Letter

Nuclear rotation at the fission limit in ²⁵⁴Rf

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(Received 19 January 2022; revised 5 April 2023; accepted 9 May 2023; published 9 June 2023)

A ground-state rotational band in the fissile nucleus ²⁵⁴Rf was observed for the first time. Levels up to spin 14 \hbar and excitation energy of 1.56 MeV were observed. The ²⁵⁴Rf nuclei were produced using the ²⁰⁶Pb(⁵⁰Ti, 2*n*) fusion-evaporation reaction. It is the weakest reaction channel ever studied using in-beam γ -ray spectroscopic methods. The reaction products were separated from the beam in the Argonne gas-filled analyzer (AGFA). The ²⁵⁴Rf nuclei were implanted into a double-sided Si strip detector at the AGFA focal plane and tagged with subsequent ground-state spontaneous fission decays using temporal and spatial correlations. Prompt γ rays in coincidence with the ²⁵⁴Rf recoils were detected in the Gammasphere array of Ge detectors. In order to identify the ground-state rotational band in ²⁵⁴Rf, a method for identifying rotational bands in low statistics γ -ray spectra was developed. The deduced ²⁵⁴Rf kinematic moment of inertia is smaller compared to neighboring even-even nuclei. This is most likely associated with a slightly lower quadrupole deformation and stronger pairing correlations in ²⁵⁴Rf. The behavior of the moment of inertia as a function of rotational frequency is similar to that of the lighter N = 150 isotones ²⁵⁰Fm and ²⁵²No.

DOI: 10.1103/PhysRevC.107.L061302

Introduction. The quest for superheavy nuclei (SHN) is one of the frontiers of nuclear physics. In recent years, SHN with atomic numbers 113-118 have been discovered [1-3]. However, because of very small production cross sections, only the basic properties such as the dominant decay modes and the lifetimes are known for SHN. On the other hand, trans-fermium nuclei located near the deformed energy gaps Z = 100 and N = 152, which similarly to the heaviest nuclei owe their existence to shell corrections, can be produced with sufficiently large cross sections to facilitate spectroscopic studies. These nuclei are among the best rotors known and are an excellent testing ground for models which are used to describe the heaviest nuclei. One of the goals of nuclear theory is to predict the location and the magnitude of major spherical shell gaps which can determine the possible occurrence of the "island of stability" of SHN. However, different theoretical approaches do not agree (see Ref. [4] and references therein). For example, the microscopic-macroscopic approach predicts

Nuclei near Z = 100, N = 152 have been studied using in-beam, isomer, and decay spectroscopic methods. See the recent review [5] of the experimental and theoretical developments in this mass region. In particular, in-beam spectroscopy has provided information about moments of inertia as a function of rotational frequency. The nuclear moment of inertia depends on deformation. However, because nucleons are paired, its value is lower than that for the rigid body and it only approaches the rigid body limit as the pairing correlations diminish. Pairing correlations depend on details of single-particle levels near the Fermi surface. In nuclei with closed shells, pairing correlations are weaker due to a lower

2469-9985/2023/107(6)/L061302(6)

Z = 114 and N = 184 whereas various self-consistent calculations give different results. In particular, the Hartree-Fock method with Skyrme forces favors Z = 126, N = 184 while the relativistic mean field theory prefers Z = 120 and N =172. It is worth noting that only the microscopic-macroscopic approach reproduces the Z = 100, N = 152 deformed shell gaps which have already been established experimentally. Interestingly, some of the orbitals which determine the spherical magic numbers for superheavy nuclei are located close to the Fermi surface in the transfermium region.

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density of states and increase as one departs from closed shells. Also, the moment of inertia undergoes a gradual rise at low rotational frequencies followed by a rapid increase when a pair of nucleons occupying an orbital with a high angular momentum breaks up and the unpaired nucleons align their angular momenta with the rotational axis of the core. In the Z = 100, N = 152 region the neutron $\nu j_{15/2}$ and/or the proton $\pi i_{13/2}$ pairs tend to align first [6]. The latter orbital is located near the Fermi surface in the heaviest nuclei.

Rotational properties of transfermium nuclei have been described within various theoretical frameworks. For example, their deformation and moments of inertia as well as the underlying single-particle structure were discussed using the macroscopic-microscopic approach with the Woods-Saxon potential and the universal set of parameters in Ref. [7]. Furthermore, results of self-consistent mean field calculations with the SLy4 interaction and a density-dependent pairing force were presented in Ref. [8], while the covariant density functional theory was used in Ref. [9] to calculate moments of inertia. The impact of higher order deformation on moments of inertia was discussed in Refs. [10] and [11] using the total-Routhian surface with the cranking shell model and the particle-number-conserving cranking shell model, respectively.

Fission plays an important role in the realm of superheavy nuclei. It determines their production cross sections and it ultimately defines the limits of their existence. The ²⁵⁴Rf ground state disintegrates rapidly by spontaneous fission with the shortest measured fission lifetime. The published ²⁵⁴Rf half-life values have shown a wide variation, ranging from 500(200) us in Ref. [12], 23(3) us [13], and 29.6(+0.7-0.6) us [14] until the value of 23.2(1.1) µs was measured recently [15] in agreement with Ref. [13]. In addition, two isomers were observed in this nucleus with half-lives of 4.7(1.1) us and 247(73) µs [15]. They were interpreted as a two-quasiparticle (qp) and a four-qp K isomer, respectively. Surprisingly, the half-life of the two-qp isomer in ²⁵⁴Rf is four orders of magnitude shorter than for the equivalent two-qp isomers in the lighter N = 150 isotones ²⁵⁰Fm and ²⁵²No. Currently, the reason for this abrupt change is not well understood. Despite the rapidly fissioning ground state, no fission events were observed for either of the two isomers. This implies that fission is hindered by a factor of at least 2 for the two-qp isomer and by a factor of at least 25 for the four-qp isomer compared to the ground-state fission. Superheavy nuclei with Z > 118 are expected to decay rapidly by spontaneous fission and fission hindrance in isomeric states could be a possible mechanism for their survival. Another interesting aspect of rotation in the proton-rich transfermium nuclei is the interplay between fission and rotation at high spin. This interplay has been extensively discussed in Ref. [16], and in Ref. [17] for the case of 254 No. In 254 Rf, fission is expected to compete with γ decay at lower excitation energies than in ²⁵⁴No due to a lower fission barrier. At present, the ²⁵⁶Rf nucleus is the heaviest nucleus which has been studied using in-beam spectroscopic methods [18]. The ground-state rotational band in ²⁵⁶Rf has been delineated up to spin 20ħ. Ground-state rotational bands are also known in the lighter N = 150 even-even isotones ²⁵⁰Fm [19] and ²⁵²No [20] as well as in ²⁵⁴No [21].

In order to trace the evolution of rotational properties away from the Z = 100 and N = 152 shells and to elucidate the role of a lower fission barrier, a search for γ -ray transitions in the fissile nucleus ²⁵⁴Rf was carried out. This paper reports first observation of the ground-state rotational band in ²⁵⁴Rf.

Experimental details. ²⁵⁴Rf nuclei were synthesized using the 206 Pb(50 Ti, 2n) reaction. The cross section for this reaction is only 2.4(2) nb [13] which is almost a factor of 7 smaller compared to the production of ²⁵⁶Rf [18]. A ⁵⁰Ti beam with an energy of 244 MeV delivered by the ATLAS linear accelerator at the Argonne National Laboratory impinged on 0.5 mg/cm²-thick ²⁰⁶Pb targets isotopically enriched to 95.9%. A 40 μ g/cm²-thick and 10 μ g/cm²-thick carbon layer was evaporated on the front and on the back of the targets, respectively. The experiment was performed in two parts. The targets were mounted on a wheel which rotated with a frequency of about 1200 rpm. During the first part, four 9 mm-wide target segments formed a circle with 17 mm average radius. Larger 11 mm-wide segments, which formed a 34 mm-radius circle, were used during the second part. The beam was wobbled horizontally with 5 Hz frequency to cover the whole target area. In total, the targets were irradiated for 94(84)h and the average beam intensity was about 20(25)pnA during the first (second) part. The beam was periodically swept away to avoid hitting the target wheel spokes. During the first part of the experiment the targets suffered significant damage and had to be replaced multiple times. This was attributed to bursts in beam intensity associated with the volatile nature of the sputtering process used to introduce the beam material into the plasma of the ion source. This problem was remedied by using larger targets during the second experiment.

Prompt γ rays were detected using Gammasphere, an array of Ge detectors arranged symmetrically around the target wheel. Gammasphere consisted of 65 and 70 detectors at the time of the first and the second part of the experiment, respectively. The average γ -ray energy resolution at energies between 100 keV and 500 keV was FWHM ≈ 3 keV. The reaction products recoiling from the target were separated from the beam in the Argonne gas-filled analyzer (AGFA). AGFA consists of a large bore vertically focusing quadrupole magnet and a combined-function horizontally focusing magnetic dipole. AGFA was filled with He gas at 0.5 Torr and was set to transport reaction products centered around $B\rho =$ 2.08 Tm. After passing through the parallel-grid avalanche counter (PGAC) located at the exit from AGFA the reaction products were implanted into a 300-µm-thick $64 \times 64 \text{ mm}^2$ double-sided Si strip detector (DSSD). The front and back side of the DSSD were divided into 160 strips each which were mutually orthogonal resulting in 25600 pixels. The DSSD energy resolution for α particles with energies around 5 MeV was FWHM ≈ 25 keV. An array of eight $4 \times 7 \text{ cm}^2$ 300 µm-thick single-sided Si strip detectors in a tunnel geometry was installed in front of the DSSD to catch escaping decay particles. Each detector was divided into seven strips perpendicular to the DSSD. A large area 5×5 cm² 300 µm-thick Si detector was placed behind the DSSD to veto energetic light particles like protons and He atoms which punched through the DSSD. All events registered in the DSSD



FIG. 1. Energies of decay events with energies larger than 50 MeV as a function of time with respect to preceding implants. The log base 10 of the decay time expressed in nanoseconds is shown on the ordinate axis.

in coincidence with the PGAC were flagged as potential recoils and the remaining ones as potential decays. The energy deposited in the DSSD and the time of flight between the PGAC and the DSSD were used to suppress scattered beam events.

Results. Figure 1 shows the energy spectrum of decay events as a function of the time with respect to preceding recoil events. A group of decay events with energies between 100 and 250 MeV corresponding to decay times around 20 µs is clearly separated from randomly correlated background decay events in Fig. 1. This group was interpreted as ²⁵⁴Rf ground-state fission events. Four high-energy events with life-times around 10 ms correspond most likely to spontaneous fission of ²⁵⁶Rf which was produced on ²⁰⁸Pb contaminant present in the targets. In Fig. 2, energies of prompt γ rays detected in coincidence with ²⁵⁴Rf recoils as a function of time between γ rays and implants are shown. A band of events corresponding to prompt coincidences is clearly visible in Fig. 2. Figure 3 contains the spectrum of γ rays in prompt coincidence with the ²⁵⁴Rf implants. Excess of counts at







FIG. 3. Energy spectrum of γ rays in prompt coincidence with ²⁵⁴Rf implants.

the energies corresponding to Rf K_{α} and K_{β} x-ray lines confirms assignment of the spectrum to ²⁵⁴Rf. Clusters of counts at higher energies correspond to discrete transitions of the ground-state rotational band in ²⁵⁴Rf. Guided by the known rotational bands in neighboring even-even nuclei, the group of five counts at 172 keV was interpreted as the 6⁺ \rightarrow 4⁺ transition in ²⁵⁴Rf. The 2⁺ \rightarrow 0⁺ and 4⁺ \rightarrow 2⁺ transitions have lower energies and undergo predominantly electron conversion. There are other clusters of counts in the spectrum but their firm assignment is difficult. In order to search for a rotational band in this low statistics spectrum a novel method was developed which takes advantage of the regular nature of rotational bands. Ground-state rotational bands in deformed even-even nuclei form a sequence of levels with even spins I = 0,2,4,... and energies given by the formula

$$E(I) = \frac{\hbar^2}{2\mathcal{J}}I(I+1), \qquad (1)$$

where \mathcal{J} is the moment of inertia. The kinematic moment of inertia $\mathcal{J}^{(1)}$ can be expressed as a function of rotational frequency ω using the Harris formula [22]

$$\mathcal{J}^{(1)} = \mathcal{J}_0 + \mathcal{J}_1 \omega^2, \qquad (2)$$

where \mathcal{J}_0 and \mathcal{J}_1 are the Harris parameters. The kinematic moment of inertia and the rotational frequency can be derived from the energies of γ -ray transitions $E_{\gamma}(I)$ in a rotational band using the expressions

$$\mathcal{J}^{(1)} = (2I - 1)/E_{\gamma}(I)$$
(3)

and

$$\omega = E_{\gamma}(I)/2, \tag{4}$$

respectively. In order to search for a rotational band in the ²⁵⁴Rf γ -ray spectrum, first, the spectrum was divided into five wide energy regions where individual ground-state band transitions starting with the 6⁺ \rightarrow 4⁺ transition are expected to be located. For each γ -ray event (*i*) in region (*j*) the quantity

$$\mathcal{J}_0^{\exp}(i) = (2I(j) + 1)/E_{\gamma}(i) - \mathcal{J}_1(E_{\gamma}(i)/2)^2, \qquad (5)$$



FIG. 4. Experimental moment of inertia (see text) calculated for the ground-state band in (a) 254 Rf and (b) 254 No.

corresponding to the \mathcal{J}_0 parameter was derived from Eq. (2), where I(j) is the spin assigned to region (*j*). The ⁴Rf moment of inertia histogram calculated using the value of $\mathcal{J}_1 = 211.2\hbar^2 \text{MeV}^{-3}$ derived from the fit to the ²⁵²No ground-state rotational band is shown in Fig. 4(a). A peak in such a spectrum indicates presence of a rotational band and its position represents the value of \mathcal{J}_0 for this band as illustrated in Fig. 4(b) where results of the same procedure applied to the ground-state rotational band in ²⁵⁴No are shown. On further inspection, the peak in Fig. 4(a), which contains 15 counts including about four background counts, was formed by the clusters of counts at 384, 339, 285, 232, 172 keV, which are marked in Fig. 3. These clusters were interpreted as the $14^+ \rightarrow 12^+ \rightarrow 10^+ \rightarrow 8^+ \rightarrow 6^+ \rightarrow 4^+ \gamma$ -ray sequence forming the 254 Rf ground-state rotational band. The 172, 232, 285, 339, 384 keV γ -ray energies were used to fit the Harris formula [Eq. (2)] and $\mathcal{J}_0 = 62.1 \,\hbar^2 \text{MeV}^{-1}$ and $\mathcal{J}_1 =$ 215.1 \hbar^2 MeV⁻³ parameters were found to fit the data best. The Harris formula results in the values of $E(2^+) = 48 \text{ keV}$ and $E(4^+) = 158$ keV for the lowest levels in the ²⁵⁴Rf band. Similarly, the energies of 420 keV and 470 keV are predicted for the $16^+ \rightarrow 14^+$ and $18^+ \rightarrow 16^+$ transitions, respectively. Interestingly, two counts were registered at 470 keV in the ²⁵⁴Rf spectrum.

Discussion. The properties of the ground-state rotational bands in the even-even nuclei near Z = 100 and N = 152 including the new data for ²⁵⁴Rf are shown in Table I.The

TABLE I. Properties of ground-state rotational bands in eveneven nuclei near Z = 100, N = 152 compared with results of microscopic-macroscopic calculations [7].

Isotope	E(2 ⁺) keV	$\mathcal{J}_0 \ \hbar^2 \mathrm{MeV}^{-1}$	\mathcal{J}_1 $\hbar^2 \mathrm{MeV}^{-3}$	$\beta_2^{\text{calc}}[7]$	$\frac{E_{\text{calc}}(2^+) [7]}{\text{keV}}$
²⁵⁴ Rf	48	62.1	215.1	0.247	46.9
²⁵² No	46.4	64.4	211.2	0.249	44.5
²⁵⁰ Fm	44.0	68.1	233.0	0.248	43.9
²⁵⁶ Rf	44.8	66.7	179.7	0.249	43.4
²⁵⁴ No	44.1	68.2	160.0	0.252	41.6



FIG. 5. Kinematic moments of inertia for N=150 and N=152 even-even isotones. The lines represent the Harris formula fitted to first 5 experimental points.

experimental ground-state band kinematic moments of inertia as a function of rotational frequency for ²⁵⁴Rf and neighboring nuclei is shown in Fig. 5. The curves in Fig. 5 represent fits to the Harris formula using the $6^+,...,14^+$ levels. The value of \mathcal{J}_0 in ²⁵⁴Rf, which corresponds to the moment of inertia at $\omega = 0$, is smaller than any of the \mathcal{J}_0 values deduced for the nuclei included in Fig. 5. A quadrupole deformation of $\beta_2 = 0.247$ was calculated for ²⁵⁴Rf using the microscopic-macroscopic model in Ref. [7]. This value is smaller than those calculated for all neighboring nuclei shown in Table I but this alone cannot explain the change in the moment of inertia. Compared to other nuclei, ²⁵⁴Rf has more nucleons outside of the Z = 100and N = 152 closed shells. This should result in stronger pairing correlations compared to other nuclei in Table I and lead to significant decrease of the moment of inertia. According to Ref. [7], the calculated energy of the 2^+ state derived from the rotational formula is 46.9 keV which agrees well with the value of 48 keV obtained from the \mathcal{J}_0 value deduced here. It is interesting to compare the microscopic-macroscopic approach with self-consistent calculations. In Ref. [23], three different energy-density-functional (EDF) models, based on covariant, Skyrme, and Gogny functionals, each with two different parameter sets were used to calculate properties of trans-fermium nuclei including the kinematic moments of inertia (see Figs. 13 and 14 in Ref. [23]). Among the functionals used, only the D1M Gogny functional predicts the decrease of the moment of inertia in ²⁵⁴Rf compared to ²⁵⁶Rf and reproduces the experimental values well.

It is also interesting to inspect evolution of the moment of inertia with rotational frequency. The moment of inertia in 254 Rf increases with rotational frequency in a similar fashion as the lighter N = 150 isotones and faster than in the N = 152isotones (see Fig. 5). The onset of alignment in the N = 150isotones was observed at around $\omega = 0.2$ MeV [18]. There is no evidence for alignment up to the maximum rotational frequency of 0.19 MeV observed in the present work.

Kinematic moments of inertia for N = 150 and N = 152 isotones of Fm, No, and Rf calculated using the cranked shell



FIG. 6. The comparison between experimental kinematic moments of inertia and values calculated using the particle-numberconserving cranked shell model (PNC-CSM) for N = 150 and N =152 isotones of Fm, No, and Rf. The experimental data are denoted by solid circles. The PNC-CSM calculations with pairing strengths $G_0 = 0.30$ and $G_0 = 0.35$ (only for ²⁵⁴Rf) are shown as solid lines and a dashed line, respectively.

model with pairing treated by the particle-number-conserving method are compared with the experimental values in Fig. 6. The calculated moments of inertia using a pairing strength of $G_0 = 0.30$, which are shown as solid lines, reproduce the experimental results very well for all nuclei except ²⁵⁴Rf. The moment of inertia for ²⁵⁴Rf is smaller than for the other nuclei which indicates stronger pairing correlations. The PNC-CSM calculations with a stronger pairing strength of $G_0 = 0.35$,

which is denoted by the dashed line in Fig. 6, agrees very well with the experimental data in 254 Rf.

Summary. The ground-state rotational band was observed for the first time in the fissile nucleus ²⁵⁴Rf. Levels up to spin 14 \hbar and excitation energy of 1.56 MeV were identified. The moment of inertia deduced for ²⁵⁴Rf is smaller than in neighboring nuclei. The moment of inertia is well reproduced by microscopic-macroscopic calculations whereas only the self-consistent approach with the D1M Gogny functional agrees with the data in ²⁵⁴Rf and ²⁵⁶Rf. The particle-numberconserving cranked shell model reproduces the kinematic moments of inertia as a function of rotational frequency for the N = 150 and N = 152 isotones of Fm, No, and Rf except for ²⁵⁴Rf where a larger moment of inertia is predicted. Good agreement between the calculations and the data was achieved for ²⁵⁴Rf when the pairing strength was increased from $G_0 = 0.30$ to $G_0 = 0.35$.

In order to shed light on the rotational alignment in 254 Rf and competition between γ -ray emission and fission, more statistics are required to observe levels with higher spins at higher excitation energy. A search for a rotational band in the neighboring nucleus 250 No, which also decays rapidly by spontaneous fission, but can be produced with a larger cross section, should be also feasible.

Acknowledgement. This material is based upon work supported by the US Department of Energy, Office of Science, Office of Nuclear Physics, under Contracts No. DE-AC02-06CH11357 (ANL) and DE-AC02-05CH11231 (LBNL) and by the National Natural Science Foundation of China under Grant No. U2032138. This research used resources of ANL's ATLAS facility, which is a DOE Office of Science User Facility.

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