Identification of a potential ultralow-*Q***-value electron-capture decay branch in 75Se via a precise Penning trap measurement of the mass of 75As**

M. Horana Gamage \bigcirc [,](https://orcid.org/0000-0002-0031-6799)<[s](https://orcid.org/0000-0002-1787-3823)up>1,*</sup> R. Bhandari,¹ G. Bollen,^{2,3} N. D. Gamage,² A. Hamaker \bigcirc ,^{4,3} D. Puentes \bigcirc ,^{4,3} M. Redshaw,^{1,4}

R. Ringle, 2 S. Schwarz, 2 C. S. Sumithrarachchi, 2 and I. Yandow^{4,3}

¹*Department of Physics, Central Michigan University, Mount Pleasant, Michigan 48859, USA*

²*Facility for Rare Isotope Beams, East Lansing, Michigan 48824, USA*

³*Department of Physics and Astronomy, Michigan State University, East Lansing, Michigan 48824, USA*

⁴*National Superconducting Cyclotron Laboratory, East Lansing, Michigan 48824, USA*

(Received 7 July 2022; accepted 15 November 2022; published 21 December 2022)

Background: Low-energy β and electron-capture (EC) decays are important systems in neutrino mass determination experiments. An isotope with an ultralow- Q -value β decay to an excited state in the daughter with $Q_{\text{es}} < 1$ keV could provide a promising alternative candidate for future experiments. ⁷⁵Se EC decay and ⁷⁵Ge β decay represent such candidates, but a more precise determination of the mass of the common daughter, ⁷⁵As, is required to evaluate whether their potential decay branches are energetically allowed and ultralow.

Purpose: Our purpose is to perform a precise atomic mass measurement of ⁷⁵As and combine the result with the precisely known atomic masses of 75 Se and 75 Ge, along with nuclear energy level data for 75 As to evaluate potential ultralow-*Q*-value decay branches in the EC decay of ⁷⁵Se and the β decay of ⁷⁵Ge.

Method: The LEBIT Penning trap mass spectrometer at the Facility for Rare Isotope Beams was used to perform a high-precision measurement of the atomic mass of ⁷⁵As via cyclotron frequency ratio measurements of ⁷⁵As⁺ to a ${}^{12}C_6{}^+$ reference ion.

Results: The ⁷⁵As mass excess was determined to be $ME(^{75}As) = -73\,035.98(43)$ keV, from which the groundstate to ground-state Q values for ⁷⁵Se EC decay and ⁷⁵Ge β decay were determined to be 866.50(44) and 1179.01(44) keV, respectively. These results were compared to energies of excited states in ⁷⁵As at 865.4(5) and 1172.0(6) keV to determine *Q* values of 1.1(7) and 7.0(7) keV for the potential ultralow EC-decay and β-decay branches of 75 Se and 75 Ge, respectively.

Conclusion: The candidate ultralow-*Q*-value EC-decay and *β*-decay branches of ⁷⁵Se and ⁷⁵Ge to excited states in 75As, identified by Gamage *et al.* [N. D. Gamage, R. Bhandari, M. H. Gamage, R. Sandler, and M. Redshaw, [Hyperfine Interact.](https://doi.org/10.1007/s10751-019-1588-5) **240**, 43 (2019)] are seen to be energetically allowed. The ⁷⁵Se EC decay to the 865.4 -keV excited state in ⁷⁵As is potentially ultralow with $Q_{\text{es}} \approx 1$ keV. However, a more precise determination of the 865.4(5)-keV level in 75 As is required to draw a more definite conclusion about the energy of this potential decay channel.

DOI: [10.1103/PhysRevC.106.065503](https://doi.org/10.1103/PhysRevC.106.065503)

I. INTRODUCTION

The ongoing development of sensitive, high-precision experimental techniques and sophisticated theoretical calculations has continued to make nuclear β decay an important tool in modern physics, for example, in tests of the standard model (SM) and searches for physics beyond the SM $[1-3]$ and in investigations of the nature and absolute mass of the neutrino [\[4–9\]](#page-5-0).

Extreme cases of β decay, e.g., highly forbidden decays and very-low-energy decays, pose significant experimental challenges in the detection of these rare processes and theoretical challenges in describing them. Ultralow (UL)-*Q*-value β decays, in which the parent nucleus decays to an excited state in the daughter with $Q_{\text{es}} < 1$ keV, provide a particularly interesting case, because they have both experimental and theoretical implications. An isotope with an ultralow-*Q*-value β decay could potentially be used as a new candidate in direct neutrino mass determination experiments [\[10–13\]](#page-5-0). They can also provide a powerful tool to test the role of atomic interference effects in nuclear β decay [\[14,15\]](#page-5-0), which become more significant for very-low-energy decays.

To date, only one ultralow-*Q*-value β decay has been observed, that of 115 In to the $3/2^+$ first excited state in 115 Sn [\[10\]](#page-5-0). The confirmation that this decay branch is energetically allowed was made by precise Penning trap measurements of the ¹¹⁵In - ¹¹⁵Sn mass difference at Florida State University [\[16\]](#page-5-0) and with JYLFTRAP at the University of Jyväskylä [\[17\]](#page-5-0), combined with the precisely known energy of the $3/2^+$ level in 115 Sn [\[18,19\]](#page-5-0). Using the results of Refs. [\[16,19\]](#page-5-0), the resulting $Q_{\text{UL}} = 0.147(10)$ keV, making this the lowest known *Q*-value β decay. Theoretical calculations of the partial half-life for 115 In decay to 115 Sn(3/2⁺) showed a significant

^{*}Corresponding author: redsh1m@cmich.edu

discrepancy compared to experimental results [\[14,15\]](#page-5-0). To identify potential candidates for future direct neutrino mass determination experiments, and to further theoretical developments, additional ultralow-*Q*-value decays are required to be identified and to have their *Q* values and partial half-lives measured.

A large number of potential ultralow-*Q*-value decays do in fact exist—see, for example, Refs. [\[12,13,20–22\]](#page-5-0), and see Ref. [\[23\]](#page-5-0) for a recent comprehensive evaluation. The identification of these ultralow-*Q*-value decay candidates is made from a comparison of the ground-state to ground-state (gs-gs) *Q* value, *Q*gs, obtained from atomic mass data for parent and daughter isotopes [\[24\]](#page-5-0), and nuclear energy level data for the daughter [\[25\]](#page-5-0). To determine if a decay to an excited state in the daughter is energetically allowed and ultralow, both *Q*gs and the energy of the final state in the daughter, *E*∗, must be known to \ll 1 keV. While nuclear energy levels are often known to precisions of ∼10–100 eV, atomic masses are usually not, except in cases where they have been measured precisely by dedicated Penning trap experiments.

In the last few years, there has been growing interest in performing precise Penning trap measurements to determine *Q*es values for potential ultralow-*Q*-value decay candidates that have been identified. Studies have been performed on 131 Cs [\[26\]](#page-5-0) at ISOLTRAP; 135 Cs [\[27\]](#page-5-0), 72 As [\[28\]](#page-5-0), and 159 Dy [\[29\]](#page-5-0) at JYFLTRAP; 112,113 Ag and 115 Cd with the Cana-dian Penning Trap [\[30\]](#page-5-0), and 89 Sr and 139 Ba at LEBIT [\[31\]](#page-5-0). These measurements have shown that $135Cs$ could have an ultralow-*Q*-value β -decay branch and that in 159 Dy EC an experimentally confirmed $[32]$ decay to the $11/2^+$ state in the daughter has a $Q_{\text{es}} = 1.2(2)$ keV. The other candidates were ruled out.

In this paper we report on precise determinations of the Q_{gs} values for ⁷⁵Ge β decay and ⁷⁵Se EC decay via a precise measurement of the mass of the common daughter, ⁷⁵As. These candidates were first identified in Ref. [\[23\]](#page-5-0), along with 89 Sr and 139 Ba [\[31\]](#page-5-0), as cases where the mass of the stable daughter isotope was known less precisely than that of the unstable parents, making them ideal cases for study with the offline ion sources available at Penning trap facilities such as the Low Energy Beam and Ion Trap (LEBIT) located at the Facility for Rare Isotope Beams (FRIB).

Figure 1 shows the energy levels and decay schemes for the $A = 75$ Ge-As-Se triplet. The potential ultralow-*Q*-value decays are indicated by dashed blue arrows. For ⁷⁵Ge, the potential ultralow- Q -value β decay is to the 1172.0(6)-keV level in ⁷⁵As, identified as a negative-parity state with spin between 1/2 and 7/2 [\[25\]](#page-5-0). For 75Se, there are two potential ultralow-*Q*value EC-decay branches: to the 865.4(5)-keV level, identified as having $J^{\pi} = 3/2^-$ or $5/2^-$, and to the 860.4(4)-keV $1/2^+$ level. The *Q*gs values listed in Fig. 1 were calculated using data from the most recent atomic mass evaluation, AME2020 [\[24\]](#page-5-0), and are limited by the 0.88 keV/ $c²$ uncertainty in the mass of 75 As. The uncertainties in the masses of 75 Ge and ⁷⁵Se are 0.052 and 0.073 keV/ c^2 , respectively, and contribute negligibly to the uncertainty in the *Q* value. The resulting potential ultralow-*Q*-value decay energies to the respective excited states in ⁷⁵As are Q_{es}^{β} (⁷⁵Ge) = 5.2(1.1) keV

FIG. 1. Decay scheme for ⁷⁵Se EC decay and ⁷⁵Ge β decay. The main decay branches are indicated by solid black arrows with corresponding branching ratios. The potential ultralow-*Q*-value decay branches are indicated by dashed blue arrows. Spin assignments in braces indicate the range of possible values for that level. The ground-state to ground-state *Q* values are obtained using data from the AME2020 [\[24\]](#page-5-0). Energy levels and *Q* values are given in units of keV.

and $Q_{\text{es}}^{\text{EC}}({}^{75}\text{Se}) = -0.7(1.0)$ and 4.3(1.0) keV. Hence, a new Penning trap measurement of the 75 As mass is called for to check the accuracy of the *Q* values and eliminate uncertainty due to current uncertainty in the 75 As mass.

II. EXPERIMENTAL DESCRIPTION

The LEBIT facility [\[33,34\]](#page-5-0), located at FRIB was used to perform precise cyclotron frequency ratio measurements of singly charged $^{75}As^{+}$ and $^{12}C_6{}^{+}$ carbon cluster ions with a similar mass-to-charge ratio. The ${}^{75}As^+$ ions were delivered to the LEBIT facility from a batch mode ion source (BMIS) where they were produced as a by-product during normal BMIS operation. The ${}^{12}C_6{}^+$ ions were produced with a laser ablation source (LAS) [\[35\]](#page-5-0) in which a Sigradur [\[36\]](#page-5-0) glassy carbon target was installed. In either case, ions first enter the LEBIT cooler and buncher [\[37\]](#page-5-0), in which they are thermalized by a helium buffer gas, confined in a linear radio-frequency (rf) Paul trap, and then released as low-emittance pulses ∼100-ns in length. Ions are then directed toward the Penning trap, which has a hyperbolic structure and is housed in a 9.4-T magnetic field. On the way to the Penning trap, ions encounter a fast electrostatic kicker at the magnetic field's entrance, which can be used to block the beam and allow only ions of a specific A/q to pass based on their time of flight [\[33\]](#page-5-0).

In the Penning trap, ions undergo three normal modes of motion: reduced-cylotron, magnetron, and axial, with frequencies f_+ , f_- , and f_z , respectively. The three normal mode frequencies can be combined to determine the free-space cyclotron frequency f_c . Pertinent to the work described here is the relationship

$$
f_c = f_+ + f_- = \frac{qB}{2\pi m},
$$
 (1)

FIG. 2. A 500-ms cyclotron frequency resonance curve for ${}^{75}As^+$ with a 100-300-100 ms Ramsey timing sequence. The solid red line is a fit to the theoretical line shape [\[44\]](#page-5-0).

where q/m is the charge-to-mass ratio of the ion, and *B* is the magnetic field strength. At the LEBIT facility, the cyclotron frequency of ions in the Penning trap is determined using the time-of-flight ion cyclotron resonance (TOF-ICR) technique [\[38,39\]](#page-5-0). Ions are parked on an initial magnetron orbit by a Lorentz steerer [\[40\]](#page-5-0) and are then driven with a quadrupolar rf pulse at a frequency *f*rf close to the free-space cyclotron frequency of the ion. Depending on the frequency difference $| f_{\text{rf}} - f_c |$, magnetron motion can be partly or fully converted to cyclotron motion. The ions are then released from the trap and impinge on a microchannel plate (MCP) detector that is used to determine the time-of-flight from the trap to the MCP, as well as the number of ions that were in the trap. An ion's time-of-flight to the MCP is minimized upon full conversion of magnetron to cyclotron motion. Hence, *f*rf is varied around *fc* and a characteristic TOF resonance curve can be mapped out.

In this work, we used the Ramsey TOF-ICR technique [\[41,42\]](#page-5-0), in which two time-separated $\pi/2$ pulses of duration T_{on} are applied on either side of a wait period, T_{off} , where no rf is applied. This scheme replaces the single π -pulse used in the traditional TOF-ICR technique to couple the magnetron and cyclotron modes. The Ramsey technique results in an interference-type pattern in the TOF resonance and has been shown to enable a determination of the cyclotron frequency to at least 3 times higher precision in the same measurement time compared to the traditional TOF-ICR technique [\[43\]](#page-5-0). We used Ramsey excitation schemes with either 500 or 1000 ms total measurement time, with a $T_{on} - T_{off} - T_{on}$ sequence of 100-300-100 ms or 200-600-200 ms, respectively. An example 500-ms Ramsey TOF-ICR resonance for ${}^{75}As^+$ from this work is shown in Fig. 2.

A single resonance like the one shown in Fig. 2 took \sim 30 min, consisted of \sim 2500–5000 ions, and enabled a determination of the ⁷⁵As⁺ or ¹²C₆⁺ frequency to ~15 mHz. To determine the cyclotron frequency ratio of ${}^{12}C_6{}^+$ to ${}^{75}As^+$ and to account for temporal magnetic field drifts, we alternated between measurements with the two ions, performing 65 measurement pairs in total.

FIG. 3. Individual cyclotron frequency ratio measurements of $f_c(^{75}As^+)/ f_c^{\text{int}}(^{12}C_6^+)$ (black squares). The red band shows the uncertainty in the average ratio, without systematic corrections and associated uncertainties applied. The blue line and band show the ratio obtained using the 75 As mass from AME2020 [\[24\]](#page-5-0) and the associated uncertainty.

Deviations from an ideal Penning trap, i.e., a magnetic field that is not perfectly uniform and an electric field that is not purely quadrupolar, combined with the finite normal mode amplitudes of ions in the trap, result in shifts to the measured frequencies. The preparation of ions of different *m*/*q* in the Penning trap can result in them having different normal mode amplitudes and hence different systematic frequency shifts. While these shifts largely cancel in the cyclotron frequency ratio, when comparing ions of different nominal *m*/*q* a residual shift to the cyclotron frequency ratio can remain. In this work, our cyclotron frequency ratio involved ${}^{12}C_6{}^+$ and $75As⁺$, which differ by three atomic mass units. To evaluate the potential shift to the ratio, we took additional cyclotron frequency ratio data for ${}^{12}C_6$ ⁺ vs ${}^{85,87}Rb$ ⁺ and ${}^{12}C_7$ ⁺. The uncertainty in the mass of the carbon cluster ions and the rubidium ions (the masses of which have been determined previously with the Florida State University Penning trap [\[45\]](#page-5-0)) are negligible compared to the statistical uncertainties in our measurements.

III. DATA AND ANALYSIS

To evaluate the cyclotron frequency ratio from our alternating measurements of ${}^{12}C_6{}^+$ and ${}^{75}As^+$, pairs of ${}^{12}C_6{}^+$ cyclotron frequency measurements were first linearly interpolated to obtain the corresponding $f_c^{\text{int}}(^{12}C_6^+)$ value at the time of the interleaved $f_c(^{75}As^+)$ measurement. This procedure eliminated potential systematic shifts to the cyclotron frequency ratio due to linear magnetic field drifts. A series of cyclotron frequency ratios, corresponding to inverse mass ratios of the two ions were found from

$$
R = \frac{f_c(^{75}\text{As}^+)}{f_c^{\text{int}}(^{12}\text{C}_6^+)} = \frac{m(^{12}\text{C}_6^+)}{m(^{75}\text{As}^+)}.
$$
 (2)

The individual measurements of *R* were then used to find the weighted average \overline{R} , the associated statistical uncertainty, and the Birge ratio [\[46\]](#page-5-0), as shown in Fig. 3. For the ${}^{12}C_6{}^+$ vs ⁷⁵As⁺ data set, consisting of 65 individual ratio measurements, the Birge ratio was found to be 1.07(6), indicating that the statistical uncertainties in the individual measurements are a reasonable estimate of the statistical spread of the data. Nevertheless, we conservatively inflated the statistical uncertainty in the weighted-average ratio by the Birge ratio.

Our auxiliary measurements of ${}^{12}C_6{}^+$ vs ${}^{85,87}Rb^+$ and ${}^{12}C_7$ ⁺ did show a systematic shift in the ratio when comparing ions of different nominal m/q of $-0.57(18) \times 10^{-9}$ per u/*e*. This result was used to apply a correction of $1.7(5) \times 10^{-9}$ to the ${}^{12}C_6{}^+$ vs ${}^{75}As^+$ ratio. We also performed a Z-class analysis to evaluate potential shifts to the cyclotron frequency from the Coulomb interaction when there was more than one ion in the trap [\[47\]](#page-5-0). This analysis showed a linear trend, but with a slope of $0.52(51) \times 10^{-9}$ per ion in the trap, consistent with no shift. In our data analysis, we limited the number of ions in the trap to between one and five. Assuming an average of three ions in the trap, a correction of $-1.6(1.5) \times 10^{-9}$ should be applied to the data. We applied half this value as a correction with an uncertainty equal to the applied correction.

In the course of the analysis we uncovered an additional systematic effect: when ions arrive at the MCP, they create a time-of-flight (TOF) distribution, the width of which depends on the energy difference between ions that are resonant with the rf quadrupole drive in the trap (i.e., magnetron motion is fully converted to cyclotron motion) and those that are not (i.e., no conversion). There is also some spread due to the variation in the initial conditions of ions in the trap. In addition, there can be background noise or ion events in the time window, including a small number of ions that are produced via the breakup of carbon cluster ions in the trap, and so a gate is placed around the TOF distribution of the ions of interest.¹ For ⁷⁵As⁺ ions, a systematic variation in the cyclotron frequency was observed as a function of location of the high-end cutoff time of the gate, i.e., on where the cutoff was made for ions with a longer TOF, which corresponds to nonresonant ions. This systematic effect was not observed for ${}^{12}C_6{}^+$ or ${}^{12}C_7{}^+$ ions produced with the LAS or ${}^{85,87}Rb^+$ ions produced with a surface ionization source. Hence, we concluded that there could be contaminant ions in the beam arriving from the BMIS that were not identified, and which could cause systematic shifts to the ${}^{75}As^+$ ions' frequency via ion-ion interactions.

To account for this effect, we systematically varied the high-end cutoff time of the gate and recorded the cyclotron frequency and χ^2 of the fit averaged over all 65 measurements. We found that χ^2 varied quadratically around its minimum—when more ion events from the longer TOF tail of the distribution were included, the fit was statistically worse, and, likewise, when the cut was made further into the TOF distribution, more data were discarded and the fit was again worse. We used the gate-setting corresponding to the minimum χ^2 (obtained from a fit of χ^2 vs gate width) in our final analysis. The cyclotron frequency we obtained as a function of the gate width also varied around a minimum value. The minimum f_c , however, did not occur at the minimum χ^2 . The

TABLE I. Error budget for ⁷⁵As⁺ vs ¹²C₆⁺ cyclotron frequency ratio measurement. $\sigma_{st}(\times BR)$ is the statistical uncertainty (inflated by the Birge ratio, BR = 1.07). ΔR_{ion} , ΔR_{SM} , and ΔR_{gate} are the systematic corrections (and corresponding uncertainties in parentheses) to account for frequency shifts due to there being more than one ion in the trap, the effect of comparing ions of dissimilar m/q , and the effect associated with how the gate on the TOF distribution was set (see text for details). \bar{R}' is the corrected ratio with statistical, systematic, and total uncertainties shown respectively in parentheses.

Source	Contribution to		
	Ratio $(\times 10^{-9})$	$M({}^{75}\text{As})$ (keV/ c^2)	
$\sigma_{\rm st}$	1.5	0.11	
$\sigma_{st} \times BR$	1.6	0.12	
$\Delta R_{\rm ion}$	$-0.8(8)$	0.06(6)	
$\Delta R_{\delta M}$	1.7(5)	$-0.12(4)$	
$\Delta R_{\rm gate}$	0.0(5.7)	0.0(0.4)	
Final ratio		\overline{R} ' = 0.961 004 378 4(16)(58)(60)	

minimum f_c was 0.007 Hz lower, and the maximum f_c was 0.015 Hz higher than our nominal value at χ^2 min. These correspond to shifts in the $^{75}As^{+}/^{12}C_{6}^{+}$ ratio of 3.7 and 7.7×10^{-9} , respectively. We used the average of these two values, 5.7×10^{-9} , to conservatively estimate the additional uncertainty to include in our final ratio to account for this systematic effect. This corresponds to an uncertainty in the mass of ⁷⁵As of about 0.4 keV/ c^2 . A summary of our ratio measurement and the systematic corrections and uncertainties applied is shown in Table I.

After accounting for systematic effects, the corrected average ratio \bar{R} ^{\prime} was used to determine the atomic mass of ⁷⁵As in atomic mass units using the mass and binding energy of the carbon cluster reference molecule and accounting for the mass of the missing electrons, *me*:

$$
M({}^{75}\text{As}) = [6M({}^{12}\text{C}) - m_e + \text{BE}({}^{12}\text{C}_6)]\bar{R}'^{-1} + m_e. \quad (3)
$$

The first ionization energies of ⁷⁵As and ¹²C₆ are both \approx eV, so the difference is $< 1 \text{ eV}$ and is negligible. The ¹²C₆ molecular binding energy, on the other hand, is -31.7 eV for the most stable cyclic structure [\[48\]](#page-6-0) and therefore is included in Eq. (3). The mass excess was then calculated from

$$
ME(^{75}As) = [M(^{75}As) - 75] \times F,
$$
 (4)

where the conversion factor $F = 931\,494.102\,42(28)$ $(\text{keV}/c^2)/\text{u}$ from Ref. [\[49\]](#page-6-0) was used.

IV. RESULTS AND DISCUSSION

Our final mass excess value for 75 As is given in Table [II](#page-4-0) and compared with the value from the AME2020 [\[24\]](#page-5-0). Our result is 1.78(98) keV/ c^2 lower than the AME2020 value and

 1 If the lighter ions resulting from the breakup are not removed from the analysis, they produce a shift of less than 0.5 ppb to the ${}^{12}C_6{}^+$ frequency.

TABLE II. Mass excess for 75 As obtained in the work along with the result from the AME2020 [\[24\]](#page-5-0) and the difference $\Delta ME =$ $ME_{LEBIT} - ME_{AME}$

Isotope	Ref.	This work (keV/c^2)	AME2020 (keV/c^2)	AME. (keV/c^2)
75 As	${}^{12}C_6$	$-73035.98(43)$	$-73034.20(88)$	$-1.78(98)$

a factor of 2 more precise. This result indicates that 75 As is more strongly bound than suggested by the AME2020 value, thus increasing the Q_{gs} values of ⁷⁵Se and ⁷⁵Ge, which can be found via

$$
Q_{\rm gs}^{\rm EC}(^{75}{\rm Se}) = [{\rm ME}(^{75}{\rm Se}) - {\rm ME}(^{75}{\rm As})]c^2 \tag{5}
$$

and

$$
Q_{\rm gs}^{\beta}({}^{75}\text{Ge}) = [\text{ME}({}^{75}\text{Ge}) - \text{ME}({}^{75}\text{As})]c^2.
$$
 (6)

Our calculated Q_{gs}^{EC} and Q_{gs}^{β} values for ⁷⁵Se and ⁷⁵Ge, obtained using our new value for $M(^{75}As)$ and the mass of the ⁷⁵Se and ⁷⁵Ge parent atoms from AME2020 [\[24\]](#page-5-0), are given in Table III.

The *Q* value for the decay to an excited state in the daughter nuclide with energy E^* can then be found via

$$
Q_{\rm es}^{\beta/\rm EC} = Q_{\rm gs}^{\beta/\rm EC} - E^*.
$$
 (7)

For the case of 75 Se EC decay, two potential excited states in ⁷⁵As with energies close to $Q_{\rm gs}^{\rm EC}(^{75}{\rm Se})$ were identified: the 860.0(4)-keV $1/2^+$ level, and the 865.4(5)-keV $3/2^+$ or $5/2^+$ level. For ⁷⁵Ge β decay, the 1172.0(6)-keV *J*[−] state with *J* = $1/2$, $3/2$, $5/2$, or $7/2$ was identified as a potential daughter state with a low-energy-decay Q value. Our Q_{gs} values are compared with the relevant E^* values in Table III. Our results indicate that all three decays listed are energetically allowed at least at the 1.5 σ level. The ⁷⁵Ge β decay to the 1172-keV state in ⁷⁵As is ruled out as not being ultralow (i.e., $Q_{\text{es}}^{\beta} > 1$ keV) at the 9σ level. Likewise, the EC decay of ⁷⁵Se to the 860-keV level in ⁷⁵As is ruled out as an ultralow decay branch, at the 8σ level. However, the EC decay of ⁷⁵Se to the 865-keV level in ⁷⁵As has $Q_{\text{es}}^{\text{EC}} \approx 1$ keV, making it a potential candidate for an ultralow-*Q*-value decay. The uncertainty in the $Q_{\text{es}}^{\text{EC}}$ value listed in Table III is predominantly $(60\% \text{ contribution})$ due to the 0.5-keV uncertainty in the 865-keV level, with the remaining 40% contribution due to the 0.4-keV uncertainty in our $M(^{75}As)$ value. Hence, a more precise determination of both the $Q_{gs}^{EC}(^{75}Se)$ value and the $E^*(865.4\text{-}keV)$ level energy is required to further evaluate this potential ultralow-*Q*-value EC decay.

TABLE III. *Q* values based on the mass measurement listed in Table II and Eqs. (5) and (6). The column E^* gives the energy of the excited state of the daughter nucleus that the potential ultralow-*Q*-value decay would go to. The result for the *Q* value of the decay branch to the excited state, *Q*es, is calculated from Eq. (7).

Parent	Decay	$Q_{\rm gs}$ (keV)	E* (keV)	$Q_{\rm es}$ (keV)
75 Se	EC.	866.50(44)	865.4(5)	1.10(67)
75 Se	EС	866.50(44)	860.4(4)	6.10(60)
75 Ge	ß	1179.01(44)	1172.0(6)	7.01(74)

In fact, a recent measurement reported by the JYFLTRAP group does provide a more precise mass for 75 As [\[50\]](#page-6-0). Our result for $ME(^{75}As)$ agrees with the JYLFTRAP result, $ME(^{75}As) = -73\,035.521(42) \text{ keV}/c^2$, which is a factor of 10 more precise, at the 1σ level. The JYFLTRAP result also indicates that 75 Se EC decay to the 865.4-keV level in 75 As is energetically allowed with a potentially ultralow-*Q* value, $Q_{\text{es}} \lesssim 1$ keV, but that a more precise determination of the 865.4-keV level is necessary.

V. CONCLUSION

Using Penning trap mass spectrometry, we have performed a precise measurement of the mass of 75As. Our result reduces the uncertainty in the mass by about a factor of 2 compared to the value adopted in the most recent atomic mass evaluation, AME2020 [\[24\]](#page-5-0), and shows a shift of–1.78 keV/ c^2 . Using this new mass and the more precisely known masses of 75 Se and 75Ge from AME2020, we were able to determine new ground-state to ground-state *Q* values for the EC decay of ⁷⁵Se and the β decay of ⁷⁵Ge. These Q values were combined with excited-state energies in ⁷⁵As to determine the Q_{es} values of potential ultralow-energy decays of 75 Se and 75 Ge. Our results indicate that the β decay of ⁷⁵Ge to the 1172.0(6)keV, negative-parity state with angular momentum between 1/2 and 7/2 is energetically allowed, but with a *Q* value of 7.0(7) keV, ruling it out as a potential ultralow-*Q*-value decay. Likewise, the EC decays of 75 Se to the 865.4(5)-keV $3/2^-$ or $5/2^-$ level and to the 860.0(4)-keV $1/2^+$ level are also energetically allowed with *Q* values of 1.1(7) and 6.1(6) keV, respectively. Our result is in agreement with a recently reported measurement by the JYFLTRAP group [\[50\]](#page-6-0). Both results indicate that the EC decay of 75 Se to the 865-keV level in ⁷⁵As with a $Q_{es} \lesssim 1$ keV is of interest as a potential ultralow-*Q*-value decay candidate and that, to further evaluate this potential decay branch, a more precise determination of the energy of the 865-keV level is required.

ACKNOWLEDGMENTS

This research was supported by Central Michigan University, Michigan State University and the Facility for Rare Isotope Beams, and the National Science Foundation under Contracts No. PHY-1565546, No. PHY-2111185, and No. PHY-2111302.

- [1] [J. C. Hardy and I. S. Towner,](https://doi.org/10.1103/PhysRevC.91.025501) Phys. Rev. C **91**, 025501 (2015).
- [2] [N. Severijns, M. Beck, and O. Naviliat-Cuncic,](https://doi.org/10.1103/RevModPhys.78.991) Rev. Mod. Phys. **78**, 991 (2006).
- [3] M. T. Burkey, G. Savard, A. T. Gallant, N. D. Scielzo, J. A. Clark, T. Y. Hirsh, L. Varriano, G. H. Sargsyan, K. D. Launey, M. Brodeur, D. P. Burdette, E. Heckmaier, K. Joerres, J. W. Klimes, K. Kolos, A. Laminack, K. G. Leach, A. F. [Levand, B. Longfellow, B. Maaß](https://doi.org/10.1103/PhysRevLett.128.202502) *et al.*, Phys. Rev. Lett. **128**, 202502 (2022).
- [4] [F. T. Avignone, S. R. Elliott, and J. Engel,](https://doi.org/10.1103/RevModPhys.80.481) Rev. Mod. Phys. **80**, 481 (2008).
- [5] S. Friedrich, G. B. Kim, C. Bray, R. Cantor, J. Dilling, S. Fretwell, J. A. Hall, A. Lennarz, V. Lordi, P. Machule, D. McKeen, X. Mougeot, F. Ponce, C. Ruiz, A. Samanta, W. K. [Warburton, and K. G. Leach,](https://doi.org/10.1103/PhysRevLett.126.021803) Phys. Rev. Lett. **126**, 021803 (2021).
- [6] K. C. Aker *et al.* [\(KATRIN Collaboration\),](https://doi.org/10.1103/PhysRevLett.123.221802) Phys. Rev. Lett. **123**, 221802 (2019).
- [7] [B. Monreal and J. A. Formaggio,](https://doi.org/10.1103/PhysRevD.80.051301) Phys. Rev. D **80**, 051301(R) (2009) .
- [8] E. C. Gastaldo et al. [\(ECHo Collaboration\),](https://doi.org/10.1140/epjst/e2017-70071-y) Eur. Phys. J.: Spec. Top. **226**, 1623 (2017).
- [9] M. Faverzani *et al.* [\(HOLMES Collaboration\),](https://doi.org/10.1007/s10909-016-1540-x) J. Low Temp. Phys. **184**, 922 (2016).
- [10] C. M. Cattadori, M. D. Deo, M. Laubenstein, L. Pandola, and V. I. Tretyak, [Nucl. Phys. A](https://doi.org/10.1016/j.nuclphysa.2004.10.025) **748**, 333 (2005).
- [11] C. M. Cattadori, M. De Deo, M. Laubenstein, L. Pandola, and V. I. Tretyak, [Phys. At. Nucl.](https://doi.org/10.1134/S1063778807010140) **70**, 127 (2007).
- [12] J. Suhonen, Phys. Scr. **89**[, 054032 \(2014\).](https://doi.org/10.1088/0031-8949/89/5/054032)
- [13] J. Kopp and A. Merle, Phys. Rev. C **81**[, 045501 \(2010\).](https://doi.org/10.1103/PhysRevC.81.045501)
- [14] [M. T. Mustonen and J. Suhonen,](https://doi.org/10.1088/0954-3899/37/6/064008) J. Phys. G: Nucl. Part. Phys. **37**, 064008 (2010).
- [15] [J. Suhonen and M. Mustonen,](https://doi.org/10.1016/j.ppnp.2009.12.018) Prog. Part. Nucl. Phys, **64**, 235 (2010) .
- [16] [B. J. Mount, M. Redshaw, and E. G. Myers,](https://doi.org/10.1103/PhysRevLett.103.122502) Phys. Rev. Lett. **103**, 122502 (2009).
- [17] J. S. E. Wieslander, J. Suhonen, T. Eronen, M. Hult, V.- V. Elomaa, A. Jokinen, G. Marissens, M. Misiaszek, M. T. [Mustonen, S. Rahaman, C. Weber, and J. Äystö,](https://doi.org/10.1103/PhysRevLett.103.122501) Phys. Rev. Lett. **103**, 122501 (2009).
- [18] J. Blachot, [Nucl. Data Sheets](https://doi.org/10.1016/j.nds.2005.03.003) **104**, 967 (2005).
- [19] V. A. Zheltonozhsky, A. M. Savrasov, N. V. Strilchuk, and V. I. Tretyak, [Europhys. Lett.](https://doi.org/10.1209/0295-5075/121/12001) **121**, 12001 (2018).
- [20] [M. T. Mustonen and J. Suhonen,](https://doi.org/10.1063/1.3527233) AIP Conf. Proc. **1304**, 401 (2010).
- [21] [M. Haaranen and J. Suhonen,](https://doi.org/10.1140/epja/i2013-13093-8) Eur. Phys. J. A **49**, 93 (2013).
- [22] [M. Mustonen and J. Suhonen,](https://doi.org/10.1016/j.physletb.2011.07.088) Phys. Lett. B **703**, 370 (2011).
- [23] N. D. Gamage, R. Bhandari, M. H. Gamage, R. Sandler, and M. Redshaw, [Hyperfine Interact.](https://doi.org/10.1007/s10751-019-1588-5) **240**, 43 (2019).
- [24] [M. Wang, W. Huang, F. Kondev, G. Audi, and S. Naimi,](https://doi.org/10.1088/1674-1137/abddaf) Chin. Phys. C **45**, 030003 (2021).
- [25] National nuclear data center, [https://www.nndc.bnl.gov/.](https://www.nndc.bnl.gov/)
- [26] J. Karthein, D. Atanasov, K. Blaum, S. Eliseev, P. Filianin, D. Lunney, V. Manea, M. Mougeot, D. Neidherr, Y. Novikov, L. [Schweikhard, A. Welker, F. Wienholtz, and K. Zuber,](https://doi.org/10.1007/s10751-019-1601-z) Hyperfine Interact. **240**, 61 (2019).
- [27] A. de Roubin, J. Kostensalo, T. Eronen, L. Canete, R. P. de Groote, A. Jokinen, A. Kankainen, D. A. Nesterenko, I. D.

[Moore, S. Rinta-Antila, J. Suhonen, and M. Vilén,](https://doi.org/10.1103/PhysRevLett.124.222503) Phys. Rev. Lett. **124**, 222503 (2020).

- [28] Z. Ge, T. Eronen, A. de Roubin, D. A. Nesterenko, M. Hukkanen, O. Beliuskina, R. de Groote, S. Geldhof, W. Gins, A. Kankainen, A. Koszorús, J. Kotila, J. Kostensalo, I. D. Moore, A. Raggio, S. Rinta-Antila, J. Suhonen, V. Virtanen, [A. P. Weaver, A. Zadvornaya](https://doi.org/10.1103/PhysRevC.103.065502) *et al.*, Phys. Rev. C **103**, 065502 (2021).
- [29] Z. Ge, T. Eronen, K. S. Tyrin, J. Kotila, J. Kostensalo, D. A. Nesterenko, O. Beliuskina, R. de Groote, A. de Roubin, S. Geldhof, W. Gins, M. Hukkanen, A. Jokinen, A. Kankainen, A. Koszorús, M. I. Krivoruchenko, S. Kujanpää, I. D. Moore, [A. Raggio, S. Rinta-Antila](https://doi.org/10.1103/PhysRevLett.127.272301) *et al.*, Phys. Rev. Lett. **127**, 272301 (2021).
- [30] N. D. Gamage, R. Sandler, F. Buchinger, J. A. Clark, D. Ray, R. Orford, W. S. Porter, M. Redshaw, G. Savard, K. S. [Sharma, and A. A. Valverde,](https://doi.org/10.1103/PhysRevC.106.045503) Phys. Rev. C **106**, 045503 (2022).
- [31] R. Sandler, G. Bollen, N. D. Gamage, A. Hamaker, C. Izzo, D. [Puentes, M. Redshaw, R. Ringle, and I. Yandow,](https://doi.org/10.1103/PhysRevC.100.024309) Phys. Rev. C **100**, 024309 (2019).
- [32] B. Myslek, Z. Sujkowski, and B. Kotlinska, *Proceedings of the Conference on the Electron Capture and Higher Order Processes in Nuclear Decays, Debrecen, Hungary, 1968* (Eoetvoe, Budapest, 1969), Vol. 1, pp. 102–107.
- [33] R. Ringle, G. Bollen, A. Prinke, J. Savory, P. Schury, S. Schwarz, and T. Sun, [Nucl. Instrum. Methods Phys. Res., Sect.](https://doi.org/10.1016/j.nima.2009.03.207) A **604**, 536 (2009).
- [34] [R. Ringle, S. Schwarz, and G. Bollen,](https://doi.org/10.1016/j.ijms.2013.04.001) Int. J. Mass Spectrom. **349-350**, 87 (2013).
- [35] C. Izzo, G. Bollen, S. Bustabad, M. Eibach, K. Gulyuz, D. J. Morrissey, M. Redshaw, R. Ringle, R. Sandler, S. Schwarz, and A. A. Valverde, [Nucl. Instrum. Methods Phys. Res., Sect. B](https://doi.org/10.1016/j.nimb.2015.12.019) **376**, 60 (2016).
- [36] HTW Sigradur, [http://www.htw-germany.com.](http://www.htw-germany.com)
- [37] S. Schwarz, G. Bollen, R. Ringle, J. Savory, and P. Schury, [Nucl. Instrum. Methods Phys. Res., Sect. A](https://doi.org/10.1016/j.nima.2016.01.078) **816**, 131 (2016).
- [38] [G. Gröff, H. Kalinowsky, and J. Traut,](https://doi.org/10.1007/BF01414243) Z. Phys. A **297**, 35 (1980).
- [39] [M. König, G. Bollen, H.-J. Kluge, T. Otto, and J. Szerypo,](https://doi.org/10.1016/0168-1176(95)04146-C) Int. J. Mass Spectrom. Ion Processes **142**, 95 (1995).
- [40] R. Ringle, G. Bollen, A. Prinke, J. Savory, P. Schury, S. Schwarz, and T. Sun, [Int. J. Mass Spectrom.](https://doi.org/10.1016/j.ijms.2006.12.008) **263**, 38 (2007).
- [41] G. Bollen, H.-J. Kluge, T. Otto, G. Savard, and H. Stolzenberg, [Nucl. Instrum. Methods Phys. Res., Sect. B](https://doi.org/10.1016/0168-583X(92)95971-S) **70**, 490 (1992).
- [42] S. George, K. Blaum, F. Herfurth, A. Herlert, M. Kretzschmar, [S. Nagy, S. Schwarz, L. Schweikhard, and C. Yazidjian,](https://doi.org/10.1016/j.ijms.2007.04.003) Int. J. Mass Spectrom. **264**, 110 (2007).
- [43] S. George, S. Baruah, B. Blank, K. Blaum, M. Breitenfeldt, U. Hager, F. Herfurth, A. Herlert, A. Kellerbauer, H.-J. Kluge, M. Kretzschmar, D. Lunney, R. Savreux, S. Schwarz, L. [Schweikhard, and C. Yazidjian,](https://doi.org/10.1103/PhysRevLett.98.162501) Phys. Rev. Lett. **98**, 162501 (2007).
- [44] M. Kretzschmar, [Int. J. Mass Spectrom.](https://doi.org/10.1016/j.ijms.2007.04.002) **264**, 122 (2007).
- [45] [B. J. Mount, M. Redshaw, and E. G. Myers,](https://doi.org/10.1103/PhysRevA.82.042513) Phys. Rev. A **82**, 042513 (2010).
- [46] R. T. Birge, Phys. Rev. **40**[, 207 \(1932\).](https://doi.org/10.1103/PhysRev.40.207)
- [47] A. Kellerbauer, K. Blaum, G. Bollen, F. Herfurth, H. J. Kluge, M. Kuckein, E. Sauvan, C. Scheidenberger, and L. Schweikhard, [Eur. Phys. J. D](https://doi.org/10.1140/epjd/e2002-00222-0) **22**, 53 (2003).
- [48] [D. P. Kosimov, A. A. Dzhurakhalov, and F. M. Peeters,](https://doi.org/10.1103/PhysRevB.78.235433) Phys. Rev. B **78**, 235433 (2008).
- [49] [E. Tiesinga, P. J. Mohr, D. B. Newell, and B. N. Taylor,](https://doi.org/10.1063/5.0064853) J. Phys. Chem. Ref. Data **50**, 033105 (2021).
- [50] M. Ramalho, Z. Ge, T. Eronen, D. A. Nesterenko, J. Jaatinen, A. Jokinen, A. Kankainen, J. Kostensalo, J. Kotila, M. I. [Krivoruchenko, J. Suhonen, K. S. Tyrin, and V. Virtanen,](https://doi.org/10.1103/PhysRevC.106.015501) Phys. Rev. C **106**, 015501 (2022).