Beta decay of ²⁴Al

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The β^+ decay of ²⁴Al has been studied with a Ge(Li)-NaI(Tl) escape-suppression spectrometer. The number of ²⁴Mg γ transitions observed following ²⁴Al decay has been increased from 19 to 44. Definite new β^+ branches were observed to ²⁴Mg levels at 105 76, 108 21, and 113 14 keV. An enhanced *E* 2 transition from the 4⁺ 9301-keV level to the 2⁺ 7349-keV level was sought for and found. A probable 8439 \rightarrow 7349 4⁺ \rightarrow 2⁺ transition was also observed. The contribution of these new experimental results to an elucidation of the band structure of ²⁴Mg is discussed.

 $\begin{bmatrix} \text{RADIOACTIVITY} & ^{24}\text{Al [from } ^{24}\text{Mg}(p,n)]; \text{ measured } E_{\gamma}, I_{\gamma}, \beta^{+} \text{ and } \gamma \text{ branchings}; \\ \text{deduced } B(E2), \log ft. \end{bmatrix}$

I. INTRODUCTION

With the possible exception of ²⁰Ne, ²⁴Mg has been the most important even-even sd-shell nucleus for testing nuclear models. In particular, early tests of various manifestations of collective motion were made in ²⁴Mg. These include the Nilsson model, SU(3) symmetries, and Hartree-Fock calculations.^{1,2}

Recently shell-model calculations have been carried out for ²⁴Mg in a complete sd-shell basis by both the Glasgow group^{3,4} and by Wildenthal and his co-workers.^{5,6} These calculations which use the Glasgow shell-model code and the Preedom-Wildenthal⁵ or Chung-Wildenthal⁶ effective interaction appear to be very successful, although very little besides energy spectra have been published so far. In the model space used in these calculations, ²⁴Mg consists of eight nucleons in the sd shell outside an inert ¹⁶O core. The number of active nucleons is large enough that a physical interpretation of the calculations is difficult. An obvious help in understanding the predictions is to express the shell-model wave functions in a collective framework such as the Nilsson model or in SU(3) symmetries. For example, expansion in terms of SU(3) symmetry has been done with several different interactions and model spaces for the low-lying levels of the K=0 ground-state band and the K = 2 band built on the 2⁺ level at 4238 keV. An example is the work of Wathne and Engeland.⁷ In brief, the theoretical evidence for K = 0 and K = 2bands in ²⁴Mg is overwhelming. At the same time the experimental information² on the states belonging to these bands is relatively complete and also strongly indicative of band structure.

Our motivation for the present experimental study was to address the question: Are there other even-parity bands based on bound levels of ^{24}Mg , and, if so, how well developed are they? In trying to answer this question one finds that both the

spectroscopic information and theoretical predictions on these bands are quite limited. First, the spectroscopic information available on the levels in question is much more scarce than for the K=0and K=2 bands.² Second, the theoretical predictions available as a guide to future experiments on the higher-lying states of these excited bands are quite meager.

The search for further even-parity bands in ²⁴Mg has a natural starting point with the first excited $J^{*} = 0^{+}$ state which is at 6432 keV. If this is the band head of a $K = 0^+$ band then possibilities for the 2^+ state of this band are at 7348 and 8653 keV. In a previous search for a band built on the 6432-keV level the possibility that the 8653-keV level was the 2⁺ member of this band was explored.⁸ A search for the 4⁺ level was unsuccessful. We shall explore the evidence for the placement of the J^{r} $= 2^+$ 7348-keV level in this band. There are two pieces of information which suggest a connection between the 2⁺ 7348- and 0⁺ 6432-keV levels. First, Feldmeier, Manakos, and Wolff⁹ in a SU(3)truncated shell-model calculation gave a prediction of 15 Weisskopf units (W.u.) (Ref. 10) for the 7348 - 6432 E2 transition. This is certainly large enough to suggest an intraband transition. Second, in their breakdown of shell-model wave functions into SU(3) symmetries, Wathne and Engeland⁷ found the composition of these two states to be remarkably similar. Unfortunately, the branching ratio of the 7348 - 6432 transition is expected to be extremely small and an experiment to measure this E2 rate-even if highly enhanced-would be very difficult.

Recent experiments¹¹ have determined $J^{r} = 4^{+}$ for both the 8437- and 9298-keV levels of ²⁴Mg. These are then the third and fourth $J^{r} = 4^{+}$ levels of ²⁴Mg with the first two belonging to the K = 0 and K = 2bands, respectively. As pointed out by Endt and van der Leun² and by Wright,¹² both of these levels are members of energy doublets/triplets and the

(1)

difficulty of untangling these multiplets has kept the properties of these levels from being well determined.

The present study is of the β^* decay of $J^{\pi} = 4^{+24}$ Al. Allowed β^* decay will take place to $J^{\#}=3^+$, 4^+ , and 5⁺ states of ²⁴Mg. We use the specificity of this reaction to select the 4⁺ levels at 8.44 and 9.30 MeV with negligible population of the other members of the doublet at 8.44 MeV and the triplet at 9.30 MeV. It has already been found that both levels are formed in ${}^{24}Al(\beta^*){}^{24}Mg.{}^{2,13,14}$ Our specific aim was to search for γ decays from these two levels to lower-lying 2^+ states.

II. EXPERIMENTAL PROCEDURE AND ANALYSIS

²⁴Al with $T_{1/2} = 2.060 \pm 0.010$ s (Ref. 15) was produced via the ²⁴Mg(p, n)²⁴Al reaction (Q = -14.661MeV).² A 99.94% pure ²⁴Mg foil, 3.8×10^{-3} cm thick, was mounted on a Delrin rabbit. The ²⁴Mg was bombarded through a 2.5×10^{-3} cm thick Ta entrance window with a 75-nA, 18-MeV proton beam. The rabbit system, described previously,¹⁶ allowed cycling of the ²⁴Mg between the bombard position and a count position which was 4 m away and on the other side of a concrete wall. A bombard-count cycle was as follows:

0-2.0 s, bombard 2.0-2.5 s, transfer

4.5-5.0 s, transfer.

Gamma-ray spectra were collected using a Ge(Li)-NaI(T1) escape-suppression (anti-Compton) spectrometer. The Ge(Li) detector had an efficiency of 19.5% and a resolution of 2.1 keV for 1332-keV γ rays. It was placed at the center of a 25.4×20.3 cm NaI(Tl) veto detector and was 9.0 cm from the ²⁴Al source. The configuration of the shield was similar to that reported by Konijn, Goudsmit, and Langeman.¹⁷

In order to stop positrons emanating from the target and to suppress low-energy γ rays and x rays, a plug consisting of 0.72 cm of brass followed by 0.66 cm of lead was inserted in the 5-cm long conical lead collimator between the ²⁴Al source and the Ge(Li) detector.

Two sets of 8192-channel spectra were recorded simultaneously by splitting the output from the Ge(Li) preamplifier. One had a dispersion of 0.51 keV/channel and extended to 4185 keV, the other had a dispersion of 1.41 keV/channel and extended to 11 588 keV. Each set consisted of the total counts (normal singles spectra) and a spectrum of those counts rejected by the NaI(Tl) shield. The anti-Compton spectra which were analyzed were

obtained by subtracting the rejected spectra from the total spectra. The reduction of the Compton background was by factors of 5-8, with maximum reduction for γ rays of energy 5-6 MeV. Retaining the rejected spectra was an aid in identifying oneand two-escape peaks which were not entirely eliminated by the anticoincidence condition.

Data were accumulated for 17.5 h. The final low-gain anti-Compton spectrum contained a total of 56×10^6 counts and the corresponding rejected spectrum contained 78×10^6 counts. The energy resolution for all the accumulated spectra varied from 2.28-keV full width at half maximum (FWHM) at 1369 keV to 5.75-keV FWHM at 7069 keV. Gamma-ray intensities and peak positions were extracted using the program SAMPO.¹⁸ Portions of the low-gain spectra are illustrated in Figs. 1-3. Figure 1 illustrates the improvement in peak/background offered by the escape suppression, while Figs. 2 and 3 illustrate the evidence for some previously unobserved cascades.

The energy calibration for the ²⁴Al spectra was arrived at in several successive steps. First, the nonlinearity of the high-gain spectrum $(E_{\star} < 4185 \text{ keV})$ was determined from a separate ⁵⁶Co calibration. The energies for $E_{\gamma} < 2900$ keV were obtained using as energy standards the ²¹Ne 350.725(6) 1 - 0 transition^{19,20} [from ²⁴Mg- $(p, \alpha)^{21} Na(\beta^{+})^{21} Ne$ and the ²⁴Mg 1 - 0 and 2 - 1 transitions of 1368.633(6) and 2754.030(14) keV as given by Greenwood, Helmer, and Gehrke.²¹ Then utilizing the relations between ²⁴Al cascade and crossover γ rays, the calibration was extended in several successive steps to $E_{\gamma} < 3900 \text{ keV}, < 5000$ keV, and finally to the whole energy region of interest. For the higher energies, one- and twoescape peaks from the rejected spectrum were also used in the energy determinations.

The efficiency versus energy curve for the anti-Compton spectrum was obtained from a ⁵⁶Co spectrum taken after the ²⁴Al data were accumulated. This gives an accurate efficiency calibration for the energy interval ~ 700-3700 keV.^{21,22} One further calibration point at 7069 keV was provided by the low-gain ²⁴Al spectrum as follows. The $J^{*} = 4^{+24}$ Al ground state must have a negligible direct branch to the 2^{+ 24}Mg 1369-keV level (i.e., $\Delta J = 2$, no β decay is twice forbidden), so that the γ -ray flux into the 1369-keV level must be equal to the γ -ray flux out. Then, anticipating our results, some 33% of the ²⁴Al γ flux is in the ²⁴Mg 7069-keV 8439 - 1369 transition and almost all of the remaining feeding of the 1369-keV level is via the 2754keV 4123-1369 transition for which the efficiency is accurately obtained from the ⁵⁶Co efficiency calibration. Thus, the relative efficiency at 7069 keV is accurately provided ($\leq \pm 2\%$) from this in-

<u>23</u>

2.5-4.5 s, count



CHANNEL NUMBER

FIG. 1. Partial γ -ray spectra from ²⁴Al $(\beta^{+})^{24}$ Mg. Singles data (upper) are shown as well as singles minus the rejects (lower). The peaks are labeled by their energies in keV and are identified in Table I. Note the reduced background in the anti-Compton spectrum, as well as the reduction of one- and two-escape peaks, which are labeled as (1) and (2), respectively.

ternal calibration. The form of the efficiency versus energy curve was assumed to be exponential for $E_{\gamma} > 3.5$ MeV. Justification for this is provided by previously published efficiency curves such as the data of Young *et al.*²³ The relative ef-



FIG. 2. Portions of the anti-Compton γ spectrum from $^{24}\mathrm{Al}(\beta^{*})^{24}\mathrm{Mg}$ illustrating the evidence for weak, previously unobserved transitions. The γ peaks are labeled by their energies in keV and are identified in Table I. For purposes of display, 12 000 counts have been subtracted from the spectrum on the left. The 1060-, 1091-, and 5340-keV transitions were not previously observed.

ficiency was assumed to have a $\pm 2\%$ accuracy for $800 < E_{\gamma} < 3500$ keV and at 7069 keV, a $\pm 10\%$ accuracy at 5300 keV, and a $\pm 20\%$ accuracy at 10000 keV. The accuracy was assumed to be given by a





Energy ^a (keV)	Assignment ^b	Relative intensity ^c (%)	Energy ^a (keV)	Assignment ^b	Relative intensity ^c (%)
426.00(10)	24 A1(426 \rightarrow 0) ^d	0.29(4)	2869,50(6)	4 238 → 1 369	1.097(28)
587.95(6)	U	0.149(8)	3203.88(8)	8439 ightarrow 5235	3.085(66)
775.40(20)	6010 ightarrow 5235	0.053(8)	3378,27(80)	7 61 6 - 4238	0.043(7)
822,00(60)	$8439 \rightarrow 7616$	0.021(8)	3493.32(N)	$7\ 616 \rightarrow 4123$	0.04(1)
860.20(60) ^e	$9301 \rightarrow 8439(Q)$	0.022(11)	3505.61(9)	$9516 \rightarrow 6010$	1.98(6)
909.09(6)	U	0.122(6)	3866.14(10)	$5235 \rightarrow 1369$	5.26(22)
996.83(10)	5235 ightarrow 4238	0.137(7)	4200,54(13)	8439 ightarrow 4238	4.02(22)
1059.78(8)	$10576 \rightarrow 9516$	0.285(17)	4237.96(6)	$4238 \rightarrow 0$	3.61(21)
1076.86(4)	9516 ightarrow 8439	14.84(31)	4280.62(13)	$9516 \rightarrow 5235$	0.66(4)
1090.67(10)	$8439 \rightarrow 7349$	0.140(7)	4316.00(12)	$8439 \rightarrow 4123$	14.20(86)
1274.71(10) ^f	$10576 \rightarrow 9301$	0.106(6)	4641.19(9)	$6010 \rightarrow 1369$	3.42(25)
1368.633(6)	$1369 \rightarrow 0(C)$	96.0(2.5)	5060.68(80)	9301 ightarrow 4238	0.036(13)
1434.11(20)	U	0.063(6)	5177,51(20)	9 301 -+ 4123	0.98(10)
1704.77(80)	$9516 \rightarrow 7812$	0.016(4)	5340.30(40)	$10576 \rightarrow 5235$	0.115(13)
1771.92(7)	$6010 \rightarrow 4238$	0.40(1)	5392.68(9)	$9516 \rightarrow 4123$	18.3(18)
1887.52(20) ^g	6010 ightarrow 4123	0.056(6)	5979.52(80)	7349 ightarrow 1369	0.093(9)
1899.70(6)	$9516 \rightarrow 7616$	0.82(2)	6246,89(11)	$7\ 616 \rightarrow 1369$	0.54(4)
1952.38(20)	9 301 → 7349	0.094(6)	7069.50(12)	8439 ightarrow 1369	43.0(1.3)
2127.51 (20)	U	0.054(7)	7348.17(90)	7 349 → 0	0.153(16)
2136.58(15)	$10 576 \rightarrow 8439$	0.168(9)	7615,17(90)	7 616 → 0	0.224(15)
2381.03(30)	$7\ 616 \rightarrow 5235$	0.037(10)	7930.87(15)	$9301 \rightarrow 1369$	1.34(10)
2428.97(15)	8439 ightarrow 6010	0.774(18)	8085.66(N)	$9456 \rightarrow 1369(Q)$	0.020(10)
2566.96(20)	U	0.065(7)	8146.04(N)	$9516 \rightarrow 1369$	0.028(7)
2577.44(80)	$7812 \rightarrow 5235(Q)$	0.030(12)	9450.09(40)	$10821 \rightarrow 1369$	0.110(20)
2754.030(14)	$4123 \rightarrow 1369(C)$	41.19(90)	9943.46(150)	$11314 \rightarrow 1369$	0.027(6)

TABLE I. Gamma rays observed from ${}^{24}Al(\beta^*){}^{24}Mg$.

^aUncorrected for nuclear recoil. The number in parentheses is the uncertainty in the last figure. An N means the γ -ray energy is calculated from the transition energy.

^{b 24}Mg $E_i \rightarrow E_f$ in keV. U denotes an unknown, Q a questionable assignment, C a γ ray used for energy calibrations. ^cNormalized such that the flux into the ²⁴Mg ground state is 100.00.

^dThe dominant decay mode of ²⁴Al^m (see text).

^ePossibly a doublet with an unknown contaminant.

^fContains an unknown contribution from the ²²Na(β)²²Ne 1274-keV γ ray.

⁸Possibly contains a contribution from the ²⁴Mg 10328 - 8439 transition.

linear interpolation between these points for E_{γ} > 3500 keV. The energies, relative intensities, and decay scheme assignments of the observed γ rays are listed in Table I.

In all, the energies and intensities of 57 definite γ rays were extracted from the anti-Compton spectrum. Forty-four of these were assigned to ²⁴Al β^+ decay. Eight others were assigned to various obvious activities such as ²³Mg(β^+)²³Na from ²⁴Mg(p, d)²³Mg or ⁴⁰K(*EC*)⁴⁰Ar. Five unidentified lines are listed in Table I.

The bombard-count cycle discriminated strongly against γ rays from the decay of ²⁴Al^m which has a half-life of 129 ± 5 ms.² The dominant decay mode of this $J^{*}=1^{*}$ metastable state is via M3 decay to the 4⁺ ground state. The energy we observe for this transition, 426.00 ± 0.10 keV, is in only fair agreement with the value of 425.8 ± 0.1 keV reported by Houkanen *et al.*¹⁴ From the observed intensity of this branch (see Table I) and previous relative intensities¹⁴ all other branches from ²⁴Al^m are inferred to be negligible in the present study.

The level energies and γ -ray branching ratios derived from the results of Table I are collected in Table II. The various γ -ray combinations contributing to a determination of the level energies were in good agreement and the uncertainties assigned to the level energies are considered to be conservative. The assignments of γ -ray peaks to ²⁴Mg transitions were first made on the basis of energy and from consideration of previous spectroscopic information. Since the energy determinations were quite accurate, the target impurities were minimal, and the number of ²⁴Mg levels directly fed via ²⁴Al is small, it is unlikely that any transition is misassigned or is a doublet. All possible transitions between the known ²⁴Mg levels below 9-MeV excitation and states for $9 < E_x < 11$ MeV with known or possible 3^+ , 4^+ , or 5^+ assign-

Л	Ei ^a (keV)	E_f (keV)	Ε _γ (keV)	Brar Pres	nching ent	ratio (%) Previous ^b	J	Ei ^a (keV)	E _f (keV)	Eγ (keV)	Branchin Present	g ratio (%) Previous ^b
2*	1 368.675(6)	0	1369	100		100			1369	7931	54.0 ±2.	57 ± 5^{d}
4+	4122.875(15)	0	4123			<10 ⁻³			4123	5177	40.0 ±1.	43 ± 5^{d}
		1369	2754	100		100			4238	5061	1.5 ± 0.3	3 <5 ^e
2*	4238.360(60)	0	4238	76.7	±0.7	78.9 ± 0.5			5235	4065	<2.0	
		1369	2870	23.3	±0.7	21.2 ± 0.5			6010	329 0	<2.4	
3+	5235.200(80)	0	5235	<0.4		<0.04			7349	1952	4.1 ± 0.4	Ł
		1369	3866	97.3	±0.3	98.2 ± 0.2			7616	1684	<0.8	
		4123	1112	<0.7		<0.5			8439	860	<0.9 ±0.	2
		4238	997	2.7	±0.3	1.8 ± 0.2	4*	9 516,210(80)	0	9514	<0.05	<5
4+	6 010.315(90)	0	6010	<0.8		<2			1369	8146	0.08 ± 0.0	04 < 0.2
		1369	4641	86.6	±1.2	87 ± 1			4123	5393	49.5 ±1.	2 50 ± 3
		4123	1887	≪2.0	±0.39	[;] <1			4238	5277	<0.7	1.0 ± 0.5
		4238	1772	10.1	±1.0	13 ± 1			5235	4281	1.8 ± 0.3	2 3.8 ± 0.4
		5235	775	1.3	±0.2	<1			6010	3506	5.3 ± 0.3	3 6±1
2+	7 349.05(40)	0	7348	62.0	±2.4	62 ± 2			7349	2167	<0.07	
		1369	5980	38.0	±2.4	38 ± 2			7 61 6	1900	2.3 ± 0.3	2 1.5±0.5
3	7 616.470(90)	0	7615	25.3	±2.0	23 ± 2			7812	1704	0.04 ± 0.0)2
		1369	6247	61.4	±2.0	72 ± 2			8439	1077	41.0 ± 1.0	38 ± 3
		4123	3493	4.3	±0.8	5 ± 1	(5)	⁺ 10 575.983(80) ^f	0	10573	<1.5	<2
		4238	3378	4.9	± 0.8	<5			1369	9205	<2	
		5235	2381	4.1	±0.7				4123	6452	<4	<10
		6010	1606	<3		<2			4238	6337	<4	<15
4*	8 439.295(90)	0	8438	<0.08	5	<2			5235	5340	17.1 ± 1.9	28 ± 10
		1369	7069	65.9	±1.5	63 ± 3			6010	4565	<3	13 ± 4
		4123	4316	21.7	±0.9	23 ± 3			7349	3227	<2	<10
		4238	4201	6.2	±0.3	7 ± 1			7616	2959	<4	$(11) \pm 4^{g}$
		5235	3204	4.8	± 0.2	6 ± 1			8439	2137	24.9 ±1.	3 <25 ^h
		6010	2429	1.2	±0.1				9301	1275	≤15.7 ±0.5	9 ⁱ <7

TABLE II. ²⁴Mg γ -ray transitions observed in the β^+ decay of ²⁴Al.

^aThe number in parentheses is the uncertainty in the least significant figure. The energies of the first two states are from Ref. 21. The other energies were calculated by combining the measured γ -ray energies as explained in the text. ^bFrom Ref. 2.

^cSome or all of this γ -ray intensity could arise from a 10328 \rightarrow 8439 transition.

< 0.2

 0.22 ± 0.02

 0.03 ± 0.02 1.0 ± 0.5

^dFrom Ref. 13. There appears to be a transcription error in Table 24.11 of Endt and van der Leun (Ref. 2) for the decay of this level.

^eFrom Ref. 34.

9 300.95(15)

^f The previous branching ratios are from Ref. 24.

7349

7616

0

1090

822

9299

gUncertain.

4+

^hExists but of unknown intensity.

ⁱ A possible ${}^{22}Na(\beta^{-}){}^{22}Ne(1274.55\text{-keV level})$ contribution cannot be resolved from this transition.

ments were considered in making the transition assignments. In almost all cases where other transitions were possibly degenerate with those of Tables I and II, the assignment could be excluded because more intense transitions from the initial level had been reported but were not observed in our spectra. Nevertheless it was felt desirable to substantiate certain aspects of the decay scheme by a $\gamma\gamma$ -coincidence experiment. The coincidence measurement was made with the escape-suppression spectrometer as described earlier and a 12.7cm diameter by 12.7-cm long NaI(T1) detector at 180° to it. A 1.2-cm Pb absorber was interposed between the rabbit and the NaI(T1). Thirty hours

of coincidence data were event-mode recorded in a 8192×4096 matrix. From subsequent analyses of these data we were able to establish almost all of the coincidence relationships necessary to support the assignments of Tables I and II.

9516

1369

1369

10820.76(40)

11 314.23(150)

1060

9449

9944

 42.3 ± 2.5

(100)

(100)

The observed yields of the γ peaks listed in Table I vary from 1.58×10^6 for the 1369-keV transition to 25 for the 9944-keV γ ray. For most of the transitions the uncertainties in the yields were small compared to that in the relative efficiency which, therefore, dominates most of the uncertainties assigned to the branching ratios in Table II.

As can be seen from the comparison to the pre-

 34 ± 12

viously adopted branching ratios of Endt and van der Leun,² our branching ratios are in good general agreement with previous results. Somewhat minor discrepancies are noted for the 4238-keV level, the 5235 + 4238 transition, the 6010 + 4123 transition, and the 9516 + 4238 and 9516 + 5235 transitions. There is only one major discrepancy which is the nonobservation of the 822-keV 8439 + 7616 transition at the level reported by Détraz¹³ in the most extensive previous ${}^{24}\text{Al}(\beta^+){}^{24}\text{Mg study}$. We can only conclude that the 822-keV γ ray observed by Détraz arose from some other source.

The present γ -ray and level energies are, in general, considerably more precise than previous measurements. The agreement with previous results² is fair.

The two highest energy γ rays observed had energies of 9450.09(40) and 9943.46(150) keV and are shown in Fig. 3. [A 9440(10)-keV γ ray was also observed by Détraz¹³ in his study of ²⁴Al β^+ decay.] These two γ rays were assigned to the decav of ²⁴Al because the energies and intensities preclude any other possibilities. (They were both too weak to show up in the coincidence spectra.) The intensities are only compatible with allowed β^{+} decay, thus the ²⁴Mg states emitting these γ rays have $J^{*}=3^{+}$, 4^{+} , or 5^{+} . For these alternatives we reject the possibility that either of these γ rays is a ground-state transition as untenable. Since the $\beta^* Q$ value is not large enough, neither can arise from a cascade to a ²⁴Mg level with $E_r > 4$ MeV. Thus we conclude both γ rays correspond to transitions to the 2⁺ 1369-keV level as shown in Table II. It is possible that the 10821- and 11314-keV levels of this work can be associated with states observed at 10824(3) and 11318(3) keV in the ${}^{24}Mg(p,p'){}^{24}Mg$ reaction.²

The observed γ -ray intensities and assignments of Table II result in the β^+ branching ratios of Table III. The normalization of the relative intensities in Table I is such that the $\% \beta^+$ branch into a given level is

$$\% \beta^+ BR = I_\gamma (decay) - I_\gamma (feeding).$$
 (2)

The β^* branching ratio results are in excellent agreement with those of Détraz¹³ with which they are compared, and also with the recent less complete results of Shibata, Imazato, Yamazaki, and Brown¹⁴ (not shown). The log*ft* values were calculated using $T_{1/2} = 2.060 \pm 0.010$ s,¹⁵ $E_{\beta max} = 12856 \pm 4$ keV,² and the currently determined branching ratios.

New ${}^{24}\text{Al}(\beta^{+}){}^{24}\text{Mg}\ \beta^{+}$ branches from the present work are those to the 10576-, 10821-, and 11314keV levels and possibly to the 7812- and 9456-keV levels.

The J^* assignments of Table III are those of Endt and van der Leun² for $E_x < 10$ MeV. The three $(3, 4, 5)^*$ assignments for $E_x > 10$ MeV follow from the allowed character of the β decay demanded by the observed log*ft* values.

The probable branch to the 7812-keV level is based on the observation of a 2577.44(80)-keV γ ray which, if it were assigned to the 7812 - 5235

²⁴ Mg Level		Positron y	rield (%)	log ft			
្រា	(keV)	Present	Previous ^a	Present	Previous		
2*	1 369	<4.0		>6.98			
4+	4123	7.7 ± 1.0	7.0 ± 4.5	6.13 ± 0.06	6.2 ± 0.3		
2*	4 2 3 8	<0.20		>7.68			
3⁺	5 2 35	1.40 ± 0.13	1.7 ± 0.8	6.59 ± 0.04	6.5 ± 0.3		
4*	6010	1.2 ± 0.1	1.3 ± 0.8	6.43 ± 0.04	6.4 ± 0.3		
2+	7 348	<0.03		>7.59			
3-	7616	<0.06		>7.19			
5+	7812	0.05 ± 0.02 ^b		7.19 ± 0.17^{b}			
4*	8 439	50.0 ± 2.0	48 ± 6	3.93 ± 0.02	3.95 ± 0.06		
4*	9 301	2.5 ± 0.2	2.4 ± 0.5	4.80 ± 0.04	4.82 ± 0.09		
3+	9 4 5 6	0.033 ± 0.017 ^b		6.63±0.30 ^b			
4*	9 516	37.0 ± 1.5	39 ± 3	3.510 ± 0.018	3.489 ± 0.034		
(3, 4, 5)*	10576	0.67 ± 0.06		4.50 ± 0.04			
(4*)	10 580 [°]	<0.2		>5.4			
(3, 4, 5)*	10821	0.11 ± 0.01^{d}	≥0.01 ±0.05	5.08 ± 0.08			
$(3, 4, 5)^+$	11 314	$[(2.6 \pm 0.8) \times 10^{-2}]^{d}$		5.19 ± 0.14			

TABLE III. Positron decay of ²⁴Al.

^a Reference 7.

^b Uncertain (see text).

^c The other member of the 10.58-MeV doublet (Refs. 8 and 24) for which decay to the 1.37-

and 4.12-MeV levels is assumed to amount to at least 40%.

^d Assumes 100% γ branching to the ²⁴Mg 1369-keV level.

100

transition (a $60 \pm 3\%$ branch²), would give 7812.79(80) keV for the excitation energy of this level. [Another possible placement of the 2577-keV γ ray is to an 11.01 \rightarrow 8.44 transition (see Ref. 11).] The probable branch to the 9456-keV level is based on the observation of a γ -ray peak (see Fig. 3) for which the energy is in agreement with that expected for the 9456 \rightarrow 1369 transition (a $61 \pm 3\%$ branch²). Both of these β^+ branches are interpreted as limits because the γ -ray peaks were very weak and because such weak intensities could very well arise from β feeding via higher-lying ²⁴Mg levels.

Our results for the 10576-keV level are a confirmation of the ²³Na(p, γ)²⁴Mg studies of Boydell and Sargood.²⁴ These authors noted a major difference in the γ -decay modes of a "10.58-MeV level" when fed by two different (p, γ) resonances. They therefore proposed a doublet with one state having the decay modes with which our results are compared in Table II. Except for the branch to the 7616-keV level, the results are in fair agreement. Boydell and Sargood proposed J=3, 4, or 5 for their 10.58-MeV level. This is consistent with the assignment $J^{\pi} = 3^+$, 4^+ , or 5^+ resulting from the log ft value of Table III. We cannot state definitively that ²⁴Al(β^+) does not feed the other member of the 10.58-MeV doublet.²⁴ However, the various cascade sums give good agreement in the calculation of the excitation energy, 10 525.983(80) keV, for the level formed. Thus, the two 10.58-MeV levels would have to be separated by ≤ 0.2 keV if indeed both are formed significantly in the β^+ decay.25

The main motivation for the $\gamma\gamma$ -coincidence measurement was to obtain additional evidence confirming the placement of the 1952- and 1091-keV transitions. The data allow a definite assignment of the 1952-keV γ ray to the 9301 – 7349 transition. Coincidence spectra are shown in Fig. 4. This is but one of several scans of the coincidence data which, taken together, establish the assignment. The 1091-keV peak (see Fig. 2) was considerably weaker, relative to the background, than the 1952keV peak. The coincidence data gave evidence supporting the assignment of Table I but fell short of being completely definite. We regard this assignment as slightly uncertain but it will be assumed definite for purposes of discussion.

III. DISCUSSION

A. β^+ decay

The experimental log*ft* values of Table III are compared to theoretical predictions of Kelvin *et* $al.^4$ in Table IV, which also includes β^* -delayed α emission results of Torgerson, Oakey, and

COINCIDENCES GATED BY 80 $E_{\gamma} = 5.8 - 7.7$ MeV 60 40 20 COUNTS PER CHANNEI SINGLES 9516 -> 7616 (1900 keV) 9000 9301 -- 7349 10576 --- 8439 (1952 keV) (2137 keV) 6000 0 1640 1720 1800 1880 CHANNEL NUMBER FIG. 4. Partial spectra illustrating the $\gamma\gamma$ -coincidence

FIG. 4. Partial spectra illustrating the $\gamma\gamma$ -coincidence data from ²⁴Al(β)²⁴Mg. Peaks are labeled in the singles spectrum (lower) by the levels (in keV) between which the corresponding transitions occur and by their energies. The three peaks which are labeled occur in both the coincidence and singles spectra as is expected from their assignments.

Macfarlane.²⁶ A comparison between these predictions and experiment was made previously by Kelvin *et al.*⁴ The small changes between our results and previous results do not affect this comparison for the states with $E_x < 9.5$ MeV. Our highly tentative association of predicted 4⁺ states with experimental states for $E_x > 10$ MeV is different from that of Kelvin *et al.*

Kelvin *et al.*⁴ did not publish predictions for ${}^{24}\text{Al}(\beta^+)$ decay to the $J^{*}=5^+$ states of ${}^{24}\text{Mg}$. This is unfortunate since a comparison of the predictions to experiment for the ${}^{24}\text{Mg}$ 7812- and 10 576-keV levels would be especially interesting in view of the great difference in their log*ft* values.

As noted by Kelvin *et al.*, the general agreement is not very good and neither is the agreement between the predictions for the two different interactions. The agreement is expected to be best for the large matrix elements (small $\log ft$) and this tends to be true. However, there are only two branches which we regard as strong ($\log ft < 4$) and only two of medium strengh ($4 \le \log ft \le 5$). The

E _f (keV)	J _{fn} a.	Expt.	log <i>ft</i> CWC ^b	PW°
4123	41	6.13	5.86	5.92
5 2 3 5	31	6.59	6.09	5.88
6010	42	6.43	5.94	5.63
7812	5 ₁	≥7.19		
8 439	4 ₃	3.93	3.92	3.95
9 301	44	4.82	4.27	4.73
9456	32	>6.63		
9 51 6	$4_5 (T=1)$	3.510	3,512	
10328	46		5.16	7.72
10 576	5 ₂	4.50		
(10 580) ^d	47	>5.4	4.10	4.82
10 821	4 ₈	5.08	5.15	
11217	4 ₉	6.6	6.54	
11 694	4 ₁₀	5.2	7.56	

TABLE IV. Comparison of the logft values of Table

III with the predictions of Kelvin et al. (Ref. 4).

^a For $E_x \leq 9516$ keV the association of shell-model states with experiment is that of Kelvin *et al.* (Ref. 4). The assumption that the 10576-keV level is the 5_2^+ state follows Fifield *et al.* (Ref. 11). The associations for the remaining states for $E_x > 10$ MeV are highly speculative. The β^+ decay to the 11217- and 11694-keV levels was observed in the β^+ -delayed α emission studies of Torgerson *et al.* (Ref. 26).

^b The Chung-Wildenthal interaction (Ref. 6) with added effective Coulomb terms.

^c The Preedom-Wildenthal interaction (Ref. 5).

^d Reference 8.

comparison for the superallowed $\Delta T = 0$ decay to the analog of the ²⁴Al ground state at 9516 keV in ²⁴Mg is interesting. The *ft* value was calculated from⁴

$$ft = \frac{6250}{\langle 1 \rangle^2 + 1.51 \langle \sigma \rangle^2} , \qquad (3)$$

where $\langle 1 \rangle$ and $\langle \sigma \rangle$ are shorthand for the Fermi and Gamow-Teller matrix elements, respectively. For $T, T_x = 1, -1 \rightarrow T, T_x = 1, 0, \langle 1 \rangle^2 = 2$ and for $\langle \sigma \rangle^2$ = 0, Eq. (3) gives $\log ft = 3.495$. The fact that the predicted $\log ft$ value of Kelvin *et al.* is larger than this value implies that their calculation gives a departure from the isospin-symmetry prediction for $\langle 1 \rangle^2$, even if $\langle \sigma \rangle^2 = 0$, which amounts to at least 4% (larger if $\langle \sigma \rangle^2 = 0$, which amounts to at least 4% (larger if $\langle \sigma \rangle^2$ is not zero). Our experimental result supports this prediction indicating $\geq 4\%$ isospin impurity in the ²⁴Mg 4⁺, T = 1 state at 9516 keV. The very small value of $\langle \sigma \rangle^2$ implied for this decay is very interesting; a possible explanation for it is offered in the next subsection.

Aside from the superallowed decay, that to the third 4⁺ state at 8439 keV is dominant. The strength of this transition and the strong 9516 \rightarrow 8439 *M*1 transition (Table II) led Détraz¹³ to propose the 8439-keV level as the antianalog of the 9516-keV state. In the next subsection we



FIG. 5. Band structure of ²⁴Mg. All the levels of ²⁴Mg for $E_x < 9$ MeV and associated levels for $E_x > 9$ MeV are grouped into bands. The bands are labeled by what is proposed to be the simplest (or one of the simplest) occupations in a Nilsson scheme and by the dominant term in a SU(3) basis as discussed in the text. Uncertain spin-parity assignments are denoted with an asterisk. The numbers in squares or circles are E2 transition strengths in Weisskopf units.

shall propose an ordering of the even-parity levels of ²⁴Mg with $E_x < 9$ MeV into bands and it will be seen that the concept of an antianalog to the J^* , $T = 4^*$, 1 9516-keV level of ²⁴Mg has a sound basis.

B. The band structure of ²⁴Mg

1. Classification in the Nilsson scheme

The unified model²⁷ based largely on the work of Nilsson²⁸ provides a convenient framework for the classification of the levels of ²⁴Mg into rotational bands. Such a classification is attempted in Fig. 5 which includes all even-parity levels of ²⁴Mg for $E_x < 9$ MeV. The levels are grouped into bands which are labeled by what is expected to be the dominant (or at least a significant) Nilsson level occupancy. We consider eight nucleons distributed among the Nilsson orbits, (#) 6, 7, 5, 9, 11, and 8. For ²⁴Mg we expect a deformation of $\beta \simeq 0.4$ and a spin-orbit parameter κ of ~0.08.²⁹ For these parameters the ordering of the orbits

in energy is as given above with #5 and #9 essentially degenerate (in ${}^{25}Mg - {}^{25}Al$ the $\frac{5}{2} + -\frac{1}{2} +$ splitting is ~0.5 MeV indicating that the #5 orbit lies below the #9 orbit by this amount).

The Nilsson occupancies for the lowest two bands are the obvious dominant ones. For the excited K=0 band there are many possibilities all involving nucleons in the #5, #8, or #11 orbits. The simplest is shown—it is not necessarily dominant.

Daum²⁹ has considered the conjugate nuclei ²⁴Na-²⁴Al in the Nilsson model. He assigned the $J^{\pi} = 4^+$, T = 1 ²⁴Al ground state to the K = 4 band based on $(\#7)^3 \#5$ as shown in Fig. 5. For this Nilsson occupancy in ²⁴Mg the nucleon in orbit #5 has an equal probability of being a neutron or a proton. Thus the T = 1 band has an antianalog T = 0band and this we assign to the band built on the 8.44-MeV level.

Again for the K = 1 band there are many possibilities. One of the simplest is obtained from the $(\#7)^3 \#5$ configuration which can generate two K = 4 bands (T = 0 and T = 1) from $\Omega_1 + \Omega_2 = \frac{5}{2} + \frac{3}{2} = 4$ and also two K = 1 bands from $\Omega_1 - \Omega_2 = \frac{5}{2} - \frac{3}{2} = 1$.

2. Classification into SU(3) symmetries

An alternate description of the band structure of ²⁴Mg is in terms of SU(3) symmetries.³⁰⁻³² Neither the Nilsson nor the SU(3) description is expected to be rigorous and it is enlightening to consider both. The bands of Fig. 5 are labeled by the expected dominant SU(3) symmetry. The theoretical evidence for the assignments to the lowest K = 0 and K = 2 bands was discussed fully previously.^{1,2} The evidence for the assignment of [44](46) to the excited K=0 band is the calculation of Wathne and Engeland.⁷ The remaining assignments are based on energetics and the connection between the Nilsson and SU(3) models.³⁰⁻³²

The connection between the SU(3) classification and the Nilsson occupancies is given in Table V for all the symmetries assigned in Fig. 5 except [431](73) which is quite complicated. The connection is for the $\beta \rightarrow \infty$ asymptotic limit or equivalently for no spin-orbit force and no 1*d*-2*s* splitting. The [44] bands of Fig. 5 all have T=0; T=1is not allowed. For [431], T=0 and T=1 are both allowed. Actually there is a multitude of $J^*=4^+$ states (each with its band) in [431](65) and [431](73). These can be specified by the quantum numbers K_s, K_L . The $K_s=1$, $K_L=3$ states, both T=0 and T=1, are expected to lie lowest³¹ and this is the state we consider in Table V.

3. Evidence for the band assignments

The lowest K = 0 and K = 2 bands have been fully discussed in the literature.¹ The main evidence for the band structure is found in the strong intraband B(E2) values and the weak interband B(E2)values. The intraband B(E2) values [in W.u. (Ref. 10)] are shown in the square boxes in Fig. 5. They are taken from the compilations of Endt^{2,33} or from the ²⁰Ne(α , γ)²⁴Mg work of Fifield *et al.*¹¹

The nuclear spectroscopy of the T=1 levels of mass 24 is very poorly known. For instance, the analogs of the ²⁴Mg 11.01- and 13.37-MeV levels have not been established in ²⁴Ne or ²⁴Al and the spin-parity assignments of these levels have not been determined. However, the values suggested by Fifield *et al.*¹¹ and shown in Fig. 5 are probably

TABLE V. Nilsson orbit occupancies for some SU(3) symmetries.

					-	
	#7 ⁴	#7 ³ #9	[44] (84)	#7 ² #9 ²	#7 #9 ³	#9 ⁴
<i>K</i> = 0	$\frac{1}{6}$			$\frac{2}{3}$		1 7
K = 2	Ū	$\frac{1}{2}$		0	$\frac{1}{2}$	6
	(#6) ⁴ (#7,	#9) ² (#5, #8, #11) ²	[44] (46)	(#6	5) ³ (#7, #9) ⁴ (#5, #8, #11)
K = 0		$\frac{1}{7}$			<u>6</u> 7	
	ŧ	7 ² #9 #11	[431] (65)	#9 ² #7 #5	#7	³ #5
<i>K</i> = 4		$\frac{2}{5}$		$\frac{2}{5}$, , , , , , , , , , , , , , , , , , ,	1 5
			[44] (73)			
		(#6) ⁴ (#7, #9) ³ (#	5, #8, #11)		(#6) ³ (#7, #9)	⁵
<i>K</i> =1		10 13			3 13	

correct. No intraband B(E2) values are known for the K=4, T=1 band since these levels decay predominantly by $\Delta T=1$ dipole transitions.

Granting that the 9.52-MeV level is the bandhead of a K=4, T=1 band, the main arguments for the K=4, T=0 band come from ${}^{24}\text{Al}(\beta^+){}^{24}\text{Mg}$. The two strongest Gamow-Teller ${}^{24}\text{Al}\beta^+$ decays are to the 8.44- and 10.58-MeV levels and since Gamow-Teller β decay has a $\Delta K \leq 1$ selection rule it is natural to assign K=4 to these two levels. The suggested spin-parity assignments shown in Fig. 5 for the 10.58- and 12.00-MeV levels are also from Fifield *et al.*¹¹

It is instructive to consider the SU(3) predictions for the β^* decay even though it is not expected that they will be obeyed in detail. For the SU(3) band structure of Fig. 5 the only allowed ²⁴Al β^* transitions are to the 4^{*} and 5^{*} levels of the K=4 bands. It is pleasing, then, to note that the transitions to three such levels are the strongest observed (see Table IV).

In more detail the Gamow-Teller matrix element vanishes identically between states of the supermultiplet [431], with $K_s = 1$, $K_L = 3$, thus the very small value for this quantity which is discussed in the previous subsection has a natural explanation in the framework of the SU(3) model.

In the same SU(3) model the predicted $\log ft$ values for decay to the 4⁺ and 5⁺ T=0 levels of the K=4 band are 3.24 and 4.16, respectively. The calculation was carried out for [431](73) SU(3) wave functions. The result for [431](65) is very closely the same and since the amplitudes for these two symmetries add incoherently the result is very insensitive to the weighting of these two symmetries in the K=4 bands. More important are the possible deviations from the assumption $K_s=1, K_L=3$. Thus, we expect that the log ftvalues quoted are approximate lower limits and the fact that the ratio of the two is given approximately correctly is perhaps all that can be expected.

The evidence for the grouping of the levels at 7.75, 8.65, and 9.46 MeV into a K=1, T=0 band is highly speculative. However, these are the only levels remaining after the other band assignments of Fig. 5 are made, thus the spin sequence is suggestive of this band structure. The spin-parity of the 9.46-MeV level has not been determined²; the 3⁺ assignment is suggested by comparison to the predictions of Kelvin *et al.*⁴

The suggested grouping of the ²⁴Mg levels at 6.43, 7.35, 9.30, and 12.86 MeV into a K=0, T=0 band is based mainly on the large B(E2) values shown in Fig. 5 for the $6^+ - 4^+$ and $4^+ - 2^+$ transitions and the theoretical evidence discussed in Sec. I. The $6^+ - 4^+$ strength is from Fifield *et al.*¹¹ while the

 $4^+ - 2^+$ strength is obtained from the present branching ratios (Table II) and previous lifetime measurements. As an aid in this discussion, the known B(E2) values for the decay of the ²⁴Mg 8.44-, 9.30-, and 9.52-MeV levels are collected in Table VI. These B(E2) values result from the branching ratios of Table II and the listed mean lives. The mean life listed for the 9.52-MeV level is the adopted value of Endt and van der Leun.² For the 8.44- and 9.30-MeV levels the only published measurement is that of Meyer *et al.*³⁴ For the 2nd through 6th excited states of ²⁴Mg, the mean life results of Meyer *et al.* average $\sim \frac{1}{2}$ of the adopted value of Endt and van der Leun.² Thus we arbitraily increase by a factor of 2 the mean lives quoted by Meyer et al. for the 8.44- and 9.30-MeV levels. This we do for purposes of present discussion but we note the importance of additional lifetime measurements for these levels.

A perusal of Table VI yields the interesting fact that only two B(E2) values are enhanced and both are $4^{+}-2^{+}$ transitions to the 7.35-MeV level. That from the 9.30-MeV level gives us the evidence for a K=0 band, while that from the 8.44-MeV level clearly indicates mixing between the K=0 and K=4 4^{+} levels at 8.44 and 9.30 MeV, i.e., E2 transitions have a $\Delta K \leq 2$ selection rule. The relatively strong

TABLE VI. Some E2 decays for $J^{\mathbf{f}} = 4^+$ levels of ²⁴Mg.

E _i (keV)	$\frac{E_f}{(\text{ke V})}$	Transition	B(E2) ^a (W.u.)	$\tau_i^{\mathbf{b}}$ (fs)
8439	1369	$4^+_3 \rightarrow 2^+_1$	0.29	26 ± 10
	4123	$4_3^+ \rightarrow 4_1^+$	≤1.1	
	4238	$4_3^+ \rightarrow 2_2^+$	0.36	
	5235	$4_3^{\star} \rightarrow 3_1^{\star}$	≤1.1	
	6010	$4_3^+ \rightarrow 4_2^+$	≤1.1	
	7349	$4^+_3 \rightarrow 2^+_3$	11 ± 5	
9301	1369	$4_4^+ \rightarrow 2_1^+$	0.17	20 ± 8
	4123	$4^+_4 \rightarrow 4^+_1$	≤1.1	
	4238	$4_4^+ \rightarrow 2_2^+$	0.04	
	7349	$4^+_4 \rightarrow 2^+_3$	15 ± 6	
9516	1369	$4_5^+ \rightarrow 2_1^+$	<1.7×10 ⁻⁴	25 ± 10
	4238	$4_5^+ \rightarrow 2_2^+$	<1.3×10 ⁻²	
	7349	$4_5^+ \rightarrow 2_3^+$	<1.1×10 ⁻¹	

^a For the $\Delta J < 2$ transitions, pure E2 was assumed, hence the limits. The uncertainties are derivable from those on the branching ratios of Table I and the mean lifetimes of column 5.

^b For the 8439- and 9301-keV levels the mean lives are twice those of Ref. 34 as discussed in the text. The value for the 9516-keV level is from Ref. 2. β^* decay of ²⁴Al to the 9.30-MeV level is further evidence of this mixing.

Other interband transitions in ²⁴Mg are present, for instance, the relatively strong 8.65 - 6.43 E2 transition shown in Fig. 5 and the decay of the 6.43-MeV level.² These remind us not to take the band structure shown in Fig. 4 too literally. The E2 decay of the 9.52-MeV level also listed in Table VI indicates sizable inhibition which goes well beyond that expected from the $\Delta T = 1$ selection rule.³⁵ The $\Delta K \leq 2$ selection rule provides an explanation for the weak decays to the K = 0 levels but not for the decay to the K = 2 4238-keV level.

IV. SUMMARY

The present study of ${}^{24}\text{Al}(\beta^{+}){}^{24}\text{Mg}$ yielded 44 ${}^{24}\text{Mg}$ γ transitions compared to the 19 previously reported.¹³ Three β^{+} branches were observed in addition to the six previously reported. One of these—to the 10 576-keV state—has the third lowest observed ${}^{24}\text{Al}\log ft$ value. In addition to this branch, the major new information obtained is the observation of $8.44 \rightarrow 7.38$ and $9.30 \rightarrow 7.38$ E2 transitions which are rather strongly collective. Because the 12.86 \rightarrow 9.30 E2 transition is also strong we suggest placement of the ${}^{24}\text{Mg}$ levels at 7.38, 9.30, and 12.86 MeV in an excited K = 0 band built on the 0⁺ 6.43-MeV level. However, the rather strong 8.44 \rightarrow 7.38 E2 transition indicates rather strong interband mixing.

Both K=4, T=1 and K=4, T=0 bands are proposed. The β^+ decay matrix elements connecting the 4^{+ 24}Al ground state (K=4, T=1) with the K=4, T=1 4⁺ level at 9.52 MeV and the 4⁺ and (5⁺) K=4, T=0 levels at 8.44 and 10.58 MeV are calculated in the extreme assumption that all have the SU(3) symmetry [431](65) with $K_s=1$, $K_L=0$. The calculation provides an explanation for the small matrix element of $\langle \sigma \rangle^2$ in the superallowed decay and predicts strong transitions to the K=4, T=0 band with the ratio of decays to the 4⁺ and 5⁺ states

in rough conformity to those for the 8.44- and 10.58-MeV levels.

The present discussion points up the need for more accurate lifetime measurements in ²⁴Mg and for additional spin-parity determinations—especially for those marked by an asterisk in Fig. 5. An expansion of the large basis shell-model wave functions³⁻⁶ in terms of SU(3) wave functions would be of great interest.

Note added in proof. J. D. Garrett, H. T. Fortune, R. Middleton, and W. Scholz [Phys. Rev. C 18, 2032 (1978)] have proposed a band structure for 24 Mg similar to that given in Fig. 5. K. Kumar [International Conference on the Structure of Medium-Heavy Nuclei, Rhodes, Greece, 1979, Inst. Phys. Conf. Ser. No. 49, 169 (1980)] has calculated results for the even J, even parity, T=0band structure of ²⁴Mg using the dynamic deformation theory. Kumar (unpublished) successfully predicts a K = 0 band (β vibration, band head at 6.01 MeV) which can be associated with the K = 0band commencing at 6.43 MeV in Fig. 5. He calculates considerable mixing of this band with the lower two bands and with the lowest K = 4 band also in agreement with results summarized in Fig. 5. However, the predicted β band has a 4⁺-0⁺ spacing of 5.5 MeV as opposed to 2.87 MeV in Fig. 5 and a close-lying second K = 0 band ($\gamma\gamma$ -vibration, band head 6.52 MeV) is predicted in disagreement with experiment (i.e., no 0^+ states other than those of Fig. 5 are known below 9 MeV in ²⁴Mg).

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reaction [implying the parity of $(-)^J$]. However, we feel that some doubt attaches to this assignment because the spin assignment of the 1416-keV resonance is not definite (Ref. 2) and because, in neither reaction, is it definite that only this member of the 10.58-MeV doublet is being populated.

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