








***Ab initio* calculation of the β -decay spectrum of ${}^6\text{He}$** G. B. King ¹, A. Baroni,² V. Cirigliano ³, S. Gandolfi ², L. Hayen ⁴, E. Mereghetti ², S. Pastore ^{1,5} and M. Piarulli ^{1,5}¹*Department of Physics, Washington University in Saint Louis, Saint Louis, Missouri 63130, USA*²*Theoretical Division, Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA*³*Institute for Nuclear Theory, University of Washington, Seattle Washington 91195-1550, USA*⁴*Department of Physics, North Carolina State University, Raleigh, North Carolina 27695, USA*⁵*McDonnell Center for the Space Sciences at Washington University in St. Louis, Missouri 63130, USA*

(Received 9 August 2022; accepted 1 December 2022; published 27 January 2023)

We calculate the β spectrum in the decay of ${}^6\text{He}$ using quantum Monte Carlo methods with nuclear interactions derived from chiral effective field theory and consistent weak vector and axial currents. We work at second order in the multipole expansion, retaining terms suppressed by $\mathcal{O}(q^2/m_\pi^2)$, where q denotes low-energy scales such as the reaction's Q value or the electron energy, and m_π is the pion mass. We go beyond the impulse approximation by including the effects of two-body vector and axial currents. We estimate the theoretical error on the spectrum by using four potential models in the Norfolk family of local two- and three-nucleon interactions, which have different cutoffs, fit two-nucleon data up to different energies, and use different observables to determine the couplings in the three-body force. We find the theoretical uncertainty on the β spectrum, normalized by the total rate, to be well below the permille level, and to receive contributions of comparable size from first- and second-order corrections in the multipole expansion. We consider corrections to the β decay spectrum induced by beyond-standard-model charged-current interactions in the standard model effective field theory, with and without sterile neutrinos, and discuss the sensitivity of the next generation of experiments to these interactions.

DOI: [10.1103/PhysRevC.107.015503](https://doi.org/10.1103/PhysRevC.107.015503)**I. INTRODUCTION**

Nuclear β decays have been instrumental in establishing the standard model (SM) as the theory of the electroweak interactions [1–3]. In the era of the Large Hadron Collider, β decays still provide very sensitive probes of physics beyond the standard model (BSM), and are highly competitive and complementary to searches at the energy frontier [4–15]. Superallowed $0^+ \rightarrow 0^+$ transitions, combined with theoretical progress in the evaluation of radiative corrections [16–21], allow for the extraction of the V_{ud} element of the Cabibbo-Kobayashi-Maskawa (CKM) quark mixing matrix with uncertainty at the level $\delta V_{ud} \approx 3 \times 10^{-4}$ [16–18,22], probing BSM scales up to 10 TeV. Improved measurements of the neutron lifetime and β asymmetry [23–26] and the percent determination of the nucleon axial charge g_A from lattice QCD [10,27–29] test right-handed charged currents at subpercent level. Global analyses of superallowed transitions, neutron decay, and mirror β decays limit the strength of nonstandard vector, axial, scalar and tensor charged-current interactions to be less than a thousandth of the weak interactions [12,15]. At the same time charged- and neutral-current Drell-Yan production at the Large Hadron Collider is starting to directly access scales of a few TeV [30–33], and the good agreement between precise theoretical predictions and Drell-Yan data allow for the exclusion of interactions at the effective scale $\Lambda = 4\text{--}5$ TeV [5,10,11,13,14].

β decay spectra provide sensitive probes of charged currents with different chiral structure from the SM. The interference of BSM currents with the $V\text{-}A$ SM interactions induces a distinctive m_e/E_e —where m_e and E_e are the electron's mass and energy, respectively—dependence in the β spectrum, the so called “Fierz interference term” [34], usually denoted by b . The first direct neutron measurements constrain the Fierz interference term to be $-0.018 < b < 0.052$ at the 90% confidence level [35,36]. The Fierz interference term induced by scalar currents is tested in Fermi transitions, with the most recent analysis of superallowed β decays yielding $b < 3.3 \times 10^{-3}$ (90% confidence level) [22]. Measurements of spectra of purely Gamow-Teller transitions, such as the decay of ${}^6\text{He}$ to ${}^6\text{Li}$, aim to push the constraint on the Fierz interference term induced by tensor and pseudoscalar currents to the level $b < 10^{-3}$ [37,38], probing tensor currents at the 10 TeV mass scale. In addition, modifications to the shape of the β spectrum can reveal the existence of sterile neutrinos, with minimal or nonminimal interactions [39–44].

With experimental sensitivity approaching the permille level, it is crucial to provide comparably accurate theoretical predictions of the β spectrum in the SM, including small corrections from the momentum dependence of nuclear matrix elements, electromagnetic corrections, and isospin breaking effects. For light to medium mass nuclei, accurate calculations of low-energy nuclear observables are currently feasible using the microscopic or *ab initio* description of nuclei. Within this approach nuclei and their properties emerge from the

underlying nucleonic dynamics and ensuing many-nucleon correlations and electroweak currents. The first *ab initio* calculation of the ${}^6\text{He} \rightarrow {}^6\text{Li}$ β decay spectrum was performed in Ref. [45], using the variational Monte Carlo method with the Argonne v_{18} two-nucleon potential, supplemented by the Urbana-IX three-nucleon interaction. References [46,47] repeated the calculation in the no-core shell model. In this paper, we compute both SM and BSM nuclear matrix elements for the ${}^6\text{He} \rightarrow {}^6\text{Li}$ decay using quantum Monte Carlo (QMC) methods [48] to solve for the structure and dynamics of the strongly correlated many-body problem for nuclei. QMC methods allow one to retain the complexity of many-nucleon dynamics whose effects are essential to explain electroweak data in a wide range of energy and momentum transfer [48–68]. Here, we base our calculations on the Norfolk two- and three-nucleon chiral effective field theory potentials and consistent electroweak currents [50–52,56,59,60,62,63,69].

This paper is structured as follows: In Sec. II, we introduce the multipole expansion of the SM weak vector and axial currents and express the differential decay rate with respect to the electron energy including terms up to second order in the multipole expansion. Section III A reports on the QMC calculations of the multipoles entering the decay rate, and in Sec. IV we discuss the uncertainties of the leading and subleading multipoles and the ensuing theoretical uncertainty in the SM decay rate, which limits the sensitivity to beyond the SM physics. We then discuss BSM signatures. In Sec. V we introduce the effective Lagrangians that mediate β decays in the presence of BSM interactions and discuss their corrections to the β spectrum. In Sec. VI we examine the implications of controlling the uncertainty on the spectrum at better than the permille level on nonstandard charged-current interactions. We conclude in Sec. VII. The appendices contain some technical details. In Appendix A we provide a list of standard and nonstandard Lagrangians which mediate β decays in the standard model EFT (SMEFT), while a sketch of the derivation of the multipole expansion for these currents is carried out in Appendix B. Formal expressions of the many-body chiral EFT current operators are given in Appendix C. The expression of the fully differential unpolarized decay rate is given in Appendix D. Appendix E is devoted to higher-order electroweak and recoil corrections.

II. DIFFERENTIAL DECAY RATE IN THE STANDARD MODEL

β decays are sensitive to a variety of physics scales, namely, the Q value of the reaction—typically a few MeV—which determines the momentum of the outgoing electron and neutrino; the nuclear binding momentum $\gamma = \sqrt{m_N B} \approx m_\pi$, where B is the binding energy, which is the relevant scale in the nuclear matrix elements; and Λ_χ , the scale at which chiral EFT breaks down. We can take advantage of the scale separation $Q \ll \gamma \ll \Lambda_\chi$ by organizing the nuclear matrix elements in a double expansion in Q/γ and γ/Λ_χ , combining chiral EFT with a multipole expansion of the weak matrix elements [70,71]. After performing the multipole expansion, the differential cross section in the SM can be expressed in

terms of few matrix elements of the axial and vector charge and current densities, which are generalized to include scalar, pseudoscalar, and tensor currents in the SMEFT.

In the SM, β decays are mediated by the exchange of a W boson between purely left-handed quarks, electrons and neutrinos. At scales much smaller than the electroweak, and focusing on the first generation of quarks, the effective Lagrangian can be expressed in terms of the local four-fermion interaction:

$$\mathcal{L}_{\text{SM}} = -\frac{4G_F}{\sqrt{2}} V_{ud} \bar{e}_L \gamma^\mu \nu_L \bar{u}_L \gamma_\mu d_L + \text{H.c.}, \quad (1)$$

where $G_F = (\sqrt{2}v^2)^{-1} = 1.166 \times 10^{-5} \text{ GeV}^{-2}$ is the Fermi constant extracted from muon decay, $v = 246 \text{ GeV}$ is the Higgs vacuum expectation value, and V_{ud} is the u - d element of the Cabibbo-Kobayashi-Maskawa mixing matrix, $V_{ud} = 0.97373(31)$ [22,72]. The Lagrangian (1) receives weak and electromagnetic corrections, which we will discuss in the following sections.

At the nuclear level, Eq. (1) leads to the weak Hamiltonian

$$H_w = \frac{G_F}{\sqrt{2}} V_{ud} \int d^3\mathbf{x} j_\mu^{\text{lept}}(\mathbf{x}) \mathcal{J}_{V-A}^\mu(\mathbf{x}), \quad (2)$$

where

$$j_\mu^{\text{lept}} = 2\bar{e}_L \gamma_\mu \nu_L. \quad (3)$$

and \mathcal{J}_{V-A}^μ denotes the hadronic realization of the quark current $\bar{u}\gamma^\mu(1-\gamma_5)d$. The derivation only assumes that such a realization exists, and we give its explicit representation in chiral EFT in Appendix C. The weak Hamiltonian can be expanded in infinite sum of multipole operators with definite total angular momentum J and parity π . The transition ${}^6\text{He}(0^+) \rightarrow {}^6\text{Li}(1^+)$ only receives contributions from operators with $J^\pi = 1^+$. The general expression for the differential decay rate, derived, for example, in Ref. [71] and reported in Eq. (D2), then contains the multipoles C_1 , L_1 , E_1 , and M_1 . These are defined in terms of the coordinate space charge and current densities in Eqs. (B15)–(B18). In chiral EFT, a momentum space representation is more convenient, and C_1 , L_1 , E_1 , and M_1 can be expressed in terms of the axial charge, $\rho(\mathbf{q}; A)$, and the vector and axial currents, $\mathbf{j}(\mathbf{q}; V)$ and $\mathbf{j}(\mathbf{q}; A)$ [45]:

$$C_1(q; A) = \frac{i}{\sqrt{4\pi}} \langle {}^6\text{Li}, 10 | \rho_+^\dagger(q\hat{\mathbf{z}}; A) | {}^6\text{He}, 00 \rangle, \quad (4)$$

$$L_1(q; A) = \frac{i}{\sqrt{4\pi}} \langle {}^6\text{Li}, 10 | \hat{\mathbf{z}} \cdot \mathbf{j}_+^\dagger(q\hat{\mathbf{z}}; A) | {}^6\text{He}, 00 \rangle, \quad (5)$$

$$E_1(q; A) = \frac{i}{\sqrt{2\pi}} \langle {}^6\text{Li}, 10 | \hat{\mathbf{z}} \cdot \mathbf{j}_+^\dagger(q\hat{\mathbf{x}}; A) | {}^6\text{He}, 00 \rangle, \quad (6)$$

$$M_1(q; V) = -\frac{1}{\sqrt{2\pi}} \langle {}^6\text{Li}, 10 | \hat{\mathbf{y}} \cdot \mathbf{j}_+^\dagger(q\hat{\mathbf{x}}; V) | {}^6\text{He}, 00 \rangle, \quad (7)$$

where the momentum carried by the current is $\mathbf{q} = \mathbf{p}_e + \mathbf{p}_\nu$, with \mathbf{p}_e and \mathbf{p}_ν the electron and antineutrino momenta, and $q = |\mathbf{q}|$. The subscript $+$ denotes the charge-changing operators $\rho_+^\dagger = \rho_x^\dagger + i\rho_y^\dagger$ and $\mathbf{j}_+^\dagger = \mathbf{j}_x^\dagger + i\mathbf{j}_y^\dagger$, where the subscripts x and y refer to the isospin space [73]. In Eqs. (4)–(7), the states are characterized by the quantum numbers J and M_J ,

denoting the total angular momentum and the projection along the z axis of the initial and final nuclear states. $\hat{\mathbf{z}}$ and $\hat{\mathbf{x}}$ denote versors in the z and x direction, respectively.

At zero momentum, the electric and longitudinal multipoles are related by

$$L_1^{(0)}(A) = \frac{1}{\sqrt{2}}E_1^{(0)}(A) = \sqrt{\frac{3}{4\pi}}g_A \text{RME}, \quad (8)$$

where the reduced matrix element denotes the standard Gamow-Teller matrix element,

$$\begin{aligned} \text{RME} &= \frac{\sqrt{2J_f+1}}{g_A} \frac{\langle J_f M | \hat{\mathbf{z}} \cdot \mathbf{j}_{\pm}(A) | J_i M \rangle}{\langle J_i M, 10 | J_f M \rangle} \\ &= -\sqrt{3} \langle J_f M | \sigma_z \tau^+ | J_i M \rangle, \end{aligned} \quad (9)$$

where $\langle J_i M, 10 | J_f M \rangle$ is a Clebsch-Gordan coefficient and, in our case, $J_i = 0$ and $J_f = 1$. M_1 encompasses the contribution from weak magnetism, while C_1 receives contributions from the induced tensor and induced pseudoscalar form factors, $d(q^2)$ and $h(q^2)$ in the notation of Ref. [74].

The momentum $q = |\mathbf{q}|$ is limited by the reaction's Q value, and, because $Q \ll m_\pi$, the matrix elements can be expanded in powers of q/m_π . From the definitions in Eqs. (B15)–(B18), it can be proved that L_1 and E_1 only have even powers of q , while C_1 and M_1 only have odd powers [71],

so that Eqs. (4)–(7) can be expanded as

$$C_1(q; A) = -i \frac{qr_\pi}{3} \left(C_1^{(1)}(A) - \frac{(qr_\pi)^2}{10} C_1^{(3)}(A) + \mathcal{O}[(qr_\pi)^4] \right), \quad (10)$$

$$L_1(q; A) = -\frac{i}{3} \left(L_1^{(0)}(A) - \frac{(qr_\pi)^2}{10} L_1^{(2)}(A) + \mathcal{O}[(qr_\pi)^4] \right), \quad (11)$$

$$M_1(q; V) = -i \frac{qr_\pi}{3} \left(M_1^{(1)}(V) - \frac{(qr_\pi)^2}{10} M_1^{(3)}(V) + \mathcal{O}[(qr_\pi)^4] \right), \quad (12)$$

$$E_1(q; A) = -\frac{i}{3} \left(E_1^{(0)}(A) - \frac{(qr_\pi)^2}{10} E_1^{(2)}(A) + \mathcal{O}[(qr_\pi)^4] \right), \quad (13)$$

where, in order to have dimensionless coefficients, we introduced $r_\pi = 1/m_{\pi^+} = 1.41382$ fm. In the equations above, the $C_1^{(i)}$, $L_1^{(i)}$, $M_1^{(i)}$, and $E_1^{(i)}$ are simply the coefficients of the expansion in qr_π , and they will be determined by the interpolation procedure discussed in Sec. IV. Their operator definitions in the impulse approximation are given, for example, in Ref. [74]. Since for the decay under consideration $qr_\pi \lesssim 0.03$, we consider terms up to second order in qr_π to reach the uncertainty goal of 10^{-4} in β decay spectra.

After summing over the lepton and nuclear spins, integrating over the neutrino energy and the angle between the electron and neutrino momentum, the SM decay rate, differential with respect to the electron energy, is given by

$$\begin{aligned} \frac{d\Gamma}{d\varepsilon} &= (1 + \Delta_R^V)(1 + \delta_R(Z, \varepsilon)) \frac{G_F^2 W_0^5 V_{ud}^2}{2\pi^3} \sqrt{1 - \frac{\mu_e^2}{\varepsilon^2}} \varepsilon^2 (1 - \varepsilon)^2 F_0(Z, \varepsilon) L_0(Z, \varepsilon) S(Z, \varepsilon) R_N(\varepsilon) \frac{4\pi}{2J_i + 1} \frac{1}{9} \\ &\times \left\{ 3 |L_1^{(0)}|^2 \left[1 + \alpha Z W_0 R \left(\frac{2}{35} - \frac{233}{630} \frac{\alpha Z}{W_0 R} - \frac{1}{70} \frac{\mu_e^2}{\varepsilon} - \frac{4}{7} \varepsilon \right) \right] + 2 W_0 r_\pi \left[\left(1 - 2\varepsilon + \frac{\mu_e^2}{\varepsilon} \right) \text{Re}(E_1^{(0)} M_1^{(1)*}) \right. \right. \\ &- \left. \left(1 - \frac{\mu_e^2}{\varepsilon} \right) \text{Re}(L_1^{(0)} C_1^{(1)*}) \right] + \frac{(W_0 r_\pi)^2}{3} \left[\left(3 - 4\varepsilon(1 - \varepsilon) - \mu_e^2 \frac{2 + \varepsilon}{\varepsilon} \right) |C_1^{(1)}|^2 \right. \\ &- \frac{3}{5} \left(1 - \frac{\mu_e^2}{\varepsilon} (2 - \varepsilon) \right) \text{Re}(L_1^{(0)} L_1^{(2)*}) + \left. \left(3 - 10\varepsilon(1 - \varepsilon) + \mu_e^2 \frac{4 - 7\varepsilon}{\varepsilon} \right) \left(|M_1^{(1)}|^2 - \frac{1}{5} \text{Re}(E_1^{(0)} E_1^{(2)}) \right) \right] \\ &\left. - \frac{4}{7} \frac{\alpha Z W_0 r_\pi^2}{R} (1 - \varepsilon) \left(\frac{E_1^{(0)} E_1^{(2)}}{2} - L_1^{(0)} L_1^{(2)} \right) \right\}, \end{aligned} \quad (14)$$

where we introduced the scaled variables $E_e = W_0 \varepsilon$ and $m_e = W_0 \mu_e$, with $W_0 = M_i - M_f = 4.016$ MeV in the case of the ${}^6\text{He} - {}^6\text{Li}$ transition. M_i and M_f denote the masses of the initial and final state nucleus. Equation (14) is accurate up to corrections of $\mathcal{O}((W_0 r_\pi)^3)$. In Eq. (14), we kept the effects of nuclear recoil at leading order in W_0/M_f . The other effect of nuclear recoil is that the electron endpoint energy shifts from $E_e = W_0$ to $E_e = W_0 - \frac{W_0^2 - m_e^2}{2M_f}$.

In addition to the leading terms in the multipole expansion, Eq. (14) includes electromagnetic effects, which are not negligible at the precision we are working. The most important contribution is from the Fermi function, given by $F_0(Z, \varepsilon)$:

$$\begin{aligned} F_0(Z, \varepsilon) &= 4(2|\mathbf{p}_e|R)^{2(\gamma_0-1)} \frac{|\Gamma(\gamma_0 + iy)|^2}{|\Gamma(2\gamma_0 + 1)|^2} e^{\pi y}, \\ \gamma_0 &= \sqrt{1 - (\alpha Z)^2}, \quad y = \frac{\alpha Z}{|v_e|}, \end{aligned} \quad (15)$$

where $R = \sqrt{5/3\langle r_{\text{ch}}^2 \rangle}$, and $\langle r_{\text{ch}}^2 \rangle$ is the charge radius of ${}^6\text{Li}$, $\sqrt{\langle r_{\text{ch}}^2 \rangle} = 2.5890(390)$ fm [75]. Z is the charge of the daughter nucleus and the electron velocity $v_e = |\mathbf{p}_e|/E_e$. Other large corrections arise from the radiative corrections Δ_R^V and $\delta_R(Z, \varepsilon)$, which contribute to the half-life at the level of a few percent. These and other higher-order corrections, encoded in the functions $L_0(Z, \varepsilon)$, $S(Z, \varepsilon)$, $R_N(\varepsilon)$ and in the explicit $\mathcal{O}(\alpha)$ and $\mathcal{O}(\alpha^2)$ terms in Eq. (14), are discussed in Appendix E. In this work, we did not attempt a rederivation of electromagnetic corrections in chiral EFT and instead followed closely the literature, as summarized in Ref. [76]. To assess the importance of higher-order terms in the multipole expansion, we define the leading contribution to the spectrum as

$$\frac{d\Gamma_0}{d\varepsilon} = \frac{G_F^2 W_0^5 V_{ud}^2}{2\pi^3} \sqrt{1 - \frac{\mu_e^2}{\varepsilon^2}} \varepsilon^2 (1 - \varepsilon)^2 \frac{4\pi}{3} |L_1^{(0)}|^2 \times F_0(Z, \varepsilon) L_0(Z, \varepsilon) S(Z, \varepsilon) R_N(\varepsilon), \quad (16)$$

where we included in the definition of $d\Gamma_0/d\varepsilon$ some radiative corrections, for which we just use results in the literature.

III. THEORETICAL FRAMEWORK

A. Quantum Monte Carlo methods

In this work, we employ Quantum Monte Carlo methods [48]—both the variational (VMC) and the Green’s function Monte Carlo (GFMC) methods—and the Norfolk chiral effective field theory many-body interactions and electroweak currents [56,59,60,62,69] to evaluate the required nuclear matrix elements. This computational scheme was most recently described in Refs. [66,67] where some of the present authors evaluated Gamow-Teller matrix elements entering β decays and electron captures in light nuclei as well as muon capture rates in $A = 3$ and 6 nuclei. Here, we will not provide the details of the computational method or the interactions. We will instead limit ourselves to briefly summarize the salient points of the calculation and defer the interested reader to Ref. [66] and references therein for additional technicalities.

The Norfolk potentials consist of local two- and three-nucleon interactions formulated in configuration space, and derived from a chiral effective field theory that retains, in addition to nucleons and pions, Δ isobars as explicit degrees of freedom [56,60,62,69]. They are denoted below as NV2+3, where the two-body interaction (NV2) is constructed up to next-to-next-to-next-to leading order ($N^3\text{LO}$) in the chiral expansion, and the three-body force (NV3) retains up to $N^2\text{LO}$ contributions [77].

In the QMC calculation, theoretical uncertainties arise from deficiencies in the nuclear wave function, i.e., from how well the QMC wave function reproduces the actual ground states for a given nuclear interaction, and from the nuclear interactions themselves. In chiral EFT, the uncertainties in the nuclear interactions stem from the two- and three-nucleon data used to determine the unknown low-energy constants (LECs) in the nuclear potential and currents, from the residual dependence of observables on the cutoff used in the calculation and from the truncation error arising from working at a finite order in the chiral expansion. In order to assign a theoret-

ical error to our estimates, we perform the calculations using four models of Norfolk interactions. These models differ in the cutoff utilized to regularize divergences, in the number of nucleon-nucleon scattering data used to constrain the LECs entering the NV2 interaction, and in the fitting procedure implemented to constrain the NV3 interaction. In particular, NV2+3 models belonging to class I (denoted with NV2+3-I) are fitted up to 125 MeV and use ≈ 2700 data points, while the NV2+3-II models are fitted up to 200 MeV and use ≈ 3700 data points. For each class, two different sets of cutoff are implemented. Specifically, the coordinate space regulators are

$$C_{R_S}(r) = \frac{1}{\pi^{3/2} R_S^3} e^{-(r/R_S)^2}, \quad (17)$$

$$C_{R_L}(r) = 1 - \frac{1}{(r/R_L)^6 e^{2(r-R_L)/R_L} + 1},$$

where C_{R_L} regulates divergences at $r \approx 0$ in the pion exchange potential, while contact interactions are regulated by C_{R_S} . Models labeled with an “a” use the combination $(R_S, R_L) = (0.8, 1.2)$ fm (models NV2-Ia and NV2-IIa), while those labeled with a “b” use $(R_S, R_L) = (0.7, 1.0)$ fm (models NV2-Ib and NV2-IIb). The NV2 models are supplemented by a three-body force at $N^2\text{LO}$. At this order, there are two LECs characterizing the NV3’s strength. They are determined by a simultaneous fit to either the trinucleon binding energy and the triton β decay reduced matrix element [62] or the trinucleon binding energy and the nd scattering length [78]. Norfolk models based on the former procedure are denoted with a “*” , that is NV2+3*.

Nuclear wave functions are constructed in two steps. First, a trial variational Monte Carlo (VMC) wave function Ψ_T , which accounts for the effect of the nuclear interaction via the inclusion of correlation operators, is generated by minimizing the energy expectation value with respect to a number of variational parameters. The second step improves on Ψ_T by eliminating excited state contamination. This is accomplished in a Green’s function Monte Carlo (GFMC) calculation which propagates the Schrödinger equation in imaginary time τ . The propagated wave function $\Psi(\tau) = \exp[-(H - E_0)\tau]\Psi_T$, for large values of τ , converges to the exact wave function with eigenvalue E_0 . Ideally, the matrix elements should be evaluated in between two propagated wave functions. In practice, we evaluate mixed estimates in which only one wave function is propagated, while the remaining one is replaced by Ψ_T . The calculation of diagonal and off-diagonal matrix elements is discussed at length in Refs. [48,79]. We will present both VMC and GFMC results. As discussed in Sec. IV A, while the latter are more accurate and are in excellent agreement with the experimental half-life, the two calculations of the spectral shape show minimal differences, well below the 10^{-3} level, justifying the use of the numerically cheaper VMC in future studies.

B. Power counting and many-body electroweak currents

Accompanying the Norfolk many-body interactions are one- and two-body axial and vector currents derived within the same chiral effective field theory formulation with pions,

TABLE I. Scaling in $Q^\nu = Q^{n-3}$ up to $\nu = 1$ of the chiral axial current, $\mathbf{j}^\nu(A)$, and charge, $\rho^\nu(A)$, operators and of the vector current operator, $\mathbf{j}^\nu(V)$. The acronyms stand for 1b = one-body, OPE = one-pion-exchange, TPE = two-pion exchange, NR = non-relativistic, RC = relativistic correction, OPE- Δ = one-pion-exchange currents with an intermediate delta excitation, and sub = subleading. “-” indicates that no contribution exists at that order, while “ \times ” that contributions of that order have not been included. We also indicate to which multipole operator, and at which order in the chiral expansion, each term contributes. See text for explanation.

Operator	LO	NLO	N ² LO	N ³ LO	N ⁴ LO
	$\nu = -3$	$\nu = -2$	$\nu = -1$	$\nu = 0$	$\nu = 1$
$\mathbf{j}^\nu(A)$	1b(NR)	-	1b(RC)	OPE(sub)	\times
Multipole	$\{L_1, E_1\}^{(i,0)}$	$\{L_1, E_1\}^{(i,1)}$	$\{L_1, E_1\}^{(i,2)}$	$\{L_1, E_1\}^{(i,3)}$	$\{L_1, E_1\}^{(i,4)}$
$\rho^\nu(A)$	-	1b(RC)	OPE	-	\times
Multipole	$C_1^{(i,0)}$	$C_1^{(i,1)}$	$C_1^{(i,2)}$	$C_1^{(i,3)}$	$C_1^{(i,4)}$
$\mathbf{j}^\nu(V)$	-	1b(NR)	OPE	1b(RC)	OPE(sub)
Multipole	$M_1^{(i,0)}$	$M_1^{(i,1)}$	$M_1^{(i,2)}$	$M_1^{(i,3)}$	$M_1^{(i,4)}$

nucleons, and Δ 's [50,51,59,60,62,63]. We use the axial and vector charges, $\rho(A)$ and $\rho(V)$, and currents, $\mathbf{j}(A)$ and $\mathbf{j}(V)$, to evaluate the SM multipoles of Eqs. (4)–(7). The current operators are arranged in powers of a second expansion parameter, namely m_π/Λ_χ or equivalently $|\mathbf{p}|/\Lambda_\chi$, where $|\mathbf{p}| \approx m_\pi$ denotes a typical nuclear physics scales such as the binding momentum, the inverse of the nuclear radius $R_A = 1.2A^{1/3}$ fm (with A being the mass number), or the typical nucleons' momenta inside nuclei. As a consequence, the coefficients of the q expansion in Eqs. (10)–(13), i.e., the multipoles $C_1^{(i)}$, $L_1^{(i)}$, $M_1^{(i)}$, and $E_1^{(i)}$, have a second expansion in Q^n , where $Q \equiv m_\pi/\Lambda_\chi$ is the chiral expansion parameter. We can then express any of the multipoles—here generically denoted with $\mathcal{M}^{(i)}$ —as

$$\mathcal{M}^{(i)} = \sum_{n=0}^{\infty} \mathcal{M}^{(i,n)}. \quad (18)$$

where the superscripts i and n indicate the orders of the multipole and chiral expansions, respectively.

The scaling in Q^n of the chiral electroweak currents derived in Refs. [50,51,59,60,62,63] is reported in Table I. The aforementioned references adopt a different convention where the counting is carried out in powers of Q^ν with $\nu = n - 3$. Without going too much into details, the operators consist of one-body contributions obtained from the non-relativistic reduction of the covariant axial and vector nucleonic four-vector currents. We denote with “1b(NR)” and “1b(RC)” leading-order and first-order terms in the nonrelativistic expansion. Two-body currents include contributions of one- and two-pion range (OPE and TPE) as well as short-range currents encoded in contact-like operators (CT). OPE currents involving nucleons' virtual excitations into a Δ are denoted with “OPE- Δ ” while those involving subleading terms in the pion-nucleon

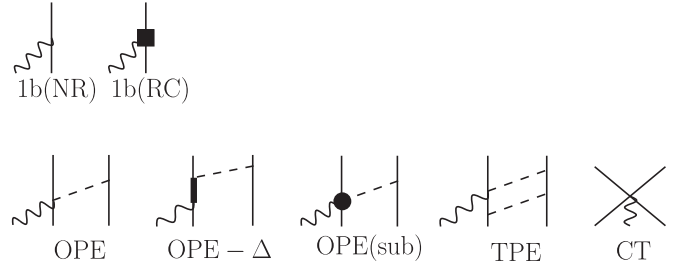


FIG. 1. Schematic representations of the types of contributions entering the one- and two-body electroweak currents from Refs. [50,51,59,60,62,63] adopted in this work. Solid, dashed, and wavy lines represent nucleons, pions, and axial and vector external fields. The square and the dot represent the relativistic corrections to the leading one-body operators and subleading terms in the pion-nucleon Lagrangian, respectively, while the thick line represents a Δ intermediate state. See Table I and Appendix C for the operators' scaling and formal expressions. Not shown in the figure are the pion-pole and tadpole diagrams (see Refs. [50,51,59,60,62,63] for details.)

chiral Lagrangian are denoted with “OPE(sub).” In addition, there are the so called pion-pole contributions where the external field couples with a pion that is then absorbed by a nucleon. These operators are schematically represented in Fig. 1 while their formal expressions are listed for convenience in Appendix C. Details on the derivation of the currents can be found in Refs. [50,51,59,60,62,63].

For our discussion, it is sufficient to focus on the leading order terms of both the charge and current operators. The E_1 and L_1 multipoles are proportional to matrix elements of the axial current, which, at leading order in the chiral expansion, is given by the usual Gamow-Teller and pion-pole contributions [see Eq. (C6) in Appendix C]. This implies that

$$E_1^{(i,0)} \quad \text{and} \quad L_1^{(i,0)} \approx \mathcal{O}(Q^0) = \mathcal{O}(1), \quad (19)$$

as can be inferred from Table I. In particular, $E_1^{(0,0)}$ and $L_1^{(0,0)}$ are determined by the zero-momentum Gamow-Teller matrix element in Eq. (8). As illustrated in Table I, two-body axial currents first contribute to $E_1^{(0,2)}$ and $L_1^{(0,2)}$ with the OPE- Δ term, while subleading OPE diagrams and contact interactions to $E_1^{(0,3)}$ and $L_1^{(0,3)}$. With the interactions and axial current that we use, E_1 and L_1 are accurate up to order $n = 3$ ($\nu = 0$) in the chiral expansion. The $E_1^{(2,0)}$ and $L_1^{(2,0)}$ multipoles are also nonzero, and they are related to the square radius of the Gamow-Teller matrix element, for which the relevant scale is the system's binding momentum. The one-body induced pseudoscalar form factor, which, as shown in Eq. (C6), in momentum space scales as q^2/m_π^2 , gives additional LO contributions to $L_1^{(2)}$. The momentum dependence of the nucleon axial form factor, on the other hand, is suppressed and contributes at N²LO in the chiral expansion to both the $E_1^{(2,2)}$ and $L_1^{(2,2)}$ multipoles. The expressions of the nucleon axial form factor used in the currents is given in Appendix C.

The C_1 multipole is induced by the axial charge, which, as shown in Table I, at lowest order receives one-body contributions from the non-relativistic expansion of the axial form factor, starting at $\mathcal{O}(1/m_N)$, and from the induced

pseudoscalar form factor. As shown in Appendix C, the latter contribution is proportional to the electron endpoint energy W_0 , which, for power counting purposes, scales as $W_0 \approx O(m_N Q^2)$. As a consequence,

$$C_1^{(1,0)}(A) \approx 0 \quad \text{and} \quad C_1^{(1,1)}(A) \approx O(Q^1). \quad (20)$$

We then expect $C_1^{(1,1)}(A)$ to be suppressed with respect to the $L_1^{(i)}$ and $E_1^{(i)}$ multipoles. The two-body axial charge operator scale as Q^2 (see Table I) and thus contributes to the $C_1^{(1,2)}$ multipole.

Finally, the magnetic multipole is induced at the one-body level by the weak magnetic form factor, or equivalently by the one-body vector current at leading order—see Eq. (C10) and Table I—and thus also in this case

$$M_1^{(1,0)}(V) \approx 0, \quad M_1^{(1,1)}(V) \approx O(Q^1), \quad (21)$$

even though the large isovector anomalous magnetic moment $\kappa_V = \kappa_p - \kappa_n \approx 3.7$ enhances the formally NLO contribution. Two-body currents start to contribute to $M_1^{(1,2)}(V)$ and are found to provide a 6% to 8% contribution to the overall matrix element.

We can use the counting of currents and potentials to give a rough estimate of the truncation error we expect for different multipoles. The coordinate space cutoff R_S can be converted into a scale $\Lambda_\chi = 2/R_S$, and, for the potentials used in this work $\Lambda_\chi \approx 500\text{--}550$ MeV. For the multipoles L_1 , E_1 , and M_1 , the first term that is missed in the calculation is of order $Q^4 = 0.6\%$, taking $\Lambda_\chi = 500$ MeV. This error is, as we will see, smaller than the uncertainties arising from using different models. For C_1 , the first missing term in the chiral expansion is $O(Q^3) \approx 2\%$. These estimates are merely indicative. This type of estimation has been the standard procedure in nuclear physics; however, for a more robust estimate, one should develop Bayesian methods to quantify the uncertainties associated with parameters entering the many-body calculations and truncation errors. Work along this line is being vigorously pursued by the community [80–85], and is beyond the scope of this work.

IV. STANDARD MODEL RESULTS AND SPECTRUM

In order to determine the matrix elements entering the β spectrum in Eq. (14) we calculate Eqs. (4)–(7) for six momenta between 0 and 0.25 fm^{-1} (0 and ≈ 50 MeV), and

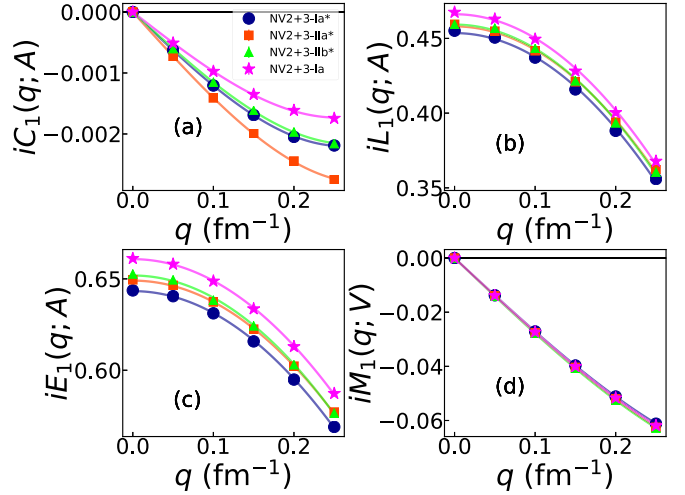


FIG. 2. VMC multipoles for the NV2+3-Ia* (blue circles), NV2+3-IIa* (orange squares), NV2+3-IIb* (green triangles), and NV2+3-Ia (magenta stars) models. The curves of best fit for each case are shown in the same color as the multipoles. Statistical errors from the Monte Carlo are included on each point but are too small to be visible in the figure.

we fit them to the functional forms in Eqs. (10)–(13). To have a realistic assessment of the theoretical uncertainty, we performed the calculation with four sets of chiral potentials and currents, NV2+3-Ia*, NV2+3-IIa*, NV2+3-IIb*, and NV2+3-Ia, using both VMC and GFMC methods.

Figure 2 shows the VMC multipoles and the associated curves of best fit for the NV2+3-Ia*, NV2+3-IIa*, NV2+3-IIb*, and NV2+3-Ia models retaining one- and two-body vector and axial current operators. The expansion coefficients obtained by fitting VMC multipoles obtained with one-body and one- and two-body operators are listed in Table II, where the error denotes the fitting error. Two-body currents have a minor effect on the L_1 and E_1 multipoles, leading to a shift in $L_1^{(0)}$, $E_1^{(0)}$, $L_1^{(2)}$, $E_1^{(2)}$ of at most $\approx 2\%$, in the case of model Ia. As expected from power counting, two-body currents are more important for M_1 and C_1 . In both cases, two-body currents contribute at $O(Q)$ compared to the LO. We see that $M_1^{(1)}$ receives an $\approx 8\%$ correction, for all the models considered here. In VMC, the corrections to $C_1^{(1)}$ are about 30–40%.

TABLE II. Expansion coefficients of the VMC multipole operators, including only one-body currents (1b) or one- and two-body currents (2b). The four columns denotes four different NV interactions, as discussed in the text. The error denotes the fitting error.

	Model Ia*		Model IIa*		Model IIb*		Model Ia	
	1b	1b+2b	1b	1b+2b	1b	1b+2b	1b	v1b+2b
$L_1^{(0)}$	1.3578(3)	1.3607(4)	1.3662(2)	1.3742(3)	1.3717(2)	1.3777(1)	1.3641(3)	1.3986(3)
$E_1^{(0)}$	1.9255(4)	1.9299(5)	1.9355(2)	1.9471(3)	1.9470(3)	1.9555(1)	1.9333(4)	1.9824(4)
$M_1^{(1)}$	-0.5487(1)	-0.5860(2)	-0.5510(1)	-0.5952(2)	-0.5550(1)	-0.60111(4)	-0.5521(1)	-0.5908(2)
$C_1^{(1)}$	-0.0182(3)	-0.0269(3)	-0.0217(3)	-0.0311(3)	-0.0151(3)	-0.0257(3)	-0.0128(3)	-0.0218(3)
$L_1^{(2)}$	23.87(5)	23.93(6)	23.51(3)	23.51(4)	23.94(3)	24.07(1)	23.64(5)	24.12(5)
$E_1^{(2)}$	17.86(7)	17.99(8)	17.35(5)	17.49(6)	18.07(5)	18.22(2)	17.45(7)	17.80(7)

TABLE III. Summary of the values of the expansion coefficients for the charge (C_1), longitudinal (L_1), electric (E_1), and magnetic (M_1) VMC multipoles for the NV2+3 models under study. The percent error due to the cutoff ϵ_C (%), energy range of the fit ϵ_E (%), the three-body force fit ϵ_{3N} (%), and the total error ϵ_{tot} (%) are also presented. Details of how the average expansion coefficients are obtained are provided in the text. The uncertainty on coefficients for individual models is fitting error only. The error is <0.001 unless otherwise noted.

Model	$C_1^{(1)}$	$L_1^{(0)}$	$L_1^{(2)}$	$E_1^{(0)}$	$E_1^{(2)}$	$M_1^{(1)}$
Ia*	-0.027	1.361	23.93(6)	1.923	17.99(8)	-0.586
IIa*	-0.031	1.374	23.51(4)	1.947	17.49(6)	-0.595
IIb*	-0.026	1.378	24.07(1)	1.955	18.22(2)	-0.601
Ia	-0.022	1.399	24.12(5)	1.982	17.80(7)	-0.591
Average	-0.026	1.378	23.869	1.954	17.847	-0.593
ϵ_E (%)	5.4	0.3	3.2	0.3	5.6	0.6
ϵ_C (%)	6.9	0.01	0.3	0.07	0.8	0.3
ϵ_{3N} (%)	7.2	0.8	7.0	0.8	13.9	0.3
ϵ_{tot} (%)	11.8	0.8	7.8	0.9	15.0	0.7

A. Uncertainty estimation

1. Variational Monte Carlo

The expansion coefficients for the NV2+3-Ia*, NV2+3-IIa*, NV2+3-IIb*, and NV2+3-Ia models are presented in Table III. From Figure 2 and Table II, it is clear that there will be some degree of uncertainty due to the choice of model. To account for the model uncertainty, four sets of average multipoles were obtained: an average between two models with the same cutoff, same determination of the three-body force, and fit to different ranges of NN scattering data (NV2+3-Ia* and NV2+3-IIa*); an average between two models fit to the same range of NN scattering data, same determination of the three-body force, and different cutoffs (NV2+3-IIa* and NV2+3-IIb*); an average between two models with consistent NV2 and cutoff but different NV3 interactions (NV2+3-Ia* and NV2+3-Ia); and an average of all four models. These average multipoles were fit to the functional forms of Eqs. (10)–(13) with the two model calculations at each q providing the upper and lower theoretical error bars. The statistical uncertainty of these averaged fits provides the estimated uncertainty on the expansion coefficients due to the energy range of the fit (ϵ_E), the choice of cutoff (ϵ_C), and the three-body force fitting procedure (ϵ_{3N}).

For the fourth case described above, a similar procedure was followed; however, when assigning the error for the multipoles at each value of q , the uncertainty was not taken as simply the spread of the model calculations. Instead, we summed in quadrature the uncertainties on each point from the other three average fits to combine the cutoff, energy range, and three-body force uncertainties. The expansion coefficients obtained from this fit are the ‘‘Average’’ results in Table III and their statistical uncertainties are cited as ϵ_{tot} . Figure 3 shows the averaged multipoles and curves of best fit for these three cases.

As evidenced by the results in Table III, the coefficient $E_1^{(2)}$ has the largest model uncertainty at 15%. The next largest

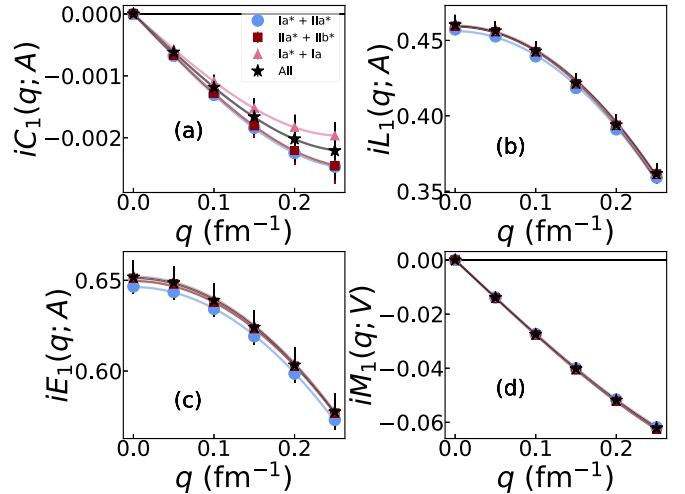


FIG. 3. Average VMC multipoles for the four NV2+3 models under study. Averages were obtained for NV2+3 potentials with the same cutoff (blue circles), the same energy range of NN scattering data used to fit the interaction (red squares), the same NN but different $3N$ force (pink triangles), and for all models (black stars). The curves of best fit for each case are shown in the same color as the average multipoles. Details on how the error bars are obtained in each case are provided in the text.

uncertainties are on the coefficients $C_1^{(1)}$ at 11.8% and $L_1^{(2)}$ at 7.8%. The remaining uncertainties are $\lesssim 1\%$ for $L_1^{(0)}$, $E_1^{(0)}$, and $M_1^{(1)}$. The main driver of the uncertainty in most of the coefficients is the choice of three-body force, though for $C_1^{(1)}$ and $M_1^{(1)}$ the uncertainties from this source and the others are comparable. In fact, it is the energy range of scattering data used to fit the interaction that provides the largest uncertainty on $M_1^{(1)}$.

We analyzed the stability of our fits by performing tests with one model, namely NV2+3-Ia*. To investigate if the range of momentum points is sufficient, we fit with the same grid size but one more point and one fewer point; i.e., we refit on the ranges $[0, 0.20] \text{ fm}^{-1}$ and $[0, 0.30] \text{ fm}^{-1}$. The leading order coefficients in the polynomials are altered by less than one percent in this procedure. In the quadratic terms, $L_1^{(2)}$ and $E_1^{(2)}$, we observe changes that are on the order of a few percent but well within the quoted model uncertainty. The effect of a more fine grid in momentum was investigated by reducing the grid spacing by half, thus doubling the number of points included in the fit. We find that this procedure changes all coefficients by less than one percent. Finally, we investigated the inclusion of higher order terms in the polynomials. This does not significantly alter the leading order coefficients; however, it does change the quadratic coefficients on the order of the model uncertainty. While this change is significant, we find that the coefficient introduced at order $\mathcal{O}(q^4)$ is highly correlated with the one at order $\mathcal{O}(q^2)$. Further, including this term reduced the residuals of the fit from $\mathcal{O}(10^{-2})$ to $\mathcal{O}(10^{-4})$. Successively adding higher order terms in the expansion leads to similar order of magnitude reductions in the residuals. This successive addition of terms was tested up to $\mathcal{O}(q^{10})$ with no convergence in the residuals observed. Thus, to avoid overfit-

TABLE IV. Expansion coefficients of the GFMC multipoles and comparison with the VMC, using three different Norfolk models. The error denotes fitting error only

Model	Method	$C_1^{(1)}$	$L_1^{(0)}$	$L_1^{(2)}$	$E_1^{(0)}$	$E_1^{(2)}$	$M_1^{(1)}$
Ia*	VMC	-0.027	1.361	23.93(6)	1.930	17.99(8)	-0.586
	GFMC	-0.047	1.308	23.98(6)	1.856	19.71(9)	-0.559
IIa*	VMC	-0.031	1.374	23.51(4)	1.947	17.49(6)	-0.595
	GFMC	-0.045(2)	1.341(7)	25(1)	1.901(9)	19(2)	-0.573(4)
Ia	VMC	-0.022	1.399	24.12(5)	1.982	17.80(7)	-0.591
	GFMC	-0.045	1.360	23.64(6)	1.929	17.77(9)	-0.575
Average	VMC	-0.026(3)	1.38(1)	24(2)	1.95(2)	18(3)	-0.593(4)
	GFMC	-0.046(2)	1.34(2)	24(3)	1.90(3)	19(5)	-0.568(9)

ting $L_1^{(2)}$ and $E_1^{(2)}$, we opted to truncate the expansion of the even polynomials at $O(q^2)$.

Another way that one might obtain the average expansion parameters with an uncertainty would be to instead use the results of Table II. The average value of the expansion coefficients for the four models differs from the values obtained in our procedure by less than one percent. Then, one could estimate the uncertainty due to the energy range used to fit NN interaction from the difference between the NV2+3-Ia* and NV2+3-IIa* results, the uncertainty due to the cutoff from the difference between the NV2+3-IIa* and NV2+3-IIb* results, and the uncertainty due to the three-body force from the difference between the NV2+3-Ia* and NV2+3-Ia results. Following this procedure, the error on the leading order expansion coefficients increase by factors of 1.5 to 1.8. The error on the dominant contributions would be $\lesssim 1.4\%$ in this approach. For the coefficients $L_1^{(2)}$ and $E_1^{(2)}$, this scheme reduces the uncertainty by a factors of 2.6 and 3.0, respectively. However, the approach that we took to obtain the uncertainty is more reasonable as each matrix element should have its own model uncertainty that then propagates to the expansion coefficients in the fitting procedure.

2. Green's function Monte Carlo

In addition to a calculation of the multipoles using VMC, we also performed calculations at the Green's function Monte Carlo level, to remove residual excited state contamination in the nuclear wave functions. Following the procedure of Ref. [79], we perform a mixed estimate extrapolation of the GFMC multipoles. We then fit the expansion coefficients of Eqs. (10)–(13) as was done for the VMC calculations. The results of these fits are presented in Table IV for models Ia*, IIa*, and Ia. The GFMC evolution reduces the leading coefficients, $L_1^{(0)}$ and $E_1^{(0)}$, by ≈ 3 –4%, which, as we will see, results in better agreement with the experimental half-life. Propagating models Ia* and Ia changes the remaining expansion coefficients at the level of a few percent while for model IIa* terms that were higher order in qr_π experienced a more significant percent change after the propagation. This can be understood by looking at the system size as a function of τ during the GFMC propagation. To further understand this, we calculate the point proton radius in GFMC and we observe that for model IIa*, the system size grows

much more rapidly in τ than models Ia and Ia*. This behavior is due to the proximity of the model IIa* ${}^6\text{He}(0^+; 1)$ [${}^6\text{Li}(1^+; 0)$] ground state energy in GFMC to the $\alpha + 2n$ [$\alpha + d$] breakup threshold. Because of terms going like $e^{-iq \cdot r_i}$ in the current operators, the monotonic increase of the system size will impact the convergence of the IIa* matrix elements needed to determine the multipoles. To account for this in the GFMC extrapolation with model IIa*, we adopt a procedure used for systems near threshold and broad resonances [57]. We note that while the system sizes grow with τ , the ground state energies of ${}^6\text{Li}$ and ${}^6\text{He}$ drop rapidly and stabilize near $\tau \approx 0.1 \text{ MeV}^{-1}$. This indicates that spurious contamination has been removed from the wave functions at that point. Under this assumption, we extract the values of the matrix elements by performing a linear fit to the matrix element in the interval $\tau = [0.1 \text{ MeV}^{-1}, 0.3 \text{ MeV}^{-1}]$ and extrapolating back to 0.1 MeV^{-1} . We determine the systematic error of this procedure by averaging in the intervals $\tau = [0.08 \text{ MeV}^{-1}, 0.3 \text{ MeV}^{-1}]$ and $\tau = [0.12 \text{ MeV}^{-1}, 0.3 \text{ MeV}^{-1}]$ to get a conservative estimate.

For the GFMC extrapolations, we assigned errors to the average matrix element arising from the energy range and three-body force following the same procedure as was done for the VMC. We also include the systematic uncertainty from the model IIa* extrapolation by summing it in quadrature with the energy range and three-body force uncertainties. For the GFMC average, the coefficients $L_1^{(0)}$, $E_1^{(0)}$, $C_1^{(1)}$, and $M_1^{(1)}$ have uncertainties ranging from $\approx 1.4\%$ to $\approx 5.2\%$. The coefficients $L_1^{(2)}$ and $E_1^{(2)}$ have uncertainties of $\approx 12\%$ and $\approx 26\%$, respectively. In the expression for the rate, the large error coefficients appear suppressed by powers of qr_π and do not contribute as strongly as the coefficients with errors of order 1%.

Because of the small cutoff uncertainty, we estimate the impact of including model IIb* by reweighting the matrix elements in the average so that model IIa* is counted twice. Averaging under this assumption provides coefficients and errors that are consistent with the average results obtained when omitting model IIb*. Thus, we conclude that we can safely neglect model IIb* in the GFMC average.

B. Experimental comparison and remaining spectral uncertainty

We can first of all check the $L_1^{(0)}$ and $E_1^{(0)}$ multipoles by comparing our calculation with experimental half-life $\tau_{1/2} =$

$807.25 \pm 0.16 \pm 0.11$ ms [86,87]. Using the VMC matrix elements in Table III, we obtain

$$\begin{aligned} \tau_{1/2}|_{\text{VMC}} &= \frac{\ln 2}{\Gamma} = \frac{1}{V_{ud}^2 g_A^2} (1175 \pm 17) \text{ ms} \\ &= (762 \pm 11 \pm 2) \text{ ms}, \end{aligned} \quad (22)$$

where we have used $V_{ud} = 0.97370 \pm 0.00031$ and $g_A = 1.2754 \pm 0.0013$. The first error is due to the nuclear matrix elements while the second is due to g_A . The error from V_{ud} is negligible. Equation (22) deviates by about 5% from the observed value. While small, the discrepancy between the VMC calculation and the experiment is not covered by the uncertainty range in Eq. (22), indicating that the errors due to either the chiral EFT truncation or the Monte Carlo method are underestimated. For this reason, we evolved the VMC wave functions in the GFMC. Using the GFMC matrix

elements, we obtain

$$\tau_{1/2}|_{\text{GFMC}} = \frac{1}{V_{ud}^2 g_A^2} (1246 \pm 37) \text{ ms} = (808 \pm 24 \pm 2) \text{ ms}, \quad (23)$$

which has a 3% error and is in perfect agreement with the observed half-life. The half-life is dominated by $L_1^{(0)}$, $E_1^{(0)}$, and by the Fermi function. The next most important correction comes from inner and outer radiative corrections, Δ_R^V and $\delta_R(Z, \varepsilon)$, which together shorten the half-life by about 4%. For Δ_R^V we use the dispersive evaluation of Refs. [16,17,88], $\Delta_R^V = 2.467(22) \times 10^{-2}$. Higher multipoles impact the half-life at the 0.1% level.

Moving on to the differential decay rate, using the VMC multipoles given in Table III, we find

$$\begin{aligned} \left. \frac{d\Gamma}{d\varepsilon} \right|_{\text{VMC}} &= \frac{d\Gamma_0}{d\varepsilon} \left\{ 1 + \left[\left(1 - 2\varepsilon + \frac{\mu_e^2}{\varepsilon} \right) (-1.16 \pm 0.01) + \left(1 - \frac{\mu_e^2}{\varepsilon} \right) (3.6 \pm 0.4) \times 10^{-2} \right] \times 10^{-2} \right. \\ &\quad - \left[\left(1 - \frac{\mu_e^2}{\varepsilon} (2 - \varepsilon) \right) (0.96 \pm 0.08) + \left(3 - 10\varepsilon(1 - \varepsilon) + \mu_e^2 \frac{4 - 7\varepsilon}{\varepsilon} \right) (0.32 \pm 0.06) \right] \times 10^{-3} \\ &\quad \left. + (4.1 \pm 1.1) \times 10^{-4} (1 - \varepsilon) \right\}, \end{aligned} \quad (24)$$

where the terms in the first line appear at NLO in the multipole expansion, and are proportional to $M_1^{(1)}$ and $C_1^{(1)}$, respectively. The terms in the second line appear at N²LO, the first proportional to $L_1^{(2)}$ and the second to a combination of $|M_1^{(1)}|^2$ and $E_1^{(0)}E_1^{(2)}$. Finally, the term in the third line is an electromagnetic correction proportional to subleading multipoles. Using the GFMC matrix elements in Table IV, the differential rate is

$$\begin{aligned} \left. \frac{d\Gamma}{d\varepsilon} \right|_{\text{GFMC}} &= \frac{d\Gamma_0}{d\varepsilon} \left\{ 1 + \left[\left(1 - 2\varepsilon + \frac{\mu_e^2}{\varepsilon} \right) (-1.15 \pm 0.02) + \left(1 - \frac{\mu_e^2}{\varepsilon} \right) (6.6 \pm 0.7) \times 10^{-2} \right] \times 10^{-2} \right. \\ &\quad - \left[\left(1 - \frac{\mu_e^2}{\varepsilon} (2 - \varepsilon) \right) (0.99 \pm 0.12) + \left(3 - 10\varepsilon(1 - \varepsilon) + \mu_e^2 \frac{4 - 7\varepsilon}{\varepsilon} \right) (0.35 \pm 0.10) \right] \times 10^{-3} \\ &\quad \left. + (4.0 \pm 1.8) \times 10^{-4} (1 - \varepsilon) \right\}. \end{aligned} \quad (25)$$

From Eqs. (24) and (25) we see that the dominant correction is, as expected [74,89], given by the magnetic multipole M_1 , which contributes to the spectrum at the percent level. The uncertainty on the ratio $M_1^{(1)}/L_1^{(0)}$, which dominates the error budget, is about 2%, and, in our calculation, it receives the main contribution from changing the energy range of the fits to NN scattering data. The next contributions come from $L_1^{(2)}$, $E_1^{(2)}$, and $C_1^{(1)}$, which affect the energy distribution at the 10^{-3} – 10^{-4} level. $L_1^{(2)}$, $C_1^{(1)}$ and $E_1^{(2)}$ have uncertainties of about 10%, 12%, and 20%, respectively, so that these terms contribute to the theory error at the 10^{-4} level.

In Fig. 4 we show the deviation of the β spectrum from the leading term in the multipole expansion, $d\Gamma_0/d\varepsilon$, defined

in Eq. (16), using both VMC and GFMC matrix elements. We see that, while GFMC and VMC differ by 6% on the total rate, the differences largely cancel in the ratio, and the corrections to the spectral shape are very similar in both cases. The bottom panel of Fig. 4 shows the error on the ratio. This is somewhat larger in GFMC, but well below 10^{-3} . Figure 5 shows the contributions of the leading $\mathcal{O}(Qr_\tau)$ correction, arising from M_1 , and of the second order terms $\mathcal{O}((Qr_\tau)^2)$ [including the formally $\mathcal{O}(Qr_\tau)$ but numerically small contribution from $C_1^{(1)}$], to the differential rate (left) and to the uncertainty (right). We see that, while M_1 dominates the correction to the spectrum, the second-order terms contribute at the 10^{-3} level and the first- and second-order terms give contributions of similar size to the uncertainty.

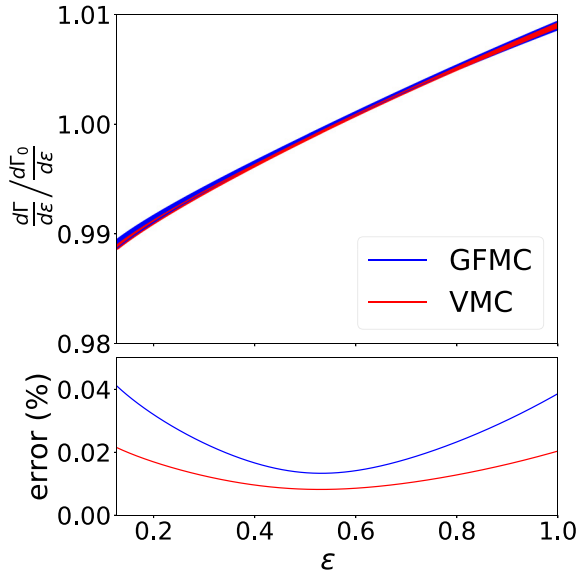


FIG. 4. Deviation of the ${}^6\text{He}$ β spectrum from the expression truncated at leading order in the multipole expansion, given in Eq. (16), and percentage error on the ratio. The blue curves use GFMC matrix elements in Table IV, while the red lines use the VMC averages in Table III. In the top panel, the width of the curves denotes the theoretical error.

I. Validation of M_1

We can further validate the calculation in several ways. The magnetic multipole M_1 can be cross-checked against data. Using the conserved vector current hypothesis, $M_1^{(1)}$ can be expressed in terms of the transition $\Gamma({}^6\text{Li}(0^+) \rightarrow {}^6\text{Li}(1^+)\gamma)$:

$$|M_1^{(1)}| = \frac{3m_\pi}{\sqrt{4\pi}} \sqrt{\frac{\Gamma_{M_1}}{\alpha E_\gamma^3}} = 0.582(6), \quad (26)$$

where the uncertainty is determined by the width, $\Gamma_{M_1} = 8.19(17)\text{eV}$ [75,90]. Using the average VMC value listed in Table III, our result agrees within 1% with the experimental

value, while individual models are consistent within 3% when including two body currents. The GFMC average is a bit lower, but still only 2% away from Eq. (26), and compatible within $\approx 1\sigma$. Our results also agree with the *ab initio* calculation of Ref. [46], which, including only the one-body piece of the current, finds $|M_1^{(1)}| = 0.565$.

Equation (26) is valid up to isospin breaking corrections. To check the level of isospin breaking that we can expect in $M_1^{(1)}$, we computed both the ${}^6\text{He}(0^+; 1) \rightarrow {}^6\text{Li}(1^+; 0)$ transition and the electromagnetic transition ${}^6\text{Li}(0^+; 1) \rightarrow {}^6\text{Li}(1^+; 0)$ with NV2+3-IIb wave functions in the impulse approximation, that is, retaining only one-body current operators at LO in the chiral expansion. Using VMC wave functions, the electromagnetic transition had a value 0.553 and the weak transition a value of 0.554. Agreement between the two M_1 calculations is thus achieved at the $\approx 0.1\%$ level. Propagating this calculation in GFMC, we find, in the electromagnetic case, $|M_1^{(1)}|_{\text{EM}} = 0.532$ and, for the weak transition, $|M_1^{(1)}| = 0.538$, showing a 0.9% difference. This analysis is performed with one nuclear interaction model and thus does not account for any model uncertainty.

The change in the level of agreement between VMC and GFMC can be understood as due to how the wave functions are generated in each method. For the VMC case, the variational parameters in the ${}^6\text{He}(0^+; 1)$ and ${}^6\text{Li}(0^+; 1)$ wave functions are minimized separately. Because of explicit isospin symmetry breaking terms in the potential, the parameters of the two wave functions differ; however, isospin symmetry breaking correlations are not turned on in the VMC wave functions. When the trial states are acted on with the imaginary time propagator, isospin breaking correlations are introduced. This, in turn, increases the disagreement for the electromagnetic and weak M_1 transitions. Because the systems differ by changing the isospin of one nucleon, the effect of this symmetry breaking is small and at the level of the experimental uncertainty in Eq. (26). Since isospin breaking corrections are smaller than the experimental error, we can use $M_1^{(1)}$ extracted from ${}^6\text{Li}(0^+; 1) \rightarrow {}^6\text{Li}(1^+; 0)$ to further

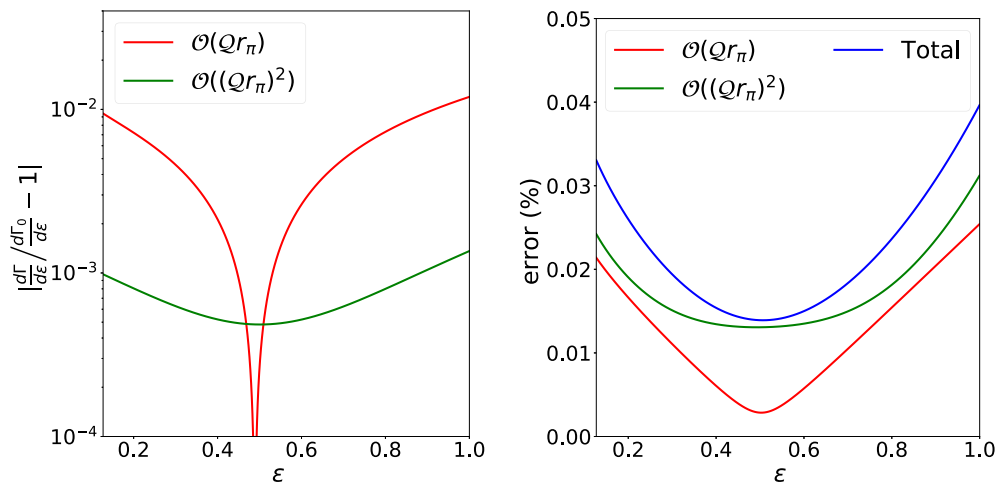


FIG. 5. Corrections to the β spectrum (*left*) and contributions to the error on the ratio between $d\Gamma/d\varepsilon$ and $d\Gamma_0/d\varepsilon$ (*right*) at next-to-leading order [$\mathcal{O}(Qr_\pi)$] and next-to-next-to leading order [$\mathcal{O}(Q^2r_\pi^2)$] in the multipole expansion. The figure uses GFMC matrix elements.

TABLE V. VMC and GFMC one-body and two-body averages for $C_1^{(1)}$ compared with Refs. [46,89]. Notice that to obtain $C_{\text{pseudo}}^{(1)}$ we rescale the results of Ref. [46] by $W_0/(W_0 + \Delta E_c)$, as we are not including Coulomb corrections to C_1 .

	VMC	GFMC	Calaprice [89]	Glick-Magid <i>et al.</i> [46]
Recoil	0.020(3)	-0.001(3)	-0.0144	-0.006
Pseudo	-0.040	-0.038		-0.039
2 body	-0.006	-0.007		
Total	-0.026(3)	-0.046(3)	-0.0144	-0.045

reduce the error on M_1 to the 1% level. Controlling isospin-breaking effects will become more important with improved measurements of the ${}^6\text{Li}(0^+) \rightarrow {}^6\text{Li}(1^+)\gamma$ transition, such as those suggested in Ref. [91].

2. Validation of C_1

The axial charge correction $C_1^{(1)}$, unlike the weak magnetism correction, is not constrained by an experimental datum. As such, its evaluation is critical for current experimental efforts aiming for the sub-0.1% level. The prefactor to $C_1^{(1)}$ from Eq. (14) is $2W_0r_\pi \approx 0.05$ and with $C_1^{(1)} \approx 1-2\%$ a relative $\approx 10-20\%$ precision is needed for a theory uncertainty to be smaller than one part in 10^4 . As can be seen in Eq. (C4), at lowest order in chiral EFT, the axial charge receives a $\mathcal{O}(1/m_N)$ recoil contribution from the coupling of the axial current to the nucleon, and a contribution from the induced pseudoscalar form factor, proportional to the energy of the electron and neutrino $E_e + E_\nu = W_0$. Naively, these would translate in an $\mathcal{O}(m_\pi/m_N)$ and $\mathcal{O}(W_0/m_\pi)$ correction to C_1 , respectively. The matrix element of the recoil component, however, vanishes for transitions between the dominant S state components of the wave function [89], making this component particularly sensitive to other wave function admixtures and two-body currents. In our calculation, indeed, C_1 is dominated by the induced pseudoscalar contribution, which, being proportional to $L_1^{(0)}$, is fairly stable. We can separate $C_1^{(1)}$ into three pieces,

$$C_1^{(1)} = C_{\text{recoil}}^{(1)} + C_{\text{pseudo}}^{(1)} + C_{\text{2-body}}^{(1)}, \quad (27)$$

where $C_{\text{recoil}}^{(1)}$ and $C_{\text{pseudo}}^{(1)}$ are given by the matrix element of $\rho_{\text{recoil}}^{-2}(\mathbf{q}, A)$ and $\rho_{\text{pseudo}}^{-2}(\mathbf{q}, A)$ in Eq. (C4), while $C_{\text{2-body}}^{(1)}$ is given by the matrix element of $\rho^{-1}(\mathbf{q}, A)$ in Eq. (C5). Results for each component, after averaging over different models as discussed in Sec. IV A, are given in Table V.

The calculation of $C_{\text{recoil}}^{(1)}$ in the decay of ${}^6\text{He}$, together with other triplet decays in the mass $A = 8, 12, 20$ systems, received a significant amount of attention over 40 years ago but ultimately remained unresolved [89,92-94]. Until recently, the only theoretical determination of the axial charge contribution to ${}^6\text{He}$ was performed by Calaprice [89] in Holstein's formalism. A direct comparison can be made by observing [95]

$$C_{\text{recoil}}^{(1)} = -\sqrt{\frac{3(2J_i + 1)}{4\pi}} \frac{d(0)}{2r_\pi M}, \quad (28)$$

where $d(q^2)$ is the so-called induced tensor form factor. Using wave functions tuned to reproduce the experimental energy levels, Calaprice obtained $d(0) = 2.4$ [89], which can be converted into $C_{\text{recoil}}^{(1)} = -0.0144$. This value is in the same ballpark of the VMC calculation, which, however, has an opposite sign. After GFMC evolution, the recoil contribution is reduced and qualitatively agrees with the result of Ref. [46].

To further track down the origin of the discrepancy, we notice that one can write

$$C_{\text{recoil}}^{(1)} = -\sqrt{\frac{3}{4\pi}} \frac{g_A}{2r_\pi m_N} (\mathcal{M}_{\sigma L} + m_N \mathcal{M}_{\sigma rp}), \quad (29)$$

where $\mathcal{M}_{\sigma L}$ and $\mathcal{M}_{\sigma rp}$ are one-body matrix elements defined in Ref. [74]. Calaprice assumed the first matrix element to be dominant and neglected $\mathcal{M}_{\sigma rp}$. To check the assumption, we calculated $\mathcal{M}_{\sigma L}$ in model NV2+3-Ia*. In the VMC, the ${}^6\text{He}$ and ${}^6\text{Li}$ wave functions are expressed in terms of the action of correlation operators on single particle wave functions with an α core and two nucleons whose wave functions are p -wave solutions of an effective $\alpha - N$ potential [45,96]. The two nucleons can be coupled in different LS channels, with strength determined by the parameter β_{LS} . In the NV2+3-Ia* model, β_{00} and β_{11} , parametrizing the 1S_0 and 3P_0 components of the ${}^6\text{He}$ wave function, are given by $\beta_{00} = 0.931$ and $\beta_{11} = -0.364$. In the case of ${}^6\text{Li}$, the 3S_1 , 1P_1 and 3D_2 components are given by $\beta_{01} = 0.967$, $\beta_{10} = 0.182$, and $\beta_{21} = 0.176$, respectively. These agree fairly well with Ref. [89], which found a smaller D wave component, $\beta_{21} = -0.03$. In this model, we obtain

$$C_{\text{recoil}}^{(1)}|_{\sigma L} = 0.013, \quad (30)$$

with statistical uncertainties < 0.001 . We observe a small, approximately linear dependence on β_{21} , which would shift the value to 0.012 for $\beta_{21} = -0.03$. To better mimic the shell model calculation of Ref. [89], we turned off the ‘‘one-pion-exchange-like’’ correlation operators in the VMC wave functions in a fashion similar to what was done in Ref. [61], observing a 10% increase of the matrix element, from 0.013 to 0.014. While the magnitude of the matrix element agrees very well with Ref. [89], we were not able to resolve the disagreement on the sign. Note that the correlation and D wave analyses above have been done only at the VMC level.

In GFMC the value of $\mathcal{M}_{\sigma L}$ is further decreased to

$$C_{\text{recoil}}^{(1)}|_{\sigma L} = 0.009. \quad (31)$$

The above values have small statistical uncertainties but are not accounting for any possible model dependencies. We conclude that (a) the magnitude of $\mathcal{M}_{\sigma L}$ agrees well with shell model calculations, but is reduced by nuclear correlations and by the GFMC evolution, and (b) the contribution of $\mathcal{M}_{\sigma rp}$ is non-negligible in the case of ${}^6\text{He}$, varying from

$$C_{\text{recoil}}^{(1)}|_{\sigma rp} = 0.007 \quad (32)$$

in VMC to

$$C_{\text{recoil}}^{(1)}|_{\sigma rp} = -0.012. \quad (33)$$

Finally, as anticipated in Refs. [89,92–94], we find substantial contributions from two-body currents for all three models.

The induced pseudoscalar contribution, captured by the $h(q^2)$ function in Holstein's formalism, was not considered in Ref. [89]. In agreement with Ref. [46], we find this contribution to dominate $C_1^{(1)}$. However, we stress that the induced pseudoscalar contribution to $C_1^{(1)}$ is partially canceled by the one to $L_1^{(2)}$, so that the single-nucleon induced pseudoscalar form factor gives corrections to the spectrum that are proportional to m_e^2 , as expected.

V. CHARGED CURRENTS IN THE SM EFFECTIVE FIELD THEORY

With the theoretical accuracy of the SM spectrum well below 0.1%, we can then study the sensitivity to physics beyond the SM. If BSM physics arises at a scale $\Lambda \gg v$, its correction to β decays can be described in the framework of the standard model effective field theory [97,98], an effective field theory that complements the SM with the most general set of gauge-invariant effective operators, expressed in terms of SM fields and organized according to their canonical dimension. SMEFT contains a single dimension-5 operator [99], which, when the Higgs gets its vacuum expectation value, gives rise to a Majorana mass term for the three left-handed neutrinos. At dimension 6, the SMEFT contains several classes of operators, which, at low energy, induce new axial, vector, scalar, pseudoscalar, and tensor semileptonic interactions between quarks, charged leptons and left-handed neutrino fields [97,98]. Since the mechanism behind the origin of neutrino masses is unknown, for generality we extend the SMEFT with a multiplet of n sterile neutrino fields ν_R (ν SMEFT)

[5,100,101]. The sterile neutrino is a singlet under the SM group. At dimension 3, ν_R has a Majorana mass term, while at dimension 4 it interacts with active neutrinos via Yukawa interactions. If one considers only these renormalizable interactions, after diagonalizing the neutrino mass matrix, the neutrino sector is characterized by $3+n$ mass eigenstates with masses m_1, \dots, m_{3+n} , and a $(3+n) \times (3+n)$ unitary mixing matrix U , which generalizes the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) mixing matrix. For simplicity, we will consider $n=1$, so that one has to consider the parameters m_4 and $U_{\alpha 4}$, with $\alpha \in \{e, \mu, \tau\}$, in addition to the SM. The inclusion of multiple light neutrino states is straightforward. From oscillation experiments and from the upper limit from the KATRIN experiment [102,103], $m_{1,2,3} \lesssim 0.8$ eV, so that they can be neglected in our analysis. We will consider m_4 as a free parameter, and mostly focus on the region in which m_4 is smaller than the Q value, so that ν_4 can be produced in the decay. In addition to renormalizable interactions, ν_R can have lepton-number-conserving nonstandard interactions with SM fields at dimension 6 [5,100,101], which, as we will see, induce new axial, vector, scalar, pseudoscalar, and tensor charged-current interactions involving sterile neutrinos.

After integrating out heavy gauge and quark fields and rotating to the neutrino mass basis, the most general low-energy Lagrangian for β decays is given in Refs. [5,104,105]. If the masses of all active and sterile neutrino states are much smaller than the nuclear scale or the electron mass, we can neglect sterile neutrino operators, whose interference with the SM is suppressed by powers of the neutrino masses. In this case, only scalar, pseudoscalar, and tensor interactions can give rise to a Fierz interference term, and, making connections with the notation of Ref. [5], we can write the relevant interactions as

$$\mathcal{L}^{(6)} = -\frac{4G_F}{\sqrt{2}} V_{ud} \left\{ \frac{1}{2} \bar{e}_R \nu_L (\epsilon_S \bar{u}d + \epsilon_P \bar{u} \gamma_5 d) + \epsilon_T \bar{e}_R \sigma^{\mu\nu} \nu_L \bar{u}_R \sigma_{\mu\nu} d_L \right\} + \text{H.c.} \quad (34)$$

If the masses of sterile neutrinos are non-negligible compared to the electron mass, there are additional interference terms. Using again the conventions of Ref. [5], we write the Lagrangian for ν_4 as

$$\begin{aligned} \mathcal{L}^{(6)} = & -\frac{4G_F}{\sqrt{2}} V_{ud} U_{e4} \left\{ \bar{e}_L \gamma_\mu \nu_4 ((1 + \epsilon_L) \bar{u}_L \gamma^\mu d_L + \epsilon_R \bar{u}_R \gamma^\mu d_R) + \frac{1}{2} \bar{e}_R \nu_4 (\epsilon_S \bar{u}d + \epsilon_P \bar{u} \gamma_5 d) \right. \\ & \left. + \epsilon_T \bar{u}_R \sigma^{\mu\nu} d_L \bar{e}_R \sigma_{\mu\nu} \nu_4 \right\} - \frac{4G_F}{\sqrt{2}} V_{ud} \left\{ \bar{e}_R \gamma_\mu \nu_4 (\tilde{\epsilon}_L \bar{u}_L \gamma^\mu d_L + \tilde{\epsilon}_R \bar{u}_R \gamma^\mu d_R) \right. \\ & \left. + \frac{1}{2} \bar{e}_L \nu_4 (\tilde{\epsilon}_S \bar{u}d + \tilde{\epsilon}_P \bar{u} \gamma_5 d) + \tilde{\epsilon}_T \bar{u}_L \sigma^{\mu\nu} d_R \bar{e}_L \sigma_{\mu\nu} \nu_4 \right\} + \text{H.c.} \end{aligned} \quad (35)$$

The conversion between the ϵ , the low-energy EFT (LEFT) couplings defined in Ref. [105] and ν SMEFT is discussed in Appendix A. The terms in the first bracket of Eq. (35) are induced by SMEFT operators involving active neutrinos and are proportional to the mixing U_{e4} . Since U_{e4} is small, we can usually neglect terms proportional to $U_{e4} \times \epsilon$. The terms in the second bracket, on the other hand, are induced by ν SMEFT operators with sterile neutrinos.

A. Beyond the SM corrections to the β spectrum

The multipole expansion can be generalized to nonstandard currents induced by SMEFT operators (see also Ref. [106]). At dimension 6 in the ν SMEFT, for both scalar/tensor and vector/axial operators, the leptonic and hadronic currents have at most spin 1. The derivation of the multipole expansion therefore proceeds essentially as in the SM. The additional vector and axial operators in Eq. (35) generate exactly the

same multipoles as in Eqs. (B15)–(B18). For scalar and pseudoscalar currents, only the C_{J_0} multipole is present, while for tensor currents only the electric, magnetic, and longitudinal multipoles are present.

Restricting again to the case of $J_i = 0$, $J_f = 1$, we can easily adapt the formulas in Eqs. (4)–(7) to the case of non-standard currents and define

$$C_1(q, P) = \frac{i}{\sqrt{4\pi}} \langle {}^6\text{Li}, 10 | \rho_{P+}^\dagger(q\hat{z}) | {}^6\text{He}, 00 \rangle, \quad (36)$$

$$L_1(q, T) = \frac{i}{\sqrt{4\pi}} \langle {}^6\text{Li}, 10 | \hat{z} \cdot \mathbf{j}_{T+}^\dagger(q\hat{z}) | {}^6\text{He}, 00 \rangle, \quad (37)$$

$$E_1(q, T) = \frac{i}{\sqrt{2\pi}} \langle {}^6\text{Li}, 10 | \hat{z} \cdot \mathbf{j}_{T+}^\dagger(q\hat{x}) | {}^6\text{He}, 00 \rangle, \quad (38)$$

$$M_1(q, T') = -\frac{1}{\sqrt{2\pi}} \langle {}^6\text{Li}, 10 | \hat{y} \cdot \mathbf{j}_{T+}^\dagger(q\hat{x}) | {}^6\text{He}, 00 \rangle. \quad (39)$$

We used

$$\rho_P(\mathbf{q}) = \int d^3\mathbf{x} e^{i\mathbf{q}\cdot\mathbf{x}} \mathcal{J}_P(\mathbf{x}), \quad \mathbf{j}_T^{(\prime)}(\mathbf{q}) = \int d^3\mathbf{x} e^{i\mathbf{q}\cdot\mathbf{x}} \mathcal{J}_T^{(\prime)}(\mathbf{x}). \quad (40)$$

The pseudoscalar density \mathcal{J}_P and tensor currents \mathcal{J}_T and $\mathcal{J}_T^{(\prime)}$ are defined in Eq. (B34), and the subscript + again refers to the isospin components. The multipoles $C_1(q, P)$, $L_1(q, T)$, $E_1(q, T)$, and $M_1(q, T')$ have a q expansion completely analogous to Eqs. (10)–(13). All corrections to the ${}^6\text{He}$ decay spectrum arising at dimension 6 in the νSMEFT can be expressed in terms of Eqs. (4)–(7) and (36)–(39).

The same power counting considerations in Sec. III B apply to multipoles induced by SMEFT charged-current operators. For purely GT transitions, tensor interactions are the most important, as they induce both $E_1^{(m,0)}(T)$ and $L_1^{(m,0)}(T)$ of order 1. From Eq. (C16), $M_1^{(1,0)}(T') = 0$, and thus we will neglect this contribution. Pseudoscalar interactions induce the multipole $C_1(q, P)$, which starts at $\mathcal{O}(qr_\pi)$. This suppression

is partially overcome by pion pole dominance of the pseudoscalar form factor, which implies

$$C_1^{(1)}(P) = \mathcal{O}(\epsilon_\chi^{-1}). \quad (41)$$

We thus include this contribution in the analysis. In the case of multipoles induced by non-SM currents, we only consider contributions at LO in both the multipole and chiral expansions, implying, in particular, that we only consider one-body currents, as given in Appendix C. To this order, from Eqs. (4)–(7), (36)–(39), and (C13)–(C16), we can see that

$$L_1^{(0)}(T) = \frac{1}{\sqrt{2}} E_1^{(0)}(T) = -\frac{2g_T}{g_A} L_1^{(0)}(A) + \mathcal{O}(\epsilon_\chi), \quad (42)$$

$$C_1^{(1)}(P) = -\frac{B}{m_\pi} L_1^{(0)}(A) + \mathcal{O}(\epsilon_\chi), \quad (43)$$

where g_T is the isovector tensor charge, $g_T = 0.989 \pm 0.033$ [107], and $B = m_\pi^2/(m_u + m_d) \approx 2.8$ GeV. B and g_T are scale dependent and given at the $\overline{\text{MS}}$ scale $\mu = 2$ GeV. Equation (43) confirms the relative enhancement of the pseudoscalar contribution.

We first consider the case in which active and sterile neutrinos have masses much smaller than the electron mass. In this particular case, nonstandard axial and vector interactions simply shift the overall normalization of Eq. (14). If BSM interactions induce not only $V - A$ but also $V + A$ right-handed currents, the only effect on the spectrum will be a shift in the relative coefficient between the $(L_1^{(0)})^2$ and $E_1^{(0)} M_1^{(1)}$ terms in Eq. (14), as the second originates from the interference of the axial and vector currents. Since this effect arises at recoil order, it will not provide strong constraints on new physics.

Tensor and scalar currents interfere with the standard model via terms proportional to the electron mass. Tensor interactions give rise to a term at $\mathcal{O}(q^0)$. The pseudoscalar contribution is formally $\mathcal{O}(q)$, but it is enhanced because of the pion-pole dominance of the pseudoscalar form factor. The differential cross section with respect to the electron energy is given by

$$\frac{d\Gamma_T}{d\varepsilon} = \frac{d\Gamma_0}{d\varepsilon} \frac{4m_e}{3E_e} \frac{1}{|L_1^{(0)}(A)|^2} \left\{ \varepsilon_T \text{Re}(E_1^{(0)}(T) E_1^{(0)*}(A) + L_1^{(0)}(T) L_1^{(0)*}(A)) - \frac{\varepsilon_P}{2} (1 - \varepsilon) W_0 r_\pi \text{Re}(C_1^{(1)}(P) L_1^{(0)*}(A)) \right\}, \quad (44)$$

where $d\Gamma_0/d\varepsilon$ is the SM decay rate at LO in the multipole expansion, defined in Eq. (16) [108]. Equation (44) shows the characteristic m_e/E_e behavior. Using Eq. (43), Eq. (44) becomes

$$\frac{d\Gamma_T}{d\varepsilon} = \frac{d\Gamma_0}{d\varepsilon} \frac{m_e}{E_e} \left\{ -8 \frac{g_T \varepsilon_T}{g_A} + \frac{2}{3} \varepsilon_P (1 - \varepsilon) \frac{W_0}{m_u + m_d} \right\}. \quad (45)$$

The tensor contributions agrees with the result of Ref. [34], while the enhancement of the pseudoscalar contribution was noted, for example, in Ref. [109]. Equation (45) is only valid at LO in chiral EFT.

Next, we consider the case of sterile neutrinos with non-negligible mass. Here we work at LO in the multipole expansion and consider $m_{\nu_4} < Q$, so that ν_4 can be produced. Considering vector and axial currents, we obtain

$$\frac{d\Gamma_{\nu_4}}{d\varepsilon} = \frac{d\Gamma_0}{d\varepsilon} \left\{ \left(\sqrt{1 - \frac{m_{\nu_4}^2}{E_v^2}} - 1 \right) (1 + \varepsilon_L - \varepsilon_R) |U_{e4}|^2 + \sqrt{1 - \frac{m_{\nu_4}^2}{E_v^2}} \frac{m_e m_{\nu_4}}{E_e E_v} U_{e4}^* (\tilde{\varepsilon}_L - \tilde{\varepsilon}_R) \right\}, \quad (46)$$

where $E_v = W_0(1 - \varepsilon)$. If we turn off all dimension-6 operators, $\varepsilon_{L,R} = 0$ and $\tilde{\varepsilon}_{L,R} = 0$, Eq. (46) is only proportional to the mixing U_{e4} . In this case the presence of sterile neutrinos has two effects, one on the normalization and one energy dependent. In the presence of nonstandard interactions of a sterile neutrino with a right-handed electron, $\tilde{\varepsilon}_L - \tilde{\varepsilon}_R$, we also get new Fierz-like

terms. For tensor interactions

$$\frac{d\Gamma_{\nu_4,T}}{d\varepsilon} = \frac{d\Gamma_0}{d\varepsilon} \frac{4}{3|L_1^{(0)}(A)|^2} \text{Re}[E_1^{(0)}(T)E_1^{(0)*}(A) + L_1^{(0)}(T)L_1^{(0)*}(A)] \left\{ \frac{m_e}{E_e} \left(\sqrt{1 - \frac{m_{\nu_4}^2}{E_v^2}} - 1 \right) \varepsilon_T |U_{e4}|^2 + \frac{m_{\nu_4}}{E_v} \sqrt{1 - \frac{m_{\nu_4}^2}{E_v^2}} \tilde{\varepsilon}_T U_{e4}^* \right\}, \quad (47)$$

so that a fourth massive neutrino would affect the standard Fierz interference term, and, more importantly, generate a new interference term, proportional to the neutrino mass. Similarly, the pseudoscalar interactions of sterile and active neutrinos give

$$\frac{d\Gamma_{\nu_4,P}}{d\varepsilon} = -\frac{d\Gamma_0}{d\varepsilon} \frac{2}{3|L_1^{(0)}(A)|^2} [(1 - \varepsilon)W_0 r_\pi \text{Re}(C_1^{(1)}(P)L_1^{(0)*}(A))] \left\{ \frac{m_e}{E_e} \left(\sqrt{1 - \frac{m_{\nu_4}^2}{E_v^2}} - 1 \right) \varepsilon_P |U_{e4}|^2 + \frac{m_{\nu_4}}{E_v} \sqrt{1 - \frac{m_{\nu_4}^2}{E_v^2}} \tilde{\varepsilon}_P U_{e4}^* \right\}. \quad (48)$$

VI. SENSITIVITY TO BSM SIGNATURES

We now consider the impact of non-standard charged-current interactions on the energy spectrum. In Fig. 6, we assume all neutrinos to be massless and we set the tensor and pseudoscalar interaction to $\varepsilon_T = 10^{-3}$ and $\varepsilon_P = 5 \times 10^{-3}$, corresponding to new physics scales $\Lambda \approx 8$ and 4 TeV, respectively. Interactions of these size lead to $\approx 10^{-3}$ corrections, which should be resolved in the next generation of experiments. For both pseudoscalar and tensor interactions, the uncertainty band includes uncertainties on the one-body parameters, g_T and B , and the nuclear uncertainties on the multipoles, but does not include the truncation to the one-body level, and it is thus slightly underestimated. High-invariant mass Drell-Yan production at the LHC currently probes ε_T at a very similar level [11,14], while a global analysis of β decays found $\varepsilon_T \in [-0.8, 1.2] \times 10^{-3}$, at the 1σ level [12]. Pseudoscalar interactions are very well constrained by the ratio $\text{BR}(\pi \rightarrow e\nu)/\text{BR}(\pi \rightarrow \mu\nu)$, which yields $-1.4 \times$

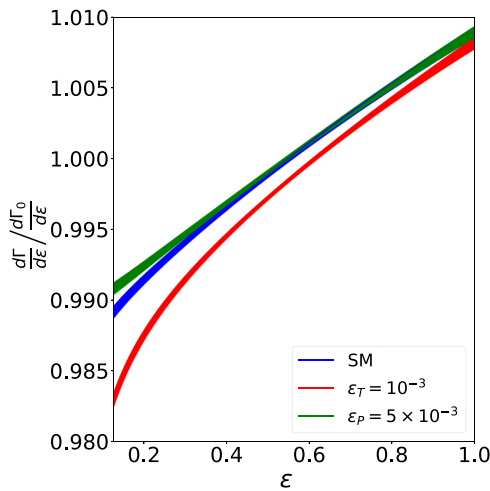


FIG. 6. Deviation of the ${}^6\text{He}$ β spectrum from the expression truncated at leading order in the multipole expansion, given in Eq. (16). The blue curve denotes the SM results, while the red and green lines include the contributions of a tensor and pseudoscalar current, respectively. The widths of the bands denote the theoretical error.

$10^{-7} < \varepsilon_P < 5.5 \times 10^{-4}$ [6]. Such values are not in reach of upcoming ${}^6\text{He}$ experiments.

We next consider the case of a massive sterile neutrino, which mixes with the electron neutrino with strength U_{e4} , and has nonstandard axial, vector, scalar, and tensor interactions. Since the corrections scale in general as m_{ν_4}/W_0 , ${}^6\text{He}$ decays can probe m_{ν_4} in the MeV range. Currently the best limit on the a sterile neutrino with mass $m_{\nu_4} = 1$ MeV come from the β spectra of ${}^{20}\text{F}$ and ${}^{144}\text{Pr}$, and, under the assumption that the neutrino interacts with the SM only via mixing, these constrain a mixing angle $U_{e4} \approx 2 \times 10^{-3}$ [40–43,110,111]. From this we see that we can always neglect terms in Eqs. (46), (47), and (48) that are proportional to nonstandard interactions of active neutrinos, $\varepsilon_{L,R,P,T}$, since they are doubly suppressed by $|U_{e4}|^2$ and v^2/Λ^2 .

In the left panel of Fig. 7 we show the corrections to the spectrum in the case sterile neutrinos interact with SM particles only via Yukawa interactions. In this case, the spectrum would show a characteristic “kink” at $\varepsilon = 1 - m_{\nu_4}/W_0$, due to the emission of a massive neutrino. With mixing $|U_{e4}|^2 = 10^{-3}$, the spectrum receives permille level corrections.

Nonstandard interactions of sterile neutrinos cause corrections to the spectrum of order $U_{e4} \tilde{\varepsilon}_J$, which could thus be relevant for $\tilde{\varepsilon}_J \approx U_{e4} \approx 3 \times 10^{-2}$, corresponding to new physics scales of 1 TeV. Sterile neutrinos with an axial coupling to quarks induce corrections proportional to $m_e m_{\nu_4}/(E_e E_v)$, which are, however, fairly small. More promising is the case of sterile neutrinos with nonstandard tensor interactions, $\tilde{\varepsilon}_T$, shown in the right panel of Fig. 7. These interactions arise, for example, in leptoquark models [105]. In this case, an interference term of the form m_{ν_4}/E_v is induced, which has a very different shape compared to tensor interactions of active neutrinos. This is particularly interesting, since the analysis of Ref. [12] found some preference for a tensor interaction involving sterile neutrinos in β decay data.

VII. CONCLUSIONS

We performed an *ab initio* calculation of the electron energy spectrum in the β decay of ${}^6\text{He}$. We used potentials derived from chiral EFT, with consistent weak vector and axial currents, and adopted quantum Monte Carlo methods to solve the many-body nuclear problem. We included terms up

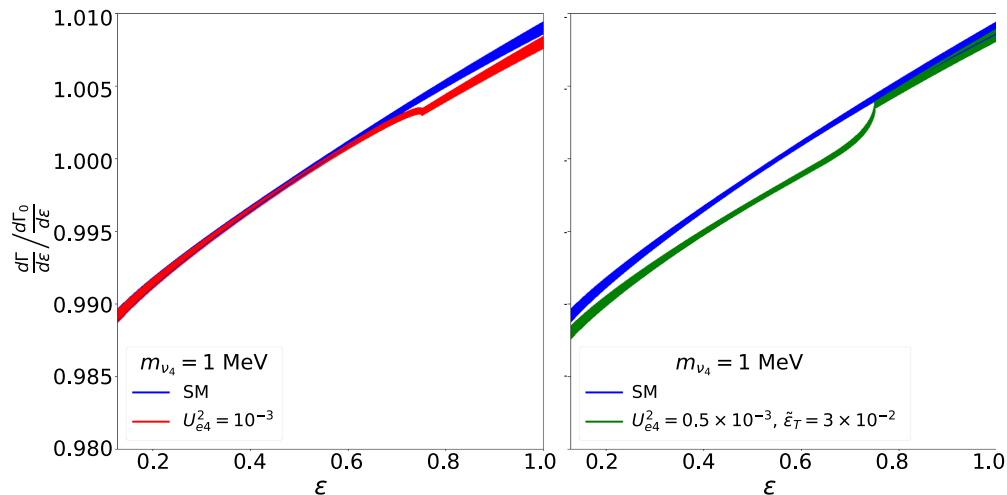


FIG. 7. Corrections to the β spectrum from sterile neutrinos with minimal (*left*) and tensor (*right*) interactions, ξ_T .

to second order in the multipole expansion [71], and state-of-the-art electromagnetic corrections, following the treatment of Ref. [76]. In particular, we included two-body currents, for the first time in an *ab initio* calculation of the spectrum. To estimate the theoretical error on the spectrum, we evaluated the matrix elements L_1 , E_1 , M_1 , and C_1 with four potential models in the Norfolk family of local two- and three-nucleon interactions, derived in chiral EFT with explicit Δ 's and including terms up to $N^3\text{LO}$ in the chiral expansion. The four interactions have different cut-off, they fit nucleon-nucleon scattering data up to different energies, and use different observables to determine the low-energy constants in the three-body force, and thus they provide a good estimate of the systematic errors in the calculation. We find the ${}^6\text{He}$ half-life to be in good agreement with experiment. The theoretical uncertainty of about 3% is dominated by the determination of the three-body force. We find the error on the spectral shape to be well below the permille level, and to receive contributions of approximately the same size from $M_1^{(1)}$, $C_1^{(1)}$, $L_1^{(2)}$, and $E_1^{(2)}$. In the case of $M_1^{(1)}$, which encodes the contribution of weak magnetism, our results agree within theoretical error with the extraction from the electromagnetic transition ${}^6\text{Li}(0^+, 1) \rightarrow {}^6\text{Li}(1^+, 0)\gamma$, which is exact in the isospin limit. We checked that isospin-breaking terms in the nuclear potential induce a 1% difference between M_1 and its electromagnetic analog, of the same size as the experimental error.

C_1 is determined by the matrix element of the axial charge density. We find this matrix element to be dominated by the induced pseudoscalar form factor, in agreement with Ref. [46]. Finally, we find that $E_1^{(2)}$ and $L_1^{(2)}$, which contribute at $N^2\text{LO}$ in the multipole expansion, give permille level corrections to the spectrum, and thus need to be included for an accuracy goal of few parts in 10^{-4} . In GFMC, $E_1^{(2)}$ and $L_1^{(2)}$ have relatively large uncertainties, 26% and 12%, respectively. Also in this case, the dominant systematic uncertainty arises from the determination of the three-body force and the linear extrapolation of model IIa*.

Two-body currents play a particularly important role for M_1 and C_1 , which receive a 8% and ≈ 20 –30% correction respectively. The effects on E_1 and L_1 are smaller.

Combining the uncertainties on different matrix elements, we reach a total error on the differential decay rate, normalized by the rate at leading order in the multipole expansion, of at most 4×10^{-4} . We discussed the consequences of such accuracy on nonstandard charged-current interactions involving active and sterile neutrinos, showing that the next generation of experiments will be sensitive to tensor interactions and, to a lesser extent, to pseudoscalar interactions of active neutrinos. Future experiments will also constrain sterile neutrinos with mass in the ≈ 1 MeV region, with both minimal and non-minimal interactions.

ACKNOWLEDGMENTS

We acknowledge stimulating conversations with A. Garcia, D. Gazit, A. Glick-Magid, M. Hoferichter, J. Menéndez, and R. Schiavilla. A.B., S.G., and E.M. are supported by the US Department of Energy through the Office of Nuclear Physics under Contract No. DE-AC52-06NA25396, and the LDRD program at Los Alamos National Laboratory. Los Alamos National Laboratory is operated by Triad National Security, LLC, for the National Nuclear Security Administration of U.S. Department of Energy (Contract No. 89233218CNA000001). The work of S.G. was also supported by the DOE Early Career research Program. L.H. is supported through the U.S. Department of Energy, Low Energy Physics Grant No. DE-FG02-ER41042 and NSF Grant No. PHY-1914133. This work is also supported by the U.S. Department of Energy under Contract No. DE-SC0021027 (G.K. and S.P.), a 2021 Early Career Award No. DE-SC0022002 (M.P.), the FRIB Theory Alliance Award No. DE-SC0013617 (S.P. and M.P.), and the U.S. DOE NNSA Stewardship Science Graduate Fellowship under Cooperative Agreement No. DE-NA0003960 (G. K.). V.C. is supported by the U.S. Department of Energy under Contract No. DE-FG02-00ER4113. The many-body calculations were performed on the parallel computers of the Laboratory Computing Resource Center, Argonne National Laboratory, and the computers of the Argonne Leadership Computing Facility via the INCITE grant “Ab-initio nuclear structure and nuclear reactions,” the 2019/2020 ALCC grant

“Low Energy Neutrino-Nucleus interactions” for the project NNInteractions, the 2020/2021 ALCC grant “Chiral Nuclear Interactions from Nuclei to Nucleonic Matter” for the project ChiralNuc, and by the 2021/2022 ALCC grant “Quantum Monte Carlo Calculations of Nuclei up to ^{16}O and Neutron Matter” for the project QMCNuc. This research also used resources provided by the Los Alamos National Laboratory Institutional Computing Program, which is supported by the U.S. Department of Energy National Nuclear Security Administration under Contract No. 89233218CNA000001.

APPENDIX A: EFFECTIVE LAGRANGIANS FOR CHARGED-CURRENT PROCESSES

1. Charged currents in SMEFT and ν SMEFT

The SMEFT Lagrangian includes all operators that are invariant under the SM $SU(3)_c \times SU(2)_L \times U(1)_Y$ gauge group and are built out of SM fields, the left-handed quark and lepton doublets, $Q = (u_L, d_L)^T$ and $L = (\nu_L, e_L)^T$, the right-handed

$SU(2)$ singlets u , d , and e , and the scalar doublet

$$H = \frac{v}{\sqrt{2}} U(x) \begin{pmatrix} 0 \\ 1 + \frac{h(x)}{v} \end{pmatrix}, \quad (\text{A1})$$

where $h(x)$ is the Higgs field, and $U(x)$ is a $SU(2)$ matrix encoding the Goldstone modes. We will also use $\tilde{H} = i\tau_2 H^*$.

At dimension 5, the only operator that can be constructed is the LNV Weinberg operator [99]

$$\mathcal{L}_{\text{SMEFT}}^{(5)} = \epsilon_{kl} \epsilon_{mn} (L_k^T C^{(5)} C L_m) H_l H_n, \quad (\text{A2})$$

where C is the charge conjugation matrix. After electroweak symmetry breaking, Eq. (A2) induces a Majorana mass term for active neutrinos. The full, nonredundant, dimension-6 Lagrangian is given in Ref. [98]. For β decays, the most important terms are quark and lepton bilinears, which modify the couplings of the W boson to left-handed quarks and leptons and induce new right-handed couplings of the W to quarks, and semileptonic four-fermion operators:

$$\begin{aligned} \mathcal{L}_{\text{SMEFT}}^{(6)} = & C_{HL3}^{(6)} (H^\dagger i \overleftrightarrow{D}_\mu^I H) (\bar{L} \tau^I \gamma^\mu L) + C_{HQ3}^{(6)} (H^\dagger i \overleftrightarrow{D}_\mu^I H) (\bar{Q} \tau^I \gamma^\mu Q) + C_{Hud}^{(6)} i (\tilde{H}^\dagger D_\mu H) (\bar{u} \gamma^\mu d) + C_{LQ3}^{(6)} (\bar{L} \gamma^\mu \tau^I L) (\bar{Q} \gamma_\mu \tau^I Q) \\ & + [C_{LedQ}^{(6)} (\bar{L}^j e) (\bar{d} Q^j) + C_{LeQu1}^{(6)} (\bar{L}^j e) \epsilon_{jk} (\bar{Q}^k u) + C_{LeQu3}^{(6)} (\bar{L}^j \sigma_{\mu\nu} e) \epsilon_{jk} (\bar{Q}^k \sigma^{\mu\nu} u) + \text{H.c.}]. \end{aligned} \quad (\text{A3})$$

All dimension-6 operators are lepton-number-conserving (LNC). LNV operators arise at dimension-seven and were constructed in Ref. [112]. Their contribution to low-energy charged-current operators were considered in Ref. [105].

In addition to the SM fields, we introduce a multiplet of sterile neutrinos ν_R , which is a singlet under the SM group. At dimension 3, this allows to write down a Majorana mass term and a Yukawa interaction, so that the renormalizable ν SMEFT Lagrangian is

$$\mathcal{L}_{\nu\text{SMEFT}} = \mathcal{L}_{\text{SM}} + \bar{\nu}_R i \gamma^\mu \partial_\mu \nu_R - \left[\frac{1}{2} \bar{\nu}_R^c \bar{M}_R \nu_R + \bar{L} \tilde{H} Y_\nu \nu_R + \text{H.c.} \right], \quad (\text{A4})$$

where \bar{M}_R is a symmetric $n \times n$ complex matrix, and Y_ν is a $3 \times n$ matrix of Yukawa couplings. The next interactions relevant to β decay appear at dimension 6,

$$\begin{aligned} \mathcal{L}_{\nu\text{SMEFT}}^{(6)} = & \mathcal{L}_{\text{SMEFT}}^{(6)} + C_{H\nu e}^{(6)} (\bar{\nu}_R \gamma^\mu e) (\tilde{H}^\dagger i D_\mu H) + C_{d\nu e}^{(6)} (\bar{d} \gamma^\mu u) (\bar{\nu}_R \gamma_\mu e) + C_{Qu\nu L}^{(6)} (\bar{Q} u) (\bar{\nu}_R L) \\ & + C_{L\nu Qd}^{(6)} (\bar{L}_i \nu_R) \epsilon_{ij} (\bar{Q}_j d) + C_{LdQ\nu}^{(6)} (\bar{L}_i d) \epsilon_{ij} (\bar{Q}_j \nu_R) + \text{H.c.} \end{aligned} \quad (\text{A5})$$

The first operator induces a coupling of the W boson to ν_R and a right-handed electron. $C_{d\nu e}^{(6)}$ is a purely right-handed semileptonic charged-current interaction, while the operators on the second line are scalar and tensor interactions of a right-handed neutrino with quarks and left-handed electrons. Finally, there are LNV operators at dimension 7 [101], which we do not consider here.

2. Charged currents in LEFT

After electroweak symmetry breaking and integrating out the W boson, the operators in Eq. (A3) and (A5) match onto a LNC β decay Lagrangian in a low-energy, $SU(3) \times U_{\text{em}}(1)$ invariant theory (LEFT). In the flavor basis, this is given by

$$\begin{aligned} \mathcal{L}^{(6)} = & -\frac{4G_F}{\sqrt{2}} \left\{ \bar{u}_L \gamma^\mu d_L [\bar{e}_L \gamma_\mu c_{\text{VL}}^{(6)} \nu_L + \bar{e}_R \gamma_\mu \bar{c}_{\text{VL}}^{(6)} \nu_R] + \bar{u}_R \gamma^\mu d_R [\bar{e}_L \gamma_\mu c_{\text{VR}}^{(6)} \nu_L + \bar{e}_R \gamma_\mu \bar{c}_{\text{VR}}^{(6)} \nu_R] + \bar{u}_L d_R [\bar{e}_R c_{\text{SR}}^{(6)} \nu_L + \bar{e}_L \bar{c}_{\text{SR}}^{(6)} \nu_R] \right. \\ & \left. + \bar{u}_R d_L [\bar{e}_R c_{\text{SL}}^{(6)} \nu_L + \bar{e}_L \bar{c}_{\text{SL}}^{(6)} \nu_R] + \bar{u}_R \sigma^{\mu\nu} d_L \bar{e}_R \sigma_{\mu\nu} c_{\text{T}}^{(6)} \nu_L + \bar{u}_L \sigma^{\mu\nu} d_R \bar{e}_L \sigma_{\mu\nu} \bar{c}_{\text{T}}^{(6)} \nu_R \right\} + \text{H.c.} \end{aligned} \quad (\text{A6})$$

Here we follow the conventions of Ref. [105] and denote with unbarred and barred lower case coefficients, such as $c_{\text{VL}}^{(6)}$ and $\bar{c}_{\text{VL}}^{(6)}$, lepton-number-conserving operators that, in the flavor basis, involve active and sterile neutrinos, and thus receive matching contributions from dimension-6 SMEFT and ν SMEFT operators, respectively. In the neutrino mass basis, and assuming one additional light state, ν_4 , the Lagrangian becomes

$$\begin{aligned} \mathcal{L}^{(6)} = & -\frac{4G_F}{\sqrt{2}} \sum_{i=1}^4 \left\{ \bar{u}_L \gamma^\mu d_L (\bar{e}_R \gamma_\mu [C_{\text{VLR}}^{(6)}]_{ei} \nu_i + \bar{e}_L \gamma_\mu [C_{\text{VLL}}^{(6)}]_{ei} \nu_i) + \bar{u}_R \gamma^\mu d_R (\bar{e}_R \gamma_\mu [C_{\text{VRR}}^{(6)}]_{ei} \nu_i + \bar{e}_L \gamma_\mu [C_{\text{VRL}}^{(6)}]_{ei} \nu_i) \right. \\ & + \bar{u}_L d_R (\bar{e}_L [C_{\text{SRR}}^{(6)}]_{ei} \nu_i + \bar{e}_R [C_{\text{SRL}}^{(6)}]_{ei} \nu_i) + \bar{u}_R d_L (\bar{e}_L [C_{\text{SLR}}^{(6)}]_{ei} \nu_i + \bar{e}_R [C_{\text{SLL}}^{(6)}]_{ei} \nu_i) \\ & \left. + \bar{u}_L \sigma^{\mu\nu} d_R \bar{e}_L \sigma_{\mu\nu} [C_{\text{TRR}}^{(6)}]_{ei} \nu_i + \bar{u}_R \sigma^{\mu\nu} d_L \bar{e}_R \sigma_{\mu\nu} [C_{\text{TLL}}^{(6)}]_{ei} \nu_i \right\} + \text{H.c.} \end{aligned} \quad (\text{A7})$$

Neglecting lepton-number-violating (LNV) operators, which only arise at dimension 7 in the ν SMEFT, $C_{\text{VLL}}^{(6)}$, $C_{\text{VRL}}^{(6)}$, $C_{\text{SLL}}^{(6)}$, $C_{\text{SRL}}^{(6)}$, and $C_{\text{TLL}}^{(6)}$ receive contributions from vector axial, scalar, pseudoscalar, and tensor interactions with left-handed neutrinos,

$$\begin{aligned} [C_{\text{VLL}}^{(6)}]_{ei} &= [c_{\text{VL}}^{(6)}]_{e\alpha} U_{\alpha i}, & [C_{\text{VRL}}^{(6)}]_{ei} &= [c_{\text{VR}}^{(6)}]_{e\alpha} U_{\alpha i}, \\ [C_{\text{SLL}}^{(6)}]_{ei} &= [c_{\text{SL}}^{(6)}]_{e\alpha} U_{\alpha i}, & [C_{\text{SRL}}^{(6)}]_{ei} &= [c_{\text{SR}}^{(6)}]_{e\alpha} U_{\alpha i}, \\ [C_{\text{TLL}}^{(6)}]_{ei} &= [c_{\text{T}}^{(6)}]_{e\alpha} U_{\alpha i}, \end{aligned} \quad (\text{A8})$$

with $i = 1, \dots, 4$, and α denotes a charged lepton flavor index, $\alpha \in \{e, \mu, \tau\}$. In practice, we will assume SMEFT operators to be diagonal in lepton flavor, and restrict our attention to the ee components. We will thus drop the flavor subscripts on the $c^{(6)}$ coefficients. The operators $C_{\text{VLR}}^{(6)}$, $C_{\text{VRR}}^{(6)}$, $C_{\text{SLR}}^{(6)}$, $C_{\text{SRR}}^{(6)}$, and $C_{\text{TRR}}^{(6)}$ involve sterile neutrinos

$$\begin{aligned} [C_{\text{VLR}}^{(6)}]_{ei} &= [\bar{c}_{\text{VL}}^{(6)}]_{eS} U_{Si}^*, & [C_{\text{VRR}}^{(6)}]_{ei} &= [\bar{c}_{\text{VR}}^{(6)}]_{eS} U_{Si}^*, \\ [C_{\text{SLR}}^{(6)}]_{ei} &= [\bar{c}_{\text{SL}}^{(6)}]_{eS} U_{Si}^*, & [C_{\text{SRR}}^{(6)}]_{ei} &= [\bar{c}_{\text{SR}}^{(6)}]_{eS} U_{Si}^*, \\ [C_{\text{TRR}}^{(6)}]_{ei} &= [\bar{c}_{\text{T}}^{(6)}]_{eS} U_{Si}^*, \end{aligned} \quad (\text{A9})$$

where S in the subscript of the PMNS matrix denote the sterile flavor state. In this case, we will absorb the factor of U_{Si} in the coefficient of the effective operators.

In the body of the paper, we adopt the ϵ notation defined in Ref. [5]. For interactions involving active neutrinos, the relation between the ϵ couplings and the couplings in Eqs. (A6) and (A7) is given by

$$\begin{aligned} V_{ud}(1 + \epsilon_L) &= c_{\text{VL}}^{(6)}, & V_{ud} \epsilon_R &= c_{\text{VR}}^{(6)}, \\ V_{ud} \epsilon_S &= c_{\text{SR}}^{(6)} + c_{\text{SL}}^{(6)}, & V_{ud} \epsilon_P &= c_{\text{SR}}^{(6)} - c_{\text{SL}}^{(6)}, \\ V_{ud} \epsilon_T &= c_{\text{T}}^{(6)}, \end{aligned} \quad (\text{A10})$$

while, in the case of sterile neutrinos, we have

$$\begin{aligned} V_{ud} \tilde{\epsilon}_L &= [\bar{c}_{\text{VL}}^{(6)}]_{eS} U_{S4}^*, & V_{ud} \epsilon_R &= [\bar{c}_{\text{VR}}^{(6)}]_{eS} U_{S4}^*, \\ V_{ud} \tilde{\epsilon}_S &= [\bar{c}_{\text{SR}}^{(6)} + \bar{c}_{\text{SL}}^{(6)}]_{eS} U_{S4}^*, & V_{ud} \tilde{\epsilon}_P &= [c_{\text{SR}}^{(6)} - c_{\text{SL}}^{(6)}]_{eS} U_{S4}^*, \\ V_{ud} \tilde{\epsilon}_T &= [\bar{c}_{\text{T}}^{(6)}]_{eS} U_{S4}^*. \end{aligned} \quad (\text{A11})$$

The matching between SMEFT, ν SMEFT, and Eq. (A6) was carried out in Ref. [105], and here we report the results. For nonstandard interactions involving active neutrinos, one finds

$$\begin{aligned} [c_{\text{VLL}}^{(6)}]_{\alpha\beta} &= V_{ud} \delta_{\alpha\beta} - v^2 [C_{\text{LQ}3}^{(6)} - C_{\text{HL}3}^{(6)}]_{\alpha\beta} + v^2 [C_{\text{HQ}3}^{(6)} \delta_{\alpha\beta}], \\ [c_{\text{VRL}}^{(6)}]_{\alpha\beta} &= \frac{v^2}{2} C_{\text{Hud}}^{(6)} \delta_{\alpha\beta}, \\ [c_{\text{SRL}}^{(6)}]_{\alpha\beta} &= -\frac{v^2}{2} [C_{\text{LedQ}}^{(6)}]_{\beta\alpha}^*, \end{aligned}$$

$$\begin{aligned} [c_{\text{SL}}^{(6)}]_{\alpha\beta} &= -\frac{v^2}{2} [C_{\text{LeQu}1}^{(6)}]_{\beta\alpha}^*, \\ [c_{\text{T}}^{(6)}]_{\alpha\beta} &= -\frac{v^2}{2} [C_{\text{LeQu}3}^{(6)}]_{\beta\alpha}^*. \end{aligned} \quad (\text{A12})$$

Here α and β denote charge lepton flavor indices, $\alpha, \beta \in \{e, \mu, \tau\}$, while we are always assuming the quark flavor indices to be u and d . The matching coefficients of LNC sterile neutrino operators are

$$\begin{aligned} [\bar{c}_{\text{VL}}^{(6)}]_{\alpha S} &= \frac{v^2}{2} [C_{\text{Hve}}^{(6)}]_{S\alpha}^*, \\ [\bar{c}_{\text{VR}}^{(6)}]_{\alpha S} &= -\frac{v^2}{2} [C_{\text{dve}}^{(6)}]_{S\alpha}^*, \\ [\bar{c}_{\text{SR}}^{(6)}]_{\alpha S} &= \frac{v^2}{2} [C_{\text{LvQd}}^{(6)}]_{\alpha S} - \frac{v^2}{4} [C_{\text{LdQv}}^{(6)}]_{\alpha S}, \\ [\bar{c}_{\text{SL}}^{(6)}]_{\alpha S} &= -\frac{v^2}{2} [C_{\text{QuvL}}^{(6)}]_{S\alpha}^*, \\ [\bar{c}_{\text{T}}^{(6)}]_{\alpha S} &= -\frac{v^2}{16} [C_{\text{LdQv}}^{(6)}]_{\alpha S}, \end{aligned} \quad (\text{A13})$$

where S denotes a sterile flavor index. Ref. [105] also reports the contribution of LNV SMEFT and ν SMEFT operators to Eq. (A7).

3. Charged currents in the chiral Lagrangian

The quark-level SM and SMEFT Lagrangians lead to interactions between pions and nucleons, which can be organized in chiral perturbation theory [113–115] and are the building blocks for the derivation of the one- and two-body nuclear currents used in this paper. Axial and pseudoscalar interactions induce couplings to a single pion,

$$\begin{aligned} \mathcal{L}_\pi &= 2G_F F_\pi \{ \partial^\mu \pi^- (\bar{e}_R \gamma_\mu [C_{\text{VLR}}^{(6)} - C_{\text{VRR}}^{(6)}]_{ei} \nu_i \\ &\quad + \bar{e}_L \gamma_\mu [C_{\text{VLL}}^{(6)} - C_{\text{VRL}}^{(6)}]_{ei} \nu_i) \\ &\quad + iB\pi^- (\bar{e}_L [C_{\text{SLR}}^{(6)} - C_{\text{SRR}}^{(6)}]_{ei} \nu_i \\ &\quad + \bar{e}_R [C_{\text{SLL}}^{(6)} - C_{\text{SRL}}^{(6)}]_{ei} \nu_i) \} + \dots, \end{aligned} \quad (\text{A14})$$

where \dots denotes terms with multiple pions and affect the nuclear currents beyond leading order. Axial and pseudoscalar interactions induce terms with an odd number of pions, while scalar and vector interactions induce terms with an even number of pions. $F_\pi = 92$ MeV is the pion decay constant, while $B = m_\pi^2 / (m_u + m_d) \approx 2.8$ GeV, at the renormalization scale $\mu = 2$ GeV. The strong coupling of the pion to pseudoscalar operators implies that this interaction dominates the nucleon pseudoscalar density. In the heavy-baryon formalism, the nucleon Lagrangian at leading order is

$$\begin{aligned} \mathcal{L}_N &= -\frac{2G_F}{\sqrt{2}} \bar{N} \tau^+ \{ v^\mu (\bar{e}_R \gamma_\mu [C_{\text{VLR}}^{(6)} + C_{\text{VRR}}^{(6)}]_{ei} \nu_i + \bar{e}_L \gamma_\mu [C_{\text{VLL}}^{(6)} + C_{\text{VRL}}^{(6)}]_{ei} \nu_i) - 2g_A S^\mu (\bar{e}_R \gamma_\mu [C_{\text{VLR}}^{(6)} - C_{\text{VRR}}^{(6)}]_{ei} \nu_i \\ &\quad + \bar{e}_L \gamma_\mu [C_{\text{VLL}}^{(6)} - C_{\text{VRL}}^{(6)}]_{ei} \nu_i) + g_S (\bar{e}_L [C_{\text{SLR}}^{(6)} + C_{\text{SRR}}^{(6)}]_{ei} \nu + \bar{e}_R [C_{\text{SLL}}^{(6)} + C_{\text{SRL}}^{(6)}]_{ei} \nu_i) \\ &\quad - 4g_T \epsilon^{\mu\nu\alpha\beta} v_\alpha S_\beta (\bar{e}_L \sigma_{\mu\nu} [C_{\text{TRR}}^{(6)}]_{ei} \nu^i + \bar{e}_R \sigma_{\mu\nu} [C_{\text{TLL}}^{(6)}]_{ei} \nu_i) \} N, \end{aligned} \quad (\text{A15})$$

where N denotes a nonrelativistic nucleon field, with velocity $v^\mu = (1, \mathbf{0})$ and spin $S^\mu = (1, \boldsymbol{\sigma}/2)$, in the nucleon rest frame. In the absence of nonstandard currents, the axial charge can be extracted from neutron decay. In this work, we adopt for the value of the axial charge [72]

$$\frac{g_A}{g_V} = 1.2754 \pm 0.0013. \quad (\text{A16})$$

This value is slightly larger and with roughly half the uncertainty of the one in the 2018 version of the PDG, $g_A|_{18} = 1.2723(23)$ [116], which is used in the code. In obtaining the half-life and β spectrum, we rescale the leading multipoles E_1 and L_1 , given in Tables III and IV, by $g_A/g_A|_{18}$. For the subleading multipoles, C_1 in particular, the difference is well within the theoretical error. In the future, one can envision using lattice QCD extraction of the axial charge [10,27,28,107,117,118], which is approaching percent level accuracy.

The scalar and tensor isovector charges have been computed in lattice QCD [10,28,107,117]. We will use for the scalar and tensor isovector charges the averages of the lattice results with $N_f = 2 + 1 + 1$ flavors of dynamical quarks, given in Ref. [28,107]

$$g_S = 1.02 \pm 0.10, \quad g_T = 0.989 \pm 0.033. \quad (\text{A17})$$

Equations (A14) and (A15) are sufficient for the construction of one-body currents at LO. The construction of vector and axial currents to subleading orders is reviewed in Refs. [51,59,119,120]. For the BSM scalar, pseudoscalar, and tensor currents, it is sufficient to work at LO.

APPENDIX B: MULTIPOLE EXPANSION FOR SM AND BSM CURRENTS

The derivation of the multipole expansion for SM currents is reviewed in Ref. [71]. The starting point is the weak Hamiltonian

$$\begin{aligned} H_w &= \frac{G_F}{\sqrt{2}} V_{ud} \int d^3\mathbf{x} j_\mu^{\text{lept}}(\mathbf{x}) \mathcal{J}_{V-A}^\mu(\mathbf{x}) \\ &= -\frac{G_F}{\sqrt{2}} V_{ud} \int d^3\mathbf{x} (\mathbf{j}^{\text{lept}}(\mathbf{x}) \cdot \mathcal{J}_{V-A}(\mathbf{x}) - j_0^{\text{lept}}(\mathbf{x}) \mathcal{J}_{V-A}^0(\mathbf{x})), \end{aligned} \quad (\text{B1})$$

where

$$j_\mu^{\text{lept}} = 2\bar{e}_L \gamma_\mu \nu_L. \quad (\text{B2})$$

where $\hat{\mathbf{e}}_{q3} = \hat{\mathbf{q}}$, $\hat{\mathbf{e}}_{q2} = \hat{\mathbf{z}} \times \mathbf{q}/|\hat{\mathbf{z}} \times \mathbf{q}|$ and $\hat{\mathbf{e}}_{q1} = \hat{\mathbf{e}}_{q2} \times \hat{\mathbf{e}}_{q3}$. The matrix element can now be expressed as

$$\langle f e \bar{\nu} | H_w | i \rangle = \frac{G_F V_{ud}}{\sqrt{2}} \left(\tilde{\ell}_0 h_a^0 - \tilde{\ell}_3 \hat{\mathbf{e}}_{q0}^\dagger \cdot \mathbf{h}_a - \sum_{\lambda=\pm 1} \tilde{\ell}_\lambda \hat{\mathbf{e}}_{q\lambda}^\dagger \cdot \mathbf{h}_a \right). \quad (\text{B13})$$

Here \mathcal{J}_{V-A}^μ denotes the hadronic realization of the quark current $\bar{u} \gamma^\mu (1 - \gamma_5) d$, and the derivation only assumes that such a realization exist. The first few orders of the explicit representation of \mathcal{J}_{V-A}^μ in chiral EFT will be given in Appendix C. Introducing the scalar and vector under rotations

$$\ell_0 = 2\langle e \bar{\nu} | \bar{e}_L \gamma_0 \nu_L | 0 \rangle \quad \text{and} \quad \boldsymbol{\ell} = 2\langle e \bar{\nu} | \bar{e}_L \boldsymbol{\gamma} \nu_L | 0 \rangle, \quad (\text{B3})$$

we can write the matrix element

$$\langle f e \bar{\nu} | H_w | i \rangle = -\frac{G_F V_{ud}}{\sqrt{2}} \int d^3\mathbf{x} (\boldsymbol{\ell} \cdot (\mathcal{J}_{V-A})_{fi} - \ell_0 (\mathcal{J}_{V-A}^0)_{fi}), \quad (\text{B4})$$

where $|i\rangle$ and $|f\rangle$ denote the initial and final nuclear states. We recall here that the leptonic tensor can be written as

$$\ell^\mu = e^{-i(\mathbf{p}_e + \mathbf{p}_\nu) \cdot \mathbf{x}} \tilde{\ell}^\mu, \quad (\text{B5})$$

with

$$\tilde{\ell}^\mu = \bar{u}_e \gamma^\mu (1 - \gamma_5) v_\nu, \quad (\text{B6})$$

with \bar{u}_e and v_ν spinors of the electron and electronic antineutrino respectively. The matrix element becomes

$$\begin{aligned} \langle f e \bar{\nu} | H_w | i \rangle &= -\frac{G_F V_{ud}}{\sqrt{2}} \int d^3\mathbf{x} e^{-i\mathbf{q} \cdot \mathbf{x}} \\ &\quad \times (\boldsymbol{\ell} \cdot (\mathcal{J}_{V-A})_{fi} - \tilde{\ell}_0 (\mathcal{J}_{V-A}^0)_{fi}), \end{aligned} \quad (\text{B7})$$

we can define now

$$(\mathcal{J}_{V-A}^\mu)_{fi}(-\mathbf{q}) = \int d\mathbf{x} e^{-i\mathbf{q} \cdot \mathbf{x}} (\mathcal{J}_{V-A}^\mu)_{fi}(\mathbf{x}). \quad (\text{B8})$$

We define $(\mathcal{J}_{V-A}^\mu)_{fi}(-\mathbf{q}) = (\mathcal{J}_{V-A}^\mu)_{fi}^\dagger(\mathbf{q})$ and, for ease of notation, we write

$$h_a^\mu = \langle f | \mathcal{J}_{V-A}^{\mu\dagger}(\mathbf{q}) | i \rangle \quad (\text{B9})$$

that leads to the following matrix element for the interaction Hamiltonian:

$$\langle f e \bar{\nu} | H_w | i \rangle = \frac{G_F V_{ud}}{\sqrt{2}} \tilde{\ell}_\mu h_a^\mu \quad (\text{B10})$$

Similarly to what has been done in Ref. [71] we can decompose the space part of the leptonic tensor $\tilde{\ell}^\mu$ in terms of spherical coordinates,

$$\tilde{\boldsymbol{\ell}} = \tilde{\ell}_3 \hat{\mathbf{e}}_{q0}^\dagger + \sum_{\lambda=\pm 1} \tilde{\ell}_\lambda \hat{\mathbf{e}}_{q\lambda}^\dagger, \quad (\text{B11})$$

where the $\hat{\mathbf{e}}_{q,\lambda}$ are defined as [45,71]

$$\hat{\mathbf{e}}_{q,\pm 1} = \mp \frac{1}{\sqrt{2}} (\hat{\mathbf{e}}_{q1} \pm \hat{\mathbf{e}}_{q2}), \quad \hat{\mathbf{e}}_{q0} = \hat{\mathbf{e}}_{q3}, \quad (\text{B12})$$

We can expand Eq. (B4) in a sum of terms with well defined total angular momentum. For the SM current one finds [71]

$$\begin{aligned} \langle f e \bar{\nu} | H_w | i \rangle = & -\frac{G_F}{\sqrt{2}} V_{ud} \langle f | \left\{ -\sum_{J \geq 1} \sqrt{2\pi(2J+1)} (-i)^J \sum_{\lambda=\pm 1} \ell_\lambda [\lambda \mathcal{M}_{J-\lambda}(q) + \mathcal{E}_{J-\lambda}(q)] \right. \\ & \left. + \sum_{J \geq 0} \sqrt{4\pi(2J+1)} (-i)^J (\ell_3 \mathcal{L}_{J0}(q) - \ell_0 \mathcal{C}_{J0}(q)) \right\} | i \rangle, \end{aligned} \quad (\text{B14})$$

The multipole operators, defined in Ref. [71], are

$$\mathcal{C}_{JM}(q) = \int d^3x j_J(qx) Y_{JM}(\Omega_x) [\mathcal{J}_V^0(x) + \mathcal{J}_A^0(x)], \quad (\text{B15})$$

$$\mathcal{L}_{JM}(q) = \frac{i}{q} \int d^3x \nabla [j_J(qx) Y_{JM}(\Omega_x)] \cdot [\mathcal{J}_V(x) + \mathcal{J}_A(x)], \quad (\text{B16})$$

$$\mathcal{E}_{JM}(q) = \int d^3x [\nabla \times j_J(qx) \mathfrak{Y}_{JJ1}^M(\Omega_x)] \cdot [\mathcal{J}_V(x) + \mathcal{J}_A(x)], \quad (\text{B17})$$

$$\mathcal{M}_{JM}(q) = \frac{i}{q} \int d^3x [j_J(qx) \mathfrak{Y}_{JJ1}^M(\Omega_x)] \cdot [\mathcal{J}_V(x) + \mathcal{J}_A(x)], \quad (\text{B18})$$

in terms of spherical Bessel function j_J , spherical harmonics Y_{JM} , and vector spherical harmonics $\mathfrak{Y}_{JJ1}^M(\Omega_x)$ (see Ref. [71] for the relevant definitions). The matrix elements L_1 , E_1 , M_1 , and C_1 that enter the decay rate in Eqs. (14) and (D2) are the reduced matrix elements of the operators (B15)–(B18). The matrix elements between a generic multipole \mathcal{T}_{JM} between initial and final nuclear states can be written as

$$\langle J_f, M_f | \mathcal{T}_{JM}^\dagger | J_i, M_i \rangle = \alpha (-)^M (-1)^{J_i - M_i} \frac{1}{\sqrt{2L+1}} \langle J_f, M_f; J_i, -M_i | J, -M \rangle T_J(q) \quad (\text{B19})$$

with $\alpha = 1$ for the multipole operator \mathcal{C} and $\alpha = -1$ for the multipole operators \mathcal{L} , \mathcal{E} , and \mathcal{M} . We denote with $T_J(q)$ the reduced matrix element associated with the generic multipole operator \mathcal{T}_{JM} . We are now in the position to obtain Eqs. (4)–(7) reported in the main text. We recall that for the problem of interest $J_i = 0$ and $J_f = 1$ therefore, using selection rules and the identities of Ref. [45] we arrive at

$$h_a^0 = -4\pi i Y_{10}(\hat{\mathbf{q}}) \frac{1}{\sqrt{3}} C_1(q), \quad (\text{B20})$$

$$\hat{\mathbf{e}}_{q0}^\dagger \cdot \mathbf{h}_a = -4\pi i Y_{1,0}(\hat{\mathbf{q}}) \frac{1}{\sqrt{3}} L_1(q), \quad (\text{B21})$$

$$\hat{\mathbf{e}}_{q\lambda}^\dagger \cdot \mathbf{h}_a = -\sqrt{2\pi} i D_{M_f, \lambda}^{1\dagger}(-\phi, -\theta, \phi) [\lambda M_1(q) + E_1(q)], \quad (\text{B22})$$

where in the last passage we recall the following definition

$$D_{M_f, \lambda}^{1\dagger}(-\phi, -\theta, \phi) = \sqrt{\frac{4\pi}{3}} Y_{1, \lambda}^*(\theta, \phi). \quad (\text{B23})$$

We can now see that taking $\hat{\mathbf{q}}$ along $\hat{\mathbf{z}}$, Eqs. (B20) and (B21) lead to Eqs. (4) and (5). Finally taking in Eq. (B22) $\hat{\mathbf{q}}$ along $\hat{\mathbf{x}}$ we obtain a linear system of two equations for two different values of λ , whose solution leads to Eqs. (6) and (7).

We can generalize Eq. (B14) to nonstandard currents induced by (ν)SMEFT operators. The most general Hamiltonian including dimension-six operator in the SMEFT has the form

$$H_6 = -\frac{G_F}{\sqrt{2}} \int d^3\mathbf{x} (-j_S^{\text{lept}} \mathcal{J}_S - j_P^{\text{lept}} \mathcal{J}_P + 2(j_T^{\text{lept}})^{i0} \mathcal{J}_T^{i0} - (j_T^{\text{lept}})_{ij} \mathcal{J}_T^{ij} + \mathbf{j}_V^{\text{lept}} \cdot \mathcal{J}_V + \mathbf{j}_A^{\text{lept}} \cdot \mathcal{J}_A - (j_0^{\text{lept}})_V \mathcal{J}_V^0 - (j_0^{\text{lept}})_A \mathcal{J}_A^0), \quad (\text{B24})$$

where the leptonic currents are

$$j_S^{\text{lept}} = 2(C_{\text{SRL}}^{(6)} + C_{\text{SLL}}^{(6)}) \bar{e}_R \nu + 2(C_{\text{SRR}}^{(6)} + C_{\text{SLR}}^{(6)}) \bar{e}_L \nu, \quad (\text{B25})$$

$$j_P^{\text{lept}} = 2(C_{\text{SRL}}^{(6)} - C_{\text{SLL}}^{(6)}) \bar{e}_R \nu + 2(C_{\text{SRR}}^{(6)} - C_{\text{SLR}}^{(6)}) \bar{e}_L \nu, \quad (\text{B26})$$

$$(j_T^{\text{lept}})^{\mu\nu} = 4C_{\text{TLL}}^{(6)} \bar{e}_R \sigma^{\mu\nu} \nu + 4C_{\text{TRR}}^{(6)} \bar{e}_L \sigma^{\mu\nu} \nu, \quad (\text{B27})$$

$$(j_V^{\text{lept}})^\mu = 2(C_{\text{VRL}}^{(6)} + C_{\text{VLL}}^{(6)}) \bar{e}_L \gamma^\mu \nu + 2(C_{\text{VRR}}^{(6)} + C_{\text{VLR}}^{(6)}) \bar{e}_R \gamma^\mu \nu, \quad (\text{B28})$$

$$(j_A^{\text{lept}})^\mu = 2(C_{\text{VLL}}^{(6)} - C_{\text{VRL}}^{(6)}) \bar{e}_L \gamma^\mu \nu + 2(C_{\text{VLR}}^{(6)} - C_{\text{VRR}}^{(6)}) \bar{e}_R \gamma^\mu \nu, \quad (\text{B29})$$

and the hadronic currents the nucleon-level realization of

$$\bar{u}d \rightarrow \mathcal{J}_S, \quad \bar{u}\gamma_5 d \rightarrow \mathcal{J}_P, \quad \bar{u}\sigma^{\mu\nu}d \rightarrow \mathcal{J}_T^{\mu\nu}, \quad \bar{u}\gamma^\mu d \rightarrow \mathcal{J}_V, \quad \bar{u}\gamma^\mu\gamma_5 d \rightarrow -\mathcal{J}_A. \quad (\text{B30})$$

Introducing the leptonic matrix elements

$$\ell_S = \langle e\bar{\nu}|j_S^{\text{lept}}|0\rangle, \quad \ell_P = \langle e\bar{\nu}|j_P^{\text{lept}}|0\rangle, \quad (\ell_0)_{V,A} = \langle e\bar{\nu}|(j_{V,A}^{\text{lept}})^0|0\rangle, \quad (\text{B31})$$

which are scalar under rotations, and the vectors

$$\ell_T = -\frac{1}{2}\varepsilon^{ijk}\langle e\bar{\nu}|(j_T^{\text{lept}})^{ij}|0\rangle, \quad \ell'_T = 2\langle e\bar{\nu}|(j_T^{\text{lept}})^{i0}|0\rangle, \quad \ell_{V,A} = \langle e\bar{\nu}|(j_{V,A}^{\text{lept}})^i|0\rangle, \quad (\text{B32})$$

the matrix element of the Hamiltonian becomes

$$\begin{aligned} \langle f e\bar{\nu}|H_6|i\rangle = & -\frac{G_F}{\sqrt{2}} \int d^3\mathbf{x} (-\ell_S(\mathcal{J}_S)_{fi} - \ell_P(\mathcal{J}_P)_{fi} - (\ell_0)_V(\mathcal{J}_V^0)_{fi} - (\ell_0)_A(\mathcal{J}_A^0)_{fi} + \ell'_T \cdot (\mathcal{J}'_T)_{fi} \\ & + \ell_T \cdot (\mathcal{J}_T)_{fi} + \ell_V \cdot (\mathcal{J}_V)_{fi} + \ell_A \cdot (\mathcal{J}_A)_{fi}), \end{aligned} \quad (\text{B33})$$

where

$$(\mathcal{J}'_T)_{fi}^k = \langle f|\mathcal{J}_T^{k0}|i\rangle, \quad (\mathcal{J}_T)_{fi}^k = \varepsilon^{klm}\langle f|\mathcal{J}_T^{lm}|i\rangle. \quad (\text{B34})$$

Since at dimension 6 in the ν SMEFT the leptonic and hadronic currents have at most spin 1 [121], for both scalar/tensor and vector/axial operators, the derivation of the multipole expansion therefore proceeds as in the SM. Additional vector and axial operators generate exactly the same multipoles as in the SM. For scalar and pseudoscalar currents, only the \mathcal{C}_{J0} multipole is present. Tensor currents generate electric, magnetic, and longitudinal multipoles, but not \mathcal{C}_J . One thus finds

$$\begin{aligned} \langle f|H_6|i\rangle = & -\frac{G_F}{\sqrt{2}}\langle f|\left\{ -\sum_{J\geq 1}\sqrt{2\pi(2J+1)}(-i)^J\sum_{\lambda=\pm 1}\ell'_\lambda[\lambda\mathcal{M}_{J-\lambda}^T(q) + \mathcal{E}_{J-\lambda}^T(q)] \right. \\ & -\sum_{J\geq 1}\sqrt{2\pi(2J+1)}(-i)^J\sum_{\lambda=\pm 1}\ell''_\lambda[\lambda\mathcal{M}_{J-\lambda}^{T'}(q) + \mathcal{E}_{J-\lambda}^{T'}(q)] + \sum_{J\geq 0}\sqrt{4\pi(2J+1)}(-i)^J[\ell_3^T\mathcal{L}_{J0}^T(q) + \ell_3^{T'}\mathcal{L}_{J0}^{T'}(q)] \\ & \left. -\sum_{J\geq 0}\sqrt{4\pi(2J+1)}(-i)^J[\ell_S\mathcal{C}_{J0}^S(q) + \ell_P\mathcal{C}_{J0}^P(q)]\right\}|i\rangle, \end{aligned} \quad (\text{B35})$$

with

$$\mathcal{C}_{JM}^S(q) = \int d^3x j_J(qx)Y_{JM}(\Omega_x)\mathcal{J}_S(x), \quad (\text{B36})$$

$$\mathcal{C}_{JM}^P(q) = \int d^3x j_J(qx)Y_{JM}(\Omega_x)\mathcal{J}_P(x), \quad (\text{B37})$$

and

$$\mathcal{L}_{JM}^{T(\prime)}(q) = \frac{i}{q} \int d^3x \nabla[j_J(qx)Y_{JM}(\Omega_x)] \cdot \mathcal{J}_T^{(\prime)}(x), \quad (\text{B38})$$

$$\mathcal{E}_{JM}^{T(\prime)}(q) = \int d^3x [\nabla \times j_J(qx)\mathcal{Y}_{J1}^M(\Omega_x)] \cdot \mathcal{J}_T^{(\prime)}(x), \quad (\text{B39})$$

$$\mathcal{M}_{JM}^{T(\prime)}(q) = \frac{i}{q} \int d^3x [j_J(qx)\mathcal{Y}_{J1}^M(\Omega_x)] \cdot \mathcal{J}_T^{(\prime)}(x). \quad (\text{B40})$$

Similar results were obtained in Ref. [106].

For the calculation of the ${}^6\text{He}(0^+) \rightarrow {}^6\text{Li}(1^+)$ transition, only multipoles with $J = 1$ and positive parity are needed. This leaves $\mathcal{C}_{10}(q, A)$, $\mathcal{L}_{10}(q, A)$, $\mathcal{E}_{1\lambda}(q, A)$, and $\mathcal{M}_{1\lambda}(q, V)$ for the SM currents and BSM axial and vector currents. For BSM scalar and tensor currents, the only nonvanishing multipoles are $\mathcal{C}_{10}^P(q)$, $\mathcal{L}_{10}^T(q)$, $\mathcal{E}_{1\lambda}^T(q)$, and $\mathcal{M}_{1\lambda}^{T'}$. The steps to express the matrix elements of the operators in Eqs. (B36)–(B40) in terms of momentum-space currents are analogous to those discussed for the SM.

APPENDIX C: CHARGED CURRENTS IN CHIRAL EFT

We report in this Appendix, for completeness, the well-known lowest-order expressions of the SM vector and axial currents [51,59,119,120,122]. We also report the currents induced by SMEFT operators. We preliminarily define, for nucleons of incoming momentum \mathbf{p}_i and outgoing momentum \mathbf{p}'_i , the center of mass and relative momenta \mathbf{K}_i and \mathbf{k}_i in the following way:

$$\mathbf{K}_i \equiv \frac{\mathbf{p}'_i + \mathbf{p}_i}{2}, \quad \mathbf{k}_i = \mathbf{p}'_i - \mathbf{p}_i, \quad (\text{C1})$$

and similarly $\mathbf{q} = \mathbf{p}_1 + \mathbf{p}_2$. We can express the charge operator up to the order of interest as

$$\rho_{5,a}(\mathbf{q}) = \sum_{\nu \in \{-2, -1\}} \rho_{5,a}^{(\nu)}(\mathbf{q}), \quad (\text{C2})$$

and similarly for the current we have

$$\mathbf{j}_{5,a}(\mathbf{q}) = \sum_{\nu \in \{-3, 0\}} \mathbf{j}_{5,a}^{(\nu)}(\mathbf{q}), \quad (\text{C3})$$

where ν is the chiral order defined as in Ref. [60]. The leading order $\nu = -2$ and next-to-leading order $\nu = -1$ axial

charge read

$$\begin{aligned}\rho_a^{(-2)}(\mathbf{q}, A) &= \rho_{\text{recoil}}^{(-2)}(\mathbf{q}, A) + \rho_{\text{pseudo}}^{(-2)}(\mathbf{q}, A) \\ &= -\frac{g_A}{2}\tau_{1,a}\left(\frac{1}{m_N}\boldsymbol{\sigma}_1 \cdot \mathbf{K}_1 - \frac{q^0}{\mathbf{q}^2 + m_\pi^2}\boldsymbol{\sigma}_1 \cdot \mathbf{q}\right)(2\pi)^3\delta(\mathbf{k}_1 - \mathbf{q}) + 1 \leftrightarrow 2,\end{aligned}\quad (\text{C4})$$

$$\rho_a^{(-1)}(\mathbf{q}, A) = i\frac{g_A}{4F_\pi^2}(\boldsymbol{\tau}_1 \times \boldsymbol{\tau}_2)_a \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{k}_2}{\omega_2^2} + 1 \leftrightarrow 2,\quad (\text{C5})$$

where the first term on the second line of Eq. (C4), suppressed by $1/m_N$, contributes to $\rho_{\text{recoil}}^{(-2)}(\mathbf{q}, A)$, while the second term, proportional to the electron-neutrino energy q_0 , is induced by the induced pseudoscalar form factor. Here we use $F_\pi \approx 92$ MeV. Similarly the leading order $\nu = -3$ and next-to-leading order $\nu = 0$ contributions to the axial current read

$$\mathbf{j}_{5,a}^{(-3)}(\mathbf{q}) = -\frac{g_A}{2}\tau_{1,a}\left[\boldsymbol{\sigma}_1 - \frac{\mathbf{q}}{q^2 + m_\pi^2}\boldsymbol{\sigma}_1 \cdot \mathbf{q}\right](2\pi)^3\delta(\mathbf{k}_1 - \mathbf{q}) + 1 \leftrightarrow 2,\quad (\text{C6})$$

$$\mathbf{j}_{5,a}^{(0)} = \tilde{\mathbf{j}}_{5,a}^{(0)}(\mathbf{q}) - \frac{\mathbf{q}}{q^2 + m_\pi^2}\mathbf{q} \cdot \tilde{\mathbf{j}}_{5,a}^{(0)} + \frac{ig_A}{4F_\pi^2 m}(\boldsymbol{\tau}_1 \times \boldsymbol{\tau}_2)_a \frac{\mathbf{q}}{q^2 + m_\pi^2}(\mathbf{K}_1 \cdot \mathbf{k}_1 + \mathbf{K}_2 \cdot \mathbf{k}_2) \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{k}_2}{\omega_2^2} + 1 \leftrightarrow 2,\quad (\text{C7})$$

where we have defined for convenience

$$\tilde{\mathbf{j}}_{5,a}^{(0)}(\mathbf{q}) = \frac{g_A^2}{2F_\pi^2} \left\{ 2c_3\tau_{2,a}\mathbf{k}_2 + (\boldsymbol{\tau}_1 \times \boldsymbol{\tau}_2)_a \left[\frac{i}{2m}\mathbf{K}_1 - \frac{c_6 + 1}{4m}\boldsymbol{\sigma}_1 \times \mathbf{q} + \left(c_4 + \frac{1}{4m} \right) \boldsymbol{\sigma}_1 \times \mathbf{k}_2 \right] \right\} \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{k}_2}{\omega_2^2} + 1 \leftrightarrow 2.\quad (\text{C8})$$

In the code, we replace g_A with the dipole parametrization of the axial form factor,

$$g_A(|\mathbf{q}|^2) = g_A \frac{1}{(1 + q^2/\Lambda_A^2)^2},\quad (\text{C9})$$

with $\Lambda_A = 1.05$ GeV. c_3 , c_4 , and c_6 are NLO low-energy constants.

Neglecting isospin-breaking effects, the charged vector current is an isospin rotation of the isovector component of the electromagnetic current. The leading term is induced by the isovector magnetic moment and by a recoil correction [51],

$$\mathbf{j}_a^{(-2)} = \frac{\tau_{1,a}}{2} \frac{1}{2m_N} (2\mathbf{K}_1 + i(1 + \kappa_V)\boldsymbol{\sigma}_1 \times \mathbf{q}),\quad (\text{C10})$$

with $\kappa_V \approx 3.7$. The NLO contributions originate from the exchange of a pion between nucleon lines, with the vector current coupling either to the pion in flight or to the nucleon. This contribution gives

$$\begin{aligned}\mathbf{j}_a^{(-1)} &= i\frac{g_A^2}{F_\pi^2}(\boldsymbol{\tau}_1 \times \boldsymbol{\tau}_2)_a \left(-\boldsymbol{\sigma}_1 \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{k}_2}{\omega_2^2} + \frac{\mathbf{k}_1 - \mathbf{k}_2}{2\omega_1^2\omega_2^2} \boldsymbol{\sigma}_1 \cdot \mathbf{k}_1 \boldsymbol{\sigma}_2 \cdot \mathbf{k}_2 \right) \\ &+ 1 \leftrightarrow 2.\end{aligned}\quad (\text{C11})$$

For currents induced by SMEFT and ν SMEFT operators, we just retain one-body contributions. Recalling that, at the quark level,

$$\mathcal{J}_S^a = \bar{q} \frac{\tau^a}{2} q, \quad \mathcal{J}_P^a = \bar{q} \gamma_5 \frac{\tau^a}{2} q, \quad \mathcal{J}_T^{\mu\nu a} = \bar{q} \sigma^{\mu\nu} \frac{\tau^a}{2} q,\quad (\text{C12})$$

where q denotes a quark doublet $q = (u, d)^T$, the scalar and pseudoscalar densities and the tensor currents are then given by

$$\mathcal{J}_{S,a}^{(-3)} = g_S \frac{\tau_{1,a}}{2} (2\pi)^3 \delta(\mathbf{k}_1 - \mathbf{q}) + 1 \leftrightarrow 2,\quad (\text{C13})$$

$$\mathcal{J}_{P,a}^{(-4)} = \frac{g_A B}{m_\pi^2 + \mathbf{q}^2} \frac{\tau_{1,a}}{2} \boldsymbol{\sigma}_1 \cdot \mathbf{q} (2\pi)^3 \delta(\mathbf{k}_1 - \mathbf{q}) + 1 \leftrightarrow 2,\quad (\text{C14})$$

$$\mathcal{J}_{T,a}^{ij(-3)} = g_T \varepsilon^{ijk} \frac{\tau_{1,a}}{2} \sigma_1^k (2\pi)^3 \delta(\mathbf{k}_1 - \mathbf{q}) + 1 \leftrightarrow 2,\quad (\text{C15})$$

$$\mathcal{J}_{T,a}^{i0(-2)} = i\frac{g'_T}{2m_N} q^i \frac{\tau_{1,a}}{2} (2\pi)^3 \delta(\mathbf{k}_1 - \mathbf{q}) + 1 \leftrightarrow 2,\quad (\text{C16})$$

and the currents defined in Appendix B are given by, for example, $\mathcal{J}_T^{ij} = \mathcal{J}_{T,x}^{ij} + i\mathcal{J}_{T,y}^{ij}$, where x and y are isospin indices. The values of the scalar and tensor charges are given in Eq. (A17). The exact value of g'_T is unknown, but should be a number of order 1.

APPENDIX D: FULLY DIFFERENTIAL UNPOLARIZED DECAY RATE

In the main text we gave the expression for the decay rate, differential in the electron energy. We give here more differential expressions, which can be used, for example, to extract corrections to the β - ν correlation a . We start from the expression given in Ref. [45,71] for the specific case of ${}^6\text{He}$. The starting point is the usual Fermi's golden rule for the unpolarized differential decay rate

$$d\Gamma = (2\pi)\delta(E_i - E_f - E_\nu - E_e) \frac{1}{2J_i + 1} \sum_{M_i M_f} \sum_{s_e s_\nu} |M|^2 \frac{d^3\mathbf{p}_e}{(2\pi)^3} \frac{d^3\mathbf{p}_\nu}{(2\pi)^3},\quad (\text{D1})$$

with $\mathbf{P}_f = -\mathbf{p}_e - \mathbf{p}_{\nu_e}$. Using the multipole expansion of the matrix element of the SM weak Hamiltonian, Eq. (B14), one gets

$$\begin{aligned}
d\Gamma &= 2\pi\delta(M_i - E_f - E_e - E_\nu)G_F^2V_{ud}^2\frac{4\pi}{2J_i+1}\left[(1+\mathbf{v}_e\cdot\mathbf{v}_\nu)\sum_{l\geq 0}|C_l(q)|^2+(1-\mathbf{v}_e\cdot\mathbf{v}_\nu+2\mathbf{v}_e\cdot\hat{\mathbf{q}}\mathbf{v}_\nu\cdot\hat{\mathbf{q}})\sum_{l\geq 0}|L_l(q)|^2\right. \\
&\quad -2\hat{\mathbf{q}}\cdot(\mathbf{v}_e+\mathbf{v}_\nu)\sum_{l\geq 0}\text{Re}[C_l(q)L_l^*(q)]+(1-\mathbf{v}_e\cdot\hat{\mathbf{q}}\mathbf{v}_\nu\cdot\hat{\mathbf{q}})\sum_{l\geq 1}[|M_l(q)|^2+|E_l(q)|^2] \\
&\quad \left.-2\hat{\mathbf{q}}\cdot(\mathbf{v}_e-\mathbf{v}_\nu)\sum_{l\geq 1}\text{Re}[M_l(q)E_l^*(q)]\right]\frac{d^3\mathbf{p}_e}{(2\pi)^3}\frac{d^3\mathbf{p}_\nu}{(2\pi)^3}, \tag{D2}
\end{aligned}$$

with $\mathbf{q} = \mathbf{p}_e + \mathbf{p}_\nu$, $\hat{\mathbf{q}} = \mathbf{q}/|\mathbf{q}|$, $\mathbf{v}_e = \mathbf{p}_e/\sqrt{p_e^2 + m_e^2}$, and $\mathbf{v}_\nu = \mathbf{p}_\nu/E_\nu$. We notice that for the transition considered $|J_i - J_f| = \pm 1, 0$ and $\pi_i\pi_f = 1$, which allow one to simplify the expressions to

$$\begin{aligned}
d\Gamma &= 2\pi\delta(M_i - E_f - E_e - E_\nu)G_F^2V_{ud}^2\frac{4\pi}{2J_i+1}[(1+\mathbf{v}_e\cdot\mathbf{v}_\nu)|C_1(q;A)|^2+(1-\mathbf{v}_e\cdot\mathbf{v}_\nu+2\mathbf{v}_e\cdot\hat{\mathbf{q}}\mathbf{v}_\nu\cdot\hat{\mathbf{q}})|L_1(q;A)|^2 \\
&\quad -2\hat{\mathbf{q}}\cdot(\mathbf{v}_e+\mathbf{v}_\nu)\text{Re}[C_1(q;A)L_1^*(q;A)]+(1-\mathbf{v}_e\cdot\hat{\mathbf{q}}\mathbf{v}_\nu\cdot\hat{\mathbf{q}})[|M_1(q;V)|^2+|E_1(q;A)|^2] \\
&\quad -2\hat{\mathbf{q}}\cdot(\mathbf{v}_e-\mathbf{v}_\nu)\text{Re}[M_1(q;V)E_1^*(q;A)]]\frac{d^3\mathbf{p}_e}{(2\pi)^3}\frac{d^3\mathbf{p}_\nu}{(2\pi)^3}. \tag{D3}
\end{aligned}$$

Retaining terms up to order q^2 we have

$$\begin{aligned}
d\Gamma &= 2\pi\delta(E_i - E_f - E_\nu - E_e)G_F^2V_{ud}^2\frac{4\pi}{2J_i+1}\frac{1}{9}\left\{(3-\mathbf{v}_e\cdot\mathbf{v}_\nu)|L_1^{(0)}(A)|^2-2\hat{\mathbf{q}}\cdot(\mathbf{v}_e+\mathbf{v}_\nu)qr_\pi\text{Re}[C_1^{(1)}(A)L_1^{(0)*}(A)]\right. \\
&\quad -2\hat{\mathbf{q}}\cdot(\mathbf{v}_e-\mathbf{v}_\nu)qr_\pi\text{Re}[M_1^{(1)}(V)E_1^{(0)*}(A)]+(1+\mathbf{v}_e\cdot\mathbf{v}_\nu)(qr_\pi)^2|C_1^{(1)}(A)|^2 \\
&\quad \left.+ (qr_\pi)^2(1-\mathbf{v}_e\cdot\hat{\mathbf{q}}\mathbf{v}_\nu\cdot\hat{\mathbf{q}})\left[|M_1^{(1)}(V)|^2-\frac{1}{5}\text{Re}(E_1^{(0)}(A)E_1^{(2)}(A))\right]\right. \\
&\quad \left.-\frac{(qr_\pi)^2}{5}(1-\mathbf{v}_e\cdot\mathbf{v}_\nu+2\mathbf{v}_\nu\cdot\hat{\mathbf{q}}\mathbf{v}_e\cdot\hat{\mathbf{q}})\text{Re}(L_1^{(0)}(A)L_1^{(2)}(A))\right\}\frac{d^3\mathbf{p}_e}{(2\pi)^3}\frac{d^3\mathbf{p}_\nu}{(2\pi)^3}, \tag{D4}
\end{aligned}$$

where we used the expansion in Eqs. (10)–(13). Integrating over the neutrino phase space and the electron angular variables, we obtain Eq. (14).

With the definitions in Eqs. (10)–(13), the coefficients of the momentum expansion of the multipole operators are real. Dropping the labels A and V and using $E_1^{(0)} = \sqrt{2}L_1^{(0)}$, we can write

$$\begin{aligned}
d\Gamma &= 2\pi\delta(E_i - E_f - E_\nu - E_e)G_F^2V_{ud}^2\frac{4\pi}{2J_i+1}\frac{1}{3}|L_1^{(0)}(A)|^2\frac{d^3\mathbf{p}_e}{(2\pi)^3}\frac{d^3\mathbf{p}_\nu}{(2\pi)^3} \\
&\quad \times \left\{1+\frac{2W_0r_\pi}{3}\left[\left(1-2\varepsilon+\frac{\mu_e^2}{\varepsilon}\right)\frac{\sqrt{2}M_1^{(1)}}{L_1^{(0)}}-\left(1-\frac{\mu_e^2}{\varepsilon}\right)\frac{C_1^{(1)}}{L_1^{(0)}}\right]\right. \\
&\quad +\frac{(W_0r_\pi)^2}{3}\left[\left(1-\frac{4}{3}\varepsilon(1-\varepsilon)-\frac{\mu_e^2}{\varepsilon}\frac{2+\varepsilon}{3}\right)\left(\frac{C_1^{(1)}}{L_1^{(0)}}\right)^2-\frac{1}{5}\left(1-\frac{\mu_e^2}{\varepsilon}(2-\varepsilon)\right)\frac{L_1^{(2)}}{L_1^{(0)}}\right. \\
&\quad \left.+\left(1-\frac{10}{3}\varepsilon(1-\varepsilon)+\frac{\mu_e^2}{\varepsilon}\frac{4-7\varepsilon}{3}\right)\frac{\tilde{M}_1^{(2)}}{L_1^{(0)}}\right] \\
&\quad +\mathbf{v}_e\cdot\mathbf{v}_\nu\left[-\frac{1}{3}-\frac{2W_0r_\pi}{3}\left((1-2\varepsilon)\frac{\sqrt{2}M_1^{(1)}}{L_1^{(0)}}+\frac{C_1^{(1)}}{L_1^{(0)}}\right)\right. \\
&\quad \left.+\frac{(W_0r_\pi)^2}{3}\left((1-\mu_e^2)\left\{\left(\frac{C_1^{(1)}}{L_1^{(0)}}\right)^2-\frac{1}{5}\frac{L_1^{(2)}}{L_1^{(0)}}\right\}-\left(1-4\varepsilon(1-\varepsilon)-\mu_e^2\right)\frac{\tilde{M}_1^{(2)}}{L_1^{(0)}}\right)\right] \\
&\quad \left.+\frac{(W_0r_\pi)^2}{3}\left((\mathbf{v}_e\cdot\mathbf{v}_\nu)^2-\frac{1}{3}\frac{|\mathbf{p}_e|^2}{E_e^2}\right)\varepsilon(1-\varepsilon)\left[2\left(\frac{C_1^{(1)}}{L_1^{(0)}}\right)^2-\frac{\tilde{M}_1^{(2)}}{L_1^{(0)}}\right]\right\}, \tag{D5}
\end{aligned}$$

where we introduced the scaled variables ε and μ_e as in Sec. II, and we defined

$$\tilde{M}_1^{(2)} = L_1^{(0)} \left(\left(\frac{M_1^{(1)}}{L_1^{(0)}} \right)^2 - \frac{\sqrt{2} E_1^{(2)}}{5 L_1^{(0)}} \right). \quad (\text{D6})$$

From Eq. (D5), we can read the corrections to the β - ν angular correlation a and to the subleading correlations $C_{(aa)}$ [4], proportional to $(\mathbf{v}_e \cdot \mathbf{v}_\nu)^2$. From these expressions, we find that the β - ν correlation coefficient, averaged over energy, is

$$\langle a \rangle = -\frac{1}{3} - 2.3(4) \times 10^{-4} - 1.2(2) \times 10^{-3}, \quad (\text{D7})$$

where we used GFMC matrix elements, and we did not include radiative corrections. The second term is the correction linear in $M_1^{(1)}$ and $C_1^{(1)}$. The term proportional to $M_1^{(1)}$ averages to almost zero, leaving behind a small contribution from $C_1^{(1)}$. The last term appears at second order in the multipole expansion. Both corrections are still below the experimental sensitivity [123].

For tensor and pseudoscalar currents, we find that the interference with the SM is given, at lowest order in the multipole expansion, by

$$d\Gamma = 2\pi \delta(E_i - E_f - E_\nu - E_e) G_F^2 V_{ud}^2 \frac{4\pi}{2J_i + 1} \frac{4m_e}{E_e} \frac{d^3\mathbf{p}_\nu}{(2\pi)^3} \frac{d^3\mathbf{p}_e}{(2\pi)^3} \\ \times \left[\epsilon_T \text{Re}[E_1(q, T)E_1^*(q, A) + L_1(q, T)L_1^*(q, A)] - \frac{\epsilon_P}{2} \hat{\mathbf{v}}_\nu \cdot \hat{\mathbf{q}} \text{Re}[C_1(q, P)L_1^*(q, A)] \right], \quad (\text{D8})$$

which, using the expansion in Eqs. (10)–(13) becomes

$$d\Gamma = 2\pi \delta(E_i - E_f - E_\nu - E_e) G_F^2 V_{ud}^2 \frac{4\pi}{2J_i + 1} \frac{4m_e}{9E_e} \frac{d^3\mathbf{p}_\nu}{(2\pi)^3} \frac{d^3\mathbf{p}_e}{(2\pi)^3} \\ \times \left[\epsilon_T (E_1^{(0)}(T)E_1^{(0)}(A) + L_1^{(0)}(T)L_1^{(0)}(A)) - \frac{\epsilon_P}{2} W_0 r_\pi ((1 - \varepsilon) + \varepsilon \mathbf{v}_e \cdot \mathbf{v}_\nu) C_1^{(1)}(P)L_1^{(0)}(A) \right], \quad (\text{D9})$$

so that tensor interactions only induce a Fierz interference term, while pseudoscalar interactions also affect a . The expressions for the decay rates in the presence of sterile neutrinos can be found in a similar way.

APPENDIX E: HIGHER-ORDER ELECTROWEAK AND KINEMATIC RECOIL CORRECTIONS

1. Coulomb and radiative corrections

The multipole expansion of the weak Hamiltonian described at the start of this work does not explicitly take into account electromagnetic interactions. The treatment of such effects in nuclear β decay has historically been divided into three parts: (i) Coulomb corrections via the outgoing β particle in the field of the final atomic state; (ii) infrared-divergent contributions from various virtual and real emission photon processes up to some $\mathcal{O}(\alpha^n Z^m)$ ($m > n$) not already contained in (i); (iii) the remainder of electroweak radiative corrections, which are independent of the process kinematics. The latter two are typically described as outer and inner radiative corrections, respectively, and are the topic of a significant body of literature [124,125] and will only be plugged into the final result. The Coulomb corrections, on the other hand, are of interest, and to first order can be understood as the elastic response of the γW box diagrams. Higher-order (i.e., $\mathcal{O}\{\alpha Z^n\}$) behavior can then be absorbed by substituting the electron plane wave by a solution of the Dirac equation in the static, spherical potential of the atomic final state, $\phi_e(\mathbf{r}, \mathbf{p}_e)$, and writing the Hamiltonian as

$$\mathcal{M}_{fi} = \int d^3r \bar{\phi}_e(\mathbf{r}, \mathbf{p}_e) \gamma^\mu (1 - \gamma^5) v(\mathbf{p}_\nu) \\ \times \int \frac{d^3s}{(2\pi)^3} e^{is \cdot r} \frac{1}{2} [\langle f(\mathbf{p}_f + \mathbf{p}_e - \mathbf{s}) | V_\mu + A_\mu | i(\mathbf{p}_i) \rangle \\ + \langle f(\mathbf{p}_f) | V_\mu + A_\mu | i(\mathbf{p}_i - \mathbf{p}_e + \mathbf{s}) \rangle]. \quad (\text{E1})$$

The multipole expansion then proceeds analogously, and the most immediate modification is the introduction of the Fermi function [Eq. (15)], i.e., the $j = 1/2$ large components of the electron wave function for a point charge. The small and $j > 1/2$ components show up as small modifications in the radial integrals folded together with the nuclear current and introduce additional small terms to the differential decay rate. These are well known in the literature and can be found in several places [74,126]. Additional correction terms are well known for moving past a point-charge model of the nucleus and introducing additional subdominant electromagnetic corrections such as screening by atomic electrons and atomic exchange processes. We use the results of Ref. [76] by translating our results into the Behrens-Bühring formalism [95,127]:

$$L_1^0 = C\sqrt{3}^A F_{101}^{(0)}, \quad E_1^0 = C\sqrt{6}^A F_{101}^{(0)}, \quad (\text{E2a})$$

$$C_1^1 = -C \frac{R}{r_\pi} {}^A F_{110}^{(0)}, \quad M_1^1 = -C \frac{R}{r_\pi} {}^V F_{111}^{(0)}, \quad (\text{E2b})$$

$$L_1^2 = C \frac{1}{\sqrt{3}} \left(\frac{R}{r_\pi} \right)^2 (5^A F_{101}^{(1)} - 2\sqrt{2}^A F_{121}^{(0)}), \quad (\text{E2c})$$

$$E_1^2 = C \sqrt{\frac{2}{3}} \left(\frac{R}{r_\pi} \right)^2 (5^A F_{101}^{(1)} + \sqrt{2}^A F_{121}^{(0)}), \quad (\text{E2d})$$

where R is the nuclear radius of the uniformly charged sphere, i.e., $R = \sqrt{5/3}(r^2)_{\text{exp}}$ and $C = \sqrt{2J_i + 1}/4\pi$. This results in

the additional terms

$$C_{\text{Coulomb}}(Z, \varepsilon) = \alpha Z W_0 R \left[|L_1^0|^2 \left(\frac{6}{35} - \frac{233}{210} \frac{\alpha Z}{W_0 R} - \frac{3}{70} \frac{\mu_e^2}{\varepsilon} - \frac{12}{7} \varepsilon \right) + \left(\frac{r_\pi}{R} \right)^2 \left(\frac{E_1^0 E_1^2}{2} - L_1^0 L_1^2 \right) \left(-\frac{20}{35} + \frac{4}{7} \varepsilon \right) \right], \quad (\text{E3})$$

where the fractional prefactors were calculated assuming a uniformly charged sphere as described in Ref. [76].

Putting everything together, the inclusion of electroweak corrections modifies the β spectrum by

$$\frac{d\Gamma}{d\varepsilon} \propto (1 + \Delta_R^V)(1 + \delta_R(\varepsilon))F_0(Z, \varepsilon)L_0(Z, \varepsilon)S(Z, \varepsilon)[C_0(\varepsilon) + C_{\text{Coulomb}}(Z, \varepsilon)] \quad (\text{E4})$$

where Δ_R^V are the inner radiative corrections to vector transitions (the difference induced by axial transitions is small and absorbed into an experimentally determined g_A value) and $\delta_R(\varepsilon)$ are outer radiative corrections to $\mathcal{O}(\alpha^3 Z^2)$. Further, $F_0 L_0$ is the Fermi function for a uniformly charged sphere, S describes the shielding of the nuclear charge by atomic electrons, and $C_{\text{Coulomb}}(Z, \varepsilon)$ are modifications due to higher-order Coulomb corrections folded with the nuclear current discussed above.

2. Kinematic recoil corrections

In Appendix B we defined the multipole expansion of the weak Hamiltonian as introduced by Donnelly and Walecka [71]. This expansion, however, is not Lorentz covariant but implicitly performed in the Breit (brick wall) system, i.e., where $\mathbf{p}_i = -\mathbf{p}_f$. In the approximation of an infinitely massive nucleus the Breit and laboratory frames agree. For

a consistent description, however, results must be Lorentz boosted back into the laboratory frame, leading to additional $\mathcal{O}(q/M)$ results. This is discussed in more detail in Ref. [95] and will not be repeated here.

The kinematic recoil corrections originating from the phase space integral can also be easily written as a multiplicative factor,

$$d\Gamma \propto |\mathcal{M}|^2 \left(1 + \frac{3E_e - W_0 - 3\mathbf{p}_e \cdot \hat{\mathbf{v}}_v}{M} \right), \quad (\text{E5})$$

where the final term gives a finite contribution in the neutrino angular integral when combined with the β - ν correlation.

Combining both leads to well-known expressions for the total kinematic recoil corrections [76,128]:

$$R_N \approx 1 - \frac{2W_0}{3M} + \frac{10E_e}{3M} - \frac{2m_e^2}{3E_e M}, \quad (\text{E6})$$

where we kept only terms to first order in q/M , thereby neglecting terms of $\mathcal{O}(10^{-6})$ at most.

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