Vibrational levels in the 86-neutron nucleus ¹⁵⁰Gd[†]

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The decay of 3.6-h ¹⁵⁰Tb to levels in ¹⁵⁰Gd has been studied by γ -ray and conversion-electron spectroscopy. A level scheme is proposed which accounts for all but 35 of the 256 γ rays assigned to this decay. Most of the low-lying levels in ¹⁵⁰Gd show a remarkably close correspondence with the predictions of the simple vibrator model. Possible candidates are observed for three-phonon quadrupole states and for the states coupling quadrupole and octupole phonons. The low-lying positive-parity levels can be adequately described by the phenomenological collective model of Gneuss and Greiner. The microscopic boson-expansion model of Kishimoto and Tamura, however, appears to predict ¹⁵⁰Gd to be more transitional than the data imply. Systematics of the 86-neutron isotones show increased stability at 64 protons. This can be correlated with the filling of the $g_{7/2}$ and $d_{5/2}$ spherical proton orbitals.

RADIOACTIVITY ¹⁵⁰Tb [from ¹⁵¹Eu(³He, 4*n*)], measured $T_{1/2}$, E_{γ} , E_{β^*} , I_{γ} , I_{∞^*} , $\gamma - \gamma$, and $\beta^* - \gamma$ coin; deduced Q, log ft. ¹⁵⁰Gd deduced levels, ICC, J, π . Enriched target, Ge(Li), and Si(Li) detectors.

I. INTRODUCTION

The even-even 86-neutron nuclei are not well understood. Presumably they have spherical ground-state shapes and vibrational collective levels since they lie between the 82-neutron closed shell and the 88- to 90-neutron shape transition at the start of the deformed "rare earth" nuclei ($150 \le A \le 190$). Recent shell-model calculations¹⁻³ support the notion that many 86-neutron isotones have spherical ground-state shapes. Little experimental data have been accumulated concerning collective levels in these nuclei. Such data would be useful in obtaining additional insight into the systematics of nuclei around the shape transition and in providing examples of vibrational nuclei with more than 82 neutrons.

In a previous study of $5.8 - \min {}^{150}$ Tb decay,⁴ we noted that the energy spacing of the lowest 2⁺, 4⁺, and 6⁺ levels in 150 Gd suggested a close correspondence with that expected for the vibrational model. Additional data on other low-spin multiphonon vibrational states are required to support this vibrational description of 150 Gd. Such states were not populated in the decay of the high-spin $5.8 - \min {}^{150}$ Tb whose J^{τ} is (9⁺).⁴ In an attempt to find these states we have studied in detail the decay of the low-spin $3.6 - h {}^{150}$ Tb.

While the work was in progress, Vylov *et al.*^{5, 6} reported that 96 γ rays were associated with 3.6-h ¹⁵⁰Tb decay. A scheme was proposed which accommodated 37 of these transitions among 17 levels. In general their experimental results are in good agreement with those obtained here, except as noted in Sec. II B. Levels in ¹⁵⁰Gd have also been observed in 5.8-min ¹⁵⁰Tb decay,^{4,7,8} in (p,t)

reactions, ⁹⁻¹¹ and with in-beam γ -ray spectroscopy.¹²⁻¹⁵ The results of these and previous 3.6h ¹⁵⁰Tb decay studies^{5, 6, 14, 16, 17} have recently been summarized.¹⁸ Preliminary data from this work were included in the summary.

In the present study 256 γ rays are assigned to 3.6-h¹⁵⁰Tb decay. The proposed decay scheme has 73 levels and accounts for all but 35 of the observed γ rays. Most of the positive- and negative-parity levels below 2 MeV can be qualitatively described in terms of the vibrational model that includes quadrupole and octuple vibrations. A more quantitative description of the positive-parity states is obtained with the phenomenological collective model of Gneuss and Greiner.¹⁹ The energy spectrum is also compared with the predictions of the microscopic boson-expansion model of Kishimoto and Tamura.²⁰

II. EXPERIMENTAL PROCEDURE AND RESULTS

A. Source preparation

Sources of 3.6-h ¹⁵⁰Tb were prepared via the ¹⁵¹Eu(³He, 4n)¹⁵⁰Tb reaction with 35-MeV ³He ions from the Texas A & M variable energy cyclotron. Targets consisted of 5 to 10 mg of Eu₂O₃ enriched to at least 95% in ¹⁵¹Eu and were irradiated from 30 to 60 min with beam currents of 3 to 5 μ A. Counting was not begun until one hour after bombardment. Thus the 5.8-min ¹⁵⁰Tb and 4.2-min ¹⁵²Tb activities were removed by decay. Significant amounts of 17-h ¹⁵¹Tb and 18-h ¹⁵²Tb as well as smaller amounts of 4.1-h ¹⁴⁹Tb, 2.3-day ¹⁵³Tb, the several isomers of ¹⁵⁴Tb, and ²⁴Na were observed in the spectra from these sources.

Complications in the assignment of γ rays to

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¹⁵⁰Tb decay arose primarily from the presence of ¹⁵¹Tb and ¹⁵²Tb. Only the most intense rays from the other activities were visible in the γ - ray spectra. Since the half-life of ¹⁵⁰Tb is 3.6 ± 0.2 h, γ rays associated with ¹⁵⁰Tb could be separated from those associated with ¹⁵¹Tb and ¹⁵²Tb decay on the basis of half-life. The decay data for ¹⁵¹Tb by Vilskii *et al.*²¹ and for ¹⁵²Tb by Zolnowski, Funk, and Mihelich²² were used to analyze unresolved γ ray multiplets consisting of an impurity line and a γ ray from ¹⁵⁰Tb.

B. Singles measurements

Singles γ -ray spectra were obtained with a 33cm³ Ge(Li) detector of resolution 1.8 keV full width at half maximum (FWHM) at 1.33 MeV. Standard modular electronics coupled with an online PDP-15/40 computer were used to acquire 8192-channel spectra. These spectra were analyzed with a modified version of the program SAM PO.²³ The calibration methods for energy and efficiency are described elsewhere.²⁴

The energies and relative intensities of the γ rays assigned to 3.6-h¹⁵⁰Tb are summarized in Table I. Included in the table are 67 weak transitions whose assignment to ¹⁵⁰Tb decay is tentative. Some of the relative intensities were deduced from γ - γ coincidence results while others have been corrected for unresolved components from either escape peaks or transitions from ¹⁵¹Tb or ¹⁵²Tb decay.

Where they overlap, these results are in excellent agreement with those reported by Vylov *et al.*^{5,6} except for the energies of γ rays above 4 MeV. In addition, the γ rays which Vylov *et al.*^{5,6} reported at 1343.1, 2488.5, 2750.5, 3083.7, 3424.0, and 3604.0 keV are interpreted here as escape peaks.

Conversion-electron spectra were taken with two different spectrometer systems, both of which employ Si(Li) detectors to determine electron energies. The spectrum from 0.4 to 1.2 MeV was obtained with a cooled Si(Li) detector which directly views the source.²⁵ The second spectrometer system uses a broad-range, trochoidal-path steering magnet which selectively transports electrons from the source to a shielded and cooled Si(Li) detector. This system has been described in detail by Gono *et al.*²⁶ The spectrum from 0.8 to 1.5 MeV was obtained with this device.

The analysis of these spectra to obtain K-conversion coefficients is summarized in Table II. In some cases the observed electron intensity represents the sum of unresolved peaks. For three of these cases, one of the components was so small that it could be neglected. The rest were analyzed by assuming a multipolarity for all but one of the components. In general these results agree with those obtained by Vylov *et al.*^{5, 6} except for transitions between 560 and 570 keV and those at 1430.51 and 1453.6 keV.

C. Coincidence measurements

The Ge(Li) detector described above and a second counter of comparable size but poorer resolution were employed for a γ - γ -t coincidence measurement. The relative time t between more or less simultaneous events in both detectors was determined by conventional electronics using extrapolated leading-edge timing for the Ge(Li) detector signals. The data were recorded in an event-byevent mode on magnetic tape with an on-line PDP-15/40 computer.

The events were sorted into a $4096 \times 4096 \times two$ channel spectrum using an IBM 7094 computer. The 4096-channel dimensions represent the energies of the coincident γ rays while the two-channel dimension separates prompt and random events. Coincidence spectra, corrected for random and Compton-background coincidences, were obtained by summing the appropriate portions of the 4096 \times 4096 \times two-channel spectrum.

The coincidence spectra were analyzed to obtain energies and intensities of the coincident γ rays. The detailed results of this analysis and a more complete description of the data reduction and analysis procedure may be found in Ref. 24.

The Q value for the decay of $3.6 - h^{150}$ Tb was determined from a $\beta^* - \gamma$ coincidence measurement. This experiment was essentially the same as the $\gamma - \gamma - t$ coincidence measurement except that the second detector was a cylindrical NE102 scintillator 4 cm long and 5 cm in diameter. The data were corrected for $\gamma - \gamma$ coincidences from a measurement in which a 4.3 g/cm² Cu absorber was placed between the source and the scintillator. This spectrum, suitably normalized, was subtracted from the measurement without absorber. Corrections for summing with annihilation radiation were made by calibrating the scintillator for energy with radioactive sources of known β^* energy.

The spectrum of β^* particles populating a particular level was obtained by summing the β^* spectra gated by the γ rays depopulating that level. Endpoint energies, determined from Kurie plots²⁸ and log ft values²⁹ for these β transitions, are summarized in Table III.

Population of the ¹⁵⁰Gd ground state was indicated in the β^* spectrum gated by 511-keV annihilation radiation. The end-point energy of the highest β^* group in this spectrum was ~625 keV higher than that of the group observed to feed the 638.05-

Energy (keV)	Relative intensity	Placement ^a	Energy (keV)	Relative intensity	Placement ^a
128.0(3) ^b	5(2)	2956 → 2828	1003.8(3)	12(2)	2521 → 1518
120.0(0)	· (⊥)	[2687 → 2559]	1000.0(8)	12(2)	[3083→2080]
153.9(3)	10(2)	1288→1134			$[4379 \rightarrow 3375]$
100.0(0)	10(2)	$[4265 \rightarrow 4111]$	1045.72(10)	175(10)	2180→ 1134
222.8(3) ^b	9(2)	[4000 4111]	1043.72(10) 1061.9(5) ^b	15(7)	1700-+638
275.6(3) ^b	8(2)		1075.3(1)	85(5)	2209 - 1134
300.4(5)b	9(3)	4322→4022	1091.2(3)	25(10)	$2521 \rightarrow 1430$
303.1(5) ^b	5(2)	1022 1022	1094.4(3)	30(5) ^d	$2687 \rightarrow 1592$
330.1(2)	15(3)	2985 → 2654	1120.1(3)	20(5)	$2408 \rightarrow 1288$
00011(2)	10(0)	$[3840 \rightarrow 3510]$	1128.2(4)	25(5)	2828 -+ 1700
338.2(2)	25(5)	2326→ 1988	1134.3(3)	15(5)	$2564 \rightarrow 1430$
378.8(5) °	20(10)	$2326 \rightarrow 1947$	1135.3(5) ^c	15(5)	$3344 \rightarrow 2209$
384.1(3)	60(10)	1518 - 1134	1157.7(5)	30(10) ^d	$2365 \rightarrow 1207$
385.5(5) °	10(5) ^d	$1592 \rightarrow 1207$	1168.7(2)	60(5) ^d	$2687 \rightarrow 1518$
411.7(4)	115(10) ^d	1700→ 1288	1176.0(2)	65(5) ^d	1814 -+ 638
412.4(2) ^e	~10	1701 - 1288	1191.1(4)	15(5) ^d	4176-2984
425.9(5)°	50(25)	2985 - 2559	1224.2(5)	12(4)	$2654 \rightarrow 1430$
437.1(1)	130(5)	1955 - 1518	()		$[4344 \rightarrow 3119]$
491.7(2)	25(5) ^d	2084 → 1592	1233.0(4)	25(5)	2521→1288
496.30(10)	2 0 6 0 (4 0) ^d	$1134 \rightarrow 638$	1253.1(3)	12(3)	$2845 \rightarrow 1592$
525.0(2)	80(10)	1955 - 1430		(-)	[3344 - 2091]
526.0(3)	17(7) ^d	1814 - 1288	1256.6(5) ^b	~7	$2687 \rightarrow 1430$
557.5(1)	50(5) ^d	1988 - 1430	1		[2956 - 1700]
565.7(1)	155(10)	$1700 \rightarrow 1130$	1274.6(2)	25(5) ^d	2408 - 1134
566.7(2)	20(5)	$1701 \rightarrow 1134$	1291.65(10)	225(10) ^d	$2426 \rightarrow 1134$
569.1(1)	350(10) ^d	1207 - 638	1317.6(3)	50(10) ^d	$1955 \rightarrow 638$
573.4(2)	45(5)	2091 - 1518	1350.1(5)	150(50)	1988 → 638
602.8(2)	30(5) ^d	$2687 \rightarrow 2084$	1351.9(5)	35(15) ^d	$2559 \rightarrow 1207$
609.3(3) ^b	10(5)	2564 → 1955	1356.1(3)	20(5) ^d	$2786 \rightarrow 1430$
		[3035→2426]	1387.0(4)	15(5)	$2521 \rightarrow 1134$
638.05(10)	≡10 000	638→ 0	1392.1(3) ^b	5(2) ^d	$2985 \rightarrow 1592$
649.5(5) °	40(20) ^d	$2080 \rightarrow 1430$	1415.0(2)	40(10) ^d	$2845 \rightarrow 1430$
650.4(2)	$560(20)^{d}$	1288-+638	1430.51(10)	300(20) ^d	1430-0
661.0(3)	≤30 ^d	2091 - 1430	1430.5(3) °	$40(20)^{d}$	$2564 \rightarrow 1134$
666.3(2) ^b	13(5) ^d	2654-1988	1442.0(5)	10(3) ^d	2080 - 638
699.4(2)	45(5) ^d	1988 → 1288	1443.6(3)	55(5) ^d	$3035 \rightarrow 1592$
743.8(2)	20(5)	$2262 \rightarrow 1518$	1446.2(3)	70(5) ^d	$2084 \rightarrow 638$
746.6(2)	10(3)	2956 → 2209	1453.62(10)	500(15)	2091 → 638
748.3(2)	65(5) ^d	1955→1207	1466.6(4)	10(2)	$2755 \rightarrow 1288$
772.6(2)	30(5) ^d	2365→1592	1493.0(5)	15(5)	$2627 \rightarrow 1134$
779.0(5)	40(5) ^d	220 9 → 14 30	1516.5(5) ^c	45(10) ^d	3035→1518
791.1(5)	35(15)	2080→ 1288	1518.5(2)	320(10) ^d	1518→0
792.5(3)	610(20)	1430638	1525.8(1)	55(5)	$2956 \rightarrow 1430$
808(1) ^b	10(4) ^d	2326→1518	1542.0(2)	55(5)	$2180 \rightarrow 638$
813.1(3)	95(20) ^d	1947 → 1134	1552.7(2)	15(5)	$2687 \rightarrow 1134$
821.1(2)	190(10) ^d	1955-+1134			[4207 → 2564]
831.5(2)	15(5) ^d	2786→1955	1554.7(2)	25(5)	$2985 \rightarrow 1430$
880.3(1)	420(10) ^d	1518-+638	1564.2(2)	15(5)	3378→1814
884.6(2)	45(10)	2091 - 1207	1571.3(3) ^b	5(2) ^d	22 09 → 63 8
895.9(3)	30(5)	2326 → 1430	1580.0(3) ^b	10(5) ^d	$2786 \rightarrow 1207$
950.0(2)	125(10)	2084 → 1134			$[4145 \rightarrow 2564]$
952.0(5)	10(5)	3035→2084	1592.7(1)	225(10) ^d	1592→0
954.5(3)	160(10) ^d	1592 → 6 38	1605.6(5)	15(5) ^d	$3035 \rightarrow 1430$
957.4(2)	110(10)	2091 → 1134	1615.4(4)	15(5)	
968.4(2)	20(5)	2956-+ 1988	1620.7(3)	20(5)	$2755 \rightarrow 1134$
978.1(3)	15(5)	$2408 \rightarrow 1430$	1624.4(3)	$40(10)^{d}$	$2262 \rightarrow 638$
				40/=\d	0045 4005
995.5(3) 997.7(4)	20(5) ^d 12(3)	$2426 \rightarrow 1430$ $2985 \rightarrow 1988$	1638.6(10) ^b 1652.7(3)	10(5) ^d 20(10)	$2845 \rightarrow 1207$ $2786 \rightarrow 1134$

TABLE I. Energies and intensities of γ rays in 3.6-h $^{150}\mathrm{Tb}$ decay.

Energy	Relative		Energy	Relative	
(keV)	intensity	Placement ^a	(keV)	intensity	Placement ^a
1670.5(10) ^b	30(10) ^d	3658→1988	2498(1) ^b	10(5) ^d	4446-+ 1947
1688.2(4)	45(10)	2326-638	2532.3(4) ^b	10(5) ^d	
1688.8(4)°	20(5)	3119→1430	2539.5(3)	65(10) ^d	3178→638
1703.1(4) ^b	10(5)	3658→1955	2552.3(5) ^b	6(2) ^d	3840→1288
1726.9(4)	30(5) ^d	2365 - 638			[4145 - 1592]
1752.2(5)	15(5) ^d	3344 → 1592	2558.9(3)	30(5)	$2559 \rightarrow 0$
1770.8(2)	75(5) ^d	$2408 \rightarrow 638$	2565.5(5) ^b	7 (3)	4745 →2180
1778.0(5) °	15(8) ^d	2985 → 1207	2592.1(3)	35(5) ^d	3726 - 1134
1778.8(5)	70(10) ^d	2416-638	2614.4(4) ^b	10(2) ^d	$4207 \rightarrow 1592$
1788.13(10)	225(10) ^d	2426 - 638	2622.1(5) ^b	5(2) ^d	$4322 \rightarrow 1700$
1796.6(2)	25(5) ^d	$4207 \rightarrow 2408$	2022.1(0)	0(2)	$[3829 \rightarrow 1207]$
1811.8(5) ^b	10(5) ^d	4176 - 2365	2661.0(3)	20(5)	4745 → 2084
1831.9(4)	15(5)	4170 - 2305 4258 - 2426	2669.0(3)	15(5)	1110 2001
	25(10) ^d		2678.6(5)	15(5)	2678→0
1884.0(3)		2521 → 638	2018.0(3)	10(0)	
1900.6(10)	70(30) ^d	$3035 \rightarrow 1134$	2690.5(5) ^b	10(3) ^d	$[4379 \rightarrow 1700]$
1914.4(2)	40(20) ^d	3344 - 1430			3329 → 638
1926.8(3)	20(5)	2564 → 638	2706.6(4)	$40(5)^{d}$	$3344 \rightarrow 638$
1949.3(3)	25(5) ^d	3083→1134	2737.5(4)	40(5)	$3375 \rightarrow 638$
1955.7(3) ^b	~5 ^d	1955 → 0	2740.0(2)	30(10) ^d	$3378 \rightarrow 638$
	(- (-) đ	[4165→ 2209]	2774.7(3)	12(3)	$4207 \rightarrow 1430$
1985.1(2)	15(5) ^d	3119→1134	2808.0(3) ^b	8 (2) ^d	
2016.6(3)	130(20) ^d	2654 → 638	2841.3(5) ^b	5(2) ^d	0045 0
2040.6(3)	50(5) ^d	2678-638	2845.6(3)	25(5) ^d	2845 → 0
2044.3(3)	10(2) ^d	3178→1134		a a (=) d	[4545 → 1700]
2056.4(5)	15(5)	$3344 \rightarrow 1288$	2872.4(3)	30(5) ^d	3510→638
2091.7(3)	200(20)	2 0 91 → 0	2876.6(3)	15(5) ^d	$4165 \rightarrow 1288$
2117.0(3)	50(5) ^d	2755-638	2895.0(5) ^b	4(2) ^d	
21 37.2(4) ^b	5(2)	3344 → 1207	2913.4(4)	7 (3) ^d	$4344 \rightarrow 1430$
		[4545→ 2408]	2936.0(4) ^b	7 (2) ^d	4530→ 1592
2148.7(3)	140(5) ^d	2786 → 63 8	2952.7(4)	8(2) ^d	$4545 \rightarrow 1592$
2169.3(5) ^b	15(3) ^d	4379→2209	2976.2(5) ^b	5(2) ^d	$4265 \rightarrow 1288$
2173.4(5)	20(5)	$3460 \rightarrow 1288$			$[4111 \rightarrow 1134]$
2180.0(3)	50(5) ^d	2180 → 0	2984.4(5)	50(10) ^d	$2985 \rightarrow 0$
		[4545-+2365]	2993.4(5)	8 (3) ^d	
2190 .3(5) ^b	8(2)	2828 → 638	3009.9(4)	6(3)	4146 - 1134
2194.9(5)	40(5) ^d	3329→1134	3024.5(5)	5(2) ^d	
2201.3(3) ^b	10(3)		3034.0(10)	15(5)	4322 → 1288
2207.8(3)	140(10)	28 45 → 6 38	3035.5(10)	25(5)	$3035 \rightarrow 0$
2230.0(5) ^b	10(3) ^d	4322 → 2091	3042.6(4) ^b	10(5) ^d	4176→1134
		[4314 → 2084]	3096.3(4)	15(5) ^d	
2233.7(5) ^b	8(2)	4314 → 2080	3102.6(4)	9(1)	4237 → 1134
2241.3(4) ^b	20(10) ^d	3375→1134	3123.9(4) ^b	10(2) ^d	4258→1134
2258.7(5) ^b	4(2) ^d		3130.5(15)	4(2) ^d	4265→1134
2263.0(4) ^b	10(5) ^d	4344 → 2080	3134.7(4)	15(5) ^d	
2283.5(5) ^b	5(3)	4545→ 226 2	3152.4(4)	9(2)	4745→1592
2296.7(5) ^b	5(2) ^d	3726-1430	3168.3(4)	8(2)	
		[4111→1814]	3191.6(4)	6(2)	3829 → 638
2318.2(3)	55(10)	2956 → 638	3197.7(4)	6(2)	4406 → 1207
2329.6(5) ^b	5(2) ^d		3202.5(4)	10(2)	3840→638
2347.2(3)	25(5) ^d	2985 → 638	3238(1)	7(2)	4446 → 1207
2365.1(3)	120(10) ^d	23650	3262.3(4) ^b	7 (3) ^d	
2376.5(4) b	13(5) ^d	3510-1134	3315.0(5)b	4(2)	4745 → 1430
2396.5(4)	80(5) ^d	3035→638	3327.8(5)	8(2) ^d	
2410.0(3)	45(10)	3840→1430	3344.7(5)	5(1)	3344→ 0
2426.3(3)	125(5) ^d	2426-0	3352.3(5)	5(1)	
2449.8(5) ^b	3(1) ^d	4406-1955	3374(1) ^b	5(2) ^d	
	• •	[4530→2080]	3383.6(5) ^b	8(4) ^d	4022 → 63 8
2460.7(4)	11(3)	4545-2084	3389.8(5) ^b	5(2)	4524 → 1134
2494.6(4)	25(5) ^d		3411.3(5) ^b	4 (2) ^d	4545-1134

TABLE I. (Continued)

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Energy (keV)	Relative intensity	Placement ^a	Energy (keV)	Relative intensity	Placement ^a
3440.8(5)	6(2)		3885(2) ^b	2(1)	4524 → 638
3448.1(5) ^b	2(1)		3927 (2) ^b	4(2)	
3460(1) ^b	4(2)	3460→0	4022(2)	5 (2) ^d	4 022 → 0
3473(1)	5(1)	4111 - 638	4107(2)	15(5)	4745 → 638
3484.5(5)	7(2)		4112(2)	10(3)	4111→ 0
3489.7(5)	7(2)		4146(2)	3(1)	4146 → 0
3509.8(5)	8 (3) ^d	3510→0	4164(2)	2(1)	4165→0
3523(1)	9(1)		4206(1)	9(2)	$4207 \rightarrow 0$
3551(2) ^b	5(3)		4236(2) ^b	~1	$4237 \rightarrow 0$
3555(2) ^b	4(2)		4246(1)	6(2)	$4246 \rightarrow 0$
3571(1) ^b	4(2) ^d	4207 → 638	4257(2)	2(1)	$4258 \rightarrow 0$
3658.2(5)	14(2)	3658 - 0	4265(2)	5(1)	$4265 \rightarrow 0$
3675(2)	8(2)	4314 → 638	4284(2)	2(1)	$4284 \rightarrow 0$
3685(2) ^b	5(2)	4322-+638	4290(1)	14(4)	4290→0
3735(2) ^b	10(4) ^d		4314(1)	4(1)	4314→0
3769(2)	5(2)	4406638	4322(1)	6(1)	4322 → 0
3774(2) ^b	5(2) ^d		4344(2) ^b	~1	$4344 \rightarrow 0$
3817(2) ^b	3(2)		4379(2) ^b	~1	4379→ 0
3829(2)	5(2)	3829→0	4406(2)	7 (2)	$4406 \rightarrow 0$
3836(2)	7 (2)		4446(1)	11(2)	$4446 \rightarrow 0$
3846(2) ^b	3(2)		4531(2) ^b	~1	$4530 \rightarrow 0$

TABLE I. (Continued)

^aBrackets indicate alternate placements in the level scheme.

^bTentatively assigned to 3.6-h ¹⁵⁰Tb from γ -ray singles measurements.

^cAssigned to 3.6-h ¹⁵⁰Tb from γ - γ coincidence results.

^dIntensity corrected for unresolved γ ray from ^{151, 152}Tb or escape peak.

^e Existence of γ ray inferred from observation of 566.7-keV γ ray and 5.8-min ¹⁵⁰Tb decay scheme (Ref. 4).

keV level. Furthermore, the end-point energy of this group is higher than the highest energy β^* group among the impurities known to be present in the ¹⁵⁰Tb sources.

III. DECAY SCHEME

The proposed decay scheme for 3.6-h ¹⁵⁰Tb contains 73 levels, 15 of which are considered tentative. The data are presented in tabular form, Table IV, because of the complexity and amount of information to be summarized. Most of the transitions are placed in the scheme using either $\gamma - \gamma$ coincidence results or energy sums and differences. Transitions with energies above 4130 keV are assumed to feed the ground state since any other placement would be inconsistent with the measured Q value. The basis for placing each transition in the scheme is indicated in Table IV. Most of the levels were constructed in a straightfoward manner from the level scheme found in 5.8-min ¹⁵⁰Tb decay⁴ and the γ - γ coincidence data in the present work. For this reason detailed arguments for the establishment of each level are not given.

Transition multipolarities and decay patterns were used to deduce J^* values. Where conversion

data were unavailable, transitions were assumed to be E1, M1, or E2. In addition to the results listed in Table II, the multipolarities determined by Vylov *et al.*^{5, 6} for the 813.1- and 821.1-keV transitions [(E2) and E1, respectively] and by Kewley *et al.*¹³ for the 1176.0- and 1350.1-keV transitions (E1 and E2, respectively) have been used in making J^* assignments. The J^* of the 1592.7-keV level was chosen to be consistent with the E1 multipolarity assigned to the 954.5-keV transition. Vylov *et al.*^{5, 6} assign the multipolarity of the 1592.7-keV transition as M1. Other differences between the present conversion-electron results and those of Vylov *et al.*^{5, 6} have already been noted.

The amount of β -decay feeding each level was calculated from transition intensity balances and the absolute intensity of the 638.05-keV γ ray as reported by Vylov *et al.*^{5,6} As indicated in Table IV, most of the levels receive <1% of the β -decay intensity. In may cases this is less than the amount of unplaced γ -ray intensity which could feed the level. Log $f_0 t$ values were given in Table III for the transitions observed in the β^+ - γ coincidence experiment. For the rest of the levels the log $f_0 t$ and log $f_1 t$ values are consistent with the transitions being allowed, first forbidden or, in

	α_{K} (units 10 ⁻⁴)					
E_{γ}			Theo. ^b			
(keV)	Ice	Exp. ^a	E1/E2/M1	Multipolarity		
411.7 + 412.3 °	3.8(8)	320(70)	62/190/460	<i>M</i> 1		
437.1	3.4(3)	260(30)	55/160/310	M1 + E2		
496.30	8.7(10)	42(15)	41/120/220	E1		
525.0 + 526.0 ^c	2.1(4)	220(20)	36/100/190	(<i>M</i> 1)		
557.5	0.44(5)	88(13)	32/87/170	E2		
565.7 + 566.7 ^e + 569.1 ^e	3.5(7)	30(50)	31/82/160	(E1)		
573.4	0.76(16)	170(40)	30/81/160	<i>M</i> 1		
638.05	≡62.4	$\equiv 62.4$	62.4	$E2^{d}$		
$649.5^{f} + 650.4$	3.3(4)	55(7)	23/60/110	E2		
791.1 ^f + 792.3	2.3(3)	36(4)	15/38/70	E2		
880.3+884.6 ^e	2.3(3)	51(5)	12/30/54	<i>M</i> 1		
950.0	0.51(14)	41(12)	11/26/45	(<i>M</i> 1)		
952.0	0.04(2)	40(30)	11/26/45			
954.5	0.12(3)	8(3)	11/26/45	E 1		
957.4	0.12(3)	11(3)	11/26/46	E1		
1045.72	0.21(11)	12(6)	9/21/36	E 1		
1075.3	0.26(4)	31(5)	9/20/33	M1		
1207.2	3.2(4)			E0		
1233.0	0.013(3)	5(2)	7/15/24	E1		
1274.6	0.018(5)	7(2)	6/14/22	Eĺ		
1291.65	0.14(2)	6(1)	6/14/22	<i>E</i> 1		
1415.0	0.03(1)	7(3)	5/12/18	<i>E</i> 1		
$1430.5^{f} + 1430.51$	0.41(6)	12(2)	5/12/17	E2		
1453.62	0.74(10)	15(2)	5/11/17	(<i>M</i> 1)		

TABLE II. K-conversion coefficients for transitions in 3.6-h 150 Tb decay.

^a The necessary γ -ray data have been taken from Table I. ^b Reference 27.

^cAssumed to be E1 to resolve multiple peak.

^dMeasured to be E2 (Ref. 13); used to normalize electron and γ -ray data. ^eAssumed to be E2 to resolve multiple peak.

^f This transition is very weak.

Level energy (keV)	β ⁺ endpoint energy (keV)	Q (keV)	<i>I</i> (β ⁺ , EC) ^a (%)	log f ₀ t ^b	$\log f_1 t^1$
g.s.	3730(120) °	4752(120)	15.4(40)	7.8	9.5
638.05	3105(100)	4765(100)	25.1(30)	7.2	8.8
1134.35	2580(120)	4736(120)	2.8(6)	8.0	9.5
1288.4	2500(155)	4811(155)	1.5(3)	8.2	9.7
1518.5	2170(100)	4711(100)	3.3(4)	7.8	9.3
1700.1	2060(100)	4782(100)	1.8(3)	7.9	9.4
1955.0	1805(95)	4783 (95)	3.4(4)	7.5	8.9
2084.4	1665(110)	4771(110)	1.1(2)	8.0	9.4
2091.7	1655(80)	4769(80)	6.6(9)	7.2	8.6
2180.1	1590(110)	4792(110)	2.0(3)	7.7	9.1
	Average	4765(35)			

TABLE III. Endpoint energies and $\log ft$ values for β^* populating particular levels in ¹⁵⁰Gd from 3.6-h ¹⁵⁰Tb decay.

^aIntensity of β -decay branch, from Table IV.

^bCalculated using tables in Ref. 29.

^cSee text for details of assigning β^* branch to ground state.

Level	Level energy		$I(\beta^*, EC)^a$	
no.	(keV)	J [#]	(%)	Depopulating transitions ^b
1	0.0	0*	15.4(40)	None
2	638.05(10)	2*	25.1(30)	638.05[1]A
3	1134.35(15)	3-	2.8(6)	496.30[2] <i>A</i>
4	1207.2(2)	0*	0.8(2)	569.1[2]A, 1207.2[1]B
5	1288.4(2)	4*	1.5(3)	153.9[3]B, 650.4[2]A
6	1430.5(2)	(1,2)*	1.9(4)	792.5[2]A, 1430.51[1]A
7	1518.5(2)	2*	3.3(4)	384.1[3]A, 880.3[2]A, 1518.5[1]A
8	1592.7(2)	1-	1.7(3)	385.5[4]A, 954.5[2]A, 1592.7[1]A
9	1700.1(2)	(3, 4)*	1.8(3)	411.7[5]A, 565.7[3]A, 1061.9[2]A
10	1700.9(2)°	5	0.2(1)	412.3[5] <i>B</i> , 566.7[3] <i>B</i>
11	1814.3(3)	3-	0.5(1)	526.0[5]A, 1176.0[2]A
12	1947.5(3)	(2,3,4 ^d) ⁻	0.5(2)	813.1[3]A
13	1955.6(2)	2*	3.4(4)	437.1[7]A, 525.0[6]A, 748.3[4]A, 821.1[3]A,
10	100010(2)	-	0.1(1)	1317.7[2]A, 1955.7[1]B
14	1988.0(3)	(2,3,4 [●]) ⁺	1.0(4)	557.5[6]A, 699.4[5]A, 1350.1[2]A
15	2080.0(5)	$(2^*, 3^{\pm}, 4^{+})^{1}$	0.5(2)	
16	2084.4(2)	(2,3)	1.1(2)	649.5[6]A, 791.1[5]A, 1442.0[2]A
10	2091.7(2)	(2,3) 2*		491.7[8]A, 950.0[3]A, 1446.2[2]A
17	2091.7(2)	2	6.6(9)	573.4[7] <i>A</i> , 661.0[6] <i>B</i> , 884.6[4] <i>A</i> , 957.4[3] <i>A</i> ,
10	0400 4/0)	2*	0.0(0)	1453.62[2]A, 2071.7[1]A
18	2180.1(2)		2.0(3)	1045.72[3]A, 1542.0[2]A, 2180.1[1]B
19	2209.5(3)	(2,3°)-	0.7(1)	779.0[6]A, 1075.3[3]A, 1571.3[2]A
20	2262.4(3)		0.4(1)	743.8[7]A, 1624.4[2]A
21	2326.3(5)	41 m	0.9(2)	338.2[14]A, 378.8[12]A, 808[7]A, 895.9[6]A, 1688.2[2]A
22	2365.1(3)	(1 [±] , 2 ⁺)	1.4(2)	777.6[8]A, 1157.7[4]A, 1726.9[2]A, 2365.1[1]A
23	2408.8(3)	(2,3,4 °)⁺	0.8(1)	978.1[6]A, 1120.1[5]B, 1274.6[3]A, 1770.8[2]A
24	2416.9(5) ^g		0.5(1)	1778.8[2] <i>A</i>
25	2426.1(3)	2*	4.2 (5)	995.5[6]A, 1291.65[3]A, 1788.13[2]A, 2426.3[1]B
26	2521.8(5)	(2*, 3*, 4*) ^f	0.7(1)	1003.8[7]B, 1091.2[6]A, 1233.0[5]A, 1387.0[3]A, 1884.0[2]A
27	2559.0(3)	(1*, 2*)	0.1(2)	1351.9[4]A, 2558.9[1]B
28	2564.9(3)		0.5(1)	609.3[12]B, 1134.3[6]A, 1430.5[3]A, 1926.8[2]A
29	2627.4(5) ^s		0.11(4)	1493.1[3]A
30	2654.5(3)		1.1(2)	666.3[14]A, 1224.1[6]B, 2016.5[2]A
31	2678.6(3)	(1 * , 2*)	0.5(1)	2040.6[2]A, 2678.6[1]B
32	2687.2(3)	(1 ⁻ , 2 [±] , 3 ⁻)	1.0(2)	128.0[27]D, 602.2[16]A, 1094.4[8]A, 1168.7[7]A
				1256.6[6]B, 1552.7[3]B
33	2755.1(3)	(2*, 3*, 4*)	0.6(1)	1466.6[5]A, 1620.7[3]A, 2117.0[2]A
34	2786.9(4)	(1-, 2+)	1.5(2)	831.5[13]A, 1356.1[6]A, 1580.0[4]B, 1652.7[3]A, 2148.7[2]A
35	2828.4(5)		0.20(4)	1128.2[9]A, 2190.3[2]B
36	2845.8(4)	(1 [±] , 2 ⁺)	1.6(2)	1253.1[8]B, 1415.0[6]A, 1638.6[4]A, 2207.7[2]A, 2845.6[1]B
37	2956.4(3)		1.2(2)	128.0[35] B , 746.6[19] B , 968.4[14] A , 1001.0[13] A ,
				1525.8[6]A, 2318.2[2]A
38	2985.0(6)		1.3(3)	330.1[28]B, 425.9[27]A, 997.7[14]A, 1392.1[8]A, 1554.7[6]A
			(0)	1778.0[4]A, 2347.1[2]A, 2984.4[1]B
39	3035.6(10)	(1-, 2+)	2.2(4)	609.3[25]D, 952.0[16]A, 1443.6[8]A, 1516.5[7]A, 1605.6[6]A
	000000000	(-,-,	(I)	1900.6[3]A, $2396.5[2]A$, $3035.5[1]B$
40	3083.7(3) ^g		0.18(4)	1000.0[3]A, $2390.5[2]A$, $3035.5[1]B1003.8[15]D$, $1949.3[3]A$
40	3119.3(3)		0.18(4)	1688.8[6]A, 1985.0[3]A
41 42	3178.3(6)			
			0.7(1)	1660.2[7] A , 2044.3[3] A , 2539.5[2] A
43	3329.0(5)	(9+)	0.4(1)	2194.9[3]A, 2690.5[2]B
44	3344.8(5)	(2*)	1.0(2)	1135.0[19]A, 1253.1[17]D, 1752.2[8]A, 1914.4[6]A,
45			0.441	2056.4[5] A , 2137.1[4] B , 2706.6[2] A , 3344.7[1] B
45	3375.7(3)		0.4(1)	2241.3[3]A, 2737.5[2]A
46	3378.3(5)	()	0.3(1)	1564.2[1]A, 2740.0[2]B
47	3961(1)	(2*)	0.17(5)	2173.4[5] A , 3460[1] B
48	3510(1)	(1-, 2+)	0.4(1)	2376.5[3]A, 2872.4[2]A, 3509.5[1]B
49	3658.5(5) ^g		0.4(1)	1670.5[14]B, 1703.1[13]B, 3658.2[1]B
50	3726.9(6)		0.3(1)	2296.7[6]B, 2592.1[3]A
51	3829.7(10) ^s	(1 * , 2 *)	0.08(2)	2622.1[4]D, 3191.6[2]B, 3829[1]B
52	3840.6(5)		0.4(1)	330.1[48]D,2410.0[6]A,2552.3[5]B,3202.5[2]B

TABLE IV. Decay scheme for $3.6-h^{150}$ Tb.

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Level no.	Level energy (keV)	J [#]	$I(\beta^*, EC)^a$ (%)	Depopulating transitions ^b
53	4022(1) ^g	(1 [±] , 2 ⁺)	0.06(3)	3383.6[2]B,4022[1]B
54	4111(1) ^g	(1*, 2*)	0.11(3)	2296.7[11]D, 2976.2[3]D, 3473[2]B, 4111[1]B
55	4145(1) ^g	(1 ⁻ , 2 ⁺)	0.06(2)	1580.0[28]D, 2552.3[3]D, 3009.9[3]B, 4146[1]C
56	4165(1)	(2*)	0.12(4)	1955.7[19]D, 2876.6[5]A, 4164[1]C
57	4176.6(5)		0.3(1)	1191.1[38]A, 1811.8[22]B, 3042.6[3]B
58	4207(2)	(1*, 2*)	0.4(1)	1552.7[30]D, 1796.6[23]B, 2614.4[8]B, 2774.7[6]B, 3571[2]B, 4207[1]C
59	4237(1) ^g	(1 ⁻ , 2 ⁺)	0.07(2)	3102.6[3]B, 4236[1]C
60	4246(2) ^g	(1 *, 2*)	0.04(2)	4246[1] <i>C</i>
61	4258(1)	(1-, 2+)	0.19(6)	1831.9[25]B, 3123.9[3]B, 4257[1]C
62	4265(1)	(2*)	0.12(3)	153.9[54]D, 2976.2[5]B, 3130.5[3]B, 4265[1]C
63	4284(2) ^g	(1*, 2*)	0.01(1)	4284[1] <i>C</i>
64	4290(2) ^g	(1*, 2*)	0.10(3)	4290[1] <i>C</i>
65	4314(1)		0.14(3)	2240.0[16]D, 2233.7[15]B, 3675[2]B, 4314[1]C
66	4322(1)	(2*)	0.4(1)	300.4[53] B , 2230.0[17] B , 2622.1[9] B , 3034.0[5] B , 3685[2] B , 4322[1] C
67	4344(1)	(1*, 2*)	0.13(5)	1224.2[41]D, 2263.1[15]B, 2913.4[6]B, 4344[1]C
68	4379(1) ^g	(1*, 2*)	0.11(3)	1003.9[45]D, 2169.3[19]B, 2678.6[9]D, 4379[1]C
69	4406(1)	(1*, 2*)	0.13(3)	2449.8[13]B, 3197.7[4]B, 3769.2[2]B, 4406[1]C
70	4446(1)		0.20(5)	2498.0[12]B, 3238[4]B, 4446[1]C
71	4524(1) ^g		0.05(2)	3389.8[3]B, 3885[2]B
72	4530(2) ^g	(1 [±] , 2 ⁺)	0.06(2)	2449.8[15]D, 2936.0[8]B, 4531[1]C
73	4545.5(10)		0.20(4)	2137.2[23]D, 2180.0[22]D, 2283.3[20]B, 2460.7[16]B, 2845.6[9]D, 2952.7[8]B, 3411.3[3]B
74	4745.5(10)		0.4(1)	2565.5[18]B, 2661.0[16]B, 3152.4[8]B, 3315.0[6]B, 4107[2]E

TABLE IV. (Continued)

^aCalculated assuming 72(9)% of the decays result in a 638.05-keV γ ray.

^bNumber in brackets indicates the level fed by the transition. Letters indicate placement by A, coincidence data; B, energy sums; C, placed feeding the ground state in order to be consistent with the measured Q value; D, alternate placements for transitions placed by energy sums.

^c The existence of this level is well established from 5.8-min ¹⁵⁰Tb decay (Ref. 4). Evidence for its population in 3.6-h ¹⁵⁰Tb decay is weak.

^dAllowed only if there is no direct β feeding to this level.

^eAllowed only if 1430.5-keV level is 2⁺.

^f 3[•] and 4⁺ allowed only if 1430.5-keV level is 2^+ .

^gTentative level.

some cases, first-forbidden unique.³⁰

The $\beta^* - \gamma$ coincidence measurement indicates that the 0^{*} ground state and the 4^{*} level at 1288.4 keV are directly populated by β decay. The log $f_1 t$ values for these transitions are consistent with their being first-forbidden unique.³⁰ Therefore, the J^{*} of the 3.6-h¹⁵⁰Tb state is most likely 2⁻. If this assignment is correct, then one would not expect direct β feeding to either the 1700.9-keV 5⁻ level or the 1947.5-keV level if it is 4⁻.

The levels in ¹⁵⁰Gd proposed in previous ¹⁵⁰Tb decay studies^{5, 6, 14, 16, 17} are in general confirmed by the present results. The placement of a few transitions and some J^r assignments have been changed. The only other difference worth noting concerns the two levels at 1700 keV. Some confusion has resulted since a distinction between them has not been made in previous work. The

1700.9-keV 5⁻ level is observed in in-beam studies¹³⁻¹⁵ and in 5.8-min ¹⁵⁰Tb decay.^{4, 7, 8} It decays to the 1288.4-keV 4* and 1134.35-keV 3" levels via 412.3-keV E1 and 566.52-keV E2 transitions, respectively.¹⁵ In 3.6-h ¹⁵⁰Tb decay a 1700.3-keV $(3, 4)^+$ level is directly populated by β decay. This level is depopulated by a 411.7-keV M1 transition to the 1288.4-keV level and a 565.7keV (E1) transition to the 1134.35-keV level. Experimental evidence that the 566.52- and 565.7-keV transitions are not the same is seen in Fig. 1 where γ -ray spectra obtained at the same gain for 3.6-h and 5.8-min ¹⁵⁰Tb are compared. The observation of a weak (566.7 ± 0.3) -keV peak as a shoulder on the 565.7-keV peak in the spectrum for 3.6-h ¹⁵⁰Tb decay provides the only evidence for the population of the 1700.9-keV level in the decay of 3.6-h¹⁵⁰Tb.

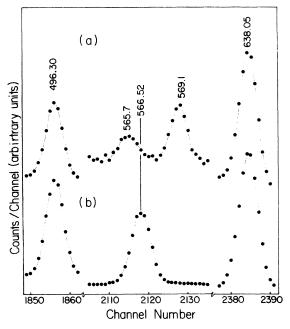


FIG. 1. γ -ray spectra from (a) 3.6-h and (b) 5.8-min ¹⁵⁰Tb decay. The gain in the two spectra is the same. The difference in energy between the peaks at 565.7 \pm 0.1 and 566.52 \pm 0.10 keV is clearly observed.

IV. DISCUSSION

A. Qualitative description of low-lying levels

The low-lying levels in 150 Gd are summarized in the level scheme in Fig. 2. Levels populated in the

decay⁴ of 5.8-min ¹⁵⁰Tb and those observed in (p,t) reaction studies⁹⁻¹¹ are also included. An in-beam γ -ray study¹⁵ using the $(\alpha, 4n)$ reaction found no additional states below 2100 keV. Of the 17 levels below 2100 keV, 15 can be described in terms of the vibrational model. The remaining two levels appear to have quasiparticle character.

A possible correspondence between the experimental levels and those predicted by the vibrational model³¹ is indicated on the left in Fig. 2. The onephonon energies of the quadrupole and octupole oscillators were assumed to be 638 and 1134 keV, respectively. Anharmonicities in the vibrational motion can be qualitatively taken into account by allowing the degenerate multiplets predicted by the model to be split. Candidates for all the quadrupole three-phonon states and at least four of the five quadrupole-octupole two-phonon states are observed in the data.

Experimental B(E2) ratios are compared with the predictions of the vibrational model in Table V. Anharmonicities in the vibrational motion will be manifested by the occurrence of multiphonon E2transitions which are strictly forbidden in the model. From the ratios listed in Table V, two-phonon transitions are hindered at least by a factor of 22 compared to one-phonon transitions, while the one three-phonon transition which was observed is hindered by a factor of ~1600. Ratios of one-phonon transitions do not show such large hindrances and are in reasonable agreement with the model predictions.

Two-phonon E0 transitions are permitted by the

TABLE V. Comparison of experimental B(E2) ratios with the predictions of models discussed in the text.

Initial level ^a	Final levels ^b		Vibrational	Phenomenological	Microsco	pic model ^e
$E, N, J^{\texttt{T}}$	$[E, \Delta N, J^{\dagger}]_1 / [E, \Delta N, J^{\dagger}]_2$	Exp.	model ^c	model ^d	I	II
1430, 2, (2)*	[0,2,0*]/[638,1,2*]	0.026(2)	0.0	0.028	0.14	0.31
1955, 3, 2*	$[0, 3, 0^{+}]/[1207, 1, 0^{+}]$	0.0006(3)	0.0	0.0012	0.32	0.067
	[638,2,2*]/[1207,1,0*]	0.045(9)	0.0	0.026	1.03	0.48
	$[1430, 1, (2)^*]/[1207, 1, 0^*]$	7(1)	0.41	0.64	6.99	
1988, 3, (3 ⁺ , 4 ⁺) ^f	$[638, 2, 2^*]/[1430, 1, (2)^*]$	0.04(1)	0.0	0.038	0.024	0.17
	$[1288, 1, 4^{+}]/[1430, 1, (2)^{+}]$	0.29(4)	0.40	0.66	0.16	0.13
2080, 3, (3 ⁺ , 4 ⁺) ^g	$[638, 2, 2^*]/[1430, 1, (2)^*]$	0.005(3)	0.0	0.0028	1.06	0.068
	[1288, 1, 4 ⁺]/[1430, 1, (2) ⁺]	0.33(22)	0.91	0.46	0.82	0.18

^aN is the number of phonons assigned to the level in the vibrational model.

 ${}^{b}\Delta N$ is the change in phonon number for a transition.

^cReference 31.

^dReference 19.

^eReference 20. Columns labeled I and II represent different ways of correlating calculated levels with experimental levels. See text for details.

 $^{\rm f}$ Theoretical values are calculated assuming state is 4^{\star} except for column labeled II where 3^{\star} is assumed.

^gTheoretical values calculated assuming state is 3^* except for column labeled II where 4^* is assumed.

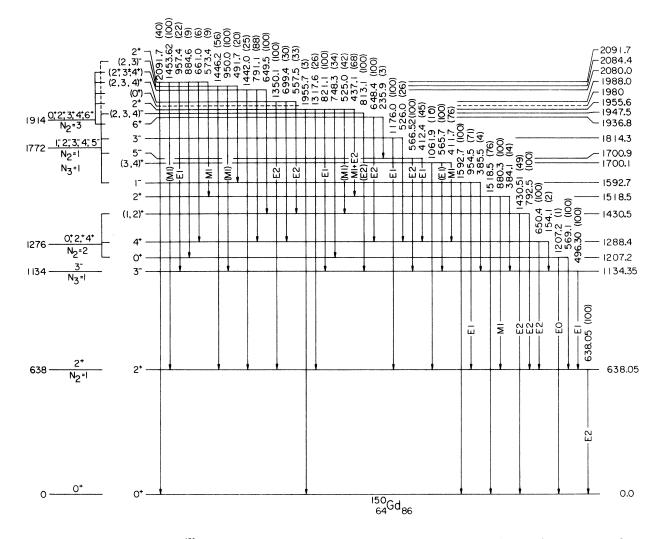


FIG. 2. Low-lying levels in 150 Gd. In addition to the results from the present study, data from Refs. 4, 9–11, and 15 have been included. The relative transition intensity for branching out of each level is given in parentheses. A possible correspondence with the predictions of the vibrational model is indicated on the left.

vibrational model. The experimental value of the X parameter [B(E0)/B(E2)] for the 1207.2-keV level is 0.020(3). The predicted value is 0.023. The theoretical X value, however, is only an estimate based on the experimental $B(E2; 2^+ \rightarrow 0^+)$ for ¹⁴⁸Sm (Ref. 32). The corresponding value for ¹⁵⁰Gd has not been measured.

The vibrational model describes the low-lying states in ¹⁵⁰Gd surprisingly well, especially when allowances are made for anharmonic effects in the vibrational motion. This vibrational description does not, however, account for two of the experimental levels below 2 MeV. These levels at 1518.5 and 1700.3 keV occur at an energy which is intermediate between the quadrupole two- and three-phonon states.

The 1518.5-keV 2⁺ level is only 88 keV above the 2⁺ level at 1430.5 keV which is assigned as the twophonon state. The assignment is based on the deexcitation characteristics of these levels to the 638.05-keV 2⁺ level. The 880.3-keV transition (1518.5 \rightarrow 638.05) is predominantly $M1 (\delta_{E2/M1}^{2} < 0.5)$ while the 792.3-keV transition (1430.5 \rightarrow 638.05) is mostly $E2 (\delta_{E2/M1}^{2} > 15)$. These limiting values of mixing ratios have been estimated from the α_{K} values in Table II. Since M1 transitions are forbidden in the vibrational model, the 1430.5-keV level is regarded as the more likely candidate to be a vibrational state. From the available data it is not possible to assign a particular structure of the 1518.5-keV level. The multipolarity of the 880.3-keV transition, however, suggests that it has a large admixture of quasiparticle components. A collective model would not be expected to account for such a state.

The 411.7-keV transition which depopulates the $1700.3 \text{-} \text{keV} (3, 4)^+$ level to the $1288.4 \text{-} \text{keV} 4^+$ level also has a large M1 component ($\delta_{E2/M1}^2 < 1.8$). A large error is associated with our value of α_{κ} for the 411.7-keV transition. If the data of Vylov et al.^{5,6} are used, $\delta_{E2/M1}^2$ is <0.57. The M1 component in the 411.7-keV transition suggests that the 1700.3-keV level also has a large admixture of guasiparticle components. Furthermore the γ decay patterns for the 1518.5- and 1700.3-keV levels are similar as can be seen in Fig. 2. Transition probabilities are summarized in Table VI. If these levels were decaying without preference to the various final states, then the experimental values of B(E1)/B(E2) and B(M1)/B(E2) ratios in Table VI should approach the values for singleparticle transition probabilities. These singleparticle ratios are listed in Table VI for comparison.

Low-lying quasiparticle states have been proposed to explain the extra 0⁺ and 2⁺ states found in in the vibrational Cd isotopes.³³ These extra states are also at an energy between the quadrupole two- and three-phonon states. The transitions from the vibrational and quasiparticle 2⁺ states to the first 2⁺ state in ¹¹⁴Cd, for example, are similar to the 792.3- and 880.3-keV transitions in ¹⁵⁰Gd in that they have $\delta_{E2/M1}^2$ values of 1.96 and 0.0025, respectively.³⁴

For the purposes of the following discussions we shall assume that the 1518.5- and 1700.3-keV levels are quasiparticle states of an unspecified structure and that they are relatively unmixed with nearby vibrational states.

B. Systematics

Level schemes of the even-even Gd isotopes with 82 to 94 neutrons are shown in Fig. 3. To facilitate comparisons the levels have been arranged in Sakai's quasirotational band representation.⁴³ Only the ground state, β , γ , and $K=0^{-}$ octupole quasibands are shown. In the present discussion it is convenient to make no distinction between rotational and quasirotational bands.

The effect of adding pairs of neutrons above the 82-neutron closed shell is seen in Fig. 3. Lowlying levels in the single-closed-shell nucleus ¹⁴⁶Gd have been described in terms of quasiproton TABLE VI. B(E1)/B(E2) and B(M1)/B(E2) ratios for the deexcitation of the 1518.5- and 1700.3-keV levels compared with single-particle values.

Level energy (keV)	B(E1)/B(E2) (fm ⁻²)	$\frac{B(M1)/B(E2)}{\left[\left(\frac{e\hbar}{2Mc}\right)^2 / e^2 \mathrm{fm}^4\right]}$
1518.5	2.1(2)×10 ⁻⁵	1.1(1)×10 ⁻³
1700.3	6(3)×10 ⁻⁵	$1.0(5) \times 10^{-2}$
Single particle	7.2×10^{-3}	3.8×10^{-2}

excitations.³⁵ The quasiground band in ¹⁴⁸Gd is similar to that in ¹⁴⁶Gd in that the spacing of the levels decreases with spin. In other 84-neutron nuclei such structure has been attributed to the coupling of two $2 f_{7/2}$ neutrons to an 82-neutron core.⁴⁴ The addition of two more neutrons produces the vibrational ¹⁵⁰Gd. A spherical-to-prolate shape transition occurs between six and eight neutrons beyond the closed shell. Both ¹⁵²Gd and ¹⁵⁴Gd can be considered to be transitional, the former being spherical but easily deformed while the latter is weakly deformed but soft to β and γ vibrations. The Gd isotopes with more than 90 neutrons have level schemes characteristic of stably deformed nuclei.

The quasibands in ¹⁵⁰Gd appear to be a systematic extension of bands found in the heavier Gd isotopes. The spacing of the states in each of the bands shows a progression from rotational to vibrational values as the shape transition region is crossed. Even the odd-even staggering of states in the quasi- γ band in ¹⁵²Gd is preserved in ¹⁵⁰Gd.

The quasi- β and quasi- γ bandhead energies also follow predictable trends. They are at a minimum in the soft transitional nuclei and increase in energy as the isotopes take on more rigid spherical or spheroidal shapes. A drop in energy of the one-phonon octupole vibrational state (3⁻ member of the quasioctupole band) has been noted in the spherical nuclei above 82 neutrons. This has been correlated with the availability of negative-parity states $\nu(f_{7/2}i_{13/2})$ at relatively low excitation energies.⁴⁵

The trends observed in these Gd isotopes are reasonably well understood and clearly show the systematics of the 88- to 90-neutron shape transition. It would be interesting to compare the adding of neutrons above the 82-neutron closed shell with the corresponding proton case. The quasiground band levels in ²⁰⁴Po (Ref. 46) and ²⁰⁶Rn (Ref. 47) show a spacing which is indicative of coupling $2h_{9/2}$ protons to a core.

In Fig. 4 the level schemes of the even-even 86neutron isotones are compared with that for 150 Gd.

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•			<u>9-</u>	<u> -</u> <u>2</u> +	l <u>4</u> ⁺		
	<u>4</u> +	<u>7-</u>					
		<u>_</u>					
			<u>7-</u>	<u>9-</u> 1 <u>0+ 7+</u>	int int		
			<u> </u>	<u>8</u> *	1 <u>2+</u> 10+		
2	-	at	<u>6+</u> 2+ <u>3+</u>	<u>7</u> <u>6</u> ⁺ <u>5</u> ⁺		I <u>2⁺</u>	-
\sim		<u>5-</u> <u>6+</u>	<u>5</u> -	<u>8</u> ⁺ <u>6</u> ⁺	<u>8</u> +		4+
Me	<u>3" 2</u> *		<u>5-</u> <u> -</u>	<u>ب</u> م+	<u>10</u> + <u>6</u> +	5+	$\frac{5^{-}}{2^{+}} \frac{4^{-}}{2^{+}} \frac{6^{+}}{2^{+}}$
E _x (MeV)		<u>4</u> +	<u>2</u> +		<u>5- 6+ ⁵⁺</u>	$\begin{array}{c} 5^{+} \\ 5^{-} \\ 5^{-} \\ 3^{-} \\ 3^{-} \\ 1^{-} \\ 1^{-} \end{array} \begin{array}{c} 5^{+} \\ 4^{+} \\ 3^{+} \\ 2^{+} \\ 2^{+} \\ 2^{+} \end{array}$	
		<u>3-</u>	<u>3-</u> ^{4⁺} <u>0</u> ⁺	$\frac{1}{3} - \frac{6^{+}}{2} + \frac{4^{+}}{2}$	$\frac{5}{3} = \frac{6^{+}}{3} \frac{5^{-}}{4^{+}}$	- 4 - 4 - 2 - 2 - 2 - 2	느 포
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		-	<u>2</u> +	<u>4</u> * 	<u>6</u> ⁺ <u>0</u> ⁺	c†	
			_			<u>6</u> +	<u>6</u> †
				<u>2</u> +	<u>4</u> +	4+	4†
					<u>2</u> *	2+	2+
0	- <u>0</u> +	0 0 ⁺	0 G β γ	0 0 ⁺ βγ	<u>2</u> + Ο Ο G β γ	4 ⁺ 2 ⁺ Ο G β γ	4 [†] 2 ⁺ Ο ⁺ Ο G β γ
	¹⁴⁶ Gd	¹⁴⁸ Gd	ο ͼ _{βγ} ¹⁵⁰ Gd	οςβγ ¹⁵² Gd	ος βγ ¹⁵⁴ Gd	οςβγ ¹⁵⁶ Gd	ος β.γ ¹⁵⁸ Gd
	Ga	Ga	Ga	Gd	····Gd	Gd	Gd

FIG. 3. Quasirotational bands in the even-even Gd isotopes. The quasigroundstate (G), quasi- β , quasi- γ , and K = 0⁻ quasioctupole (O) bands are shown. In addition to the results of the present study, data are also taken from Refs. 35, 36 (¹⁴⁶Gd); 37 (¹⁴⁸Gd); 4, 15 (¹⁵⁰Gd); 22, 38 (¹⁵²Gd); 39, 40 (¹⁵⁴Gd); 41 (¹⁵⁶Gd); and 42 (¹⁵⁸Gd).

Most of these nuclei have not been studied in detail; thus comparisons can be made for only a few levels.

The energy of the first 2^* state is at a maximum in ¹⁵⁰Gd while the energy of the first-excited 0^* state (quasi- β bandhead) drops ~ 0.5 MeV from ¹⁴⁶Nd to ¹⁵⁰Gd. The energy of the second 2^* state (quasi- γ bandhead) remains approximately constant. The first excited 4^* and 6^* states follow the trends of the first 2^* state.

From a survey of positive-parity two-phonon states in even-even nuclei from Zn to Te, Hadermann and Rester⁵² suggest that a drop in the twophonon 0⁺ energy coupled with a rise in the onephonon 2⁺ energy can be correlated with the filling of spherical proton or neutron orbitals. There is some evidence from α -decay studies⁵³ that the spherical 64-proton nuclei have an extra stability which could be associated with the filling of the $1g_{7/2}$ and $2d_{5/2}$ proton orbitals. The systematics of the positive-parity states in the 86-neutron isotones appear to provide further evidence of orbital filling at 64 protons.

The systematics of the negative-parity states may also be understood in the same way. The en-

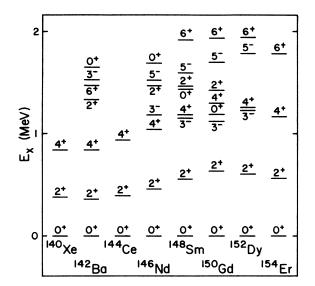


FIG. 4. Level schemes of the even-even 86-neutron isotones from Xe to Er. In addition to the results of the present study, data are also taken from Refs. 48 (140 Xe, 144 Ce); 43 (142 Ba); 36 (146 Nd); 49 (148 Sm); 4, 15 (150 Gd); 50 (152 Dy); and 51 (154 Er).

ergy of the first 3⁻ state also reaches a minimum at ¹⁵⁰Gd. A similar trend has been noticed in the 82- and 84-neutron isotones and ascribed to the lowering in energy of the $\pi(d_{5/2}h_{11/2})$ and $\pi(g_{7/2}h_{11/2})$ states as the spherical $1g_{7/2}$ and $2d_{5/2}$ orbitals are filled.⁴⁵ The spacing of the 5⁻ level above the 3⁻ level suggests that it is possibly a quadrupoleoctupole two-phonon vibration.

It is obvious from Fig. 4 that not all the 86neutron isotones are examples of vibrational nuclei. From Xe to Nd the isotones exhibit a more transitional character as evidenced by the relatively high energies of the quasi- β and quasi- γ bandheads with respect to the 4⁺ member of the ground band. The data for ¹⁵⁴Er suggest a return to a more transitional character in the heavier isotones. Wilhelmy et al.48 have drawn a similar conclusion from the analysis of energy ratios of the first 4^+ and 2^+ states for nuclei in this region. The best candidates for vibrational nuclei among the 86-neutron isotones appear to be those nearest the orbital closure at 64 protons. It would be interesting to see whether ¹⁴⁸Sm and ¹⁵²Dy can be described within a vibrational framework.

C. Comparison with theory

1. Phenomenological model

The experimental data for ¹⁵⁰Gd have already been compared with the predictions of the phenomenological vibrational model. The agreement was remarkably good considering the simplicity of this model. For a more quantitative description of the data a phenomenological model which includes anharmonicity in the vibrational motion is required.

One such model for quadrupole collective motion is based on a Hamiltonian for an anharmonic quadrupole oscillator. All terms are kept which do not violate time reversal or rotational invariance. Gneuss and Greiner¹⁹ have described this model in detail and demonstrated that it is capable of reproducing collective level spectra in a wide variety of nuclei. For the calculation performed here, the Hamiltonian contained only five free parameters, since the potential energy terms above fourth order were not used. In all other respects the calculation is equivalent to the one described by Gneuss and Greiner.

The experimental levels in ¹⁵⁰Gd and those predicted by this model are compared in Fig. 5. The parameters P_2 , P_3 , C_2 , C_3 , and C_4 [see Eq. (2), Ref. 19] for this calculation were 0.00487, 0.0193, 16.62, 32.55, and 91.38 MeV, respectively. These parameters were not determined from a leastsquares fit to the data. Undoubtedly better results could be obtained if such a procedure were used.

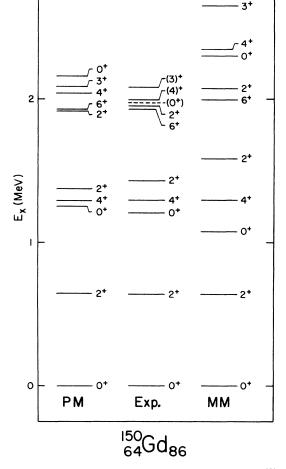


FIG. 5. Comparison of the experimental levels in ¹⁵⁰Gd with those predicted by the phenomenological model (PM) of Gneuss and Greiner and the microscopic model (MM) of Kishimoto and Tamura.

The calculation reproduces all but the three-phonon 0^* state reasonably well.

An ambiguity exists in attempting to correlate the experimental and theoretical results since the J^{r} values for the 1988.0- and 2080.0-keV levels are not uniquely determined. The calculation indicates that they should be 4⁺ and 3⁺, respectively, while the systematics discussed above suggests the opposite ordering. Unfortunately, either ordering is permitted by the currently available data. Until this question can be experimentally resolved, the J^{r} values for these states will be assumed to be those which are most consistent with what is being compared. This ambiguity also exists in comparisons with the microscopic model in the next section.

Transition probabilities were calculated under the assumption that the $B(E2; 2^* \rightarrow 0^*)$ value for ¹⁵⁰Gd is the same as that for ¹⁴⁸Sm (Ref. 32). Experimental and calculated B(E2) ratios are shown in Table V. The agreement is good. Except for one value all the calculated ratios are within a factor of 2.3 or less of the experimental ratios.

2. Microscopic model

Two different microscopic models have been used to describe the 88- to 90-neutron shape transition in the Sm isotopes. Kumar and Baranger^{54, 55} employ the microscopic pairing-plus-quadrupole model to determine the parameters of Bohr's collective Hamiltonian. Kishimoto and Tamura²⁰ start with a fermion Hamiltonian having monopole pairing, quadrupole pairing, and quadrupole particle-hole residual interactions. The parameters for a collective boson Hamiltonian are obtained through the use of boson-expansion techniques. A major difference in the two models results from the way in which the coupling to noncollective modes of excitation is treated. Kumar and Baranger adopt the adiabatic assumption while Kishimoto and Tamura include this coupling in an approximate way. Since the effects of this coupling should be important in describing vibrational nuclei, the boson-expansion model may be more appropriate for ¹⁵⁰Gd.

The calculation which has been made is essentially the same as that described in detail in Ref. 20. The two adjustable parameters of the model representing the strengths of the quadrupole pairing and quadrupole particle-hole interactions were taken to be 0.65 and 0.765, respectively, in units of $240A^{-5/3}$ MeV. The parameters were chosen to reproduce the energies of the first 2⁺ and 4⁺ states in ¹⁵⁰Gd.

The calculated results may be compared with experiment in two ways. In the first method the correspondence between calculated and experimental levels is based on energies, spins, and parities. For example, the 2^+_3 state observed at 1955.6 keV is to be compared with the 2⁺₂ state calculated to occur at 2200 keV. This is shown in Fig. 5. The experimental spin sequence is reproduced if the 1988.0- and 2080.0-keV levels are assumed to be 4^+ and 3^+ , respectively. The average deviation of the experimental and theoretical levels is ~ 190 keV. For the phenomenological model the average deviation was ~ 30 keV. Calculated B(E2)ratios for this method of correlating the calculated and experimental results are given in Table V in the column labeled I. The agreement is relatively poor.

Alternatively the calculated and experimental results can be compared on the basis of quasirotational bands. The grouping of states into bands is based on B(E2) values and values of static quad-

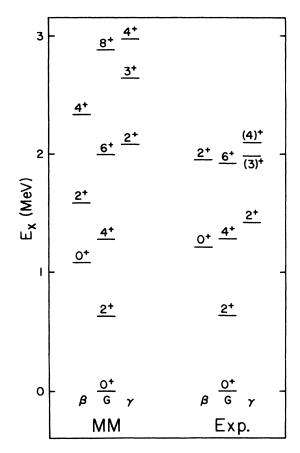


FIG. 6. Comparison of experimental quasirotational bands for ¹⁵⁰Gd with those predicted by the microscopic model (MM) of Kishimoto and Tamura.

rupole moments.²⁰ In this case the 2_3^* state at 1955.6 keV is to be compared with the calculated 2^* state at 1600 keV which is a member of the quasi- β band. Energy spectra are shown in Fig. 6 for this representation and in Table V for B(E2)ratios (column labeled II). In this case, the 1988.0- and 2080.0-keV levels are assumed to be 3^* and 4^* , respectively. The calculated quasi- β band is found to be too low in energy while the quasi- γ band lies too high. The spacing of the states within each band is reasonably consistent with experiment and there is even a slight oddeven bunching of the quasi- γ band states. However, the calculated B(E2) ratios again do not reproduce the data very well.

The boson-expansion model does not satisfactorily reproduce either the experimental level spectrum or B(E2) ratios. The model predicts ¹⁵⁰Gd to be more transitional than the data would imply. This is observed in three ways. First, if one compares the relative positions of the excited quasibands with respect to the quasiground band states,

Initial level ^a	Final levels ^a	¹⁵⁰ Gd	¹⁵² Gd ^b	¹⁵⁴ Gd ^c	Theo.
2	0 _g /2 _g	0.013(6)	0.0192(19)	0.121(4)	0.139
2 ₈	$4_{g}/2_{g}$		2.04(27)	2.72(8)	1.78
	$0_{\rm g} / 0_{\rm g}$	1670(840)	107(11)	125(6)	15.0
4 ₈	$2_{g}^{2}/4_{g}^{2}$ $6_{g}/4_{g}$ $0_{g}/2_{g}$			0.0855(25)	0.148
٢	$6_{,}/4_{,}$			5.91(18)	2.45
2 ₇	$0_{r}/2_{r}$	0.026(2)	0.142(14)	0.464(11)	0.307
•	$4_{,}/2_{,}$			0.145(5)	1.74
3,	$\frac{4_g}{2_g}$	0.14(4)	0.45(5)	1.032(31)	1.28
,	$2_{y}/2_{g}$	25 (6)	<28	16.5(8)	5.83
4 ₇		0.015(10)	0.071(15)	0.138(7)	0.377
,	2 _g /4 _g 2 _y /2 _g	200(120)	48(21)		14.8

TABLE VII. Experimental B(E2) ratios for ^{150,152,154}Gd and predictions of the microscopic boson-expansion model for ¹⁵⁰Gd.

^a Experimental levels are labeled by their quasirotational band assignments.

^bReference 22.

^cReferences 22 and 39.

then the calculated spectrum for ¹⁵⁰Gd resembles the experimental spectrum for the transitional nucleus ¹⁵²Gd (see Fig. 3). Second, in Table VII, experimental B(E2) ratios for ^{150,152,154}Gd are compared with the model predictions. Many of the ratios calculated for ¹⁵⁰Gd are more consistent with the experimental values for the transitional nuclei (particularly¹⁵⁴Gd) than with those for ¹⁵⁰Gd. Third, the potential wells from the phenomenological and microscopic calculations are compared in Fig. 7. To make this comparison more meaningful, the Hamiltonian for each model has been transformed so that all anharmonicities are included only in the potential energy terms. A prescription for this canonical transformation is given in Appendix I of Ref. 20. One would expect that the potential well for a vibrational nucleus would be centered around $\beta = 0$. While neither well exhibits this property, the well for the phenomenological model is less asymmetric about $\beta = 0$ than the microscopically calculated well.

These comparisons do not make clear whether a basic problem exists with the boson-expansion model or whether the problem lies in some of the "fixed" parameters (e.g., single-particle energies). In order to describe the level scheme for ¹⁵⁰Gd with a microscopic model it may be vital to include the effects of the extra states at 1518.5 and 1700.3 keV. At present such a model is not available. Kishimoto and Tamura have indicated that noncollective states may be explicitly included within the framework of the boson-expansion model and some preliminary results from such a calculation have been reported.³³ It would be very interesting to see if this model could reproduce the experimental data for ¹⁵⁰Gd.

3. Negative-parity states

Theoretical models for describing the negativeparity states resulting from the coupling of quadrupole and octupole vibrations are not well developed. One reason for this lies in the fact that such models must first adequately describe the quadrupole collective states before treatments of octupole vibrations or the quadrupole-octupole

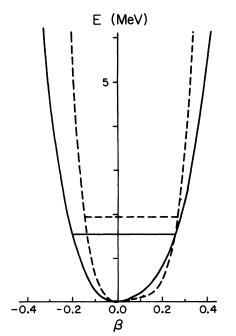


FIG. 7. Potential wells from the phenomenological (solid line) and microscopic (dashed line) models for $^{150}\mathrm{Gd}.$

coupling can be trusted. Furthermore not much experimental data are available for such states.

Vogel and Kocbach⁵⁶ have developed a microscopic model for calculating the two-phonon states resulting from the coupling of one quadrupole and one octupole vibration. Basically this model uses techniques similar to those developed by Kumar and Baranger^{54, 55} for describing quadrupole collective states. Numerical results have been reported for ¹⁴⁶Nd and ¹⁴⁸Sm. In general the energies of the calculated levels were too high. This was attributed in part to an inadequate treatment of the quadrupole motion in the model.

Calculations for the negative-parity states in ¹⁵⁰Gd will not be presented here. A comparison of the experimental data with the predictions of a phenomenological quadrupole-octupole coupling model³⁸ can be found in a discussion of the (α, xn) reaction¹⁵ in which negative-parity states are populated to much higher spins.

V. CONCLUSIONS

The decay of $3.6-h^{150}$ Tb is found to populate a large number of low-spin levels in 150 Gd. The levels below ~2 MeV have properties which strongly resemble those expected for a vibrational nucleus. Two of the low-lying levels which are not accounted for as vibrational states may have large

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admixtures of quasiparticle components. The proposed vibrational levels in ¹⁵⁰Gd can be reproduced with the phenomenological model of Gneuss and Greiner¹⁹ but the microscopic boson-expansion model of Kishimoto and Tamura²⁰ predicts ¹⁵⁰Gd to be more transitional than the data indicate.

The vibrational interpretation of the levels in ¹⁵⁰Gd fits well into the systematics of the 88- to 90-neutron region in the Gd isotopes. The systematics of the known 86-neutron isotones are found to be influenced by the closure of the spherical $g_{7/2}$ and $d_{5/2}$ proton orbitals at Z = 64. The isotones near the orbital closure are vibrational while those farther away are more transitional.

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