} and the energies are $\omega(k) = (\mu^2 + \vec{k}^2)^{1/2}$ and $E_{\tau}(\vec{p})$ $=m_{\tau}+\vec{p}^2/2m_{\tau}$. Explicit expressions for the interactions h, v, and V_0 are given by the partial wave expanded forms Eqs. (3.2) - (3.6) and (4.36) in Ref. 6. We depict these interactions in Fig. 1. The $\Delta\Delta$ and $N\Delta \rightleftharpoons N\Delta$ parts of V_0 are omitted to simplify the model. Note that the vertex interaction h describes the $\pi N P_{33}$ interaction only. Other weaker interactions are represented by the two-body interaction v. The strategy of the BL model is to determine h, v, and V_0 phenomenologically by fitting the experimental phase shifts of πN scattering up to 300 MeV and NN scattering up to about 1 GeV, and some available data of π -deuteron reactions. The model is apparently suitable to the study of the Δ -nucleus potential.

The Δ spreading potential is parametrized⁴ as

$$W_{\rm sp}(\vec{r}) = W_0 f(r) + 2V_{LS} g(r) \vec{S}_{\Delta} \cdot \vec{L}_{\Delta} , \qquad (4)$$

where f(r) is assumed to be proportional to the nuclear density, g(r) is a surface-peaked function, \vec{S}_{Δ} and \vec{L}_{Δ} are, respectively, the spin and orbital angular momenta of Δ . Clearly, $W_{sp}(\vec{r})$ is only a part of the Δ -nucleus potential. Within the framework of the isobar-hole formalism, the Δ motion is also influenced by a binding potential and by its decay mechanism $\Delta \rightleftharpoons \pi N$. The Δ -nucleus potential or the self-energy of Δ can be calculated from the $N\Delta$ Bruckner G matrix in nuclear matter in a quite similar manner as the N-nucleus potential is calculated from the NN G matrix.⁷ Our main task is therefore to compute the $N\Delta$ G matrix in nuclear matter.

It will be helpful to give an insight into the origin



FIG. 2. The simplest absorption mechanism.



FIG. 3. Rescattering corrections to absorption.

of the spreading potential and to mention our essential assumption before going into detail. The most important contribution to the imaginary part of W_0 absorption. The simplest pion absorption is mechanism contained in the BL Hamiltonian is $\pi N \not \rightarrow N \Delta \not \rightarrow N N$ (Fig. 2). We assume that this mechanism is the dynamical origin of the spreading potential $W_{\rm sp}(\vec{r})$. One might therefore assume that the Δ self-energy corresponding to $W_{\rm sp}(\vec{r})$ should be the sum of all contributions involving 2 particle-1 hole (2p-1h) intermediate states during $N\Delta$ collisions in nuclear matter. However, in the BL model 2p-1h states can also be reached via more complicated pion absorption mechanisms as illustrated in Fig. 3. They are characterized by the decay of Δ into pion and 1 particle states and subsequent pion rescatterings before reaching 2p-1h states. We do not consider these rescattering effects on absorption.

The above assumption can be stated in mathematical expressions as follows: The starting point is the Dyson equation for the one-particle Green's function $G(\vec{p}_{\Delta}, w_{\Delta})$ of a Δ with momentum \vec{p}_{Δ} and energy w_{Δ} propagating in ordinary nuclear matter,

$$G(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) = G_0(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) + G_0(\vec{\mathbf{p}}_{\Delta}, w_{\Delta})$$
$$\times \Sigma(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) G(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) , \qquad (5)$$

where $\Sigma(\vec{p}_{\Delta}, w_{\Delta})$ is the sum of all proper self-energy diagrams in the Feynman-Dyson perturbation series.⁸ It is then grouped according to the number of nucleon-hole lines in intermediate states of each diagram. We are concerned only with terms in which at most one nucleon is excited above the Fermi sea (1*h*-line diagrams). Figure 4 is the only 0hole (0*h*) line contribution given by

$$\Sigma_{\rm res}^{(0)}(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) = \int d^{3}p \, d^{3}k \, \langle \vec{\mathbf{p}}_{\Delta} \frac{3}{2} \mid h^{\dagger} \mid \vec{\mathbf{p}} \frac{1}{2}, \vec{\mathbf{k}} \rangle$$

$$\times \frac{Q_{1}}{w_{\Delta} - E_{N}(\vec{\mathbf{p}}) - \omega(\vec{\mathbf{k}}) + i\epsilon}$$

$$\times \langle \vec{\mathbf{p}} \frac{1}{2}, \vec{\mathbf{k}} \mid h \mid \vec{\mathbf{p}}_{\Delta} \frac{3}{2} \rangle , \qquad (6)$$

where Q_1 is the Pauli operator for the πN intermediate state.



FIG. 4. 0*h*-line Δ self-energy. The nucleon particle and hole lines are, respectively, represented by \uparrow and \downarrow .

The 1h-line contributions can be separated into two different parts,

$$\Sigma^{(1)}(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) = \Sigma^{(1)}_{\text{abs}}(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) + \Sigma^{(1)}_{\text{res}}(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) .$$
(7)

 $\Sigma_{abs}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$ is the sum of all the graphs in which the external Δ lines couple directly to 2p-1h states through the interaction V_0 . Some examples of these graphs are shown in Fig. 5. These are contributions from the direct absorption mechanism indicated in Fig. 2. No $\Delta \rightleftharpoons \pi N$ vertex interactions act on the initial or final Δ . All other graphs in which at least one of the external Δ lines decays through the $\Delta \rightleftharpoons \pi N$ interaction are summarized as $\Sigma_{res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$ and some of them are illustrated in Fig. 6. From its definition $\Sigma_{\rm res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$ can be considered higher order terms of the simplest self-energy Eq. (6) and can be renormalized into the latter. Indeed all of these rescattering diagrams can be cast in the 1-h line contributions to the renormalization effects on the $\Delta \rightleftharpoons \pi N$ vertex, excepting Figs. 6(a) and (b) which are nucleon and pion propagator renormalizations. In particular, all the contributions to the Δ selfenergy coming from absorption as described by Fig. 3 are $\Delta \rightleftharpoons \pi N$ vertex corrections. One sees that the absorption mechanism of Fig. 3 is mixed with the direct absorption Fig. 2 to make up the vertex correction diagram Fig. 6(c).

It is seen above that the Δ self-energy contains two distinctive dynamics; rescattering and absorp-



FIG. 5. 1*h*-line Δ self-energy coming from the sum of all the interactions between two nucleon particles in the 2p-1h states sandwiched by V_0 .



FIG. 6. 1*h*-line Δ rescattering diagrams which are renormalization corrections to the 0*h*-line diagram Fig. 4. The blob in (a) is for the NN G matrix defined by Eq. (10d).

tion. We calculate each of them to lowest order in hole-line expansion, namely, rescattering to $\Sigma_{\rm res}^{(0)}(\vec{p}_{\Delta}, w_{\Delta})$ and absorption to $\Sigma_{\rm abs}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$. Inclusion of $\Sigma_{\rm res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$ does not necessarily produce an improvement upon $\Sigma_{\rm res}^{(0)}(\vec{p}_{\Delta}, w_{\Delta})$. If we include $\Sigma_{\rm res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$, it introduces modifications of the $\Delta \rightleftharpoons \pi N$ vertex so that the isobar-hole interaction Was well as the Pauli blocking term δW of $\Sigma_{\rm res}^{(0)}(\vec{p}_{\Delta}, w_{\Delta})$ must also be modified. To be consistent with the analyses^{3,4} using the simplest onepion exchange isobar-hole interaction $W + \delta W$, it seems reasonable to neglect $\Sigma_{\rm res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$. The nucleon propagator renormalization term Fig. 6(a) contributes chiefly to the real part of the Δ binding potential. We also ignore this effect (we will discuss this term later).

The 1h-line absorption self-energy as defined above can be written as

$$\Sigma_{abs}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$$

$$= \int_{p < p_{F}} d^{3}p \langle \vec{p}_{\Delta} \frac{3}{2}, \vec{p} \frac{1}{2} | G(w = w_{\Delta} + \epsilon_{N}(\vec{p})) |$$

$$\times \vec{p}_{\Delta} \frac{3}{2}, \vec{p} \frac{1}{2} \rangle. \qquad (8)$$

G(w) is the sum of all interactions between two baryons sandwiched by V_0 . The summation is performed above the Fermi sea applying Bruckner's method. The resulting equation for G(w) in the two-baryon subspace $NN \oplus N\Delta$ is

$$G(w) = V_0 + V_0 \frac{Q_2}{w - H_0 - V_D(w)} G(w) + V_0 \frac{Q_2}{w - H_0 - V_D(w)} G_c(w) \times \frac{Q_2}{w - H_0 - V_D(w)} G(w)$$
(9a)

with

$$G_{c}(w) = \widetilde{V}_{c}(w) + \widetilde{V}_{c}(w) \frac{Q_{2}}{w - H_{0} - V_{D}(w)} G_{c}(w) ,$$
(9b)

$$\tilde{V}_{c}(w) = V_{E}(w) + V_{B}(w)$$
 (9c)

Here Q_2 is the Pauli operator of the subspace $NN \oplus N\Delta$ and the effective $N\Delta$ interactions $V_D(w)$, $V_E(w)$, and $V_B(w)$ are, respectively, defined as

$$V_D(w) = \sum_i h_i^{\dagger} \frac{Q_3}{w - H_0} h_i , \qquad (10a)$$

$$V_E(w) = \sum_{i \neq j} h_i^{\dagger} \frac{Q_3}{w - H_0} h_j , \qquad (10b)$$

$$V_B(w) = \sum_i h_i^{\dagger} \frac{Q_3}{w - H_0} G_{NN}(w) \frac{Q_3}{w - H_0} h_i , \quad (10c)$$

where h_i stands for the $\Delta \rightleftharpoons \pi N$ vertex of the *i*th baryon. Q_3 is the Pauli operator for πNN intermediate states and $G_{NN}(w)$ is the NN G matrix calculated in the presence of a spectator pion,

$$G_{NN}(w) = V_0 + V_0 \frac{Q_3}{w - H_0} G_{NN}(w)$$
 (10d)

The effective interactions $V_D(w)$, $V_E(w)$, and $V_B(w)$ are shown by the diagrams in Fig. 7. Within the BL model, it is found that the effect of $V_B(w)$ is small except in the $N\Delta$ S wave. Furthermore, $V_B(w)$ is the interaction appearing in calculating the graph Fig. 6(a) in $\Sigma_{\rm res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$. We drop this interaction for evaluation of G(w) in accordance with the neglect of $\Sigma_{\rm res}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$.

The structure of Eqs. (9) and (10) is identical with that of the coupled two-baryon scattering equation in Ref. 6 [see its Eqs. (4.11) and (4.18)], the only difference being the presence of the Pauli operators. The angular-averaged Pauli operators⁹ are used in our calculation. Then the numerical method of Ref. 6 can be used to solve the *G* matrix equation in the partial wave representation. No details will therefore be given here except to note that the *G* matrix depends on the total momentum of the two baryons as well as the starting energy.

Following the current practice,⁷ we calculate the central part of $W_{\rm sp}(\vec{r})$ from the 1*h*-line absorption self-energy in infinite nuclear matter as $W_0 = \Sigma_{\rm abs}^{(1)}(\vec{p}_{\Delta}, w_{\Delta})$. In terms of the partial wave expanded G matrix (using the convention of Ref. 6), W_0 as a function of p_{Δ} is written as

$$W_{0}(p_{\Delta}) = (2\pi)^{3} \frac{3}{4\pi p_{F}^{3}} \rho_{0} \sum_{L_{\Delta}S_{\Delta}JT} \frac{(2J+1)(2T+1)}{64} \frac{1}{4\pi} \int_{p < p_{F}} d^{3}p \, G_{L_{\Delta}S_{\Delta},L_{\Delta}S_{\Delta}}^{JT}(q,q;W(\vec{p}_{\Delta},\vec{p})) , \qquad (11)$$

where

$$q = \left| \frac{m_{\Delta} \vec{p} - m \vec{p}_{\Delta}}{m + m_{\Delta}} \right|, \qquad (12)$$

$$W(\vec{p}_{\Delta},\vec{p}) = \epsilon_N(\vec{p}) + \epsilon_{\Delta}(\vec{p}_{\Delta}) - \frac{(\vec{p}+\vec{p}_{\Delta})^2}{2(m+m_{\Delta})} .$$
(13)

For an arbitrary choice of f(r) in Eq. (4), ρ_0 is given by

$$\rho_0 = A / 4\pi \int_0^\infty r^2 dr f(r) .$$
 (14)

If f(r) is taken to be the standard Woods-Saxon form, ρ_0 corresponds to the density of nuclear matter with the Fermi momentum p_F . The single particle energies of N and Δ are specified as follows: Following the standard nuclear matter theory,¹⁰ we choose

$$\epsilon_{N}(\vec{p}) = m + \frac{\vec{p}^{2}}{2m^{*}} + U_{0}, \quad p \le p_{F} ,$$

$$= m + \frac{\vec{p}^{2}}{2m}, \quad p > p_{F} . \qquad (15)$$

The Δ -single particle energy is determined selfconsistently⁷

$$\epsilon_{\Delta}(\vec{p}_{\Delta}) = m_{\Delta} + \frac{\vec{p}_{\Delta}^{2}}{2m_{\Delta}} + \operatorname{Re}\left[\Sigma(\vec{p}_{\Delta},\epsilon_{\Delta}(\vec{p}_{\Delta}))\right].$$
(16a)



FIG. 7. Effective $N\Delta$ interactions defined by Eqs. (10a)-(10c).

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Note that the total Δ self-energy

$$\Sigma(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) = \Sigma_{\text{res}}^{(0)}(\vec{\mathbf{p}}_{\Delta}, w_{\Delta}) + \Sigma_{\text{abs}}^{(1)}(\vec{\mathbf{p}}_{\Delta}, w_{\Delta})$$
(16b)

enters into the self-consistent equation (16a).

In order to determine the strength of the spin-

orbit potential, we must derive the Δ self-energy $\Sigma(\vec{p}'_{\Delta}, \vec{p}_{\Delta}; w_{\Delta})$ in finite nuclei and extract a term proportional to $\vec{S}_{\Delta} \cdot \vec{p}'_{\Delta} \times \vec{p}_{\Delta}$. If we assume that the momentum transfer $\vec{p}'_{\Delta} - \vec{p}_{\Delta}$ caused by the nuclear density variation near the nuclear surface is small, it is straightforward to express the spin-orbit strength in terms of G(w) using a Thomas-Fermi type approximation.¹¹ The result is

$$V_{LS}(p_{\Delta}) = (2\pi)^{3} \frac{3}{4\pi p_{F}^{3}} \frac{\rho_{0}}{20r_{0}^{2}p_{\Delta}^{2}} \sum_{L_{\Delta}S_{\Delta}JT} \frac{(2J+1)(2T+1)}{64} \frac{S_{\Delta}(S_{\Delta}+1)+3}{S_{\Delta}(S_{\Delta}+1)} [J(J+1)-L_{\Delta}(L_{\Delta}+1)-S_{\Delta}(S_{\Delta}+1)] \\ \times \frac{1}{4\pi} \int_{P < p_{F}} d^{3}p \frac{\vec{p}_{\Delta} \cdot \vec{q}}{q^{2}} G_{L_{\Delta}S_{\Delta},L_{\Delta}S_{\Delta}}^{JT} (q,q; W(\vec{p}_{\Delta},\vec{p})),$$
(17a)

with

$$g(r) = \frac{r_0^2}{r} \frac{d}{dr} f(r)$$
 (17b)

 r_0 is any constant length to make g(r) dimensionless.

The spreading potential thus obtained is dependent on the Δ momentum p_{Δ} , namely, it is nonlocal. To compare with the phenomenological local potential determined by Horikawa, Thies, and Lenz,⁴ we must construct an equivalent local potential. As Negele and Yazaki¹² have recently pointed out, the nonlocality of the optical potential influences the equivalent local potential. However, since the resulting W_0 is not a rapidly varying function of p_{Δ} in the relevant momentum region $p_{\Delta} \leq 400$ MeV/c, we do not consider effects of the nonlocality here but obtain a local potential by choosing the incident pion momentum

$$k_{\pi} = [(E_{\pi} + \mu)^2 - \mu^2]^{1/2}$$

as an average value of p_{Δ} (E_{π} is the pion kinetic energy).

The parameters used in the calculation are $\rho_0 = 0.17 \text{ fm}^{-3}$, $p_F = 1.36 \text{ fm}^{-1}$, $r_0 = 1.1 \text{ fm}$, $U_0 = -80 \text{ MeV}$, and $m^* = 0.68 \text{ m}$. The results of $W_0(p_{\Delta})$ and $V_{LS}(p_{\Delta})$ are shown in Figs. 8 and 9. Only the imaginary part of W_0 can be directly compared to that of the isobar-hole model.⁴ Comparison of other parts of the potential can meaningfully be made only in the shape-independent quantities

$$\Omega = \frac{4\pi}{A-1} \int_0^\infty r^2 dr \ W_0 f(r)$$

= $\frac{A}{A-1} \frac{W_0}{\rho_0}$, (18)

$$S = \int_{0}^{\infty} r \, dr \, V_{LS}g(r)$$

= $-V_{LS}r_{0}^{2}f(0)$, (19)

where f(0)=1. These are shown in Figs. 10 and 11.

From Fig. 8 it is seen that the calculated strength of $\text{Im}W_0$ agrees qualitatively with the phenomenological one.⁴ Theoretical values amount to about 85% of the empirical ones on the average. As shown in Refs. 3 and 4, the imaginary spreading potential can account for the measured total pionabsorption cross section. The agreement obtained here indicates that the dominant absorption



FIG. 8. Strength of the Δ central potential, $W_0(p_{\Delta})$. The empirical values are taken from Ref. 4 by estimating the Δ momentum p_{Δ} from the pion momentum k_{π} .



FIG. 9. Strength of the Δ spin-orbit potential, $V_{LS}(p_{\Delta})$.

mechanism must be the process shown in Fig. 2. This interpretation must be carefully distinguished from that of many-nucleon absorption implied essentially by a kinematic analysis of inclusive pion absorption data.¹³ There involves a nontrivial and model-dependent problem of separating initial and final state interactions from the absorption mechanism. The only way to clarify the situation seems to be to study the consequences of the BL Hamiltonian (1)-(3) using the two-nucleon absorption approach of this work and to analyze the data of Ref. 13 and others. The research in this direction will be discussed elsewhere.

The calculated Ω for ¹²C and ¹⁶O (Fig. 10) agrees qualitatively with that of Ref. 4. The discrepancy for ⁴He is significant but the use of the *G* matrix in nuclear matter and the local density approximation for such a light nucleus is very questionable. It is



FIG. 10. Volume integral Ω of the Δ central potential. The empirical values are taken from Ref. 4.



FIG. 11. Surface integral S of the Δ spin-orbit potential strength function. The empirical values are taken from Ref. 4.

important to note that in our calculation the real part of Ω comes from the 1*h*-line absorption diagram, while that of Ref. 4 also includes the contribution from a real Δ binding potential. It would seem that the 1*h*-line absorption contains the most important physics of the central part of the Δ -nucleus potential.

Contrary to this, the calculated S for the spinorbit interaction (Fig. 11) is much smaller than the empirical values. According to Ref. 4, the Δ nucleus spin-orbit interaction is roughly the same as the nucleon-nucleus spin-orbit interaction. Our result $2V_{LS} \approx 8$ MeV at $p_{\Lambda} = 262$ MeV/c is less than half of the N-nucleus spin-orbit strength $\approx 17 \text{ MeV}$ (Ref. 14). The phenomenological procedure for extracting the spin-orbit interaction suffers from larger uncertainties (see discussions in Ref. 4) and the discrepancy we have found will not easily be resolved until the freedom of the analyses is reduced by imposing additional constraints. Of course approximations we have made might be incorrect for calculating the spin-orbit interaction. The neglect of Fig. 6(a) could be partially responsible for the disagreements. The second-order diagrams with respect to the G matrix can also be sources of the Δ spin-orbit coupling as in the nucleon case.¹⁵ The origin of the nucleon-nucleus spin-orbit force is one of the most fundamental and important problems in nuclear physics but is most poorly understood. Much work will have to be done for the consistent understanding of both of the N- and Δ -spin-orbit interactions.

TABLE I. Contributions of important partial waves to W_0 and W_{LS} at $p_{\Delta} = 261.74$ MeV/c.									
Channels		ReW ₀	Im W ₀	ReV _{LS}	Im V _{LS}				
NΔ	NN	(MeV)	(MeV)	(MeV)	(MeV)				

Channels		ReW_0	$\operatorname{Im} W_0$	${\rm Re}V_{LS}$	$Im V_{LS}$
NΔ	NN	(MeV)	(MeV)	(MeV)	(MeV)
⁵ S ₂	${}^{1}D_{2}$	-3.28	-18.30	0	0
${}^{3}P_{0}$	${}^{3}P_{0}$	-8.20	-2.30	2.13	0.59
${}^{3}P_{1} + {}^{5}P_{1}$	$^{3}P_{1}$	-4.86	-0.66	1.01	0.14
${}^{3}P_{2} + {}^{5}P_{2}$	${}^{3}P_{2}+{}^{3}F_{2}$	-11.18	-5.14	-0.19	0.03
⁵ P ₃	${}^{3}F_{3}$	-4.17	-4.74	-0.66	-0.74
${}^{5}D_{0}$	${}^{1}S_{0}$	-4.36	-0.76	2.05	0.35
⁵ D ₄	${}^{1}G_{4}$	-0.72	-0.73	-0.23	-0.23
	Total	- 36.76	-32.62	4.12	0.14

To examine the absorption mechanism in more detail, we present in Table I the contributions of important $N\Delta$ partial waves to W_0 and V_{LS} . An essential point emerges here. Although the $N\Delta$ partial wave ${}^{5}S_{2}$ is the biggest contribution (-18.30 MeV) to $Im W_0$, the sum of other partial waves (-14.30 MeV) is of the same magnitude. This result is somewhat unexpected but is actually not surprising because the inelasticities of the NN partial waves listed in Table I are not small (see Fig. 3 of Ref. 6). Our results indicate the importance of treating absorption consistently with the twonucleon data. Although some geometrical factors favor absorption through the $N\Delta$ ⁵S₂ wave, calculations of absorption must be carried out with a careful treatment of all partial waves. Equal weight of all partial waves are also clearly exhibited in their contributions to $\text{Re}W_0$ and V_{LS} shown in Table I.

We have provided a microscopic picture of the

phenomenological spreading potential $W_{\rm sp}(\vec{r})$ entering the isobar-hole model. We should mention here that the isobar-hole model^{3,4} can rigorously be formulated from the BL many-body Hamiltonian. In this sense the BL model furnishes us with an internally consistent description of π -nucleus interaction with the πN and NN data as inputs to the model. It is hoped that in the future the π -nucleus data can be understood in a deeper theoretical approach by combining the ideas behind the construction of the isobar-hole model and the many-body Hamiltonian.

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