Decay of ²⁰Na to y-ray emitting states of ²⁰Ne[†]

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²⁰Ne γ rays following ²⁰Na decay identified new branches to states at 9.87, 10.88, and 11.26 MeV, as well as known branches to states at 1.63 and 10.27 MeV. Branching ratios and Ft values based on these observations are compared to nuclear structure calculations, and the β^{\pm} *Stasymmetry for A = 20 is discussed.*

RADIOACTIVITY ²⁰Na [from ²⁰Ne(p, n), E = 22.9 MeV]; measured E_γ , I_y; deduce
 β^+ branching ratios, $\mathfrak{F}t$ values, ²⁰Ne levels J^{*}, mirror β^+ asymmetry; enriched target; Ge(Li) detector.

I. INTRODUCTION

The positron decay of 20 Na to excited states of ²⁰Ne is important for questions of mirror β^* -decay symmetry, weak vector current conservation (CVC}, isospin purity of the lowest isobaric analog state in 20 Ne, and for comparison with nuclearlog state in ²⁰Ne, and for comparison with nucl
structure calculations.^{1–14} This article report an experiment on ²⁰Na β ⁺ delayed ²⁰Ne γ rays folan experiment on that p delayed the p rays for-
lowing the ²⁰Ne(*p*, *n*) ²⁰Na production reaction. The original intent in this experiment was to improve the observational basis of an indirect computation of the ²⁰Na(β ⁺) ²⁰Ne(1.63 MeV) branching ratio that is of interest for the mirror β^* -decay comparison. When the systematic comparison of allowed mirror β^{\dagger} decays was first made by Wilkin- $\sin \theta$ is a matrice θ and $\sin \theta$ is the matrix $\sin \theta$ where $\sin \theta$ eter measurement of this branching ratio¹⁵ and to son,¹ it was necessary to reject a β-ray-spectrom
eter measurement of this branching ratio¹⁵ and to
calculate it indirectly,^{1,2} using the total decay rate and subtracting the decay to α -emitting 2⁺ states, which were normalized to an assumed strength for the superallowed branch. Allowed transitions to 1^* and 3^* states were unknown and could not have been discovered by measurements of delayed α emission, which is forbidden by conservation of angular momentum and parity. No branches to γ -ray emitting states other than ²⁰Ne(1.63 MeV) were known, although experimental branching ratio limits of $1-2\%$ ^{15, 16} still allowed significant new branches and γ -decay competition from known α -emitting states. γ competition of 10% in the isospin-inhibited α decay of ²⁰Ne(10.27 MeV, $T=1$), for example, would change the indirectly computed rate of the important $^{20}Ne(1.63 \text{ MeV})$ branch by an amount comparable to the uncertainty from other sources, but would correspond to a 20 Na decay branching ratio of only 0.3% for the 8.64- MeV deexcitation γ ray, well below those experidecay branching ratio of only 0.3% for the
MeV deexcitation γ ray, well below those experimental limits. Recent experiments, $2^{-5.7-9}$ mental limits. Recent experiments, 2^{-5} , 7^{-9} includ-

ing an observation of this latter transition, $7^{1.8}$ have greatly clarified matters. Early measurements^{15, 16} implied a 50% inhibition of the superallowed transition to ²⁰Ne(10.27 MeV, $T = 1$), a large violation of isospin symmetry, and a small half-life, but they are contradicted by recent results. The 20 Na half-life and the 20 Na- 20 Ne mass difference are now well determined as 445.7 ± 2.9 msec^{3,7-9,12} and 13.892 ± 0.007 MeV,³ respectively. The relative strengths of all α -emitting states and their ratio to the strong 1.63-MeV delayed γ transition are also well determined.⁸ Direct measurements now indicate that the superallowed transition has a strength consistent with $CVC⁷$ and that the St values of the mirror β^{\pm} decays of ²⁰Na and ²⁰F to values of the mirror β^{\pm} decays of ²⁰Na and ²⁰F to ²⁰Ne(1.63 MeV) differ by only a few percent.^{7,8,12}

Results from the present work on 20 Na decay to γ -emitting states allow an improved determination of $\Gamma_{\gamma}/\Gamma_{\alpha}$ for ²⁰Ne(10.27 MeV, $T=1$). This has already been used¹⁷ in connection with a study of the ¹⁶O(α , γ) ²⁰Ne reaction to limit the isospin impurity in that state, and to predict the weak magnetism strength in the 20 Na and 20 F decays, which can be used in a test of CVC and in studies of secondclass-current effects in β decay. The present work also updates the comparison of the directly and indirectly determined $\mathfrak{F}t$ values for the superallowed and first-excited-state transitions. A preallowed and first-excited-state transitions. A _I
liminary account of this work has been given.¹⁸

Il. EXPERIMENT

²⁰Na activity was produced by the ²⁰Ne(p, n)²⁰Na reaction. Isotopically enriched (99.95%) ²⁰Ne target gas¹⁹ was contained in a 1.3-cm long, 0.6 -cmdiam. brass-walled target volume between $13-\mu m$ Ta foil beam windows. Natural abundance neon, which contains 22 Ne, was avoided as target gas to eliminate an 11-sec component in the 1.63-MeV γ ray which would arise from ${}^{22}Ne(\rho, {}^{3}He) {}^{20}F(\beta^-)$ -

FIG. 1. High-energy portion of the ²⁰Ne γ -ray spectrum from 20 Na decay. The labels identifying observed 20 Ne transitions are used also in Table I and in the text.

 20 Ne(1.63 MeV) reactions and which would yield incorrect relative intensity measurements unless separated.⁹ After 460-msec bombardments by 22.9-MeV protons from the University of Colorado cyclotron, the target "rabbit" was repeatedly shuttled 6 m to a shielded Ge(Li) γ -ray detector and, starting 400 msec after bombardment, four successive 320-msec, 1024-channel energy spectra were accumulated. A 5.2-sec bombard-count cycle with $3-\mu A$ bombardment currents was maintained for 10 h. Besides the strong 1.63-MeV γ ray, several much weaker high-energy γ rays were also observed, and all may be attributed to reactions of protons and 20 Ne, as none were present in empty-target experiments. Because of background from activity induced in the Ta foils and brass walls of the target rabbit, this experiment is not sensitive to weak γ rays with energies below ⁵ MeV. Figure I shows the high-energy region of the summed energy spectrum, with

markers indicating the observed transitions. Two γ rays were strong enough for meaningful halflife determinations, and these agree with that of the strong 1.63-MeV line. Energies and intensities of the observed γ rays relative to the 1.63-MeV line are given in Table I and are discussed below. The indicated uncertainties arise about equally, for the strongest lines, from the statistics of the γ -ray spectrum and from the uncertainties in the relative detection efficiency for low and high energies. The intensities have been used to determine branching ratios BR and $\mathfrak{F}t$ values for transitions to states of 20 Ne, as given in Table II, and illustrated in Fig. 2.

A. Detector configuration

 γ rays were detected with a commercial closedend coaxial lithium-drifted germanium [Ge(Li}] $\rm{detector}^{20}$ of approximately 34-cm³ active volume. During counting periods the distance between the Ge(Li) crystal face and the source center was 13.0 cm. γ rays from the ²⁰Na source passed through a collimator, a 5.1-cm-thick Lucite β -ray absorber, and a 2.4-cm-thick Pb hardener before reaching the detector. The collimator (a 1.9-cmdiam hole in a 2.5-cm-thick Pb plate near the source) reduced the incidence of positron annihilation quanta arising outside the source, without affecting direct radiation from the source. The (low-atomic-number) Lucite absorbed the intense high-energy β rays from the source with minimal production of high-energy bremsstrahlung background. Finally, the Pb hardner preferentially attenuated low-energy γ rays, and allowed the use of higher bombardment currents and higher counting rates for the high-energy γ rays of interest.

Energy ^a (keV)	Assignment in ²⁰ Ne		Label ^b	Relative intensity ^c
1633 ± 2	$1633 - 0$	$2^+ \rightarrow 0^+$		1.000
8240 ± 5 ^d	$9873 \rightarrow 1633 \quad 3^{+} \rightarrow 2^{+}$		D	$(2.66 \pm 1.33) \times 10^{-4}$
8641 ± 3	$10274 - 1633$	$\begin{array}{c}\nT = 1 \\ 2^+ \rightarrow 2^+ \n\end{array}$	A	$(1.26 \pm 0.16) \times 10^{-3}$
9251 ± 3 ^d	$10884 - 1633$	$T=1$ $3^+ \rightarrow 2^+$	E	$(3.80 \pm 1.31) \times 10^{-4}$
9628 ± 5 ^d	$11\,261 - 1633$	$T=1$ $1^+ \rightarrow 2^+$	C	$(4.26 \pm 1.50) \times 10^{-4}$
11261 ± 5	11 261 \rightarrow 0	$T = 1$ $T^+ \rightarrow 0$	в	$(2.13 \pm 0.28) \times 10^{-3}$

TABLE I. ²⁰Ne γ -ray transitions following ²⁰Na decay.

^a Nuclear energy-level differences are given. Recoil corrections of up to 3.4 keV have been made.

 b The labels used here are the same as in Fig. 1.</sup>

 c_{γ} -ray intensities are given relative to the intensity of the 1.63-MeV γ ray, which was itself determined to be $79.38 \pm 1.58\%$ of ²⁰Na decay, including cascades. (See text).

^d Too weak for half-life check.

C. γ -ray intensity measurements

The ν -ray intenstities reported here are based on high-energy double-escape-peak efficiencies measured relative to the full-energy-peak efficiency at 1.63 MeV. The collimator, attenuators, and geometry of the main experiment were used also for the efficiency experiments. At low energies the energy dependence of the full-energypeak efficiency was determin established intensity ratios of 56 Co γ rays. 22 high energies the double-escape-peak efficiency is based on measurements of γ -ray cascades from ²³Na(β , γ)²⁴Mg resonances. γ rays were observed at 55° with respect to the beam, where the dominant anisotropi and 1417-keV resonances give equal-intensity
 $(\pm 2\%)$ cascades,²³ ²⁴Mg(12.957 – 1.369 – 0.0 MeV) and 24 Mg(13.052 \div 4.123 \div 1.369 MeV), respectively, which determine the detection efficiency at 11.59 elative to low energies. The efficiency ratios were obtained from th

TABLE II. Decay of ²⁰Na to γ -ray emitting states of ²⁰Ne.

^a $\overline{\mathfrak s}$ values include radiative, nuclear-size, lepton-wavelength, electronic-screening, and electron-capture corrections.

 $^{\rm b}$ Based on the 1.63-MeV γ -ray intensity with corrections for cascade contributions from higher states.

 $\frac{3}{2}$ method of the observed 9873 \rightarrow 1633 γ ray, to include other branches. dIncludes α emission, the observed 10 274 \rightarrow 1633 γ ray, and the known 1 $_{10.27}/\Gamma_{8.64} = 0.034 \pm 0.009$.

 $i_{10,27}$, $i_{8,64}$ – 0.054 ± 0.009).
30 ± 0.10 times the strength of the observed 10 884 \rightarrow 1633 γ ray, to include other branches

r.30 \pm 0.10 times the strength of the observed 10 sold \rightarrow 1633 γ ray,
f Based on the sum of 11 261 \rightarrow 0 and 11 261 \rightarrow 1633 γ -ray intensities

The lowest $T = 1$ state of ²⁰Ne also emits α particles as shown. Decay to other α -emitting states is not shown. from the 9.87- and 10.88-MeV states which were too (See Ref. 8.) The dashed line indicates known transitions

B. γ -ray energy measurements An energy calibration was established using ²⁸Si γ rays observed in a ²⁸Si(p, n)²⁸P(β ⁺)²⁸Si(γ) experiment which was conducted immediately follow-

using the same experimental configuration. Tran-

ing the main ²⁰Ne(*p*, *n*)²⁰Na(β ⁺)²⁰Ne(γ)

weak to be observed here.

ratios after small corrections (2%) for relative attenuation in the target backings.

III. ANALYSIS AND DISCUSSION

A. Assignment of observed γ rays

The observed γ rays and their assignments are given in Table I. The 1.63- and 8.64-MeV transitions have previously been observed in $^{20}\mathrm{Na}$ decay. The latter, denoted A in Fig. 1, is the $\frac{1}{2}$ in Tal
 $\frac{1}{7}$, $\frac{1}{5}$, $\frac{1}{15}$ first-excited-state transition from the isobaric analog state 20 Ne(10.27 MeV, $T = 1$). The relative intensity $(1.26 \pm 0.16) \times 10^{-3}$ (or $\approx 1/794$) is consistent with the value $1/(1700 \pm 1100)$ reported originally.⁸ The energy observed here, 8641 ± 3 keV, also agrees with the previous value⁸ 8646 \pm 4 keV, and with the α energy²⁴ and γ -ray energy²⁵ in the $^{16}O(\alpha, \gamma)$ ²⁰Ne reaction. A 10.27-MeV γ -ray transition $3.4 \pm 0.9\%$ as strong²⁴ would not have been observed here. The 11.26- and 9.63-MeV transitions 8 and ^C are the ground-state and firstexcited-state transitions from a known 1⁺ state at
11.26 MeV.²⁶ The first-excited-state transition C 11.26 MeV.²⁶ The first-excited-state transition C from this state allows an energy calibration check to the highest energies, because its double-escape peak is close to the full-energy peak of the 8.64- MeV transition A. From 8.641 MeV, 1.633 MeV, the 33 ± 5 -keV separation of these peaks, and $2m_e c^2 = 1.022$ MeV, a 1⁺ state excitation energy 11.263 MeV is determined, in agreement with the preferred value 11.261 ± 0.005 MeV determined directly from the isolated higher-energy transition B. Transitions D and E correspond to firstexcited-state transitions from states at 9.87 and 10.88 MeV which have been observed in the ¹⁹F(³He, $d\gamma$)²⁰Ne reaction.^{27,28} Accompanying lower-energy γ rays from these states of the reported intensity²⁷ are beyond the sensitivity of this experiment, as is the ground-state 10.27 -MeV transition from 20 Ne(10.27 MeV, $T = 1$), but their presence is assumed in calculating the branching ratios and St values for the decay of 20 Na given in Table II.

B. Branching ratio calculations

The second forbidden 2^+ – 0⁺ ground-state β decay is expected to have negligible strength (10^{-7}) , and the experimental limit $< 5 \times 10^{-4}$ agrees with and the experimental limit <5×10⁻⁴ agrees with
this.¹⁵ The strength of all observed β^+ delayed α groups⁸ is 7.198 ± 0.050 times the strength of the group from ²⁰Ne(10.27 MeV, $T = 1$), which itself is $(27.94 \pm 2.66)^{-1}$ times as strong as the 1.63-MeV γ -ray strength,⁸ so that α emission is 0.2576 ± 0.0246 times the 1.63-MeV γ -ray strength. The known strength of γ -ray-emitting states that is not already included in the 1.63-MeV γ -ray strength (noncascade transitions from the 11.26- and 10.27- MeV states) is 0.0022 ± 0.0003 times the $1.63-MeV$ γ -ray strength. Combining all of this information

gives $79.38 \pm 1.58\%$ of ²⁰Na decay as the strength of the 1.63-MeV γ ray, including cascades. (This number normalizes the strength of branches to γ ray emitting states.)

Since 0.25% of the 1.63-MeV γ -ray strength results from cascades of the 11.26-, 10.88-, 10.27-, and 9.87-MeV states (including known unobserved branches), $^{24.27}$ decay directly to $^{20}Ne(1.63$ MeV) is 79.18 \pm 1.58% of ²⁰Na decay, only slightly smaller than the value $79.47 \pm 1.57\%$ used before the presthan the value $13.3121.31\%$ used before the press-
ent γ -ray observations,⁸ and in the direction of a larger $\mathfrak{F}t$ value. At the same time, the y-ray competition in the decay of $^{20}Ne(10.27 \text{ MeV}, T=1)$ to 16 O + α renormalizes the α strength of the analog state [when a given 20 Ne(10.27 MeV, T = 1) ft value is assumed and of all α -emitting 2⁺ states measured relative to it, which increases the indirectly computed branching ratio for the first excited state, decreasing the St value. The effect of the present observations is thus to narrow the difference of the direct and indirect $\mathfrak{F}t$ calculations for the 20 Ne(1.633 MeV) branch, and to reduce the uncertainty in the mirror β ⁺ $\mathfrak{F}t$ asymmetry

The branching ratios (for positron emission plus electron capture) given in Table II, with the exception of that for the 10.27-MeV state, are based on the observed γ -ray intensities and known unobserved transitions and 79.38% for the intensity of the 1.63-MeV γ ray. For the α -emitting state ²⁰Ne(10.27 MeV, $T = 1$), we use the present $\Gamma_{8.64}$ / $_{63}$ = (1.26 ± 0.16) $\times 10^{-3}$, $\Gamma_{\alpha}/\Gamma_{1.63}$ = (27.94 ± 2.66) of Torgerson et al.,⁸ and $\Gamma_{10.27}/\Gamma_{8.64} = 0.034$ ± 0.009 of Pearson and Spear²⁴ to compute the branching ratios $\Gamma_{8.64}/\Gamma$ = 0.034 ± 0.005 and $\Gamma_{\alpha}/$ $\Gamma = 0.965 \pm 0.005$. Then the ²⁰Na(β ⁺)²⁰Ne(10.27 MeV, $T = 1$) branching ratio is $(27.94 \pm 2.66)^{-1}$ $(0.0965$ ± 0.005 ⁻¹ times the strength of the 1.63-MeV γ ray, or $2.944 \pm 0.224\%$ of ^{20}Na decay. The previou value was $2.89 \pm 0.23\%$.⁸ Thus, the higher strengt determined in the present work for the 8.64-MeV transition results in a slight increase in the calculated rate of the superallowed transition, and a slightly smaller $\mathfrak{F}t$ value.

The observations of Adelberger and Marrs of the ¹⁹F(³He, $d\gamma$)²⁰Ne reaction²⁷ give the relative intensities for γ rays from ²⁰Ne states at 9.87, 10.88, and 11.58 MeV. Ratios of the total decay of those states to the observed transitions to 20 Ne(1.63 MeV) were taken as (1.29 ± 0.10) , $(1.30$ ± 0.10), and (1.60 ± 0.20) , respectively, in calculating 20 Na decay branching ratios or upper limits for these states.

C. Ft-value calculations

Comparative half-lives $\mathfrak{F}t$ for the observed 20 Na β^* -decay transitions to y-ray emitting ²⁰Ne states are given in Table II, and are based on the excitation energies and branching ratios BR(EC+ β^+)

of Table II, the 20 Na- 20 Ne mass difference 13.892 $+0.007 \text{ MeV}$, and the ²⁰Na half-life $t_{1/2} = 0.4457$ ± 0.0029 sec.¹² The statistical rate function f^s , including the effect of screening by atomic electrons, mas calculated with a code developed by trons, was calculated with a code developed by
Bahcall and Zimmerman,²⁹ and additional smal corrections mere calculated in approximations developed by Wilkinson^{30, 31} which are accurate to 0.1% for light elements and energetic decays. Corrections have also been made for electron-

capture contributions³² to the decay rates, so that
\n
$$
\mathfrak{F}t = f^{s}(1 + \epsilon_{1,2} + \epsilon_{3})(1 + \delta^{R})
$$

$$
\times (1+f_0^\epsilon/f_0^{\beta+})\, t_{1/2}/\mathrm{BR}(\mathrm{EC}+\beta^+)
$$
 ,

with f^s calculated as in Ref. 29, the nuclear-size and lepton-wavelength correction $(1 + \epsilon_{1,2} + \epsilon_3)$ from Ref. 30, the outer radiative correction $(1+\delta^R)$ from Ref. 31, and the electron-capture-to-positron ratio $f_0^{\epsilon}/f_0^{\beta+}$ from Ref. 32. As a check, similar calculations were made for carefully analyzed 0^+ – 0^+ pure Fermi decays of 14 O and 26 ^mAl, with results which agreed to $\sim 0.2\%$ with published reresults which agreed to $\sim 0.2\%$ with published results based on the same data.³³ Accuracy of this order is important only for the mirror β^* -decay Ft comparison, and for the discussion of superallomed transitions.

D. Unobserved transitions

Besides the transitions presented in Tables I and II, the marginally too-large single-escape peak of transition D suggests a possible underlying transition of energy 8.74 ± 0.02 MeV, and ²⁰Na decay branching ratio $\leq 2 \times 10^{-4}$. A (1⁻) state at 8.74 MeV was reported in early ${}^{20}Na(\beta^+\alpha)$ measurements, but mas not observed in a recent high-precision but was not observed in a recent ingu-precisive experiment.⁸ If a transition to $^{20}Ne(8.74 \text{ MeV})$ exists, it has $\log 5t > 6.5$, consistent with firstforbidden character.

There is likewise no definite evidence for either the >90% 8.32-MeV γ ray³⁴ from ²⁰Ne(9.95 MeV) [a suggested (1⁺) state]^{26,34} or the 60% 9.95-MeV γ ray²⁷ from ²⁰Ne(11.58 MeV). Upper limits for these transitions require $\log 3t > 6.0$ for $^{20}Ne(9.95)$ MeV) and $log 5t \ge 4.0$ for ²⁰Ne(11.58 MeV). These limits do not rule out allowed transitions to these states.

E. 1⁺ state at 11.26 MeV

The allowed $\Im t$ value and the γ decay of the state observed here require $J^{\pi} = 1^+$ or 3^+ . The log $\mathcal{F}t$ value 3.732 ± 0.053 and excitation energy 11.261 ± 0.005 MeV are remarkably close to the values 3.73 and 11.27 MeV, respectively, predicted for a 1^* , $T = 1$ state in extended-shell-model calculaa $1^{\ast}, T$ = 1 state in extended-shell-model calcula
tions of Lanford and Wildenthal.¹⁴ The observe

11.26-MeV γ -ray transiton B corresponds to the giant M1 transition exciting a 1^+ state at this energy (nearly exhausting the energy-weighted sun
rule) in 180° inelastic electron scattering.³⁵ The rule) in 180° inelastic electron scattering.³⁵ The relative strength 0.20 ± 0.08 for the transitions C and B to $^{20}Ne(1.63 \text{ MeV})$ and $^{20}Ne(0.0 \text{ MeV})$ (see Table I) corresponds to γ -ray branching ratios 17% and 83%, and to reduced transition probabilities $B(M1)$ in the ratio 0.32 ± 0.12 . These agree well with the shell-model calculations of Maripuu well with the shell-model calculations of Maripuu
and Wildenthal,¹³ which predicted branching ratio: 18% and 82% $[B(M1)$ ratio 0.353]. A brief communication³⁶ has reported observation of this state in the $^{19}F(d, n\gamma)$ ²⁰Ne reaction at a slightly lower excitation energy 11.252 ± 0.002 MeV and with relative γ -ray intensities corresponding to a $B(M1)$ ratio 0.53 ± 0.07 which, however, favored the rotational-model prediction 0.50.

There may be another unnatural-parity state nearby. Excitation energies which range from 11.23 to 11.27 MeV have been reported and are not entirely consistent.²⁶ Representative values not entirely consistent.²⁶ Representative values
include 11.259 ± 0.010 MeV,³⁷ 11.233 ± 0.010 MeV,²
11.239 ± 0.015 MeV,²⁸ 11.252 ± 0.002 MeV,³⁶ and th $11.239\pm0.015\,\,{\rm MeV,}^{28}\,\,11.252\pm0.002\,\,{\rm MeV,}^{36}$ and the present value 11.261 ± 0.005 MeV. Examination of Fig. 1 suggests a possible γ -ray peak approximately 40 keV lower in energy than the doubleescape peak of transition C, but this is not definitely established (and is not included in calculating the intensity of C here).

F. 3+ states at 10.SS and 9.87 MeV

The allowed $\mathfrak{F}t$ values and γ decay of these states again require $J^{\pi} = 1^+$ or 3^+ . Neither state
is excited in 180° inelastic electron scattering,³⁵ is excited in 180° inelastic electron scattering, 35 but this does not eliminate 1' assignments, because theoretical considerations indicate that the $M1$ strength in ²⁰Ne should be concentrated in a single 1⁺ state [as is observed³⁵ for ²⁰Ne(11.26) MeV)], although several 1^* states may be pres-MeV)], although several 1⁺ states may be pres-
ent.^{13,38} For each state the ¹⁹F(³He, dy)²⁰Ne results of Adelberger and Marrs²⁷ show that the γ decay proceeds to the 2' first-excited state (the transition E and ^D observed here} and to the 4' second-excited state, but not to the 0^+ ground state. This favors a 3' assignment for each state. A $T = 1$ assignment for ${}^{20}Ne(10.88 \text{ MeV})$ is favored
by the existence of the 3^{+} state ${}^{20}F(0.656 \text{ MeV})$, 26 by the existence of the 3^* state $^{20}F(0.656 \text{ MeV})$, 26 its probable analog. $^{20}Ne(9.87 \text{ MeV})$ is below the 10.27-MeV analog of the 20 F and 20 Na ground states, and is assigned $T = 0$. Shell-model calculations¹⁴ predicted a 3^+ , $T=1$ state at 10.75 MeV with $\log ft = 4.89$, in close agreement with $log5t = 4.838 \pm 0.132$ of ²⁰Ne(10.88 MeV). The $log5t = 4.90$ for a predicted 3^+ , $T = 0$ state at 10.50 MeV does not agree with the $\log \mathfrak{F}t = 5.783 \pm 0.178$ observed here for $^{20}Ne(9.87 \text{ MeV})$.

 20 Ne(10.27 MeV) was identified as the lowest T = 1 state, isobaric analog of ²⁰Na and ²⁰F, from measurements of the isospin-forbidden measurements of the isospin-forbidden
 $^{16}O(\alpha, \gamma)^{20}$ Ne reaction.²⁴ The excitation energy 10.274 ± 0.003 MeV determined here agrees with 10.279 ± 0.004 MeV determined originally from the β^+ delayed γ -ray energy,⁸ with 10.278 ± 0.005 MeV from the β^* delayed α energy,⁸ with 10.272 ± 0.009 MeV from the ¹⁶O(α , γ)²⁰Ne resonance energy (E_{α}) MeV from the ¹⁶O(α , γ)²⁰Ne resonance energy (E_{α}
= 6.930 ± 0.010 MeV),²⁴ and with 10.271 ± 0.003 MeV
from the ¹⁶O(α , γ)²⁰Ne γ -ray energy.²⁵ The *M*1 from the ¹⁶ $O(\alpha, \gamma)$ ²⁰Ne γ -ray energy.²⁵ The M1 strength of the 8.64-MeV γ -ray transition from this state shows this transition to be predominantly orbital, being approximately 4.5 times the strength computed for the spin part from the analogous β^{\pm} decays.^{17,24} This large Γ_{γ} value¹⁷ is well approximated by extended-shell-model calculations,³⁹ as are the β^{\pm} decays to the same
state.¹⁴ From the small α -decay width¹⁷ Γ_{α} state.¹⁴ From the small α -decay width¹⁷ Γ_{α} =116 ± 20 eV for ²⁰Ne(10.27 MeV, $T = 1) \div {}^{16}O + \alpha$, 100 times smaller than Γ_{α} for nearby states of the same spin and parity, an isospin impurit
limit of about 1% may be determined.¹⁷ Thes limit of about 1% may be determined.¹⁷ These matters were discussed in greater detail in con-
nection with the ¹⁶O(α , γ)²Ne reaction.¹⁷ nection with the $^{16}O(\alpha, \gamma)^{20}$ Ne reaction.¹⁷

H. State at I.63 MeV—mirror St comparison for ²⁰ Na(β ⁺) and ²⁰ F(β ⁻) decays

For ${}^{20}Na(\beta^*){}^{20}Ne(1.633 \text{ MeV})$, the experimental branching ratio $79.18 \pm 1.58\%$ and $\mathfrak{F}t$ value 97250 ± 2060 sec were given in Table II. An alternate, indirect computation of these quantities starting from a theoretical value for the superallowed from a theoretical value for the superallowed
transition is also of interest, $1-3 \cdot 12$ and is justified by the extremely well known normal Fermi decay strength, 33 and by the high isospin purity of the analog state (deduced from its small α -particle analog state (deduced from its small α -partic
decay width).¹⁷ The Ft value 2800 ± 80 sec assumed for the superallowed transition is a revision of a previous value 2780 ± 80 sec³ reflecting the increased precision in the normal Fermi the increased precision in the normal Fermi
strength.³³ This value was obtained by combinin the Fermi strength with a range of possible values for the small Gamow-Teller component.³ The assumed ft value for the ²⁰Ne(10.27 MeV, $T = 1$) branch agrees with direct knowledge (Table II), so the indirect ft value for the ²⁰Ne(1.633 MeV) branch which results from it will also agree. The expected uncertainty in the superallowed transition strength is smaller for the theoretical value, however, so that the indirect computation may well give the best ft value for ${}^{20}\text{Na}(\beta^+){}^{20}\text{Ne}(1.633)$ MeV). Shell-model calculations¹⁴ gave $log5t = 4.98$, very close to both the direct and indirect caleulations for the $^{20}Ne(1.63 \text{ MeV})$ branch.

Assuming $\mathfrak{F}t = 2800 \pm 80$ sec and using Γ_{α}/Γ $= 0.965 \pm 0.005^{17}$ for 20 Ne(10.274 \pm 0.003 MeV, T = 1), and using the relative α intensities of Torgerso *et al.*,⁸ β^* delayed α groups amount to 21.63 \pm 0.72% of ²⁰Na decay. Decay to γ rays amounts to an additional 0.37%. Thus the indirectly computed branching ratio for ${}^{20}Na(\beta^+){}^{20}Ne(1.633 \text{ MeV})$ is $78.00 \pm 0.72\%$, about 1.5% smaller than the direct experimental value $79.18 \pm 1.58\%$. The corresponding (larger) indirect 20 Na $\mathfrak{F}t$ value is 98728 ± 1154 sec. For the 20 F side of the mirror the $\mathfrak{F}t$ value is 94134 ± 174 sec, based on maximum β ⁻ energy, 5.3959 \pm 0.0008 MeV¹² and branching ratio 99.983 $\pm 0.003\%$ ¹² for ²⁰F(β ⁻)²⁰Ne(1.633 MeV), and ²⁰F half-life 10.999 ± 0.019 sec (a weighted average of 11.03 ± 0.06 sec² and 10.996 ± 0.020 sec⁴⁰). The $A = 20$ mirror asymmetry $\delta = [\mathfrak{F}t(\beta^*)/\mathfrak{F}t(\beta^-)] - 1$ is thus 0.033 ± 0.022 computed directly, and 0.049 ± 0.013 computed indirectly starting from the assumed strength of the superallowed transition.

Since 1970 when the comparison of allowed mirror β^{\pm} decays underwent renewed scrutiny¹ in connection with the possibility of second-class weak currents, the δ values computed for $A = 20$ have changed considerably. With successive experimental improvements the best value of δ has passed through the values -0.067 ± 0.032 , $+0.062$ passed through the values $-0.067 + 0.032$, $+0.037$, $+0.054 + 0.023$, $+0.026 + 0.023$, $+0.037$, $+0.054 + 0.023$, $+0.037$ the present direct and indirect results +0.033 ± 0.022 and $+0.049 \pm 0.013$, respectively. Many improvements have been made for other mirror improvements have been made for other mirro
pairs as well.¹² Simultaneously a better under standing of nuclear-structure effects has developed, so that the binding-energy-induced wave function differences between the nominal mirror nuclei are now thought to account for most of the nuclei are now thought to account for most of the experimentally observed β^+ -decay $\mathfrak{F}t$ asymmetry.¹² For the $^{20}Na-^{20}F$ comparison, the experimental asymmetry δ is indistinguishable from δ bind $= 0.029 \pm 0.025$ calculated in a recent survey.¹²

Interest in the possibility of second-class currents continues, but it now seems likely that searches using phenomena less sensitive to nuclear structure effects will prove more definitive. 41

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