Effective shell-model interaction for nuclei "southeast" of ¹⁰⁰Sn

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We construct an effective shell-model interaction for the valence space spanned by single-particle neutron and single-hole proton states in ¹⁰⁰Sn. Starting from chiral nucleon-nucleon and three-nucleon forces and single-reference coupled-cluster theory for ¹⁰⁰Sn we apply a second similarity transformation that decouples the valence space. The particle-particle components of the resulting effective interaction can be used in shell model calculations for neutron deficient tin isotopes. The hole-hole interaction can be used to calculate the N = 50 isotones south of ¹⁰⁰Sn, and the full particle-hole interaction describes nuclei in the region southeast of ¹⁰⁰Sn. We compute low-lying excited states in selected nuclei southeast of ¹⁰⁰Sn, and find reasonable agreement with data. The presented techniques can also be applied to construct effective shell-model interactions for other regions of the nuclear chart.

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I. INTRODUCTION

The shell model is the paradigm to understand the structure of atomic nuclei and to compute their properties [1-4]. In this model, valence nucleons move within the mean field produced by the inert core and interact via a residual effective interaction. For a recent review of the long and winding road to effective shell-model interactions, we refer the reader to Ref. [5] and briefly summarize some developments in what follows. For a given model space, the matrix elements of the effective interaction are often determined phenomenologically by fit to data [6-8], or by tweaking the monopole terms [9] of a microscopically derived G matrix [10]. More microscopic approaches [11,12] are based on many-body perturbation theory which, however, poses its own challenges [13,14] and requires resummations [15,16]. Recent extensions of such approaches Coraggio et al. [17] sum up the irreducible two-body diagrams to infinite order and subtract the one-body components. The starting point typically are low-momentum nucleon-nucleon potentials [18] with a suitably chosen cutoff [19,20].

Effective interactions can also be derived without tuning any adjustable parameters when starting from nucleonnucleon and three-nucleon potentials [21–27]. Several of these approaches require one to decouple the two-body interaction in a small shell-model space from the (infinite) Hilbert space. Here, similarity transformations play a key role. These can be based on the Lee-Suzuki approaches [28,29], coupled-cluster theory [22,30,31], or the in-medium similarity renormalization group (IMSRG) [32–34].

In this paper, we develop an effective shell-model interaction for nuclei southeast of ¹⁰⁰Sn based on similarity transformations. This region of the nuclear chart is interesting because its cornerstone—the doubly magic N = Z = 50 nucleus ¹⁰⁰Sn—is close to the proton dripline, the endpoint of α decays [35,36], and exhibits one of the strongest Gamow-Teller matrix elements [37–39]. Several recent experiments studied the structure of nuclei in this part of the nuclear chart [40–44]. Traditionally, shell-model computations of nuclei in this region start from a ⁸⁸Sr or ⁹⁰Zr core [45,46]. In such an approach, the proton shell is almost full once elements close to tin have been reached, in addition, the computational cost quickly becomes a bottleneck (see, e.g., [46]), which makes it attractive to compute a particle-hole shell-model interaction starting from ¹⁰⁰Sn. We note that ¹⁰⁰Sn can be computed from scratch using coupled-cluster theory [46], and this nucleus is predicted to be a "better," i.e., more strongly bound, core than ⁸⁸Sr [45].

Recent nonperturbative approaches to effective interactions include methods that are based on many-body wave functions calculated with ab-initio methods such as no-core shell model (NCSM) [25,47] and coupled-cluster theory [22,23]. Once the wave functions of $A_{core} + 1$ and $A_{core} + 2$ are obtained, the effective interaction can be extracted through a Lee-Suzuki transformation [28,29,48,49]. Effective interactions derived through these approaches decouple the model space from the excluded space on the two- and three-body levels without introducing additional parameters. In principle, these nonperturbative methods also work for multishell spaces. However, it is challenging to obtain converged full-space wave functions for all allowed quantum numbers defined by the valence space, and for heavy nuclei. In the coupled-cluster implementation, for instance, some observables are sensitive to small contributions from high-lying excited states.

This computational problem is, to some extent, overcome by the valence-space IMSRG (VS-IMSRG) [34,50,51], and

the shell-model coupled-cluster (SMCC) [26] approaches. Instead of employing the *A*-body wave functions in Hilbert space, VS-IMSRG and SMCC are straightforward extensions of IMSRG and coupled-cluster theory, respectively, to openshell nuclei. Similarity transformations are applied to the normal-ordered Hamiltonian to decouple the model space and these approaches are suitable to calculate medium-heavy nuclei. Recently the VS-IMSRG was successfully applied to compute nuclei in the island of inversion region by decoupling a multishell valence space [52].

In this paper we select the core nearest to the target nuclei and perform a shell model calculation with both particles and holes. The effective interaction is derived through the particle-hole shell-model coupled-cluster approach (ph-SMCC). In ph-SMCC we treat the three-body correlations in a single-reference coupled-cluster approach instead of in the shell-model effective interaction. This is an advantage because the coupled-cluster method is more accurate than SMCC in dealing with the induced three-body correlations. The ph-SMCC is an extension of SMCC, in which a secondary similarity transformation is applied to the coupled-cluster effective Hamiltonian. The resulting Hamiltonian decouples the excluded space and the valence space, and the latter is spanned by particle and hole states. Once the Hamiltonian is decoupled, the valence effective interaction consists of particle-particle (pp), hole-hole (hh), and particle-hole (ph) channels. The pp channel is a conventional shell model effective interaction, and particle-removed nuclei can be calculated with the *hh* interaction. The *ph* channel can be used to calculate nuclei located in the southeast and northwest of a double magic nucleus, i.e., as protons removed, and neutrons attached, or vice versa. The particle-hole decoupling is equivalent to a conventional decoupling with the core in the valence space, and this is similar to the ensemble normal-ordering used in the VS-IMSRG [51]. The resulted effective particlehole interaction can be re-normal-ordered with respect to a smaller core to obtain a more conventional pp interaction.

The paper is organized as follows. In Sec. II, we briefly introduce the single-reference coupled-cluster method and the resulting similarity-transformed Hamiltonian. This is followed by the derivation of the particle-hole decoupling. In Sec. III we apply the *ph*-SMCC method to the 100 Sn region using two different chiral nucleon-nucleon and three-nucleon potentials. Finally, we summarize our results.

II. SHELL-MODEL COUPLED-CLUSTER PARTICLE AND HOLE INTERACTION

A. Shell-model coupled cluster

The coupled-cluster method [30,31,53-56] is useful for *ab-initio* calculations of medium-mass nuclei [46,57]. Our coupled-cluster calculations start from the intrinsic Hamiltonian,

$$H = \left(1 - \frac{1}{A}\right) \sum_{i=1}^{A} \frac{p_i^2}{2m} + \sum_{i < j=1}^{A} \left(v_{ij} - \frac{\overrightarrow{p_i} \, \overrightarrow{p_j}}{mA}\right) + \sum_{i < j < k}^{A} v_{ijk}.$$
(1)

Here, *m* is the mass of the nucleon, p_i and p_j are the singleparticle momentum, *A* is the mass number, v_{ij} is the nucleonnucleon potential, and v_{ijk} is the three-body potential defined in the laboratory coordinates. The Hamiltonian is henceforth normal-ordered with respect to a reference state $|\Phi_0\rangle$, e.g., a Hartree-Fock (HF) state or a product state of natural orbitals. We denote the energy expectation value of the reference state as $E_0 = \langle \Phi_0 | H | \Phi_0 \rangle$. The normal-ordered Hamiltonian is

$$\mathcal{H} = E_0 + \sum_{pq} f_{pq} \{ p^{\dagger}q \} + \frac{1}{4} \sum_{pqrs} V_{pqrs} \{ p^{\dagger}q^{\dagger}sr \}$$
(2)

$$+\frac{1}{36}\sum_{pqrstu}V_{pqr,stu}\{p^{\dagger}q^{\dagger}r^{\dagger}stu\}.$$
(3)

Here, we use p^{\dagger} as the particle or hole creation operator on state $|\phi_p\rangle$, and p is the annihilation operator. The one-body Fock matrix has elements f_{pq} , while V_{pqrs} and $V_{pqr,stu}$ denote the two-body and three-body matrix elements, respectively. To avoid dealing with the three-body diagrams after normal-ordering, Eq. (2) is usually truncated at the two-body level and the residual three-body terms are discarded. This approximation is accurate in ⁴He [58], ¹⁶O [59], and nuclear matter [60] (albeit only for three-nucleon forces with nonlocal regulators).

Coupled cluster theory is based on the exponential ansatz

$$|\Phi\rangle = e^T |\Phi_0\rangle,\tag{4}$$

where T is the cluster operator

$$T = T_1 + T_2 + T_3 + \cdots,$$
 (5)

that introduces one-particle-one-hole (1p-1h), 2p-2h, 3p-3h,...,Ap-Ah excitations. The np-nh excitation operator is

$$T_n = \frac{1}{(n!)^2} \sum_{\substack{i_1, i_2, \dots, i_n \\ a_1, a_2, \dots, a_n}} t_{i_1, i_2, \dots, i_n}^{a_1, a_2, \dots, a_n} \{a_1^{\dagger} a_2^{\dagger} \cdots a_n^{\dagger} i_1 i_2 \cdots i_n\}.$$
 (6)

The Schrödinger equation is then written as

$$\mathcal{H}e^{T}|\Phi_{0}\rangle = Ee^{T}|\Phi_{0}\rangle. \tag{7}$$

Left multiplication with e^{-T} on both sides of Eq. (7) yields

$$e^{-T}\mathcal{H}e^{T}|\Phi_{0}\rangle = E|\Phi_{0}\rangle. \tag{8}$$

Equation (8) indicates that the reference state $|\Phi_0\rangle$ is an eigenstate of the similarity-transformed Hamiltonian

$$\overline{\mathcal{H}} \equiv e^{-T} \mathcal{H} e^T.$$
⁽⁹⁾

In other words, $\overline{\mathcal{H}}$ generates no *ph* excitations of the reference, and

$$\langle \Phi^{a_1 a_2 \cdots a_n}_{i_1 i_2 \cdots i_n} | \overline{\mathcal{H}} | \Phi_0 \rangle = 0.$$
 (10)

The T amplitudes fulfill Eq. (10) and thereby decouple the reference state from all excited states. The similarity transformed Hamiltonian can be expanded as

$$\overline{\mathcal{H}} = E_{CC} + \sum_{pq} \overline{\mathcal{H}}_{pq} \{ p^{\dagger}q \} + \sum_{pqrs} \overline{\mathcal{H}}_{pqrs} \{ p^{\dagger}q^{\dagger}sr \} + \sum_{pqrstu} \overline{\mathcal{H}}_{pqrstu} \{ p^{\dagger}q^{\dagger}r^{\dagger}uts \} + \cdots, \qquad (11)$$

and contains—when *T* is not truncated—*A*-body operators, even when the normal-ordered Hamiltonian \mathcal{H} has lower rank. Here, $E_{CC} = \langle \Phi_0 | \overline{\mathcal{H}} | \Phi_0 \rangle$ is the correlation energy. In practice, the cluster operator in Eq. (5) is truncated to avoid the exponential computational expense of full configuration mixing. Truncation to $T = T_1 + T_2$ yields the coupled-cluster singles and doubles (CCSD) approximation.

Single-reference coupled-cluster theory is most efficient to calculate closed-shell nuclei. Open-shell systems can be computed by starting from a deformed, symmetry-breaking reference state [61], and the broken symmetry needs to be restored through angular momentum projection [62,63]. Another approach to open-shell nuclei employs equation-ofmotion methods (EOM-CC) [23,64–66]; these are restricted to neighbors of closed-shell nuclei

The SMCC method is an alternative approach to treat the open-shell problem with coupled-cluster methods by creating an effective interaction to be used in shell-model calculations. In such a calculation, the effective interaction is defined in a model space spanned by one or two major shells, whereas $\overline{\mathcal{H}}$ is defined in the full space defined by several shells in the harmonic-oscillator basis. To construct a shell-model effective interaction, the Hilbert space is split into the model space *P*, and its complement *Q*, where

$$P + Q = \mathbf{1}.\tag{12}$$

The task is then to decouple the P space from the Q space for the Hamiltonian. To accomplish this, a secondary similarity transformation is applied to the coupled-cluster Hamiltonian,

$$\overline{\overline{\mathcal{H}}} = e^{-S}\overline{\mathcal{H}}e^S,\tag{13}$$

and the decoupling between P and Q space requires

$$Q\overline{\mathcal{H}}P = 0. \tag{14}$$

Equation (14) is an A-body equation and S is also an A-body operator. In what follows, we use S_{pp} to represent the S operator in particle-particle decoupling, S_{hh} for that in the hole-hole decoupling, S_{hp} for particle and hole decoupling, and S with no subscript for any combination of them. On the two-body level, S_{pp} has the form

$$S_{pp} = \sum_{av_1} s_{v_1}^a \{a^{\dagger}v_1\} + \sum_{abv_1v_2} s_{v_1v_2}^{ab} \{a^{\dagger}b^{\dagger}v_1v_2\} + \sum_{abhv} s_{hv}^{ab} \{a^{\dagger}b^{\dagger}hv\} + \cdots$$
(15)

Here, we used v to represent valence state, while a, b, c, ...and i, j, k, ... denote particle and hole states, respectively. The first term in Eq. (15) represents a valence particle v_1 coupled to a state outside the model space a through a onebody operator [see diagram (a) in Fig. 1], the second term couples two particles inside the model space to a state which has at least one particle outside the model space [see diagram (b) in Fig. 1], and the third term couples the valence state to particle-hole configurations [see diagram (c) in Fig. 1]. One can write down three-body terms (and operators of higher rank) in a similar manner.

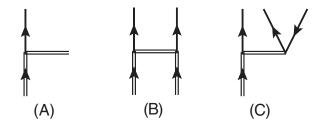


FIG. 1. Diagrammatic representation of the decoupling generator S_{pp} . The horizontal line is the S_{pp} operator, with the particles indicated by incoming and outgoing arrow lines. The model-space particles indicated by incoming double lines

In this paper, we truncate Eqs. (14) and (15) at the twobody level and refer to this as the SMCC(2) approximation. The two-body level decoupling condition (14) becomes

$$\langle a|\overline{\mathcal{H}}|v_1\rangle = 0, \tag{16}$$

$$\langle ab|\overline{\overline{\mathcal{H}}}|v_1v_2\rangle = 0,$$
 (17)

$$\langle ab|\overline{\mathcal{H}}|iv_1\rangle = 0. \tag{18}$$

Once decoupled, the effective Hamiltonian is

$$H_{\rm eff} = P \overline{\mathcal{H}} P, \tag{19}$$

which is defined only inside the model space. In principle, the Hamiltonian H_{eff} reproduces a subset of the eigenvalues of $\overline{\mathcal{H}}$. In practice, however, the truncation of many-body terms leads to a discrepancy of eigenvalues between H_{eff} and $\overline{\overline{\mathcal{H}}}$, because the truncation of many-body terms breaks the similarity of the transformation. The quality of the effective interaction thus depends on whether the neglected many-body terms are small compared to the retained two-body matrix elements [26].

The evaluation of $\overline{\mathcal{H}}$ via Eq. (13) is nontrivial due to the properties of *S*. In single-reference coupled-cluster theory, the operator *T* only consists of excitation operators, and we have

$$[T_m, T_n] = 0, (20)$$

i.e., different excitations commute. Thus, H can only contract with T from the right side, and Eq. (8) terminates exactly on the fourth nested commutator in CCSD for two-body Hamiltonians. In contrast, the operator S in SMCC consists of both excitation and de-excitation operators, and these do not commute with each other, i.e.,

$$[S_m, S_n] \neq 0. \tag{21}$$

Consequently, $\overline{\mathcal{H}}$ can contract with *S* from both the left and the right side and the evaluation of $\overline{\overline{\mathcal{H}}}$ does not terminate within the Baker-Campbell–Hausdorff (BCH) expansion,

$$\overline{\overline{\mathcal{H}}} = \overline{\mathcal{H}} + [\overline{\mathcal{H}}, S] + \frac{1}{2!} [[\overline{\mathcal{H}}, S], S] + \cdots .$$
(22)

Equation (22) is a set of nonlinear coupled equations, meanwhile, the nonterminating BCH expansion complicates the evaluation of $\overline{\overline{H}}$. We note that a truncation of Eq. (22) is not accurate. In this present work we adopt an approach that is similar to the Magnus method [67] from the IMSRG [68]. The idea is to make S dependent on a flow parameter t and then to replace Eq. (13) by the differential equation

$$\frac{\partial \overline{\mathcal{H}}(t)}{\partial t} = \left[\overline{\overline{\mathcal{H}}}(t), \frac{\partial S}{\partial t}\right].$$
(23)

Using the initial condition $\overline{\overline{\mathcal{H}}}(0) = \overline{\mathcal{H}}$ and choosing S(0) as the off-diagonal part of $\overline{\mathcal{H}}$ then requires us to numerically integrate Eq. (23). We obtain the operators S(t) and $\overline{\mathcal{H}}(t)$ as solutions, and the corresponding matrix elements become constant for sufficiently large values of the flow parameter t. When performing the integration, $\overline{\mathcal{H}}$ develops more than two-body terms (even when $\overline{\mathcal{H}}$ is of two-body rank), because the commutator of two two-body operators is in general a three-body operator. The induced three-body force can link the valence space and the excluded space via an off-diagonal three-body term S_{3b} . The diagonal three-body terms also contribute to the BCH expansion through $[H_{3h}, S_{2h}]_{2h}$. The explicit evaluation of the three-body terms is challenging and numerically expensive [69]. We therefore follow the nonperturbative IMSRG(2) approximation of Ref. [34] and truncate each commutator at the two-body level. This approximation is motivated by the success of the coupled-cluster and IMSRG computations when contributions from three-body operators are truncated at the normal-ordered two-body level [34,58,59]. In the BCH expansion nested commutators are computed to very high order until they become numerically very small; this makes the expansion nonperturbative. The rank of operators kept in the expansion, however, is truncated at two, i.e., only normal-ordered two-body operators are kept. The solution of Eq. (23) thus becomes $\overline{\overline{\mathcal{H}}} = \overline{\overline{\mathcal{H}}}_{1b} + \overline{\overline{\mathcal{H}}}_{2b}$.

When solving Eq. (23) we compute nested commutators recursively via

$$D^{(n)} = [D^{(n-1)}, S]_{2b}.$$
 (24)

Here, $D^{(0)} = \overline{\mathcal{H}}$. This allows us to rewrite Eq. (22) as

$$\overline{\overline{\mathcal{H}}} = \sum_{n=0}^{\infty} \frac{1}{n!} D^{(n)}.$$
(25)

Let us discuss the contribution of the neglected higher-body terms. To make estimates we turn to the BCH expansion (22) and assume that any corrections of *S* are small. (We note that *S* is dimensionless.) The leading contribution from three-body terms to two-body terms are $[[D, S_{2b}]_{3b}, S_{2b}]_{2b}$. Sun *et al.* [26] found that this commutator yielded a much smaller contribution to $\overline{\overline{H}}$ than the leading term $[D, S_{3b}]_{2b}$ from the three-body operator S_{3b} . Thus, we approximate the leading induced three-body force as

$$\overline{\overline{\mathcal{H}}}_{3b} = [\overline{\mathcal{H}}_{2b}, S_{2b}]_{3b}.$$
(26)

We note that this approximation, while well motivated, is in the end a pragmatic choice. Below we gauge its accuracy by comparison to more precise computations. The main task is then to calculate the contraction $[D^{(n)}, S]$. We find

$$D^{(n+1)ia}_{bc} = (1 - P_{bc}) \left(\sum_{e} D^{(n)ia}_{be} S^{e}_{c} + \sum_{je} D^{(n)ij}_{be} S^{ea}_{jc} \right)$$
$$- \sum_{e} \left(D^{(n)ie}_{bc} S^{a}_{e} - D^{(n)i}_{e} S^{ea}_{bc} \right)$$
$$+ \frac{1}{2} \sum_{ef} D^{(n)ia}_{ef} S^{ef}_{bc}, \qquad (27)$$

$$D^{(n+1)ij}_{\ \ ka} = \frac{1}{2} \sum_{ef} D^{(n)ij}_{\ \ ef} S^{ef}_{ka} + \sum_{e} D^{(n)ij}_{\ \ ke} S^{e}_{a},$$
(28)

$$D^{(n+1)ia}_{jb} = \sum_{e} \left(D^{(n)ia}_{je} S^{e}_{b} - D^{(n)ie}_{jb} S^{a}_{e} + D^{(n)i}_{e} S^{ea}_{jb} \right) + \frac{1}{2} \sum_{ef} D^{(n)ia}_{ef} S^{ef}_{jb} + \sum_{ke} D^{(n)ik}_{je} S^{ea}_{kb}, \qquad (29)$$

$$D^{(n+1)_{cd}^{ab}} = \frac{1}{2} \sum_{ef} \left(D^{(n)_{ef}^{ab}} S^{ef}_{cd} - S^{ab}_{ef} D^{(n)_{cd}^{ef}} \right) + \sum_{e} (1 - P_{cd}) \left(D^{(n)_{cd}^{ab}} S^{e}_{d} - S^{ab}_{ce} D^{(n)_{ed}^{e}} \right) - \sum_{e} (1 - P_{ab}) \left(D^{(n)_{cd}^{eb}} S^{a}_{e} - S^{eb}_{cd} D^{(n)_{e}^{a}} \right) - \sum_{k} (1 - P_{cd}) D^{(n)_{cd}^{k}} S^{ab}_{kd} + \sum_{je} (1 - P_{ab}) (1 - P_{cd}) D^{(n)_{ec}^{ja}} S^{eb}_{jd}, \qquad (30)$$

$$D^{(n+1)l_j}_{kl} = 0, (31)$$

and

$$D^{(n+1)ab}_{ic} = \frac{1}{2} \sum_{ef} \left(D^{(n)ab}_{ef} S^{ef}_{ic} - S^{ab}_{ef} D^{(n)ef}_{ic} \right) + \sum_{e} \left(D^{(n)ab}_{ie} S^{e}_{c} - D^{(n)e}_{c} S^{ab}_{ie} \right) + \sum_{e} (1 - P_{ab}) \left(D^{(n)b}_{e} S^{ae}_{ic} - S^{b}_{e} D^{(n)ae}_{ic} \right) - \sum_{k} S^{ab}_{kc} D^{(n)k}_{i} - \sum_{je} (1 - P_{ab}) D^{(n)ja}_{ie} S^{eb}_{jc}, \quad (32)$$

$$D^{(n+1)ia}_{\ jk} = -\sum_{e} D^{(n)ie}_{\ jk} S^a_e,$$
(33)

$$D^{(n+1)ij}_{\ ab} = \frac{1}{2} \sum_{ef} D^{(n)ij}_{\ ef} S^{ef}_{ab} + \sum_{e} (1 - P_{ab}) D^{(n)ij}_{\ ae} S^{e}_{b}.$$
 (34)

We make an initial guess of the *S* operator, compute the operator $\overline{\overline{\mathcal{H}}}$ and update *S* accordingly. By repeating this procedure we integrate out the off-diagonal pieces of $\overline{\mathcal{H}}$ until the decoupling condition is fulfilled.

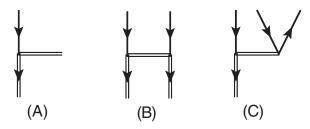


FIG. 2. Diagrammatic representation of S_{hh} . The horizontal line is the S_{hh} operator, with the particles and holes indicated by incoming and outgoing arrow lines. The model-space holes indicated by outgoing double lines

B. Particle-hole shell model coupled cluster

Before moving to the particle-hole decoupling, we derive the effective hole-hole interaction. A hole-hole decoupling yields an effective interaction for valence holes. On the twobody level, the operator S for hole-hole decoupling is

$$S_{hh} = \sum_{iv_1} s_i^{v_1} \{v_1^{\dagger}i\} + \sum_{v_1v_2ij} s_{ij}^{v_1v_2} \{v_1^{\dagger}v_2^{\dagger}ij\} + \sum_{ijv_1a} s_{ij}^{v_1a} \{v_1^{\dagger}a^{\dagger}ij\}.$$
(35)

Here, v_i denotes a valence-hole state. The first two terms link valence holes and holes outside the model space through a one- and two-body operator, and the last term is the single hole coupled to the particle-hole excitation configurations. Diagrams representing Eq. (35) are shown in Fig. 2.

Similar to the particle-particle decoupling, S_{hh} is determined through solving the decoupling equation (14). The two-body level decoupling equation for the hole-hole effective interaction is

$$\langle v_1 | \overline{\mathcal{H}} | i \rangle = 0, \tag{36}$$

$$\langle v_1 v_2 | \overline{\mathcal{H}} | ij \rangle = 0, \qquad (37)$$

$$\langle v_1 a | \overline{\mathcal{H}} | i j \rangle = 0. \tag{38}$$

Once decoupled, the hole-hole sector of $\overline{\mathcal{H}}$ should reproduce a subset of eigenvalues of particle removed systems.

Using the two-body level S for the hole-hole decoupling in Eq. (35), the single commutator yields

$$D^{(n+1)ia}_{bc} = \frac{1}{2} \sum_{kl} D^{(n)kl}_{bc} S^{ia}_{kl} - \sum_{k} D^{(n)ka}_{bc} S^{i}_{k}, \qquad (39)$$
$$D^{(n+1)ij}_{ka} = -\sum_{l} (1 - P_{ij}) D^{(n)lj}_{ka} S^{i}_{l} + \sum_{bl} (1 - P_{ij}) D^{(n)lj}_{ba} S^{ib}_{kl}$$

$$+\sum_{l} \left(D^{(n)ij}_{\ \ la} S^{l}_{k} - S^{ij}_{kl} D^{(n)l}_{\ \ a} \right) + \frac{1}{2} \sum_{lm} D^{(n)lm}_{\ \ ka} S^{ij}_{lm},$$
(40)

$$D^{(n+1)ia}_{jb} = \sum_{k} \left(-D^{(n)ka}_{jb} S^{i}_{k} + D^{(n)ia}_{kb} S^{k}_{j} - D^{(n)k}_{b} S^{ia}_{jk} \right) + \frac{1}{2} \sum_{kl} D^{(n)kl}_{jb} S^{ia}_{kl} + \sum_{ck} D^{(n)ka}_{cb} S^{ic}_{jk},$$
(41)

$$D^{(n+1)ab}_{\ \ cd} = 0, (42)$$

and

$$D^{(n+1)ij}_{kl} = \frac{1}{2} \sum_{mn} \left(D^{(n)mn}_{kl} S^{ij}_{mn} - S^{mn}_{kl} D^{(n)ij}_{mn} \right) + \sum_{m} (1 - P_{ij}) \left(-D^{(n)mn}_{kl} S^{j}_{m} + S^{im}_{kl} D^{(n)j}_{m} \right) + \sum_{m} (1 - P_{kl}) \left(D^{(n)ij}_{km} S^{m}_{l} - S^{ij}_{km} D^{m}_{l} \right) + \sum_{a} (1 - P_{ij}) D^{(n)j}_{a} S^{ia}_{kl} + \sum_{am} (1 - P_{ij}) (1 - P_{kl}) D^{(n)m}_{ka} S^{ja}_{lm}, \qquad (43)$$

$$D^{(n+1)ab}_{ic} = \sum_{k} D^{(n)ab}_{kc} S^{k}_{i}, \qquad (44)$$

$$D^{(n+1)ia}_{jk} = \frac{1}{2} \sum_{mn} \left(D^{(n)mn}_{jk} S^{ia}_{mn} - S^{mn}_{jk} D^{(n)ia}_{mn} \right) + \sum_{l} \left(-D^{(n)la}_{jk} S^{i}_{l} + S^{la}_{jk} D^{(n)i}_{l} \right) + \sum_{l} (1 - P_{jk}) \left(D^{(n)ia}_{lk} S^{l}_{j} - S^{ia}_{lk} D^{(n)l}_{j} \right) + \sum_{b} D^{(n)a}_{b} S^{ib}_{jk} - \sum_{cl} (1 - P_{jk}) D^{(n)la}_{kc} S^{ic}_{jl}, \qquad (45)$$

$$D^{(n+1)ij}_{\ ab} = \frac{1}{2} \sum_{kl} D^{(n)kl}_{\ ab} S^{ij}_{kl} - \sum_{k} (1 - P_{ij}) D^{(n)ik}_{\ ab} S^{j}_{k}.$$
 (46)

Equations (31) and (42) imply that the hole-hole interaction is not affected by a particle-particle decoupling, meanwhile the particle-particle interaction is not changed by a hole-hole decoupling. This is true only in a single commutator, and the interplay between the *hh* and *pp* decoupling occurs in the nested commutator. This suggests that the *pp* interaction is weakly affected by the *hh* decoupling, because of the diminishing importance of increasingly nested commutators. In practical calculations, the effective *pp* interaction from a particle-hole decoupling is not identical to that from a *pp* decoupling, but as expected they do reproduce the same eigenvalues.

In the particle-hole decoupling, valence particles and holes need to be treated on an equal footing. The possibility of particle-hole de-excitation in the valence space complicates the decoupling, as each sector of the Hamiltonian need to be decoupled, see Fig. 3. In most shell-model calculations, we do not usually need protons or neutrons to be particle and hole states simultaneously. In the current work, we restrict ourselves to a model space that contains only neutron particles and proton hole states (or verse vice). The off-diagonal Hamiltonian contains no particle-hole de-excitations due to isospin conservation, and only the last term in Fig. 3 contributes. The *S* operator for the particle-hole decoupling is

$$S = S_{hh} + S_{pp} + S_{hp}.$$
 (47)

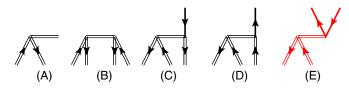


FIG. 3. Diagrammatic representation of S_{hp} . The horizontal line is the S_{hp} operator, with the particles indicated by incoming and outgoing arrow lines. The valence space particles and holes are indicated by incoming double lines. (A) represents a particle-hole de-excitation in the valence space, (B) a valence space 2p-2h de-excitation, (C) a valence-space 2h-1p coupled to one hole, (D) a valence-space 2p-1hcoupled to one particle, and (E) represent valence-space particle-hole coupled to a particle-hole in the excluded space.

Here, S_{pp} and S_{hh} have already been discussed above and we will focus on the additional terms. Diagram E in Fig. 3 is

$$S_{hp} = \sum_{iajv} s_{jv_2}^{v_1 a} \{ v_1^{\dagger} a^{\dagger} j v_2 \}.$$
(48)

Here, S_{hp} couples the particle-hole states between the model space and the excluded space.

Additional terms contributing to the single commutator $[D^{(n)}, S_{hp}]$ are

$$D^{(n+1)ia}_{bc} = \sum_{kd} (1 - P_{bc}) \left(D^{(n)ka}_{cd} S^{id}_{kb} - S^{ka}_{cd} D^{(n)id}_{kb} \right) - \sum_{i} (1 - P_{bc}) D^{(n)j}_{b} S^{ia}_{jc},$$
(49)

$$D^{(n+1)ij}_{ka} = \sum_{dl} (1 - P_{ij}) \left(-D^{(n)il}_{kd} S^{jd}_{la} + S^{il}_{kd} D^{(n)jd}_{la} \right) + \sum_{b} (1 - P_{ij}) D^{(n)j}_{b} S^{ib}_{ka},$$
(50)

$$D^{(n+1)ia}_{\ jb} = \sum_{kc} \left(-D^{(n)ka}_{\ jc} S^{ic}_{kb} + S^{ka}_{jc} D^{(n)ic}_{\ kb} \right) + \sum_{c} \left(D^{(n)a}_{\ c} S^{ic}_{jb} - D^{(n)c}_{\ b} S^{ia}_{jc} \right) + \sum_{k} \left(-D^{(n)k}_{\ j} S^{ia}_{kb} + D^{(n)i}_{\ k} S^{ka}_{jb} \right),$$
(51)

$$D^{(n+1)ij}_{\ ab} = -\sum_{kc} (1 - P_{ij})(1 - P_{ab}) D^{(n)ik}_{\ ac} S^{jc}_{kb}.$$
 (52)

Once decoupled, the effective interaction consists of three sectors. The first sector is the typical *pp* interaction, for which the low-lying excited states of particles attached to the core can be calculated. The second sector is the hole-hole interaction, which can be used in hole-hole shell model calculations for particle removed systems. Another sector is the particle-hole channel, which can be used to calculate nuclei with protons removed and neutrons attached depending on the model space selected.

For a hole-hole shell-model calculation, one can use standard shell-model codes. The expectation values of the one-body interaction are calculated via

$$\langle ij|f_{pq}\{p^{\dagger}q\}|kl\rangle = -f_{ki}\delta_{jl} - f_{lj}\delta_{ik} + f_{li}\delta_{jk} + f_{kj}\delta_{il}, \quad (53)$$

and two-body expectation values are

$$\langle ij|f_{pqrs}\{p^{\dagger}q^{\dagger}rs\}|kl\rangle = f_{klij}.$$
(54)

The particle-hole interaction is more complicated. The onebody operator is applied as

$$\langle ai^{-1}|f_{pq}\{p^{\dagger}q\}|bj^{-1}\rangle = f_{ab}\delta_{ij} - f_{ji}\delta_{ab}.$$
 (55)

The two-body operator needs an explicit Pandya transformation [70]

$$\langle ai^{-1}|V|bj^{-1}\rangle^J = -\sum_{J'} J'^2 \begin{cases} j_a & j_i & J\\ j_b & j_j & J' \end{cases} \langle aj|V|bi\rangle^{J'}.$$
 (56)

The interaction can also be re-normal-ordered with respect to a lighter core where all single-particle orbitals are particle states. The shell-model effective interaction from SMCC is non-Hermitian. To be used in standard shell model codes, the effective interaction can be made Hermitian [49] through

$$H_{\rm eff}^{\rm hm} = [\omega^{\dagger}\omega]^{1/2} H_{\rm eff}^{\rm nhm} [\omega^{\dagger}\omega]^{-1/2}.$$
 (57)

Here, ω is the matrix that diagonalizes the non-Hermitian Hamiltonian $H_{\text{eff}}^{\text{nhm}}$. The Hermitian $H_{\text{eff}}^{\text{hm}}$ contains one and two-body terms and can be used in conventional shell-model calculations.

The truncation at the two-body level is a reasonable first approximation in SMCC and VS-IMSRG. However, The off-diagonal three-body force S_{3b}^{od} induced by the similarity transformation cannot be neglected for an accurate description of nuclei. The full three-body decoupling is computation-ally very expensive and challenging to implement. However, the leading contribution from S_{3b}^{od} is accessible through a linearized approximation in the BCH expansion (22) and captures the important parts of three-body correlations [26]. We discuss this correction next.

The leading contributions from induced three-body terms comes from Eq. (26), and that approximation neglects induced three-body terms from any of the nested commutators. The operator $\overline{\overline{\mathcal{H}}}_{3b}$ can be split into diagonal $\overline{\overline{\mathcal{H}}}_{3b}^{da}$ and off-diagonal $\overline{\overline{\mathcal{H}}}_{3b}^{od}$ contributions. Suppose $\overline{\overline{\mathcal{H}}}_{3b}^{od}$ is driven to zero in a third similarity transformation

$$Qe^{-S_{3b}}\overline{\overline{\mathcal{H}}}e^{S_{3b}}P = 0,$$
(58)

where S_{3b} couples the valence space and excluded space at the three-body level. This operator would feed back to the two-body decoupling equation, breaking the two-body level decoupling condition $Q\overline{\overline{H}}P = 0$. As an approximation, we assume this feedback can be neglected, because in a proper decoupling, the induced $\overline{\overline{H}}_{3b}$ should be small compared to $\overline{\overline{H}}_{2b}$. The left-hand side of Eq. (58) can therefore be written as

$$Qe^{-S_{3b}}\overline{\overline{\mathcal{H}}}e^{S_{3b}}P \approx Q(1-S_{3b})\overline{\overline{\mathcal{H}}}(1+S_{3b})P$$
(59)

$$\approx Q\overline{\overline{\mathcal{H}}}P + Q[\overline{\overline{\mathcal{H}}}, S_{3b}]P.$$
(60)

Substitution of Eq. (60) into Eq. (58) yields the decoupling equation that determines S_{3b} ; these are

$$Q\overline{\overline{\mathcal{H}}}_{1b}P + Q[\overline{\overline{\mathcal{H}}}, S_{3b}]_{1b}P \approx 0, \tag{61}$$

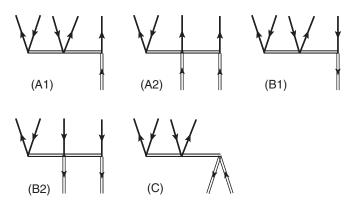


FIG. 4. Diagrammatic representation of S_{3b} . The horizontal line is the S_{3b} operator, with the particles indicated by incoming and outgoing arrow lines. The model-space particles and holes are indicated by incoming and outgoing double lines respectively. A1 and A2 are used in particle-particle decoupling, B1 and B2 for hole-hole decoupling, and all A, B, and C diagrams are used for particle-hole decoupling.

$$Q\overline{\overline{\mathcal{H}}}_{2b}P + Q[\overline{\overline{\mathcal{H}}}, S_{3b}]_{2b}P \approx 0, \tag{62}$$

$$Q\overline{\overline{\mathcal{H}}}_{3b}P + Q[\overline{\overline{\mathcal{H}}}, S_{3b}]_{3b}P \approx 0.$$
(63)

As we have assumed that $Q\overline{\mathcal{H}}P = 0$ is preserved on the twobody level, Eqs. (61) and (62) are trivial, whereas S_{3b} can be determined by Eq. (63). If we keep only the diagonal onebody part in $\overline{\mathcal{H}}$, the operator S_{3b} has the simple form

$$\langle pqu|S_{3b}|rst\rangle = \frac{\langle pqu|\overline{\mathcal{H}}_{3b}|rst\rangle}{\overline{\overline{\mathcal{H}}}_{pp} + \overline{\overline{\mathcal{H}}}_{qq} + \overline{\overline{\mathcal{H}}}_{uu} - \overline{\overline{\mathcal{H}}}_{rr} - \overline{\overline{\mathcal{H}}}_{ss} - \overline{\overline{\mathcal{H}}}_{tt}}.$$
(64)

Figure 4 shows the diagrams included in S_{3b} . The diagrams A1 and A2 couple the valence particles to the excluded space. Diagrams B1 and B2 couple the valence holes and the excluded space, and diagram C couples the particle-hole states to the excluded space in the isospin $T_z = \pm 1$ channels.

With S_{3b} , the effective shell model interaction becomes

$$H_{1b}^{\text{eff}} = P\overline{\overline{\mathcal{H}}}_{1b}P + P[\overline{\overline{\mathcal{H}}}, S_{3b}]_{1b}P,$$
(65)

$$H_{2b}^{\text{eff}} = P\overline{\overline{\mathcal{H}}}_{2b}P + P[\overline{\overline{\mathcal{H}}}, S_{3b}]_{2b}P.$$
(66)

The final effective interaction is thereby decoupled at the two-body level and three-body corrections are included in a linearized approximation.

III. RESULTS

We use two chiral interactions to compute the corresponding shell-model effective interactions. The first is the 1.8/2.0(EM) potential [71] which results from a similarityrenormalization-group (SRG) transformation [72] at cutoff $\lambda = 1.8 \text{ fm}^{-1}$ of the N³LO(EM) nucleon-nucleon potential from [73], and the three-body potential is given at N²LO in chiral EFT [74–76] with a cutoff $\Lambda_{NNN} = 2.0 \text{ fm}^{-1}$. The second potential is the Δ NNLO_{GO} with a cutoff $\Lambda = 394 \text{ MeV}$ TABLE I. Single-particle energies calculated with the 1.8/2.0(EM) and Δ NNLO_{GO} potentials. $p_{3/2}$, $p_{1/2}$, and $g_{9/2}$ are hole states with respect to ¹⁰⁰Sn, and the remaining states are particle states.

s.p. energy (MeV)		1.8/2.0(EM)	ΔNNLO _{GO}
	<i>p</i> _{3/2}	-6.502	-4.798
holes	$p_{1/2}$	-4.871	-3.159
	89/2	-3.106	-2.267
particles	87/2	-10.832	-10.072
	$d_{5/2}$	-10.548	-9.375
	$d_{3/2}$	-8.054	-6.811
	<i>s</i> _{1/2}	-7.789	-6.220
	$h_{11/2}$	-5.596	-5.012

[77]. This nucleon-nucleon and three-nucleon interaction includes Δ degrees of freedom. With both interactions we perform a Hartree-Fock (HF) calculation in the harmonic oscillator basis with a frequency of $\hbar \omega = 16$ MeV. The model space is spanned by 13 shells ($N_{\text{max}} = 12$), and the three-body matrix elements are truncated at $E_{3\text{max}} = 16\hbar\omega$. The two-body and three-body interaction is then normal ordered with respect to the ¹⁰⁰Sn core. The CC calculations are performed with the normal-ordered two-body interaction.

We use ¹⁰⁰Sn as the core, with the model space spanned by neutron-particle and proton-hole states. Using the effective shell-model interactions, we studied the nuclei southeast of ¹⁰⁰Sn by removing protons and attaching neutrons. To decouple the core, we employed the CCSDT-1 approximation [78] for both potentials. Using the highly optimized nuclear tensor contraction library (NTCL) [79] we are able to perform these calculations in the full space (without truncating the number of triples amplitudes) on SUMMIT, the supercomputer of the U.S. Department of Energy with a peak performance of 200 petaflops which is operated by the Oak Ridge Leadership Computing Facility at Oak Ridge National Laboratory.

The resulting binding energy of ¹⁰⁰Sn is -837 MeV and -816 MeV for 1.8/2.0(EM) and Δ NNLO_{GO}, respectively, compared to the datum of -825.3 MeV. The proton model space includes the $p_{3/2}$, $p_{1/2}$, and $g_{9/2}$ orbitals, and the neutron space consists of $g_{7/2}$, $d_{5/2}$, $d_{3/2}$, $s_{1/2}$, and $h_{11/2}$. We decouple the Hamiltonian at the two-body level, and include the off-diagonal three-body terms approximately according to Eqs. (65) and (66).

Table I shows the calculated single-particle and single-hole energies from the two chiral potentials. For ¹⁰¹Sn, the experimental splitting is a 171 keV between the $5/2^+$ and the $7/2^+$ states [35,36]. In our calculation the 1.8/2.0(EM) yields the $E(g_{7/2})$ 284 keV lower than $E(d_{5/2})$, which agrees with the PA-EOM-CC and VS-IM-SRG calculations [46]. Meanwhile, the splitting from Δ NNLO_{GO} is 700 keV, which is a bit larger than the datum. Both potentials predict a $9/2^+$ ground state of ⁹⁹In, which has a single-hole configuration. Matrix elements of these interactions are available in the Supplemental Material [80].

Figure 5 shows the calculated binding energies of twoparticle (102 Sn), two-hole (98 Cd), and one-particle–one-hole

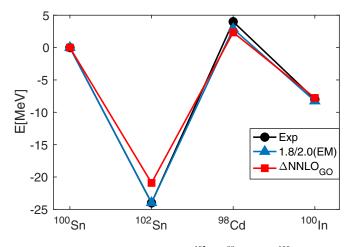


FIG. 5. Ground-state energies of 102 Sn, 98 Cd, and 100 In relative to 100 Sn, calculated with 1.8/2.0(EM) and Δ NNLO_{GO} potential and compared to experimental data. We used an arbitrary energy shift of 90 keV between the (6⁺) and (5⁺) levels.

(¹⁰⁰In) nuclei with respect to the ground state of ¹⁰⁰Sn. The 1.8/2.0(EM) potential is in agreement with the experimental data. The Δ NNLO_{GO} somewhat underestimates ¹⁰²Sn and overbinds the two-proton system ⁹⁸Cd. The two potentials coincide in ¹⁰⁰In. In a simple non-interacting shell-model picture, the ground state of ¹⁰⁰In depends on $E(\pi g_{7/2}) - E(\nu g_{9/2})$, and the coincidence in ¹⁰⁰In is indicated from our calculated single-particle energies.

The spectra of the two-hole state nucleus ⁹⁸Cd and the particle-hole nucleus ¹⁰⁰In are shown in Fig. 6. For cadmium, the 1.8/2.0(EM) interaction yields an accurate spectrum while the Δ NNLO_{GO} potential yields a spectrum that is too compressed. The dominant configuration for ⁹⁸Cd is two holes in $\pi g_{9/2}$. The Δ NNLO_{GO} generates a small matrix element $\langle \pi g_{9/2}\pi g_{9/2}|V|\pi g_{9/2}\pi g_{9/2}\rangle^J$, and this leads to the compressed spectrum. The odd-odd nucleus ¹⁰⁰In has a high level density close to the ground state with quasidegenerate $J^{\pi} = 2^+ \cdots 7^+$ states, stemming from the coupling of a $\pi g_{9/2}$ hole to the almost degenerate $\nu g_{7/2}$ and $\nu d_{5/2}$ particle states. While the experimental spectrum is unknown [37] regarding spin assignments and the energy shift with respect to the

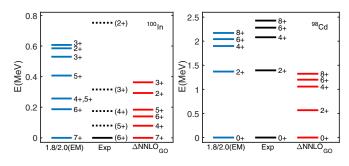


FIG. 6. Low-energy spectra of 98 Cd (right) and 100 In (left) calculated with 1.8/2.0(EM) and Δ NNLO_{GO} potentials, respectively, and compared to experimental data. For 100 In, the energies of the experimentally known excited states (shown as dotted lines) have an unknown shift with respect to the ground state [37].

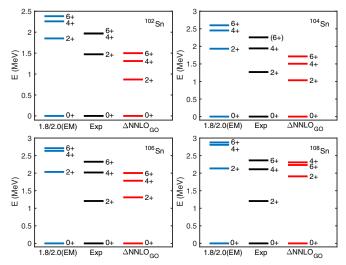


FIG. 7. Calculated low-lying excited states in Sn isotopes with the 1.8/2.0(EM) (blue) and $\Delta \text{NNLO}_{\text{GO}}$ (red) potentials, and compared to data (black).

ground state, the level density from the 1.8/2.0(EM) interaction is more accurate than that from the Δ NNLO_{GO} potential. We also note that the level density of the 1.8/2.0(EM) interaction agrees with the coupled-cluster results shown in Fig. 4 of the Supplementary Information of Ref. [38]. The ground-state spin of ¹⁰⁰In is not known experimentally [81], and both interactions yield a 7⁺ ground state. However, our estimated uncertainties of about 0.2 MeV are too large to make a prediction here.

The tin isotopes depend only on the neutron-neutron interaction, i.e., the particle-particle interaction constructed in this work. Figure 7 shows the low-lying excited states calculated for ^{102–104}Sn. The 1.8/2.0(EM) potential yields a near constant 2⁺ energy, in agreement with the VS-IMSRG calculation of Ref. [46]. The Δ NNLO_{GO} potential also gives a near constant $E(2^+)$ energy (with the exception of ¹⁰⁸Sn), but generally lower in excitation energy than obtained using the 1.8/2.0(EM) potential. Data fall in-between the two employed potentials for ^{102,104}Sn, for ¹⁰⁶Sn the results obtained with Δ NNLO_{GO} agrees well with data, while for ¹⁰⁸Sn both potentials overestimate the $E(2^+)$ energy. The difference in the $E(2^+)$ energies obtained from the two potentials stems mostly from the associated single-particle energies given in Table I.

The cadmium isotopes are obtained by removing two protons from ¹⁰⁰Sn core and adding neutrons. Our calculations, shown in Fig. 8, indicate that the removed neutrons are mainly from $g_{9/2}$ and the spectra of cadmium isotopes are mostly determined by the neutron configurations. Similar to the tin isotopes, the 2⁺ states for even-even cadmium from 1.8/2.0(EM) are higher than the data, and the Δ NNLO_{GO} shows a slight increase of $E(2^+)$ as the neutron number increases.

The two potentials produce more differences in the indium isotopes, which remove one proton and add neutrons to the ¹⁰⁰Sn core. The odd-mass indium nuclei, shown in Fig. 9, have $9/2^+$ ground states and a first $1/2^-$ low-lying excited state (with the exception of ¹⁰¹In) according to the experimental

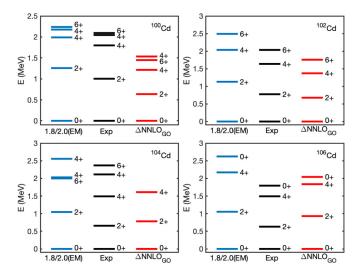


FIG. 8. Calculated low-lying excited states in Cd isotopes with the 1.8/2.0(EM) (blue) and $\Delta \text{NNLO}_{\text{GO}}$ (red) potentials, and compared to data (black).

data. The 1.8/2.0(EM) potential reproduces the correct order of the ground states and the $1/2^-$ excited state in $^{101-107}$ In. The computed $1/2^-$ state is generally higher than the data, which may be due to a too large gap between the $g_{9/2}$ and $p_{1/2}$ orbitals. The Δ NNLO_{GO} reproduces the correct level ordering in 101,103 In, but fails to reproduce the correct ground state of 105,107 In. In contrast to tin and cadmium isotopes, the indium isotopes exhibit more proton and neutron correlations. Finally, we note that uncertainties stemming from model-space truncations and the approximate treatment of normal-ordered three-body forces away from the 100 Sn core will impact the results. In the recent work [44] we estimated that empirical pairing gaps in this mass region computed with the SMCC and employing the same Hamiltonians carried an uncertainty of about ± 0.2 MeV.

IV. SUMMARY

We presented a systematic derivation of the particle-hole variant of the shell-model coupled-cluster method to compute nuclei in the vicinity of ¹⁰⁰Sn. The shell-model effective interaction is defined in a model space consisting of both particles and holes. The decoupling of the model space from the excluded space is accomplished at the two-body level, and the induced off-diagonal three-body terms are included in a linearized approximation. For nuclei in the vicinity of ¹⁰⁰Sn the particle-hole effective interaction benefits from a

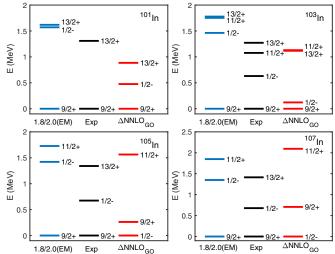


FIG. 9. Calculated low-lying excited states in In isotopes with the 1.8/2.0(EM) (blue) and $\Delta \text{NNLO}_{\text{GO}}$ (red) potentials, and compared to data (black).

more favorable reference state and more realistic mean-field than, e.g., taking a ⁸⁸Sr core. The computational resources required for nuclei close to the core are also reduced by introducing explicit hole states in the shell model. The method is validated through cross-benchmark with the existing *abinitio* methods. We derived effective interactions from two sets of chiral potentials that include nucleon-nucleon and three-nucleon forces, and computed properties of tin, indium, and cadmium isotopes. Binding energies and spectra exhibit a good to fair agreement with data.

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