Isoscalar giant monopole resonance in ²⁴Mg and ²⁸Si: Effect of coupling between the isoscalar monopole and quadrupole strength

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Background: In highly deformed nuclei, there is a noticeable coupling of the isoscalar giant monopole resonance (ISGMR) and the K = 0 component of the isoscalar giant quadrupole resonance (ISGQR), which results in a double peak structure of the isoscalar monopole (IS0) strength (a narrow low-energy deformation-induced peak and a main broad ISGMR part). The energy of the narrow low-lying IS0 peak is sensitive to both the incompressibility modulus K_{∞} and the coupling between IS0 and isoscalar quadrupole (IS2) strength.

Purpose: This study aims to investigate the two-peaked structure of the ISGMR in the prolate ²⁴Mg and oblate ²⁸Si nuclei and identify among a variety of energy density functionals based on Skyrme parametrizations the one which best describes the experimental data. This will allow for conclusions regarding the nuclear incompressibility. Because of the strong IS0/IS2 coupling, the deformation splitting of the ISGQR will also be analyzed.

Methods: The ISGMR was excited in ²⁴Mg and ²⁸Si using α -particle inelastic scattering measurements acquired with an $E_{\alpha} = 196$ MeV beam at scattering angles $\theta_{Lab} = 0^{\circ}$ and 4° . The K600 magnetic spectrometer at iThemba LABS was used to detect and momentum analyze the inelastically scattered α particles. An experimental energy resolution of \approx 70 keV (FWHM) was attained, revealing fine structure in the excitation-energy region of the ISGMR. The ISO strength distributions in the nuclei studied were obtained with the difference-of-spectra (DoS) technique. The theoretical comparison is based on the quasiparticle random-phase approximation (QRPA) with a representative set of Skyrme forces.

Results: IS0 strength distributions for ²⁴Mg and ²⁸Si are extracted and compared to previously published results from experiments with a lower energy resolution. With some exceptions, a reasonable agreement is obtained. The IS0 strength is found to be separated into a narrow structure at about 13–14 MeV in ²⁴Mg, 17–19 MeV in ²⁸Si, and a broad structure at 19–26 MeV in both nuclei. The data are compared with QRPA results. The results of the calculated characteristics of IS0 states demonstrate the strong IS0/IS2 coupling in strongly prolate ²⁴Mg and oblate ²⁸Si. The narrow IS0 peaks are shown to arise due to the deformation-induced IS0/IS2 coupling and strong collective effects. The cluster features of the narrow IS0 peak at 13.87 MeV in ²⁴Mg are also discussed. The best description of the IS0 data is obtained using the Skyrme force SkP⁸ with an associated low nuclear incompressibility $K_{\infty} = 202$ MeV allowing for both the energy of the peak and integral IS0 strength in ²⁴Mg and ²⁸Si to be reproduced. The features of the ISGQR in these nuclei are also investigated. An anomalous deformation splitting of the ISGQR in oblate ²⁸Si is found. The observed structure of ISGQR in ²⁴Mg is described.

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Conclusions: The ISGMR and ISGQR in light deformed nuclei are coupled and thus need to be described simultaneously. Only such a description is relevant and consistent. The deformation-induced narrow IS0 peaks can serve as an additional sensitive measure of the nuclear incompressibility.

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I. INTRODUCTION

For many years, the isoscalar giant monopole resonance (ISGMR) served as an important source of information on the nuclear incompressibility [1–3]. In deformed nuclei, this resonance demonstrates a remarkable coupling of the isoscalar monopole and quadrupole (K = 0) modes, which leads to a double peak structure of the ISO strength [4-10] (here, K stands for the projection of the total angular momentum on the symmetry axis of an axially deformed nucleus [11]). Strongly deformed light nuclei, like prolate ²⁴Mg and oblate ²⁸Si, are especially attractive for an exploration of the nuclear incompressibility K_{∞} and the role of the ISO/IS2 coupling. Indeed, in these nuclei, one may expect a deformation-induced strong narrow peak in the ISO strength located below the main ISGMR region [12–14]. The energy of this peak should be sensitive to both the incompressibility K_{∞} and the ISO/IS2 coupling strength.

The ISGMR in ²⁴Mg has been explored in (α, α') and (⁶Li, ⁶Li') experiments since the 1980s (see Ref. [2] for the extensive review and discussion). More recent (α , α') experiments were performed at the Research Center for Nuclear Physics (RCNP) for ²⁴Mg [12,14,15] and ²⁸Si [16]. In 2021, the results of an RCNP (⁶Li, ⁶Li') experiment on IS0 strength in ²⁴Mg were published [17]. The ISO/IS2 coupling was mentioned in Refs. [12,17] in order to explain the existence of the ISO structure located around 16–19 MeV excitation in ²⁴Mg. The existence of a prominent ISO peak at about 18 MeV in oblate ²⁸Si was noted but not yet addressed [16]. These results generally confirm previous findings from (α, α') experiments performed by the Texas A&M University (TAMU) group on ²⁴Mg [18] (for the dataset where the out-of-plane angle was measured) and ²⁸Si [19]. For ²⁸Si, discrepancies between RCNP [16] and TAMU [19] results in the high excitationenergy region above 20 MeV were noted and attributed to the phenomenological background subtraction method employed by TAMU [3].

The ISO/IS2 coupling was earlier considered within the quasiparticle random phase approximation (QRPA) in ²⁴Mg [12–14,17,20,21] and ²⁸Si [21]. However, in the study [21], the narrow ISO peaks in ²⁴Mg and ²⁸Si were not specifically discussed. Furthermore, one can dispute whether the low-energy narrow ISO peak in ²⁴Mg has been correctly assigned (and used in the QRPA analysis) in the RCNP studies [12,14,17] (see discussion in Sec. IV B).

In this paper, the ISGMR in prolate ²⁴Mg and oblate ²⁸Si is explored using α -particle inelastic scattering measurements with an $E_{\alpha} = 196$ MeV beam measured at scattering angles $\theta_{Lab} = 0^{\circ}$ and 4° using the high energy-resolution K600 magnetic spectrometer at the iThemba Laboratory for Accelerator Based Science (LABS), South Africa. The study focuses on the deformation-induced splitting of the ISGMR into two parts: a narrow peak arising due to ISO/IS2 coupling, and a wide structure representing the main ISGMR. The former exists only in deformed nuclei while the latter appears in both spherical and deformed nuclei.

By comparing the obtained experimental data with QRPA calculations, we will identify the narrow IS0 peak in ²⁴Mg, explain the origin of the narrow IS0 peaks in both prolate ²⁴Mg and oblate ²⁸Si, and suggest a Skyrme force suitable for the description of the IS0 strength in these nuclei. Moreover, we will analyze in detail the IS2 strength, which determines the energy of the narrow IS0 peak, and show that, in strongly deformed nuclei, only a simultaneous description of the IS0 and IS2 strengths can be considered as consistent. The simultaneous availability of data for a prolate (²⁴Mg) and an oblate (²⁸Si) case allows the conclusions to be generalized.

The paper is organized as follows: in Sec. II, the experimental details are given. The method of extraction and the resulting ISO strength distributions are presented in Sec. III while, in Sec. IV, details of the QRPA calculations are given and the comparison with the data is shown and discussed. In Sec. V, conclusions are drawn and finally, in the Appendix, the predicted deformation splitting of the ISGQR in ²⁴Mg and ²⁸Si is presented.

II. EXPERIMENTAL DETAILS

The experimental procedure followed in this study is similar to that described elsewhere [21,22]. As such, only salient details are provided here. The experiment was performed at the Separated Sector Cyclotron (SSC) facility of iThemba LABS using a beam of 196 MeV α particles. Inelastically scattered α particles were momentum analyzed by the K600 magnetic spectrometer after interacting with either a 0.23 mg/cm² thick ²⁴Mg or a 0.23 mg/cm² thick ²⁸Si foil. The horizontal and vertical positions of the scattered α particles in the focal plane of the spectrometer were measured using two multiwire drift chambers. Energy deposition in the plastic scintillators in the focal plane as well as time-of-flight measurements relative to the cyclotron radio frequency were used for particle identification.

Spectra were acquired with the spectrometer set at scattering angles 0° and 4° . In the former, scattering angles $\theta_{Lab} = 0^{\circ} \pm 1.91^{\circ}$ and in the latter, scattering angles from $\theta_{Lab} = 2^{\circ}$ to 6° were covered by a circular spectrometer aperture, respectively. The trajectory of the scattered particles through the focal plane was used to reconstruct the scattering angle. A multihole collimator was used to calibrate the measured focalplane angles to scattering trajectories into the spectrometer.

In the zero-degree mode, both the unscattered beam and the inelastically scattered particles are transported through the K600 magnetic spectrometer with the beam exiting the spectrometer only a few centimeters away from the drift chamber position in the high-dispersion focal plane mode of the

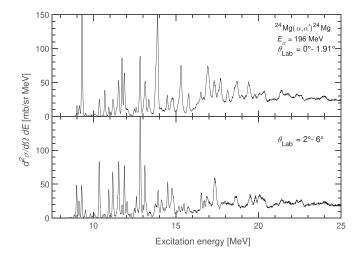


FIG. 1. Double-differential cross sections for the ²⁴Mg(α, α') reaction at $E_{\alpha} = 196$ MeV with $\theta_{\text{Lab}} = 0^{\circ}-1.91^{\circ}$ (top) and $\theta_{\text{Lab}} = 2^{\circ}-6^{\circ}$ (bottom).

K600. For such measurements, suppression of beam-induced background is critical. In order to subtract the instrumental background, a standard method used at the iThemba LABS K600 [23] and the RCNP magnetic spectrometers [24] was employed that exploits differences in the vertical distribution of real and background events. This method establishes background spectra from the regions of the focal plane above and below the vertically focused band of true events and allows a direct subtraction of these components from the central region of interest. In contrast to the zero-degree mode, improved instrumental background conditions in the small-angle mode allows to operate the spectrometer in vertical off-focus mode and also use the medium-dispersion focal plane. In the offfocus mode, the vertical position on the focal plane relates to the vertical component of the scattering angle into the spectrometer aperture and allows for its reconstruction [24].

The same techniques as employed in Ref. [22] were adopted for the analysis of the data. This includes software corrections of kinematic effects and optical aberrations in the horizontal focal-plane position that depend on the vertical focal-plane position and the scattering angle into the spectrometer. The energy calibration was based on well-known states in ²⁴Mg [15,25]. An energy resolution of \approx 70 keV (FWHM) was attained. Figures 1 and 2 show the inelastic scattering cross sections extracted at 0°–1.91° and 2°–6° for ²⁴Mg and ²⁸Si, respectively. The broad structure seen in Fig. 2 below 10 MeV is due to hydrogen contaminants in the targets. However, it does not affect the excitation-energy ranges relevant for the extraction of the ISO strength as demonstrated in the next section.

III. EXPERIMENTAL RESULTS

The multipole decomposition analysis (MDA) technique was employed in numerous studies to extract multipole strength distributions in nuclei, including the ISO strength distributions [3,12,14,16–19,26,27]. However, due to the limited number of angular data points in the present study, only the

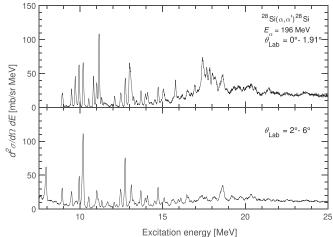


FIG. 2. Same as Fig. 1, but for ²⁸Si.

ISO strength distributions could be determined in this study via the difference-of-spectra (DoS) method [28]. In the DoS method, an excitation-energy spectrum for the angular region associated with the first minimum of the L = 0 angular distribution is subtracted from the excitation-energy spectrum taken at 0°. This method, therefore, requires the determination of the suitable angle cut from the measurement at $\theta_{Lab} = 2^{\circ}-6^{\circ}$, which is obtained from distorted-wave born approximation (DWBA) calculations.

In the present study, the DWBA calculations were performed according to the method described in Ref. [29]. A density-dependent single-folding model for the real part of the potential U(r), obtained with a Gaussian α -nucleon potential, and a phenomenological Woods-Saxon potential for the imaginary term of U(r) were used, so that the α -nucleus potential can be written as

$$U(r) = V_{\text{fold}}(r) + i \frac{W}{\{1 + \exp[(r - R_{\text{I}})/a_{\text{I}}]\}},$$
 (1)

with radius $R_{\rm I} = r_{0\rm I}(A_{\rm p}^{1/3} + A_{\rm t}^{1/3})$ and diffuseness $a_{\rm I}$. The subscripts p and t refer to projectile and target, respectively, and A denotes the mass number. The potential $V_{\rm fold}(r)$ is obtained by folding the ground-state density with a density-dependent α -nucleon interaction

$$V_{\text{fold}}(r) = -V \int d^3 r' \rho(r') [1 - \beta \rho(r')^{2/3}] \exp(-z^2/t^2), \quad (2)$$

where z = |r - r'| is the distance between the center of mass of the α particle and a target nucleon, and $\rho(r')$ is the groundstate density of the target nucleus at the position r' of the target nucleon. The parameters $\beta = 1.9$ fm² and range t = 1.88 fm were taken from Ref. [29]. The ground-state density $\rho(r)$ of the target nucleus at the position r is given by

$$\rho(r) = \frac{\rho_0}{1 + \exp\left(\frac{r-c}{a}\right)},\tag{3}$$

where the Fermi-distribution parameters c and a describe the half-density radius and the diffuseness, respectively. Numerical values for ²⁴Mg and ²⁸Si were taken from Ref. [30].

TABLE I. Optical model parameters taken from the TAMU group publications.

Nucleus	V (MeV)	W (MeV)	r_{0I} (fm)	$a_{\rm I}$ (fm)	Ref.
²⁴ Mg	41.02	35.39	0.934	0.614	[34]
²⁸ Si	44	32.5	0.9306	0.687	[35]

DWBA calculations were carried out using the computer code PTOLEMY [31,32]. The optical model parameters (OMP) used in the DWBA calculations, viz. r_{0I} , a_I , V, and W, for each nucleus are usually obtained by fitting elastic scattering cross sections. The procedure is fully described in Ref. [33]. However, due to the lack of elastic scattering data in this study, the parameters were taken from studies of the TAMU group on the same nuclei. This is justified by the similarity of their beam energy (240 MeV) compared to this study. The OMP parameters are summarized in Table I.

The measured cross sections can be converted to fractions (a_0) of the energy-weighted sum rule EWSR(IS0) = $2\hbar^2/m \times A\langle r^2 \rangle$ by comparing with DWBA calculations assuming 100% EWSR, which are shown in Fig. 3. The IS0 strength is then calculated using the a_0 values and is expressed as

$$S_0(E_{\rm x}) = \frac{\rm EWSR(IS0)}{E_{\rm x}}a_0(E_{\rm x}) = \frac{2\hbar^2 A \langle r^2 \rangle}{mE_{\rm x}}a_0(E_{\rm x}), \quad (4)$$

where *m* is the nucleon mass, E_x is the excitation energy corresponding to a given state or energy bin, and $\langle r^2 \rangle$ is the second moment of the ground-state density. We use $\langle r^2 \rangle = 9.345 \text{ fm}^2$ for ²⁴Mg and 9.753 fm² for ²⁸Si [30].

The DoS technique relies on the subtraction of the contributions of all multipoles except L = 0 from the 0° spectrum using data from an angle range defined by the first minimum of the L = 0 contribution. It was shown that the sum of angular distributions of all multipoles L > 0 is similar at the maximum and first minimum of the L = 0 distribution [3,28]. Hence, the subtraction of the inelastic spectrum associated with the flat angular distribution (where L = 0 is at a minimum from the 0° spectrum) is assumed to represent essentially the IS0 component excited in α inelastic scattering close to 0°.

The direct subtraction of the angle cut spectrum from the spectrum at $0^{\circ} \leq \theta_{\text{Lab}} \leq 1.91^{\circ}$ (blue) yields the black difference spectra shown in the upper and lower panels of Fig. 4 for ²⁴Mg and ²⁸Si, respectively. The angle cut for each nucleus indicated by the highlighted yellow area in Fig. 3 and listed in Table II was chosen to be statistically significant while allowing for minimal inclusion of the L = 0 component. The

TABLE II. Angle cuts implemented in the 4° dataset to define the angular region around the first L = 0 minimum.

Nucleus	²⁴ Mg	²⁸ Si
Angle cut (θ_{Lab})	3.39°-4.4°	3.03°-4.07°
Angle cut ($\theta_{c.m.}$)	4°-5.2°	3.5°-4.7°
$L = 0$ minimum ($\theta_{c.m.}$)	4.6°	4.1°

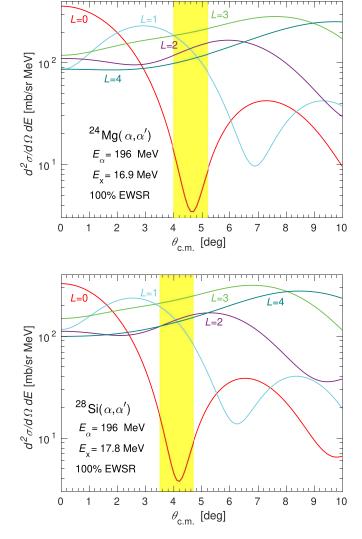


FIG. 3. DWBA calculations of the differential cross sections for the ²⁴Mg(α , α') (top panel) and ²⁸Si(α , α') (bottom panel) reaction at $E_{\alpha} = 196$ MeV for various isoscalar electric multipoles. The calculations have been normalized to 100% of the appropriate EWSR at an excitation energy of 16.9 MeV in ²⁴Mg and 17.8 MeV in ²⁸Si, representative of the maximum of the ISO strength distributions.

angle corresponding to the minimum of the L = 0 distribution is also indicated in Table II.

ISO strength distributions in ²⁴Mg and ²⁸Si were determined from the difference spectra. The fraction of ISO strength in EWSR per MeV was obtained by dividing the extracted experimental ISO component by the corresponding integral value of L = 0, 100% EWSR obtained at the average angles $\theta_{c.m.}^{av.} = 1.13^{\circ}$ (²⁴Mg) and 1.12° (²⁸Si) in the 0° spectra. The results are shown in Figs. 5 and 6 for 0.5 MeV wide energy bins. The uncertainties shown include both systematic and statistical errors. For comparison purposes, the ISO strengths from the RCNP experiments ([12,14,15] for ²⁴Mg and [16] for ²⁸Si) and the TAMU experiments ([18] for ²⁴Mg and [19] for ²⁸Si) are also shown. The results from Ref. [18] are for the dataset where the out-of-plane angle was measured. In Refs. [18,19], the ISGMR data were given as fractions

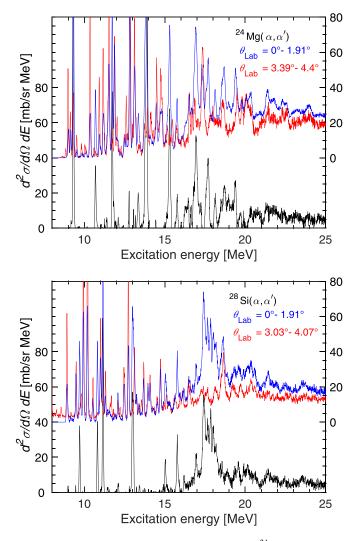


FIG. 4. Double-differential cross sections for ²⁴Mg(α , α') (top panel) and ²⁸Si(α , α') (bottom panel) at $E_{\alpha} = 196$ MeV. The blue and red spectra represent the data acquired at $0^{\circ} \leq \theta_{\text{Lab}} \leq 1.91^{\circ}$ and at $3.39^{\circ} \leq \theta_{\text{Lab}} \leq 4.4^{\circ}$ for ²⁴Mg and $3.03^{\circ} \leq \theta_{\text{Lab}} \leq 4.07^{\circ}$ for ²⁸Si, respectively. For clarity, these spectra are shifted by 40 mb/sr MeV. The black spectra represent the difference calculated using the DoS technique (see text).

 a_0 of the sum rule EWSR(IS0). We converted these data into ISO strength using the relation (4). For better visibility the comparison of ISO strength distributions with the present results in Fig. 5 for ²⁴Mg is split into two parts. The upper part shows results from Refs. [12,14,17] and the lower those from Refs. [15,18]. Data for ²⁴Mg from the present study generally agree within error bars with the RCNP and TAMU data. We note that the strengths given in Ref. [15] were extracted with a density-independent interaction for the real part of the transition potential known to lead to much larger values compared to the approach applied in the present work [37]. When appropriately normalized to the present results, good agreement of the two distributions is obtained. Above 17 MeV the ISO strength from the TAMU group is consistently higher than both the iThemba LABS and RCNP strengths. These discrepancies might be attributed to

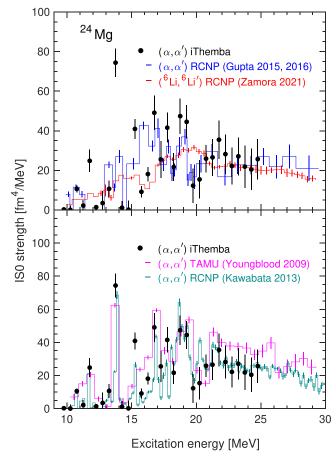


FIG. 5. ISO strength distributions in ²⁴Mg. The present iThemba LABS data are shown as black filled circles. Also shown are the (α, α') [12,14] and (⁶Li, ⁶Li') data [17] from the RCNP shown as blue and red histograms, respectively (top panel). The bottom panel displays results from TAMU [18] (magenta histogram) and a different measurement at RCNP [15] shown by dark cyan histogram.

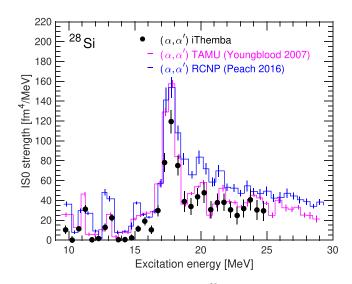


FIG. 6. ISO strength distribution in ²⁸Si. The present iThemba LABS data are shown as black filled circles. Data shown as blue and magenta histograms with vertical error bars are from RCNP [16] and TAMU [19].

the background-subtraction method employed by the TAMU group. At the same time, Fig. 5 shows that the RCNP [12,14], TAMU [18], and present iThemba LABS data for ²⁴Mg give somewhat different strength distributions below the main IS-GMR region. In particular, the RCNP experiments [12,14] suggest the highest ISO structures at 16-20 MeV and a two times lower peak at 13-16 MeV. A similar structure was observed in the recent (⁶Li, ⁶Li') experiment [17]. Instead, the TAMU and iThemba LABS data (bottom panel of Fig. 5) give the highest ISO peak at about 13.8 MeV and a broader peak with a maximum at about 18 MeV. It is remarkable that, unlike the RCNP data [12,14], another (α, α') experiment at RCNP with superior energy resolution [15] does show the strong narrow peak at 13.8 MeV (bottom panel of Fig. 5). Moreover, this peak was clearly observed in the early (α , α' [38] and (⁶Li, ⁶Li') [39] experiments (see the discussion in Sec. 4.2.3 of Ref. [2]). So most of the experimental datasets confirm the existence of the strong narrow ISO peak at 13.8–13.9 MeV in ²⁴Mg. As can be seen in the next section, this peak is of crucial importance for our present study. The differences between the various datasets, especially for strong narrow peaks of ISO strength, can most likely be ascribed to differences in the experimental energy resolution and a subtle interplay between binning effects and differences in the calibration of the excitation energy.

As for ²⁸Si (Fig. 6), the iThemba LABS results are mainly consistent with the previous findings by the TAMU group, except for a slightly lower value of the ISO strength at the peak of the distribution. The results from RCNP, on the other hand, exhibit almost a factor of 2 larger ISO strength above 19 MeV than both the iThemba LABS and the TAMU results. One should also note that some of the ISO strength at lower excitation energy might not belong to the ISGMR. It was suggested that, in ²⁸Si, the 0⁺ states at 9.71, 10.81, 11.14, and 12.99 MeV, i.e., most of the ISO strength below 15 MeV, can be considered as potential bandheads for superdeformed bands [22].

With the good energy resolution of the present experiment, $J^{\pi} = 0^+$ states could be resolved up to 16 MeV in ²⁴Mg and 15 MeV in ²⁸Si. The corresponding strengths *S*(ISO) and % EWSR exhausted by the strongest discrete states evident in Fig. 4 are summarized in Tables III and IV, respectively. The values obtained in this study are in good agreement with results from Ref. [21] which analyzed angular distributions of individual states, employing a different DWBA code as well as OMPs.

IV. QRPA CALCULATIONS AND COMPARISON WITH EXPERIMENT

A. Details of the calculations

The calculations were performed within the QRPA model [40–43] based on the Skyrme functional [44,45]. The model is fully self-consistent. Both the mean field and the residual interaction are derived from the initial Skyrme functional. The residual interaction takes into account all the terms following from the Skyrme functional and Coulomb (direct and exchange) parts. Both particle-hole and particle-particle

TABLE III. Excitation of prominent $J^{\pi} = 0^+$ states in ²⁴Mg. *S*(IS0) is the experimental IS0 strength exhausted by the state.

E _x (MeV) ^a	E _x (MeV) ^b	E _x (MeV) ^c	% EWSR ^a	% EWSR ^e	S(IS0) $(fm^4)^a$
9.31(1)	9.30539(24)		1.0(1)	1.4(3)	19.9(20)
10.68(1)	10.6797(4)		0.31(5)	0.29(6)	5.32(61)
11.73(1)	11.7281(10)		0.77(11)	1.0(2)	12.1(16)
13.36(2)		13.37(1)	0.40(5)	0.5(1)	4.32(71)
		13.79(1)		1.7(3)	
13.87(2) ^d	13.884(1)	13.89(1)	2.8(3) ^d	2.6(5)	37.7(38) ^d
15.32(2)		15.33(3)	1.7(2)	1.9(4)	20.7(25)
15.78(2)		15.79(3)	0.40(5)	1.1(2)	4.6(10)

^aPresent experiment.

^bFrom Ref. [36].

^cFrom Ref. [21].

^dThe integration region for the 13.87(2) state from the present experiment includes both the 13.79 and 13.89 MeV states from Ref. [21].

channels are included [41]. Pairing is treated at the BCS level [41]. Spurious admixtures caused in ISO strength by pairinginduced violation of the particle number and in IS2 strength by violation of the rotational invariance are removed using the method from Ref. [43].

A representative set of Skyrme forces is used, see Table V. For our aims, the most important characteristics of the forces are the nuclear incompressibility K_{∞} and the isoscalar effective mass m_0^*/m affecting ISGMR and ISGQR energies, respectively. The corresponding values are listed in Table V. We employ the standard force SkM* [46] and the most recent force SV-bas [47]. These forces were previously used in the analysis presented in [12–14,16]. Further, we use the early force SkT6 [48] and more recent force SV-mas10 [47] which both have a large effective mass, $m_0^*/m = 1$, and so are favorable for the description of the ISGQR [10,49,50]. Finally, we exploit the force SkP^{δ} [51] with a low incompressibility $K_{\infty} = 202$ MeV. In a recent study of the ISGMR in Mo isotopes, this force gave the best results [10].

The nuclear mean field and pairing are computed by the code SKYAX [52] using a two-dimensional grid in cylindrical coordinates. The calculation box extends up to three nuclear radii and the grid step size is 0.4 fm. The axial quadrupole equilibrium deformation is obtained by minimization of the energy of the system. As shown in Fig. 7, both ²⁴Mg and ²⁸Si

TABLE IV. Same as in Table III but for excitation of $J^{\pi} = 0^+$ states in ²⁸Si.

$\overline{E_{\rm x}~({\rm MeV})^{\rm a}}$	$E_{\rm x}~({\rm MeV})^{\rm b}$	% EWSR ^a	% EWSR ^b	$S(IS0) (fm^4)^a$
9.70(2)	9.71(2)	0.22(4)	0.38(8)	5.2(6)
10.81(2)	10.81(3)	0.27(4)	0.35(7)	5.7(6)
11.14(2)	11.142(1)	0.8(1)	0.9(2)	15.3(17)
13.00(2)	12.99(2)	0.95(12)	0.8(2)	16.9(18)
15.03(3)	15.02(3)	0.40(9)	0.8(2)	6.0(15)
15.77(3)		0.7(1)		9.7(16)

^aPresent experiment.

^bFrom Ref. [22].

TABLE V. Incompressibility K_{∞} and isoscalar effective mass m_0^*/m for the Skyrme forces SV-bas, SkM*, SkP^{δ}, SkT6, and SV-mas10 used in the present analysis.

	SV-bas	SkM*	SkP^δ	SkT6	SV-mas10
$\overline{K_{\infty}}$ (MeV)	234	217	202	236	234
m_0^*/m	0.9	0.79	1	1	1

have oblate and prolate minima. However, in the ground state, ²⁴Mg is prolate and ²⁸Si is oblate. As shown in Table VI, the calculated deformation parameters β somewhat underestimate the experimental values [36]. This is typical for light deformed nuclei; see, e.g., Ref. [53]. For all the applied Skyrme forces, the pairing in ²⁴Mg and ²⁸Si was found to be very weak. This is explained by the sparse single-particle spectrum in light nuclei.

The calculations use a large configuration space. The single-particle spectrum extends from the bottom of the potential well up to 30 MeV. For example, SkP^{δ} calculations use 404 proton and 404 neutron single-particle levels for ²⁴Mg and 422 proton and 422 neutron single-particle levels for ²⁸Si. The two-quasiparticle (2qp) basis in QRPA calculations with SkP^{δ} extends up to 70 MeV. It includes 1056 proton and 956 neutron $K^{\pi} = 0^+$ states in ²⁴Mg and 1142 proton and 1002 neutron $K^{\pi} = 0^+$ states in ²⁸Si. The isoscalar energy-weighted sum rules EWSR(IS0) and EWSR(IS2)= $25\hbar^2/(4\pi m) \times A\langle r^2 \rangle$ are exhausted by 97–99%. QRPA strength functions for isoscalar monopole (L = 0) and quadrupole (L = 2) transitions have the form

$$S_L(E) = \sum_{K=0}^{L} (2 - \delta_{K,0}) \sum_{\nu \in K} |\langle \nu | \hat{O}_{LK} | 0 \rangle|^2 \xi_{\Delta}(E - E_{\nu}), \quad (5)$$

where ν labels QRPA eigenvalues $|\nu\rangle$ with energies E_{ν} , and $|0\rangle$ is the QRPA ground state. The monopole and quadrupole

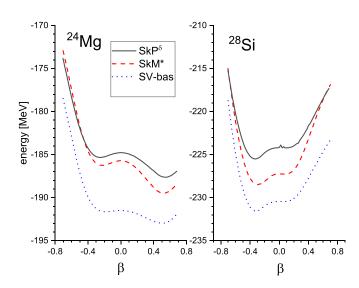


FIG. 7. Total energy of ²⁴Mg (left) and ²⁸Si (right), calculated in the framework of HF + BCS with the forces SkP^{δ}, SkM*, and SVbas as a function of the deformation parameter β .

TABLE VI. Experimental and calculated deformation parameters β in ²⁴Mg and ²⁸Si.

	Expt. ^a	SV-bas	SkM*	\mathbf{SkP}^{δ}	SkT6	SV-mas10
²⁴ Mg ²⁸ Si			0.522 -0.291			$0.518 \\ -0.295$
^a From	Ref [36]					

isoscalar transition operators are $\hat{O}_{00} = \sum_{i}^{A} r_{i}^{2}$ and $\hat{O}_{20} = \sum_{i}^{A} r_{i}^{2} Y_{20}(\hat{r}_{i})$, respectively. For convenient comparison with the experimental data, the strength is smoothed by a Lorentz function $\xi_{\Delta}(E - E_{\nu}) = \Delta/(2\pi[(E - E_{\nu})^{2} - \Delta^{2}/4])$ with an averaging parameter $\Delta = 1$ MeV. The dimension of the strength functions is fm⁴/MeV. We also present 2qp strength functions calculated without the residual interaction. In this case, the number ν in Eq. (5) labels 2qp states.

B. Strength distribution and IS0/IS2 coupling

The QRPA ISO strength functions from Eq. (5) are compared with the experimental data in Figs. 8–11. The iThemba LABS experimental data are shown by black filled circles where each data point accumulates the ISO strength in an energy interval of 0.5 MeV.

In Fig. 8, the strength functions calculated with the force SV-bas are shown. In our set of Skyrme forces, SV-bas has a large (though rather typical for Skyrme forces) incompressibility $K_{\infty} = 234$ MeV. The calculated QRPA IS0 strengths (blue solid line) in ²⁴Mg and ²⁸Si are separated into a narrow structure at ≈ 15 MeV in ²⁴Mg and ≈ 19 MeV in ²⁸Si and a broad structure at higher energy, which is a typical picture

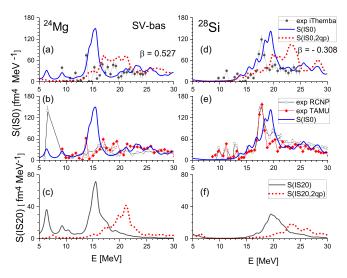


FIG. 8. Upper panels: QRPA (solid blue line) and 2qp (shortdash red line) IS0 strength functions calculated with the force SVbas $(K_{\infty} = 234 \text{ MeV})$ in ²⁴Mg (left) and ²⁸Si (right) are compared with the present iThemba LABS experimental data shown as black filled circles. Middle panels: comparison with RCNP [14,16] (black empty squares) and TAMU [18,19] (red filled diamonds) experimental data. Bottom panels: QRPA (solid black line) and 2qp (short-dashed red line) IS20 strengths.

for deformed nuclei [9]. It is instructive to compare QRPA IS0 strength with the 2qp IS0 one (red dashed line) obtained without the residual interaction. We see a strong collective effect: the residual interaction essentially shifts down the IS0 strength and creates collective peaks at ≈ 15 MeV in ²⁴Mg and ≈ 19 MeV in ²⁸Si.

The bottom panels of Fig. 8 show the LK = 20 branches of the ISGQR (see the Appendix for a discussion of the deformation splitting of the ISGQR in prolate and oblate nuclei) in the QRPA and 2qp cases. Here, we also observe a strong collective effect. Additionally, one observes that the main IS20 QRPA peak lies at the same energy as the narrow IS0 QRPA peak. Thus, both peaks are related, i.e., we observe a clear deformation-induced coupling of the IS0 and IS20 modes. The larger the nuclear deformation, the more strength is concentrated in the narrow IS0 peak [9]. Altogether, we see that the narrow IS0 peaks in ²⁴Mg and ²⁸Si have a dual (collective effect + deformation-induced IS0/IS20 coupling) origin.

Now the question arises which structures in the experimental data should be associated with the calculated ISO narrow peak. To inspect this point, let us consider in Fig. 8 the experimental data from iThemba LABS (upper panel), RCNP [14,16] and TAMU [18,19] (middle panel). In 28 Si, all the data give a distinctive strong narrow peak at 17-19 MeV. In ²⁴Mg, the issue is more complicated. Here, the iThemba LABS and TAMU data [18] give a distinctive narrow peak at 13-14 MeV but the RCNP data [14] do not. At the same time, a hump at 13–14 MeV also appears in the (⁶Li, ⁶Li') data [17]. Moreover, it is clearly observed at 13.8 MeV in the RCNP (α , α') data of Ref. [15], which have an energy resolution comparable to the present experiment. Finally, the narrow peak at 13–14 MeV was observed in early (α, α') [38] and (⁶Li, ⁶Li') [39] experiments. So, in ²⁴Mg, the peak at 13-14 MeV is a good candidate for the comparison with the narrow ISO peak in our QRPA calculations and the analysis of IS0/IS2 coupling.

Note that the proper choice of the experimental ISO peak for the comparison with theory is of crucial importance in the present study. As argued above, this peak arises due to ISO/IS2 coupling in deformed nuclei and its energy location can be used to find the optimal Skyrme force and thereby determine the relevant incompressibility K_{∞} and isoscalar effective mass m_0^*/m . In Ref. [14], the local maximum at 16–17 MeV was assumed to be the narrow peak, which resulted in the preference of the force SkM* ($K_{\infty} = 217$ MeV, $m_0^*/m = 0.79$). Below we will show that in ²⁴Mg, the experimental peak at 13–14 MeV in the iThemba LABS and TAMU data is a better suited candidate. This will change the predicted values of K_{∞} and m_0^*/m .

To justify the latter, let us compare the calculated and experimental IS0 strengths in Fig. 8. It is seen that the calculated narrow IS0 peaks lie higher by 1.5-2 MeV than the experimental peaks at 13.8 MeV (²⁴Mg) and about 18 MeV (²⁸Si). If we additionally upshift the calculated IS0 strength in both nuclei, we could match the RCNP structure at 16–17 MeV in ²⁴Mg but, simultaneously, we would essentially worsen the description of the strength in ²⁸Si. Instead, if we assume the 13.8 MeV peak as the relevant narrow structure in ²⁴Mg,

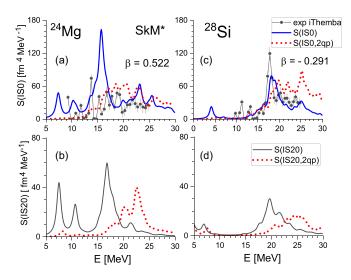


FIG. 9. Same as upper and lower panels of Fig. 8 but for the force SkM* with $K_{\infty} = 217$ MeV.

then it remains possible to describe ISO results in ²⁴Mg and ²⁸Si consistently by an equal downshift of the QRPA narrow peaks.

Figure 8 shows that SV-bas generally well describes the main ISGMR located at 16–25 MeV in ²⁴Mg and 20–25 MeV in ²⁸Si. The only noticeable discrepancy is that the theory underestimates the ISO strength at 16–20 MeV in ²⁴Mg where the experiment shows a sequence of prominent peaks. This discrepancy can be caused by the disregard in our calculations of the coupling with complex configurations (CCC). The CCC could redistribute the strength, reducing the height of the narrow ISO peak and shifting a part of its strength to the range 16–20 MeV.

Previous work [3,9] has shown that the lower K_{∞} is, the smaller the ISGMR centroid energy. As seen in Fig. 8, SV-bas gives too large energies for the ISO narrow peaks. So it is worth trying a force with a smaller K_{∞} . In Fig. 9, results for the SkM* force with $K_{\infty} = 217$ MeV are shown. As in Fig. 8, we see a strong collective effect. The energy of the narrow ISO peak again coincides with the energy of the quadrupole LK = 20 branch, which means that this peak is produced by the ISO/IS2 coupling. However, using SkM* does not improve the results: energies of the ISO narrow peaks are still noticeably overestimated. The downshift of the ISO strength due to the smaller $K_{\infty} = 217$ MeV is countered by an upshift due to the smaller value of the effective mass $m_0^*/m = 0.79$ (see the discussion of the impact of the effective mass in the Appendix).

Next, in Fig. 10, we present results for the force SkP^{δ} with a low incompressibility $K_{\infty} = 202$ MeV and high effective mass $m_0^*/m = 1$. Now both factors downshift the IS0 strength in comparison to that of the SV-bas case. An improved agreement with the experimental IS0 strength distributions is observed: in both ²⁴Mg and ²⁸Si, the calculated IS0 narrow peaks match the experimental structures. Note that SkP^{δ} also provides the best description of the ISGMR in deformed Mo isotopes [10]. In Fig. 10, the narrow peaks are again produced

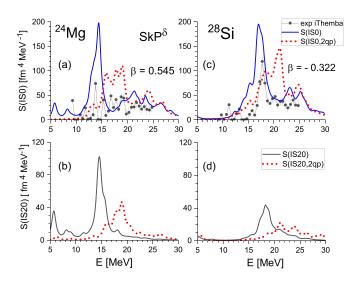


FIG. 10. Same as in Fig. 9 but for the force SkP^{δ} with $K_{\infty} = 202 \text{ MeV}$.

by IS0/IS2 coupling and the strong collective effect takes place.

In Figs. 8–10, the quadrupole IS20 peak in ²⁴Mg is stronger and has a lower energy than in ²⁸Si. As a result, the narrow IS0 peak in ²⁴Mg is more separated from the main ISGMR than in ²⁸Si. This is explained by the larger absolute quadrupole deformation of ²⁴Mg in comparison to ²⁸Si, which results in a stronger downshift of the IS20 peak in ²⁴Mg. Besides, Figs. 8–10 show that IS0/IS2 coupling is relevant also for lower-energy states.

Due to the ISO/IS2 coupling, the energy of the narrow peak is sensitive to the isoscalar effective mass m_0^*/m which affects the ISGQR energy [10,49]. In general, the larger m_0^*/m is, the smaller the ISGQR centroid energy [49]. Then one can try to downshift the energy of the narrow ISO peak solely using a large of m_0^*/m . In Fig. 11, the forces SkT6 ($K_{\infty} = 236$ MeV) and SV-mas10 ($K_{\infty} = 234$ MeV), both with $m_0^*/m = 1$, are applied for the calculation of the ISO strength. It is found that these two forces still overestimate the energies of the narrow ISO peaks like in the cases of SV-bas and SkM*. Thus, the optimum Skyrme force also needs a lower value of K_{∞} .

Note that, for the QRPA calculation of the ISGMR strength in 24 Mg with a Gogny force [17], a reasonable agreement with the data was achieved only after applying an *ad hoc* 2 MeV upshift of the ISO strength distribution. In our analysis, we do not use any *ad hoc* shifts.

C. IS2 strength

Since the energy of the narrow monopole peak is determined by the K = 0 branch of ISGQR, it is important to consider ISGMR and ISGQR simultaneously. In particular, we should see how well we can describe ISGQR and its deformation splitting in ²⁴Mg and ²⁸Si. This is studied in the Appendix for various values of the deformation parameter β . For ²⁴Mg, only prolate deformations are inspected. For ²⁸Si, where oblate and prolate deformations for excited states were predicted [22,54–56], both types of deformations

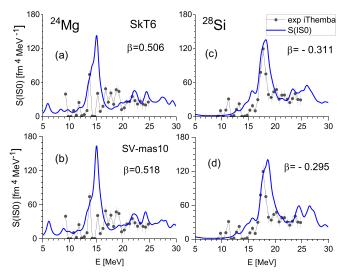


FIG. 11. QRPA ISO strength functions (solid lines) calculated with the Skyrme forces SkT6 (upper panels) and SV-mas10 (bottom panels).

are considered. In the Appendix, the calculations for ²⁴Mg display the expected behavior for prolate deformation splitting where the energies of the branches LK = 20, 21, 22 follow the relation $E_{20} < E_{21} < E_{22}$. The branch LK = 20 is strong. It is downshifted by 4–5 MeV when increasing the deformation from $\beta = 0.1$ to 0.522. The same splitting features are seen in prolate ²⁸Si. However, in oblate ²⁸Si, the deformation splitting of ISGQR looks surprising. The branch LK = 20 is not upshifted, as would be expected, but downshifted (by 3–4 MeV from $\beta = -0.1$ to -0.412) and becomes rather weak.

For prolate ²⁴Mg, these features are illustrated in Fig. 12, where SkM*, SV-bas, and SkP^{δ} QRPA results are compared to the ISGQR data obtained in reactions (⁶Li, ⁶Li') [17] and (α , α') [14,15,18]. In Refs. [18], the ISGQR data were given as fractions a_2 of the sum rule EWSR(IS2). We converted these data into IS2 strength using the relation

$$S(\text{IS2}) = \frac{\text{EWSR(IS2)}}{E_{\text{x}}} a_2(E_{\text{x}})$$
(6)

similar to expression (4) for the ISO strength. Because of the different approach for the conversion from cross sections to strength (cf. Sec. III), the results of Ref. [15] are multiplied by a factor of 2.5 to roughly match the other data sets. Figure 12 demonstrates that all four experimental datasets [14,15,17,18] give similar distributions, especially RCNP [14] and TAMU [18] data. Left panels of Fig. 12 show that the humps of the experimental (⁶Li, ⁶Li') strength at 14, 19, and 22 MeV are produced by the LK = 20, 21, and 22 branches, respectively. The forces SV-bas and SkP^{δ} give the most reasonable description. We also see a clear dependence of the calculated ISGQR strength on the isoscalar effective mass m_0^*/m running from 0.79 (SkM*) to 0.9 (SV-bas) and 1.0 (SkP^{δ}). The larger m_0^*/m is, the stronger the deformation (and deformation splitting) is and the lower the ISGQR energy.

In the right panels of Fig. 12, the calculated summed IS2 strength is compared to the (α, α') experimental data

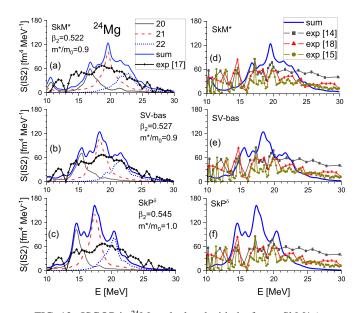


FIG. 12. ISGQR in ²⁴Mg calculated with the forces SkM* (upper panels), SV-bas (middle panels), and SkP^{δ} (bottom panels). For each force, the equilibrium deformation β and isoscalar effective masses m_0^*/m are shown. Left panels show strength functions *S*(IS2) for the branches *LK* = 20 (thin solid black line), 21 (dashed red line), and 22 (dotted blue line) as well as for the summed 20 + 21 + 22 strength (bold blue line). The calculated *S*(IS2) is compared to the experimental (⁶Li, ⁶Li') data from Zamora *et al.* [17] (black rhombuses). Right panels compare the summed *S*(IS2) to the experimental (α , α') data from Gupta *et al.* [14], Youngblood *et al.* [18] and Kawabata [15].

[14,15,18]. The best description of the data is obtained with the force SkP^{δ}, even the fine structure of the ISGQR from measurements [14,18] is reproduced. It is seen that the experimental bumps at 14–15, 17–18, and 19–20 MeV are produced just by *LK* = 20, 21, and 22 branches of the ISGQR.

What is important for this study, both the IS0 (present iThemba + Refs. [15,18]) and IS2 [14,15,17,18] experimental data for ²⁴Mg show prominent peaks at 14–15 MeV. These peaks are well reproduced by our QRPA calculations with SkP^{δ}. The coexistence of the IS0 and IS2 peaks at 14–15 MeV, obtained both in theory and experiment, justifies the strong IS0/IS2 coupling in ²⁴Mg.

In Fig. 13, a similar analysis of the ISGQR strength is done for the oblate ²⁸Si. Left panels of the figure show that the oblate deformation results in unexpected sequence of the IS-GQR branches, $E_{22} < E_{20} < E_{21}$, discussed in the Appendix. In ²⁸Si, the main peak at 19 MeV is produced by the LK = 22branch while the branches LK = 20 and 21 contribute to the right shoulder of the peak. Similar to ²⁴Mg, we see a clear dependence of the peak energy on the effective mass m_0^*/m : the larger m_0^*/m is, the lower the peak energy. In ²⁸Si, the forces SkM* and SV-bas somewhat better describe the ISGQR than SkP^{δ}. The description is not satisfactory since the energy position of the main peak and strong structure at 12–14 MeV are not reproduced. Perhaps this is caused by a softness and corresponding shape coexistence in ²⁸Si (in contrast to the stiff prolate character of ²⁴Mg).

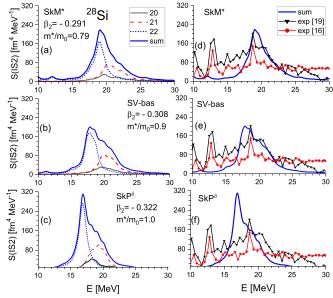


FIG. 13. The same as in Fig. 12 but for ²⁸Si. In the right panels, the calculated summed S(IS2) is compared to the experimental (α, α') data from Youngblood *et al.* [19] and Peach *et al.* [16].

Note that our QRPA IS2 strength functions are computed for band heads $I^{\pi}K = 0^{+}0$, $1^{+}1$, $2^{+}2$ while IS2 experimental data are obtained for the states $I^{\pi}K = 2^{+}0$, $2^{+}1$, $2^{+}2$, i.e., for rotational states in K = 0 and 1 cases. So, for a more accurate comparison with the experiment, the calculated IS2 (K = 0, 1) strengths should be shifted to a higher energy by the rotational correction. It can be estimated roughly from the energy of the $2^{+}0$ state in the ground-state rotational band, which is 1.368 MeV for 24 Mg and 1.779 MeV for 28 Si [36]. Though the corrections are significant, they do not affect the conclusions drawn from Figs. 12 and 13.

D. Integral IS0 strength

For further comparison, it is worth considering the summed IS0 strengths $\sum_{\nu} B_{\nu}(IS0)$ in the energy range 9–25 MeV covered by the present experiment. Here, $B_{\nu}(IS0) = |\langle \nu | \hat{O}_{2K} | 0 \rangle|^2$ is the reduced probability of the IS0 transition from the ground to ν th QRPA state. The results for the main three forces used in the present study (SkP^{δ}, SkM^{*}, and SV-bas) are shown in Table VII. One observes an anticorrelation between K_{∞} and

TABLE VII. Peak energy E_p of the narrow IS0 resonance and summed strength $\sum_{\nu} B_{\nu}$ (IS0) in the energy interval 9–25 MeV from various Skyrme forces compared with the present experimental data.

		24	Mg	28	³ Si
	K_{∞} (MeV)	E _p (MeV)	$\frac{\sum B(\text{IS0})}{(\text{fm}^4)}$	E _p (MeV)	$\frac{\sum B(\text{IS0})}{(\text{fm}^4)}$
Exp.		13.75(2)	728(41)	17.75(3)	895(40)
SkP ^δ	202	14.3	796	17.0	908
SkM*	217	15.6	706	18.0	780
SVbas	234	15.4	634	19.4	685

TABLE VIII. Energy-weighted strengths $\sum_{\nu} E_{\nu}B_{\nu}$ (IS0) for ²⁴Mg and ²⁸Si, summed in the energy intervals 0–16 MeV, 16–25 MeV, 25–70 MeV, and 0–70 MeV. The strengths are given in % of the monopole EWSR (see text for more detail). In the interval 0–16 MeV, the experimental strength is measured only in the interval 9–16 MeV.

	²⁴ Mg				28	Si		
	0–16	16–25	25–70	0–70	0–16	16–25	25–70	0–70
Expt.	15(1)	55(4)	0	70(4)	7(1)	68(4)	0	75(4)
\mathbf{SkP}^{δ}	38	35	25	98	11	63	25	97
SkM*	31	37	29	97	5	64	29	98
SV-bas	35	25	37	97	6	55	38	99

 $\sum_{\nu} B_{\nu}(IS0)$: the calculated $\sum_{\nu} B_{\nu}(IS0)$ steadily decreases from SkP^{δ} to SV-bas, i.e., with increasing K_{∞} from 202 to 234 MeV. As demonstrated for the strength distributions in Figs. 8–11, the best agreement with experiment is obtained using the force SkP^{δ}. Indeed, this force gives a reasonable reproduction of the summed strength in ²⁴Mg (together with the force SkM*) and the best reproduction of the summed strength in ²⁸Si. Besides, SkP^{δ} describes the peak energy E_p of the narrow ISO resonance in ²⁴Mg well and reasonably well in ²⁸Si. So, with this force, we simultaneously describe the energy of the narrow ISO peak and integral strength $\sum_{\nu} B_{\nu}(ISO)$. To the best of our knowledge, this is done for the first time.

It is of interest to inspect contributions of ISO strength from different energy intervals to the energy-weighted sum rule EWSR(IS0). They are summarized in Table VIII. It is seen that the experimental strengths exhaust about 70% of the EWSR in 24 Mg and 77% in 28 Si. The forces SkP^{δ}, SkM* and SV-bas give rather similar results. All calculations exhaust close to 100% EWSR in the model space 0-70 MeV. In ²⁴Mg, the calculations overestimate the experimental strength in the 0-16 MeV range (low-energy IS0 strength and the narrow peak) and underestimate it in the 16-25 MeV range (main part of ISGMR). It seems that the Skyrme forces redistribute too much of the strength from the main ISGMR to the narrow ISO peak. A large part of the strength, 25–37%, is located above 25 MeV, i.e., beyond the scope of the iThemba LABS experiment. For ²⁸Si, the agreement of the experiment and theory is much better. The calculations reproduce the weaker strength in the 0-25 MeV interval and the major strength in the 16-25 MeV interval rather well.

E. Features of individual IS0 transitions

The energy range 9–25 MeV embraces many QRPA 0⁺ states. For example, in the SkP^{δ} case, we have 631 states in ²⁴Mg and 537 states in ²⁸Si. It is instructive to consider in more detail QRPA states which form the narrow IS0 peak and strong IS0 excitations at a lower energy. In Table IX, the characteristics of IS0 states in the energy range 5–18 MeV with a strength *B*(IS0) > 40 fm⁴ are presented. In ²⁴Mg, there are two transitions with a large monopole strength below 13 MeV. The three transitions at energies between 13.5 and 14.5 MeV form the strong narrow IS0 peak shown in Fig. 10. Following

TABLE IX. Characteristics of low-energy $K^{\pi} = 0^+$ states with excitation energy $E_{\nu} > 5$ MeV and reduced transition probability $B(IS0) > 40 \text{ fm}^4$ in ²⁴Mg and ²⁸Si, calculated within QRPA with the force SkP⁸. For each state, we show the excitation energy, the reduced transition probabilities B(IS0) and B(IS20), and the main 2qp components (contribution to the state norm in %, and structure in terms of Nilsson asymptotic quantum numbers).

Nucleus	E_{ν} (MeV)	<i>B</i> (IS0) (fm ⁴)	<i>B</i> (IS20) (fm ⁴)	Ma (%)	in 2qp components $[N, n_z, \Lambda]$
²⁴ Mg	5.8	51	56.5	49	<i>pp</i> [220 ↑ −211 ↓
c	12.9	75	11.7	96	$pp [211 \uparrow -431 \uparrow]$
	13.5	77	26.5	92	<i>pp</i> [220 ↑ -440 ↑
	14.3	166	77.1	40	<i>pp</i> [211 ↑ -411 ↑
	14.5	81	40.0	48	<i>pp</i> [211 ↑ -411 ↑
²⁸ Si	15.3	42	3.0	91	<i>pp</i> [200 ↑ -431 ↓
	16.6	59	7.1	90	<i>pp</i> [202 ↑ -413 ↓
	16.9	91	10.5	68	$nn [200 \uparrow -411 \downarrow]$
	17.3	47	4.3	39	<i>nn</i> [200 ↑ -420 ↑
	17.6	62	6.0	75	<i>pp</i> [202 ↓ -431 ↑
	17.8	82	18.8	22	<i>pp</i> [200 ↑ -600 ↑

our calculations, two of these states, at 14.3 and 14.5 MeV, are very collective (composed by many 2qp components with the largest 2qp contributions to the state norm being 40% and 48%, respectively), which justifies the collective effect mentioned above. Most probably, just these two states with close energies produce the 13.87 MeV state seen in the iThemba LABS data (see Table III). It is difficult to establish one-to-one correspondence between the measured transitions in Table III and calculated transitions in Table IX since our Skyrme QRPA calculations neglect some important effects (coupling with complex configurations and shape coexistence) and thus have not enough accuracy for the detailed comparison. For ²⁴Mg and ²⁸Si, there were attempts of one-to-one comparison of the experimentally observed discrete 0⁺ levels and results of Gogny QRPA [17] and antisymmetrized molecular dynamics (AMD) [57] calculations. However, these calculations achieve a reasonable agreement with the data only by using 2 MeV ad hoc upshift of the ISO strength.

All the states in Table IX have large values for $0_{gs}^+ \rightarrow 2^+0_\nu$ isoscalar quadrupole transitions with $\Delta K = 0$. This once more confirms the strong ISO/IS2 coupling in strongly prolate ²⁴Mg. The main IS0 resonance at 18–25 MeV (not considered in Table IX) is formed by a large number of mildly collective states with $B(IS0) < 1 \text{ fm}^4$. The results for strongly oblate ²⁸Si generally agree with those for ²⁴Mg. However, as seen from Table IX, in ²⁸Si our calculations do not predict monopole states with a high B(IS0) and B(IS2, K = 0) below 15 MeV.

F. Cluster features of the narrow IS0 peak in ²⁴Mg

The strong narrow peak at 13.87 MeV in ²⁴Mg lies close to the thresholds of the ¹²C + ¹²C and ¹⁶O + 2α cluster configurations and, following the coincidence experiment [15], should exhibit cluster features. To check this point, we calculated the density distribution of the QRPA 14.3 MeV 0⁺ state

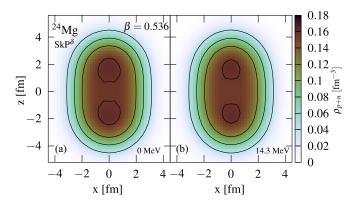


FIG. 14. Density profiles in the *x*-*z* plane, calculated with the force SkP^{δ} in ²⁴Mg for the ground state (left) and excited QRPA $K^{\pi} = 0^+$ state at 14.3 MeV (right).

(assuming its correspondence to the observed 13.87 MeV excitation) and compared it with the density of the ground state. In Fig. 14, both plots exhibit the densities in the x-z plane

where z is the symmetry axis. Another relevant collective 0^+ state at 14.5 MeV has a similar density contribution and so is not considered here.

We see that, in the ground state, prolate ²⁴Mg has two areas of the enhanced density in the pole regions. Although smaller, these areas of high density persist in the 14.3 MeV 0^+ state. These areas can be considered as precursors of clustering. They can be treated as predecessors of two ¹²C clusters or, to a lesser extent, of two α particles with ¹⁶O. Note that the plot for 14.3 MeV state agrees well with the ¹²C + ¹²C density plots in Ref. [57] where the cluster features of ²⁴Mg are analyzed within the AMD method. The similarity of the density plots for the ground and excited states can be explained by a dual nature (mean field + clustering) of the ground state of light nuclei; see the discussion and further citation in Ref. [57]. Altogether we see that the strong narrow ISO peak in ²⁴Mg might combine both mean-field (ISO/IS2 coupling) and cluster features. The nucleus ²⁸Si also demonstrates cluster properties (see, e.g., discussions in Refs. [21,54]). However, analysis of these properties is complicated by the shape coexistence in this nucleus and so is skipped here.

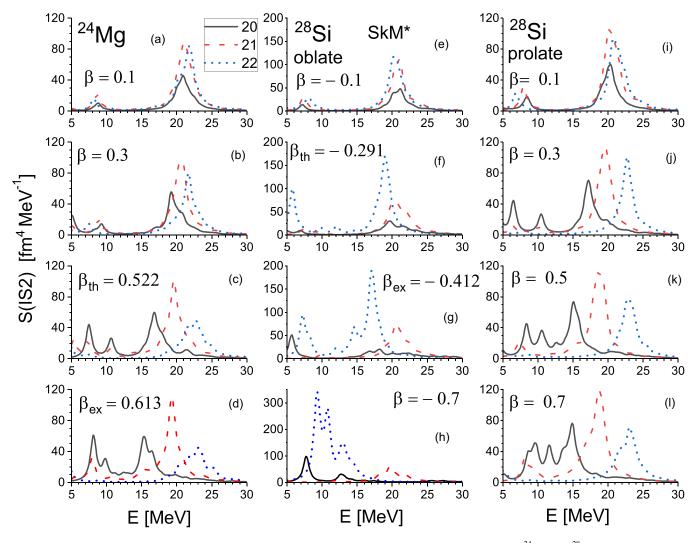


FIG. 15. ISGQR branches LK = 20 (solid black line), 21 (dashed red line), and 22 (dotted blue line) in ²⁴Mg and ²⁸Si, calculated with the force SkM* for different quadrupole deformations. For ²⁸Si, both oblate (middle panels) and prolate (right panels) deformations are considered.

V. CONCLUSIONS

The isoscalar monopole strength in the energy interval $9 \leq E_x \leq 25$ MeV in ²⁴Mg and ²⁸Si has been investigated using α -particle inelastic scattering with a 196 MeV beam at scattering angles $\theta_{\text{Lab}} = 0^\circ$ and 4° . The K600 magnetic spectrometer and the Separated Sector Cyclotron (SSC) at iThemba LABS were used.

The DoS technique was applied in order to extract the ISO strength distributions. The difference spectrum obtained for each nucleus was converted to the fraction of the ISO EWSR per bin and compared with MDA analysis results obtained by the TAMU [18,19] and RCNP [12,14–17] groups. Overall, the strength distributions obtained in this study show a reasonable agreement with results from both groups, with the exception of the RCNP data [12,14] for ISO structures in the energy range 13–18 MeV in ²⁴Mg. At the same time, we find good agreement for this energy range with another (α , α') experiment performed at RCNP [15].

The extracted ISO strength distributions were compared to Skyrme QRPA calculations. A representative set of Skyrme forces (SkM*, SV-bas, SkP^{δ}, SkT6, and SV-mas10) with different incompressibilities K_{∞} and isoscalar effective masses m_0^*/m was used. Similar to a recent study for Mo isotopes [10], the present iThemba LABS experimental data for the ISGMR in ²⁴Mg and ²⁸Si are best described by the force SkP^{δ} with a low incompressibility $K_{\infty} = 202$ MeV. This force allows one to reproduce both (i) the energy of the narrow ISO peaks at 13.8 MeV in ²⁴Mg and 18 MeV in ²⁸Si and (ii) the integral ISO strengths $\sum B(ISO)$.

In the theoretical analysis, the main focus was on the narrow IS0 peak constituting a convenient and sensitive test case to determine K_{∞} and m_0^*/m in deformed nuclei. The comparison of IS0 and IS2 strength distributions justifies that the narrow IS0 peak appears due to the deformation-induced coupling between the ISGMR and the K = 0 branch of the IS-GQR. The important role of a large collective downshift of the IS0 strength (due to the isoscalar residual interaction) in the production of the narrow IS0 peak lies close to the ${}^{12}C + {}^{12}C$ threshold and, in accordance with our analysis and observations [15], can demonstrate cluster features.

In connection with the strong IS0/IS2 coupling effect, we also performed a detailed analysis of the ISGQR in ²⁴Mg and ²⁸Si. An unusual deformation splitting in oblate ²⁸Si was found. The calculations show that the quadrupole K = 0 branch is strongly downshifted in prolate ²⁴Mg and slightly downshifted in oblate ²⁸Si. This explains why the narrow IS0 peak is well separated from the main ISGMR in ²⁴Mg but superimposed on the broad structure of the main ISGMR in ²⁸Si. The fine structure of the ISGQR in ²⁴Mg was described and explained using the force SkP⁸.

Our analysis demonstrates that the ISGMR in deformed nuclei must be described simultaneously with the ISGQR. Only then the description may be considered as relevant and consistent. In our calculations, the Skyrme force SkP^{δ} gives the best results for the simultaneous reproduction of the ISGMR and ISGQR. However, even this force has some shortcomings in the description of the ISGQR in ²⁸Si. There is

a need for new Skyrme parametrizations with $K_{\infty} < 217$ MeV and $m_0^{*}/m \approx 0.9$.

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APPENDIX: ISGQR IN ²⁴Mg and ²⁸Si

As shown in the main text, the K = 0 branch of the ISGQR determines the energy of the narrow IS0 peak. The energy of this branch is governed by the deformation splitting of the ISGQR into K = 0, 1, and 2 components. So, it is worth investigating the deformation splitting of the ISGQR in ²⁴Mg and ²⁸Si in more detail. Figure 15 shows the strength distributions of ISGQR branches for different axial quadrupole deformations. The calculations are performed with the force SkM* but other forces lead to the same qualitative conclusions.

In the plots (a)–(d), the ISGQR strength distribution in prolate ²⁴Mg is shown at $\beta = 0.1$, 0.3, 0.522 (calculated equilibrium value) and 0.613 (experimental value). We see a splitting picture typical for prolate nuclei: the K = 0 branch is downshifted from 21 MeV ($\beta = 0.1$) to 16 MeV ($\beta = 0.613$), the K = 1 branch shows less downshift from 21 to 19 MeV, and the K = 2 branch is upshifted from 22 to 23 MeV. The energy of the K = 0 branch is most sensitive to the deformation. However, even at $\beta = 0.613$, this branch does not approach the excitation energy 13–14 MeV where experiment gives the narrow monopole peak. So, to describe the experimental data we need a force like SkP^{δ} with a smaller incompressibility.

In the plots (e)–(h), dependence of the ISGQR in ²⁸Si on *oblate* quadrupole deformation is shown. The deformations are $\beta = -0.1$, -0.291 (calculated equilibrium value), -0.412 (experimental value), and -0.7. The deformation splitting demonstrates some surprising features. The K = 0 branch is not upshifted, as would be expected from analytic estimates [9], but downshifted from 21 to ≈ 13 MeV. The K = 0 strength is redistributed to a lower-energy region and significantly suppressed (see also Table X). The K = 1

TABLE X. EWSR exhaustion for the LK = 20, 21, and 22 branches of the ISGQR, calculated with the force SkM*.

	24 Mg, $\beta = 0$).522	28 Si, $\beta = -0.291$		
	(fm ⁴ MeV)	(%)	(fm ⁴ MeV)	(%)	
20	4468	26.5	3650	17.5	
21	7396	42.0	6998	33.7	
22	5534	31.5	10144	48.8	

branch keeps almost the same energy. The K = 2 branch is strongly downshifted from 21 to 8–15 MeV and its strength is significantly enhanced (see also Table X). At calculated $\beta_{\text{th}} = -0.291$ and experimental $\beta_{\text{ex}} = -0.412$ deformations, the K = 0 peak approaches the energy 17–18 MeV of the narrow IS0 peak. So the force SkM* can in principle describe the narrow IS0 peak in ²⁸Si once we take a proper value of the oblate quadrupole deformation.

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In the plots (i)–(1), the dependence of ISGQR in ²⁸Si on a *prolate* quadrupole deformation is inspected. This case is interesting since some studies predict superdeformed rotational bands based on prolate 0⁺ bandheads [22,54–56] at 9–13 MeV in ²⁸Si. As seen from Fig. 7, the total energy of ²⁸Si indeed has a local prolate minimum. The plots (i)–(1) show that the deformation splitting in prolate ²⁸Si would be rather similar to that in prolate ²⁴Mg. The K = 0 branch in prolate ²⁸Si covers the energy interval 8–16 MeV. So, our calculations do not contradict the suggestion [22,54–56] that ²⁸Si can possess highly prolate states at 9–13 MeV. Moreover, at $\beta = 0.3$, K = 0 strength covers the energy interval 16–19 MeV where the experimental narrow ISO peak is located.

In summary, from the present analysis we may conclude that the description of the experimental narrow IS0 peaks (i) needs a Skyrme force like SkP^{δ} ($K_{\infty} = 202$ MeV) with a low nuclear incompressibility in the case of ²⁴Mg and (ii) can be achieved with both SkP^{δ} and SkM* ($K_{\infty} = 217$ MeV) forces in the case of ²⁸Si once we allow the coexistence for oblate and prolate quadrupole deformations.

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