# Nuclear states projected from a pair condensate

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(Received 15 January 2022; accepted 4 March 2022; published 17 March 2022)

Atomic nuclei exhibit deformation, pairing correlations, and rotational symmetries. To meet these competing demands in a computationally tractable formalism, we revisit the use of general pair condensates with good particle number as a trial wave function for even-even nuclei. After minimizing the energy of the condensate, allowing for general triaxial deformations, we project out states with good angular momentum with a fast projection technique. To show applicability, we present example calculations from pair condensates in several model spaces and compare against angular-momentum projected Hartree-Fock and full configuration-interaction shell-model calculations. This approach successfully generates spherical, vibrational, and rotational spectra, demonstrating potential for modeling medium- to heavy-mass nuclei.

DOI: 10.1103/PhysRevC.105.034317

# I. INTRODUCTION

Medium- and heavy-mass nuclides exhibit a rich portfolio of behavior, not only the classic rotational, vibrational, and superfluid (spherical) spectra [1,2], but also triaxial deformation and wobbling [3] and chiral doublet bands [4]. An accurate description of the many-body wave functions of heavy nuclides is important to understanding alpha cluster preformation in  $\alpha$  decay [5] as well as nucleosynthesis in astrophysical environments [6,7].

There are numerous microscopic approaches to modeling heavy nuclei, including the relativistic mean-field method [8], the Hartree-Fock Bogoliubov method [2,9–16], the random phase and related approximations [2], and algebraic models (e.g., the interacting boson model [17]), as well as the projected shell model [18] and powerful Monte Carlo methods [19].

Here we focus on the configuration-interacting shell model, which has been successful in reproducing low-lying states and transition rates of light and medium-heavy nuclei [20,21]. Within a given valence space and with a good effective interactions, the configuration-interaction shell model reproduces large amounts of data. The downside of configuration-interaction methods, however, is they are at the mercy of the valence space. While any calculation in the 1s0d shell is nowadays nearly trivial [22] and much of the

1p0f shell is accessible [23,24], as one reaches up to the  $2s1d0g_{7/2}0h_{11/2}$  space (nuclides with  $N, Z \in [50, 82]$ ), calculations quickly become intractable due to the large dimension of the many-body basis.

Among many different configuration-interaction truncation schemes, one is to construct the basis from products of a restricted set of fermion pairs coupled to angular momentum  $0, 2, 4, 6, \dots$  (S, D, G, I pairs). In low-lying states, such pairs are favored energetically. An example of this is the general seniority model, which focuses on S-pair condensates and their breaking [25–28] and is especially useful in semimagic nuclei. Going further, the nucleon-pair approximation (NPA) truncates the configuration space with angular-momentum coupled fermion pairs and has proved to be effective in medium-heavy and heavy nuclei with schematic pairingplus-quadrupole-quadrupole interactions [29,30]. In the NPA matrix elements and overlaps between many-pair states with good total angular momentum (i.e., J scheme) are computed using commutation relations, which become numerically burdensome as the number of pairs increase, in practice limiting calculations to a maximum of about eight valence pairs. To address this burden, M-scheme NPA codes have been developed recently, pushing the upper limit of valence particles in the NPA to  $Z_{\text{valence}}$ ,  $N_{\text{valence}} \leq 12$  [31,32].

An alternate is to replace fermion pairs by bosons. The interacting boson model (IBM) [17] typically starts from S and D bosons and constructs Lie group reduction chains, classifying nuclei into different group limits. There have been efforts to map shell-model calculations to boson models or

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pair models [33–35]. While bosons models have the advantage of simpler commutation relations than fermion pairs, rigorous mapping of fermion pairs into bosons is far from trivial.

Related to the challenge of mapping fermion pairs to bosons is the problem of choosing the structure of the fermion pairs. Pairs determined from the variational principle, either from varying pairs by direct conjugate gradient descent, or extracted from a deformed Hartree-Fock (i.e., single-particle variation) state, can produce good quality spectra and transitions of transitional and deformed nuclei [36–38]. In other words, it can be useful to combine a variational approach (for optimization) and pair truncations (for small dimension).

Variational methods start from a given trial wave function (or Ansatz) and minimizes the expectation value of the given Hamiltonian. Hartree-Fock calculations use a Slater determinant, an antisymmetrized product of single-particle states, as the ansatz, and varies the single-particle states to minimize the energy. The Hartree-Fock-Bogoliubov scheme utilizes a quasiparticle (Bardeen-Cooper-Schrieffer or BCS) vacuum constructed from a canonical basis as the trial wave function, and the canonical basis as well as BCS occupancies are varied [2]. Slater determinants used in Hartree-Fock calculations have the advantage of fast evaluation of Hamiltonian matrix elements, while the HFB scheme benefits from inherent simplicity of the quasiparticle vacuum.

Exploration of Ansätze is important and can lead to new approaches [39]. In this paper we focus on a pair condensate Ansatz, well known in quantum chemistry as "antisymmetrized geminal power" wave functions [40],

$$(\hat{A}^{\dagger})^{n}|0\rangle, \quad \hat{A}^{\dagger} = \frac{1}{2} \sum_{\alpha\beta} A_{\alpha\beta} \hat{c}^{\dagger}_{\alpha} \hat{c}^{\dagger}_{\beta},$$
(1)

where  $|0\rangle$  is the bare fermion vacuum, and  $\hat{c}^{\dagger}_{\alpha}$ ,  $\hat{c}^{\dagger}_{\beta}$  are fermion creation operators with  $\alpha$ ,  $\beta$  labeling single-particle states [i.e.,  $\alpha$  stands for  $(n_{\alpha}l_{\alpha}j_{\alpha}m_{\alpha})$ ], and, finally,  $A_{\alpha\beta}$  are the "pair structure coefficients." This is also called a "collective" pair, although sometimes that term is restricted to pairs with good angular momentum; we make no such restriction here. When we find a variational minimum, we restrict ourselves to real pair structure coefficients with skew symmetry  $A_{\alpha\beta} = -A_{\beta\alpha}$ ; however, later, under rotation we will have complex skewsymmetric structure coefficients. With a proper choice of  $A_{\alpha\beta}$ , such a condensate wave can be a Slater determinant as in Hartree-Fock theory [2,36], or a seniority-0 state in Talmi's generalized seniority model [25], or particle-numberprojected component of a BCS or HFB vacuum with particle number 2n [41]. Therefore the pair condensate is a more general Ansatz than any one of those limits.

The pair condensate Ansatz has not been widely used in nuclear structure theory, however, because the evaluation of Hamiltonian matrix elements to be varied can be complicated [38,41,42], and the angular-momentum projection from a pair condensate has not been previously implemented for nuclear structure. In recent work [32], we rederived formulas (see Refs. [34,35] and references therein) for general *M*-scheme nucleon-pair bases,

$$\hat{A}_1^{\dagger} \cdots \hat{A}_n^{\dagger} |0\rangle, \qquad (2)$$

with the pair condensate (1) as a special case. In this work we present the explicit formulas for pair condensates and implement angular-momentum projection after variation. Variation after angular-momentum projection (VAP) leads to lower energies but is more complicated and time-consuming; hence we restrict ourselves to projection after variation (PAV). Hereinafter, we label this approach "projection after variation of a pair condensate" (PVPC). It is straightforward to introduce more pair condensates into the configuration space, such as in the generator coordinate method [2], but for now we only consider one pair condensate.

In Sec. II, we outline the formalism for variation of a pair condensate, with details in the Appendix. We show typical timing to find the minimum, as variation is the most time consuming part of the calculation, as exemplified by the descent of <sup>132</sup>Dy in the  $2s1d0g_{7/2}0h_{11/2}$  shell. In Sec. II A, we briefly introduce the formalism for angular-momentum projection, following a recently developed recipe: linear algebraic projection (LAP) [43,44]. In Sec. II B, we decompose a pair condensate into fractions with good angular momentum and show typical probability distributions for rotational nuclei and spherical nuclei. In Sec. IIC, we outline the formalism for electromagnetic transitions between projected wave functions. In Sec. III, we present benchmark calculations of nuclides in the 1s0d, 1p0f, and  $2s1d0g_{7/2}0h_{11/2}$  valence spaces with discussions and analysis. Finally, we summarize and discuss future directions in Sec. IV.

### **II. FORMALISM**

For the nuclear many-body Hamiltonian, we use shellmodel effective interactions in occupation space, with oneand two-body parts, in proton-neutron format,

$$\begin{aligned} \hat{H} &= \hat{H}_{0} + \hat{H}_{pp} + \hat{H}_{nn} + \hat{H}_{pn} \\ &= \sum_{a \in \pi} \epsilon_{a} \hat{n}_{a} + \sum_{a \in \nu} \epsilon_{a} \hat{n}_{a} + \sum_{abcd \in \pi} \frac{\sqrt{(1 + \delta_{ab})(1 + \delta_{cd})}}{4} \sum_{I} V_{pp}(abcd; I) \sum_{M} \hat{A}^{\dagger}_{IM}(ab) \hat{A}_{IM}(cd) \\ &+ \sum_{abcd \in \nu} \frac{\sqrt{(1 + \delta_{ab})(1 + \delta_{cd})}}{4} \sum_{I} V_{nn}(abcd; I) \sum_{M} \hat{A}^{\dagger}_{IM}(ab) \hat{A}_{IM}(cd) + \sum_{ac \in \pi, bd \in \nu, I} V_{pn}(abcd; I) \sum_{M} \hat{A}^{\dagger}_{IM}(ab) \hat{A}_{IM}(cd), \end{aligned}$$
(3)

where  $\hat{A}_{IM}^{\dagger}(ab)$  and  $\hat{A}_{IM}(cd)$  are so-called "noncollective" pair creation and annihilation operators [in contrast with the collective pair creation operator of Eq. (1)]:  $\hat{A}_{IM}^{\dagger}(ab) = [\hat{c}_a^{\dagger} \otimes \hat{c}_b^{\dagger}]_{IM}$ , where  $\otimes$  signals coupling via Clebsch-Gordan coefficients, and  $\hat{A}_{IM}(ab) = [\hat{A}_{IM}^{\dagger}(ab)]^{\dagger}$ .

We work in several model spaces: in the  $1s_{1/2}$ - $0d_{3/2}$ - $0d_{5/2}$  or 1s0d shell, using the universal 1s0d-shell interaction version B, or USDB [22]; in the  $1p_{1/2}$ - $1p_{3/2}$ - $0f_{5/2}$ - $0f_{7/2}$  or 1p0f shell, using a monopole-modified *G*-matrix interaction version 1A, or GX1A [23,24]; and, finally, for the space between magic numbers 50 and 82, comprising the orbits  $2s_{1/2}$ - $1d_{3/2}$ - $1d_{5/2}$ - $0g_{7/2}$ - $0h_{11/2}$ , or  $2s1d0g_{7/2}0h_{11/2}$ , with a monopole-modified *G*-matrix interaction [45].

To carry out our calculations, we need a variety of matrix elements. To begin with, for variation (before projection) we need for a given state  $|\Phi\rangle$ : the normalization  $\langle \Phi | \Phi \rangle$ , the expectation value of a Hamiltonian or the energy of that state  $E(\Phi) = \langle \Phi | \hat{H} | \Phi \rangle / \langle \Phi | \Phi \rangle$ , and the variation of the energy with respect to some parameter  $\partial E(\Phi) / \partial \lambda$ . To carry out projection, we need the overlap or norm kernel of the state  $\Phi$  with a rotated state  $\Phi'$ ,  $\langle \Phi | \Phi' \rangle$ , the Hamiltonian matrix element or Hamiltonian kernel between these two states,  $\langle \Phi | \hat{H} | \Phi' \rangle$ , and, finally, in order to calculate electromagnetic and other transitions represented by a one-body operators  $\hat{Q}$ , we need  $\langle \Phi | \hat{Q} | \Phi' \rangle$ .

All of these are easily found for, say, Hartree-Fock and Hartree-Fock-Bogoliubov calculations. For the pair condensate calculations the formalism is more complicated and less intuitive. We present the detailed formalism in the Appendix but sketch out the application here.

Restricting ourselves to even-even nuclei, our trial Ansatz is a direct product of proton and neutron pair condensates

$$|\Phi\rangle = (\hat{A}_{\pi}^{\dagger})^{N_{\pi}/2} (\hat{A}_{\nu}^{\dagger})^{N_{\nu}/2} |0\rangle, \qquad (4)$$

where  $N_{\pi}$ ,  $N_{\nu}$  are valence numbers for protons and neutrons. We vary  $|\Phi\rangle$ , or more exactly, vary the proton and neutron pair structure coefficients, to minimize the energy

$$\frac{\langle \Phi | \hat{H} | \Phi \rangle}{\langle \Phi | \Phi \rangle}.$$
(5)

To find minima, we apply the GNU scientific library (GSL) [46] conjugate gradient tools. We do variation before projection; that is, we first find the minimum energy of the unprojected pair condensate and afterwards project out states of good angular momentum.

Figure 1(a) shows the central processing unit (CPU) time for one iteration for N = Z nuclei in different major shells, with increasing valence particles. Our timing data were taken on a 44-core workstation using OpenMP parallelization. In the 1s0d, 1p0f,  $2s1d0g_{7/2}0h_{11/2}$  major shells, the numbers of free parameters in variation of one pair condensate basis are 132, 380, and 992, respectively. The computation complexity of variation has polynomial dependence on the valence particle numbers. In Fig. 1(b), an example is given for variation of <sup>132</sup>Dy in the  $2s1d0g_{7/2}0h_{11/2}$  shell, with interactions based on the CD-Bonn potential with the G-matrix approach and monopole optimizations [45]. With 992 free parameters, the iteration tends to converge after 200 steps. We note that the descent is steep for the first 50 steps, and less so afterwards. The descent terminates when the modulus of the gradient vector is smaller than a given criterion. If we set a stricter criterion on the gradient, the descent will continue beyond 200 steps, but with only a slight additional lowering of the

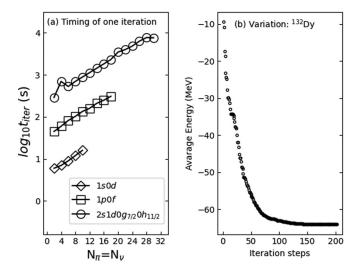


FIG. 1. (a) Logarithm CPU time (s) of one iteration in the variation for N = Z nuclei, using conjugate gradient method [47] wrapped in the GNU Scientific Library function *gsl\_multimin\_fdf\_minimizer* [46]. 1*s*0*d*, 1*p*0*f*, 2*s*1*d*0*g*<sub>7/2</sub>0*h*<sub>11/2</sub> are three valence shells. All these data are collected on a 44-core small workstation, in parallel with simple OpenMP. (b) The descent of the average energy  $\langle \Phi | \hat{H} | \Phi \rangle / \langle \Phi | \Phi \rangle$  for <sup>132</sup>Dy as a test case in the 2*s*1*d*0*g*<sub>7/2</sub>0*h*11/2 shell. There are 32(32 - 1) = 992 free parameters for the variation of the proton and neutron pairs.

energy. Starting from several independent random pairs, we typically find two or three local minima.

#### A. Angular-momentum projection

In this work we allow for triaxial deformations. While there is extensive literature on the standard quadrature approach for angular-momentum projection [2,48–50], we adopt the recently developed "linear algebraic projection"(LAP) technique [43,44]. Although both approaches are based on standard quantum theory of angular momentum [51], i.e., rotational properties of irreducible tensors, for our application we found LAP to be ten times faster than quadrature.

For the standard development of angular-momentum projection, see Refs. [2,51]. The basic idea is that a pair condensate state, or any state  $|\Phi\rangle$  for that matter, can be decomposed with normalized spherical tensor components, with total angular momentum *I* and angular-momentum projection on the third axis *K*,

$$|\Phi\rangle = \sum_{IK} c_{IK} |IK\rangle.$$
(6)

In any finite space the sums are not infinite but restricted to a maximum *I*, *K*, and in practice the coefficients  $c_{IK}$  restricts the sum even further [43]. The projection operator  $\hat{P}_{MK}^{J}$  picks out the (J, K) spherical tensor component and rotates it to a (J, M) spherical tensor,

$$\hat{\rho}^{J}_{MK}|\Phi\rangle = c_{JK}|J, K \to M\rangle.$$
(7)

Then one diagonalizes the Hamiltonian in the truncated space of

$$\left\{\hat{P}^{J}_{M,-J}|\Phi\rangle,\ \hat{P}^{J}_{M,-J+1}|\Phi\rangle,\ \ldots,\ \hat{P}^{J}_{M,J}|\Phi\rangle\right\}.$$
(8)

Because the Hamiltonian is a scalar operator, thus commuting with the projection operator, and because  $(\hat{P}_{MK}^J)^{\dagger} = \hat{P}_{KM}^J$ ,  $\hat{P}_{K'M}^J \hat{P}_{MK}^J = \hat{P}_{K'K}^J$ , we only need

$$\langle \Phi | \hat{H} \hat{P}^J_{MK} | \Phi \rangle \equiv \mathcal{H}^J_{MK}, \tag{9}$$

$$\langle \Phi | \hat{P}^{J}_{MK} | \Phi \rangle \equiv \mathcal{N}^{J}_{MK}. \tag{10}$$

We then solve a generalized eigenvalues problem, i.e., the Hill-Wheeler equation,

$$\sum_{K} \mathcal{H}^{J}_{K'K} g^{r}_{JK} = \epsilon_{r,J} \sum_{K} \mathcal{N}^{J}_{K'K} g^{r}_{JK}.$$
 (11)

where *r* denotes different states with angular momentum *J*, at most 2J + 1 states, and  $g_{JK}^r$  values determine the projected wave function.

To project out states, either by quadrature or by LAP, one needs to rotate states about the Euler angles  $\alpha$ ,  $\beta$ ,  $\gamma$  using the rotation operator

$$\hat{R}(\alpha,\beta,\gamma) = \exp\left(\frac{i\gamma}{\hbar}\hat{J}_z\right)\exp\left(\frac{i\beta}{\hbar}\hat{J}_y\right)\exp\left(\frac{i\alpha}{\hbar}\hat{J}_z\right).$$
 (12)

The matrix elements of the rotation operator between spherical tensors; that is, states with good total angular momentum *j* and *z* component *m*, is the Wigner *D* matrix [51],  $\langle j'm'|\hat{R}(\alpha, \beta, \gamma)|jm\rangle = \delta_{j'j}D^{j}_{m'm}(\alpha\beta\gamma).$ 

For a given set of Euler angles  $\Omega = (\alpha, \beta, \gamma)$ , the matrix element of the rotational operator  $\hat{R}(\Omega)$  is

$$\begin{split} \langle \Phi | \hat{R}(\Omega) | \Phi \rangle &= \sum_{IMK} c_{IK}^* c_{IM} D_{KM}^I(\Omega) \langle IK | IM \to K \rangle \\ &= \sum_{IMK} D_{KM}^I(\Omega) \mathcal{N}_{KM}^I. \end{split}$$
(13)

For a series of Euler angles  $\Omega_1, \Omega_2, \ldots$ , we calculate the norm kernels  $\langle \Phi | \hat{R}(\Omega_i) | \Phi \rangle$ , as well as the *D* functions  $D_{KM}^l(\Omega_i)$ , and then find  $\mathcal{N}_{KM}^l$  as solutions to a set of linear equations. Similarly the Hamiltonian kernel  $\mathcal{H}_{KM}^l$  can be also found.

The traditional projection procedure relies on the orthogonality of the Wigner *D* functions and thus upon quadrature; therefore, the more abscissas (Euler angles) one takes, the more accurate are the projected energies. The basis for LAP, Eq. (13), is exact, but in practice one truncates the sum over *I* with  $I \leq J_{\text{max}}$ . To find  $J_{\text{max}}$ , one computes  $f_J$ , the fraction of *J* components defined in Eq. (21) with  $f_J = \sum_M |c_{JM}|^2$  and sets  $J_{\text{max}}$  such that  $f_J \geq f_{tol}$  for all  $J \leq J_{\text{max}}$ . We follow Ref. [44] and take  $f_{tol} \approx 10^{-4}$  typically. The error in projection is due to this truncation. As a result of finite  $J_{\text{max}}$ , in LAP the number of Euler angles is smaller. We implemented both methods, and found that LAP is more than 10 times faster than quadrature, with same accuracy of the energies of low-lying states.

To carry out this projection we need to be able to rotate pair condensates. This is almost as straightforward as for Slater determinants in angular-momentum projected Hartree-Fock [43]: one simply rotates the single-particle states, defining the matrix element of the rotation operator on singleparticle bases  $\alpha$ ,  $\beta$  as

$$\mathcal{D}_{\alpha\beta}(\Omega) = \langle n_{\alpha}l_{\alpha}j_{\alpha}m_{\alpha}|\hat{R}(\Omega)|n_{\beta}l_{\beta}j_{\beta}m_{\beta}\rangle$$
$$= \delta_{n_{\alpha}n_{\beta}}\delta_{l_{\alpha}l_{\beta}}\delta_{j_{\alpha}j_{\beta}}D^{j_{\alpha}}_{m_{\alpha}m_{\beta}}(\Omega).$$
(14)

Under rotation, a nucleon-pair (1) then becomes

$$\hat{R}(\Omega)\hat{A}^{\dagger} = \frac{1}{2} \sum_{\alpha\beta\gamma\delta} A_{\alpha\beta} \mathcal{D}_{\gamma\alpha}(\Omega) \mathcal{D}_{\delta\beta}(\Omega) \hat{c}_{\gamma}^{\dagger} \hat{c}_{\delta}^{\dagger}$$
$$= \frac{1}{2} \sum_{\gamma\delta} A'_{\gamma\delta} \hat{c}_{\gamma}^{\dagger} \hat{c}_{\delta}^{\dagger} \equiv \hat{A}'^{\dagger}, \qquad (15)$$

with new pair coefficients

$$A'_{\gamma\delta} = \sum_{\alpha\beta} A_{\alpha\beta} \mathcal{D}_{\gamma\alpha}(\Omega) \mathcal{D}_{\delta\beta}(\Omega) = (\mathcal{D}(\Omega)A\mathcal{D}(\Omega)^{\top})_{\gamma\delta}.$$
 (16)

Therefore, the matrix element of rotational operator  $\hat{R}(\Omega)$  on a pair condensate is

$$\langle \hat{A}^{n} | \hat{R}(\Omega) | (\hat{A}^{\dagger})^{n} \rangle = \langle \hat{A}^{n} | (\hat{A'}^{\dagger})^{n} | 0 \rangle, \quad A' = \mathcal{D}(\Omega) A \mathcal{D}(\Omega)^{\top},$$
(17)

and similarly

$$\langle \hat{A}^{n} | \hat{H} \hat{R}(\Omega) | (\hat{A}^{\dagger})^{n} \rangle = \langle \hat{A}^{n} | \hat{H} (\hat{A'}^{\dagger})^{n} | 0 \rangle, \quad A' = \mathcal{D}(\Omega) A \mathcal{D}(\Omega)^{\top}.$$
(18)

Parity projection is similar and straightforward, with a given wave function  $|\Psi\rangle$ , the positive- and negative-parity components are ( $\hat{\mathcal{P}}$  is the space inversion operator)

$$\frac{1}{2}(1\pm\hat{\mathcal{P}})|\Psi\rangle.$$
(19)

To invert a many-body wave function, is equivalent to inverting all its single-particle orbits. When a pair is operated on by the space inversion operator, an additional phase factor is generated for the pair structure coefficients,

$$\hat{\mathcal{P}}\hat{A}^{\dagger} = \frac{1}{2} \sum_{\alpha\beta} A_{\alpha\beta} (\hat{\mathcal{P}}\hat{c}^{\dagger}_{\alpha}) (\hat{\mathcal{P}}\hat{c}^{\dagger}_{\beta}) = \frac{1}{2} \sum_{\alpha\beta} A_{\alpha\beta} (-1)^{l_{\alpha}+l_{\beta}} \hat{c}^{\dagger}_{\alpha} \hat{c}^{\dagger}_{\beta}.$$
(20)

In this work, we do both angular-momentum and parity projections for the results in the  $2s1d0g_{7/2}0h_{11/2}$  shell.

### B. Tensor decomposition of a pair condensate

As expressed by Eq. (6), a pair condensate can be decomposed into components with different angular-momentum quantum numbers (J, M). In other words, an optimized pair condensate is considered as a mixture of low-lying nuclear states. We define the fraction of probability  $f_J$  for total angular momentum J in a pair condensate, as in Ref. [52],

$$f_J = \sum_M |c_{JM}|^2,$$
 (21)

where  $c_{JM}$  is the amplitude of (J, M) tensor component in the pair condensate, as in Eq. (6). To illustrate the decompositions for rotational and spherical nuclei, we show <sup>52</sup>Fe and <sup>48</sup>Ca, respectively, in Fig. 2. For <sup>52</sup>Fe, in Fig. 2(a), the decomposition is distributed across J = 0, 2, 4, 6, as expected for rotational nuclei. Conversely, for <sup>48</sup>Ca in Fig. 2(b), the wave function is almost entirely contained in the J = 0 component; while there are nonzero  $f_J$  for J > 0 values, the magnitudes are too small to be seen in Fig. 2(b). Therefore, the optimized condensate for semimagic nuclei is mostly the spherical J = 0 ground state, while tiny nonspherical components of the condensate are projected out to be the nonzero spectrum, which turn out

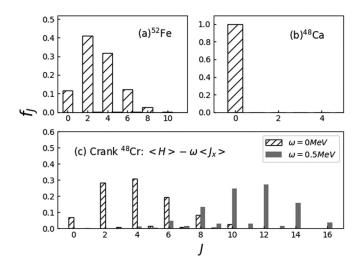


FIG. 2. (a) Fractions of *J*-components in the optimized pair condensate for <sup>52</sup>Fe, as an example of rotational nuclei. (b) Fractions for <sup>48</sup>Ca as an example of spherical nuclei. (c) Fractions for <sup>48</sup>Cr as an example of cranking. All these cases are computed with the GX1A interactions in the 1p0f shell.

to be surprisingly accurate in comparison with shell-model results, as will be shown in Fig. 5.

However, if cranking, or variation with constraints, is involved, one can increase the fraction of components with higher J. As in the usual procedure [2], one optimizes the trial wave function using the effective Hamiltonian

$$\hat{H}' = \hat{H} - \omega \langle J_x \rangle. \tag{22}$$

Figure 2(c) shows the decomposition of <sup>48</sup>Ca at  $\omega = 0$  and  $\omega = 0.5$  MeV; the latter increases the components with J > 0. Nonetheless, for simplicity we restrict ourselves to optimization without cranking for the rest of this paper.

# C. Electromagnetic transitions

In this section we consider electromagnetic transitions between projected wave functions. As this is addressed in detail elsewhere, e.g., Ref. [50], we briefly review the formalism. Consider a state  $|\Phi\rangle$  such as the optimized pair condensate; after diagonalization in the subspace of kets for angular momentum J,

$$\hat{P}^J_{MK}|\Phi\rangle, \quad K = -J, -J+1, \dots, J, \tag{23}$$

the eigenstate is

$$\psi_{JM}^{r} = \sum_{K} g_{JK}^{r} \hat{P}_{MK}^{J} |\Phi\rangle = \sum_{K} g_{JK}^{r} \left| \hat{P}_{MK}^{J} \Phi \right\rangle, \qquad (24)$$

where  $g_{JK}^r$  is the solution to the Hill-Wheeler equation (11), and  $|\hat{P}_{MK}^J \Phi\rangle = \hat{P}_{MK}^J |\Phi\rangle$ . The one-body electromagnetic operator is denoted as (with proton and neutron parts),

$$\hat{Q}^s_{\sigma} = \hat{Q}^{\pi s}_{\sigma} + \hat{Q}^{\nu s}_{\sigma}, \qquad (25)$$

where s is the angular-momentum rank of the operator and  $\sigma$  is the z component. The matrix element of the transition operator on projected kets is

$$\langle \Phi | (\hat{P}_{M'K'}^{J'})^{^{\intercal}} \mathcal{Q}_{\sigma}^{s} \hat{P}_{MK}^{J} | \Phi \rangle$$

$$= (JM, s\sigma | J'M') \langle (\hat{P}_{*K'}^{J'} \Phi) | | \mathcal{Q}^{s} | | \hat{P}_{*K}^{J} \Phi \rangle, \qquad (26)$$

where  $\langle (\hat{P}_{*K'}^{J'} \Phi) || Q^s || \hat{P}_{*K}^J \Phi \rangle$  is the reduced matrix element and  $(JM, s\sigma | J'M')$  is the Clebsch-Gordan coefficient [51]. As  $(\hat{P}_{M'K'}^{J'})^{\dagger} = \hat{P}_{K'M'}^{J'}$ , and the projection operator  $\hat{P}_{K'M'}^{J'}$  just picks out the (J', M') component and rotates it to (J', K'),

$$\begin{split} \langle \Phi | (\hat{P}_{M'K'}^{J'})^{\mathsf{T}} \mathcal{Q}_{\sigma}^{s} \hat{P}_{MK}^{J} | \Phi \rangle \\ &= \langle \Phi | (\hat{P}_{M'K'}^{J'})^{\mathsf{T}} \sum_{J''M''} (s\sigma, JM | J''M'') (\mathcal{Q}^{s} \hat{P}_{*K}^{J} | \Phi)_{M''}^{J''}), \\ &= (s\sigma, JM | J'M') \langle \Phi | (\mathcal{Q}^{s} \hat{P}_{*K}^{J})_{K'}^{J'} | \Phi \rangle. \end{split}$$
(27)

Therefore, the reduced matrix element can be computed by

$$\begin{split} \left\langle \left( \hat{P}_{*K'}^{J'} \Phi \right) \middle| \left| \mathcal{Q}^{s} \right| \middle| \hat{P}_{*K}^{J} \Phi \right\rangle \\ &= (-1)^{s+J-J'} \langle \Phi \middle| \left( \mathcal{Q}^{s} \hat{P}_{*K}^{J} \right)_{K'}^{J'} \middle| \Phi \rangle \\ &= \sum_{\sigma} \langle J, K' - \sigma; s, \sigma \middle| J'K' \rangle \langle \Phi \middle| \mathcal{Q}_{\sigma}^{s} \hat{P}_{K'-\sigma,K}^{J} \middle| \Phi \rangle. \end{split}$$
(28)

Here the needed projections are also carried out using the linear algebra method.

Between the initial and final eigenfunctions  $\psi_{JM}^r = \sum_K g_{JK}^r \hat{P}_{MK}^J |\Phi\rangle$ , the reduced matrix element of  $\hat{Q}^s$  is then finally

$$\begin{aligned} \left\langle \psi_{J'}^{r'} \middle| \left| \hat{Q}^{s} \right| \middle| \psi_{J}^{r} \right\rangle &= \sum_{KK'} g_{JK}^{r} g_{J'K'}^{r'} \left\langle \left( \hat{P}_{*K'}^{J'} \Phi \right) \middle| \left| \hat{Q}^{s} \right| \left| \left( \hat{P}_{*K}^{J} \Phi \right) \right\rangle \\ &= \sum_{KK'} g_{JK}^{r} g_{J'K'}^{r'} \sum_{\sigma M} (JM, s\sigma | J'K') \langle \Phi | Q_{\sigma}^{s} \hat{P}_{MK}^{J} | \Phi \rangle. \end{aligned}$$

$$(29)$$

The reduced probability, or B value, is

$$B(F, J_{r} \to J_{r'}') = \frac{2J' + 1}{2J + 1} |\langle \psi_{J'}^{r'} || \hat{Q}^{s} || \psi_{J}^{r} \rangle|^{2}$$
  
$$= \frac{2J' + 1}{2J + 1} \left| \sum_{KK'\sigma} g_{JK}^{r} g_{J'K'}^{r'} (JK' - \sigma, s\sigma | J'K') \right|^{2} \times \langle \Phi | Q_{\sigma}^{s} \hat{P}_{K'-\sigma,K}^{J} | \Phi \rangle \right|^{2}.$$
(30)

### **III. BENCHMARK RESULTS AND ANALYSIS**

In Sec. III A, we present ground-state energies of all 1s0d-shell even-even nuclei, with comparisons against full configuration-interaction shell-model and projected Hartree-Fock calculations, and show ground-band spectra as well as electric-quadrupole transitions strengths, or B(E2) values, of silicon isotopes as an example. In Sec. III B, we show 1p0f-shell results, including shape evolution shown in lowest spectrum of <sup>46,48</sup>Ca, <sup>48,50</sup>Ti, <sup>50,52</sup>Cr and the backbending phenomenon in <sup>52</sup>Fe. Finally, in Sec. III C, we illustrate the applicability to medium-heavy nuclides with <sup>104</sup>Sn, <sup>106</sup>Te, <sup>108,124,126</sup>Xe, and <sup>126,128</sup>Ba as examples in the  $2s1d0g_{7/2}0h_{11/2}$  shell.

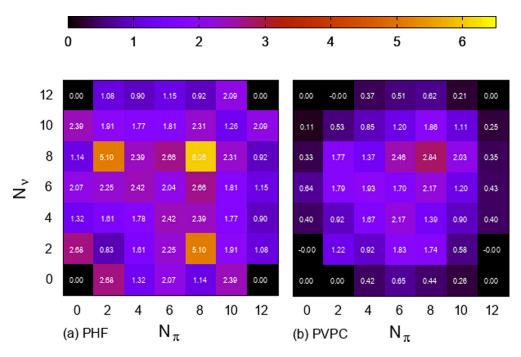


FIG. 3. Ground-state energy differences in MeV between approximate calculations and "exact diagonalization" from full configurationinteraction shell model,  $\Delta E_{g.s.}^{approx} = E_{g.s.}^{approx} - E_{g.s.}^{SM}$ , for all even-even nuclei in 1*s*0*d* shell. Panel (a) is for angular-momentum-projected Hartree Fock results, and panel (b) is for the PVPC in this work. See text for more details.

### A. The 1s0d shell

In variational methods, the ground-state energy is often used as a marker of the quality of the approximation. Because the 1*s*0*d* shell computations are easy and fast for shell-model codes, we compute ground-state energies of all even-even nuclei in the 1*s*0*d* shell, comparing PVPC against full configuration-interaction shell-model diagonalization, using the BIGSTICK code [53] and angular-momentum projected Hartree-Fock (PHF) calculations, e.g., Ref. [54], using the exact same shell-model space and interaction input. In particular, we characterize the discrepancy in the ground-state energy relative to the full configuration shell model (SM) results, which for us are "exact," by

$$\Delta E_{\text{g.s.}}^{\text{approx}} = E_{\text{g.s.}}^{\text{approx}} - E_{\text{g.s.}}^{\text{SM}}.$$
(31)

Here "approx" can be either PHF or PVPC. In Fig. 3,  $\Delta E_{g.s.}^{PHF}$ and  $\Delta E_{g.s.}^{PVPC}$  are shown as so-called "heat maps" indexed by colors. For most cases, the PHF ground-state energy is about 2 MeV higher than the exact ground-state energy. For three nuclei,  $\Delta E_{g.s.}^{PHF}$  is more than 5 MeV. For example, <sup>32</sup>S is a rotational nucleus, while the Hartree-Fock minimum generates a spherical shape for it, therefore only a 0<sup>+</sup> is projected out, as is also pointed out in Ref. [55]. PVPC is rather accurate for ground states of semimagic nuclei, with  $\Delta E_{g.s.}^{PVPC}$  typically less than 0.5 MeV, which is unsurprising, as generalized seniority-0 configurations are contained within the pair condensate Ansatz. For rotational nuclei,  $\Delta E_{g.s.}^{PVPC}$  is typically 1–2 MeV, i.e., PVPC has ground states of rotational nuclei slightly lower than that of PHF. These are benchmark results for one reference state, i.e., one pair condensate for PVPC or one Slater determinant for PHF. We expect inclusion of more reference states as done in generator-coordinate calculations [49] will reduce discrepancies.

Aside from the ground-state energies, it is interesting to see how well the excitation energies are approximated by such "projection after variation" methods, as was done recently for PHF [54]. As a test we investigate yrast bands of even-even isotopes of silicon, using results from PVPC, PHF, and full configuration-interaction shell model. Figure 4 demonstrates

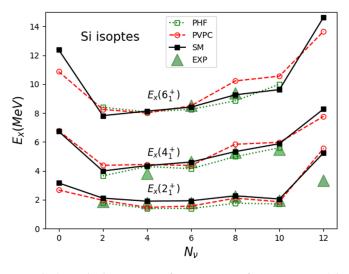


FIG. 4. Excitation energy of yrast states of even-even nuclei  $^{22-34}$ Si. "PHF" stands for angular-momentum-projected Hartree-Fock, "PVPC" stands for angular-momentum projection after variation of a pair condensate (this work), and "SM" stands for the full configuration-interaction shell model. All three methods use the same occupation space 1s0d and the same interaction, USDB [22]. Experimental data are from Refs. [56–61].

TABLE I. B(E2) values for silicon isotopes in units of  $e^2$  fm<sup>4</sup>. SM stands for full configuration shell model with exact diagonalization, while PVPC stands for projection and variation of one pair condensate. Both methods use the USDB interaction in the 1*s*0*d* shell. Effective charges are set as  $e_p = 1.5e$ ,  $e_n = 0.5e$ .

Nuclide	Method	$2^+_1 \rightarrow 0^+_1$	$4^+_1 \rightarrow 2^+_1$	$6^+_1 \rightarrow 4^+_1$
<sup>22</sup> Si	SM	33.32	16.96	13.07
	PVPC	32.28	14.10	14.81
<sup>24</sup> Si	SM	44.19	23.89	17.58
	PVPC	37.13	25.60	41.66
<sup>26</sup> Si	SM	44.64	11.29	41.94
	PVPC	37.34	60.31	70.56
<sup>28</sup> Si	SM	78.48	110.4	94.71
	PVPC	83.47	116.3	119.2
<sup>30</sup> Si	SM	46.52	15.94	32.70
	PVPC	48.33	33.68	78.15
<sup>32</sup> Si	SM	43.52	67.11	52.70
	PVPC	37.54	64.06	66.85
<sup>34</sup> Si	SM	34.91	16.80	6.14
	PVPC	30.78	14.93	1.50

that both PHF and PVPC generate good yrast-band excitation spectra for open-shell nuclei. For semimagic nuclei <sup>22,34</sup>Si, Hartree-Fock minima are spherical and thus PHF trivially generates only the ground state  $0_1^+$ . The PVPC generates yrast bands of <sup>22,34</sup>Si close to that of the shell model. For <sup>34</sup>Si both the PVPC and the shell model predict the  $2_1^+$  excitation energy higher than the experimental value [56]. We note that approximate excitation energies do not follow a variational principle and can be lower than that of the full configuration shell model. Although it is not shown here, the yrare and higher bands of PVPC can be worse, and likely need more reference states to improve, such as from coexisting local minima [54].

As a test of the approximate wave functions, we present in Table I the  $B(E2, I + 2 \rightarrow I)$  values for silicon isotopes corresponding to Fig. 4. While projected states can reproduce the transition strengths qualitatively, there can still be discrepancies, for example,  $B(E2, 4_1^+ \rightarrow 2_1^+)$  of <sup>26</sup>Si from the PVPC is about 1/6 of that from the shell model; again, this may be improved by additional reference condensates.

### B. The 1p0f shell

We use the GX1A interaction in the 1*p*0*f* shell [23,24]. We find the PVPC ground-state energy is typically 1–2 MeV above the exact shell-model value. Here we give results of Ca, Ti, Cr isotones with N = 26, 28, to show that PVPC generates spectrum of pairing-like and vibrational-like spectrum selfconsistently, as the number of valence particles increase. Then we demonstrate backbending in the rotational yrast spectrum of <sup>52</sup>Fe, with the pair condensate as the intrinsic wave function.

# *1.* <sup>46,48</sup>Ca, <sup>48,50</sup>Ti, <sup>50,52</sup>Cr

Figure 5 presents the low-lying spectra of N = 26, 28 isotones  ${}^{46,48}$ Ca,  ${}^{48,50}$ Ti,  ${}^{50,52}$ Cr, using results from PVPC,

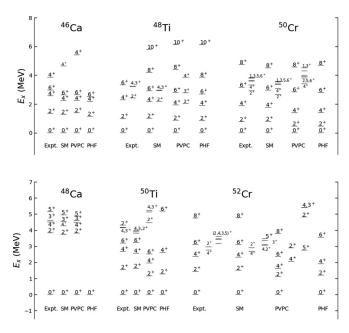


FIG. 5. Low-lying spectrum of N = 26, 28 isotones: <sup>46,48</sup>Ca, <sup>48,50</sup>Ti, <sup>50,52</sup>Cr. All theoretical results are produced with the interaction GX1A [23,24] in the 1*p*0*f* shell. Experimental results are from Refs. [62–65].

PHF, and the full configuration shell model, as well as comparison with experiment which illustrates the quality of the shell-model interaction. For the spherical nucleus <sup>48</sup>Ca, the fraction  $f_J$  for J = 0 is more than 99.9%, as shown in Fig. 2; nonetheless, the J > 0 states projected out with  $f_J < 0.1\%$ turn out to be unexpectedly accurate. We see seniority-like spectra with a substantial pairing gap between  $0_1^+$  and  $2_1^+$ , and the 3+, 4+, 5+ levels clustered close to  $2_1^+$ . Although it is not shown in Fig. 5, the  $0_1^+$  of <sup>46,48</sup>Ca is  $\approx 0.3$  MeV above the exact shell-model ground state, while other nuclei in the picture have ground-state energies 1–2 MeV above shell-model results. The excitation spectra of the "ground band" from all three theoretical methods agree qualitatively, except that PHF does not generate J > 0 states for <sup>48</sup>Ca, with a spherical Hartree-Fock energy about 0.4 MeV above the exact shell-model value.

Table II provides the corresponding B(E2) values for the nuclides in Fig. 5. The agreement between PVPC and the full configuration shell model supports the quality of PVPC wave functions, although we note that  $B(E2; 6_1^+ \rightarrow 4_1^+)$  of <sup>48</sup>Ca by PVPC is far from the shell-model result. We note that the shell-model spectra of <sup>48</sup>Ca has four 6<sup>+</sup> states with an excitation energy between 7.7 and 9 MeV, while PVPC yields only one 6<sup>+</sup> state in the same energy range. Therefore, the  $6_1^+$  by PVPC is likely a mixture of those shell model 6<sup>+</sup> states, a possible origin for the underestimation of  $B(E2; 6_1^+ \rightarrow 4_1^+)$ .

# 2. Backbending in <sup>52</sup>Fe

While many methods using angular-momentum projection produce good rotational bands, e.g., projected Hartree-Fock [43] and the projected shell model [48], a more stringent test is backbending. For a simple rotor picture, the excitation energy of a yrast state  $E_x(J_1^+)$  is proportional to J(J + 1), but

Nucleus	Method	$2^+_1 \rightarrow 0^+_1$	$4^+_1 \rightarrow 2^+_1$	$6^+_1 \rightarrow 4^+_1$	Nucleus	Method	$2^+_1 \rightarrow 0^+_1$	$4^+_1 \rightarrow 2^+_1$	$6^+_1  ightarrow 4^+_1$
<sup>46</sup> Ca	SM PVPC	7.72 7.31	6.02 4.10	3.08 2.16	<sup>48</sup> Ca	SM PVPC	10.12 9.10	1.99 1.76	4.03 0.36
<sup>48</sup> Ti	SM PVPC	87.89 84.67	126.06 116.36	51.28 113.57	<sup>50</sup> Ti	SM PVPC	85.20 65.80	83.10 58.77	39.84 19.87
<sup>50</sup> Cr	SM PVPC	184.85 190.33	261.43 257.78	221.87 266.68	<sup>52</sup> Cr	SM PVPC	142.31 102.8	91.94 89.00	45.91 56.39

TABLE II. B(E2) value of <sup>48</sup>Ca, <sup>50</sup>Ti, <sup>52</sup>Cr, <sup>52</sup>Fe in  $e^2$  fm<sup>4</sup>. The PVPC and the shell model both use the GX1A interactions [23,24], with effective charges  $e_p = 1.5e$ ,  $e_n = 0.5e$ , and oscillator length  $A^{1/6}$  fm.

both experiment and theory see abrupt changes in the moment of inertia and thus the yrast spectrum deviates from the simple J(J + 1) rule. One explanation of this is pair breaking due to the increasing Coriolis force [2]. In Fig. 6 we show yrast energies of <sup>52</sup>Fe, from PVPC with a pair condensate playing the role of the intrinsic state, and that from exact diagonalization with the shell model. The PVPC energies are about 1 MeV or more above the full configuration shell-model energies, but both PVPC and the shell model exhibits two backbends. The PVPC curve shows more abrupt changes than the shell model. The change for J > 12 is likely because six valence protons and six valence neutrons restricted to the  $0f_{7/2}$  orbit have a maximum angular momentum of 12.

In Fig. 6, we also show yrast states projected from an uncranked Hartree-Fock minimum, denoted as "PHF," as well as those projected from cranked Hartree-Fock minima, denoted as "PCHF." Without cranking, the Hartree-Fock minimum has components with  $J \leq 12$ ; with cranking, i.e., finding the state that minimizes the expectation value of  $\hat{H}' = \hat{H} - \omega \langle \hat{J}_x \rangle$ , components with J > 12 are populated. At  $\omega = 0.4$  MeV, the cranked Hartree Fock results produces a J = 12 state that is still too high energetically; but at  $\omega = 0.45$  MeV a

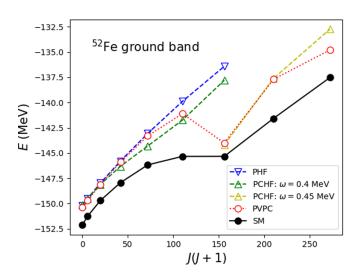


FIG. 6. Backbending of yrast band of  ${}^{52}$ Fe. "PVPC" is projected states from one pair condensate, "SM" is the full configuration shell model, "PHF" is the projected Hartree-Fock method, and "CPHF" is the projection from a cranked Hartree-Fock state. The GX1A interaction is used for the 1p0f shell.

J = 12 much lower in energy is produced. The PVPC yrast band projected from a single pair condensate follows the cranked Hartree-Fock results and significantly improved upon the band projected from the uncranked HF state. Because the PVPC states are projected out of a single pair condensate, our results appear to challenge the idea that backbending arises from the Coriolis-induced breaking of pairs [2,66]. We note that triaxial projection naturally induces mixing of components with different *K*, as also indicated in schematic studies of triaxial quantum rotors [67]. While our uncranked HF calculation has a natural axial symmetry arising without constraint, the cranked HF minima have a small but nonetheless non-negligible triaxiality.

## C. The $2s1d0g_{7/2}0h_{11/2}$ shell

Figure 7 shows the ground-state band of three N = 54 isotones. For <sup>104</sup>Sn, <sup>106</sup>Te we take the experimental data from Refs. [68,69], while the low-lying spectrum of <sup>108</sup>Xe has not been not measured. PVPC generates a ground band rather close to that of the shell model. The discrepancy  $\Delta E_{g.s.}^{PVPC} \equiv E_{g.s.}^{PVPC} - E_{g.s.}^{SM}$  is about 0.2, 0.8, 0.9 MeV, respectively for <sup>104</sup>Sn, <sup>106</sup>Te, <sup>108</sup>Xe. As in the 1*s*0*d* or 1*p*0*f* shell, PVPC does better for semimagic nuclei than PHF, as one might anticipate, but PHF approaches PVPC for the rotational nucleus <sup>108</sup>Xe, although the PVPC ground-state energy is about 0.5 MeV lower than that of PHF. In Table III, the  $B(E2, I_1^+ \rightarrow (I-2)_1^+)$  values are presented, in correspondence to Fig. 7. The B(E2) values of PVPC agree qualitatively with those of

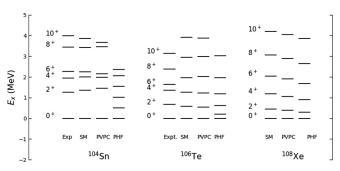


FIG. 7. Ground-band excitation energies of N = 54 isotones <sup>104</sup>Sn, <sup>106</sup>Te, <sup>108</sup>Xe. Both the shell model and the PVPC use a monopole-adjusted *G*-matrix interaction [45], and the experimental data are from Refs. [68,69].

Nucleus	Method	$2^+_1 \rightarrow 0^+_1$	$4^+_1 \rightarrow 2^+_1$	$6^+_1  ightarrow 4^+_1$	$8^+_1  ightarrow 6^+_1$	$10^+_1 \rightarrow 8^+_1$
<sup>104</sup> Sn	SM	42.9	40.0	14.2	18.1	30.8
	PVPC	36.1	24.5	5.4	10.8	19.6
<sup>106</sup> Te	SM	444.8	588.0	573.0	478.6	504.8
	PVPC	430.3	597.8	619.7	565.8	485.0
<sup>108</sup> Xe	SM	943.3	1277	1373	1449	1369
	PVPC	755.6	1044	963.2	901.1	1044

TABLE III. B(E2) value of N = 54 isotones <sup>104</sup>Sn, <sup>106</sup>Te, <sup>108</sup>Xe in  $e^2$  fm<sup>4</sup>. Both methods use a monopole-modified *G*-matrix interaction [45] in the  $2s1d0g_{7/2}0h_{11/2}$  shell. Effective charges are set as  $e_p = 1.5e$ ,  $e_n = 0.5e$ .

the shell model. We also show the ground band of  $^{124,126}$ Xe,  $^{126,128}$ Ba in Fig. 8 generated with PVPC and PHF, although  $0_1^+$  of both these methods appear not low enough. For all four nuclei  $^{124,126}$ Xe,  $^{126,128}$ Ba, the PVPC generates ground-state energies less than 0.5 MeV lower than PHF does, and the excitation energies of PVPC and PHF are mostly alike.

#### **IV. SUMMARY**

In summary, we derived formulas for projection after variation of pair condensates (PVPC), implemented those formulas into codes, and present one-reference (i.e., one pair condensate for one nucleus) benchmark results. The variation of a pair condensate is time-consuming because it takes several hours to optimize a pair condensate for <sup>132</sup>Dy in the  $2s1d0g_{7/2}0h_{11/2}$  space on a 44-core machine. After finding the variational minimum, triaxial projection with the LAP technique is very quick. Because the pair condensate as a variation Ansatz is more general than the Hartree-Fock, or the Hartree-Fock-Bogolyubov method, the PVPC generates semimagic or open-shell nuclei more consistently. We benchmark results of nuclei in 1s0d, 1p0f,  $2s1d0g_{7/2}0h_{11/2}$  shells by comparing results of PVPC with PHF and shell model when possible. For semimagic nuclei the PVPC generates seniority-type spectrum, satisfactorily close to shell-model results. For rotational nuclei, the PVPC generates rotational ground band, similar to results from the projected Hartree-Fock method, although less than 1 MeV lower, and with a few more states.

In future work, a benchmark comparison with the Hartree-Fock-Bogoliubov method can be interesting because the HFB and PVPC both have intrinsic pairing structure. PVPC allows for triaxial deformation, but due to constraint of one reference

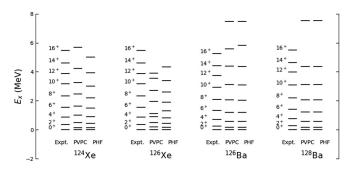


FIG. 8. Ground-band excitation energy of <sup>124,126</sup>Xe, <sup>126,128</sup>Ba.

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(one pair condensate), triaxial phenomena, such as a low  $0_2^+$  in <sup>76</sup>Ge, are not reproduced yet. In addition, inclusion of more than one pair condensate à la generator coordinate methods can be implemented. It is also straightforward to extend the formalism to proton-neutron pair condensates, or odd valence protons or neutrons, or separate projection of proton or neutron parts as related to the scissors mode.

## ACKNOWLEDGMENTS

We would like to thank useful discussions with Fang-Oi Chen, Jie Meng, Guan Jian Fu, and part of shell-model results from Yi Xin Yu. Yi Lu acknowledges support by the National Natural Science Foundation of China (Grants No. 11705100 and No. 12175115), the Youth Innovations and Talents Project of Shandong Provincial Colleges, and Universities (Grant No. 201909118), Higher Educational Youth Innovation Science and Technology Program Shandong Province (Grant No. 2020KJJ004). Yang Lei is grateful for the financial support of the Sichuan Science and Technology Program (Grant No. 2019JDRC0017), and the Doctoral Program of Southwest University of Science and Technology (Grant No. 18zx7147). This material is also based upon work (Calvin W. Johnson) supported by the U.S. Department of Energy, Office of Science, Office of Nuclear Physics, under Award No. DE-FG02-03ER41272.

# APPENDIX A: FORMULAS FOR HAMILTONIAN MATRIX ELEMENTS AND THEIR DERIVATIVES

To carry out the variation and projection of a pair condensate state  $|\Phi\rangle \equiv (\hat{A}^{\dagger})^n |0\rangle$  we need a variety of matrix elements. In particular we need the normalization  $\langle \Phi | \Phi \rangle$ , then expectation of the Hamiltonian of that state  $\langle \Phi | \hat{H} | \Phi \rangle$ , and the partial derivatives of these two quantities:  $\partial \langle \Phi | \Phi \rangle / \partial A_{\alpha\beta}$ and  $\partial \langle \Phi | \hat{H} | \Phi \rangle / \partial A_{\alpha\beta}$ . To do projection, we also need to rotate the condensate  $\hat{R}(\Omega)(\hat{A}^{\dagger})^n | 0 \rangle = (\hat{A}'^{\dagger})^n | 0 \rangle$ , with coefficient matrix  $A' = \mathcal{D}(\Omega)A\mathcal{D}(\Omega)^{\top}$ , and to calculate the norm and Hamiltonian kernels:  $\langle \hat{A}^n | \hat{R}(\Omega) | (\hat{A}^{\dagger})^n \rangle = \langle \hat{A}^n | (\hat{A}'^{\dagger})^n \rangle$  and  $\langle \hat{A}^n | \hat{H} \hat{R}(\Omega) | (\hat{A}^{\dagger})^n \rangle = \langle \hat{A}^n | \hat{H} | (\hat{A'}^{\dagger})^n \rangle$ . In this Appendix we give the expressions for Hamiltonian matrix elements and their derivatives in terms of condensate overlaps including what we term impurity pairs. In Appendix B, we derive all necessary overlap formulas of condensates, with or without "impurity" pairs, completing the necessary formalism.

## 1. Hamiltonian matrix elements

The Hamiltonian matrix element,  $\langle (\hat{A}_1)^n | \hat{H} | (\hat{A}_2^{\dagger})^n \rangle$ , consists of one-body parts  $\langle (\hat{A}_1)^n | \hat{Q} | (\hat{A}_2^{\dagger})^n \rangle$ , such as single-particle energies, and two-body parts  $\langle (\hat{A}_1)^n | \hat{A}_3^{\dagger} \hat{A}_4 | (\hat{A}_2^{\dagger})^n \rangle$ . The one-body operator  $\hat{Q} \equiv \sum_{\alpha\beta} Q_{\alpha\beta} \hat{c}_{\alpha}^{\dagger} \hat{c}_{\beta}$  has structure coefficients  $Q_{\alpha\beta}$ . Through commutation relations, the matrix element of a onebody operator can be expressed as an overlap,

$$\begin{aligned} \langle (\hat{A}_1)^n | \hat{Q} | (\hat{A}_2^{\dagger})^n \rangle &= n \langle (\hat{A}_1)^{n-1} [ \hat{A}_1, \hat{Q} ] | (\hat{A}_2^{\dagger})^n \rangle \\ &= n \langle (\hat{A}_1)^{n-1} \hat{A}_3 | (\hat{A}_2^{\dagger})^n \rangle, \end{aligned}$$
(A1)

where  $\hat{A}_3 \equiv [\hat{A}_1, \hat{Q}]$ , with pair structure coefficients  $A_3 = A_1Q + Q^{\top}A_1$ . We term  $\hat{A}_3$  an "impurity pair" because it replaces one  $\hat{A}_1$  in the bra vector  $\langle (\hat{A}_1)^n |$  to arrive at Eq. (A1).

Similarly, the matrix element of a general two-body interaction  $\hat{A}_{3}^{\dagger}\hat{A}_{4}$  can be written as

$$\begin{split} \langle (\hat{A}_{1})^{n} | \hat{A}_{3}^{\dagger} \hat{A}_{4} | (\hat{A}_{2}^{\dagger})^{n} \rangle &= \langle \hat{A}_{1}^{n} \hat{A}_{4} | \hat{A}_{3}^{\dagger} (\hat{A}_{2}^{\dagger})^{n} \rangle + \langle \hat{A}_{1}^{n} [ \hat{A}_{3}^{\dagger}, \hat{A}_{4} ] (\hat{A}_{2}^{\dagger})^{n} \rangle \\ &= \langle \hat{A}_{1}^{n} \hat{A}_{4} | \hat{A}_{3}^{\dagger} (\hat{A}_{2}^{\dagger})^{n} \rangle + \frac{1}{2} \operatorname{tr} (A_{3} A_{4}) \langle \hat{A}_{1}^{n} | (\hat{A}_{2}^{\dagger})^{n} \rangle \\ &- n \langle \hat{A}_{1}^{n-1} \hat{A}_{5} | (\hat{A}_{2}^{\dagger})^{n} \rangle, \end{split}$$
(A2)

where tr( $A_3A_4$ ) is the trace of the matrix product  $A_3A_4$ , and  $A_5 = A_1A_3A_4 + A_4A_3A_1$ . The first term on the right of Eq. (A2),  $\langle \hat{A}_1^n \hat{A}_4 | \hat{A}_3^\dagger (\hat{A}_2^\dagger)^n \rangle$ , with two impurity pairs, is the most complicated overlap to evaluate in Eq. (A2).

## 2. Derivatives of Hamiltonian matrix elements

To calculate derivatives  $\partial \langle \hat{A}_1^n | \hat{H} | (\hat{A}_2^{\dagger})^n \rangle / \partial \langle A_1 \rangle_{\alpha\beta}$  and  $\partial \langle \hat{A}_1^n | (\hat{A}_2^{\dagger})^n \rangle / \partial \langle A_1 \rangle_{\alpha\beta}$ , we borrow from Ref. [19] a general formalism for derivatives. Let  $\hat{O}$  be an operator, e.g., 1 for overlap or  $\hat{H}$  for Hamiltonian matrix elements. The matrix element of  $\hat{O}$  on two pair condensates is

$$\langle \hat{A}_1^n | \hat{O} | (\hat{A}_2^{\dagger})^n \rangle.$$
 (A3)

Taking  $\hat{A}_1(\lambda) = \hat{A}_1 + \lambda \hat{\Gamma}$ , where  $\lambda$  is small (and, we assume, real), yields to first order

$$\langle (\hat{A}_1 + \lambda \hat{\Gamma})^n | \hat{O} | (\hat{A}_2^{\dagger})^n \rangle$$
  
=  $\langle \hat{A}_1^n | \hat{O} | (\hat{A}_2^{\dagger})^n \rangle + n\lambda \langle \hat{A}_1^{n-1} \hat{\Gamma} | \hat{O} | (\hat{A}_2^{\dagger})^n \rangle, \quad (A4)$ 

hence

n

$$\frac{\partial}{\partial\lambda} \langle \hat{A}_1^n(\lambda) | \hat{O} | (\hat{A}_2^{\dagger})^n \rangle \bigg|_{\lambda=0} = n \langle \hat{A}_1^{n-1} \hat{\Gamma} | \hat{O} | (\hat{A}_2^{\dagger})^n \rangle.$$
(A5)

Now  $\hat{\Gamma}$  plays the role of an impurity pair. The derivative of an overlap between pair condensates is simply

$$\frac{\partial}{\partial\lambda}\langle (\hat{A}_1(\lambda))^n | (\hat{A}_2^{\dagger})^n \rangle = n \langle (\hat{A}_1)^{n-1} \Gamma | (\hat{A}_2^{\dagger})^n \rangle, \qquad (A6)$$

while the derivative of a one-body matrix element [Eq. (A1)] is

$$\frac{\partial}{\partial\lambda} \langle (\hat{A}_1)^n | \hat{\mathcal{Q}} | (\hat{A}_2^{\dagger})^n \rangle = n \langle (\hat{A}_1)^{n-1} \Gamma | [\hat{\mathcal{Q}}, \hat{A}_2^{\dagger}] (\hat{A}_2^{\dagger})^{n-1} \rangle = n \langle (\hat{A}_1)^{n-1} \Gamma | \hat{A}_3^{\dagger} (\hat{A}_2^{\dagger})^{n-1} \rangle, \tag{A7}$$

where  $\hat{A}_3^{\dagger} \equiv [\hat{Q}, \hat{A}_2^{\dagger}]$ , with pair structure coefficients  $A_3 = A_2 Q^{\top} + Q A_2$ . Finally, the derivative of two-body matrix element [Eq. (A2)] is

$$\frac{\sigma}{\partial\lambda} \langle (\hat{A}_{1})^{n} | \hat{A}_{3}^{\dagger} \hat{A}_{4} | (\hat{A}_{2}^{\dagger})^{n} \rangle = n \langle \hat{A}_{1}^{n-1} \hat{\Gamma} | \hat{A}_{3}^{\dagger} \hat{A}_{4} | (\hat{A}_{2}^{\dagger})^{n} \rangle 
= n \langle \hat{A}_{1}^{n-1} \hat{\Gamma} \hat{A}_{4} | \hat{A}_{3}^{\dagger} (\hat{A}_{2}^{\dagger})^{n} \rangle 
+ n tr(A_{3}A_{4}) \langle \hat{A}_{1}^{n-1} \hat{\Gamma} | (\hat{A}_{2}^{\dagger})^{n} \rangle 
- n(n-1) \langle \hat{A}_{1}^{n-2} \Gamma \hat{A}_{5} | (\hat{A}_{2}^{\dagger})^{n} \rangle,$$
(A8)

where  $A_5 = A_1A_3A_4 + A_4A_3A_1$ . On the right side, the most complicated overlap in this work is the first term, with three impurity pairs. Thus the derivatives of Hamiltonian matrix elements can be computed from overlaps of pair condensates with impurity pairs, as discussed in Appendix B.

#### **APPENDIX B: OVERLAP FORMULAS**

In this Appendix, we give formulas for overlaps between pair condensates each with up to three impurity pairs; these are needed for Appendix A and Sec. II. We begin with a general formula that expresses the overlap for n-pair states in terms of matrix traces [35,70],

$$\langle \hat{A}_{1} \cdots \hat{A}_{n} | \hat{B}_{1}^{\dagger} \cdots \hat{B}_{n}^{\dagger} \rangle = \sum_{k=1}^{n} (-2)^{-k} \sum_{m_{1} \ge m_{2} \ge \cdots \ge m_{k}}^{m_{1} + \cdots + m_{k} = n} \prod_{w \in \{m_{1}, \cdots, m_{k}\}}^{k} (R_{w}!)^{-1}$$

$$\times \sum_{\{j_{1}, \dots, j_{n}\}, \{i_{1}, \dots, i_{n}\}} \operatorname{tr} (A_{i_{1}} B_{j_{1}} \cdots A_{i_{m_{1}}} B_{j_{m_{1}}}) \cdots \operatorname{tr} (A_{i_{n+1}-m_{k}} B_{j_{n+1}-m_{k}} \cdots A_{i_{n}} B_{j_{n}}).$$
(B1)

Here  $\{i_1, \ldots, i_n\}$  and  $\{j_1, \ldots, j_n\}$  are rearrangements of  $\{1, 2, \ldots, n\}$ ,  $m_1 \ge m_2 \ge \cdots \ge m_k$  is a partition of *n*, and  $R_i$  is the degeneracy of  $m_i$  in a partition, i.e., the number of times  $m_i$  is repeated.

While one can carry out Eq. (B1) either by commutation relations [32,70] or by a generating function [35], here we consider the important special case of a collective pair condensate, i.e., all *n* pairs in a state are identical:

$$\langle (\hat{A}_1)^n | (\hat{A}_2^{\dagger})^n \rangle = (n!)^2 \sum_{k=1}^n (-2)^{-k} \sum_{m_1 \ge m_2 \ge \cdots \ge m_k}^{m_1 + \cdots + m_k = n} \prod_{w \in \{m_1, \cdots, m_k\}}^k m_w^{-1} \prod_{w \in \{m_1, \cdots, m_k\}} (R_w!)^{-1} \cdots$$

$$\times \operatorname{tr}[(A_1 A_2)^{m_1}] \cdots \operatorname{tr}[(A_1 A_2)^{m_k}].$$
(B2)

Therefore, to calculate one overlap between condensates as in Eq. (B2), we calculate  $tr(A_1A_2), \dots, \dots tr[(A_1A_2)^n]$ , and sum over all ordered partitions  $\{m_1, \dots, m_k\}$ . The computation time is polynomial in *n*. Note when *n* is large, e.g., n = 16 in  $2s1d0g_{7/2}0h_{11/2}$  shell ( $N, Z \in [50, 82]$ ), the summation of (B2) can have cancellations of large numbers, requiring careful consideration of round-off error.

Similarly, from Eq. (B1), we can obtain overlap formulas with impurity pairs. Our work here requires overlap formulas with at most three different impurity pairs:

$$\begin{split} \langle (\hat{A}_{1})^{n-1}A_{3}|(\hat{A}_{2}^{\dagger})^{n} \rangle &= (n-1)!n! \sum_{k=1}^{N} (-2)^{-k} \sum_{m_{1} \geq m_{2} \geq m_{2} \geq m_{2} \leq m_{1} \leq m_{1}}^{m_{1}} m_{w}^{-1} \prod_{w \in [m_{1}, \dots, m_{1}]}^{m_{1}} (R_{w}!)^{-1} \sum_{r=1}^{k} m_{r} \\ &\times \operatorname{tr}[(A_{1}A_{2})^{m_{1}}] \cdots \operatorname{tr}[(A_{1}A_{2})^{m_{r}-1}A_{3}A_{2}] \cdots \operatorname{tr}[(A_{1}A_{2})^{m_{1}}]. \tag{B3} \\ \langle (\hat{A}_{1})^{n-1}A_{3}|A_{4}^{\dagger}(\hat{A}_{2}^{\dagger})^{n-1} \rangle &= [(n-1)!]^{2} \sum_{k=1}^{n} (-2)^{-k} \sum_{m_{1} \geq m_{2} \geq m_{2} \geq m_{2} \leq m_{1} \leq m_{1}}^{m_{1} \leftarrow m_{2} = m_{1}} \prod_{w \in [m_{1}, \dots, m_{k}]}^{m_{1}} (R_{w}!)^{-1} \\ &\times \left\{ \sum_{r=1}^{k} m_{r} \sum_{s=0}^{m_{1}-1} \cdots \operatorname{tr}[A_{3}(A_{2}A_{1})^{s}A_{4}(A_{1}A_{2})^{m_{r}-1-s}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{s} \cdots \operatorname{tr}[(A_{1}A_{2})^{m_{1}-1}A_{3}A_{2}] \cdots \operatorname{tr}[A_{1}A_{4}(A_{1}A_{2})^{m_{1}-1-s}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{s} \cdots \operatorname{tr}[(A_{1}A_{2})^{m_{1}-1}A_{3}A_{2}] \cdots \operatorname{tr}[A_{1}A_{4}(A_{1}A_{2})^{m_{1}-1-s}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{s} \cdots \operatorname{tr}[(A_{1}A_{2})^{m_{1}-1}A_{3}A_{2}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1-s}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{s} \cdots \operatorname{tr}[(A_{1}A_{2})^{m_{1}-1}A_{3}A_{2}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1-s}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{r} \sum_{s = 0}^{m_{1}-2} \cdots \operatorname{tr}[A_{3}A_{2}(A_{1}A_{2})^{s}A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{r} \sum_{s = 0}^{m_{1}-2} \cdots \operatorname{tr}[A_{3}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{r} \sum_{s = 0}^{m_{1}-1} \cdots \operatorname{tr}[A_{4}(A_{2}A_{1})^{l}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{r} \sum_{s = 0}^{m_{1}-1} \cdots \operatorname{tr}[A_{4}(A_{2}A_{1})^{l}A_{2}(A_{1}A_{2})^{m_{1}-1}] \cdots \operatorname{tr}[A_{4}A_{4}(A_{1}A_{2})^{m_{1}-1}] \cdots \\ &+ \sum_{l \leq r, s \leq k}^{r \geq m_{1}} m_{r} \sum_{s = 0}^{m_{1}-1} \cdots \operatorname{tr}[A_{4}(A_{2}A_{1})^{l}A_{2}(A_{1}A_{2})^{m_{1$$

where  $[A_5]_{2t}$  means  $A_5$  is at position 2t, and  $[A_4]_{2s+1}$  means  $A_4$  is at position 2s + 1 in the matrix product.

- A. N. Bohr and B. R. Mottelson, *Nuclear Structure* (World Scientific Publishing Company, Singapore, 1998), in two volumes.
- [2] P. Ring and P. Shuck, *The Nuclear Many-Body Problem* (Springer-Verlag, Berlin, Heidelberg, New York, 1980).
- [3] S. Frauendorf and F. Dönau, Phys. Rev. C 89, 014322 (2014).
- [4] A. D. Ayangeakaa, U. Garg, M. D. Anthony, S. Frauendorf, J. T. Matta, B. K. Nayak, D. Patel, Q. B. Chen, S. Q. Zhang, P. W. Zhao, B. Qi, J. Meng, R. V. F. Janssens, M. P. Carpenter, C. J. Chiara, F. G. Kondev, T. Lauritsen, D. Seweryniak, S. Zhu, S. S. Ghugre, and R. Palit, Phys. Rev. Lett. **110**, 172504 (2013).
- [5] C. Xu, G. Röpke, P. Schuck, Z. Ren, Y. Funaki, H. Horiuchi, A. Tohsaki, T. Yamada, and B. Zhou, Phys. Rev. C 95, 061306(R) (2017).
- [6] M. R. Mumpower, R. Surman, G. C. McLaughlin, and A. Aprahamian, Prog. Part. Nucl. Phys. 86, 86 (2016).
- [7] T. Kajino, W. Aoki, A. B. Balantekin, R. Diehl, M. A. Famiano, and G. J. Mathews, Prog. Part. Nucl. Phys. 107, 109 (2019).
- [8] J. Meng, International Review of Nuclear Physics: Relativistic Density Functional for Nuclear Structure (World Scientific, 2016), Vol. 10.
- [9] B. Bally and M. Bender, Phys. Rev. C 103, 024315 (2021).
- [10] A. Sánchez-Fernández, B. Bally, and T. R. Rodríguez, Phys. Rev. C 104, 054306 (2021).
- [11] B. Bally, A. Sánchez-Fernández, and T. R. Rodríguez, Phys. Rev. C 100, 044308 (2019).
- [12] N. Shimizu, T. Mizusaki, K. Kaneko, and Y. Tsunoda, Phys. Rev. C 103, 064302 (2021).
- [13] J. L. Egido and L. M. Robledo, Phys. Rev. Lett. 70, 2876 (1993).
- [14] J. Berger, M. Girod, and D. Gogny, Nucl. Phys. A 428, 23 (1984).
- [15] M. Bender and P.-H. Heenen, Phys. Rev. C 78, 024309 (2008).
- [16] N. Schunck, J. Dobaczewski, W. Satuła, P. Bączyk, J. Dudek, Y. Gao, M. Konieczka, K. Sato, Y. Shi, X. Wang, and T. Werner, Comput. Phys. Commun. 216, 145 (2017).
- [17] F. Iachello and A. Arima, *The Interacting Boson Model* (Cambridge University Press, 2006).
- [18] Y. Sun, K. Hara, J. A. Sheikh, J. G. Hirsch, V. Velázquez, and M. Guidry, Phys. Rev. C 61, 064323 (2000).
- [19] T. Otsuka, M. Honma, T. Mizusaki, N. Shimizu, and Y. Utsuno, Prog. Part. Nucl. Phys. 47, 319 (2001).
- [20] E. Caurier, G. Martínez-Pinedo, F. Nowacki, A. Poves, and A. P. Zuker, Rev. Mod. Phys. 77, 427 (2005).
- [21] B. Brown, Prog. Part. Nucl. Phys. 47, 517 (2001).
- [22] B. A. Brown and W. A. Richter, Phys. Rev. C 74, 034315 (2006).
- [23] M. Honma, T. Otsuka, B. A. Brown, and T. Mizusaki, Phys. Rev. C 65, 061301(R) (2002).
- [24] M. Honma, T. Otsuka, B. A. Brown, and T. Mizusaki, Eur. Phys. J. A 25, 499 (2005).
- [25] I. Talmi, Simple Models of Complex Nuclei (Contemporary Concepts in Physics) (CRC Press, 1993).
- [26] L. Y. Jia, Phys. Rev. C 99, 014302 (2019).
- [27] L. Y. Jia, Phys. Rev. C 93, 064307 (2016).
- [28] L. Y. Jia, J. Phys. G 42, 115105 (2015).
- [29] Y. M. Zhao and A. Arima, Phys. Rep. 545, 1 (2014).
- [30] K. Higashiyama, N. Yoshinaga, and K. Tanabe, Phys. Rev. C 65, 054317 (2002).
- [31] B. C. He, L. Li, Y. A. Luo, Y. Zhang, F. Pan, and J. P. Draayer, Phys. Rev. C 102, 024304 (2020).
- [32] Y. Lei, Y. Lu, and Y. M. Zhao, Chin. Phys. C 45, 054103 (2021).
- [33] T. Otsuka, A. Arima, F. Iachello, and I. Talmi, Phys. Lett. B 76, 139 (1978).

- [34] C. W. Johnson and J. N. Ginocchio, Phys. Rev. C 50, R571 (1994).
- [35] J. N. Ginocchio and C. W. Johnson, Phys. Rev. C 51, 1861 (1995).
- [36] G. Fu and C. W. Johnson, Phys. Lett. B 809, 135705 (2020).
- [37] G. J. Fu and C. W. Johnson, Phys. Rev. C 104, 024312 (2021).
- [38] Y. Lei, H. Jiang, and S. Pittel, Phys. Rev. C 102, 024310 (2020).
- [39] T. Mizusaki and P. Schuck, Phys. Rev. C 104, L031305 (2021).
- [40] A. J. Coleman, J. Math. Phys. 6, 1425 (1965).
- [41] D. J. Rowe, T. Song, and H. Chen, Phys. Rev. C 44, R598 (1991).
- [42] H. Chen, T. Song, and D. Rowe, Nucl. Phys. A 582, 181 (1995).
- [43] C. W. Johnson and K. D. O'Mara, Phys. Rev. C 96, 064304 (2017).
- [44] C. W. Johnson and C. Jiao, J. Phys. G 46, 015101 (2019).
- [45] C. Qi and Z. X. Xu, Phys. Rev. C 86, 044323 (2012).
- [46] M. Galassi, J. Davies, J. Theiler, B. Gough, G. Jungman, P. Alken, M. Booth, F. Rossi, and R. Ulerich, *GNU Scientific Library* (gnu, 2019).
- [47] R. Fletcher, SIAM Rev. 26, 143 (1984).
- [48] Y. Sun, Phys. Scr. 91, 043005 (2016).
- [49] J. M. Yao, J. Meng, P. Ring, and D. Vretenar, Phys. Rev. C 81, 044311 (2010).
- [50] R. Rodríguez-Guzmán, J. Egido, and L. M. Robledo, Nucl. Phys. A 709, 201 (2002).
- [51] A. R. Edmonds, Angular Momentum in Quantum Mechanics (Princeton University Press, Princeton, New Jersey, 1996).
- [52] J. A. Sheikh, P. Ring, E. Lopes, and R. Rossignoli, Phys. Rev. C 66, 044318 (2002).
- [53] C. W. Johnson, W. Ormand, and P. G. Krastev, Comput. Phys. Commun. 184, 2761 (2013).
- [54] S. M. Lauber, H. C. Frye, and C. W. Johnson, J. Phys. G 48, 095107 (2021).
- [55] C. W. Johnson, K. A. Luu, and Y. Lu, J. Phys. G 47, 105107 (2020).
- [56] N. Nica and B. Singh, Nucl. Data Sheets 113, 1563 (2012).
- [57] S. Kanno, T. Gomi, T. Motobayashi, K. Yoneda, N. Aoi, Y. Ando, H. Baba, K. Demichi, Z. Fulop, U. Futakami, H. Hasegawa, Y. Higurashi, K. Ieki, N. Imai, N. Iwasa, H. Iwasaki, T. Kubo, S. Kubono, M. Kunibu, Y. U. Matsuyama, S. Michimasa, T. Minemura, H. Murakami, T. Nakamura, A. Saito, H. Sakurai, M. Serata, S. Shimoura, T. Sugimoto, E. Takeshita, S. Takeuchi, K. Ue, K. Yamada, Y. Yanagisawa, A. Yoshida, and M. Ishihara, Prog. Theor. Phys. **146**, 575 (2002).
- [58] M. Basunia and A. Hurst, Nucl. Data Sheets 134, 1 (2016).
- [59] M. Shamsuzzoha Basunia, Nucl. Data Sheets **114**, 1189 (2013).
- [60] M. Shamsuzzoha Basunia, Nucl. Data Sheets 111, 2331 (2010).
- [61] J. P. Dufour, R. Del Moral, A. Fleury, F. Hubert, D. Jean, M. S. Pravikoff, H. Delagrange, H. Geissel, and K. H. Schmidt, Z. Phys. A 324, 487 (1986).
- [62] S.-C. Wu, Nucl. Data Sheets **91**, 1 (2000).
- [63] T. Burrows, Nucl. Data Sheets 107, 1747 (2006).
- [64] J. Chen and B. Singh, Nucl. Data Sheets 157, 1 (2019).
- [65] Y. Dong and H. Junde, Nucl. Data Sheets 128, 185 (2015).
- [66] F. Stephens and R. Simon, Nucl. Phys. A 183, 257 (1972).
- [67] J. L. Wood, A.-M. Oros-Peusquens, R. Zaballa, J. M. Allmond, and W. D. Kulp, Phys. Rev. C 70, 024308 (2004).
- [68] J. Blachot, Nucl. Data Sheets **108**, 2035 (2007).
- [69] D. De Frenne and A. Negret, Nucl. Data Sheets 109, 943 (2008).
- [70] B. Silvestre-Brac and R. Piepenbring, Phys. Rev. C 26, 2640 (1982).