Inelastic scattering E4 transition probabilities in the 0d, 1s shell

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Theoretical E4 transition matrix elements calculated in the full $d_{5/2}$ - $s_{1/2}$ - $d_{3/2}$ shell-model space are presented for excitations of some ground states of stable nuclei for $18 \le A \le 38$. These matrix elements have been used to calculate inelastic transition probabilities for electromagnetic and hadronic probes. Available data on the relative magnitudes and energy distributions of E4 excitation strengths are found to be well reproduced theoretically thereby. Based on several experimentally measured transition strengths, the empirical isoscalar E4 effective charge is found to be $e_p + e_n = (2.0 \pm 0.2)e$. This is similar in magnitude to the analogous isoscalar E2 effective charge, but is much larger than existing theoretical estimates.

NUCLEAR STRUCTURE ¹⁸O, ¹⁹F, ²⁰Ne, ²²Ne, ²⁴Mg, ²⁵Mg, ²⁶Mg, ²⁸Si, ³⁰Si, ³²S, ³⁴S, ³⁶Ar, ³⁸Ar; calculations for the strengths of *E*4 inelastic scattering transitions and their proton and neutron components; complete $0d_{5/2}-1s_{5/2}-0d_{3/2}$ shellmodel wave functions; Chung-Wildenthal Hamiltonians.

I. INTRODUCTION

In this paper we present calculations for the strengths of $\Delta J^{\pi} = 4^{*}(E4)$ excitations of ground states in the region $18 \le A \le 38$. The one-body transition density matrix elements for these excitations were calculated from the $d_{5/2}$ - $s_{1/2}$ - $d_{3/2}$ shell-model wave functions of Chung and Wildenthal.^{1,2} These transition densities were then used to calculate matrix elements for electromagnetic excitation of these transitions and for the relative strengths of these same transitions as they are excited by various hadronic probes. The present calculations systematically employ harmonic-oscillator singleparticle wave functions. The effects of using different (e.g., Hartree-Fock) radial dependences for the single-particle states were examined and found to be significant, but no alternative clearly preferable to harmonic-oscillator dependence was determined.

Our present study focuses on three aspects of *E*4 phenomena. The first issue we address presumes reasonably good theoretical-experimental agreement between *relative* values of several large measured transition strengths. In such a context it is meaningful to speak of the "effective charges" appropriate to the model space which serve to produce agreement between calculation and experiment in the absolute as well as relative sense. By comparing values from the extant body of measured E4 rates to our calculated values we are able to extract a value of the isoscalar E4 effective charge for the sd-shell space of $e_p + e_n$ = $(2.0 \pm 0.2)e$.

We then proceed to deal with the general questions of what type of E4 phenomena are predicted by the Chung-Wildenthal wave functions and how well these accord with experimental facts. There are several aspects to be considered, such as the dependence of the aggregate E4 strength on the A value, the distribution of strength within a given nucleus as a function of excitation energy, and the A dependence of this distribution.

Lastly, we consider the individual neutron and proton components of the transitions in the $T \neq 0$ nuclei. From the values of these components the relative strengths by which a given state is excited by various hadronic probes, as well as electronmagnetically, can be predicted. We consider in particular, alphas, protons, and pions.

In Sec. II the calculation of the shell-model transition-density matrix elements is described as well as their combination with the single-particle matrix elements (which incorporate the radial wave function dependence) to obtain the theoretical B(E4) values. In Sec. III these values are compared with the strong experimental B(E4) values to obtain E4 effective charges. In addition, the

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distributions of calculated E4 strengths for all stable even-even nuclei are presented in comparison with available experimental data, together with some selected results for odd-even nuclei. In Sec. IV a method for calculating the hadronic excitation strengths is presented. The conclusions of the present study constitute Sec. V.

II. SHELL-MODEL CALCULATIONS OF THE TRANSITION DENSITIES

The reduced matrix element for a one-body operator between initial and final *sd*-shell wave functions for *N* valence particles is expressed as a linear combination of reduced matrix elements between single-particle states ρ_j and the one-body transition densities $D_{\Delta T_i,\Delta T}^{N,JT,J',T'}$ by the relation

$$\left\langle \psi^{NJT} \middle| \middle| \left| \sum_{i=1}^{N} O(\Delta J, \Delta T) \middle| \left| \middle| \psi^{NJ^{*}T^{*}} \right\rangle$$

$$= \sum_{jj'} D^{N,JT,J^{*}T^{*}}_{\Delta T,\Delta T} (j,j') \langle \rho_{j} \middle| \left| O(\Delta J, \Delta T) \right| \left| \left| \rho_{j'} \right\rangle,$$

$$(1)$$

where

$$D_{\Delta T,\Delta T}^{N,JT,J'T'}(j,j') = \frac{\langle \psi^{NJT} | | | (a_j^* \times \tilde{a}_{j'})_{\Delta J\Delta T} | | | \psi^{NJ'T'} \rangle}{[(2\Delta J+1)(2\Delta T+1)]^{1/2}}.$$
(2)

The operator $O(\Delta J, \Delta T)$ has associated with it a unit operator in isospin space when $\Delta T = 0$ and the operator τ in isospin space when $\Delta T = 1$; our isospin convention is such that

$$O(\Delta J)_{p/n} = \frac{1}{2} [1(+/-)\tau_3] O(\Delta J).$$
(3)

The proton and neutron matrix elements for these operators between the sd-shell wave functions are given by

$$A_{p/n}(T_{z}) = \frac{(-1)^{T_{f}-T_{z}}}{2} \begin{pmatrix} T_{f} & 0 & T_{i} \\ -T_{z} & 0 & T_{z} \end{pmatrix} \left\langle \psi^{J_{f}T_{f}} \right| \left| \left| \sum_{i=1}^{N} O(\Delta J, \Delta T = 0)_{i} \right| \left| \left| \psi^{J_{i}T_{i}} \right\rangle \right\rangle$$

$$(+ /-) \frac{(-1)^{T_{f}-T_{z}}}{2} \begin{pmatrix} T_{f} & 1 & T_{i} \\ -T_{z} & 0 & T_{z} \end{pmatrix} \left\langle \psi^{J_{f}T_{f}} \right| \left| \left| \sum_{i=1}^{N} O(\Delta J, \Delta T = 1)_{i} \right| \left| \psi^{J_{i}T_{i}} \right\rangle.$$

$$(4)$$

It is well known in the case of E2 transitions that core-polarization corrections are important and we anticipate that they will also be important for E4 transitions. Thus we introduce proton and neutron effective charges e_p and e_n which relate the total proton matrix element M_p to the proton and neutron matrix elements in the model space A_p and A_n :

$$M_{p} = A_{p}e_{p} + A_{n}e_{n} = A_{p}(1 + \delta e_{p}) + A_{n}\delta e_{n}.$$
 (5)

In terms of M_{b} the B(E4) value is given by

$$B(E4, J_i - J_f) = (2J_i + 1)^{-1}M_{\rho}^2.$$
(6)

The quantity δe_p represents the effect of virtual excitation of core protons by the valence protons and the quantity δe_n represents the effect of the virtual excitation of core protons by the valence neutrons. If we denote δe_p by δ_{pp} and δe_n by δ_{pn} , then the generalizations of Eq. (5) for both the proton and neutron matrix elements take the form

$$M_{\rho} = A_{\rho} (1 + \delta_{\rho\rho}) + A_{n} \delta_{\rho n} ,$$

$$M_{n} = A_{n} (1 + \delta_{nn}) + A_{\rho} \delta_{n\rho} .$$
(7)

(The neutron matrix elements will be needed to calculate strengths for hadronic excitations.) For nuclei with approximately equal numbers of neutrons and protons $\delta_{np} = \delta_{pn}$ and, since the nuclear interaction is charge symmetric, $\delta_{nn} = \delta_{pp}$. Thus

we can express the neutron matrix element in terms of the proton and neutron effective charges,

$$M_n = A_n e_p + A_p e_n \,. \tag{8}$$

The values of the one-body transition densities D for some of the transitions discussed in the following sections are given in Table I. These quantities constitute the total information content of the shell-model wave functions insofar as these E4 excitations are concerned. The matrix elements

$$\left\langle \psi^{NJT} \left| \left| \right. \right| \left| \right. \sum_{i=1}^{N} O\left(\Delta J , \Delta T \right)_{i} \left| \right. \left| \left. \left| \right. \psi^{NJ'T} \right. \right\rangle \right\rangle$$

or the equivalent $A_{p/n}$ can be calculated from these entries once a set of single-particle matrix elements are specified.

The triply reduced single-particle matrix elements are

$$\langle \rho_j | | O(\Delta J, \Delta T = 0) | | \rho_{j'} \rangle = (2)^{1/2} \langle \rho_j | O(\Delta J) | \rho_{j'} \rangle$$

and

$$\langle \rho_j | \left| \left| O(\Delta J, \Delta T = 1) \right| \right| \left| \rho_{f'} \rangle = (6)^{1/2} \langle \rho_j | O(\Delta J) \right| \left| \rho_{j'} \rangle.$$
(9)

The electric multipole operator is given by

$$O(\Delta J)_i = r_i^{\Delta J} Y_m^{(\Delta J)}(\hat{r}_i) .$$
⁽¹⁰⁾

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			<i>j</i> - <i>j</i> ′					
A	$T_i \rightarrow T_f$	$J_i \rightarrow J_f$	$\frac{\Delta T}{\frac{5}{2}-\frac{5}{2}} = 0$	$\frac{5}{2}$ $\frac{3}{2}$	$\frac{3}{2}$ $\frac{5}{2}$	$\frac{\Delta T}{\frac{5}{2} - \frac{5}{2}} = 1$	$\frac{5}{2}$ $\frac{3}{2}$	$\frac{3}{2}$ $\frac{5}{2}$
18	1-+1	$0^+ - 4^+(1)$	-0.8266	-0.0726	0.1817	-0.6749	-0.0593	0.1483
19	$\frac{1}{2} \rightarrow \frac{1}{2}$	$\frac{1}{2}^{+} \rightarrow \frac{9}{2}^{+}(1)$	0.4483	0.3411	-0.1715	0.2395	-0.0377	-0.1721
20	0-0	0 ⁺ -+4 ⁺ (1)	0.4033	0.2714	-0.3061			
		$0^+ \rightarrow 4^+(2)$	0.0141	-0.1054	0.0695			
22	1-1	0 ⁺ →4 ⁺ (1)	-0.0210	-0.2009	0.3206	0.6233	0.0303	-0.0638
		$0^+ - 4^+(2)$	0.8423	0.0269	-0.2217	0.3384	-0.0610	-0.0218
24	0-0	$0^+ - 4^+(1)$	-0.3057	-0.0076	-0.0426			
		0 ⁺ → 4 ⁺ (2)	0.2402	0.3082	-0.4845			
		0+-+4+(3)	0.3469	-0.0976	0.0327			
25	$\frac{1}{2} \rightarrow \frac{1}{2}$	$\frac{5}{2}^+ \rightarrow \frac{9}{2}^+ (1)$	0.4962	-0.2212	0.0778	-0.1159	-0.0182	0.0257
		$\frac{5}{2}^{+} \rightarrow \frac{9}{2}^{+}(2)$	0.2396	0.4615	-0.2587	-0.0123	0.0056	0.0323
		$\frac{5}{2}^+ \rightarrow \frac{13}{2}^+$ (3)	-0.2156	-0.2610	0.5347	0.1647	0.0585	-0.0318
26	1-1	$0^+ - 4^+(1)$	0.7231	0.1225	-0.3194	-0.5609	0.0920	-0.1292
		0 ⁺ → 4 ⁺ (2)	-0.0417	-0.3206	0.7093	0.0043	-0.0235	0.1286
		$0^+ \rightarrow 4^+(3)$	0.1833	0.1277	-0.2988	0.3690	-0.0097	-0.1744
28	0-+0	0+4+(1)	-0.0065	-0.4561	0.2478			
		$0^+ \rightarrow 4^+$ (2)	0.2267	0.4496	-0.1517			
		0 ⁺ → 4 ⁺ (3)	-0.2008	0.1694	-0.1190			
30	1-1	0 ⁺ -+4 ⁺ (1)	0.2305	0.6219	-0.3076	-0.0978	0.3469	-0.1662
		0 ⁺ → 4 ⁺ (2)	-0.1606	-0.8941	0.3455	-0.0534	0.2664	-0.0551
		$0^+ - 4^+(3)$	-0.0122	-0.0787	-0.0178	0.3036	0.3685	-0.0363
32	0-+0	0+ -+ 4+ (1)	-0.0911	-0.5652	0.3328			
		$0^+ - 4^+(2)$	0.0917	0.3303	-0.1418			
	0-1	0 ⁺ →4 ⁺ (2)				0.0288	0.4918	-0.0858
34	1-1	0 ⁺ →4 ⁺ (1)	0.1361	0.8449	-0.4448	0.0049	-0.0445	-0.0372
		$0^+ - 4^+(2)$	-0.1047	0.1241	-0.0464	-0.0690	-0.3905	0.1371
36	00	$0^+ \rightarrow 4^+(1)$	-0.1252	-0.4596	0.3211			
	0-1	0+-+4+(3)				-0.0089	-0.4011	0.1070
38	1-1	0 ⁺ -+4 ⁺ (1)	-0.0345	-0.8156	0.1947	0.0281	0.6659	-0.1589

TABLE I. Calculated values of the one-body transition densities for some selected E4 transitions.

For $\Delta J = 4$ there are only two possible independent single-particle matrix elements in the *sd*-shell space and these are given by

$$\langle d_{5/2} | | r^{4}Y^{(4)} | | d_{5/2} \rangle = \left(\frac{9}{7\pi}\right)^{1/2} \langle d_{5/2} | r^{4} | d_{5/2} \rangle ,$$

$$\langle d_{5/2} | | r^{4}Y^{(4)} | | d_{3/2} \rangle = \left(\frac{18}{7\pi}\right)^{1/2} \langle d_{5/2} | r^{4} | d_{3/2} \rangle .$$

$$(11)$$

Since the E4 matrix elements are weighted by

 r^4 , the relation between experiment and theory is very sensitive to the specification of the radial integrals in Eqs. (11). We have chosen to evaluate these integrals independently for each nucleus we consider by choosing harmonic-oscillator wave functions parametrized to reproduce the individual measured values of the rms charge radii. The rms charge radii $r_{\rm ch}$ for essentially all stable *sd*shell nuclei are now known to high accuracy, a rather recent development. These values are

	$r_{\rm ch}$	b	πω
	(fm)	(fm)	(MeV)
¹⁸ O	2.794(3) ^a	1.820	12.50
¹⁹ F	2.898(10) ^b	1.833	12.34
²⁰ Ne	3.020(20) ^c	1.869	11.87
²² Ne	2.949(20) ^d	1.821	12.50
²⁴ Mg	3.035(18) ^e	1.813	12.61
²⁵ Mg	3.003(11) ^f	1.793	12.90
²⁶ Mg	3.017(32) ^g	1.802	12.77
²⁸ Si	3.127(3) ^{b, h}	1.827	12.42
^{30}Si	3.139(3) ⁱ	1.834	12.33
^{32}S	3.264(2) ^b	1.881	11.72
^{34}S		1.881	11.72
³⁶ Ar	3.399(5) ⁱ	1.938	11.04
³⁸ Ar	$3.414(10)^{j}$	1.947	10.94

TABLE II. Rms charge radii and the extracted harmonic oscillator parameter b.

^a Reference 26.

^bReference 27.

^c Reference 28, Table I.

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<sup>d</sup> Reference 10; r_{ch}(^{22}Ne) - r_{ch}(^{20}Ne) = -0.071 fm.
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^e Reference 29.

^f Reference 30.

^gReference 28, Table II.

^h For natural Si, $r_{ch} = 3.129$ (3).

ⁱ Reference 31, Table IV.

^j Reference 31, Table III.

listed in Table II.

For the harmonic-oscillator potential $V(r) = \frac{1}{2}m\omega^2 r^2$, the point proton rms radius for nuclei in the *sd* shell is given by

$$r_{p}^{2} = \left[\frac{18 + (Z - 8)^{7}_{2}}{Z}\right]b^{2} - \frac{3b^{2}}{2A},$$

where

 $b^2 = \hbar / m \omega \,. \tag{12}$

The last term is the correction for center-of-mass motion.^{3,4} The charge radius is obtained by folding r_p with the rms charge radii for the protons and neutrons and adding relativistic corrections. Ignoring the relativistic spin-orbit correction, which is important only for nuclei with a large neutron or proton excess of spin unsaturated nucleons, the rms charge radii are given by³

$$r_{\rm ch}^{2} = r_{p}^{2} + r_{\rm proton}^{2} + \frac{N}{Z} r_{\rm neutron}^{2} + \frac{3}{4} \left(\frac{\hbar}{mc}\right)^{2}, \qquad (13)$$

where $r_{\text{proton}^2} = (0.86)^2$ and $r_{\text{neutron}^2} = -(0.35)^2$. The values of *b* and $\hbar \omega$ extracted from the experimental charge radii are given in Table II. In terms of *b*, the radial matrix elements of Eqs. (11) are expressed as

$$\langle d_{5/2} | r^4 | d_{5/2} \rangle = \langle d_{5/2} | r^4 | d_{3/2} \rangle = 15.75b^4.$$
 (14)

The procedure we have adopted obviously takes

into account all of the new precisely known variations in the sizes of *sd*-shell ground states. The conventional prescriptions for either harmonicoscillator or Saxon-Woods potentials, which assume some smoothly varying mass dependence, e.g., $\hbar \omega = 41 A^{-1/3}$, fail to account for these variations, sometimes by significant amounts. In our opinion it is appropriate to remove this source of noise in the comparison of theoretical to experimental elements by abandoning size formulas altogether and using instead the individual measured radii.

We utilize harmonic-oscillator radial dependence rather than Saxon-Woods or Hartree-Fock prescriptions because these latter, supposedly more "realistic," presciptions do not in fact offer improvements over the harmonic oscillator sufficient to justify the additional complexity and ambiguity their use introduces into the problems at hand. The problem inherent in specifying singleparticle wave functions for open-shell nuclei such as concern us here is that the experimental separation energies are inconsistent with the energies of the shell-model single-particle potential, e.g., for ²⁸Si the $1s_{1/2}$ separation energy is greater than that for $0d_{5/2}$, and $0d_{3/2}$ is greater than $1s_{1/2}$, in inverse order to the shell-model sequence. In the full shell-model calculation this effect comes out of the two-body part of the Hamiltonian, but simple single-particle models are helpless to deal with it. Moreover, by going from the infinite harmonic-oscillator well to finite wells, one runs the risk of obtaining more realistic surface behavior in some wave functions at the cost of introducing more serious errors in others.

In our judgement it is advisable to retain the advantages of harmonic-oscillator dependence, where possible, until rigorous treatments of the single-particle problem, treatments which will presumbably follow along the paths explored in the treatment of single-nucleon transfer reactions,⁵ become available. At the same time, it is also advisable to ascertain the degree to which conclusions based on the harmonic oscillator are subject to change when alternate models are used. To this end we have explored the consequence of using the radial-matrix elements of r^4 which are obtained from the Hartree-Fock calculations of Brown and Massen.⁶ Some results from this study are presented in the Appendix.

III. RESULTS AND COMPARISON WITH EXPERIMENT

We first consider the strongest experimentally measured E4 transitions together with the matrix elements A_p and A_n in order to determine the empirical effective charges from the relationship

$$\begin{split} B(E4) &= (2J_i + 1)^{-1} M_{p}^{2} \\ &= (2J_i + 1)^{-1} (A_{p} e_{p} + A_{n} e_{n})^{2}. \end{split}$$

This comparison is made in Table III in which the matrix elements M_p have been rewritten in the form

$$M_{p} = \frac{1}{2}(A_{p} + A_{n})(e_{p} + e_{n}) + \frac{1}{2}(A_{p} - A_{n})(e_{p} - e_{n})$$
(15)

in order to show that the isovector quantity $A_p - A_n$ is usually small relative to the isoscalar quantity $A_p + A_n$. (An exception to this occurs in 22 Ne.) For this reason $e_p - e_n$ cannot be reliably determined from the existing experimental data and we have set it equal to the bare value $e_p - e_n = e$.

The values of $e_p + e_n$ which serve to equate the shell-model predictions to the measured value for each experimentally observed transition are listed as the last column in Table III. All cluster around a value of $e_p + e_n = 2.0e$. This empirical value is much larger than the theoretical value of $e_p + e_n = 1.15e$ calculated by Horikawa *et al.*⁷ Part of this discrepancy may be due to the fact that Horikawa *et al.* ssumed unperturbed energies ($\Delta E \simeq 2\hbar\omega$, $4\hbar\omega$,...) for the one-particle-one-hole states and if there were a collective "giant" 4⁺, T = 0 state it could occur lower in energy, which would thus enhance its admixture into the model-space wave functions and make the isoscalar effective charge larger.

Analyses of the experimentally measured B(E4)for the $\frac{19}{2} \rightarrow \frac{11}{2}$ transition in ⁵³Fe have suggested that the E4 effective charge is small and the E6 effective charge is even negative.⁸ The present results suggest that these "anomalies" in ⁵³Fe may originate within the valence (fp) shell configurations via effects which have not been accurately calculated in previous shell-model calculations because of the large dimensions involved, rather than being a consequence of systematic, state-independent quenching of the effective charge.

We present in Tables IV-VII calculated energies and values of A_p , A_n , M_p , M_n , and B(E4) for ΔJ^{π} = 4* excitations, along with what experimental information exists. Rather complete sets of predictions are given for $\Delta T = 0$ and $\Delta T = 1$ excitations in N = Z nuclei (Tables IV and V, respectively) and for the $\Delta T = 0+1$ excitations of the T = 1states in $T_z = -1$ nuclei (Table VI). For stable even-odd nuclei we present (in Table VII) results only for ¹⁹F and ²⁵Mg, for which some experimental data exist. The values of M_p , M_n , and B(E4) are based upon the values of the added effective charges $\delta_{pp} = \delta_{nn} = \delta_{np} = \delta_{pn} = 0.5$ as motivated by the results of Table III.

In comparing theoretical and experimental values of B(E4) it is important to note that there are several theoretical sources of uncertainty in the values quoted for the experimental B(E4). The B(E4) values are obtained from inelastic electron scattering data by an extrapolation of the form factors measured at finite values of the momentum transfer q to the value for $q = \omega$ by using a model for the shape of the form factor. All results which we quote have been obtained with the Tassie model, ⁹ which is probably quite adequate for the strong E4 transitions but potentially questionable for the weak E4 transitions, where unusual shapes in the transition density may result from cancellation between single-particle form factors. In

	$J_i \rightarrow J_f$	$[B(E4)]^{1/2}(e^{2} \mathrm{fm}^{4})$ Theo ry	Experiment ^a	$\frac{e_{p}+e_{n}^{b}}{e}$
¹⁹ F	$\frac{1}{2}^{+} \rightarrow \frac{9}{2}^{+} (1)$	$46.8(e_p + e_n) \\ -17.2(e_p - e_n)$	69 ± 8	1.8±0.2
20 Ne	0 ⁺ → 4 ⁺ (1)	$106(e_p + e_n)$	195 ± 20	1.8 ± 0.2
²² Ne	$0^{+} \rightarrow 4^{+}(1)$	$34.4(e_p + e_n) + 42.0(e_p - e_n)$	130 ± 15	2.6 ± 0.4
^{24}Mg	$0^+ - 4^+(2)$	$105(e_p + e_n)$	207 ± 15	2.0 ± 0.2
²⁶ Mg	$0^+ \rightarrow 4^+(2)$	$65.0(e_p + e_n) \\ -11.2(e_p - e_n)$	161 ± 22	2.7 ± 0.3
^{28}Si	0 ⁺ - 4 ⁺ (1)	$79.6(e_p + e_n)$	${\bf 134\pm7}$	1.7 ± 0.1

TABLE III. Isoscalar effective charges for the stronger experimentally measured *E*4 transitions.

^a For experimental references see Tables IV, VI, and VII.

^b $(e_p - e_n) = e$ assumed.

		E_f (1	MeV)	A, ^a	M_{b}^{b}	B(E4)	$(e^2 \text{fm}^8 \times 10^3)$	Method of	
	$J_i \rightarrow J_f$	th	exp ^c	(fm ⁴)	(fm ⁴)	th	exp	exp. analysis	
20 Ne	0 ⁺ → 4 ⁺ (1)	4.13	4.25 ^d	106.1	212	45	38 ± 8^{e}	DUELS ^h	
	4+ (2)	9.86	9.03 ^d	-20.2	-40	1.6			
	4+(3)	10.97		-3.8	-8	0.06			
	4+ (4)	11.70		-1.1	-2	0.004			
	4+(5)	14.29		-1.4	-3	0.009			
	4 ⁺ (6)	14.63		13.3	27	0.7			
^{24}Mg	$0^+ \rightarrow 4^+(1)$	4.42	4.12	-19.7	39	1.6	2.0 ± 0.3^{f}	DUELSh	
	4+(2)	5.89	6.01	104.9	210	44	43 ± 6^{f}	DUELSh	
	4+ (3)	8.79		12.5	25	0.6			
	4+(4)	9.62		-0.8	1.7	0.003			
	4 ⁺ (5)	11.15		12.8	26	0.7			
	4+ (6)	12.18		-9.1	-18	0.3			
	4 ⁺ (7)	12.25		-14.1	-28	0.8			
	4+ (8)	12.63		-6.8	-14	0.2			
$^{28}{ m Si}$	$0^+ \rightarrow 4^+(1)$	4.88	4.62	-79.6	-159	25	27±5 ^g	g	
	4+(2)	7.49	6.89	85.5	171	2 9			
	4+ (3)	9.70		16.4	33	1.1			
	4+ (4)	10.15		16.3	33	1.1			
	4+(5)	10.50		-13.0	-26	0.7			
	4 ⁺ (6)	11.06		1.3	3	0.009			
	4+(7)	11.29		-3.8	-8	0.06			
	4*(8)	11.93		7.8	16	0.3			
^{32}S	$0^+ \rightarrow 4^+(1)$	4.82	4.46	-121.4	-243	5 9			
	4+(2)	6.65	6.41	67.7	135	18			
	4+(3)	7.55		20.0	40	1.6			
	4+(4)	7.94		-8.7	-17	0.3			
	4+(5)	9.00		17.2	34	1.2			
	4 ⁺ (6)	9.36		-14.8	-30	0.9			
	4+(7)	9.98		19.6	39	1.5			
	4+(8)	10.40		7.8	16	0.3			
³⁶ Ar	$0^+ \rightarrow 4^+(1)$	4.74	4.41	-123.5	-248	62			
	4*(2)	6.81		3.3	6.5	0.04			
	4+(3)	9.07		-30.5	-61	3.7			
	4+(4)	9.98		18.6	37	14			
	4+(5)	11.75		16.6	33	1.1			
	4+(6)	12.64		22.7	45	2.0			

TABLE IV. $\Delta T = 0$ E4 transitions in N-Z nuclei.

 $^{a}A_{n}=A_{p}.$

 $b M_n = M_p$.

^c From Ref. 21 except where noted.

^d Reference 32.

e Reference 10.

^f Reference 11.

^g The results given in Ref. 17 for the PWBA B(E4) values in ²⁴Mg and ²⁸Si are 28.7×10³ and (17.8±2.0)×10³ e^{2} fm⁸, respectively, whereas a DWBA value of $(43\pm6)\times10^{3} e^{2}$ fm⁸ is given in Ref. 11 for ²⁴Mg. For ²⁸Si we have used $B(E4) = [(43\pm6)/28.7] \times (17.8\pm2.0)\times10^{3} e^{2}$ fm⁸. ^h Reference 16.

Reference 10.

addition, the E4 transition strengths have been extracted from simple one-step, plane-wave Born approximation (PWBA) calculations. The results we quote from Refs. 10-15 have been corrected for distorted-wave effects with the computer code DUELS.¹⁶ For a few cases, only the PWBA analysis is available (Refs. 17-19). B(E2) values extracted from a distorted-wave Born approximation (DWBA) analysis are typically about 30% larger than those extracted from a PWBA analysis¹⁷ and the difference is probably larger for B(E4) values (see Sec. 5.5 in Ref. 15). Two-step excitations such as $0 \rightarrow 2 \rightarrow 4$ have not been considered and this must be important on some level for the transi-

		E_f (MeV)	A, ^b	M _p ^c	$B(E4)$ ($e^2 \mathrm{fm}^8 \times 10^3$)
	$J_i \rightarrow J_f$	th	exp	(fm ⁴)	(fm ⁴)	th
²⁰ Ne	$0^+ \rightarrow 4^+(1) T = 1$	10.86	11.07 ^d	27.7	27.7	0.77
	$4^{+}(2)T = 1$	13.88		8.4	8.4	0.07
	$4^{+}(3)T = 1$	15.74		-3.9	-3.9	0.015
	$4^{+}(4)T = 1$	16.22		-9.7	-9.7	0.09
^{24}Mg	$0^+ \rightarrow 4^+(1) T = 1$	9.74	9.51^{e}	16.4	16.4	0.27
-	$4^{+}(2)T = 1$	12.28		26.5	26.5	0.70
	$4^{+}(3) T = 1$	12.89		25.1	25.1	0.63
	$4^{+}(4) T = 1$	13.36		-1.1	-1.1	0.001
^{32}S	$0^+ \rightarrow 4^+(1) T = 1$	10.08	10.15^{f}	-20.6	-20.6	0.42
	$4^{+}(2) T = 1$	10.78		-75.5	-75.5	5.7
	$4^{+}(3)T = 1$	11.85		15.7	15.7	0.25
	$4^{+}(4) T = 1$	12.41		26.3	26.3	0.69
³⁶ Ar	$0^+ \rightarrow 4^+(1) T = 1$	10.32		-7.5	-7.5	0.06
	$4^{+}(2)T = 1$	12.50		31.0	31.0	0.96
	$4^{+}(3)T = 1$	12.76		73.1	73.1	5.3
	$4^{+}(4) T = 1$	14.67		13.3	13.3	0.18

TABLE V. $\Delta T = E4$ transitions in N = Z nuclei.^a

^a The theoretical results for ²⁸Si have not been calculated. ^b $A_n = -A_p$. ^c $M_n = -M_p$. ^dReference 32.

^e Based on the energy of the $1^+T = 1$ state in ³²S and the $4^+T = 1$ state in ³²P (Ref. 21). ^f Reference 21.

TABLE VI. E	transitions	for $T_{z} = -1$	nuclei	(T = 1)	-T = 1).
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2 2 1	$J_i \rightarrow J_f$	E_f th	(MeV) exp ^a	А _р (fm ⁴)	A_n (fm ⁴)	<i>M_p</i> (fm ⁴)	M _n (fm ⁴)	<i>B</i> (E4 th) (e ² fm ⁸ ×10 ³) exp	Method of exp analysis
¹⁸ O	$0^+ \rightarrow 4^+(1)$ $4^+(2)$	$\begin{array}{c} 3.52\\ 8.21\end{array}$	3.56 ^b	0 0	-107.1 - 81.5	- 54 -41	-161 -122	2.9 1.7	$1.04 \pm 0.20^{\circ}$	DUELS ^f
²² Ne	$0^+ \rightarrow 4^+(1)$ $4^+(2)$ $4^+(3)$ $4^+(4)$	3.42 5.40 6.22 7.21	3.36 5.52 6.34 (7.34)	-76.3 38.4 -50.9 -13.9	7.6 69.7 -51.6 62.9	-111 92 -102 11	-27 124 -103 87	$12.3 \\ 8.5 \\ 10.4 \\ 0.3$	17 ± 4^{d}	DUELS ^f
²⁶ Mg	$0^+ \rightarrow 4^+(1)$ $4^+(2)$ $4^+(3)$ $4^+(4)$	4.59 5.41 5.87 6.13	4.32 4.90 5.47 5.72	71.6 -53.8 2.1 -60.0	45.3 -76.2 66.1 -8.2	130 -119 36 -94	104 -141 100 -42	$16.9 \\ 14.1 \\ 1.3 \\ 8.8$	26 ± 7^{e} 3.4 ^e 13 ± 5 ^e	DUELS ^f DUELS ^f DUELS
³⁰ Si	$0^+ \rightarrow 4^+(1)$ $4^+(2)$ $4^+(3)$ $4^+(4)$	5.40 6.02 7.52 8.23	5.28 5.95 7.22	36.1 111.9 54.5 15.4	107.6 66.2 45.3 -9.2	108 201 59 18	180 155 41 -6	$11.7 \\ 40 \\ 3.5 \\ 0.3$		
³⁴ S	$0^+ \rightarrow 4^+(1)$ $4^+(2)$ $4^+(3)$ $4^+(4)$	4.88 6.90 7.23 7.76	4.69 6.25	$101.3 \\ 58.4 \\ 37.6 \\ 56.8$	100.6 -44.4 59.1 -34.6	202 65 86 68	201 37 107 23	41 4.2 7.4 4.6		
³⁸ Ar	$0^+ \rightarrow 4^+(1)$ $4^+(2)$	7.79 14.88	5.35	-172.9 -11.4	0 0	-259 -17	-86 -6	67 0.3		

^a From Ref. 21 except as noted. ^b Reference 32. ^c Reference 12.

^d Reference 13.

^e Reference 14.

^f Reference 16.

		E_f (MeV)					B (E4)	$(e^2 \text{fm}^8 \times 10^3)$	
	$J_i \rightarrow J_f$	th	exp ^a	A _p (fm ⁴)	A_n (fm ⁴)	Μ _p (fm ⁴)	<i>M_n</i> (fm ⁴)	th	exp	Method of exp analysis
19 _F	$\frac{1}{2}^+ \rightarrow \frac{7}{2}^+ (1)$	4.64	4.38 ^b	-13.1	-2.3	-21	-10	0.44		
	$\frac{7}{2}^{+}(2)$	5 .9 5	5.46 ^b	-36.6	-66.2	-88	-118	7.7		
	$\frac{7}{2}^{+}(3)$	6.43	6.07 ^b	-26.6	-35.1	-57	-66	3.2		
	$\frac{7}{2}^{+}(4)$	8.36		10.7	23.2	28	40	0.8		
	$\frac{1}{2}^+ \rightarrow \frac{9}{2}^+(1)$	2.56	2.78^{b}	42.3	91.2	109	158	5.9	4.7 ± 1.1^{c}	PWBA
	$\frac{9}{2}^{+}(2)$	6.73	6.59 ^b	-16.0	-33.6	41	58	0.8		
	$\frac{9}{2}^{+}(3)$	8.27		14.7	-5.9	18	-1.7	0.17		
	$\frac{9}{2}^{+}(4)$	10.74		-7.9	43.1	9.6	61	0.05		
$^{25}\mathrm{Mg}$	$\frac{5}{2}^+ - \frac{3}{2}^+$ (1)	1.17	0.97	-16.7	-43.5	-47	-74	0.37		
	$\frac{3}{2}^{+}(2)$	3.08	2.80	-16.9	19.2	-16	20	0.043	$(1.5 \pm 0.4)^{d_{\bullet}e}$	PWBA
	$\frac{3}{2}^{+}(3)$	4.71	4.36	-2.9	-10.1	-9	-17	0.014		
	$\frac{3}{2}^{+}(4)$	6.04		12.8	18.6	28	34	0.13		
	$\frac{5^+}{2} \rightarrow \frac{5^+}{3}$ (2)	2.06	1.96	-27.3	16.4	-33	11	0.18		
	$\frac{5}{2}$ (3)	4.24	3.91	65.2	91.5	144	170	3.5	1.9 ± 0.4^{d}	PWBA
	$\frac{5}{2}^{+}(4)$	5.06		-6.6	0.6	-10	-2.4	0.017		
	$\frac{5}{2} \rightarrow \frac{7}{2} (1)$	1.75	1.61	-27.0	-20.3	-51	-44	0.43		
	$\frac{7^{+}}{2}$ (2)	3.02	2.74	-59.1	-101.8	-140	-182	3.3	$1.5 \pm 0.4^{d,e}$	PWBA
	$\frac{7}{2}$ (3)	5.18	5.01	-42.7	12.1	-58	-3.2	0.56		
	$\frac{7}{2}^{+}(4)$	5 .9 4	5.98	-14.6	-4.3	-24	-14	0.010		
	$\frac{5^+}{2} \rightarrow \frac{9^+}{2}(1)$	3.62	3.41	13.0	-5.4	17	-1.7	0.05		_
	$\frac{9}{2}^{+}(2)$	4.08	4.06	68.1	63.0	135	127	3.0	1.3 ± 0.3^{d}	PWBA
	$\frac{9}{2}^{+}(3)$	4.85	4.71	-21.0	15.9	-24	13	0.09		
	$\frac{9^+}{2}(4)$	6.56	5.97	54.1	66.1	115	125	2.2		
	$\frac{5}{2} \rightarrow \frac{11}{2} (1)$	5.26	(5.25)	139.5	123.3	271	255	12.2	5.8 ± 1.2^{d}	PWBA
	11 +								13 ± 6 ^f	DUELS ^g
	$\frac{11}{2}$ (2)	5.80	5.53	-6.1	6.2	-6	6	0.006		
	$\frac{11}{2}$ (3)	6.28	(6.04)	20.5	-19.3	21	-19	0.074		
	$\frac{11}{2}$ (4)	7.70		6.1	-11.7	3.3	-14	0.003		
	$\frac{1}{2} \rightarrow \frac{1}{2}$ (1)	5.73	5.46	37.0	-0.2	56	17	0.5		
	$\frac{10}{2}$ (2)	7.87		17.1	19.1	34	37	0.2		
	$\frac{13}{2}$ (3)	8.0 9		-85.0	-54.6	-154	-125	3 .9		
	$\frac{13}{2}$ (4)	8 .9 6		-38. 9	-26.7	-71	-76	0.8		

TABLE VII. Selected E4 transitions in odd-even nuclei ($T = 1/2 \rightarrow T = 1/2$).

^a From Ref. 21 except where noted. ^b Reference 32. ^c Reference 18. ^d Reference 19. ^e The $\frac{3^+}{2}$ (2) and $\frac{7}{2}$ (2) states were not resolved in the electron scattering experiment of Ref. 19. ^f Reference 15.

^gReference 16.



FIG. 1. Theoretical values for the $[B(E4), 0^* \rightarrow 4^*]^{1/2}$ matrix elements as a function of excitation energy for N=Z nuclei in the *sd* shell. Both $\Delta T = 0$ and $\Delta T = 1$ transitions are shown. The numbers are taken from Tables IV and V; $\Delta T = 1$ transitions in ²⁸S have not been calculated.

tions which have a very weak one-step E4 matrix element.

The calculated E4 strength distributions for N=Z nuclei are shown in Fig. 1 and compared with experiment in Tables IV and V. The identities of the observed strong transitions in the T = 0 nuclei ²⁰Ne, ²⁴Mg, and ²⁸Si are correctly predicted, in particular the shift of the dominant E4 strength from the lowest 4⁺ state in ²⁰Ne to the second 4⁺ state in ²⁴Mg and back (partially) to the lowest 4* state in ²⁸Si. Moreover, the magnitudes of these three strong transitions are reproduced to within experimental error. To pursue this particular aspect of the E4 picture further it is important to obtain experimental measurements of the second 4⁺ state in ²⁸Si, of the first two 4⁺ states in ³²S, and of the first 4^{+} state in 36 Ar. It would also be valuable to measure or to set experimental upper

bounds on the strength of the higher excited 4⁺ states in the systems. Future experimental studies on these T = 0 nuclei should also concentrate efforts on attempting the formidable task of extracting strengths for $\Delta T = 1$ excitations to J^{π} , $T = 4^{+}$, 1 states, the predictions for which are given in Table IV.

The predicted distributions of E4 strength in the $T_z = -1$ nuclei are more complex than for the $T_z = 0$ systems and the individual values are smaller. The experimental data, single values for ¹⁸O and ²²Ne and three values for ²⁶Mg, confirm these predictions about fragmentation and average magnitudes, but except for ¹⁸O the experimental B(E4) values tend to be larger than the predicted values by about 1 to 1.5 standard deviations. However, because of the varying structures of these excitations, this trend is not to be attributed to any simple isovector effective charge renormalization. More experimental measurements of E4 strength values of these systems would go far toward firmly validating the general picture of E4 strengths in T = 1 systems predicted by the present stage of nuclear structure theory as well as to provide specific and unique information on the salient characteristics of the individual $J^{\pi} = 4^+$ states.

The spectra of E4 excitations which can be obtained with odd-mass targets are, of course, potentially richer and more complex than for the $T_z = -1$, let alone the $T_z = 0$ systems. Experimental electron scattering data have been reported for 19 F (Ref. 18) and 25 Mg (Refs. 15 and 19); results of our calculations for these two nuclei are presented in Table VII. Experiment and theory are in good agreement for the $\frac{1}{2}$ + $-\frac{9}{2}$ +(1) transition in ¹⁹F; other $\frac{1}{2}$ + $\frac{9}{2}$ + E4 transitions up to 11 MeV in excitation are predicted to be weak. Only a few of the $\frac{5}{2}$ + J^{T} transitions in ²⁵Mg are predicted to be strong and these are in general not to the lowest level of each spin. The relative strengths of the B(E4) values reported by Okazaki *et al*. (Ref. 19) are in excellent agreement with the calculations. There is an overall factor of 2 discrepancy between experiment and theory. However, this is probably associated with the fact that a PWBA analysis was used by Okazaki et al. to extract the E(E4) values (see the remarks in Sec. 5.5 of Ref. 15); the experimental B(E4) which we have used for other nuclei have been obtained using a DWBA analysis.¹⁶ Experimental results for two other transitions which are predicted to be strong in ²⁵Mg have not been reported [the $\frac{5}{2}$ * $+\frac{9}{2}$ *(2) and $\frac{5}{2}$ * $+\frac{13}{2}$ *(3) transitions]. There is obviously an extensive amount of work to be done on this component of the E4 picture. Ultimately, the experimental delineation of the distributions

and magnitudes of E4 strength in odd-mass nuclei as a function of A, excitation energy, and final state spin, and the concurrent testing of the analogous nuclear structure model predictions will significantly enlarge our understanding both of nuclear wave functions and the foundations of the E4 excitation mode.

IV. STRENGTH' RATIOS FOR EXCITATION BY ELECTRO-MAGNETIC AND VARIOUS HADRONIC PROBES

In this section we derive some estimates of the matrix elements associated with inelastic hadron scattering and present some sample results for proton, alpha, and pionic probes. The calculations are based upon the assumption, which has proved to be remarkably successful in describing alpha scattering,²⁰ that the cross sections for hadronic excitation are proportional to the same matrix elements which enter into the calculations of electromagnetic transition rates. Thus these cross sections are proportional to

$$B'(a) = (2J_i + 1)^{-1} [C_{\flat}(a)M_{\flat} + C_n(a)M_n]^2, \qquad (16)$$

where $C_{p/n}$ represents the interaction strengths of probe (a) with the protons/neutrons of the nucleus.

Since we are concerned here with nuclear structure effects rather than the details of the probespecific absolute interaction strengths we choose to work with normalized strength factors B(a) defined so as to yield a value for $T = 0 \rightarrow T = 0$ transitions (for which $M_p - M_n = 0$) which is unique, independent of probe (a):

$$B(a) = B'(a) / [C_{p}(a) + C_{n}(a)]^{2}$$

= $(2J_{i} + 1)^{-1} [\frac{1}{2} (M_{p} + M_{n}) + \frac{1}{2}R(a)(M_{p} - M_{n})]^{2}, (17)$

where

$$R(a) = \frac{C_{p}(a) - C_{n}(a)}{C_{p}(a) + C_{n}(a)}.$$
 (18)

For the electromagnetic probes, $C_{p} = e$ and $C_{n} = 0$ and hence R(a=e)=1, and Eq. (17) reduces to Eq. (6),

$$B(E4) = B(e) = (2J_i + 1)^{-1}M_b^2.$$
(19)

For the alpha particle, $C_p = C_n$ and hence R(a = alpha) = 0 and

$$B(alpha) = (2J_i + 1)^{-1} \left[\frac{1}{2} (M_p + M_n) \right]^2.$$
 (20)

For pions at energies near the (3,3) resonance $R(a=\pi^*)=\pm 0.5$ and

$$B(\pi^{\pm}) = (2J_i + 1)^{-1} [\frac{1}{2} (M_p + M_n) \pm \frac{1}{4} (M_p - M_n)]^2 . \quad (21)$$

Pion inelastic scattering experiments are currently designed to yield the ratios of excitation strengths with π^- and π^+ projectiles. This quantity is expressed in our model as

$$\frac{B(\pi^{-})}{B(\pi^{+})} = \left(\frac{M_{p} + 3M_{n}}{3M_{p} + M_{n}}\right)^{2}.$$
(22)

This precise result depends, of course, upon the assumption that near resonance the π^- -n and π^+ -p interaction amplitudes are 3.0 times stronger than the π^- -p and π^+ -n amplitudes, but analogous expressions are easily obtained for different as-sumptions about these elementary-particle features.

Many strongly excited 4^* states at high excitation energy have been reported in alpha inelastic scattering experiments. In Table VIII are listed some strengths from alpha scattering on ²⁴Mg (Ref. 20) in comparison with the present theoretical predictions for the first nine 4^* states. Above 9.0 MeV there is complete disagreement between experiment and theory. This could be due to experimental errors in the spin assignments (in Endt and van der Leun²¹ no definite 4^* assignments have been made above 9.0 MeV) or to two-step contributions to the inelastic scattering cross sections which were not considered in the DWBA analysis, rather than a failure of the nuclear structure wave functions.

Some predicted values for $B(\alpha)$ and $B(\pi^{\pm})$ in $T_z = -1$ nuclei are given in Table IX. It should be quite interesting to carry out the corresponding experimental measurements since several quite strong deviations from isoscalar behavior are predicted. Inelastic alpha scattering to 4⁺ states in ²²Ne and ²⁶Mg as well as in several N=Z nuclei has been measured by Rebel *et al.*²² but the extraction of $B(\alpha)$ values via the methods of Ref. 20 have not been carried out.

As a final example of hadronic excitations we compare some experimentally measured proton

TABLE VIII. Alpha scattering transition strengths for $0^+ \rightarrow 4^+$ transitions in ²⁴Mg.

$J_f(T_f)$	Energy th	v (MeV) exp ^a	B(alpha th) (fm ⁸ ×10 ³) exp ^a	
4+(0)	4.42	4.12	1.6	· ·	
4+(0)	5.89	6.01	44.0	21.4 ± 3.2	
4+(0)	8.79	8.64	0.6	3.0 ± 0.5	
4 *		9.12		8.1 ± 1.4	
(4+)		9.30		12.2 ± 1.9	
4+		10.03		16.0 ± 2.4	
4*(0)	9.62		0.002		
4 ⁺ (1)	9.74				
4*(0)	11.15		0.6		
4+(0)	12.17		0.3		
4+(0)	12.24		0.7		
4+(1)	12.28				

^a $B(alpha) = (A/2Z)^2 B(IS)$ where B(IS) are taken from Table IX in Ref. 20.

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Nucleus	$J_i \rightarrow J_f$	$B(\text{alpha}) \text{ (fm}^8 \times 10^3)$	$\frac{B\text{(alpha)}}{B(E4)}$	$\frac{B(\pi^{-})}{B(\pi^{+})}$
¹⁸ O	$0^+ - 4^+(1)$	11.5	4.00	2.78
22 Ne	$0^+ - 4^+(1)$	4.7	0.38	0.28
	4+ (2)	11.7	1.37	1.34
²⁶ Mg	$0^+ - 4^+(1)$	13.7	0.80	0.80
	4+ (2)	16.9	1.20	1.19
^{30}Si	$0^+ \rightarrow 4^+(1)$	20.7	1.77	1.65
	4+ (2)	31.8	0.78	0.77
^{34}S	0 ⁺ -+4 ⁺ (1)	40.8	1.00	1.00
	4+(2)	0.2	0.05	0.09
^{38}Ar	$0^{+} \rightarrow 4^{+}(1)$	30.0	0.44	0.36

TABLE IX. Alpha scattering and pion scattering strengths for the $0^+ \rightarrow 4^+$ transitions in $T_Z = -1$ nuclei.

scattering strengths in ²⁶Mg (Ref. 23) with estimates for 30 MeV proton scattering. From the microscopic optical model of Brieva and Rook²⁴ we obtain values for the interaction strengths at this energy of $C_p = 0.3$ and $C_n = 0.7$. This yields

 $B(\text{proton}, 30 \text{ MeV}) = (2J_i + 1)^{-1} (0.3M_p + 0.7M_n)^2.$ (23)

The predicted and measured *B* values are presented in Table X. The experimental $0^* - 4_2^*$ strength is twice as large as the predicted value. This is probably a real discrepancy since, in addition, the experimental electron scattering B(E4) value is larger than predicted (see Table V). The predicted ratio B(p)/B(E4) is consistent with experiment. The situation for the weaker $0^* - 4_1^*$ transition is not so clear since the B(E4) value has not yet been extracted from the experimental data; the calculated B(p) value is in fair agreement with experiment. The proton scattering strength to the third 4^* state is predicted to be much stronger than the electromagnetic excitation strength.

V. SUMMARY AND CONCLUSIONS

The Chung-Wildenthal wave functions for the sd-shell nuclei have been used to calculate hexadecapole transition densities. These have been combined with harmonic-oscillator radial wave functions to predict proton and neutron E4 matrix elements within the sd-shell model space. There is at present no good theoretical prescription for obtaining radial wave functions for open-shell nuclei which are more realistic than those of the harmonic-oscillator potential.

The effects of model-space truncation on the E4 operator have been taken into account by introducing empirical effective charges. The experimental values for strong electromagnetic excitations are all consistent with a isoscalar effective charge of $e_p + e_n \simeq 2.0e$. This is much larger than the results of microscopic core-polarization calculations of Horikawa *et al.*,⁷ which gave $e_b + e_n \simeq 1.15e$.

These empirical effective charges were used to calculate the proton and neutron matrix elements M_p and M_n ; electromagnetic excitation strengths are determined by M_p while hadronic scattering strengths depend on both M_p and M_n . Predicted electromagnetic B(E4) values are in good agreement with existing experiments over a wide range of magnitudes. However, experimental electron scattering results are lacking for the upper sd shell as well as for pure $\Delta T = 1$ excitations in N = Z nuclei.

It is more difficult to extract nuclear structure information from the hadronic scattering cross sections. For transitions which are characterized by small direct matrix elements, two-step contributions are certainly important; this may be responsible for the disagreement between the very small isoscalar excitation strengths predicted to 4^* states above 9 MeV in ²⁴Mg compared with experiment. A systematic analysis of alpha, proton, or pion scattering for $0^* - 4^*$ transitions in many *sd*-shell nuclei is needed to confirm our predic-

TABLE X. Proton scattering transition strengths B(p) = B (proton, 30 MeV) for the $0^+ \rightarrow 4^+$ transitions in ²⁶Mg.

Energy	7 (MeV)	B(p) (f	$m^8 \times 10^3$)	B(p)/B(E4)
th	exp ^a	th ^b	exp ^c	th	exp^d
4.59	4.32	12.5	18±2	0.74	
5.41	4.90	18.1	41 ± 2	1.28	1.6 ± 0.3
5.87	5.47	6.5		5.0	
6.13	5.72	3.3		0.38	

^a Reference 21.

 ${}^{b}B(p) = (0.3M_{p} + 0.7M_{n})^{2}.$

^cReference 23.

^d For $B(E4)_{exp}$ see Table VI.

	Harmonic	Hartree-Fock		
	oscillator	Skyrme III	Skyrme IV	
$\langle d_{5/2} r^4 d_{5/2} \rangle_p (\mathrm{fm}^4)^a$	179	181	198	
$\langle d_{5/2} r^4 d_{3/2} \rangle_p (\text{fm}^4)^a$	179	236	274	
$r_p (fm)^b$	2.905	2.905	2.902	
$BE(a_{5/2})_p$ (MeV) ^c		-5.42	-5.62	
$BE(d_{3/2})_p$ (MeV) ^c		-0.02	0.78	

TABLE XI. Hartree-Fock results for ²⁰Ne.

^a Proton radial-matrix elements.

^b Point proton rms radii.

^c Proton single-particle energies.

tions for the relative sizes and phases of the proton and neutron matrix elements.

APPENDIX

In this section we discuss the Hartree-Fock calculations for ²⁰Ne and ³⁶Ar which have recently been carried out by Brown and Massen.⁶ For these calculations a Skyrme interaction²⁵ was used and a self-consistent spherical potential was constructed out of the monopole densities obtained by adding the squares of the radial wave functions weighted by the orbit occupation probabilities, which in turn had been obtained from the *sd*-shellmodel calculations. The spin-orbit parameter was chosen to reproduce the $d_{5/2}$ - $d_{3/2}$ spin-orbit splitting of about 5 MeV in A = 17.

The results using Skyrme III and Skyrme IV interactions for ²⁰Ne are given in Table XI. Under the column labeled "harmonic oscillator" we give the r^4 matrix elements which we would obtain if we took the theoretical value of $r_p = 2.905$ fm and interpreted it as a sum of harmonic oscillators to extract b and then to calculate $\langle r^4 \rangle = 15.75b^4$. For the $d_{5/2}$ orbit the Hartree-Fock results are very close to the harmonic-oscillator results, but the Hartree-Fock results for the off-diagonal, $d_{5/2}-d_{3/2}$ matrix element are much larger than the harmonic-oscillator result and they depend strongly on the value of the binding energy. (The radial wave function for the unbound $d_{3/2}$ state obtained with the Skyrme IV interaction was obtained by the method described by Beiner *et al.*²⁵)

The experimental separation energies of the $d_{5/2}$ and $d_{3/2}$ orbits in ²⁰Ne, -16.87 and -18.42 MeV, respectively, are very different from the Hartree-Fock single-particle energies because of the residual interactions. Thus, the tails of the wave functions, which are important for the r^4 matrix elements and which are determined by the separation energy, cannot be reliably calculated with spherical Hartree-Fock or any other smoothly mass dependent single-particle wave functions. These large separation energies may even result in $\langle r^4 \rangle$ matrix elements which are smaller than the harmonic-oscillator results. There is at present no rigorous theoretical prescription for taking the residual interactions into account.

The situation is different when the orbits are more deeply bound. In Table XII the Hartree-Fock

	Harmonic oscillator	Hartree-Fock	
		Skyrme III	Skyrme IV
$\langle d_{5/2} r^4 d_{5/2} \rangle_p$ (fm ⁴)	220	201	196
$\langle d_{5/2} r^4 d_{3/2} angle_{p}$ (fm ⁴)	220	213	219
r_p (fm)	3.318	3.318	3.298
$BE(d_{5/2})_p$ (MeV)		-13.51	-17.52
$BE(d_{3/2})_{p}$ (MeV)	·	-7.26	-6.93

TABLE XII. Hartree-Fock results for ³⁶Ar.^a

^a See footnotes to Table XI.

results for ³⁶Ar are given. For both Skyrme III and Skyrme IV the results are close to the harmonic-oscillator values, even though the $d_{5/2}$ single-particle energies are quite different in the two cases. In this case the experimental separation energies for the $d_{5/2}$ and $d_{3/2}$ orbits of -17.01 and -15.25 MeV, respectively, are fairly close to the Hartree-Fock single-particle energies.

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