Mass and β decay of the new neutron-rich isotope ω_{Mn}

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The new isotope ⁶⁰Mn was produced by bombarding ⁴⁸Ca with ¹⁸O ions at $E_{18₀}$ = 50-70 MeV. The β -decay scheme of ⁶⁰Mn was determined from γ -ray singles and γ - γ coincidence experiments using large Ge(Li) detectors. The half life of ⁶⁰Mn was found to be 1.79 \pm 0.10 s. A new level with excitation energy 2792 keV was observed in the daughter, 60 Fe. The ground-state spin and parity of 60 Mn was determined to be 3^+ and its total decay energy Q_β was measured to be 8.51 \pm 0.10 MeV. This corresponds to a mass excess of $-52.89+0.10$ MeV, which is compared with various predictions. The structure of ⁶⁰Fe as determined from the present work and from previous experiments is compared with a shell-model calculation.

RADIOACTIVITY 60 Mn: measured $T_{1/2}$, E_{γ} , I_{γ} , E_{β} , γ - γ coin, β - γ coin; de-duced decay scheme, logft, mass excess, g.s.J^T. 60 Fe: deduced levels. Enriched targets, Ge(Li) detectors, plastic scintillator, multiple rabbit.

I. INTRODUCTION

A program to study the β -decay properties of nuclei far from stability in the $1f-2p$ shell has been underway at Argonne National Laboratory for several years. These studies provide information that is useful to both nuclear physics and astrophysics. Spectroscopic information is obtained on the daughter nuclides, and it is usually possible to determine, or at least set limits on, the J^{π} values of the parent ground states. The measured mass excesses and $\log ft$ values are useful for testing mass predictions and for establishing β -decay systematics, both of which are needed in nucleosynthesis calculations. Studies of the $T_z = \frac{9}{2}$ nuclei ⁵¹Sc (Ref. 1), ⁵³Ti (Ref. 2), ⁵⁷Cr (Ref. 3), and ⁵⁹Mn (Ref. 4) have been previously reported. The present study deals with the β -decay of the new neutron-rich isotope 60 Mn. A preliminary report on the decay of 60 Mn has appeared in Ref. 5.

Until recently, little was known about the structure of the 60 Mn decay daughter, 60 Fe. The results of recent $^{48}Ca(^{18}O, \alpha^2n\gamma)^{60}$ Fe (Refs. 6 and 7), $^{48}Ca(^{15}N, 2np\gamma)^{60}Fe$ (Ref. 7), $^{58}Fe(t, p\gamma)^{60}Fe$ (Ref. 7), and ${}^{58}Fe(t, p)^{60}Fe$ (Ref. 8) experiments have provided new information on the energies, spins, and parities of states in ⁶⁰Fe. The last reaction also provided an accurate mass excess for ${}^{60}Fe$ of -61.404 ± 0.005 MeV, which was needed to calculate the 60 Mn mass excess from the present Q₈ measurement.

In the present work, we have measured the γ -ray singles and γ - γ coincident spectra following the β -decay of ⁶⁰Mn. γ rays have been attributed to the decay of 60 Mn by comparisons of their energies with those of previously reported levels in ${}^{60}Fe$ and by their coincidence relations with known γ rays in 60 Fe. From these measurements, the half life, ground-state spin, parity, and decay scheme of ⁶⁰Mn were determined. In addition, a $\beta-\gamma$ coincidence experiment was performed to measure the total decay energy Q_{β} . The resulting mass excess of ⁶⁰Mn is compared with various predictions. The level structure of 60 Fe deduced from the present and previous measurements is compared with a shell-model calculation.

II. EXPERIMENTAL METHOD

The isotope 60 Mn was produced by bombarding 48 Ca targets with 50-70-MeV 18 O ions. The predominant reaction responsible for the production of 60 Mn at these energies is most likely ⁴⁸Ca(18 O, α *pn*)⁶⁰Mn. All of the data presented in this paper were obtained using a bombarding energy of 56 MeV, because an initial survey indicated that this energy produced the highest yield of ${}^{60}\text{Mn}$ relative to other contaminant radioactive species. The compound-nucleus evaporation code ALICE^9 predicts a $\rm ^{60}M$ n production cross section of 80 $\rm \mu k$ at this bombarding energy. Beam currents of 50-200 nA (electrical) of $^{18}O^{6+}$ ions were supplied by the Argonne National Laboratory tandem ac-

celerator. Each target assembly consisted of a 96.8% enriched 48 Ca rolled foil 1.17 mg/cm² thick placed on a 0.05 mm tantalum backing which was mounted on a Delrin holder. Approximately 100 μ g/cm² of gold was flashed over the calcium to inhibit oxidation. The targets were irradiated in vacuo and then transferred to a shielded counting area using a multiple-target pneumatic transfer system (multiple rabbit) described in detail by
Parks *et al*.¹⁰ This device can hold up to eight Parks $et\ al.^{10}$ This device can hold up to eight targets. After a target is activated, transferred, counted, and returned, a new target is rotated into place. This allows time between bombardment/count cycles which reduces the buildup of unwanted long-lived activities. Control of bombardment, target transfer, and computer gating functions was accomplished by means of a crystalcontrolled sequence timer.

Delayed γ rays were measured using one or two 15% efficient Ge(Li) detectors. In order to reduce the number of β particles entering the detectors, Lucite absorbers of 0.65-cm thickness were placed between the sample and the detector(s). A 4.4-cm deep by 7-cm diameter NE102 plastic scintillator coupled to an RCA8575 phototube was used to detect β particles in coincidence with decay γ rays. Energy and efficiency calibrations of the γ -ray detectors were performed using standard γ -ray sources.

In the singles measurement, γ spectra were routed into eight 4096-channel time bins, each of 1-s duration. Separate scaling of a pulser, gated by the busy signal from the analog-to-digital converter, provided dead-time information. The overall resolution was typically 2.5 keV (full width at half-maximum) for the 1332 -keV line of ${}^{60}Co$.

Coincidence experiments mere performed using constant fraction timing techniques with commercially available electronics. In the $\beta-\gamma$ coincidence experiment, a pile-up rejector was used with the plastic scintillator. Data for both γ - γ and β - γ coincidence experiments were recorded on magnetic tape for subsequent off-line analysis. Coincident γ ray or β ⁻ spectra were produced from events in one detector by setting windows on γ -ray peaks of interest in the other detector. These spectra mere corrected for background-coincident events by subtracting spectra obtained by setting windows on background regions both above and below the peaks of interest.

III. RESULTS

A. Half-life and decay scheme of 60 Mn

 60 Mn was identified by the observation of an 823.6-keV γ ray which decayed with a half life of approximately ² s. This energy agreed with previous measurements of the energy of the first ex-

NEUTRON-RICH ISOTOPE...
cited state of ${}^{60}Fe.{}^{6,7}$ Other strong γ rays with energies of 492.9 and 1968.8 keV were also observed to decay with half-lives near ² s.

The half-life of ⁶⁰Mn was determined by observing the decays of the 823.6- and 1968.8-keV γ rays. After correction for dead time, the composite decay curve for these two γ rays yielded a halflife of 1.79 ± 0.10 s.

More than one hundred γ rays were observed in the singles spectra following the bombardment of 48 Ca with 56-MeV 18 O. In order to make shortlived activities stand out among such a large number of γ rays, a useful technique is to subtract the sum of the spectra in time bins 5, 6, 7, and 8 from the sum of the spectra in time bins 1, 2, 3, and 4 after correcting for system dead time. The result of this operation is shown in Fig. 1 to illustrate the singles data. The dead-time correction procedure used to produce this spectrum was to make the yields of the 1163.5-keV γ ray from the decay of the 13.9-m 62°C equal in all time bins before the appropriate additions and subtractions were performed. This spectrum has also been approximately corrected for detector efficiency by multiplying the contents of each channel by the channel number.

In order to determine the decay scheme of 60 Mn, a γ - γ coincidence experiment was performed. Two examples of the coincidence spectra obtained in this experiment are shown in Fig. 2. Four previously unassigned γ rays were identified in the coincidence gates. γ rays with energies of 1290.2, 1475.3, and 2299.3 keV, which had been obscured in the singles spectra by contaminant γ rays of nearly the same energies, appeared in the coincidence data. In addition, a 678.2-keV transition, which had been seen weakly in the singles spectra, appeared in the 1290-keV coincidence gate. A 271.8-keV γ ray was observed in the singles data to decay with a half-life of 1.84 ± 0.23 s, which is approximately equal to that of the 60 Mn decay γ rays. This γ ray, however, did not appear in any of the coincidence gates that were set. Furthermore, the intensity of this γ ray relative to 60 Mn decay γ rays did not remain constant when the 18 O bombarding energy was varied. As a result, this γ ray was not assigned to the decay of 60 Mn.

The energies and relative intensities of the γ rays attributed to the decay of 60 Mn are summarized in Table I. The γ -ray intensities were normalized with respect to the intensity of the 823.6 keV γ ray and were corrected for coincidence summing. The relative intensity of the 1290.2-keV γ ray had to be determined from the 824-keV coincidence spectrum because of the interference from other γ rays in the singles spectra. For

FIG. 1. β -delayed γ -ray singles spectrum observed following the bombardment of ⁴⁸Ca with 56-MeV ¹⁸O. A spectrum obtained during a 4-s period has been subtracted from that obtained in the previous 4 s. The spectrum has been approximately corrected for detector efficiency by multiplying the contents of each channel by the channel number.

FIG. 2. Backgound-corrected γ spectra observed in coincidence with the 493- and 824-keV γ rays.

I_{ν} (Relative)	
24.4 ± 1.3	
3.0 ± 0.7	
100	
13.9 ± 1.3	
14.6 ± 1.4	
71.1 ± 3.6	
17.6 ± 2.0	

TABLE I. Energies and relative intensities of γ rays in 60 Fe observed following the β decay of 60 Mn.

this transition, it was assumed that the relative detection efficiency was unaffected by the coincidence circuitry. The relative intensities of all the other 60 Mn decay γ rays were determined from the singles measurements.

Using the results of the γ -ray singles and γ - γ

 $(3)^+$

\2'

9.3 ± 1.3 5.75 $4 + 1.3 + 5.75$

 2^{+}

 $O⁺$

60 $_{\rm F}$ 26^{re}34

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 E_{x} (keV) 6.8 2792.3

2299

He₃₃₆ (ego)

- ი

823.6

 1.8 ± 0.1 s

Log ft 4.56

60
25^{Mn}35 $Q_B = 8.5 \pm 0.1$ MeV

 6.7 ± 2.0 5.83

 0.3 ± 3.5 >6.50

 I_B (%) 83.7 ± 3.6

FIG. 3. The decay scheme of 60 Mn. Only those 60 Fe levels involved in the decay of 60 Mn are shown. The γ -ray energies and relative intensities were measured in the present work. The spin and parity assignments are discussed in the text. The numbers enclosed by parentheses are the γ -ray yields per 100 60 Mn decays.

coincidence experiments, the decay scheme of ' 60 Mn shown in Fig. 3 was constructed. All seven γ rays attributed to the decay of 60 Mn have been placed in this scheme. The levels at 823.6, 2114, and 2299 keV were previously observed in in-beam γ ray^{6,7} and in charged-particle⁸ studies of ⁶⁰Fe. The 2792.3-keV level was not previously reported. The excitation energies of the 823.6-, 2114-, 2299-, and 2792.3-keV levels have estimated uncertainties of 0.3, 1, 1, and 0.⁶ keV, respectively.

The β -branching ratios shown in Fig. 3 were calculated from the intensity balance of γ rays for each state in ${}^{60}Fe$. In these calculations, it was assumed that there is no β branch to the ⁶⁰Fe ground state. Log ft values have been determined from the tables of Gove and Martin 11 by using a total decay energy of 8.51 MeV as discussed below.

B. Mass of 60 Mn

The total β -decay energy of 60 Mn was measured in a $\beta-\gamma$ coincidence experiment. Data were cyclically accumulated in 4-s counting periods following 3-s irradiations of the targets. Digital gates were set on the 492.9- and 1968.8-keV γ rays and on background regions both above and below these peaks. The background-corrected β spectra coincident with the 492.9- and 1968.8keV γ rays, respectively, were added together and the resulting spectrum is shown in Fig. 4.

The end-point energy of this β spectrum was determined by using the shape-fitting technique determined by using the shape-fitting technique de- $\,$ scribed by Davids $e t \, a l.^{12} \,$ This method utilizes a least-squares minimization procedure to stretch or compress a standard β spectrum along the energy (horizontal) axis and to normalize the resulting spectrum in order to fit other β spectra. β spectra with well-known end-point energies from the decays of other isotopes produced in the same $\beta-\gamma$ experiment are fitted by this method. A similar fit is made to the β spectrum whose end-point energy is to be determined. The stretch factors calculated for the known isotopes are then plotted as a function of end-point energy and a linear least-squares fit is made to the data. The energy calibration provided by this fit, together with the stretch factor determined for the unknown spectrum, is used to calculate the unknown endpoint energy.

A major reason for using this technique is that most of the data points in the spectrum may be used in the fitting procedure. A second reason is that experimental problems such as poor detector resolution, losses of β particles from the edges of the detector, backscattering, and energy

FIG. 4. Background-corrected β spectrum from the decay of ⁶⁰Mn to the 2792.3-keV state in ⁶⁰Fe. The smooth curve is a fit made to the data as described in the text. The decay scheme shown in the inset includes only those 60 Fe levels and γ rays that are involved in producing this β spectrum.

losses in the material between the source and the detector all affect the observed shapes of β spectra. By using an experimentally obtained β spectrum as a standard, the major influences of these effects are automatically included in the end-point determination.

The standard used in the present measurement was taken from the β spectrum coincident with the 1779-keV γ ray from the decay of ²⁸Al. This isotope was produced by the ${}^{12}C({}^{18}O, pn){}^{28}Al$ reaction on hydrocarbon contaminants which accumulated on the targets. The resulting spectrum shown in Fig. 5 is a pure β ⁻ spectrum with an end-point

energy of 2865 ± 2 keV. The smooth curve, drawn through the data by hand, was then used as the standard shape. The accuracy of this hand-drawn fit was checked by fitting the original 28 Al data with this standard.

Similar γ -ray-coincident β spectra from the decays of other isotopes were also produced. These spectra were fitted by using the techniques described above. Table II lists the isotopes whose β spectra were fitted, the energies of the coincident γ rays, the known β end-point energies, and the resulting stretch factors. The fitting region and χ_v^2 , the reduced χ^2 , of each fit is also

Nucleus	γ ray (keV)	Region fit (Channel Nos.)	Stretch factor	χ_{ν}^2	β end point (keV)
$51T_1$	320	$10 - 38$	0.735 ± 0.003	1.62	2146 ± 4
49Ca	3084	$8 - 36$	0.750 ± 0.005	0.93	2184 ± 7
20 O	1057	$8 - 35$	0.965 ± 0.045	1.12	2759 ± 9
28 Al	1779	$5 - 48$	1.000 ± 0.001	0.70	2865 ± 2
50 _{Ca}	$1519 + 1591$	$8 - 50$	1.079 ± 0.007	1.49	3120 ± 18
19 _O	1357	$8 - 50$	1.142 ± 0.008	0.95	3265 ± 4
50 Sc	524	$9 - 61$	1.298 ± 0.007	1.55	3694 ± 17
15 _C	4277	$10 - 64$	1.571 ± 0.012	2.14	4473 ± 2
	(D.E. of 5299)				

TABLE II. Stretch fit results for nuclei with known β end-point energies.

FIG. 5. Background-corrected β spectrum from the decay of 28 Al to the 1779-keV state in 28 Si. The smooth hand-drawn curve through the data is the "standard shape" described in the text.

listed. All of the spectra used in the present analysis are pure β spectra produced by allowed transitions. As described above, a linear leastsquares fit was made to a plot of the calculated stretch factors versus end-point energies. The results of this procedure are shown in Fig. 6. This fit was used as the calibration for determining the end-point energy of the β -decay branch of 60 Mn to the 2792.3 -keV level in 60 Fe.

The effects of coincident γ rays and of multiple β feedings can be included in the present method of analysis. The solid curve shown in Fig. 4 is the result of a fit to the data made by assuming

FIG. 6. β end-point energy calibration for the known β spectra obtained in the $\beta-\gamma$ coincidence experiment. The straight line is a linear least-squares fit to the data which are listed in Table II.

the decay scheme shown in the inset. The peak in the curve near channel 11 represents the Compton edge of the coincident 824-keV γ ray. The value of the stretch parameter deduced from this fit to the β spectrum of 60 Mn is listed in Table III. Combining this stretch factor with the results of the calibration yields the end-point energy also given in Table III. The total β -decay energy Q_8 is obtained by adding the 2792.3-keV excitation energy in ⁶⁰Fe to this end-point energy, and the result is 8.51 ± 0.10 MeV. Using the value of the mass excess of ${}^{60}\text{Fe}$ measured by Norman *et al.*,⁸ the resulting mass excess of 60 Mn is -52.89 ± 0.10 MeV. The quoted uncertainty includes a 50-keV systematic uncertainty for the shape-fitting procedure.

C. Ground-state spin and parity of 60 Mn

The spins and parities of the levels at 2114- and 2299-keV excitation energy in 60 Fe have been shown by previous workers^{7,8} to be 4^+ and 2^+ , respectively. The measured $\log ft$ values of the transitions to these states are both < 5.9 . Ac-

cording to the rules deduced by \rm{Raman} and $\rm{Gove,}^{13}$ these β branches thus are allowed transitions. Therefore, in these transitions, J can change by 0 or \pm 1 with no change in parity. These considerations uniquely determine the ground-state spin and parity of 60 Mn to be 3^* .

D. Spin and parity of the 2792.3-keV level in ⁶⁰ Fe

As was mentioned above, the only level in ${}^{60}Fe$ observed in the decay of 60 Mn that had not been previously reported is the 2'792.3-keV level. The measured logft value of the β branch to this state is 4.56. Thus, according to the rules o<mark>f Ra</mark>man
and Gove,¹³ this transition is definitely allowed and $\rm{Gove,}^{13}$ this transition is definitely allowed Knowing that the spin and parity of 60 Mn is 3⁺ then restricts the spin and parity of the 2792.3-keV state to be 2^* , 3^* , or 4^* .

If the spin and parity of this state were 2^+ , then it might be expected to decay to the (0^+) level at 1975 keV and/or to the 0^+ ground state. However, neither transition is observed. If the spin and parity were 4^+ , then 492.9 - and 1968.8 -keV transitions would both be $E2$. The estimated $E2$ single-particle rate of the 1968.8-keV transition is more than 1000 times faster than that of the 492.9 keV transition. Thus, in order to explain the observed branching ratios, the 492.9-keV transition would have to be enhanced by a factor of more than 340 relative to the 1968.8-keV transition. Such a large relative enhancement is unlikely. The fact that the 2792.3-keV level was not observed in $(t, p\gamma)$ (Ref. 7) or (t, p) (Ref. 8) studies of 60 Fe suggests that this state could have unnatural parity, for it is well known that the (t, p) reaction tends to strongly populate only naturalparity states. The only unnatural-parity state that could be fed by an allowed transition is a 3' state. It has been observed that the decay of 58 Mn has a 76% branch to ^a 3' state at 2134-keV exhas a 76% branch to a 3^+ state at 2134-keV ex-
citation in 58 Fe.¹⁴ Perhaps an analogous situatio exists in the decay of 60 Mn. As a result of these considerations, we suggest a spin and parity of $(3)^+$ for the 2792.3-keV level in ${}^{60}Fe$, although the 2' and 4' possibilities are not ruled out.

IV. DISCUSSION

A. Mass of ⁶⁰Mn

In Table IV, the measured value of the mass excess of ⁶⁰Mn is compared with various mass predictions. The measured Q_8 is also compared with the values predicted by these models. The Q_8 values not enclosed by parentheses were calculated using the mass excess of ⁶⁰Fe measured by Norman et al^8 . The Q_8 values in parentheses were calculated using the 60 Fe mass excess pre-

W. D. Myers, Ref. 15.

^b H. V. Groote, E. R. Hilf, and K. Takahashi, Ref. 15.

 c P. A. Seeger and W. M. Howard, Ref. 15.

^d S. Liran and N. Zeldes, Ref. 15.

^e J. Jänecke, Ref. 15.

E. Comay and I. Kelson, Ref. 15.

dicted by the various models. These comparisons show that the predictions vary from one another by as much as 1.6 MeV and from the measured value by as much as 0.9 MeV. The updated Garvey-Kelson prediction of Jänecke¹⁵ and the droplet model prediction of Myers¹⁵ show the best agreement with the measured mass excess value.

B. Ground-state spin and parity of 60 Mn

A single-particle shell model predicts that the ground-state configuration of ${}^{60}_{25}Mn_{35}$ should be $\pi(1f_{7/2})^{-3}\nu(1f_{5/2})^3$. The ground-state spins and parities of the odd- A Mn isotopes indicate that the $\pi(f_{7/2})^{-3}$ configuration is coupled to $\frac{5}{2}$ and that of the adjacent odd-A $N=35$ isotone, 61 Fe, indicates that the $\nu(f_{5/2})^3$ configuration can be coupled to $\frac{3}{2}$ (Ref. 16). If one assumes the same couplings to be present in ⁶⁰Mn, then Nordheim'
rules, as modified by Brennan and Bernstein,¹⁷ rules, as modified by Brennan and Bernstein,¹⁷ suggest a ground-state spin and parity of 3' for ${}^{60}\text{Mn}$. It should be noted that ${}^{54}\text{Mn}$, ${}^{56}\text{Mn}$, and probably 58 Mn all have spins and parities of 3^+ (Ref. 16). By analogy, it could be expected that the spin and parity of 60 Mn would also be 3⁺. However, it should also be pointed out that although the above-mentioned rules predict the correct spins and parities for 54 Mn and 60 Mn, they fail to predict the correct values for ⁵⁶Mn and ⁵⁸Mn.

C. Level structure and electromagnetic transitions in ^{60}Fe

A shell-model calculation of the ${}^{60}Fe$ level scheme and electromagnetic transition branching ratios was performed several years ago by Larratios was performed several years ago by Lar
son, Raman, and McGrory.¹⁸ In this calculation an inert⁴⁸Ca core was assumed. The protons were restricted to the $1f_{7/2}$ shell, and the neutrons were allowed to occupy any vacant $1f-2p$ orbital.

FIG. 7. A comparison of the experimental level structure and electromagnetic transitions below 3.1 MeV in 60 Fe with the shell-model calculations of Larson, Raman, and McGrory (Ref. 18). Dashed lines indicate probable correspondences. The experimental level scheme is a compilation of the results of the present work and those of previous studies (Refs. 7 and 8) of 60 Fe. The numbers enclosed by parentheses are the branching percentages for each transition.

The form of the interactions and the matrix elements used in this calculation are described by McGrory¹⁹ in his previous studies of ⁵⁶Cr and ⁵⁸Fe. Figure 7 shows a comparison of the calculated level structure and electromagnetic transitions in ⁶⁰Fe with the structure as determined by the present and previous experiments.^{7,8}

For excitation energies below 2.4 MeV, the agreement between the calculated and the observed positions of the energy levels is very good. The calculated energies of the first four excited states are all within 75 keV of the observed values, and the predicted ordering of the $0^+, 4^+, 2^+$ triplet is in agreement with experiment. As expected, the

 $0⁺$ and $4⁺$ members of this triplet decay only to the first excited state. The 2⁺ member, however, is predicted, and is observed to decay strongly to both the ground state and first excited state. The spin, parity, and decay modes of the observed 2.358-MeV level have not been experimentally determined. As a result, no correspondence has been made between this level and any of the calculated levels.

Above 2.4-MeV excitation energy, there is still reasonable agreement between the calculated and observed level structures. However, the predicted ordering of levels is not in complete agreement with experiment, and some of the calculated levels

2183

have not been observed experimentally. The observed $(3)^+$ level at 2.792-MeV probably corresponds to the calculated 2.507-MeV state. As predicted, its predominant decay mode is to the first 2' excited state. The observed branching ratios for the decays to the second 2' state and to the first 4' state, however, appear to be interchanged with respect to the calculations. Above 2.8-MeV excitation, no correspondences have been made between the observed and the calculated levels.

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