Significance of the compound nucleus surface energy coefficients in the synthesis of superheavy nuclei with Z = 112-120

R. Zargini^{®*} and S. A. Seyyedi^{®†}

Department of Physics, Payame Noor University (PNU), P.O. Box 19395-4697, Tehran, Iran

(Received 2 July 2023; revised 2 August 2023; accepted 16 August 2023; published 5 September 2023)

This paper investigates the impacts of the different surface energy coefficients on the compound nucleus decay modes during heavy ion fusion reactions, with focus given to the superheavy nuclei (SHN) in the range of Z = 112-118. The evaporation-residue (ER) cross sections were calculated for different surface asymmetric constants, k_s and surface energy constants, γ_0 . In these calculations, the dinuclear system model and proximity potential, along with considering deformed nuclei, were employed. Comparing the experimental data and this theoretical approach, the best values of k_s and γ_0 are 0.7546 and 0.9180 MeV fm⁻², respectively. Furthermore, this new model was used to investigate the probability of synthesis of experimentally unknown heavier systems with Z = 119 and 120. There exist five promising combinations to synthesize SHN with Z = 119: (a) ²⁴⁹Cf(⁴⁵Sc, 3n)²⁹¹119 with the ER cross section, $\sigma_{3n} = 417.1$ fb at the incident energy $E_{c.m.} =$ 219 MeV($E^* = 39.84$ MeV), (b) ²⁴⁹Cf(⁴⁵Sc, 4n)²⁹⁰119 with the ER cross section, $\sigma_{4n} = 138.5$ fb at the incident energy $E_{c.m.} = 221 \text{ MeV}(E^* = 41.84 \text{ MeV})$, (c) ²⁴⁷Bk(⁵⁰Ti, 3*n*)²⁹⁴119 with the ER cross section, $\sigma_{3n} = 11.2 \text{ fb}$ at the incident energy $E_{\text{c.m.}} = 226 \text{ MeV}(E^* = 35.15 \text{ MeV})$, (d) $^{254}\text{Es}(^{48}\text{Ca}, 3n)^{299}119$ with the ER cross section, $\sigma_{3n} = 9115.15$ fb at the incident energies $E_{c.m.} = 208 \text{ MeV}(E^* = 32.14 \text{ MeV})$, and e) $^{254}\text{Es}(^{48}\text{Ca}, 4n)^{298}$ 119 with the ER cross section, $\sigma_{4n} = 735.46$ fb at the incident energies $E_{c.m.} = 210 \text{ MeV}(E^* = 34.14 \text{ MeV})$. In addition, it is found that the best combinations to synthesize SHN with Z = 120 are ²⁴⁹Cf(⁵⁰Ti, 3*n*)²⁹⁶120 with the ER cross section, $\sigma_{3n} = 51.19$ fb at the incident energy $E_{c.m.} = 228 \text{ MeV}(E^* = 33.19 \text{ MeV})$, and ${}^{251}\text{Cf}({}^{50}\text{Ti}, 3n){}^{298}120$ with the ER cross section, $\sigma_{3n} = 43.17$ fb at the incident energy $E_{c.m.} = 227 \text{ MeV}(E^* = 33.82 \text{ MeV})$.

DOI: 10.1103/PhysRevC.108.034606

I. INTRODUCTION

Superheavy nuclei (SHN) investigations have attracted a great deal of attention among researchers in the field of nuclear physics. Fusion reactions are among the important tools in the synthesis of SHN. Nuclei with Z < 112 are produced via cold fusion reactions using double magic target ²⁰⁸Pb whereas nuclei with Z = 112-118 are generated via hot fusion reactions and employing double magic projectile ⁴⁸Ca and actinide targets [1-16]. Theoretical studies could provide extremely valuable insights into the optimization of expensive experimental efforts towards the synthesis of SHN [17-43]. Different models, such as the two step model [44], the dinuclear system (DNS) model [17-20], the fusion by diffusion model [28,45,46], and the nuclear collectivization model [37] have been previously employed to calculate the evaporationresidue (ER) cross sections. In the DNS model, the calculation of total potential is of great importance for which different models such as double folding, Woods-Saxon, Skyrme energy density formalism, and proximity have been developed [43,47]. Proximity potential is employed as an analytical description for colliding nuclei (an approach based on the phenomenology model). All of these potentials' foundations are based on the proximity theorem which the nuclear part of the interaction potential will produce various factors. These factors are independent of the colliding nuclei masses and are dependent on the inverted harmonic oscillator and universal function. This concept provides an important simplification in the study of nuclear potential [43]. One of the important parameters in the potential proximity is the surface energy coefficient, γ . 13 coefficients for surface asymmetric constant, k_s and surface energy constant, γ_0 have been introduced [48–55], which contribute to factors such as the height of the potential barrier, the position of the potential barrier, and inverted harmonic oscillator potential, which consequently impact the calculated ER cross sections.

To study SHN synthesis with the DNS model, one needs to know the total potential of colliding nuclei. For the calculation of the ER cross section, there are three fundamental stages: a) capturing the projectile by the target and overcoming potential barriers, b) combination of the projectile and the target nuclei and formation of excited compound nucleus, and c) evaporation of one or more neutrons from the excited compound nucleus reaching a state of equilibrium and survive against fission [17]. In the past four decades, different theoretical models predicted that there is island of stability on the top of the nuclear chart with a relatively high half-life [56,57]. These models have predicted that beyond closed proton shell number with Z = 82, proton numbers with Z = 114, 120, and 124 are also closed-proton shells. Similarly, beyond closed neutron shell number with N = 126, neutron number with N = 184 can be introduced as a new closed neutron shell in

^{*}reza.zargini@gmail.com; r.zargini@student.pnu.ac.ir

[†]a.seyyedi@pnu.ac.ir

the nuclear chart [58]. At the moment, the californium (Cf) nucleus is the heaviest available target that can be used for SHN synthesis, and the production of targets heavier than Cf is very difficult. For example, einsteinium (²⁵⁴Es) can be produced in microgram quantities, approximately three orders of magnitude less than typically required for SHN synthesis [59]. Therefore, to produce nuclei with Z = 119 and 120, the projectile heavier than ⁴⁸Ca such as ⁵⁰Ti, ⁵⁴Cr, ⁵⁸Fe, and ⁶⁴Ni should be used [29]. In 2008, one experiment was set up to synthesize nuclei with Z = 120, by the combination of 238 U(64 Ni, xn) ${}^{302-x}$ 120, where the ER cross section, $\sigma =$ 90 fb at the excited compound nucleus (CN) energy $E^* =$ 36.4 MeV, was reported [32]. In 2009, another attempt was performed in order to synthesize nuclei with Z = 119 by the combination of ${}^{249}\text{Bk}({}^{50}\text{Ti}, xn)^{299-x}119$, where the maximal measured ER cross section, 50 fb was reported [60]. In the same year, for the synthesis of Z = 120 SHN, the combination of 244 Pu(58 Fe, xn) ${}^{302-x}$ 120, the ER cross section was measured, $\sigma = 400$ fb [61]. In 2015 Hofmann *et al.* tried to use the combination of 248 Cm(54 Cr, xn) ${}^{302-x}$ 120, to synthesize Z = 120, for which they reported the maximum ER cross section equal to $5.3^{+3.6}_{-2.1}$ pb [62]. In 2020, another group tried to produce nuclei with Z = 119, 120 by combinations of 249 Bk $({}^{50}$ Ti, $xn)^{299-x}$ 119 and 249 Cf $({}^{50}$ Ti, $xn)^{299-x}$ 120. In those attempts, the ER cross sections were reported equal to 65 fb and 200 fb at the excited CN energies $E^* =$ 43.2 MeV and $E^* = 37.6$ MeV, respectively [63]. No evidence confirming the formation of SHN with Z = 119and 120 was observed. Some more theoretical studies have been performed on the synthesis of SHN with Z = 119and 120 [19,22,26,27,29,31,32,34,38-41,45,64-68]. In these researches, different combinations have been suggested to synthesize Z = 119, 120; the most suggested combination for Z = 119 is 249 Bk $({}^{50}$ Ti, $xn)^{299-x}$ 119 with ER cross section in x = 3, 4 neutron channels and ranges from 12 fb to 480 fb and the most suggested combination for Z = 120 is 249 Cf(50 Ti, xn) $^{299-x}$ 120 with ER cross section in x = 3, 4 neutron channels and ranges from 1.5 fb to 760 fb.

This current paper presents the model and employed theory in Sec. II. The results of our study on the probability of the synthesis of SHN in the range of $112 \le Z \le 118$ are described in Sec. III A. Moreover, using the obtained model, our results of the probability of the synthesis of nuclei with Z = 119, 120 will be discussed in Sec. III B.

II. MODEL AND THEORY

A. Total interaction potential

The potential barrier characteristics, such as barrier height, position, curvature, and shape, are important parameters that should be calculated. Total potential is the function of two nuclei radius distance, colliding angle, and deformation parameters of two nuclei. Total potential is equal to the summation of the coulomb long range repulsive potential, nuclear short range attractive potential, and centrifugal potential [43]:

$$V_{Total} = V_C(R.Z_i, \beta_{\lambda i}, \theta_i) + V_N(R, A_i, \beta_{\lambda i}, \theta_i) + V_l(R, A_i, \beta_{\lambda i}, \theta_i),$$
(1)



FIG. 1. Interaction angles of two nuclei during fusion.

where $V_C(R.Z_i, \beta_{\lambda i}, \theta_i)$ is the Coulomb potential, $V_N(R, A_i, \beta_{\lambda i}, \theta_i)$ shows the nuclear potential, and $V_l(R, A_i, \beta_{\lambda i}, \theta_i)$ denotes the centrifugal potential. Considering the multipole deformations, the Coulomb potential is [69]

$$V_C(R.Z_i, \beta_{\lambda i}, \theta_i) = \frac{Z_1 Z_2 e^2}{r} + 3Z_1 Z_2 e^2 \sum_{\lambda, i=1,2} \frac{R_i^{\lambda}(\alpha_i, T)}{(2\lambda + 1)R^{\lambda + 1}} \times Y_{\lambda}^{(0)}(\theta_i) \bigg[\beta_{\lambda i} + \frac{4}{7} \beta_{\lambda i}^2 Y_{\lambda}^{(0)}(\theta_i) \bigg].$$
(2)

In Eq. (2), Z_1Z_2 is the Coulomb factor and is related to projectile and target, *r* is the center of two nuclei distance, $R_i^{\lambda}(\alpha_i, T)$ is the deformed radius of the projectile and target, $\beta_{\lambda i}$ represents multipole deformations of the projectile and target, θ_i is the angle between the nuclear symmetry axis and the collision *Z* axis, and α_i is the angle between the symmetry axis and the radius vector of the colliding nucleus [21,25]. It should be noted that α_i is measured in the clockwise direction from the symmetry axis while θ_i is measured in the counterclockwise direction, as shown in Fig. 1.

For the calculation of the nuclear potential, the proximity potential is employed as follows [47]:

$$V_N(R, A_i, \beta_{\lambda i}, \theta_i) = 4\pi \gamma \ bR\phi\left(\frac{z}{b}\right),\tag{3}$$

where $b = (\frac{\pi}{\sqrt{3}})0.55 \approx 0.99 \text{ fm}$, $R = \frac{C_1.C_2}{C_1+C_2}$, $C_i = R_i(1 - (\frac{b}{R_i})^2)_{i=1,2}$ are Susmann central radii for projectile and target, and $\phi(\frac{z}{b})$ is the universal function, where in $\xi = \frac{z}{b} = \frac{r-C_1-C_2}{b}$. *z* is the minimum distance of projectile and target during colliding. In Eq. (3), the CN surface energy coefficient is obtained from [47]

$$\gamma = \gamma_0 \left[1 - k_s \frac{(N-Z)^2}{A^2} \right]. \tag{4}$$

In Eq. (4), γ_0 is the surface energy constant which is obtained from $\gamma_0 = \frac{a_s}{4\pi r_0^2}$, where a_s is the usual liquid drop model surface energy coefficient and r_0 is the nuclear radius constant, and k_s is the surface asymmetric constant [70]. N, Z, and A are the neutron number, proton number, and mass number of CN, respectively. The universal function is obtained as [47]

$$\phi(\xi) = -1.7817 + 0.9270\xi + 0.143\xi^2 - 0.09\xi^3,$$

for $\xi \le 0$.

034606-2

Considering the nuclear deformations, the projectile and target radii are given as [69]

$$R_i(\alpha_i, T) = R_{0i}(T) \left[1 + \sum_{\lambda} \beta_{\lambda i} Y_{\lambda}^{(0)}(\alpha_i) \right].$$
(6)

In Eq. (6), $R_{0i}(T) = (1.28A_i^{\frac{1}{3}} - 0.76 + 0.8A_i^{-\frac{1}{3}})_{i=1,2}$ is the nuclear radii of the colliding participant, where the nuclear deformations are not taken into the account. Centrifugal potential is given as

$$V_l = \frac{\hbar^2 l(l+1)}{2I_{NS}}.$$
 (7)

Here, $I_{NS} = \mu r^2$ is the nonsticking moment of inertia, and *l* is the angular momentum. As *l* increases, the depth of the saddle point will decrease in the graph of the scattering potential versus distance. In other words, the quasifission barrier will continue to decrease, until the saddle point in the potential graph vanishes. This results in reduced values for the survival probability of CN and consequently decay of the compound nucleus to fission fragments.

The barrier penetration model developed by Wong has been widely used to describe the fusion reactions at the energies close to, or greater than that of, the barrier [69]. The capture cross section is expressed as the sum of the cross sections for each partial wave, l [17],

$$\sigma_{cap} = \frac{\pi \hbar^2}{2\mu E_{\text{c.m.}}} \sum_{l=0}^{l_{max}} (2l+1)T_l(E_{\text{c.m.}},l).$$
(8)

In Eq. (8), μ is the reduced mass of the interacting nuclei, $E_{\text{c.m.}}$ is the center of mass of the colliding systems, l is the angular momentum, and $T_l(E_{\text{c.m.}}, l)$ denotes the penetration probability for the *l*th partial wave and calculated according to the Hill-Wheeler equation [71]

$$T_l(E_{\rm c.m.}) = (1 + \exp\left[2\pi (E_l - E_{\rm c.m.})/\hbar\omega_l\right])^{-1}.$$
 (9)

For calculating the capture cross section, the Wong formula is used [24,69]:

$$\sigma_{cap}(E) = \frac{10R_0^2\hbar\omega_0}{2E_{\text{c.m.}}}\ln\left(1 + \exp\left[\frac{2\pi(E_{\text{c.m.}} - E_0)}{\hbar\omega_0}\right]\right), \quad (10)$$

where E_0 is the height of the total potential and $\hbar\omega_0$ shows the inverted harmonic oscillator potential and is given as

$$\hbar\omega_0 = \frac{\hbar}{\sqrt{\mu}} \sqrt{\left|\frac{d^2 V(r)}{dr^2}\right|_{R_l}}.$$
(11)

Fusion cross section is one of the main parameters, which is described as

$$\sigma_{fus} = \frac{\pi \hbar^2}{2\mu E_{\rm c.m.}} \sum_{l=0}^{l_{max}} (2l+1)T_l(E_{\rm c.m.}, l)P_{CN}(E_{\rm c.m.}, l).$$
(12)

In Eq. (12), $P_{CN}(E_{c.m.}, l)$ is the fusion probability for which various models have been used [18,22,40]. In this work to calculate the fusion probability, the developed model by Za-grebaev and Greiner [40] was employed:

$$P_{CN}(E^*, l) = \frac{P_{CN}^0}{1 + \exp\left[\frac{E_B^* - E^*}{\Delta}\right]},$$
(13)

$$P_{CN}^{0} = \frac{1}{1 + \exp\left[\frac{Z_{1}Z_{2}-\zeta}{\tau}\right]}.$$
 (14)

In Eqs. (13) and (14), E^* is the CN excitation energy, E_B^* is the energy of the CN at the center-of-mass beam energy equal to the Bass barrier, Δ is an adjustable parameter and is often set around 4 MeV, $\zeta \approx 1760$ and $\tau \approx 45$ are the fitted parameters [40]. The compound nucleus excitation energy which also defines the damping of the shell correction to the fission barrier of CN, is obtained from

$$E_{CN}^* = E_{\text{c.m.}} + Q_{val} - E_{rot},$$
 (15)

where E_{rot} is the rotational energy of CN and is obtained from $E_{rot} = \frac{\hbar^2 l(l+1)}{2 \mu r^2}$ and Q_{val} is obtained from [35]

$$Q_{val} = \Delta M(A, Z) - \Delta M(A_1, Z_1) - \Delta M(A_2, Z_2).$$
 (16)

In Eq. (16), $\Delta M(A, Z)$, $\Delta M(A_1, Z_1)$, and $\Delta M(A_2, Z_2)$ are mass excess values of CN, projectile, and target, respectively. In this work, to calculate Q_{val} , the mass excess values were obtained from Möller *et al.*'s [72] tables.

B. The evaporation residue cross section

Once the excited CN is formed the equilibrium condition is achieved via the evaporation of one or more neutrons before decay. Therefore, one should consider the competition between neutron emission and decay to fission fragments in CN. In cold fusion reactions, the formed CN has an excitation energy of $E_B^* < 15$ MeV, where the equilibrium condition is reached by the evaporation of one or two neutrons, where, as in hot fusion reactions with the excitation energy of $E_B^* > 15$ MeV, the evaporation of three or four neutrons is involved [36]. The evaporation residue cross section is given by

$$\sigma_{ER}^{xn} = \frac{\pi \hbar^2}{2\mu E_{\text{c.m.}}} \sum_{l=0}^{l_{max}} (2l+1)T_l(E_{\text{c.m.}},l) \times P_{CN}(E_{\text{c.m.}},l)W_{sur}^{xn}(E^*,l),$$
(17)

where $W_{sur}^{sn}(E^*, l)$ denotes the survival probability of CN. Because of the high Coulomb barrier for the emission of charged particles from the excited heavy nucleus, the widths for the emission of a proton or an α particle are much smaller than the neutron emission width Γ_n . Under these circumstances, $\Gamma_t \approx \Gamma_n + \Gamma_f$ was set and the survival probability $W_{sur}^{sn}(E^*, l)$ reflects the competition between neutron evaporation and fission of the excited CN. The survival probability is given by [42]

$$W_{sur}^{xn}(E^*, l) = P_{xn}(E_{CN}^*) \prod_{i=1}^{i_{\max}=x} \left(\frac{\Gamma_n}{\Gamma_n + \Gamma_f}\right)_{i, E^*}, \quad (18)$$



FIG. 2. The distribution of total potential versus distance for 249 Cf(45 Sc, xn) $^{294-x}$ 119 combination, along with deformations, β_{2i} and rotations, α_i , of projectile and target.

where Γ_n and Γ_f are partial neutron emission width and partial fission width, respectively. At $E_{CN}^* > 10$ MeV, the Γ_n is much smaller than the Γ_f . Therefore, the survival probability $W_{sur}^{xn}(E^*, l)$ is mainly determined by the realization probability $P_{xn}(E_{CN}^*)$ and the ratio $\frac{\Gamma_n}{\Gamma_c}$ [20,73]:

$$\frac{\Gamma_n}{\Gamma_f} = \frac{4A^{\frac{2}{3}}a_f(E_{CN}^* - B_n)}{K_0 a_n \left[2a_f^{\frac{1}{2}}[(E_{CN}^* - B_f)]^{\frac{1}{2}} - 1)\right]} \times \exp\left[2a_n^{1/2}(E_{CN}^* - B_n)^{\frac{1}{2}} - 2a_f^{1/2}(E_{CN}^* - B_f)^{\frac{1}{2}}\right].$$
(19)



FIG. 3. The distribution of total potential versus distance for 249 Cf(45 Sc, xn) $^{294-x}$ 119 combination, along with deformations, β_{2i} and directions, θ_i , of projectile and target.

In Eq. (19), B_n denotes the neutron binding energy, B_f is the fission barrier, K_0 is set to a constant value of 10 MeV, $a_n = A/10$ and $a_f = 1.1a_n$ are the level density parameters of the fissioning nucleus at the ground state and saddle configurations, respectively [73]. The fission barrier is given by [18]

$$B_f(E^*) = B_f^{LD} + S \exp\left(\frac{-E^*}{E_D}\right).$$
(20)

In Eq. (20), B_f^{LD} is the liquid drop (LD) fission barrier (macroscopic) and *S* is the shell correction term [34]. The liquid drop fission barrier is very low or equal to zero for heavy elements with $Z \ge 109$ [19,42]. The shell damping energy is given by

$$E_D = \frac{5.48A^{\frac{1}{3}}}{1+1.3A^{-\frac{1}{3}}}.$$
 (21)

To calculate $P_{xn}(E_{CN}^*)$, an equation, which was developed by Jackson [74], was employed. $P_{xn}(E_{CN}^*)$ represent the probability of emitting an exact number of neutrons, *x*, from the CN and is given by

$$P_{xn}(E_{CN}^*) = I(\Delta_x, 2x - 3) - I(\Delta_{x+1}, 2x - 1), \quad (22)$$

where I(z, n) is the Pearson's incomplete γ function, and is obtained by $I(z, n) = (\frac{1}{n!}) \int_0^z m^n e^{-m} dm$, in Eq. (22), $\Delta_x = \frac{(E_{CN}^* - \sum_i^x B_n(i))}{T}$, $B_n(i)$ is the binding energy of the *i*th evaporated neutron, and *T* is the compound nucleus temperature that is given by $E_{CN}^* = E_{c.m.} + Q_{val} = \frac{1}{a}AT^2 - T$. It should be noted that Eq. (22) is valid for calculation of the evaporation probability of two neutrons and more.

III. DISCUSSION AND CONCLUSION

A. Significance of the compound nucleus surface energy coefficients on the ER cross sections

Four parameters are involved in the calculation of the potential: a) The nuclear deformations, b) the surface energy coefficient, c) the universal function, and d) the nucleus radius. Nuclear deformations, along with collision angles, play important roles to the obtained ER cross sections. Figures 2 and 3 show the distribution of total potential versus distance,

TABLE I. The different surface energy coefficients available from the literature.

r_0 (fm)	a_s (MeV)	k_s	$\gamma_0 ({\rm MeV}{\rm fm}^{-2})$	Ref.
1.2049	18.56	1.79	1.01734	[48]
1.2249	17.9439	1.7826	0.9517	[<mark>49</mark>]
1.16	24.7	4.0	1.460734	[<mark>50</mark>]
1.18	21.7	3.0	1.2402	[51]
1.18	20.57	2.2	1.1756	[52]
1.16	21.53	2.5	1.27326	[52]
1.16	21.14	2.4	1.2502	[52]
1.16	21.13	2.3	1.2496	[52]
1.2249	17.9439	2.6	0.9517	[53]
1.16	21.18466	2.345	1.25284	[54]
1.18995	19.3859	1.9830	1.08948	[55]
1.21610	17.0603	0.7546	0.9180	[55]
1.21725	16.9707	2.2938	0.911445	[55]

TABLE II. Selected combinations along with the calculated values and experimental data for the surface energy coefficient γ , the CN excitation energy E_{CN}^* , and the ER cross section, σ_{ER} in different neutron channels.

			Calculated values			Experimental eata			
Combination	k_s	$\gamma_0 ({\rm MeVfm^{-2}})$	γ (MeV fm ⁻²)	E_{CN}^{*} (MeV)	σ_{ER} (pb)	$\overline{E_{CN}^{*}(\text{MeV})}$	σ_{ER} (pb)	Ref.	
$^{48}Ca + ^{238}U$	4.0	1.460734	1.8614	31.6	$\sigma_{3n} = 6.83$	29.3-33.5	$\sigma_{3n} = 0.5^{+1.15}_{-0.41}$	[1]	
				34.6	$\sigma_{3n} = 15.77$	34.6	$\sigma_{3n} = 0.7^{+0.58}_{-0.35}$	[12]	
				35.6	$\sigma_{3n} = 14.34$	32.9-37.2	$\sigma_{3n} = 2.5^{+1.8}_{-1.1}$	[<mark>1</mark>]	
				35.6	$\sigma_{4n} = 8.61$	32.9-37.2	$\sigma_{4n}=0.8$	[<mark>1</mark>]	
				39.6	$\sigma_{4n}=6.16$	37.7-41.9	$\sigma_{4n} = 0.6^{+1.6}_{-0.5}$	[<mark>1</mark>]	
	0.7546	0.9180	0.8854	31.6	$\sigma_{3n} = 0.005$	29.3-33.5	$\sigma_{3n} = 0.5^{+1.15}_{-0.41}$	[1]	
				34.6	$\sigma_{3n} = 0.51$	34.6	$\sigma_{3n} = 0.7^{+0.58}_{-0.35}$	[12]	
				35.6	$\sigma_{3n} = 1.50$	32.9-37.2	$\sigma_{3n} = 2.5^{+1.8}_{-1.1}$	[<mark>1</mark>]	
				35.6	$\sigma_{4n} = 0.90$	32.9-37.2	$\sigma_{4n}=0.8$	[<mark>1</mark>]	
				39.6	$\sigma_{4n}=2.81$	37.7-41.9	$\sigma_{4n} = 0.6^{+1.6}_{-0.5}$	[<mark>1</mark>]	
${}^{48}Ca + {}^{237}Np$	4.0	1.460734	1.2103	39.3	$\sigma_{3n} = 3.93$	36.9-41.2	$\sigma_{3n} = 0.9^{+1.6}_{-0.6}$	[2]	
	0.7546	0.9180	0.8883	39.3	$\sigma_{3n} = 2.03$	36.9-41.2	$\sigma_{3n} = 0.9^{+1.6}_{-0.6}$	[2]	
${}^{48}Ca + {}^{240}Pu$	4.0	1.460734	1.2071	38.1	$\sigma_{3n} = 4.70$	36.5-41.1	$\sigma_{3n} = 2.5^{+2.9}_{-1.4}$	[13]	
				43.1	$\sigma_{4n} = 1.01$	40.9-45.4	$\sigma_{4n} = 2.6^{+3.3}_{-1.7}$	[13]	
	0.7546	0.9180	0.8879	38.1	$\sigma_{3n} = 2.04$	36.5-41.1	$\sigma_{3n} = 2.5^{+2.9}_{-1.4}$	[13]	
				43.1	$\sigma_{4n}=0.65$	40.9-45.4	$\sigma_{4n} = 2.6^{+3.3}_{-1.7}$	[13]	
${}^{48}Ca + {}^{242}Pu$	4.0	1.460734	1.1936	32.8	$\sigma_{2n} = 9.38$	30.4-34.7	$\sigma_{2n} = 0.5^{+1.4}_{-0.4}$	[<mark>1</mark>]	
				39.8	$\sigma_{3n} = 3.81$	38-42.4	$\sigma_{3n} = 1.4^{+3.2}_{-1.2}$	[<mark>10</mark>]	
				40.8	$\sigma_{4n} = 3.90$	38-42.4	$\sigma_{4n} = 1.4^{+3.2}_{-1.2}$	[<mark>10</mark>]	
				39.8	$\sigma_{3n} = 3.81$	38-42.4	$\sigma_{3n} = 3.6^{+3.4}_{-1.7}$	[<mark>1</mark>]	
				40.8	$\sigma_{4n} = 3.90$	38-42.4	$\sigma_{4n} = 4.5^{+3.6}_{-1.9}$	[<mark>1</mark>]	
				49.8	$\sigma_{4n} = 0.19$	48-52	$\sigma_{4n} = 0.6^{+0.9}_{-0.5}$	[<mark>9</mark>]	
				50.8	$\sigma_{5n}=0.008$	48-52	$\sigma_{5n} = 0.6^{+0.9}_{-0.5}$	[<mark>9</mark>]	
	0.7546	0.9180	0.8863	32.8	$\sigma_{2n} = 0.02$	30.4-34.7	$\sigma_{2n} = 0.5^{+1.4}_{-0.4}$	[<mark>1</mark>]	
				39.8	$\sigma_{3n} = 1.77$	38-42.4	$\sigma_{3n} = 1.4^{+3.2}_{-1.2}$	[<mark>10</mark>]	
				40.8	$\sigma_{4n} = 2.02$	38-42.4	$\sigma_{4n} = 1.4^{+3.2}_{-1.2}$	[<mark>10</mark>]	
				39.8	$\sigma_{3n} = 1.77$	38-42.4	$\sigma_{3n} = 3.6^{+3.4}_{-1.7}$	[<mark>1</mark>]	
				40.8	$\sigma_{4n} = 2.02$	38-42.4	$\sigma_{4n} = 4.5^{+3.6}_{-1.9}$	[1]	
				49.8	$\sigma_{4n} = 0.14$	48-52	$\sigma_{4n} = 0.6^{+0.9}_{-0.5}$	[<mark>9</mark>]	
				50.8	$\sigma_{5n} = 0.006$	48-52	$\sigma_{5n} = 0.6^{+0.9}_{-0.5}$	[<mark>9</mark>]	
${}^{48}Ca + {}^{244}Pu$	4.0	1.460734	1.1800	37.4	$\sigma_{3n} = 11.15$	36.1-39.5	$\sigma_{3n} = 8.0^{+7.4}_{-4.5}$	[11]	
				41.4	$\sigma_{3n} = 2.86$	39.8-43.9	$\sigma_{3n} = 3.5^{+3.3}_{-2.0}$	[7]	
				42.4	$\sigma_{3n} = 1.95$	39-43	$\sigma_{3n} = 1.7^{+2.5}_{-1.1}$	[15]	
				37.4	$\sigma_{4n} = 19.09$	36.1-40.1	$\sigma_{4n} = 2.8^{+4.2}_{-2.1}$	[7]	
				40.4	$\sigma_{4n} = 9.42$	39.8-43.9	$\sigma_{4n} = 11^{+15}_{-7}$	[7]	
				41.4	$\sigma_{4n} = 6.94$	39.8-43.9	$\sigma_{4n} = 9.8^{+3.9}_{-3.1}$	[11]	
				42.4	$\sigma_{4n} = 5.01$	39-43	$\sigma_{4n} = 5.3^{+3.6}_{-2.1}$	[15]	
				52.4	$\sigma_{5n}=0.01$	50.4-54.7	$\sigma_{5n} = 1.1^{+2.6}_{-0.9}$	[15]	
	0.7546	0.9180	0.8847	37.4	$\sigma_{3n} = 1.93$	36.1-39.5	$\sigma_{3n} = 8.0^{+7.4}_{-4.5}$	[11]	
				41.4	$\sigma_{3n} = 1.41$	39.8-43.9	$\sigma_{3n} = 3.5^{+3.3}_{-2.0}$	[7]	
				42.4	$\sigma_{3n} = 1.06$	39-43	$\sigma_{3n} = 1.7^{+2.5}_{-1.1}$	[15]	
				37.4	$\sigma_{4n} = 3.32$	36.1-40.1	$\sigma_{4n} = 2.8^{+4.2}_{-2.1}$	[7]	
				40.4	$\sigma_{4n} = 4.10$	39.8-43.9	$\sigma_{4n} = 11^{+15}_{-7}$	[7]	

R. ZARGINI AND S. A. SEYYEDI

			TABLE	II. (Continued.)				
			Calculated values			Experimental eata		
Combination	k_s	$\gamma_0 ({\rm MeVfm^{-2}})$	γ (MeV fm ⁻²)	E_{CN}^{*} (MeV)	σ_{ER} (pb)	$\overline{E_{CN}^{*}}$ (MeV)	σ_{ER} (pb)	Ref
				41.4	$\sigma_{4n} = 3.46$	39.8-43.9	$\sigma_{4n} = 9.8^{+3.9}_{-3.1}$	[11]
				42.4	$\sigma_{4n} = 2.73$	39-43	$\sigma_{4n} = 5.3^{+3.6}_{-2.1}$	[15]
				52.4	$\sigma_{5n} = 0.008$	50.4-54.7	$\sigma_{5n} = 1.1^{+2.6}_{-0.9}$	[15]
${}^{48}\text{Ca} + {}^{243}\text{Am}$	4.0	1.460734	1.2039	36.5	$\sigma_{3n} = 12.35$	34-38.3	$\sigma_{3n} = 8.5^{+6.4}_{-3.7}$	[<mark>16</mark>]
				39.5	$\sigma_{3n} = 4.07$	38-43.2	$\sigma_{3n} = 3.7^{+1.3}_{-1.0}$	[3]
				40.5	$\sigma_{3n} = 2.74$	38-43.2	$\sigma_{3n} = 2.7^{+4.8}_{-1.6}$	[3]
				44.5	$\sigma_{4n} = 0.78$	42.4-46.5	$\sigma_{4n} = 0.9^{+3.2}_{-0.8}$	[3]
	0.7546	0.9180	0.8875	36.5	$\sigma_{3n} = 3.89$	34-38.3	$\sigma_{3n} = 8.5^{+6.4}_{-3.7}$	[<mark>16</mark>]
				39.5	$\sigma_{3n} = 2.06$	38-43.2	$\sigma_{3n} = 3.7^{+1.3}_{-1.0}$	[3]
				40.5	$\sigma_{3n} = 1.51$	38-43.2	$\sigma_{3n} = 2.7^{+4.8}_{-1.6}$	[3]
				44.5	$\sigma_{4n} = 0.53$	42.4-46.5	$\sigma_{4n} = 0.9^{+3.2}_{-0.8}$	[3]
${}^{48}\text{Ca} + {}^{245}\text{Cm}$	4.0	1.460734	1.2074	32.7	$\sigma_{2n} = 12.97$	30.9-35	$\sigma_{2n} = 0.9^{+1.1}_{-0.6}$	[15]
				37.7	$\sigma_{2n} = 2.06$	35.9-39.9	$\sigma_{2n} = 0.7^{+2.0}_{-0.6}$	[4]
				32.7	$\sigma_{3n} = 197.3$	30.9-35	$\sigma_{3n} = 1.3^{+1.2}_{-0.7}$	[15]
				41.7	$\sigma_{3n} = 6.45$	40.7-44.8	$\sigma_{3n} = 1.9^{+2.1}_{-1.0}$	[4]
				42.7	$\sigma_{3n} = 4.18$	40.7-44.8	$\sigma_{3n} = 3.7^{+3.6}_{-1.8}$	[4]
				42.7	$\sigma_{4n} = 0.97$	40.7-44.8	$\sigma_{4n} = 1.0^{+0.0}_{-0.0}$	[4]
	0.7546	0.9180	0.8879	32.7	$\sigma_{2n} = 0.76$	30.9-35	$\sigma_{2n} = 0.9^{+1.1}_{-0.6}$	[15]
				37.7	$\sigma_{2n} = 0.96$	35.9-39.9	$\sigma_{2n} = 0.7^{+2.0}_{-0.6}$	[4]
				32.7	$\sigma_{3n} = 11.63$	30.9-35	$\sigma_{3n} = 1.3^{+1.2}_{-0.7}$	[15]
				41.7	$\sigma_{3n} = 4.06$	40.7-44.8	$\sigma_{3n} = 1.9^{+2.1}_{-1.0}$	[4]
				42.7	$\sigma_{3n} = 2.75$	40.7-44.8	$\sigma_{3n} = 3.7^{+3.6}_{-1.8}$	[4]
				42.7	$\sigma_{4n} = 0.64$	40.7-44.8	$\sigma_{4n} = 1.0^{+0.0}_{-0.0}$	[4]
${}^{48}Ca + {}^{248}Cm$	4.0	1.460734	1.1875	30.4	$\sigma_{3n} = 19.51$	30.5	$\sigma_{3n} = 0.9$	[1]
				33.4	$\sigma_{3n} = 26.96$	33	$\sigma_{3n} = 0.5^{+0.5}_{-0.26}$	[<mark>1</mark>]
				38.4	$\sigma_{3n} = 5.25$	36.8-41.1	$\sigma_{3n} = 1.1^{+1.7}_{-0.7}$	[14]
				40.4	$\sigma_{3n} = 2.32$	40.9	$\sigma_{3n} = 0.9^{+2.1}_{-0.7}$	[14]
				33.4	$\sigma_{4n} = 22.13$	33	$\sigma_{4n} = 0.3$	[1]
				38.4	$\sigma_{4n} = 9.60$	36.8-41.1	$\sigma_{4n} = 3.3^{+2.5}_{-1.4}$	[14]
				40.4	$\sigma_{4n} = 5.12$	40.9	$\sigma_{4n} = 3.4^{+2.7}_{-1.6}$	[14]
	0.7546	0.9180	0.8856	30.4	$\sigma_{3n} = 0.014$	30.5	$\sigma_{3n} = 0.9$	[1]
				33.4	$\sigma_{3n}=0.9$	33	$\sigma_{3n} = 0.5^{+0.5}_{-0.26}$	[1]
				38.4	$\sigma_{3n} = 2.38$	36.8-41.1	$\sigma_{3n} = 1.1^{+1.7}_{-0.7}$	[14]
				40.4	$\sigma_{3n} = 1.29$	40.9	$\sigma_{3n} = 0.9^{+2.1}_{-0.7}$	[14]
				33.4	$\sigma_{4n} = 0.7$	33	$\sigma_{4n} = 0.3$	[<mark>1</mark>]
				38.4	$\sigma_{4n} = 4.36$	36.8-41.1	$\sigma_{4n} = 3.3^{+2.5}_{-1.4}$	[14]
				40.4	$\sigma_{4n}=2.84$	40.9	$\sigma_{4n} = 3.4^{+2.7}_{-1.6}$	[14]
${}^{48}Ca + {}^{249}Bk$	4.0	1.460734	1.1978	32.8	$\sigma_{3n} = 35.24$	30.4-34.7	$\sigma_{3n} = 0.7^{+1.7}_{-0.57}$	[<mark>6</mark>]
				34.8	$\sigma_{3n} = 18.11$	33.2-37.4	$\sigma_{3n} = 0.5^{+1.1}_{-0.4}$	[5]
				34.8	$\sigma_{3n} = 18.11$	35	$\sigma_{3n} = 3.6^{+6.1}_{-2.5}$	[<mark>8</mark>]
				34.8	$\sigma_{3n} = 18.11$	32.8-37.5	$\sigma_{3n} = 1.1^{+1.2}_{-0.6}$	[<mark>6</mark>]
				38.8	$\sigma_{3n} = 3.57$	39	$\sigma_{3n}=0.7$	[5]
				38.8	$\sigma_{3n} = 3.57$	39	$\sigma_{3n} = 0.32$	[<mark>6</mark>]
				34.8	$\sigma_{4n} = 11.41$	35	$\sigma_{4n} = 0.8$	[5]

			IABLE	II. (Continuea.)			
			Calculated values			Experimental eata		
Combination	k_s	$\gamma_0 ({\rm MeVfm^{-2}})$	γ (MeV fm ⁻²)	E_{CN}^{*} (MeV)	σ_{ER} (pb)	$\overline{E_{CN}^{*} (\mathrm{MeV})}$	σ_{ER} (pb)	Ref.
				34.8	$\sigma_{4n} = 11.41$	35	$\sigma_{4n} = 0.59$	[6]
				38.8	$\sigma_{4n} = 4.20$	37.2-41.4	$\sigma_{4n} = 1.3^{+1.5}_{-0.6}$	[5]
				38.8	$\sigma_{4n} = 4.20$	39	$\sigma_{4n} = 2.0^{+2.2}_{-1.1}$	[<mark>8</mark>]
				38.8	$\sigma_{4n} = 4.20$	37-41.9	$\sigma_{4n} = 1.5^{+1.1}_{-0.5}$	[<mark>6</mark>]
				42.8	$\sigma_{4n} = 1.06$	40.3-44.8	$\sigma_{4n} = 2.4^{+3.3}_{-1.4}$	[<mark>6</mark>]
				45.8	$\sigma_{4n}=0.34$	43.8-48.3	$\sigma_{4n} = 2.0^{+1.8}_{-1.1}$	[<mark>6</mark>]
	0.7546	0.9180	0.8868	32.8	$\sigma_{3n} = 3.70$	30.4-34.7	$\sigma_{3n} = 0.7^{+1.7}_{-0.57}$	[<mark>6</mark>]
				34.8	$\sigma_{3n} = 5.45$	33.2-37.4	$\sigma_{3n} = 0.5^{+1.1}_{-0.4}$	[5]
				34.8	$\sigma_{3n} = 5.45$	35	$\sigma_{3n} = 3.6^{+6.1}_{-2.5}$	[<mark>8</mark>]
				34.8	$\sigma_{3n} = 5.45$	32.8-37.5	$\sigma_{3n} = 1.1^{+1.2}_{-0.6}$	[<mark>6</mark>]
				38.8	$\sigma_{3n} = 1.94$	39	$\sigma_{3n}=0.7$	[5]
				38.8	$\sigma_{3n} = 1.94$	39	$\sigma_{3n} = 0.32$	[<mark>6</mark>]
				34.8	$\sigma_{4n} = 3.43$	35	$\sigma_{4n}=0.8$	[5]
				34.8	$\sigma_{4n} = 3.43$	35	$\sigma_{4n} = 0.59$	[<mark>6</mark>]
				38.8	$\sigma_{4n} = 2.29$	37.2-41.4	$\sigma_{4n} = 1.3^{+1.5}_{-0.6}$	[5]
				38.8	$\sigma_{4n} = 2.29$	39	$\sigma_{4n} = 2.0^{+2.2}_{-1.1}$	[<mark>8</mark>]
				38.8	$\sigma_{4n}=2.29$	37-41.9	$\sigma_{4n} = 1.5^{+1.1}_{-0.5}$	[<mark>6</mark>]
				42.8	$\sigma_{4n} = 0.71$	40.3-44.8	$\sigma_{4n} = 2.4^{+3.3}_{-1.4}$	[<mark>6</mark>]
				45.8	$\sigma_{4n} = 0.25$	43.8-48.3	$\sigma_{4n} = 2.0^{+1.8}_{-1.1}$	[<mark>6</mark>]
⁴⁸ Ca + ²⁴⁹ Cf	4.0	1.460734	1.2142	29.6	$\sigma_{3n} = 168.2$	26.6-31.7	$\sigma_{3n} = 0.3^{+1.0}_{-0.27}$	[4]
				34.6	$\sigma_{3n} = 38.06$	32.1-36.6	$\sigma_{3n} = 0.5^{+1.6}_{-0.3}$	[4]
	0.7546	0.9180	0.8887	29.6	$\sigma_{3n} = 3.77$	26.6-31.7	$\sigma_{3n} = 0.3^{+1.0}_{-0.27}$	[4]
				34.6	$\sigma_{3n} = 15.93$	32.1-36.6	$\sigma_{3n} = 0.5^{+1.6}_{-0.3}$	[4]

TABLE II. (Continued.)

along with collision angles for ${}^{249}Cf({}^{45}Sc, xn){}^{294-x}119$ combination. Figure 2 shows the potential distribution for the rotation of colliding nuclei, α_i (see Fig. 1). One can consider for collisions of $\alpha_1 = 90^\circ$ and $\alpha_2 = 90^\circ$, the maximum potential, and for collisions of $\alpha_1 = 0^\circ$ and $\alpha_2 = 0^\circ$, the minimum potential occurs. However, for collisions of $\alpha_1 = 90^\circ$ and $\alpha_2 = 90^\circ$ the quasifission barrier (B_{qf}) is very low, therefore the survival probability of CN is very small. It means that in competition between the neutron evaporation and changing to the fission fragments, the fission fragments would occur, and the ER cross section will be low. Figure 3 shows the potential distribution for the direction of colliding nuclei, θ_i . It is clear for collisions of $\theta_1 = 0^\circ$ and $\theta_2 = 0^\circ$, the maximum potential, and for collisions of $\theta_1 = 90^\circ$ and $\theta_2 = 90^\circ$, the minimum potential occurs. It is considered for collisions of $\theta_1 = 0^\circ$ and $\theta_2 = 0^\circ$ the B_{qf} is smaller than collisions of $\theta_1 = 90^\circ$ and $\theta_2 =$ 90°. It means that for a collision of $\theta_1 = 0^\circ$ and $\theta_2 = 0^\circ$ the probability of changing to fission fragments is more than the neutron evaporation, and of course, the ER cross section also will be low. In actual conditions, nuclei can collide with each other, in different rotation and direction, therefore the average of the capture cross section will be considered.

The surface energy coefficient is given by Eq. (4). Based on previous studies, 13 different values have been provided for k_s and γ_0 [48–55]. Those coefficients are listed in Table I. The combinations which lead to the synthesis of SHN with Z = 112-118 were chosen. These combinations included 238 U(48 Ca, xn) ${}^{286-x}$ 112, 237 Np(48 Ca, xn) ${}^{285-x}$ 113, 242 Pu(48 Ca, xn) $^{290-x}$ 114, 243 Am(48 Ca, xn) $^{291-x}$ 115, 248 Cm(48 Ca, xn) $^{291-x}$ 116, 240 Pu(48 Ca, *xn*) $^{288-x}$ 114, 244 Pu(48 Ca, *xn*) $^{292-x}$ 114, 245 Cm(48 Ca, xn) $^{293-x}$ 116, 249 Bk $(^{48}$ Ca, xn $)^{297-x}$ 117, 249 Cf $(^{48}$ Ca, xn $)^{297-x}$ 118. Using the values listed for k_s and γ_0 in Table I, the values for the surface energy coefficient, the CN excitation energy, and the ER cross section were calculated. Some selected calculated results along with the experimental data are listed in Table II. The selected results represent the values of k_s and γ_0 by which the highest agreement ($k_s = 0.7546$ and $\gamma_0 = 0.9180 \,\mathrm{MeV}\,\mathrm{fm}^{-2}$) and lowest agreement ($k_s = 4.0$ and $\gamma_0 = 1.460734 \,\mathrm{MeV}\,\mathrm{fm}^{-2}$) between the experimental data and calculated values were achieved.

From Table II, one can conclude that the calculated ER cross sections with the defined values of $k_s = 0.7546$ and $\gamma_0 = 0.9180 \text{ MeV } \text{fm}^{-2}$ is in good agreement with the experimental data. Figure 4 represents the ER cross sections versus the excitation energies for combinations of $^{238}\text{U}(^{48}\text{Ca}, xn)^{286-x}112$, $^{237}\text{Np}(^{48}\text{Ca}, xn)^{285-x}113$, $^{242}\text{Pu}(^{48}\text{Ca}, xn)^{290-x}114$,



FIG. 4. The ER cross section vs. the CN excitation energy for combinations of (a) 238 U(48 Ca, xn) $^{286-x}112$, (b) 237 Np(48 Ca, xn) $^{285-x}113$, (c) 240 Pu(48 Ca, xn) $^{288-x}114$, (d) 242 Pu(48 Ca, xn) $^{290-x}114$, (e) 244 Pu(48 Ca, xn) $^{292-x}114$, (f) 243 Am(48 Ca, xn) $^{291-x}115$, (g) 245 Cm(48 Ca, xn) $^{293-x}116$, (h) 248 Cm(48 Ca, xn) $^{296-x}116$, (i) 249 Bk(48 Ca, xn) $^{297-x}117$, and (j) 249 Cf(48 Ca, xn) $^{297-x}118$. The thin lines included 2*n* solid (blue), 3*n* dash (magenta), 4*n* dot (olive), and 5*n* short dash dot (orange), denote the calculated cross sections for values of $k_s = 4.0$ and $\gamma_0 = 1.460734$ MeV fm⁻² and the thick lines included 2*n* dash dot dot (blue), 3*n* short dash (magenta), 4*n* short dot (olive), and 5*n* solid (orange), show the calculated cross sections for values of $k_s = 0.7546$ and $\gamma_0 = 0.9180$ MeV fm⁻². The experimental data are given by Refs. [1-16].



FIG. 5. The ER cross sections vs. the excitation energies of the CN for combinations of (a) ${}^{249}Cf({}^{45}Sc, xn)^{294-x}119$, (b) ${}^{249}Bk({}^{50}Ti, xn)^{299-x}119$, (c) ${}^{247}Bk({}^{50}Ti, xn)^{297-x}119$, (d) ${}^{248}Cm({}^{51}V, xn)^{299-x}119$, (e) ${}^{243}Am({}^{54}Cr, xn)^{297-x}119$, (f) ${}^{242}Pu({}^{55}Mn, xn)^{297-x}119$, (g) ${}^{254}Es({}^{48}Ca, xn)^{302-x}119$.

²⁴⁴Pu(⁴⁸Ca, xn)^{292-x}114, ²⁴³Am(⁴⁸Ca, xn)^{291-x}115, ²⁴⁵Cm(⁴⁸Ca, xn)^{293-x}116, ²⁴⁸Cm(⁴⁸Ca, xn)^{296-x}116, ²⁴⁹Bk(⁴⁸Ca, xn)^{297-x}117, ²⁴⁹Cf(⁴⁸Ca, xn)^{297-x}118.

B. Probability of synthesis of SHN with Z = 119,120

Many attempts are in progress to synthesize SHN with Z = 119, 120 [63]. In this part, using the obtained model in Sec. III A, the synthesis probability of SHN with Z = 119, 120 was investigated by calculating the ER cross sections for several combinations. To synthesize nuclei beyond Z = 118, targets heavier than californium (Cf) should be used. To produce the nuclei with Z = 119 the seven combinations, namely 249 Cf (45 Sc, xn) $^{294-x}119$, 249 Bk(50 Ti, xn) $^{299-x}119$, 247 Bk(50 Ti, xn) $^{297-x}119$, 242 Pu(55 Mn, xn) $^{299-x}119$, and 254 Es(48 Ca, xn) $^{302-x}119$ were initially selected. In these

combinations, the projectiles heavier than ⁴⁸Ca, with stable nuclei (⁴⁵Sc, ⁵⁰Ti, ⁵¹V, ⁵⁴Cr, and ⁵⁵Mn), were chosen. The actinide targets which had already been employed to the synthesis of SHN with Z = 114-118 were used. The ²⁵⁴Es(⁴⁸Ca, xn)^{302-x}119 combination was chosen due to the potential interest for ²⁵⁴Es as a new target in the near future [59]. The results are presented in Fig. 5 and the optimum conditions are summarized in Table III. From Table III one can consider that the most promising combinations for synthesis of SHN with Z = 119 are ²⁴⁹Cf(⁴⁵Sc, 3n)²⁹¹119, ²⁴⁹Cf(⁴⁵Sc, 4n)²⁹⁰119, ²⁴⁷Bk(⁵⁰Ti, 3n)²⁹⁴119, ²⁵⁴Es(⁴⁸Ca, 3n)²⁹⁹119, and ²⁵⁴Es(⁴⁸Ca, 4n)²⁹⁸119.

For SHN with Z = 120, six combinations included ${}^{254}\text{Es}({}^{45}\text{Sc}, xn){}^{299-x}120$, ${}^{249}\text{Cf}({}^{50}\text{Ti}, xn){}^{299-x}120$, ${}^{251}\text{Cf}({}^{50}\text{Ti}, xn){}^{301-x}120$, ${}^{249}\text{Bk}({}^{51}\text{V}, xn){}^{300-x}120$, ${}^{248}\text{Cm}({}^{54}\text{Cr}, xn){}^{302-x}120$, and ${}^{243}\text{Am}({}^{55}\text{Mn}, xn){}^{298-x}120$

TABLE III. Combinations, excitation energies of the CN, and the maximum ER cross sections in three and four neutron channels to synthesis of nuclei with Z = 119.

Combination	E_{CN}^{*} (MeV)	σ_{ER} (fb)
249 Cf(45 Sc, $3n$) 291 119	39.84	$\sigma_{3n} = 417.1$
249 Cf (45 Sc, $4n$) 290 119	41.84	$\sigma_{4n} = 138.5$
²⁴⁹ Bk(⁵⁰ Ti, 3n) ²⁹⁶ 119	35.60	$\sigma_{3n} = 7.96$
²⁴⁹ Bk(⁵⁰ Ti, 4n) ²⁹⁵ 119	37.60	$\sigma_{4n} = 3.31$
²⁴⁷ Bk(⁵⁰ Ti, 3n) ²⁹⁴ 119	35.15	$\sigma_{3n} = 11.2$
²⁴⁷ Bk(⁵⁰ Ti, 4 <i>n</i>) ²⁹³ 119	37.15	$\sigma_{4n} = 2.46$
248 Cm $(^{51}$ V $, 3n)^{296}$ 119	40.37	$\sigma_{3n} = 0.17$
248 Cm $(^{51}$ V $, 4n)^{295}$ 119	41.37	$\sigma_{4n} = 0.13$
243 Am $(^{54}$ Cr, $3n)^{294}$ 119	33.34	$\sigma_{3n} = 0.8$
243 Am $(^{54}$ Cr, $4n)^{293}$ 119	36.34	$\sigma_{4n} = 0.12$
²⁴² Pu(⁵⁵ Mn, 3 <i>n</i>) ²⁹⁴ 119	37.1	$\sigma_{3n} = 0.02$
²⁵⁴ Es(⁴⁸ Ca, 3 <i>n</i>) ²⁹⁹ 119	32.14	$\sigma_{3n} = 9115.15$
254 Es(48 Ca, $4n$) 298 119	34.14	$\sigma_{4n} = 735.46$

TABLE IV. Combinations, excitation energies of the CN, and the maximum ER cross sections in three and four neutron channels to synthesize nuclei with Z = 120.

Combination	E_{CN}^{*} (MeV)	σ_{ER} (fb)
254 Es(45 Sc, $4n$) 295 120	49.81	$\sigma_{4n} = 0.05$
²⁴⁹ Cf(⁵⁰ Ti, 3 <i>n</i>) ²⁹⁶ 120	33.19	$\sigma_{3n} = 51.19$
²⁴⁹ Cf(⁵⁰ Ti, 4n) ²⁹⁵ 120	36.19	$\sigma_{4n} = 1.05$
²⁵¹ Cf(⁵⁰ Ti, 3 <i>n</i>) ²⁹⁸ 120	33.82	$\sigma_{3n} = 43.17$
²⁵¹ Cf(⁵⁰ Ti, 4 <i>n</i>) ²⁹⁷ 120	36.82	$\sigma_{4n} = 1.14$
249 Bk(51 V, $3n$) 297 120	38.86	$\sigma_{3n} = 0.06$
249 Bk(51 V, $4n$) 296 120	40.86	$\sigma_{4n} = 0.03$
²⁴⁸ Cm(⁵⁴ Cr, 3 <i>n</i>) ²⁹⁹ 120	33.58	$\sigma_{3n} = 0.19$
²⁴³ Am(⁵⁵ Mn, 3 <i>n</i>) ²⁹⁵ 120	36.26	$\sigma_{4n} = 0.007$

previously reported by other research teams are summarized in Table V.

IV. SUMMARY

were investigated, The results are presented in Fig. 6 and the ideal conditions are summarized in Table IV. From Table IV one can consider that the most promising combinations to synthesize nuclei with Z = 120 are ²⁴⁹Cf(⁵⁰Ti, 3*n*)²⁹⁶120 and ²⁵¹Cf(⁵⁰Ti, 3*n*)²⁹⁸120.

The obtained ER cross sections (in femtobarn) from this work in comparison with some common combinations

In this research the roles of the different surface energy coefficients on the compound nucleus decay modes were investigated. The fuscous was given to the superheavy nuclei in the range of Z = 112-118 which were synthesized via heavy ion fusion reactions. With employing a dinuclear system model, proximity potential, and considering deformed nuclei, the evaporation residue cross sections were



FIG. 6. The ER cross section vs. the excitations energies of the CN for combinations of (a) ${}^{254}\text{Es}({}^{45}\text{Sc}, xn)^{299-x}120$, (b) ${}^{249}\text{Cf}({}^{50}\text{Ti}, xn)^{299-x}120$, (c) ${}^{251}\text{Cf}({}^{50}\text{Ti}, xn)^{301-x}120$, (d) ${}^{249}\text{Bk}({}^{51}\text{V}, xn)^{300-x}120$, (e) ${}^{248}\text{Cm}({}^{54}\text{Cr}, xn)^{302-x}120$, (f) ${}^{243}\text{Am}({}^{55}\text{Mn}, xn)^{298-x}120$.

45 Sc + 249 Cf	50 Ti + 247 Bk	50 Ti + 249 Bk	$^{50}\text{Ti} + ^{249}\text{Cf}$	$^{50}{ m Ti} + {}^{251}{ m Cf}$	Ref.
417.1(3 <i>n</i>), 138.5(4 <i>n</i>)	11.2(3 <i>n</i>), 2.46(4 <i>n</i>)	7.96(3 <i>n</i>), 3.31(4 <i>n</i>)	51.2(3 <i>n</i>), 1(4 <i>n</i>)	43.2(3 <i>n</i>), 1.1(4 <i>n</i>)	This work
5.47(3 <i>n</i>), 0.72(4 <i>n</i>)		48.2(3 <i>n</i>), 5.67(4 <i>n</i>)	7.7(3 <i>n</i>)	17.2(3 <i>n</i>)	[38]
	43(3 <i>n</i>)				[29]
		12(3 <i>n</i>), 64(4 <i>n</i>)			[26]
		40(3), 55(4 <i>n</i>)	40(3 <i>n</i>), 46(4 <i>n</i>)		[40]
		35(3 <i>n</i>), 35(4 <i>n</i>)	20(3n), 4.5(4n)		[65]
		60(3 <i>n</i>), 110(4 <i>n</i>)	50(3n), 3.5(4n)		[<mark>66</mark>]
		480(3 <i>n</i>), 310(4 <i>n</i>)	150(3 <i>n</i>), 50(4 <i>n</i>)		[<mark>67</mark>]
		33(3 <i>n</i>), 15(4 <i>n</i>)			[29]
		40(3 <i>n</i>), 150(4 <i>n</i>)			[27]
		340(3 <i>n</i>), 160(4 <i>n</i>)	34(3 <i>n</i>), 2(4 <i>n</i>)	120(3 <i>n</i>), 4(4 <i>n</i>)	[26]
		30(3 <i>n</i>), 36(4 <i>n</i>)	5.5(3 <i>n</i>), 6.1(4 <i>n</i>)		[39]
		105.1(3 <i>n</i>), 40.9(4 <i>n</i>)			[34]
			100(3 <i>n</i>), 2.5(4 <i>n</i>)		[31]
			760(3 <i>n</i>), 28(4 <i>n</i>)		[46]
			20(3 <i>n</i>), 21(2 <i>n</i>)		[41]
			102.6(3 <i>n</i>), 2.3(4 <i>n</i>)		[32]

TABLE V. The comparison between the results from current study with previously reported results employing different models.

calculated. Comparing results from performed calculations and experimental data, the best values for surface asymmetric constant, k_s and surface energy constant, γ_0 were obtained to be 0.7546 and 0.9180 MeV fm⁻², respectively. Furthermore, this new model was used to investigate the probability of synthesis of superheavy nuclei with Z = 119, 120. The most promising combinations to synthesize nuclei with Z =119 are ²⁴⁹Cf(⁴⁵Sc, 3n)²⁹¹119 with the ER cross section, $\sigma_{3n} = 417.1$ fb at the incident energy $E_{c.m.} = 219$ MeV($E^* =$ 39.84 MeV), ²⁴⁹Cf(⁴⁵Sc, 4n)²⁹⁰119 with the ER cross section, $\sigma_{4n} = 138.5$ fb at the incident energy $E_{c.m.} = 221$ MeV($E^* =$ 41.84 MeV), ²⁴⁷Bk(⁵⁰Ti, 3n)²⁹⁴119 with the ER cross section, $\sigma_{3n} = 11.2$ fb at the incident energy $E_{c.m.} = 226 \text{ MeV}(E^* = 35.15 \text{ MeV})$, $^{254}\text{Es}(^{48}\text{Ca}, 3n)^{299}119$ with the ER cross section, $\sigma_{3n} = 9115.15$ fb at the incident energies $E_{c.m.} = 208 \text{ MeV}(E^* = 32.14 \text{ MeV})$, and $^{254}\text{Es}(^{48}\text{Ca}, 4n)^{298}119$ with the ER cross section, $\sigma_{4n} = 735.46$ fb at the incident energies $E_{c.m.} = 210 \text{ MeV}(E^* = 34.14 \text{ MeV})$. In addition, it was found out that the best combinations to synthesis of SHN with Z = 120 are $^{249}\text{Cf}(^{50}\text{Ti}, 3n)^{296}120$ with the ER cross section, $\sigma_{3n} = 51.19$ fb at the incident energy $E_{c.m.} = 228 \text{ MeV}(E^* = 33.19 \text{ MeV})$, and $^{251}\text{Cf}(^{50}\text{Ti}, 3n)^{298}120$ with the ER cross section, $\sigma_{3n} = 43.17$ fb at the incident energy $E_{c.m.} = 227 \text{ MeV}(E^* = 33.82 \text{ MeV})$.

- Y. T. Oganessian, V. K. Utyonkov, Y. V. Lobanov, F. S. Abdullin, A. N. Polyakov, I. V. Shirokovsky, Y. S. Tsyganov, G. G. Gulbekian, S. L. Bogomolov, B. N. Gikal, A. N. Mezentsev, S. Iliev, V. G. Subbotin, A. M. Sukhov, A. A. Voinov, G. V. Buklanov, K. Subotic, V. I. Zagrebaev, M. G. Itkis, J. B. Patin, K. J. Moody, J. F. Wild, M. A. Stoyer, N. J. Stoyer, D. A. Shaughnessy, J. M. Kenneally, P. A. Wilk, R. W. Lougheed, R. I. Il'kaev, and S. P. Vesnovskii, Phys. Rev. C 70, 064609 (2004).
- [2] Y. T. Oganessian, V. K. Utyonkov, Y. V. Lobanov, F. S. Abdullin, A. N. Polyakov, R. N. Sagaidak, I. V. Shirokovsky, Y. S. Tsyganov, A. A. Voinov, G. G. Gulbekian, S. L. Bogomolov, B. N. Gikal, A. N. Mezentsev, V. G. Subbotin, A. M. Sukhov, K. Subotic, V. I. Zagrebaev, G. K. Vostokin, M. G. Itkis, R. A. Henderson, J. M. Kenneally, J. H. Landrum, K. J. Moody, D. A. Shaughnessy, M. A. Stoyer, N. J. Stoyer, and P. A. Wilk, Phys. Rev. C **76**, 011601(R) (2007).
- [3] Y. T. Oganessian, V. K. Utyonkov, S. N. Dmitriev, Y. V. Lobanov, M. G. Itkis, A. N. Polyakov, Y. S. Tsyganov, A. N. Mezentsev, A. V. Yeremin, A. A. Voinov, E. A. Sokol, G. G.

Gulbekian, S. L. Bogomolov, S. Iliev, V. G. Subbotin, A. M. Sukhov, G. V. Buklanov, S. V. Shishkin, V. I. Chepygin, G. K. Vostokin, N. V. Aksenov, M. Hussonnois, K. Subotic, V. I. Zagrebaev, K. J. Moody, J. B. Patin, J. F. Wild, M. A. Stoyer, N. J. Stoyer, D. A. Shaughnessy, J. M. Kenneally, P. A. Wilk, R. W. Lougheed, H. W. Gäggeler, D. Schumann, H. Bruchertseifer, and R. Eichler, Phys. Rev. C **72**, 034611 (2005).

- [4] Y. T. Oganessian, V. K. Utyonkov, Y. V. Lobanov, F. S. Abdullin, A. N. Polyakov, R. N. Sagaidak, I. V. Shirokovsky, Y. S. Tsyganov, A. A. Voinov, G. G. Gulbekian, S. L. Bogomolov, B. N. Gikal, A. N. Mezentsev, S. Iliev, V. G. Subbotin, A. M. Sukhov, K. Subotic, V. I. Zagrebaev, G. K. Vostokin, M. G. Itkis, K. J. Moody, J. B. Patin, D. A. Shaughnessy, M. A. Stoyer, N. J. Stoyer, P. A. Wilk, J. M. Kenneally, J. H. Landrum, J. F. Wild, and R. W. Lougheed, Phys. Rev. C 74, 044602 (2006).
- [5] Y. T. Oganessian, F. S. Abdullin, P. D. Bailey, D. E. Benker, M. E. Bennett, S. N. Dmitriev, J. G. Ezold, J. H. Hamilton, R. A. Henderson, M. G. Itkis, Y. V. Lobanov, A. N. Mezentsev, K. J. Moody, S. L. Nelson, A. N. Polyakov, C. E. Porter,

A. V. Ramayya, F. D. Riley, J. B. Roberto, M. A. Ryabinin, K. P. Rykaczewski, R. N. Sagaidak, D. A. Shaughnessy, I. V. Shirokovsky, M. A. Stoyer, V. G. Subbotin, R. Sudowe, A. M. Sukhov, R. Taylor, Y. S. Tsyganov, V. K. Utyonkov, A. A. Voinov, G. K. Vostokin, and P. A. Wilk, Phys. Rev. C 83, 054315 (2011).

- [6] Y. T. Oganessian, F. S. Abdullin, C. Alexander, J. Binder, R. A. Boll, S. N. Dmitriev, J. Ezold, K. Felker, J. M. Gostic, R. K. Grzywacz, J. H. Hamilton, R. A. Henderson, M. G. Itkis, K. Miernik, D. Miller, K. J. Moody, A. N. Polyakov, A. V. Ramayya, J. B. Roberto, M. A. Ryabinin, K. P. Rykaczewski, R. N. Sagaidak, D. A. Shaughnessy, I. V. Shirokovsky, M. V. Shumeiko, M. A. Stoyer, N. J. Stoyer, V. G. Subbotin, A. M. Sukhov, Y. S. Tsyganov, V. K. Utyonkov, A. A. Voinov, and G. K. Vostokin, Phys. Rev. C 87, 054621 (2013).
- [7] J. M. Gates, C. E. Düllmann, M. Schädel, A. Yakushev, A. Türler, K. Eberhardt, J. V. Kratz, D. Ackermann, L. L. Andersson, M. Block, W. Brüchle, J. Dvorak, H. G. Essel, P. A. Ellison, J. Even, U. Forsberg, J. Gellanki, A. Gorshkov, R. Graeger, K. E. Gregorich, W. Hartmann, R. D. Herzberg, F. P. Heßberger, D. Hild, A. Hübner, E. Jäger, J. Khuyagbaatar, B. Kindler, J. Krier, N. Kurz, S. Lahiri, D. Liebe, B. Lommel, M. Maiti, H. Nitsche, J. P. Omtvedt, E. Parr, D. Rudolph, J. Runke, H. Schaffner, B. Schausten, E. Schimpf, A. Semchenkov, J. Steiner, P. Thörle-Pospiech, J. Uusitalo, M. Wegrzecki, and N. Wiehl, Phys. Rev. C 83, 054618 (2011).
- [8] Y. T. Oganessian, F. S. Abdullin, C. Alexander, J. Binder, R. A. Boll, S. N. Dmitriev, J. Ezold, K. Felker, J. M. Gostic, R. K. Grzywacz, J. H. Hamilton, R. A. Henderson, M. G. Itkis, K. Miernik, D. Miller, K. J. Moody, A. N. Polyakov, A. V. Ramayya, J. B. Roberto, M. A. Ryabinin, K. P. Rykaczewski, R. N. Sagaidak, D. A. Shaughnessy, I. V. Shirokovsky, M. V. Shumeiko, M. A. Stoyer, N. J. Stoyer, V. G. Subbotin, A. M. Sukhov, Y. S. Tsyganov, V. K. Utyonkov, A. A. Voinov, and G. K. Vostokin, Phys. Rev. Lett. 109, 162501 (2012).
- [9] P. A. Ellison, K. E. Gregorich, J. S. Berryman, D. L. Bleuel, R. M. Clark, I. Dragojević, J. Dvorak, P. Fallon, C. Fineman-Sotomayor, J. M. Gates, O. R. Gothe, I. Y. Lee, W. D. Loveland, J. P. McLaughlin, S. Paschalis, M. Petri, J. Qian, L. Stavsetra, M. Wiedeking, and H. Nitsche, Phys. Rev. Lett. **105**, 182701 (2010).
- [10] L. Stavsetra, K. E. Gregorich, J. Dvorak, P. A. Ellison, I. Dragojević, M. A. Garcia, and H. Nitsche, Phys. Rev. Lett. 103, 132502 (2009).
- [11] C. E. Düllmann, M. Schädel, A. Yakushev, A. Türler, K. Eberhardt, J. V. Kratz, D. Ackermann, L. L. Andersson, M. Block, W. Brüchle, J. Dvorak, H. G. Essel, P. A. Ellison, J. Even, J. M. Gates, A. Gorshkov, R. Graeger, K. E. Gregorich, W. Hartmann, R. D. Herzberg, F. P. Heßberger, D. Hild, A. Hübner, E. Jäger, J. Khuyagbaatar, B. Kindler, J. Krier, N. Kurz, S. Lahiri, D. Liebe, B. Lommel, M. Maiti, H. Nitsche, J. P. Omtvedt, E. Parr, D. Rudolph, J. Runke, B. Schausten, E. Schimpf, A. Semchenkov, J. Steiner, P. Thörle-Pospiech, J. Uusitalo, M. Wegrzecki, and N. Wiehl, Phys. Rev. Lett. 104, 252701 (2010).
- [12] S. Hofmann, D. Ackermann, S. Antalic, H. G. Burkhard, V. F. Comas, R. Dressler, Z. Gan, S. Heinz, J. A. Heredia, F. P. Heßberger, J. Khuyagbaatar, B. Kindler, I. Kojouharov, P. Kuusiniemi, M. Leino, B. Lommel, R. Mann, G. Münzenberg, K. Nishio, A. G. Popeko, S. Saro, H. J. Schött, B. Streicher,

B. Sulignano, J. Uusitalo, M. Venhart, and A. V. Yeremin, Eur. Phys. J. A **32**, 251 (2007).

- [13] V. K. Utyonkov, N. T. Brewer, Y. T. Oganessian, K. P. Rykaczewski, F. S. Abdullin, S. N. Dmitriev, R. K. Grzywacz, M. G. Itkis, K. Miernik, A. N. Polyakov, J. B. Roberto, R. N. Sagaidak, I. V. Shirokovsky, M. V. Shumeiko, Y. S. Tsyganov, A. A. Voinov, V. G. Subbotin, A. M. Sukhov, A. V. Sabel'nikov, G. K. Vostokin, J. H. Hamilton, M. A. Stoyer, and S. Y. Strauss, Phys. Rev. C **92**, 034609 (2015).
- [14] S. Hofmann, S. Heinz, R. Mann, J. Maurer, J. Khuyagbaatar, D. Ackermann, S. Antalic, W. Barth, M. Block, H. G. Burkhard, V. F. Comas, L. Dahl, K. Eberhardt, J. Gostic, R. A. Henderson, J. A. Heredia, F. P. Heßberger, J. M. Kenneally, B. Kindler, I. Kojouharov, J. V. Kratz, R. Lang, M. Leino, B. Lommel, K. J. Moody, G. Münzenberg, S. L. Nelson, K. Nishio, A. G. Popeko, J. Runke, S. Saro, D. A. Shaughnessy, M. A. Stoyer, P. Thörle-Pospiech, K. Tinschert, N. Trautmann, J. Uusitalo, P. A. Wilk, and A. V. Yeremin, Eur. Phys. J. A 48, 62 (2012).
- [15] Y. T. Oganessian, V. K. Utyonkov, Y. V. Lobanov, F. S. Abdullin, A. N. Polyakov, I. V. Shirokovsky, Y. S. Tsyganov, G. G. Gulbekian, S. L. Bogomolov, B. N. Gikal, A. N. Mezentsev, S. Iliev, V. G. Subbotin, A. M. Sukhov, A. A. Voinov, G. V. Buklanov, K. Subotic, V. I. Zagrebaev, M. G. Itkis, J. B. Patin, K. J. Moody, J. F. Wild, M. A. Stoyer, N. J. Stoyer, D. A. Shaughnessy, J. M. Kenneally, P. A. Wilk, R. W. Lougheed, R. I. Ilkaev, and S. P. Vesnovskii, Phys. Rev. C 69, 054607 (2004).
- [16] Y. T. Oganessian, F. S. Abdullin, S. N. Dmitriev, J. M. Gostic, J. H. Hamilton, R. A. Henderson, M. G. Itkis, K. J. Moody, A. N. Polyakov, A. V. Ramayya, J. B. Roberto, K. P. Rykaczewski, R. N. Sagaidak, D. A. Shaughnessy, I. V. Shirokovsky, M. A. Stoyer, V. G. Subbotin, A. M. Sukhov, Y. S. Tsyganov, V. K. Utyonkov, A. A. Voinov, and G. K. Vostokin, Phys. Rev. Lett. **108**, 022502 (2012).
- [17] G. G. Adamian, N. V. Antonenko, W. Scheid, and V. V. Volkov, Nucl. Phys. A 633, 409 (1998).
- [18] G. G. Adamian, N. V. Antonenko, and W. Scheid, Nucl. Phys. A 678, 24 (2000).
- [19] G. G. Adamian, N. V. Antonenko, and W. Scheid, Eur. Phys. J. A 41, 235 (2009).
- [20] Z.-Q. Feng, G.-M. Jin, F. Fu, and J.-Q. Li, Nucl. Phys. A 771, 50 (2006).
- [21] R. K. Gupta, N. Singh, and M. Manhas, Phys. Rev. C 70, 034608 (2004).
- [22] W. Loveland, Phys. Rev. C 76, 014612 (2007).
- [23] Z. H. Liu and J.-D. Bao, Phys. Rev. C 80, 034601 (2009).
- [24] I. Dutt and R. K. Puri, Phys. Rev. C 81, 064609 (2010).
- [25] O. N. Ghodsi and F. Lari, Phys. Rev. C 89, 054607 (2014).
- [26] X. J. Bao, Y. Gao, J. Q. Li, and H. F. Zhang, Phys. Rev. C 91, 011603(R) (2015).
- [27] N. Ghahramany and A. Ansari, Eur. Phys. J. A 52, 287 (2016).
- [28] K. Hagino, Phys. Rev. C 98, 014607 (2018).
- [29] L. Liu, C. Shen, Q. Li, Y. Tu, X. Wang, and Y. Wang, Eur. Phys. J. A 52, 35 (2016).
- [30] Z.-H. Liu and J.-D. Bao, Phys. Rev. C 87, 034616 (2013).
- [31] A. K. Nasirov, G. Mandaglio, G. Giardina, A. Sobiczewski, and A. I. Muminov, Phys. Rev. C 84, 044612 (2011).
- [32] K. P. Santhosh and V. Safoora, Phys. Rev. C 94, 024623 (2016).
- [33] K. P. Santhosh and V. Safoora, Phys. Rev. C 95, 064611 (2017).
- [34] K. P. Santhosh and V. Safoora, Phys. Rev. C 96, 034610 (2017).
- [35] K. Santhosh and V. Safoora, Eur. Phys. J. A 54, 80 (2018).

- [36] K. Santhosh and V. Safoora, Braz. J. Phys. 51, 90 (2021).
- [37] V. I. Zagrebaev, Phys. Rev. C 64, 034606 (2001).
- [38] X.-J. Lv, Z.-Y. Yue, W.-J. Zhao, and B. Wang, Phys. Rev. C 103, 064616 (2021).
- [39] K. Siwek-Wilczyńska, T. Cap, M. Kowal, A. Sobiczewski, and J. Wilczyński, Phys. Rev. C 86, 014611 (2012).
- [40] V. Zagrebaev and W. Greiner, Phys. Rev. C 78, 034610 (2008).
- [41] L. Zhu, W.-J. Xie, and F.-S. Zhang, Phys. Rev. C 89, 024615 (2014).
- [42] A. S. Zubov, G. G. Adamian, N. V. Antonenko, S. P. Ivanova, and W. Scheid, Phys. Rev. C 65, 024308 (2002).
- [43] S. K. P. Amritanshu Shukla, *Nuclear Structure Physics* (CRC Press, Boca Raton, FL, 2021).
- [44] C. Shen, Y. Abe, D. Boilley, G. Kosenko, and E. Zhao, Int. J. Mod. Phys. E 17, 66 (2008).
- [45] Z. H. Liu and J.-D. Bao, Phys. Rev. C 80, 054608 (2009).
- [46] K. Siwek-Wilczyńska, T. Cap, and J. Wilczyński, Int. J. Mod. Phys. E 19, 500 (2010).
- [47] J. Blocki and W. J. Świątecki, Ann. Phys. 132, 53 (1981).
- [48] W. D. Myers and W. J. Swiatecki, Nucl. Phys. 81, 1 (1966).
- [49] W. D. Myers and W. J. Swiatecki, LBNL Report No. UCRL-17070, Lawrence Berkeley National Laboratory, 1966, https:// escholarship.org/uc/item/3g69c9s2.
- [50] P. Möller and J. R. Nix, Nucl. Phys. A 272, 502 (1976).
- [51] H. J. Krappe, J. R. Nix, and A. J. Sierk, Phys. Rev. C 20, 992 (1979).
- [52] P. Möller and J. R. Nix, Nucl. Phys. A 361, 117 (1981).
- [53] G. Royer and B. Remaud, J. Phys. G 10, 1057 (1984).
- [54] P. Möller and J. Nix, At. Data Nucl. Data Tables 59, 185 (1995).
- [55] K. Pomorski and J. Dudek, Phys. Rev. C 67, 044316 (2003).
- [56] A. Sobiczewski, F. A. Gareev, and B. N. Kalinkin, Phys. Lett. 22, 500 (1966).
- [57] H. Meldner, LBNL Report No. UCRL-16843, Lawrence Berkeley National Laboratory, 1966, https://escholarship.org/uc/item/ 28q2j00w.
- [58] M. A. Stoyer, Nature (London) 442, 876 (2006).
- [59] J. B. Roberto, C. W. Alexander, R. A. Boll, J. D. Burns, J. G. Ezold, L. K. Felker, S. L. Hogle, and K. P. Rykaczewski, Nucl. Phys. A 944, 99 (2015).
- [60] V. I. Zagrebaev and W. Greiner, Nucl. Phys. A 944, 257 (2015).
- [61] Y. T. Oganessian, V. K. Utyonkov, Y. V. Lobanov, F. S. Abdullin, A. N. Polyakov, R. N. Sagaidak, I. V. Shirokovsky, Y. S. Tsyganov, A. A. Voinov, A. N. Mezentsev, V. G. Subbotin, A. M. Sukhov, K. Subotic, V. I. Zagrebaev, S. N. Dmitriev, R. A. Henderson, K. J. Moody, J. M. Kenneally, J. H. Landrum,

D. A. Shaughnessy, M. A. Stoyer, N. J. Stoyer, and P. A. Wilk, Phys. Rev. C **79**, 024603 (2009).

- [62] S. Hofmann, S. Heinz, R. Mann, J. Maurer, G. Münzenberg, S. Antalic, W. Barth, H. G. Burkhard, L. Dahl, K. Eberhardt, R. Grzywacz, J. H. Hamilton, R. A. Henderson, J. M. Kenneally, B. Kindler, I. Kojouharov, R. Lang, B. Lommel, K. Miernik, D. Miller, K. J. Moody, K. Morita, K. Nishio, A. G. Popeko, J. B. Roberto, J. Runke, K. P. Rykaczewski, S. Saro, C. Scheidenberger, H. J. Schött, D. A. Shaughnessy, M. A. Stoyer, P. Thörle-Pospiech, K. Tinschert, N. Trautmann, J. Uusitalo, and A. V. Yeremin, Eur. Phys. J. A 52, 180 (2016).
- [63] J. Khuyagbaatar, A. Yakushev, C. E. Düllmann, D. Ackermann, L. L. Andersson, M. Asai, M. Block, R. A. Boll, H. Brand, D. M. Cox, M. Dasgupta, X. Derkx, A. Di Nitto, K. Eberhardt, J. Even, M. Evers, C. Fahlander, U. Forsberg, J. M. Gates, N. Gharibyan, P. Golubev, K. E. Gregorich, J. H. Hamilton, W. Hartmann, R. D. Herzberg, F. P. Heßberger, D. J. Hinde, J. Hoffmann, R. Hollinger, A. Hübner, E. Jäger, B. Kindler, J. V. Kratz, J. Krier, N. Kurz, M. Laatiaoui, S. Lahiri, R. Lang, B. Lommel, M. Maiti, K. Miernik, S. Minami, A. K. Mistry, C. Mokry, H. Nitsche, J. P. Omtvedt, G. K. Pang, P. Papadakis, D. Renisch, J. B. Roberto, D. Rudolph, J. Runke, K. P. Rykaczewski, L. G. Sarmiento, M. Schädel, B. Schausten, A. Semchenkov, D. A. Shaughnessy, P. Steinegger, J. Steiner, E. E. Tereshatov, P. Thörle-Pospiech, K. Tinschert, T. Torres De Heidenreich, N. Trautmann, A. Türler, J. Uusitalo, M. Wegrzecki, N. Wiehl, S. M. Van Cleve, and V. Yakusheva, Phys. Rev. C 102, 064602 (2020).
- [64] G. Giardina, G. Fazio, G. Mandaglio, M. Manganaro, A. K. Nasirov, M. V. Romaniuk, and C. Saccá, Int. J. Mod. Phys. E 19, 882 (2010).
- [65] N. Wang, J. Tian, and W. Scheid, Phys. Rev. C 84, 061601(R) (2011).
- [66] N. Wang, E.-G. Zhao, W. Scheid, and S.-G. Zhou, Phys. Rev. C 85, 041601(R) (2012).
- [67] J. Zhang, C. Wang, and Z. Ren, Nucl. Phys. A 909, 36 (2013).
- [68] Z. Gan, X. Zhou, M. Huang, Z. Feng, and J. Li, Sci. China Phys., Mech. Astro. 54, 61 (2011).
- [69] C. Y. Wong, Phys. Rev. Lett. 31, 766 (1973).
- [70] I. Dutt, Pramana 76, 921 (2011).
- [71] D. L. Hill and J. A. Wheeler, Phys. Rev. 89, 1102 (1953).
- [72] P. Möller, A. J. Sierk, T. Ichikawa, and H. Sagawa, At. Data Nucl. Data Tables 109–110, 1 (2016).
- [73] R. Vandenbosch and J. R. Huizenga, *Nuclear Fission*, Vol. 29 (Academic Press, New York, 1974), pp. 864–865.
- [74] J. D. Jackson, Can. J. Phys. 34, 767 (1956).