Decays of the lowest $T = 2$ states in $A = 4N$ nuclei from ⁸Be to ⁴⁴Ti

.S. J. Freedman,* C. A. Gagliardi, M. A. Oothoudt,[†] A. V. Nero,[†] R. G. H. Robertson,[§] and F. J. Zutavern Princeton University, Princeton, New Jersey 08540

E. G. Adelberger

University of Washington, Seattle, Washington 98195

A. B. McDonald

Atomic Energy of Canada Limited, AECL, Chalk River Nuclear Laboratories, Chalk River, Canada, KOJ 1JO (Received 11 December 1978)

Branching ratios for isospin-forbidden decays have been measured for the lowest $T = 2$ levels in ${}^{8}Be$, ${}^{12}C$, Drancining ratios for isospin-foroident decays have been measured for the lowest $t = 2$ levels in Be, C, 16 O, 24 Mg, 28 Si, 32 S, 36 Ar, 40 Ca, and 44 Ti. The $T = 2$ levels were populated by the isos reaction at $E_n = 42$ or 46 MeV. Coincidences were observed between tritons detected in a quadrupole-dipoledipole-dipole magnetic spectrometer and decay particles detected in a surface-barrier detector telescope, a time-of-flight neutron detector, or an NaI (Tl) γ -ray detector. Systematics of isospin-forbidden decay widths are discussed.

NUCLEAR STRUCTURE ⁸Be, ¹²C, ¹⁶O, ²⁴Mg, ²⁸Si, ³²S, ³⁶Ar, ⁴⁰Ca, ⁴⁴Ti; lowest $J^* = 0^*$, $T = 2$ level; measured isospin-forbidden particle-decay branching ratios.

I, INTRODUCTION

The breaking of isospin symmetry in nuclei is an interesting topic in nuclear physics. The principal concern is whether or not the observed breaking can be explained by the effect of the Coulomb interaction alone. Isospin symmetry breaking is directly probed by studying isospin forbidden decays. If the interaction responsible for the decay conserves isospin then one measures the isospin mixing of the states involved. Moreover, since the final state is generally at low excitation energy in the residual nucleus, it is reasonable to attribute the observed breaking to isospin mixing in the initial state. In those circumstances where the only isospin allowed decay channels are electromagnetic, isospin forbidden particle decay usually dominates. Such systems are particularly useful since experiments can detect very small isospin admixtures. Examples can be found in light and medium weight nuclei; in particular, favorable conditions occur for the lowes $T = \frac{3}{2}$ levels in $A = 4N+1$ nuclei and the lowest $T = 2$ levels in $A = 4N$ nuclei.

The decays of $T = \frac{3}{2}$ levels have been studied extensively and absolute decay widths have been obtained by combining the measured values of branching ratios and total widths. Relatively complete information on partial decay widths for T_z $=(N-Z)/2=\frac{1}{2}$ and $T_z=-\frac{1}{2}$ members of $T=\frac{3}{2}$ multiplets have been assembled.¹⁻⁴ This work has revealed a number of striking and unexpected regularities. Most significant is the discovery that the

variation of decay widths from nucleus to nucleus is not random but displays regular behavior. This indicates that the widths reflect basic features of nuclear charge dependence rather than the properties of "accidental" $T = \frac{1}{2}$ levels which happen to be nearly degenerate with the $T = \frac{3}{2}$ state. It is also found that a $T = \frac{3}{2}$ level with $T_{z} = +\frac{1}{2}$ tends to have greater reduced width^{3,5,6} than the corresponding $T_z = -\frac{1}{2}$ state. This establishes that $\Delta T = 2$ and $\Delta T = 1$ mixing amplitudes are comparable,¹ at least up to $A = 25$ which is the heaviest system for which the neutron decay channel is open. The particle decays predominantly feed excited states of the residual nuclei; presumably this reflects the particle-hole structure of the isospin impurity. However, the ratios of reduced widths for decay to the various excited states are different for T_z $=\frac{1}{2}$ and $T_z=-\frac{1}{2}$, implying that the isospin admixtures include more than one nuclear configuration. Finally, there is a very surprising periodicity⁴ in the proton decays of $T_z = -\frac{1}{2}$ levels to the ground states of the $A=4N$ nuclei. The $A=4N+1$, $T=\frac{3}{2}$ levels with N odd have systematically larger widths than those with N even. This oscillation with a period of eight nucleons was completely unexpected and has not yet been satisfactorily explained. Possible clues to the many unexplained regular

ities associated with the $T=\frac{3}{2}$ decays may be found by examining other related but very different systems, such as the $T = 2$ states of the $A = 4N$ nuclei. In contrast to the well defined decay width for the $T = \frac{3}{2}$ levels only scattered information had been obtained for $T = 2$ states when this work was

19 1907 C 1979 The American Physical Society

begun. Decay branching ratios of the lowest $T = 2$ states in 16 O had been reported by Koh *et al.*,⁷ while McGrath ${et}$ ${al.},^8$ had measured the decay branching ratios in ²⁰Ne, ²⁴Mg, ²⁸Si, ³²S, and ⁴⁰Ca.
Total widths of the $T = 2$ states in ⁸Be.^{9,10}¹⁶O.¹¹ branching ratios in ²⁰Ne, ²⁴Mg, ²⁸Si, ³²S, and
Total widths of the $T = 2$ states in ⁸Be, ^{9,10,16}O, and 20 Ne (Ref. 12) had been obtained from resonance studies. Isospin allowed gamma decays of the lowest $T = 2$ states in ²⁰Ne, ²⁴Mg, ²⁸Si, ³²P, the lowest $T = 2$ states in ²⁰Ne, ²⁴Mg, ²⁸Si, ³²P,
³²S, and ⁴⁴Ti had been observed.¹²⁻¹⁸ It was known that proton ($\Delta T = 1$ or $\Delta T = 2$) and alpha ($\Delta T = 2$) decays occurred with roughly equal frequencies. However, because all the resonance measurements of particle widths were confined to $A \le 20$ it was not possible to determine if the isospin-forbidden particle widths of $T=2$ states had the same A departicle widths of $T = 2$ states had the same $\frac{1}{4}$
pendence as did the $T = \frac{3}{2}$ decays, or if similar periodicities of eight nucleons occurred for the $T = 2$ decays.

In this paper we report on one phase of a program to determine the systematics of the decays of the lowest $T=2$ states in $A=4N$ nuclei, namely, measurements of the branching ratios for charged particle decays in 8 Be, 12 C, 16 O, 24 Mg, 28 Si, 32 S, ^{36}Ar , ^{40}Ca , and ^{44}Ti , neutron decays in ^{8}Be , and ¹⁶O and γ decays in ⁴⁴Ti. Relevant resonance studies have been performed at Seattle and Chalk River, and the decay widths of some of these $T=2$
states will be reported elsewhere.¹⁹⁻²¹ states will be reported elsewhere.¹⁹⁻²¹

II. EXPERIMENTAL METHOD A. Technique

The experiments were performed with beamsof 42 to 46 MeV protons from the Princeton University AVF Cyclotron. The lowest $J^{\pi} = 0^+, T = 2$ levels of 8 Be, 12 C, 16 O, 24 Mg, 28 Si, 32 S, 36 Ar, ⁴⁰Ca, and ⁴⁴Ti were populated by (p, t) reactions and coincidences were observed between tritons detected at $\theta_{lab} \approx +23^\circ$ or $+27^\circ$ in a quadrupoledipole-dipole-dipole (QDDD) magnetic spectrometer and various charged decay particles detected in a surface-barrier detector telescope. The detector telescope was mounted at angles of -90° or -120° in a 20 cm diameter scattering chamber (see Fig. 1). Various surface barrier detectors were used; the telescope was typically about 4 cm from the target. For 8 Be, 12 C, and 16 O, neutron decays were observed in a time-offlight detector having a 12.5 cm diameter by ⁵ cm thick NE102 plastic scintillator and a 69 cm flight path. For 44 Ti, decay γ rays were observed in a 12.5 cm diameter by 10 cm thick NaI (Tl) detector.

After passing through the target, the beam was focused by a quadrupole doublet to a well shielded beam dump located within the concrete wall, about 7 meters from the target. This system was so effective that there was a negligible contribution

FIG. 1. Schematic diagram of the experimental apparatus.

to the counting rates in the neutron and γ -ray detectors from sources other than the target.

The focal plane detection system for the QDDD consisted of a 50 cm long resistive-wire positionsensitive proportional counter, followed by a plastic scintillator. The energy loss in the proportional counter and the total-energy signal from the scintillator were used to distinguish between tritons and other particles of similar rigidity. In addition, an electrostatic deflector between D2 and D3 reduced the background from deuterons and protons. The full solid angle of the QDDD $(\approx 14$ msr} was used. In each case the various multipole elements of the spectrometer were adjusted to correct for kinematic aberration. The typical triton resolution was less than 25 keV full width at half maximum (FWHM). In most cases the resolution was limited by the finite target thickness.

Multiparameter coincidences were recorded event by event on magnetic tape by a Sigma 2 computer. Spectra corresponding to various coincident combinations were monitored on line. In addition, single parameter events corresponding to the focal plane position for identified tritons were recorded but prescaled by a factor of 10 or 100 to reduce computer dead time. By simultaneously recording singles and coincidence information, any portion of the data could be ana-

lyzed to determine. branching ratios. To first order, the results are independent of beam intensity variations or target thickness changes, depending only on the overall detection efficiency of ihe decay-particle detectors, which was measured before and after each run.

All levels studied have $J=0$, hence the decay angular distributions are isotropic in the center of mass of the decaying nucleus and measurement at a single angle for the decay detector is sufficient to determine the branching ratios.

For the charged particle decay measurements coincidence events were composed of a time to amplitude converter (TAC) signal; energy signals from the two or three surface barrier detectors in the telescope; and the focal plane position signal for a properly identified triton. The focal plane position signal was obtained by analog division of the signals from the two ends of the resistive wire detector and triton identification was accomplished by single channel analyzers placed on the scintillator signal and on the summed signals from the two ends of the proportional counter.

The TAC was started by a signal derived from a fast beam-burst pickup, in fast coincidence with signals from the focal plane scintillator, and was

FIG. 2. Time to amplitude converter spectra for coincidences between tritons detected in the focal plane of the QDDD and (a) charged particles detected in the surface-barrier detector telescope (b) γ 's and neutrons detected in the NE102 plastic scintillator.

stopped by a timing signal from any of the three surface barrier detectors. This choice of start signal minimized the TAC dead time because the scintillator rate (-1 kHz) was always less than the rate in the charged particle detector. The technique of using the beam burst for timing yielded a time resolution better than 1 ns (FWHM); it avoided the time slewing due to triton transit time in the QDDD ($\Delta t \approx 70$ ns). The better timing resolution provided an enhanced signal to noise ratio for charged particles and γ rays and allowed us to make high resolution time-of-flight measurements for neutrons.

For the γ decay measurement three parameter events consisted of triton position, TAC and γ -ray energy. For neutron decay, two parameter events consisted of triton position and TAC. A discriminator rejected events in the neutron detector due to proton recoils with energies less than 1 MeV. For the neutron flight path of 0.6S m this restricted neutron flight times to ≤ 50 ns and confined real coincident events within the beam burst repetition time (50 ns). However, the large variation in flight times through the QDDD for tritons ($\Delta t \approx 70$ ns) resulted in real events occurring in two successive periods. Figure 2 shows typical TAC spectra for charged particles and neutrons. Timeof-flight spectra for neutrons were obtained by summing together the two periods containing real events after aligning them by means of the well defined prompt γ -ray produced peaks in the TAC spectrum.

B. Efficiency calibrations

To accurately determine and monitor the effective detection efficiency for charged particle decays we made measurements on a 100% particle decaying state at regular intervals during the ex-'periments. We chose the 8.92 MeV, $J^r = \frac{1}{2}$ leve of ¹³N which was populated via the ¹⁴N(p, d)¹³N reaction at the same energies used in the primary measurements. The 8.92 MeV level with $\Gamma = 230$ keV decays by isospin allowed proton emission to the ground and first excited states of 12 C. Because of the large width it is reasonable to neglect γ emission and assume 100% proton decay for efficiency measurements. Since the spin is $\frac{1}{2}$ the proton decays will be isotropic in the center of mass. A melamine target was used for calibration of the solid target geometry and nitrogen gas was used to calibrate the gas cell geometry. Of course, the magnetic fields of the QDDD and the electronic windows on the focal plane detector had to be readjusted for calibration runs. However, branching ratios were always determined with respect to singles events in the QDDD so that procedure did not increase the uncertainty.

FIG. 3. (a) Focal plane position spectrum of deuterons from the ¹⁴N(*p*,*d*) ¹³N(8.92 MeV) reaction for $E_p = 42$ MeV, $\theta = 23^{\circ}$. (b) Spectrum of protons from decay of ¹³N(8.92 MeV) observed in the particle telescope at $\theta_{1ab} = 120^\circ$.

Figure 3(a) is a spectrum of deuterons from the focal plane detector; the broad, but clearly defined peak corresponding to the ^{13}N , 8.92 MeV level is evident in the figure. Figure 3(b) shows a spectrum of decay protons observed in coincidence with deuterons in the region of the 8.92 MeV level. The peaks corresponding to decays to the ground state and 4.43 MeV level of ^{12}C are well defined, and the sum of the counts in these peaks with kinematic corrections were used to define the overall detection efficiency for charged particles. In all cases, the efficiency obtained agreed within $\pm 20\%$ with the geometrically defined value, indicating relatively small effects due to electronic pileup and dead times. To apply a final correction, however, counting rates were monitored during actual runs and during the associated efficiency calibrations and corrections were made for the difference in dead times and pile up in the two runs. These corrections were always less than 10%. The pileup corrections were checked by studying the number of "singles" events on tape which also contained random coincidences in other ADC's.

The branching ratio for proton decay of the 13 N (8.92 MeV) level to ¹²C (4.43 MeV) was also determined from the particle telescope spectra. We found $R = \Gamma_{b0}/\Gamma_{b1} = 1.6 \pm 0.1$. This result was then used to calibrate the γ -ray detector using the 4.43 MeV γ ray from the first excited state of ¹²C. The γ -ray efficiency was extrapolated to other energies using the tables of Marion and Young²² as discussed in Ref. 23.

The neutron detection efficiency for a similar plastic scintillator was previously measured' for a pulse-height threshold corresponding to a proton recoil energy of 450 keV. The higher threshold $(E_{n} > 1 \text{ MeV})$ used in the present work was set relative to the Compton edge for γ rays from a 22 Na source. We used the tables of Verbinskii et al^{24} to define the relative light output of neutrons and γ rays. The change in detection efficiency caused by the increased threshold was also estimated from these tables.

III. RESULTS

A. 8 Be(27.48 MeV, $T = 2$)

The 27.483 MeV, $(J^{\pi}=0^{+}, T=2)$ level of ⁸Be has a large number of energetically allowed but isospin-forbidden decay channels as shown in Fig. 4. We attempted to identify as many charged particle

FIG. 4. Partial energy level diagram for 8 Be, showing \cdot possible decay modes of 8 Be (27.48 MeV, T = 2).

FIG. 5. Typical particle identification spectrum for the surface-barrier detector telescope.

types as possible by employing a three-element telescope consisting of 16, 35, and 1000 μ m thick surface barrier detectors. (Runs ⁸Be-2 and ⁸Be-3). Protons and tritons penetrate the first thin E detector for energies down to ≈ 1 MeV, while ³He and α particles penetrate down to laboratory energies of about 3.3 MeV. In principle these energies become the lower threshold for particle identification but in practice the threshold for unambiguous identification is somewhat higher because of noise in the second detector. The three element telescope allowed us to identify the triton and 'He decays of the $T = 2$ state as well as α decays to the 20.2 MeV $(0^+, 0)$ excited state of ⁴He. The time-of-flight neutron detector was used to search for decays to ⁷Be. An initial run (Run ⁸Be-1) with a two-element charged particle detector telescope high statistics for energetic charged particle de-(35 and 275 μ m) and no neutron detector provided cays.

Charged particle identification was achieved by comparing energy losses in the various elements of the telescope with known energy-range data b^y a "table-lookup" procedure. A typical identification spectrum is shown in Fig. 5 for the simple two element telescope. For the three-element

FIG. 6. Foeal plane position spectrum of tritons for 45.8 MeV protons incident on the 10 Be O target.

FIG. 7. (a) Spectrum of coincidences between tritons identified in the QDDD focal plane detector and deuterons reflective the surface-barrier detector telescope. The Be $(27.48 \text{ MeV}, T = 2)$ level. Counts to the left of the enhancements near triton channel 35 are due to the dashed line correspond to particles which stop in the $t-d$ coincidences between the indicated solid lines corfront detector and cannot be identified. (b) Spectrum of responding to the ¹⁰Be $(p, td_1)^6$ Li(3[†]) reaction, projected on the triton energy axis (actually focal plane position). The peak corresponds to decays of the 8 Be (27.48 MeV, $T=2$) level to ${}^{6}Li(36)$. (c) As in (b) but for the 10 Be(p,td)⁶Li(1₀) reaction. (d) Focal plane position singles spectrum for identified tritons. Note the unreso shoulder due to ${}^{16}O(b,t){}^{14}O(7.78 \text{ MeV})$. (e) Spectrum of counts between the indicated dot and dash lines, projected on the deuteron energy axis. (f) As in (e) but for the region including the $T=2$ level. (g) As in (e) but for the indicated background region.

telescope, low energy particles were identified using the 16 μ m ΔE detector and higher energy particles with the 35 μ m ΔE detector.

The $T = 2$ level was populated via the 10 Be(p, t)⁸Be reaction using 46 MeV protons and a 200 μ g/cm² 10 BeO target on a 1 mg/cm² Pt backing. The target was provided by Goosman; the 10 Be isotopic enrichment was $>94\%$.

A portion of a typical triton singles spectrum observed in the QDDD at $\theta_{lab} = 23^{\circ}$ (Runs $^{8}Be-2$ and B^8 Be-3) is shown in Fig. 6. The large peak corresponding to the $T = 2$ state in ⁸Be is evident. Tritons from the reaction $^{16}O(p, t)^{14}O(7.78 \text{ MeV})$ account for the structure at lower triton energy noted in the figure. The distorted shape of the contaminant peak is due to imperfect focusing of the tritons from this reaction for the settings used in the experiment. In run ⁸Be-1 the spectrograph angle was different (θ_{lab} = 27°) and the contaminant was not completely resolved [see Fig. 7(d)]. Since the ^{14}O (7.78 MeV) level decays by proton emission to $^{13}N(g,s)$ the low energy decay protons from $B\text{Be}(T=2)$ were obscured [particularly the decay to ${}^{7}Li(\frac{5}{2}$ ₀, 7.467 MeV)]. In addition the error in determining branching ratios for other decays was increased significantly because of the uncertainty in determining the singles yield. The results in

Table I are obtained by a weighted average of Run 8 Be-1, 8 Be-2, and 8 Be-3 for all decay channels where decay particle identification was unambiguous. Otherwise only the results of 8 Be-2 and 8 Be-3 were included. In Runs ⁸Be-1 and ⁸Be-2 the telescope was at -120° and in Run 8 Be-3 at -90° .

Figure 7(a) shows a two dimensional histogram of events corresponding to coincidences between tritons in the QDDD and decay particles in the deutron identification peak (see Fig. 5 from Run 8 Be-1). The axes of the plot correspond to triton momentum (position in the focal plane of the QDDD} and deuteron energy (as measured by the particle telescope).

The dashed verticle line indicates the minimum deuteron energy which passes through the 51 μ m detector. Counts below this threshold are from particles which stop in this detector and could arise from low energy d, α , ³He, or triton decay particles. In subsequent figures such unidentified decay particles will not be plotted except for a few cases as noted in the figure captions. The enhancements seen in Fig. 7 cluster along kinematic lines determined by energy and momentum conservation in the three body final state. The particularly significant enhancements around channel 35 (triton energy) correspond to decays of ${}^{8}Be(T=2)$ to

TABLE I. Decay properties of the 27.49 MeV, $J^{\pi} = 0^{+}$, $T = 2$ level of 8 Be.

Branching

 $\texttt{I} = 5.5 \pm 2 \text{ keV}$, Ref. 10. b Reference 37.

FIG. 8. Spectra of protons identified in the surfacebarrier detector telescope in coincidence with tritons in three regions of the focal plane as indicated in Fig. $7(e) - (g)$.

 ${}^6Li(1_0^+, g.s.)$ and ${}^6Li(3_0^+, 2.185 \text{ MeV})$. Projections of the two dimensional spectra onto the triton momentum and the deuteron energy axes are shown in Figs. $7(b) - 7(g)$.

Spectra projected onto the triton energy axis were ordinarily used to determine branching ratios (in some cases the other projection was used to help subtract overlapping background). The backgrounds subtracted in each case were determined by linear interpolation between the background levels on either side of the peaks. Branching ratios were calculated from the ratio of counts in the " $T = 2$ " peaks in the coincident and singles triton spectra, with corrections applied for overall detection efficiency, including pileup dead time effects, and for solid angle transformation from laboratory to center-of-mass frames.

FIG. 9. α particle decay spectra projected on the decay particle detector axis as in Fig. 7 (e)- (g) .

FIG. 10. 3 He (a) and t (b) decay spectra corresponding to 8 Be (T = 2), projected on the decay particle detector axis with background spectra subtracted.

Figures $7(e) - 7(g)$ show the spectra of deuteron decay particles projected onto the decay energy axis. The decay properties of the $T = 2$ state are clearly dominated by the deuteron decay to the ground state and first excited state of ⁶Li.

Similar projections on the decay axis for runs ⁸Be-2 and ⁸Be-3 are shown for decay protons in Fig. $8(a) - 7(c)$. Here one can observe significant

FIG. 11. Time-of-flight neutron spectrum corresponding to the ⁸Be ($T=2$) peak after background subtraction (histogram). The filled circles correspond to a normalized random spectrum (see text).

decay for regions on either side of the $T = 2$ peak. The large peak at about 3.2 MeV in Fig. 8(c) arises from proton decay of the $^{14}O(7.78 \text{ MeV})$ level, clearly observable in the triton spectrum of Fig. 6.

Figure $9(a)$ -(c) shows similar projections for α particle decay. Whereas little or no enhancement in the region of the $T = 2$ peak is observed for decay to the ground state of ⁴He ($E_{\alpha} \approx 18.5$) MeV), there is clear evidence for decay to the 20.5 MeV, $J^{\pi} = 0^+$ first excited state $(E_{\alpha} \approx 6 \text{ MeV})$.

Projected decay spectra for 3 He and t decay, corresponding to the $T=2$ peak after subtraction of a kinematically corrected background spectrum, are shown in Fig. 10(a) and 10(b), respectively. Broad continua are observed, expected for the unbound residual nuclei 5 He and 5 Li and branching ratios were evaluated for the total number of counts above threshold.

The time-of-flight neutron spectrum corresponding to the $T = 2$ peak after background subtraction is shown as a histogram in Fig. 11, together with the normalized. random spectrum (filled circles) corresponding to the average of seven beam repetition periods. The large peak at the left corresponds to prompt γ rays from the target and was used to align the random spectra for averaging. The time scale in ns/m was derived from the known beam repetition frequency and the target to detector distance of 0.66 m. The random coincidence rate was sufficiently large (due to the thick Pt backing) that the statistical accuracy of the neutron branching ratios was very poor, but small net excess of counts were observed for decays to the channels indicated.

Final branching ratios were obtained as a weighted average from a11 the runs in which the particular decay particles were unambiguously identified and are listed in Table I. These are the first branching ratio measurements for the 27.49 MeV $T = 2$ state. This state has been observed as a resonance in ${}^6\text{Li}(d, p_1)$ and ${}^6\text{Li}(d, \alpha)$ by Black *et al.*,⁸
in ${}^4\text{He}(\alpha, p)$ and ${}^4\text{He}(\alpha, d)$ by Hinterberger *et al.*²⁵ in ⁴He(α , β) and ⁴He(α , d) by Hinterberger *et al.*²⁵ in ⁴He(α , p) and ⁴He(α , d) by Hinterberger *et al.*²⁵
and in ⁶Li(d , γ) by Noé *et al.*¹⁰ These different experiments have yielded inconsistent values for $\Gamma_{c.m.}:$ (10 ± 3) keV, (14.7 ± 4.0) keV, and (5.5 ± 2.0) keV, respectively. We have adopted the value 5.5 ± 2.0 keV given in Ref. 10. Although the charged particle resonance studies have indicated finite decay widths for the p_1 , d_0 , and α_0 channels, quantitative partial widths were not obtained previously due to the complications introduced by nonzero channel spins. Partial widths in the c.m. system, obtained by combining our branching ratios with $\Gamma_{\rm cm}$ from Ref. 10 are listed in Table I, along with the corresponding reduced widths in single particle units, $\theta_i^2 = \gamma_i^2 / \gamma_{WL}^2 = \Gamma_i / 2P \gamma_{WL}^2$

where Γ_i is the particle decay width, P is the Coulomb penetrability and $\gamma_{WL}^2 = 2\hbar^2/\mu R$ is the Wigner limit width. In all cases in this paper, P and γ_{WL}^2 were calculated using a radius $R = 1.4$ $\times (A_1^{1/3} + A_2^{1/3})$ fm.

The radiative width Γ_{M_1} is obtained by combining the value $\Gamma_{M_1}\Gamma_{d_0}/\Gamma$ = (6.12 ± 0.64) eV from Ref. 10 with our value for Γ_{d_0}/Γ .

B. ${}^{12}C$ (27.6 MeV, $T=2$)

The 27.6 MeV $(0^+, 2)$ level of ¹²C was populated in the ¹⁴C(p, t)¹²C reaction using 46 MeV protons. The 300 μ g/cm^{2 14}C target on a 1 mg/cm² Au

FIG. 12. QDDD focal plane position spectra of tritons from the ¹⁴C(p, t)¹²C reaction at $E_p = 46$ MeV for (a) singles, (b) coincidences with proton decays to 11 B(0.0 MeV), (c) coincidences with proton decays to $^{11}B(2.12 \text{ MeV})$, (d) coincidences with alpha decays to 8 Be (0.0 MeV). The peak corresponds to the 27.6 MeV, $T=2$ level of ^{12}C .

backing was made by cracking acetylene (enriched to $\approx 94\%$ ¹⁴C).

The low cross section for the ${}^{12}C(T=2)$ level $(-10 \mu b/sr)$ and the large background (due in part to reactions with the thick gold backing) made it difficult to obtain good statistical accuracy for branching ratios. The high random coincidence rate precluded any measurement of neutron branching ratios. Figure 12 shows the singles triton spectrum for $\theta_{lab} = 27^\circ$ as well as the corresponding real coincidence triton spectra associated with proton decays to $^{11}B(g.s.)$ and $^{11}B(2.12 \text{ MeV})$, and α decays to ⁸Be(g.s.). These data were obtained with the two-element telescope described above at -120° to the beam (Run ¹²C-1). Additional data were taken with the three element telescope at -90° which allowed us to measure the branching ratio for deuteron decays to $^{10}B(g.s.)$. The results for these two runs are combined in Table II.

Figure 13 shows the background subtracted real coincidence α particle spectra for Runs ¹²C-1 and 12 C-2. It was constructed by subtracting projected spectra above and below the $T = 2$ peak from the spectrum on resonance. A peak is observable corresponding. to decays to. the ground state of ⁸Be, but for higher excitation energies a continuum is observed.

The measured branching ratios are listed in Table II, including an entry labeled "continuum, " associated with the net counts below channel 40 in Fig. 13. The latter branching ratio was calculated as though the decay proceeded sequentially through excited states of 8 Be and only the α particle from the primary decay was detected. Since a significant proportion of this yield may arise from $^{12}C(T=2)$ - 3 α directly, with the possibility of detecting any of the three alphas, the branching ratio in this channel may be smaller by up to a factor of 3. The present results are in reasonable agree-

FIG. 13. Surface-barrier detector telescope energy spectrum for alpha particles in coincidence with tritons from ${}^{14}C(p,t) {}^{12}C(27.6 \text{ MeV}, T=2)$. An average background spectrum has been subtracted.

ment with and more precise than the previous measurement by Ashery *et al.*²⁶ surement by Ashery et $al.^{26}$

The 27.6 MeV, $T = 2$ level of ¹²C has been sought extensively as an isospin-forbidden resonance in the ${}^{11}B(p,\gamma\gamma)$, ${}^{10}B(d,p)$, ${}^{10}B(d,\alpha)$, ${}^{10}B(d,\gamma\gamma)$, and the ${}^{11}B(p,\gamma\gamma)$, ${}^{10}B(d,p)$, ${}^{10}B(d,\alpha)$, ${}^{10}B(d,\gamma\gamma)$, and ${}^{9}Be({}^{3}\text{He},\gamma\gamma)$ reactions.²⁷⁻³⁰ A tentative identifica tion of the level as a resonance in the $^{9}Be(^{3}He, \gamma\gamma)$ $reaction²⁸$ was not verified in two other searches tion of the level as a resonance in the Bey He
reaction²⁸ was not verified in two other search
in the same channel.^{29,30} Searches in the other channels found no resonance. The present branching ratio measurements indicate why these measurements have not been successful. The $T = 2$ level decays primarily via α -particles or protons to excited states, with very small ground state branches for protons or deuterons. In our experiment ³He particles from a possible decay branch to 9 Be are too low in energy (~1 MeV) to have been

Branching ratio $(\%)$					
Channel	Present work	Ref. 26	Γ_i^{a} (eV)	$\gamma^2/\gamma_{\rm w}^2$ (×10 ⁻⁶) ^a	
$\alpha + \binom{8}{10}(0^+_0)$	10.5 ± 3.0	10	$<$ 4050	< 285	
+ "Continuum"	9.1 ± 3.5		3750	\bullet \bullet \circ	
$p + {^{11}B(\frac{3}{20})}$	3.0 ± 2.2		< 1650	91	
+ $^{11}B(\frac{1}{20})$	8.0 ± 2.3	Σ = 30 \pm 10	3090	≤ 196	
$+$ ¹¹ B($\frac{5}{2}$ ⁰)	$\mathbf{0}$ ± 3.3		< 990	316	
$+$ ¹¹ B($\frac{3}{2}$)	8.4 ± 3.2		3450	290	
+ ${}^{11}B(\frac{7}{20}+\frac{1}{20})$	8 ± 5		$<$ 3900	000	
$d + {^{10}B(3_0^{\dagger})}$	2.8 ± 2.0		1440	.<973	

TABLE II. Decay properties of the 27.6 MeV, $J^{\pi} = 0^{+}$, $T = 2$ level of ¹²C.

^a Upper limit based on Γ < 30 keV (Ref. 31).

identif ied.

A recent study of the $T = 2$ state via the ¹⁴C(p, t) reaction³¹ has yielded an upper limit of 30 keV for the width of the level. This result was used to define upper limits for reduced widths listed in Table H.

C. ${}^{16}O(22.72 \text{ MeV}, T= 2)$

The lowest $T = 2$ level in ¹⁶O was populated by the $^{18}O(p, t)^{16}O$ reaction with 42.4 MeV protons incident upon a target consisting of 200 μ g/cm² of WO₃ (enriched to >90% in ¹⁸O) on a 10 μ g/cm² carbon backing. The spectrum of tritons observed in the focal plane of the QDDD spectrometer at θ_{lab} $=+23^{\circ}$ is shown in Fig. 14(a). The $T=2$ level is clearly observed with an energy resolution of 25 keV FWHM. Decay particles were observed at laboratory angles of -90° and -120° in a threeelement detector telescope consisting of 16-, 51-, and $1000 - \mu$ m surface-barrier detectors. Figures 14(b) and 14(c) show the spectra of α particles and protons observed in coincidence with tritons

FIG. 14. (a) ^A QDDD focal plane position spectrum for tritons from the $^{18}O(p,t)^{16}O$ reaction for E_p = 42.4 MeV. (b) Surface-barrier detector telescope energy spectrum of α particles in coincidence with tritons in the $T = 2$ peak of (a). The crosses indicate events wherein the particles stopped in the front detector and could not be unambiguously identified as α particles. (c) As in (b) but for proton decays.

associated with the production of the $T = 2$ level. Clear indication of decay strength is seen for both particle types. Events in the α particle spectrum below about 4 MeV (plotted as crosses) correspond to particles which stop in the $16-\mu m$ front detector and include some carbon, nitrogen or oxygen ions from light particle decays. Figure 15 shows timeof-flight spectra for real (solid line) and random (dots) coincidence between tritons in the peak associated with the $T = 2$ level and neutrons or γ rays detected in the plastic scintillator. The real spectrum is the sum of two rf periods and the random spectrum is the normalized sum of eight rf periods. Statistically significant peaks are observed in the expected regions for neutrons decaying to the first two excited states of ¹⁵O and for γ rays (probably from gamma decay of states populated by particle decays).

Branching ratios from the present work are compared with those from previous studies in Table III. There is excellent agreement between the III. There is excellent agreement between the present work and the results of Nero $et \ al. ,¹¹$ but some discrepancies exist in comparison with the results of Koh et al ,⁷ particularly for the proton branching ratios. This may be due to the much poorer energy resolution in Ref. 7 (400 keV FWHM). Recently this $T = 2$ level was sought as a resonance in high resolution studies of the ${}^{15}N(p, p)$ resonance in high resolution studies of the $^{15}N(p, p)$
reaction and not observed,³² in agreement with the small branching ratio measured in the present work.

A total width of 12.5 ± 3.5 keV was obtained for A total width of 12.5 ± 3.5 keV was obtained this level by Nero *et al.*¹¹ and confirmed by McDonald et al.³² using the ¹²C(α , α ₂) reaction.

FIG. 15. Time-of-flight neutron spectrum corresponding to the ¹⁶O(T=2) peak after background subtraction (histogram). The filled circles correspond to a normalized random spectrum (see text).

	Branching ratio $(\%)$				
Channel	Present work	Ref. a	Ref. b	Γ_i (eV) ^c	$\gamma^2/\gamma_{\rm WL}^2$ (×10 ⁻⁶)
$\alpha + {}^{12}C(0_0^+)$	1.6 ± 0.7	2.0 ± 0.6	1 ± 4	190 ± 100	$19 + 10$
$+{}^{12}C(2_0^+)$	$1.9 = 0.7$		0 ± 3	230 ± 110	30 ± 14
+ ${}^{12}C(0_1^+)$	14 ± 2	13 ± 5		1680 ± 550	260 ± 85
$p + {^{15}{\rm N}(\frac{1}{20})}$	7 ± 2	≥ 6	25 ± 6	840 ± 343	55 ± 22
$+\frac{15}{N(\frac{5}{20}^+ + \frac{1}{20}^+)}$	11 \pm 2		21 ± 5	1320 ± 454	\cdots
+ $^{15}N(\frac{3}{20})$	$5 -$ \pm 2		14 ± 5	600 ± 300	$87 + 44$
$n + {^{15}O(\frac{1}{20})}$	≤ 15			< 2325	<171
+ ${}^{15}O(\frac{1}{20}^+$ + $\frac{5}{20}^+)$	23 ±15			2760 ± 1970	0 ₀

TABLE III. Decay properties of the 22.72 MeV, $J^{\pi} = 0^{+}$, $T = 2$ level of ¹⁶O.

^a Reference 11.

'

^b Reference 7.

 $C_T = 12 \pm 3.5$ keV (Refs. 11 and 32).

D. 24 Mg (15.43 MeV, $T=2$)

Figure 16 shows the spectrum of tritons at θ_{lab} =23° from the ²⁶Mg(p, t)²⁴Mg reaction for 42.4 MeV protons incident on a 200 μ g/cm^{2 26}Mg (enriched to >95%) layer on a 10 μ gm/cm² carbon backing. Also shown are the spectra of protons and α particles corresponding to particle decays of the $T = 2$ state to energy levels of ²³Na and ²⁰Ne. The decay particles were detected in a $51-\mu m-$ 1000- μ m particle detector telescope at -120° . The charged particle decay branching ratios are sum-

FIG. 16. Singles triton spectrum and coincidences spectra for proton and alpha decays of the 24 Mg (15.43 MeV, $T = 2$) level (see caption of Fig. 14).

marized in Table IV. The branching ratios for α decay are in reasonable agreement with the pre-'vious results of McGrath $et al.,⁸$ but there is some discrepancy for the protonbranches. Partial widths and reduced widths for the decays have been deduced by combining our results with the recently measured $\Gamma = 500 \pm 200$ eV obtained by the ²³Na(p, p) (Ref. 19) and ²³Na(p, γ) (Ref. 20) reactions. Previous measurements by the ²³Na(p, γ) actions. Previous measurements by the $^{23}Na(p, r)$
reaction^{14,33} have defined the gamma decay properties of the level.

E. 28 Si (15.22 MeV, $T=2$)

The ²⁸Si (15.22 MeV, $J^{\pi} = 0^+$, $T = 2$) level was produced with 42 MeV protons incident on a 200 μ g/cm² SiO₂, target enriched to >90% in ³⁰Si. Figure 17 shows the singles triton spectrum observed in the QDDD spectrometer at $\theta_{lab} = 27^\circ$ as well as spectra of protons and α particles corresponding to the various decay modes of the $T=2$ level. The decay particles were detected in a 50 and 1000- μ m particle detector telescope at -120°. The resultant branching ratios shown in Table IV are in excellent agreement with those of McGrath et al.,⁸ but have better precision. The $T = 2$ level in $28Si$ was previously observed as a resonance in in ²⁸Si was previously observed as a resonance if the ²⁴Mg(α , γ) reaction,^{27,15} and an upper limit of 2 keV was obtained for the total width. The observation of a resonance in the gamma yield from ²⁴Mg(2⁺) was used in the latter work¹⁵ to infer Γ_{α} , $=35 \pm 20$ eV, which may be combined with the present branching ratio measurement to yield Γ_{α_0} $=301 \pm 190$ eV, in agreement with a recent precise measurement $\Gamma_{\alpha_0} = 222 \pm 25$ eV, by the ²⁴Mg(α, α)
reaction.²¹ reaction.

Channel	Branching ratio $(\%)$ Present work	Ref. 8	$\Gamma_i(\text{eV})$	$\gamma^2/\gamma_{\rm WL}^2$ (×10 ⁻⁶)
24 Mg(1543 MeV)	$(\Gamma = 435 \pm 135 \text{ eV}^a)$			
$\alpha + {}^{20}\text{Ne}(0_0^+)$	± 1.3 6	3 ± 3	30 ± 14	9.4 ± 4.4
$^{20}Ne(2_1^+)$	± 2 18	$22 + 4$	90 ± 37	92 ± 38
$p + {}^{23}\text{Na}(\frac{3}{20}^+)$	49 ± 6	$71 + 7$	245 ± 103	135 ± 57
$^{23}Na(\frac{5}{20}^+)$	± 2 7	1 ± 4	35 ± 17	$27 + 13$
$^{23}Na(\frac{7}{20}^+)$	± 1 1		5± - 5	15000 ± 15000
$^{23}Na(\frac{1}{20}^+)$	4.6 ± 1		23 ± 14	31 ± 19
$^{23}Na(\frac{1}{20}+\frac{9}{20}^+)$	$7 -$ ± 1		35 ± 15	
28 Si(15.22 MeV)	$(\Gamma_{\alpha 0} = 222 \pm 25 \text{ eV}^{\text{b}})$			
$\alpha + {}^{24}Mg(0_0^+)$	69 ± 6	72 ± 8	222 ± 25	135 ± 15
$+{}^{24}Mg(2_0^+)$	8 ± 2	9 ± 4	$26 \pm$ 8	135 ± 41
$p + {}^{27}Al(\frac{5}{20}^+)$	3.5 ± 2	4 ± 8	11± 6	7.9 ± 4.3
+ $^{27}Al(\frac{1}{2}^+$ + $\frac{3}{2}^+$	2	4 ± 5	$<$ 7	$\cdot \cdot \cdot$
$+{}^{27}$ Al(2.2 to 2.9 MeV)	< 16		59	$\cdot \cdot \cdot$
32 S(12.05 MeV)	$(\Gamma < 170 \text{ eV})$ ^c			
$\alpha + {}^{28}\text{Si}(\,0_0^+)$	4 ±4	$18 + 7$	< 14	< 16
$+$ ²⁸ Si(2 ⁺)	2	11 ± 5	< 3.4	< 283
$p + {}^{31}P(\frac{1}{20}^+)$	±13 100	$86 + 17$	< 192	<60
+ ${}^{31}P(\frac{3}{20}^+)$	$\overline{4}$ ±4		< 14	< 194
44 Ti(9.34 MeV)	$(\Gamma = 1.10 \pm 0.28 \text{ eV})^d$			
$\alpha + {}^{40}\text{Ca}(0_0^+)$	32 ± 5		0.35 ± 0.07	$57 + 11$
$p + {}^{43}\text{Sc}(\frac{7}{20})$	$<\!\!4$	000	< 0.055	\bullet \circ \circ
γ + 44 $\text{Ti}(1^+)$	54 ±11		0.75 ± 0.19	

TABLE IV. Particle decay properties of lowest $T=2$ levels $(J^{\pi}=0^{+})$.

 a References 19 and 20.

 b Reference 21.</sup>

^c Reference 17 for $\Gamma_b/\Gamma = 100\%$.

 d Reference 23.

F. $32S$ (12.05 MeV, $T=2$)

The spectrum of tritons at 27° from the ${}^{34}S(\rho, t){}^{32}S$ reaction at $E_b = 42$ MeV is shown in Fig. 18(a). The target consisted of 200 μ g/cm² of ³⁴S, isotopically enriched to greater than 95%, deposited on a 20 μ g/cm² carbon foil and covered with a second 20 μ g/cm² carbon foil. The only significant decay branch observed was proton decay to the ground state of ${}^{31}P$, as may be seen in the coincident spectra in Figs. $18(b)$ and $18(c)$ and the resultant branching ratios listed in Table IV. The present results are in qualitative agreement with those of McGrath et al , but the alpha branching ratios are somewhat smaller. This $T = 2$ state has been

studied¹⁷ as a resonance in the ³¹P(p, γ) reaction and the ${}^{31}P(p, p)$ reaction has been used to set an upper limit of 170 eV for its total width (with the assumption of 100% p_0 branching ratio).

G. ${}^{36}Ar$ (10.86 MeV, $T = 2$)

The branching ratio for particle decay of the lowest $T = 2$ state in ³⁶Ar had not been previously measured. The state was populated by the ${}^{38}\text{Ar}(p, t)$ ³⁶Ar reaction with 42 MeV protons incident on 38 Ar gas (isotopically enriched to >95%). The gas cell was a 1 cm diameter cylinder 7 cm long. Tritons passed through a 2 mm diameter exit window covered by a 500μ g/cm² Ni foil at 23° to the axis of the gas cell and in line with the en-

FIG. 17. Singles triton spectrum and coincidence spectra for proton and alpha decays of the 28 Si (15.22 MeV, $T = 2$) level (see caption of Fig. 14).

trance aperture of the QDDD. Decay particles passed through a 1 cm diameter 625 μ g/cm² Ni window at -90% viewed by a two-element telescope consisting of a 35- and a 500- μ m surface barrier detector. The beam entered through a 2.5 μ m Al window covered by 7μ g/cm² of Formvar and exited through a 4 μ m Havar foil. The pressure in the cell was maintained at 50 Torr during the experiment. The overall detection efficiency

FIG. 18. Singles triton spectrum and coincidence spectra for proton and alpha decays of the ^{32}S (12.05 MeV, $T = 2$) level. (See caption of Fig. 14).

FIG. 19. QDDD focal plane position spectra for identified tritons from the ${}^{38}\text{Ar}(p,t){}^{36}\text{Ar}$ reaction (c) in singles mode, (b) in coincidence with protons corresponding to decays to the ground state of 35 Cl, and (a) in coincidence with α particles corresponding to decays to the ground state of 32S.

was measured as described in Sec. IIB, using the ¹⁴N(p , d)¹³N reaction on nitrogen gas.

The $T = 2$ state in ³⁶Ar was observed to decay only by proton emission to ${}^{35}Cl$ (g.s.) as indicated in Fig. 19 and Table V. The γ -decay properties of this level have been studied^{34,35} by the ³⁵Cl(p, γ) reaction but the total width has not been determined. From the experimental resolution in the QDDD the width must be <35 keV.

TABLE V. Branching ratios (%) of lowest $T = 2$ levels $(J^{\pi} = 0^+).$

Channel	Present work	Ref. 8
	$36Ar(10.86 \text{ MeV}, T=2)$	
$\alpha + {}^{32}S(0_0^+)$	<6	000
$p + {}^{35}Cl(\frac{3}{20}^+)$	109 ± 15	\cdots
$+{}^{35}Cl(\frac{1}{20}^+)$.	<4	
	40 Ca(11.98, T = 2)	
$\alpha + {}^{36}Ar(0_0^+)$	93 ± 9	$116 + 11$
$+{}^{36}Ar(2_0^+)$	$<$ 3	-1 ± 2
$p+{}^{39}K({3\over 20})$	< 5	$-3 = 9$

H. 40 Ca (11.98, $T = 2$)

The spectrum of tritons observed at 27° from the ⁴²Ca(p, t)⁴⁰Ca reaction at $E_p = 42$ MeV is shown in Fig. 20. The target consisted of \sim 200 μ g/cm² of ^{42}Ca (isotopically enriched to >95%) evaporated on a 20 μ g/cm² carbon foil and transferred to the target chamber in a vacuum lock. The only significant decay branch observed was alpha decay to the ground state of ${}^{36}\text{Ar}$, as may be seen from the coincidence spectra in Fig. 20 and the resultant branching ratios listed in Table V. The present results are in agreement with those of McGrath et al.⁸ This $T = 2$ state has never been observed as an isospin forbidden resonance. The observed width in the present measurement may be used to define Γ < 20 keV.

I. 44 Ti(9.34 MeV, $T = 2$)

Details of the measurement of branching ratios for alpha and gamma decay of the lowest $T = 2$ level of 44Ti have been given in a previous publicalevel of ⁴⁴Ti have been given in a previous put
tion.²³ These results may be combined with a measurement³⁶ of $\Gamma_{\alpha} \Gamma_{\gamma}/\Gamma = 0.24 \pm 0.05$ eV to determine the total and partial widths for this level as listed in Table IV.

FIG. 20. QDDD focal plane position spectra for identified tritons from the ${}^{42}Ca(p,t)^{10}Ca$ reaction (c) in singles mode, (a) in coincidence with protons corresponding to decays to the ground state of $39K$, and (b) in coincidence with α particles corresponding to decays to the ground state of ${}^{36}\mathrm{Ar}$.

IV. DISCUSSION

Very few calculations of isospin forbidden particle decay widths of $T = 2$ levels have been performed. However, for ⁸Be, where the spectrum and structure of possible admixed levels is particularly simple, Barker and Kumar 37 have been remarkably successful in predicting partial and total decay widths, as may be seen in Table I. They find that the isospin mixing due to the Coulomb interaction is dominated by the third $T = 0$ state which is predominantly the LS coupled $[22]$ ¹¹S configuration which has a space wave function similar to the $T = 2$ state. The orbital symmetry of this state favors decay by deuteron emission to the ground and first excited states of ${}^{6}Li$ and inhibits proton, neutron, and alpha decays to low-lying states. The close agreement in ⁸Be between the measured decay widths and those calculated from Coulomb mixing alone does not support the conjecture of a significant $\Delta T = 1$ charge dependence of the short range nuclear force which has been invoked³⁸ to explain the Nolen-Schiffer anom
alv.³⁹ The close agreement for the absolute magaly.³⁹ The close agreement for the absolute magnitudes of the widths may. be fortuitous, however, since the isospin admixture depends on the relative spacing of the 0^+ , $T = 0$ and 0^+ , $T = 2$ states, which has not been determined experimentally. In fact, there is evidence⁴⁰ from the isobaric multiplet mass equation that the admixing $T = 0$ state may lie below the $T = 2$ state, rather than above, as in Barker 's calculation.

The influence of structural symmetry is particularly striking for the α_0 decay, which is sensitive only to the [4] configuration. The decay is highly inhibited with $\theta_{\alpha}^2 = (0.5 \pm 1.0) \times 10^{-6}$. On the other hand, the orbital symmetry does not inhibit decays to excited states of ⁴He, and the most impressive feature of the 8 Be($T = 2$) decays is the relative strength of the branch to the 20.2 MeV 0^+ , $T = 0$ state of ⁴He. This branch has $\theta_{\alpha}^2 = (54 \pm 9) \times 10^{-6}$, at least a factor of 30 greater than the α_0 branch. Barker³⁶ did not consider decays to the 20.2 MeV state, since his wave functions contained only an inert Os shell core. The strength of the α , branch presumably arises because the 4He excited state has a [22] configuration to which decays are favored.

The α decays of ¹⁶O(T=2) show an interesting similarity to those of ⁸Be ($T = 2$). The α_0 decays are again weak, $\theta_{\alpha}^2 = (19 \pm 10) \times 10^{-6}$, while decays to the excited 0^+ state of 12 C are much stronger with θ_{α}^2 = (260 ± 85) × 10⁻⁶. This presumably again indicates that the Coulomb interaction readily mixes T_{\leq} states with $(2p - 2h)$ space configurations, which are similar to those of the $T = 2$ state. This trend may persist in heavier nuclei, since $20Ne$

has a strong decay branch to the unresolved $3⁻$ and 0^+ states in ¹⁶O at $E_r \sim 6$ MeV.

The available information on partial decay widths for $T = 2$ levels is now sufficiently complete that it is appropriate to look for systematic trends with 'mass number such as were observed for $T = \frac{3}{2}$ it is appropriate to look for systematic trends
mass number such as were observed for $T = \frac{3}{2}$
levels.^{2,3,4} Figures 21 and 22 show the reduce widths, θ^2 , for proton and alpha decays of the $T = 2$ levels, plotted as a function of mass. Although the present results are generally not as accurate as those for the $T = \frac{3}{2}$ levels (Fig. 23) a number of trends are worth noting.

First, the general magnitudes of the reduced widths, and also their slope as a function of A are quite similar for the decays of the $T = \frac{3}{2}$ and $T = 2$ levels. This is further evidence for the regularity of the T -forbidden decays and suggests that we can learn general properties of isospin mixing from such decays.

Second, reduced widths are similar for alpha $(T=0)$ and proton $(T=0 \text{ or } 1)$ decay channels indicating that isotensor $(\Delta T = 2)$ and isovector $(\Delta T = 1)$ mixing matrix elements are of similar magnitude. This result has recently been discussed 41 in terms of the isospin mixing of $T = 0$ and $T = 1$ antianalog configurations due to the Coulomb interaction. These mixing matrix elements involve differences of single proton-core and two-proton Coulomb matrix elements which can be derived from Coulomb energy differ ences, using a simple schematic model for the dominant configurations. Although single proton-core (isovector) Coulomb matrix elements are much larger than two-proton Coulomb matrix elements, the mixing of $T = 0$ and $T = 1$ anti-analog configurations invariably involves differences of these matrix elements for protons in different valence orbits. These differences are similar in magnitude producing significant isovector and iso-

FIG. 21. Reduced widths $\theta^2 = \gamma^2 / \gamma_{WL}^2$ for isospin-for bidden proton (alpha) decays of $A = 4N$, $T = 2$ levels to the ground state of $A=4N-1$ ($A=4N-4$) nuclei.

FIG. 22. Reduced widths $\theta^2 = \gamma^2/\gamma_{WL}^2$ for isospin-for bidden proton (alpha) decays of $A = 4N$, $T = 2$ levels to the first excited state of $A = 4N - 1$ ($A = 4N - 4$) nuclei.

tensor mixing. The similarity of isovector and isotensor matrix elements also results in significant differences in analogous neutron and proton decay widths for $T = \frac{3}{2}$, $T_g = \pm \frac{1}{2}$ nuclei, as is observed experimentally³ (see Fig. 23).

More detailed calculations of isospin forbidden decay widths must determine the distribution of $T₁$ levels with appropriate J^{π} and their allowed decay probability for the various channels. However, the relatively regular variation with mass of the isospin forbidden reduced widths of the $T = 2$ and $T = \frac{3}{2}$ levels and their similarity in magnitude, lends credence to the assumption that the mixing may be dominated by some general configuration, such as the antianalog states. For the alpha decays, the reduced widths to excited states are as large as or larger than those to the ground state.

FIG. 23. Reduced widths $\theta^2 = \gamma^2/\gamma_{WL}^2$ for isospin-forbidden proton (dots) and neutron (crosses) decay of the lowest $T = \frac{3}{2}$ levels of $A = 4N+1$ nuclei to the ground state and lowest 2^+ level of $A = 4N$ nuclei

This is compatible with mixing of the antianalog configurations, having $2p - 2h$ structure similar to the $T = 2$ level.

The $\Delta A = 8$ periodicity observed for the proton reduced widths of $T = \frac{3}{2}$, $T_g = -\frac{1}{2}$ levels⁴ may be observable in the reduced widths for p_0 decay of the T = 2 levels, but is not evident in the α_0 or α_1 widths. This is similar to the $T = \frac{3}{2}$ case where the oscillation is not observed in the reduced widths for proton decay to the lowest 2⁺ levels or in neutron decays from $T = \frac{3}{2}$, $T_z = +\frac{1}{2}$ levels to ground states or excited states. This suggests that the oscillation is not representative of an overall variation in isospin impurity amplitudes, but is dependent both on isospin mixing and decay probabilities to the particular channels involved.

Further improvements in the accuracy of isospin forbidden particle decay widths of $T = 2$ levels must come primarily from resonance studies. It is interesting to predict total widths for the presently unobserved cases assuming reduced widths similar to those in Figs. 21 and 22.

For ¹²C if θ^2 (p_0 or α_0) \approx 10 \times 10⁻⁶, then $\Gamma \approx$ 6 or keV, respectively. However, a total width of 30 keV would be consistent with the neighboring re-

- *Present address: Department of Physics, Stanford University, Stanford, California 94305.
- (present address: LAMPF, Los Alamos Scientific Lab., Los Alamos, N. M. 87545.
- fPresent address: Lawrence Berkeley Laboratory, University of California, Berkeley, California 94720.
- 5present address: Department of Physics, Michigan State University, East Lansing, Michigan 48824.
- ¹E. G. Adelberger, A. B. McDonald, C. L. Cooke, C. N. Davids, A. P. Shukla, H. B. Mak, and D. Ashery, Phys. Rev. C7, 889 (1973).
- ${}^{2}E$. G. Adelberger, in Proceedings of the International Conference on Nuclear Structure and Spectroscopy, Amsterdam, edited by H. P. Blok and A. E. L. Dieperink (Scholar's Press, Amsterdam, 1974), p. 641.
- 3A. B. McDonald, H. B. Mak, H. C. Evans, G. T. Ewan, and H. B. Trautvetter, Nucl. Phys. A273, 477 (1976) and references therein.
- ${}^{4}P$. G. Ikossi, W. J. Thompson, T. B. Clegg, W. W. Jacobs, and E.J. Ludwig, Nucl. Phys. A274, ¹ (1976).
- 5 A. B. McDonald, T. K. Alexander, and O. Hausser, Nucl. Phys. A37, 545 (1969).
- 6 H. Weigmann, R. L. Macklin, and J. A. Harvey, Phys. Rev. C14, 1328 (1976).
- ⁷T. K. Koh, W. R. Falk, N. E. Davison, and J. M. Nelson, Phys. Rev. C7, 50 (1973).
- 8R. L. McGrath, J. Cerny, J. C. Hardy, G. Goth, and
- A. Arima, Phys. Rev. C1, 184 (1971).
⁹J. L. Black, W. J. Caelli, D. L. Livesey, and R. B.
- Watson, Phys. Lett. $30B$, 100 (1969).
¹⁰J. W. Noé, D. F. Geesaman, P. Paul, and M. Suffert, Phys. Lett. 65B, 125 (1976).
- $¹¹A$. V. Nero, R . E. Pixley, and E. G. Adelberger, Phys.</sup>

duced widths. For $^{36}\mathrm{Ar},~\theta^2$ $(\not\! p_0) \stackrel{\scriptstyle <}{\scriptstyle \sim} 50 \times 10^{-6}$ implie duced widths. For \mathbf{H} , θ ($p_0 \approx 30 \times 10^{10}$ mm
 $\Gamma \le 6$ eV and for ⁴⁰Ca, $\theta^2(\alpha_0) \le 50 \times 10^{-6}$ implie $T \leq 7$ eV. These widths are small enough that gamma branching ratios may be measurable. (For ^{36}Ar , Γ_{μ_1} = 0.08 ± 0.02 eV).

Although such experiments to determine the widths of $T = 2$ states with $A \le 36$ are difficult, they are definitely worth pursuing in order to see if the completely unexplained $\Delta A = 8$ oscillations seen in p_{0} decay widths of $T=\frac{3}{2}$ states are present in decays of the $T = 2$ states as well.

In summary, the information now available on isospin-forbidden particle decay widths of $T=\frac{3}{2}$ and $T = 2$ levels in light nuclei indicates relatively regular trends as a function of mass. This suggests the dominance of a general isospin-mixing mechanism. Relatively large isotensor mixing is observed, which might be thought to arise from a charge dependent nuclear inter action. However, a simple model indicates 41 that such isotensor matrix elements can arise from the Coulomb interaction so that the identification of specifically nuclear charge-dependent effects in particle decay widths will be very difficult, and will require detailed nuclear structure calculations for each case.

Rev. C6, 679 (1972).

- ^{12}R , Block, R. G. Pixley, and P. Truol, Phys. Lett. 25B, 215 (1967).
- 13 H. M. Kuan et al., Phys. Lett. 25B, 217 (1967).
- ¹⁴F. Reiss, W. J. O'Connell, D. W. Heikkinen, H. M.
- Kuan, and S. S. Hanna, Phys. Rev. Lett. 19, 307 (1967). ¹⁵N. A. Jelley, N. Anyas-Weiss, M. R. Wormald, B. T.
- Underwood, and K. W. Allen, Phys. Lett. 40B, 200 (1972). 16 D. P. Balamuth and E. G. Adelberger, Phys. Rev. C16,
- 928 (1977).
- 17 J. Vernotte, S. Gales, M. Langevin, and J. M. Maison, Phys. Rev. C8, 178 (1973).
- $18W$. R. Dixon, R. S. Storey, J. J. Simpson, and R. D. Lawson, Phys. Rev. C13, 1745 (1976).
- 19 A. B. McDonald, E. D. Earle, W. McLatchie, H. B. Mak, D. Martin, and P. G Ikossi, Nucl. Phys. A305, 151 (1978).
- 20 J. L. Osborne, E. G. Adelberger, and K. A. Snover, Nucl. Phys. A305, 144 (1978).
- $21P.$ G. Ikossi, K. A. Snover, J. L. Osborne, E. G. Adelberger, and A. B. McDonald, Bull. Am. Phys. Soc. 23, 521 (1978); and (to be published).
- $22\overline{J}$. B. Marion and F. C. Young, Nuclear Reaction Analysis (North-Holland, Amsterdam, 1968).
- ²³S. J. Freedman, C. A. Gagliardi, M. A. Oothoudt,
- A. V. Nero, R. G. H. Robertson, F.J. Zutavern,
- E. G. Adelberger, and A. B. McDonald, Phys. Rev. C17, 2071 (1978).
- $24\overline{V}$. V. Verbinskii, W. R. Burrus, T. A. Love, W. Zobel, and N. W. Hill, Nucl. Instrum. Methods 65, 8 (1968).
- 25_{F.} Hinterberger, P. V. Rossen, E. Lebsanft, P. Lutticken, J. Bisping, R. Jahn, and B. Schuller, Nucl. Phys. A259, 385 (1976).
-
- ²⁶D. Ashery, M. S. Zisman, G. W. Goth, G. J. Wosniak, R. B. Weisenmiller, and J. Cerny, Phys. Rev. C13, 1345 (1976),
- $27K$. A. Snover, Ph.D. thesis, Stanford University, 1969 (unpublished).
- 28 J. L. Black, W. J. Caelli, and R. B. Watson, Phys. Rev. Lett. 25, 877 (1970).
- ²⁹E. K. Warburton, H. M. Kuan, D. E. Alberger,
- P. Paul, and K. A. Snover, Phys. Rev. C6, 375 (1972). 30S. S. Hanna, M. Hass, Z. Shkedi, and Y. S. Horowitz, Nucl. Phys. A195, 91 (1972).
- ³¹R. G. H. Robertson, T. L. Khoo, G. M. Crawley, A. B. McDonald, E. G. Adelberger, and S.J. Freedman, Phys. Rev. C17, 1535 (1978).
- 32A. B. McDonald, P. G. Ikossi, H. B. Mak, D. Martin, and E. D. Earle, unpublished.
- 33 J. Szücs, B. Y. Underwood, T. K. Alexander, and N. Anyas-Weiss, Nucl. Phys. A212, 293 (1973).
- ³⁴D. J. Martin, H. C. Evans, and J. A. Szücs, Phys. Rev. C14, 1320 (1976).
- 35A. Huck, G. J. Costa, G. Walter, M. M. Aleonard,
- J. Dalmas, P. Hubert, F. Lecuo, P. Mennrath, J. Vernotte, M. Langevin, and J. M. Maison, Phys. Rev. C13, 1786 (1976).
- $36J. L.$ Simpson, W. R. Dixon, and R. S. Storey, Phys. Rev. Lett. 29, 1472 (1972); W. R. Dixon, R. S. Storey, J.J. Simpson, and R. D. Lawson, Phys. Rev. C13, 1945 (1976).
- 37 F. C. Barker and N. Kumar, Phys. Lett. $30B$, 103 (1969).
- 38 J. W. Negele, in Proceedings of the International Conference on Nuclear Structure and Spectroscopy, Amsterdam, edited by H. P. Blok and A. E. L. Dieperink
- (Scholar's Press, Amsterdam, 1974). ³⁹J. A. Nolan Jr. and J. P. Schiffer, Annu. Rev. Nucl.
- Sci. 19, 471 (1969).
- ${}^{40}R$, G. H. Robertson, W. Benenson, E. Kashy, and D. Mueller, Phys. Rev. C13, 1018 (1976); R. G. H. Robertson, E. Kashy, W. Benenson, and A. Ledebuhr, ibid. 17, 4 (1978).
- 41A. B. McDonald and E. G. Adelberger, Phys. Rev. Lett. 40, 1692 (1978).