β decay of the very neutron-deficient 60 Ge and 62 Ge nuclei

```
S. E. A. Orrigo , <sup>1,*</sup> B. Rubio, <sup>1</sup> W. Gelletly, <sup>1,2</sup> P. Aguilera, <sup>1,3</sup> A. Algora, <sup>1,4</sup> A. I. Morales, <sup>1</sup> J. Agramunt, <sup>1</sup> D. S. Ahn, <sup>5</sup> P. Ascher, <sup>6</sup> B. Blank, <sup>6</sup> C. Borcea, <sup>7</sup> A. Boso, <sup>8</sup> R. B. Cakirli, <sup>9</sup> J. Chiba, <sup>10</sup> G. de Angelis, <sup>11</sup> G. de France, <sup>12</sup> F. Diel, <sup>13</sup>
P. Doornenbal,<sup>5</sup> Y. Fujita,<sup>14</sup> N. Fukuda,<sup>5</sup> E. Ganioğlu,<sup>9</sup> M. Gerbaux,<sup>6</sup> J. Giovinazzo,<sup>6</sup> S. Go,<sup>15</sup> T. Goigoux,<sup>6</sup> S. Grévy,<sup>6</sup> V. Guadilla,<sup>1</sup> N. Inabe,<sup>5</sup> G. G. Kiss,<sup>4,5</sup> T. Kubo,<sup>5</sup> S. Kubono,<sup>5</sup> T. Kurtukian-Nieto,<sup>6</sup> D. Lubos,<sup>16</sup> C. Magron,<sup>6</sup> F. Molina,<sup>3</sup> A. Montaner-Pizá,<sup>1</sup> D. Napoli,<sup>11</sup> D. Nishimura,<sup>17</sup> S. Nishimura,<sup>5</sup> H. Oikawa,<sup>10</sup> V. H. Phong,<sup>5,18</sup> H. Sakurai,<sup>5,19</sup> Y. Shimizu,<sup>5</sup>
    C. Sidong, P.-A. Söderström, T. Sumikama, H. Suzuki, H. Takeda, Y. Takei, M. Tanaka, M. Tanaka, M. Wu, and S. Yagi<sup>10</sup>
                                 <sup>1</sup>Instituto de Física Corpuscular, CSIC-Universidad de Valencia, E-46071 Valencia, Spain
                                <sup>2</sup>Department of Physics, University of Surrey, Guildford GU2 7XH, Surrey, United Kingdom
                                               <sup>3</sup>Comisión Chilena de Energía Nuclear, Casilla 188-D, Santiago, Chile
                           <sup>4</sup>Institute of Nuclear Research of the Hungarian Academy of Sciences, Debrecen, H-4026, Hungary
                                              <sup>5</sup>RIKEN Nishina Center, 2-1 Hirosawa, Wako, Saitama 351-0198, Japan
       <sup>6</sup>Centre d'Etudes Nucléaires de Bordeaux Gradignan, CNRS/IN2P3, Université de Bordeaux, 33175 Gradignan Cedex, France
                ^{7}National Institute for Physics and Nuclear Engineering IFIN-HH, P. O. Box MG-6, Bucharest-Magurele, Romania
                           <sup>8</sup>INFN Sezione di Padova and Dipartimento di Fisica, Universitá di Padova, I-35131 Padova, Italy
                                                 <sup>9</sup>Department of Physics, Istanbul University, Istanbul 34134, Turkey
                                     <sup>10</sup>Department of Physics, Tokyo University of Science, Noda, Chiba 278-8510, Japan
                                              <sup>11</sup>Laboratori Nazionali di Legnaro INFN, I-35020 Legnaro, Padova, Italy
        <sup>12</sup>Grand Accélérateur National d'Ions Lourds (GANIL), CEA/DRF-CNRS/IN2P3, Bvd Henri Becquerel, 14076 Caen, France
                                       <sup>13</sup>Institute of Nuclear Physics, University of Cologne, D-50937 Cologne, Germany
                                         <sup>14</sup>Department of Physics, Osaka University, Toyonaka, Osaka 560-0043, Japan
                       <sup>15</sup>Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996-1200, USA
                                  <sup>16</sup>Physik Department E12, Technische Universität München, D-85748 Garching, Germany
               <sup>17</sup>Department of Natural Sciences, Tokyo City University, 1-28-1 Tamazutsumi, Setagaya-ku, Tokyo 158-8557, Japan
                             <sup>18</sup>Faculty of Physics, VNU University of Science, 334 Nguyen Trai, Thanh Xuan, Hanoi, Vietnam
                               <sup>19</sup>Department of Physics, University of Tokyo, 7-3-1 Hongo, Bunkyo, Tokyo 113-0033, Japan
```

(Received 25 August 2020; revised 30 November 2020; accepted 18 December 2020; published 29 January 2021; corrected 3 March 2022)

We report here the results of a study of the β decay of the proton-rich Ge isotopes, 60 Ge and 62 Ge, produced in an experiment at the RIKEN Nishina Center. We have improved our knowledge of the half-lives of 62Ge [73.5(1) ms] and 60 Ge [25.0(3) ms] and its daughter nucleus, 60 Ga [69.4(2) ms]. We measured individual β delayed proton and γ emissions and their related branching ratios. Decay schemes and absolute Fermi and Gamow-Teller transition strengths have been determined. The mass excesses of the nuclei under study have been deduced. A total β -delayed proton-emission branching ratio of 67(3)% has been obtained for ⁶⁰Ge. New information has been obtained on the energy levels populated in 60 Ga and on the $1/2^-$ excited state in the βp daughter ⁵⁹Zn. We extracted a ground state-to-ground state feeding of 85.3(3)% for the decay of ⁶²Ge. Eight new γ lines have been added to the deexcitation of levels populated in the ⁶²Ga daughter.

DOI: 10.1103/PhysRevC.103.014324

I. INTRODUCTION

The investigation of nuclear structure close to the limits of nuclear existence is one of the frontiers of modern nuclear physics. The study of the properties of exotic nuclei is crucial to provide tests of the predictions of nuclear models at extreme values of isospin. The experimental challenge involved with the production of such unstable nuclei has resulted in a worldwide effort to build next-generation facilities producing and accelerating radioactive ion beams (RIBs) of very high intensity. Heavy neutron-deficient nuclei can now be produced up to the proton drip-line, enabling the realization of detailed decay studies and the observation of new exotic decay modes [1-4].

The structural properties of exotic proton-rich nuclei are also important for Nuclear Astrophysics, because many of them lie on the rapid proton-capture process (*rp*-process) reaction pathway leading to the production of the heavy elements in the Universe in explosive stellar environments [5–8].

Decay spectroscopy experiments with implanted RIBs are a powerful tool to explore the properties of exotic nuclei and provide rich spectroscopic information: β -decay half-lives, delayed emission of γ rays, particle-decay branching ratios

^{*}sonja.orrigo@ific.uv.es

(commonly protons are emitted in the case of proton-rich nuclei), excited states populated in the daughter nucleus, and so on. Furthermore β decay provides direct access to the absolute values of the Fermi B(F) and Gamow-Teller B(GT) transition strengths.

In this paper we present new results on the β decay of the neutron-deficient Ge isotopes, 62 Ge and 60 Ge. They were produced and implanted in unprecedented numbers at the Radioactive Ion Beam Factory (RIBF) of the RIKEN Nishina Center (Japan) thanks to the availability of a high-intensity beam of 78 Kr. 62 Ge is a $T_z=-1$ nucleus about which little was known at the time of our experiment. 60 Ge is a $T_z=-2$ nucleus whose decay is almost unknown, apart from a first measurement of its half-life with 28 events [9]. 60 Ge is a semimagic N=28 isotone and the heaviest one for which the mirror 60 Ni nucleus is stable, allowing one to explore mirror symmetry. The 60 Ga daughter lies right at the proton drip-line and nothing is known of its level scheme. Hence our study provides brand new information.

In a more general context, this work is part of a systematic study of proton-rich nuclei which we have carried out at different RIB facilities [3,4,10–12], also in comparison with mirror charge-exchange experiments done on the mirror stable target [13]. One focus of interest for the $T_z=-2$ nuclei such as ⁶⁰Ge is to explore the competition between the γ deexcitation and (isospin-forbidden) proton emission from the T=2 isobaric analog state (IAS), that is populated in the daughter nucleus by the β decay. This feature has been observed in all the lighter $T_z=-2$ systems already studied [3,4]. The $T_z=-1$ nuclei such as ⁶²Ge are also of interest to check whether the suppression of the isoscalar γ transitions between $J^{\pi}=1^+$, T=0 states (Warburton and Weneser quasirule [14,15]) observed in previous $T_z=-1$ cases [12] persists.

The paper is organized as follows. Section II describes the experiment and the setup. The details of the data analysis are described in Sec. III. The results obtained on the β^+ decay of 60 Ge are given in Sec. IV, together with new results on the β^+ decay of the 60 Ga daughter. Section V shows the results on the β^+ decay of 62 Ge. Finally, Sec. VI summarizes the conclusions.

II. THE EXPERIMENT

The ⁶⁰Ge and ⁶²Ge proton-rich nuclei were produced with unprecedented intensity in an experiment performed at RIBF, RIKEN Nishina Center (Japan). A high-intensity (up to 250 pnA) ⁷⁸Kr primary beam was accelerated to 345 MeV/nucleon and fragmented on a ⁹Be target with thickness of 5 mm. The exotic fragments produced in this way were separated, selected and identified in the BigRIPS separator [16,17] by means of the $B\rho - \Delta E$ – ToF method [18]. A series of parallel-plate avalanche counters, multisampling ionization chambers and plastic scintillators constitute the detection setup of BigRIPS and are employed to measure the position of the transmitted ions at different focal planes (which is related to the magnetic rigidity $B\rho$), their energy loss ΔE and time of flight (ToF), respectively. Standard particleidentification procedures, together with additional off-line cuts on the BigRIPS variables to remove background events

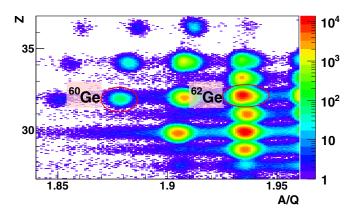


FIG. 1. Z versus A/Q identification plot for the BigRIPS setting optimized for 65 Br, in which both 60 Ge and 62 Ge are implanted. The positions of these nuclei are shown.

thoroughly, allow one to identify the ions by their atomic number Z and mass-to-charge ratio A/Q with high resolution [18,19]. An example of the two-dimensional identification matrix obtained for the BigRIPS setting optimized for 65 Br is shown in Fig. 1, where the positions of 60 Ge and 62 Ge are indicated. It should be noted that the odd 59 Ge and 61 Ge isotopes were also produced and their decay was studied in Refs. [19,20].

The selected and identified fragments were then transmitted to the exit of the ZeroDegree spectrometer (ZDS) [17] where a setup for β -decay spectroscopy was installed. This setup consisted of the Wide-range Active Silicon Strip Stopper Array for Beta and ion detection (WAS3ABi) [21] and the EUroball-RIKEN Cluster Array (EURICA) [22]. A common timestamp with a frequency of 10^8 Hz was used to correlate the data acquisition systems of BigRIPS, WAS3ABi, and EURICA.

The WAS3ABi array was used to detect both the implanted heavy-ions and subsequent charged-particle (β particles and protons) decays, measuring their energy, position and time. WAS3ABi consisted of three 1-mm-thick double-sided silicon-strip detectors (DSSSD), each one having a 6×4 cm² area. Each DSSSD was segmented into 60 vertical (X) and 40 horizontal (Y) strips with a pitch of 1 mm, defining a total of 2400 pixels. Each strip was read individually by standard analog electronics providing energy and time signals. A plastic scintillator was placed behind WAS3ABi acting as a veto, i.e., tagging the fragments that were not stopped inside WAS3ABi and also those undergoing secondary reactions in WAS3ABi.

In order to optimize the resolution for the β particles, the gain was set to achieve a full energy range of 4 MeV for the X strips, while the energy range of the Y strips was increased to 10 MeV to allow the detection of high-energy protons. Known β -delayed proton emitters (57 Zn, 61 Ge, and 65 Se) produced during the experiment, together with conversion electrons from a 207 Bi source, were used for the energy calibration of WAS3ABi [20].

An implantation signal in WAS3ABi is characterised by an overflow energy signal. This is because the implanted fragments reach the array with energies well above 1 GeV and, with the gain settings mentioned above, they saturate the electronics. Since it is not only the implantation strip that saturates, but also the neighboring ones, the position (X, Y) at which the implantation took place was determined by the fastest X and Y strips to fire.

WAS3ABi was surrounded by the EURICA array [22], employed to detect both the prompt and β -delayed γ rays after the implantation event and measure their energies and times. In the present experiment EURICA used a total of 84 high-purity germanium detectors, arranged in 12 clusters containing 7 Ge crystals each. The γ -ray signals were acquired up to 100 μ s after WAS3ABi provided a trigger in order to allow for the detection of the decay of isomeric states. EURICA was calibrated in energy using γ rays from ¹³³Ba, ¹⁵²Eu and ²⁴⁴Cm sources [22]. The γ efficiency was calibrated using the ¹³³Ba and ¹⁵²Eu sources, together with known γ rays from ¹²⁹Cd and ¹³²Sn observed in a previous experiment [23]. The array had an absolute detection efficiency of 10% at 662 keV [24].

III. THE DATA ANALYSIS

During the experimental campaign BigRIPS was optimized for different settings [20,24]. In the present paper we report the analysis of data from two of them. The first was a setting optimised for 65 Br, where both 60 Ge and 62 Ge were implanted. The second was optimised for 64 Se, where only 62 Ge was implanted. The WAS3ABi deadtime fraction τ , determined by comparing free and accepted triggers recorded for each run, was 22(3)% for the first setting (60 Ge) and 26(3)% for the combination of both settings (62 Ge).

Provided that the ions were identified in BigRIPS as explained in Sec. II, an implantation event had also to simultaneously satisfying the following conditions: (a) a signal in the last fast-plastic scintillator (F11) of the BigRIPS+ZDS spectrometer setup, (b) an overflow energy signal in WAS3ABi and (c) no signal in the veto plastic behind WAS3ABi. A β -decay event was defined as an event simultaneously satisfying the following conditions: (A) no signal in F11, (B) no overflow energy signal in the Y strips of WAS3ABi and (C) an energy signal above threshold (typically 50 keV) in WAS3ABi.

Standard techniques, extensively described in Sec. III of Ref. [4], were employed to perform the data analysis, starting with the time correlations between implantation and decay events, the construction of the charged-particle decay-energy spectrum and the β -delayed γ -ray energy spectrum for decay events correlated with implantations of a given nuclear species and the determination of the Fermi and Gamow-Teller transition strengths.

In particular, in the present case, for a specified DSSSD in WAS3ABi each implantation event was correlated in time with any decay event occurring before and after it in the same pixel and in the eight pixels surrounding it. This is because given the smaller size of the pixels, namely 1 mm², the truly correlated decay event may also occur in one of the closest pixels. This procedure accounts for the true time correlations, but at the price of introducing many random correlations. These latter will contribute to a flat background in the correlation-time spectrum which adds to the typical decay curve described by the Bateman equations [25] and can easily be taken into account in the half-life fit [4]. Here

the time correlations were performed over a period of time of [-1 s, +1 s], where the backward-correlations part [-1 s, 0] clearly includes only random correlations and is used in the background estimation procedure. The half-life results are reported with their statistical uncertainties.

For a given nuclear species, the charged-particle decayenergy spectrum measured in WAS3ABi was obtained as the sum of the spectra from all the pixels. As in Ref. [4] a clean decay-energy spectrum was formed by subtracting a background energy spectrum, constructed by setting a gate of [-1 s, 0] on the correlation time, from the decay-energy spectrum constructed in [0, 1 s].

The γ -ray energy spectrum correlated with implantations of a given nuclear species was obtained by summing all the Ge crystals of EURICA, producing spectra both with and without addback. It should be noted that in this paper we show those with addback. A similar background subtraction procedure was employed to obtain a clean γ -energy spectrum, using the same gates of [-1 s, 0] and [0, 1 s] on the correlation time. In addition, the γ -ray time was restricted to [0, 800 ns] with respect to the decay signal in WAS3ABI.

IV. β DECAY OF ⁶⁰Ge

The exotic neutron-deficient 60 Ge nucleus is a special system to study. It is a semimagic N=28 isotone and the heaviest one for which the mirror nucleus (60 Ni) is stable, which makes it easier to investigate mirror symmetry. Furthermore, the decay of 60 Ge is almost unknown. Nothing is known about the level scheme of the 60 Ga daughter, which is a nucleus lying right at the proton drip-line. There exists a first measurement of the half-life of 60 Ge [9] with 28 implantation events (19 of them with observation of β -delayed protons). The RIKEN high-intensity 78 Kr beam allowed us to achieve the unprecedented statistics of 1.5×10^4 implants of 60 Ge. Moreover, 60 Ge is a $T_z = -2$ nucleus, hence its study may shed light on the possible competition between the β -delayed proton emission and γ deexcitation from the IAS in the daughter, as already observed in the decay of lighter $T_z = -2$ systems [3,4,26].

The charged-particle energy spectrum for decay events correlated with ⁶⁰Ge implants is shown in Fig. 2. The bump visible in Fig. 2(a) at low energy is due to the detection of β particles and is fitted by an exponential function (green line). Eight peaks are observed at higher energy, corresponding to β -delayed proton emission. Their fit is shown in Fig. 2(b). The best candidate for the deexcitation of the IAS of ⁶⁰Ge in ⁶⁰Ga is the most intense peak at $E_p = 2522 \text{ keV}$ which, as explained later in the text, corresponds to an excitation energy E_X 2612 keV in ⁶⁰Ga. As expected in this kind of experiment, the summing of the proton signals with those of the coincident β particles affects the line shape of the peaks. Following our procedure [3,4], the line shape was determined by Monte Carlo simulations and checked by fitting the well-known ⁵⁷Zn peaks. The decay energies E_p and their intensities I_p per 100 decays are shown in Table I. It should be noted that E_p is the total energy released in the proton decay including the energy of the recoiling nucleus.

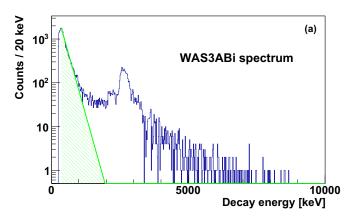
Figure 3 shows the correlation-time spectrum obtained for 60 Ge selecting only the β -delayed proton decays, by setting

TABLE I. Results for the β^+ decay of 60 Ge. The first two columns show information on the β -delayed proton emission: The decay energies E_p and their intensities I_p (normalized to 100 decays). The last two columns report the γ -ray energies E_{γ} and their intensities I_{γ} (normalized to 100 decays of 60 Ge). The observed γ lines belong to different nuclei in the decay chain of 60 Ge (see the text).

β -delayed proton emission		β -delayed γ deexcitation		
$E_p \text{ (keV)}$	I_p (%)	E_{γ} (keV)	Ι _γ (%)	
820(13)	2.8(4)	463.3(1)	8(1)	
1076(23)	4.0(5)	491.2(2)	7(1)	
1359(19)	5.1(4)	837.2(2)	9(1)	
1684(17)	4.2(3)	1003.3(2)	15(2)	
2067(15) ^a	10.2(5)	1332.3(4)	4(1)	
2522(15)b	33(1)	1774.6(9)	2(1)	
2981(23)	3.2(3)	3848.0(9)	13(3)	
3490(22)	1.9(2)	. ,	. ,	

^aIAS proton emission to the first excited state of ⁵⁹Zn.

^bIAS proton emission to the ground state of ⁵⁹Zn.



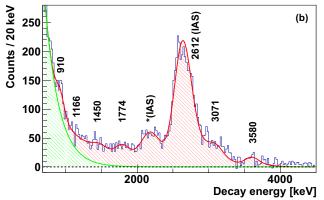


FIG. 2. (a) Charged-particle energy spectrum for decay events correlated with 60 Ge implants. The peak due to the detection of β particles is seen at low energy and fitted by an exponential function (in green). (b) Fit of the peaks related to the proton emission following the β decay of 60 Ge (in red). Peaks are labeled according to the corresponding excitation energies in the β daughter 60 Ga. The peak marked by an asterisk corresponds to the transition from the IAS to the first excited state in 59 Zn (see text).

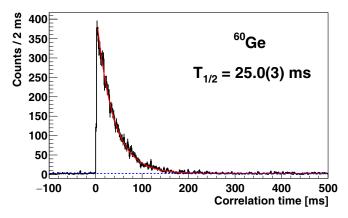


FIG. 3. Spectrum of the time correlations between each proton decay (WAS3ABi energy $E_p > 1$ MeV) and all 60 Ge implants. The dashed blue line is the background (fixed by fitting the backward-correlations part) and the red line is the result of the fit.

the condition WAS3ABi energy $E_p > 1$ MeV. The half-life is determined by a least-squares fit to the data including the parent activity and a linear background, fixed by a fit to the backward-correlations part, as in Ref. [4]. A value of $T_{1/2} = 25.0(3)$ ms was obtained for ⁶⁰Ge. A fit using a maximum likelihood minimization method gives a consistent result. We have improved the precision on the half-life value in comparison with the only existing measured value of 20^{+7}_{-5} ms [9].

The total proton-emission branching ratio B_p is determined, as in Ref. [4], by comparing the total number of protons, N_p (obtained from the fit shown in Fig. 3), with the total number of implanted nuclei, N_{imp} , according to:

$$B_p = \frac{N_p}{N_{\text{imp}} (1 - \tau)},\tag{1}$$

where τ is the deadtime fraction (see Sec. III). A value $B_p = 67(3)\%$ is obtained, where the uncertainty takes into account both the statistical and systematics ones. The latter was estimated as previously [4,26] by repeating the determination of B_p changing the condition on the energy E_p by ± 100 keV.

Figure 4 shows the γ -ray energy spectrum from the β decay of 60 Ge ($T_z=-2$) to 60 Ga ($T_z=-1$). Besides the 511 keV γ line associated with the annihilation of the positrons emitted in the β decay, seven γ lines are observed at 463, 491, 837, 1004, 1332, 1775, and 3848 keV. These γ transitions belong to different nuclei in the decay chain of 60 Ge, as we discuss in detail below. The γ lines which we observe for the first time are underlined in the figure. In particular, the γ lines at 837, 1332, and 1775 keV are associated with the γ deexcitation of states in the 60 Ga daughter. The energies and intensities of all the γ rays observed in Fig. 4 are given in Table I. The intensities I_{γ}^i are obtained as:

$$I_{\gamma}^{i} = \frac{N_{\gamma}^{i}}{\epsilon_{\gamma}^{i} \epsilon_{\text{decay}} N_{\text{imp}} (1 - \tau)}, \tag{2}$$

where N_{γ}^{i} represents the number of counts obtained from the integral of the *i*th γ peak and ϵ_{γ}^{i} is the corresponding efficiency for γ detection (Sec. II). The data acquisition system was triggered by either an implantation event or a decay event

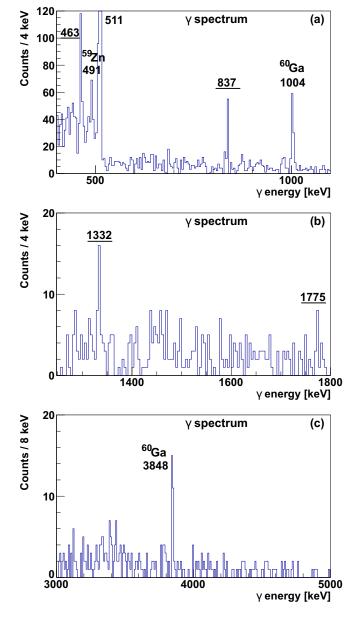


FIG. 4. γ -ray energy spectrum for decay events correlated with 60 Ge implants. The observed γ lines belong to different nuclei in the decay chain of 60 Ge. Underlined values indicate γ lines which are seen for the first time in the present work (see the text). Known γ lines are labeled according to the parent nucleus. (a) The region [400,1100] keV is shown with bins of 4 keV/channel. (b) The region [1250,1800] keV is shown with bins of 4 keV/channel. (c) The region [3000,5000] keV is shown with bins of 8 keV/channel.

and the γ rays were not included in the trigger. They were acquired in coincidence with decay events (β or β p) and thus affected by the decay detection efficiency $\epsilon_{\rm decay}$. This efficiency was estimated as in Ref. [4] and it was 58(2)% for γ rays coincident with pure β events, and 1 for the 463-keV γ ray coincident with β p events.

The reader may refer to Fig. 8 for a summary of the decays that intervene in the following discussions. Below we discuss

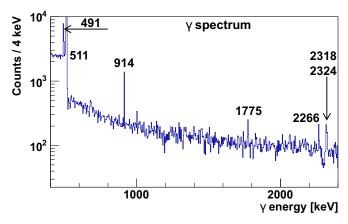


FIG. 5. γ -ray energy spectrum for decay events correlated with 59 Zn implants, showing the γ deexcitation in 59 Cu. 59 Zn is also populated in the β -delayed proton decay of 60 Ge.

in detail the origin of all the observed γ lines. As shown in Fig. 2, β -delayed proton emission takes place in the β decay of ⁶⁰Ge. This process, passing through proton-unbound excited states of ⁶⁰Ga, populates states in ⁵⁹Zn ($T_z = -1/2$). ⁵⁹Zn then β decays to its mirror nucleus, ⁵⁹Cu ($T_z = +1/2$) with $T_{1/2} = 182.4(4)$ ms (see below). The 491-keV γ line is known and is the transition connecting the $1/2^-$ first excited state of ⁵⁹Cu to its $3/2^-$ ground state (g.s.) [27,28].

Analyzing the direct implantation of 59 Zn (2.5×10⁶ implants) we obtained $T_{1/2} = 182.4(4)$ ms, that may be compared with the previous values of 210(20) [27], 182.0(18) [28], 173(14) [29], and 174(2) [11] ms. We also obtained the γ ray energy spectrum correlated with ⁵⁹Zn implants (shown in Fig. 5) where, as expected, we observe known γ lines deexciting levels in 59 Cu. Comparing the γ lines of Figs. 4 and 5, the new γ lines (463, 837, and 1332 keV) are not present in Fig. 5, while in both figures there is a line at 1775 keV. A known 1775-keV γ ray deexcites the 2267 level in the ⁵⁹Cu daughter nucleus, populating the 491-keV level. We think that the 1775-keV line of Fig. 4 belongs to a different transition, because otherwise we should have also observed the more intense 914-keV line in Fig. 4. The proton separation energy in ⁵⁹Zn is 2836.8(7) keV [30] and the total β -delayed proton-emission branching ratio is 0.023(8)% [27], hence the contribution of the proton emission from ⁵⁹Zn is negligible in the decay chain of ⁶⁰Ge.

As mentioned above, the first excited state in 59 Cu is $1/2^-$ with 491 keV excitation energy and deexcites to its $3/2^-$ ground state through a γ transition of the same energy [27,28]. The ground state of 59 Zn is also $3/2^-$ [27] but the energy of the expected $1/2^-$ excited state is not known. The isopin symmetry and the expected similarity between 59 Cu and 59 Zn suggest that the 463-keV γ line of Fig. 4 may be associated with the transition between the expected $1/2^-$ excited state of 59 Zn and its $3/2^-$ ground state. An additional hint at this conclusion comes from the energy difference between the proton peaks at $E_p = 2067(15)$ and 2522(15) keV (Table I) which agrees, within the uncertainty, with the value of 463 keV. This suggests that the two proton peaks are indeed a doublet corresponding to the transitions from the IAS in 60 Ga

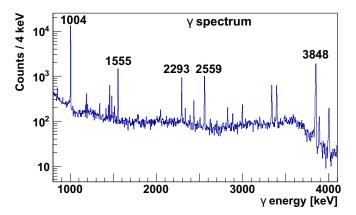


FIG. 6. γ -ray energy spectrum for decay events correlated with 60 Ga implants, showing the γ deexcitation in 60 Zn. The 60 Ga nucleus is also populated in the β decay of 60 Ge.

to the first excited and ground states of 59 Zn, respectively. It should be noted that the parent of 59 Zn, 59 Ga, is probably unbound, therefore the β -delayed proton decay from 60 Ge is presumably the best way to populate the first excited state at 463 keV in 59 Zn.

On the other hand, the β decay of ⁶⁰Ge populates directly states in the 60 Ga daughter. The latter then β decays to 60 Zn (with $T_z = 0$ and being its own mirror since Z = N = 30) populating its IAS at 4852 keV, which deexcites mainly by the cascade of 3848- and 1004-keV γ rays [31] that are both observed in Fig. 4. Looking at the direct implantation of ⁶⁰Ga $(7.6 \times 10^5 \text{ implants})$, we obtained the γ -ray energy spectrum of Fig. 6, showing the γ deexcitation in 60 Zn. In this spectrum we observe all the most intense γ transitions at 1004, 1555, 2293, 2559, and 3848 keV, already seen in Ref. [31], and several other γ lines listed in Table II. The intensities of the five known γ transitions agree with those observed in Ref. [31]. None of the new γ lines (837, 1332, and 1775 keV) of Fig. 4 is observed in Fig. 6. A value $B_p = 1.6(7)\%$ was obtained for the β -delayed proton emission from ⁶⁰Ga [31], populating ⁵⁹Cu. Thus this process could also make a minor contribution to the observed 491 keV γ line (Fig. 4), even if we do not see this line in Fig. 6. Analyzing the charged-particle spectrum for decay events correlated with ⁶⁰Ga implants, we found the contribution of the β -delayed proton emission from 60 Ga to be of the order of 1% in comparison to that from ⁶⁰Ge (Fig. 2). Last, we have improved our knowledge of the ⁶⁰Ga half-life, obtaining $T_{1/2} = 69.4(2)$ ms from the fit shown in Fig. 7, that may be compared with the previous values of 70(15) [31] and 76(3) ms [11].

The mass of 59 Zn is known [32], but the mass of 60 Ga has not been measured, which makes difficult to estimate properly the excitation energy E_X of the IAS and other nuclear levels populated in the 60 Ga daughter based on the observation of the proton peaks. The mass excess of 60 Ge has also not been measured. From systematics (#) the most recent atomic mass evaluation (AME) [32] gives a ground-state mass excess of $-27090^{\#}(300^{\#})$ keV for 60 Ge and $-39590^{\#}(200^{\#})$ keV for 60 Ga. From these masses the value $Q_{\beta} = 12500^{\#}(360^{\#})$ keV is obtained [30]. The proton separation energy in 60 Ga is

TABLE II. γ rays observed in the β^+ decay of 60 Ga. The γ -ray energies E_{γ} and intensities I_{γ} (normalized to 100 decays of 60 Ga) from the present study are reported in the first two columns. For comparison, the last columns show E_{γ} and I_{γ} from Ref. [31], where the fourth column shows I_{γ} relative to the 1004 keV γ line, while the fifth column shows absolute I_{γ} for 100 decays of 60 Ga (calculated using our absolute intensity for the 1004 keV γ ray).

E_{γ} (keV)	I_{γ} (%)	$E_{\gamma} (\text{keV})^{a}$	$I_{\gamma}^{\mathrm{rel a,b}}$	$I_{\gamma} (\%)^{\mathrm{a,c}}$
669.3(3)	0.36(9)			
850.8(2)	0.75(10)			
913.9(3)	0.3(1)			
1003.5(1)	62(3)	1003.7(2)	100(17)	62(11)
1028.6(3)	0.38(8)			
1188.4(2)	1.6(1)			
1201.8(3)	0.29(7)			
1413.7(3)	0.35(7)			
1442.1(3)	0.40(8)			
1481.4(3)	1.3(1)			
1554.7(3)	7.0(5)	1554.9(6)	12(5)	7.4(31)
1780.8(7)	0.2(1)			
2047.2(6)	0.7(2)			
2293.2(4)	6.3(5)	2293.0(10)	10(5)	6.2(31)
2334.2(5)	0.8(2)			
2434.2(4)	1.8(2)			
2558.7(4)	8.5(6)	2559.0(8)	13(5)	8.1(31)
2624.3(7)	0.3(1)			
2826.0(5)	1.3(2)			
2884.0(5)	0.8(2)			
2996.8(5)	2.0(3)			
3337.4(4)	7.1(6)			
3394.8(4)	7.0(6)			
3848.5(4)	38(3)	3848.3(7)	57(13)	35(8)
3889.1(5)	2.8(8)			
4000.9(5)	2.8(4)			
4805.0(5)	0.4(1)			
4850.2(6)	0.2(1)			
4891.9(5)	0.4(1)			

^aFrom Ref. [31].

also not known experimentally. Systematics gives us $S_p = -340^{\#}(200^{\#})$ keV [30], while a semiempirical estimate gives 40(70) keV [31]. A systematic value for the excitation energy of the IAS is $E_X^{\rm IAS} = 2540^{\#}(50^{\#})$ keV [32]. Reference [9] calculated a value of 2520(280) keV based on Coulomb energy difference.

From our experimental data we can obtain the mass excess of the IAS in 60 Ga. The information from the β -delayed proton emission gives us a proton energy $E_p^{\rm IAS} = 2522(15)$ keV (Table I), remembering that $E_X = E_p + S_p$. Using our $E_p^{\rm IAS}$ and the measured value for the mass excess of the $^{59}{\rm Zn}_{\rm g.s.}$, -47215.6(8) keV [32], one can determine the mass excess of $^{60}{\rm Ga}^{\rm IAS}$ being -37405(15) keV. Considering the latter mass together with the measured masses for the three members of the T=2 isospin multiplet ($^{60}{\rm Zn}^{\rm IAS}$, $^{60}{\rm Cu}^{\rm IAS}$)

^bRelative intensity.

^cAbsolute intensity (calculated using our measured intensity for the $1004 \text{ keV } \gamma$ ray).

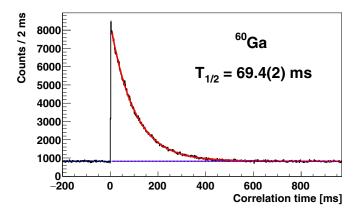


FIG. 7. Fit of the correlation-time spectrum obtained for the β decay of 60 Ga (red line). The fit includes the parent activity (green line practically coincident with the red line), the daughter activity (purple line) and the background (dashed blue line) fixed by fitting the backward-correlations part.

and 60 Ni_{g.s.}, listed in Table III), one can apply the isobaric multiplet mass equation (IMME) [33,34] to determine the 60 Ge_{g.s.} mass excess, obtaining a value of -27678(22) keV. The 2003 AME value of $-27770^{\#}(230^{\#})$ keV [35] agrees with our value better than the most recent AME mass excess. In Sec. V we observe similar behavior for the case of 62 Ge. Issues with AMEs subsequent to that in 2003 have been reported already for proton-rich nuclei in this region of the mass chart [3,4,36] and we believe that more mass measurements would be very useful to constrain the future AME in this mass region.

The IAS in 60 Cu, which is the mirror nucleus of 60 Ga, lies at 2536.0(6) keV and deexcites through a γ cascade with γ energies 1866 and 670 keV [37]. Based on mirror symmetry, together with the observed intensities for the 837 and 1775 keV γ rays (Table I), we think that the IAS in 60 Ga deexcites through a γ cascade consisting of the 1775 and 837 keV γ rays, in this order. From the energies of these γ rays we can determine the excitation energy of the 60 Ga IAS , $E_X^{IAS} = 2611.8(9)$ keV, which agrees with the expected values [9,32]. We can also deduce other properties of 60 Ga, such as the ground-state mass excess, -40016(15) keV, and $S_p = 90(15)$ keV. Finally, we calculate the value $Q_\beta = 12338(27)$ keV for the β^+ decay of 60 Ge.

TABLE III. Mass excesses (in keV) of the four members of the A=60, T=2 mass multiplet that are used as the input for the IMME calculations. Mass excess (in keV) of the $T_z=-2$, $^{60}\mathrm{Ge_{g.s.}}$ nucleus obtained by the IMME calculations and compared with the 2016 and 2003 AME systematics [32,35].

A = 60, T = 2 input values				
$T_z = +2 -64473.1(4)^a$	$T_z = +1 \\ -55804(5)^a$	$T_z = 0$ -46807(24) ^a	$T_z = -1 -37405(15)^{\mathbf{b}}$	
IMME result -2767	~	2016 AME [32] -27090°(300°)	2003 AME [35] -27770°(230°)	

^aFrom Ref. [32].

TABLE IV. Summary of the results for the β^+ decay of ⁶⁰Ge. Level excitation energies E_X in ⁶⁰Ga, β feedings I_{β} , Fermi B(F), and Gamow-Teller B(GT) transition strengths.

E_X (keV)	I_{β} (%)	B(F)	B(GT)
3580(27)	1.9(2)		0.14(1)
3071(28)	3.2(3)		0.18(2)
2611.8(9) ^a	45.3(20)	3.1(1)	. ,
1774(23)	4.2(3)	. ,	0.11(1)
1450(25)	5.1(4)		0.11(1)
1166(28)	4.0(5)		0.074(9)
910(20)	2.8(4)		0.044(6)
837.2(2)	7(2)		0.11(3)

aIAS.

The quantities deduced above are used to determine the excitation energies E_X of the levels in 60 Ga, the Fermi B(F) and Gamow-Teller B(GT) transition strengths. They are reported in Table IV together with the β feeding I_{β} to each level, inferred from the intensities measured for both the β -delayed proton and γ emissions (Table I). The level scheme deduced from the results of the present experiment is shown in Fig. 8.

We obtain a value B(F) = 3.1(1) that is smaller than the expected B(F) = |N - Z| = 4. From Table IV, $I_{\beta}^{IAS} = 45.3\%$ but we would need a feeding of 58% in order to get the correct Fermi strength. A possible explanation could be that we are missing weak γ branches deexciting the IAS. As a matter of fact, we do not have enough information to place the 1332 keV γ ray in the level scheme. If it deexcites the IAS, then this latter would gain an additional feeding of 4% (see Table I) out of the missing 12.7%.

Another possible scenario would be to invert the order of the γ rays in the cascade deexciting the IAS, i.e., first the 837- and then the 1775-keV γ rays. In this way the IAS would gain an additional feeding of 7% and we would obtain a higher value for the strength, B(F) = 3.6(2). According to this possibility, the 837-keV γ ray would populate the level at $E_X = 1774$ keV, which we already observed from the proton emission, but the balance of feeding populating and deexciting this level would be negative [-3.0(15)%]. The 837-keV level would be removed from the decay scheme of Fig. 8 and for this reason we indicated it with a dashed line. If this second scenario is confirmed, then the emission of the 837-keV γ ray followed by the emission of the proton from the 1774-keV level would be a new observation of β -delayed γ -proton decay, an exotic decay mode that we already observed in ⁵⁶Zn [3]. A new experiment with higher statistics would be desirable to shed light on these aspects.

In the decay of the $T_z=-2$ nuclei the β -delayed proton emission from the IAS populated in the daughter nucleus is isospin-forbidden, but usually it is observed and this is attributed to a T=1 isospin impurity in the IAS wave function [3,4]. As in other $T_z=-2$ cases, in the decay of ⁶⁰Ge we observe competition between the β -delayed proton and γ emissions from the IAS populated in ⁶⁰Ga. However here the proton emission is found to be dominant, being 95% of the observed IAS deexcitation (74.5% if we account for the

 $^{{}^{\}mathrm{b}}\mathrm{From}$ the present β -delayed proton emission data.

^cValues obtained from systematics.

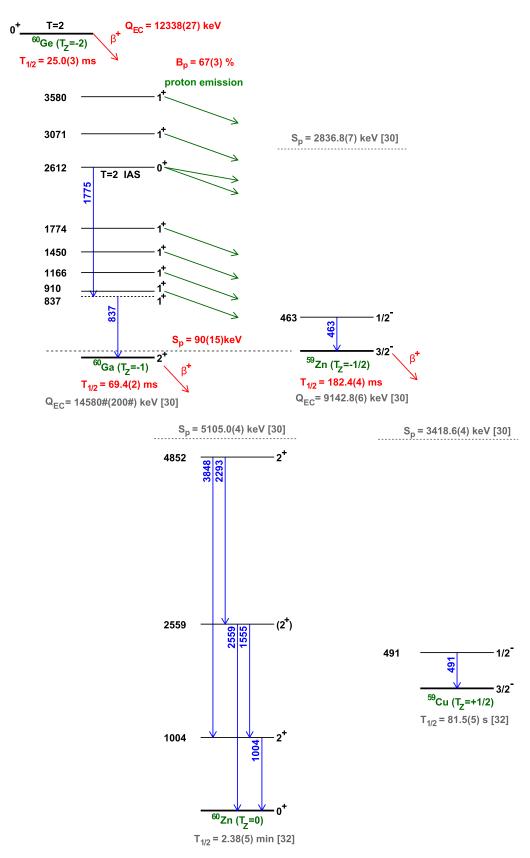


FIG. 8. Partial decay scheme of 60 Ge deduced from the results of the present experiment. Evidence is seen of population of levels in four nuclei in the decay chain of 60 Ge: 60 Ga, 60 Zn, 59 Zn, and 59 Cu (see the text for more details). The observed γ transitions are indicated by blue arrows, while green arrows indicate the observed proton emission. The quantities in red are deduced from the present data.

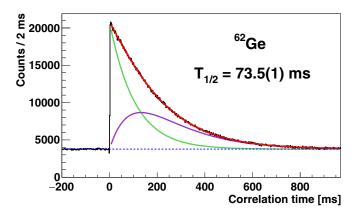


FIG. 9. Fit of the correlation-time spectrum obtained for the β decay of 62 Ge. The green line indicates the parent activity, the purple line is the daughter activity, the dashed blue line is the background (fixed by fitting the backward-correlations part of the spectrum) and the red line is the sum of all.

missing γ feeding). The explanation for this behavior may lie in nuclear structure reasons.

V. β DECAY OF 62Ge

 62 Ge is a $T_z=-1$ proton-rich nucleus which was only poorly known at the time of the present experiment. The high-intensity 78 Kr beam available at RIKEN made it possible to record the unprecedented statistics of 2.1×10^6 implants of 62 Ge.

Figure 9 shows the correlation-time spectrum obtained for the β decay of 62 Ge. The 62 Ge half-life was determined by fitting the correlation-time spectrum using the Bateman equations [25] including the β decay of 62 Ge, the growth and decay of its daughter 62 Ga (with $T_{1/2}=116.121(21)$ ms [32]) and the random correlation background. A half-life of $T_{1/2}=73.5(1)$ ms is obtained. The maximum likelihood and least-squares minimization methods gave the same result. This half-life value agrees with and improves the precision of our previous measurement (76(6) ms obtained with 6.1×10^3 implants of 62 Ge [11]). As already discussed in Ref. [11], there is an older measurement, 129(35) ms [29], lying inside two standard deviations mainly because of the poorer statistics. Additionally, another value of 82.9(1.4) ms is reported in Ref. [38] again with much poorer statistics.

The charged-particle energy spectrum from the decay of 62 Ge is shown in Fig. 10. The bump at low energy is due to the detection of β particles, while no discrete peaks from β -delayed proton emission are observed. The proton separation energy in the 62 Ga daughter is 2927(16) keV [30], above this value proton emission is allowed. From our data we estimated an upper limit of 5‰ for the total proton-emission branching ratio.

The γ -ray energy spectrum from the β decay of ⁶²Ge is shown in Fig. 11. We observed twelve γ lines at 571, 978, 1017, 1117, 1360, 1899, 2011, 2164, 2642, 2967, 3340, and 3595 keV. Eight of them were observed for the first time, while the four most intense lines (571, 978, 1017, and 2164 keV) were also seen in Ref. [38] together with two other γ lines at

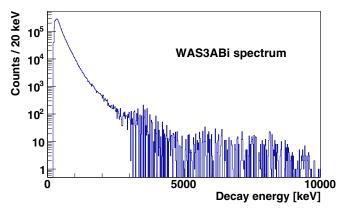


FIG. 10. Charged-particle energy spectrum for decay events correlated with ⁶²Ge implants.

1247 and 2414 keV which we did not see. The γ lines which we observed for the first time are underlined in Fig. 11.

The decay scheme of 62 Ge deduced from the results of the present experiment is shown in Fig. 12. The β^+ decay of the $J^{\pi}=0^+$, T=1 ground state of 62 Ge ($T_z=-1$) populates the 0^+ , T=1 IAS in the 62 Ga daughter ($T_z=0$) through a super-allowed Fermi transition and 1^+ , T=0 states through Gamow-Teller transitions. According to the quasirule of Warburton and Weneser, namely that the M1 transitions between $J^{\pi}=1^+$, T=0 states are strongly suppressed [14,15], one

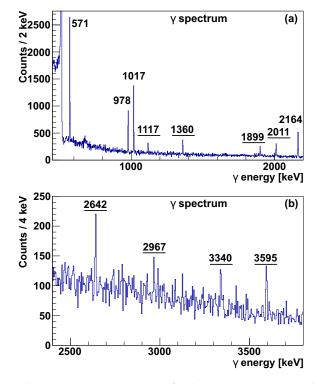


FIG. 11. γ -ray energy spectrum for decay events correlated with 62 Ge implants. Underlined values indicate γ lines which are seen for the first time in the present work. (a) The region [450,2200] keV is shown with bins of 2 keV/channel. (b) The region [2400,3800] keV is shown with bins of 4 keV/channel.

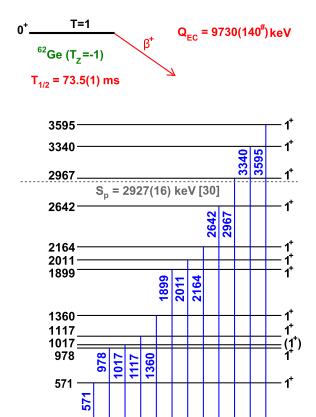


FIG. 12. Partial decay scheme of 62 Ge deduced from the results of the present experiment. The observed γ transitions are indicated by blue arrows. The quantities in red are deduced from the present data.

T=1 IAS

 62 Ga ($T_z = 0$)

 $T_{1/2} = 116.121(21) \text{ ms } [32]$

can expect that all the populated 1^+ , T=0 states decay predominantly to the 0^+ , T=1 ground state of 62 Ga. We have found evidence for this quasirule in other $T_z=-1$ cases [12]. In order to test this rule for the present case, we looked for small branches from the upper 1^+ states to the lower 1^+ levels. Even considering the most favourable case, i.e., a transition from the 1^+ state at 2164 keV to the first 1^+ state at 571 keV, the deexciting γ ray was not observed. We set an upper limit of 1 ‰ on its intensity.

In a fusion-evaporation-reaction study of 62 Ga [42] a $^{2+}$ level was proposed at 1017 keV based on the observation of a weak γ ray at 446 keV in coincidence with the 571-keV γ ray, with no observation of the 1017-keV γ ray. Considering two following in-beam studies, Ref. [39] also reported the 446-keV γ ray, but Ref. [40] did not observe it; moreover in both references the 1017-keV γ ray was not seen. In the present β -decay experiment and also in those of Ref. [38] an intense γ ray was seen at 1017 keV, while the 446-keV γ ray was not observed. Hence we think that the level we see at 1017 keV is not the same level as in Ref. [42] and therefore it is a

TABLE V. Summary of the results for the β^+ decay of 62 Ge. The first two columns show the γ -ray energies E_{γ} and their intensities I_{γ} (normalized to 100 decays). The last four columns give the level excitation energies E_X in 62 Ga, β feedings I_{β} , Fermi B(F) and Gamow-Teller B(GT) transition strengths to the 62 Ga levels. The B(GT) values are calculated by imposing B(F)=2.0 (see the discussion in the text).

E_{γ} (keV)	I_{γ} (%)	E_X (keV)	I_{β} (%)	B(F)	B(GT)
3594.7(5)	0.6(1)	3594.7(5)	0.6(1)		0.07(1)
3339.6(5)	0.30(6)	3339.6(5)	0.30(6)		0.030(7)
2966.8(5)	0.35(6)	2966.8(5)	0.35(6)		0.028(5)
2641.8(5)	0.4(1)	2641.8(5)	0.4(1)		0.029(7)
2164.1(4)	2.6(2)	2164.1(4)	2.6(2)		0.13(1)
2010.9(4)	0.96(8)	2010.9(4)	0.96(8)		0.045(5)
1899.3(4)	0.58(6)	1899.3(4)	0.58(6)		0.025(3)
1359.7(2)	0.70(5)	1359.7(2)	0.70(5)		0.022(2)
1117.4(2)	0.41(4)	1117.4(2)	0.41(4)		0.011(2)
1017.1(1)	2.6(1)	1017.1(1)	2.6(1)		0.067(6)
978.3(1)	1.8(1)	978.3(1)	1.8(1)		0.047(4)
571.3(1)	3.4(1)	571.3(1)	3.4(1)		0.068(6)
		g.s. ^a	85.3(3) ^b	2.0	

aIAS.

 1^+ state, like the other levels populated in the decay. This is further supported by the non-observation of weak γ branches from the upper excited 1^+ levels to the 1017-keV level which, if it were a 2^+ level, would not be affected by the Warburton and Weneser quasirule.

Since there is no appreciable β -delayed proton emission from ⁶²Ge, Eq. (2) is simplified and the intensity I_{γ}^{i} of each β -delayed γ ray can be calculated as follows:

$$I_{\gamma}^{i} = \frac{N_{\gamma}^{i}}{\epsilon_{\gamma}^{i} N_{\beta}},\tag{3}$$

with the advantage of not depending on ϵ_{decay} . Here N_{β} represents the total number of β events correlated with ⁶²Ge and is obtained from the half-life fit in Fig. 9.

Table V summarizes our results for 62 Ge. The γ -ray energies E_{γ} and intensities I_{γ} are reported in the first two columns of the table. Assuming that all the populated 1^+ states decay directly to the 0^+ ground state of 62 Ga (as discussed above), the excitation energies E_X of the levels in 62 Ga are deduced directly from the corresponding E_{γ} . Similarly, the β feeding I_{β} to each level is inferred from the respective I_{γ} . The third and fourth columns of Table V give E_X and I_{β} , respectively. Finally, the last two columns give the B(F) and B(GT) transition strengths, where the B(GT) values are obtained by setting B(F) = 2.0 (see the discussion below).

The Fermi transition connects the $^{62}\text{Ge}_{\text{g.s.}}$ to its IAS, the $^{62}\text{Ga}_{\text{g.s.}}$. The IAS β feeding (ground state–to–ground state) is calculated by subtracting all the indirect γ feeding to the $^{62}\text{Ga}_{\text{g.s.}}$, $I_{\beta}^{\text{IAS}} = (100 - \Sigma_i I_{\gamma}^i) = 85.3(3)\%$. Then, in order to calculate the β -decay strengths, one needs to know the mass excesses of the nuclear states involved in the transition. The mass excess of the $^{62}\text{Ga}_{\text{g.s.}}$ is known experimentally and has

^bThe ground state–to–ground state feeding is $I_{\beta}^{\text{IAS}} = (100 - \Sigma_i I_{\gamma}^i)$.

a value of -51986.9(6) keV [32]. However the ground-state mass of ⁶²Ge is not known. Since the A = 62, T = 1 isospin multiplet is a triplet (62Ge, 62Ga, and 62Zn), we cannot use the IMME [33,34] to determine the ⁶²Ge_{g.s.} mass, hence we have to rely on systematics. The most recent AME gives a value of $-41740^{\#}(140^{\#})$ keV [32], which gives a Fermi strength of 1.5(1), i.e., much smaller than the expected B(F) = |N - Z|2. If we make the Fermi strength equal to 2, then we obtain a mass excess of -42 258 keV for the ⁶²Ge_{g.s.} and, assuming for this mass a conservative uncertainty of 140 keV as in the mass evaluation systematics, we obtain the value Q_{β} = 9730(140[#]) keV shown in Fig. 12. The 2003 AME value of $-42240^{\#}(140^{\#})$ keV [35] is in agreement with our deduced mass. In addition to the cases of ⁶²Ge and ⁶⁰Ge (Table III), similar issues with the AMEs subsequent to the 2003 AME have been already reported for proton-rich nuclei in this region of the mass chart [3,4,36]. More mass measurements are required in this region to constrain the future AME.

From our deduced ground-state mass excesses for 60 Ge and 62 Ge we can extract the two-neutron separation energy in 62 Ge, $S_{2n} = 30723(142)$ keV. It should be noted that the values from both the 2016 (30970 $^{\#}$ (330 $^{\#}$) keV [30]) and 2003 (30620 $^{\#}$ (270 $^{\#}$) keV [41]) AME systematics agree well with our value.

The B(GT) values shown in Table V agree with those observed in Ref. [38], confirming the absence of enhanced low-lying Gamow-Teller strength in 62 Ga and, hence, the negligible role of the T=0 proton-neutron pairing condensate in A=62 [38].

VI. CONCLUSIONS

We have studied the β decay of the very neutron deficient 60 Ge and 62 Ge isotopes. We measured half-lives of $T_{1/2} = 25.0(3)$ ms for 60 Ge, 69.4(2) ms for the daughter 60 Ga and 73.5(1) ms for 62 Ge. In all cases we have improved the precision in comparison with values in the literature [9,11,31,38].

The decay of ⁶⁰Ge proceeds mainly by β -delayed proton emission populating ⁵⁹Zn with a total branching ratio of 67(3)%. Through this process we have observed for the first time the $1/2^-$ first excited state in ⁵⁹Zn and determined its excitation energy as 463.3(1) keV. The β -delayed γ emission is also seen in competition, populating energy levels previ-

ously unknown in the drip-line 60 Ga nucleus. Three new γ lines (837, 1332, and 1775 keV) have been observed and the level scheme deduced.

The decay of 62 Ge proceeds mostly by direct population of the 62 Ga ground state with a ground state–to–ground state feeding of 85.3(3)%. A total of 12 γ lines have been observed, eight of them for the first time. The Warburton and Weneser quasirule [14,15], already observed in previous $T_z = -1$ cases [12], has again been confirmed here. Moreover, even if there are a few differences in the observed γ rays with respect to Ref. [38], our study confirms that there is no evidence of enhanced low-lying Gamow-Teller strength in 62 Ga due to isoscalar proton-neutron pairing.

Finally, the ground-state mass excesses of 60 Ge, 60 Ga, and 62 Ge have been deduced and the absolute β -decay transition strengths have been determined.

ACKNOWLEDGMENTS

This work was supported by the Spanish MICINN Grants No. FPA2014-52823-C2-1-P and No. FPA2017-83946-C2-1-P (MCIU/AEI/FEDER); Centro Excelencia Severo Ochoa del IFIC SEV-2014-0398; Junta para la Ampliación de Estudios Programme (CSIC JAE-Doc contract) cofinanced by FSE; MCIU Grant No. IJCI-2014-19172. We acknowledge the support of the Generalitat Valenciana Grant No. PROMETEO/2019/007. We acknowledge the support of STFC (UK) Grant No. P005314/1 and ANID/CONICYT FONDECYT Regular 1171467. R.B.C. acknowledges support from the Max-Planck Partner group. We acknowledge the support from NKFIH (NN128072), and by the ÚNKP-20-5-DE-2 New National Excellence Program of the Ministry of Human Capacities of Hungary. G.G.K. acknowledges support from the János Bolyai research fellowship of the Hungarian Academy of Sciences. This work was carried out at the RIBF operated by RIKEN Nishina Center, RIKEN and CNS, University of Tokyo. This work was partially supported by KAKENHI (Grant No. 25247045). The authors acknowledge the EUROBALL Owners Committee for the loan of the germanium detectors and the PreSpec Collaboration for the readout electronics of the cluster detectors.

^[1] B. Blank and M. J. G. Borge, Prog. Part. Nucl. Phys. 60, 403 (2008).

^[2] M. Pfützner, M. Karny, L. V. Grigorenko, and K. Riisager, Rev. Mod. Phys. 84, 567 (2012).

^[3] S. E. A. Orrigo, B. Rubio, Y. Fujita, B. Blank, W. Gelletly, J. Agramunt, A. Algora, P. Ascher, B. Bilgier, L. Cáceres *et al.*, Phys. Rev. Lett. 112, 222501 (2014).

^[4] S. E. A. Orrigo, B. Rubio, Y. Fujita, W. Gelletly, J. Agramunt, A. Algora, P. Ascher, B. Bilgier, B. Blank, L. Cáceres *et al.*, Phys. Rev. C 93, 044336 (2016).

^[5] H. Schatz, A. Aprahamian, J. Görres, M. Wiescher, T. Rauscher, J. F. Rembges, F.-K. Thielemann, B. Pfeiffer, P. Möller, K.-L. Kratz *et al.*, Phys. Rep. **294**, 167 (1998).

^[6] J. L. Fisker, H. Schatz, and F.-K. Thielemann, Astrophys. J. Suppl. Ser. 174, 261 (2008).

^[7] A. Parikh, J. José, G. Sala, and C. Iliadis, Prog. Part. Nucl. Phys. 69, 225 (2013).

^[8] D. K. Galloway, A. Cumming, E. Kuulkers, L. Bildsten, D. Chakrabarty, and R. E. Rothschild, Astrophys. J. 601, 466 (2004).

^[9] A. A. Ciemny, W. Dominik, T. Ginter, R. Grzywacz, Z. Janas, M. Kuich, C. Mazzocchi, K. Miernik, M. Pfützner, M. Pomorski et al., Eur. Phys. J. A 52, 89 (2016).

^[10] S. E. A. Orrigo, B. Rubio, W. Gelletly, B. Blank, Y. Fujita, J. Giovinazzo, J. Agramunt, A. Algora, P. Ascher, B. Bilgier et al., Phys. Rev. C 94, 044315 (2016).

- [11] L. Kucuk, S. E. A. Orrigo, A. Montaner-Pizá, B. Rubio, Y. Fujita, W. Gelletly, B. Blank, Y. Oktem, T. Adachi, A. Algora et al., Eur. Phys. J. A 53, 134 (2017).
- [12] F. Molina, B. Rubio, Y. Fujita, W. Gelletly, J. Agramunt, A. Algora, J. Benlliure, P. Boutachkov, L. Cáceres, R. B. Cakirli et al., Phys. Rev. C 91, 014301 (2015).
- [13] Y. Fujita, B. Rubio, and W. Gelletly, Prog. Part. Nucl. Phys. 66, 549 (2011).
- [14] G. Morpurgo, Phys. Rev. 110, 721 (1958).
- [15] D. H. Wilkinson, *Isospin in Nuclear Physics* (North-Holland Pub. Co., Amsterdam, 1969).
- [16] T. Kubo, Nucl. Instrum. Methods B 204, 97 (2003).
- [17] T. Kubo, D. Kameda, H. Suzuki, N. Fukuda, H. Takeda, Y. Yanagisawa, M. Ohtake, K. Kusaka, K. Yoshida, N. Inabe et al., Prog. Theor. Exp. Phys. 2012, 03C003 (2012).
- [18] N. Fukuda, T. Kubo, T. Ohnishi, N. Inabe, H. Takeda, D. Kameda, and H. Suzuki, Nucl. Instrum. Methods B 317, 323 (2013).
- [19] B. Blank, T. Goigoux, P. Ascher, M. Gerbaux, J. Giovinazzo, S. Grévy, T. Kurtukian Nieto, C. Magron, J. Agramunt, A. Algora et al., Phys. Rev. C 93, 061301 (2016).
- [20] T. Goigoux, P. Ascher, B. Blank, M. Gerbaux, J. Giovinazzo, S. Grévy, T. Kurtukian Nieto, C. Magron, P. Doornenbal, G. G. Kiss *et al.*, Phys. Rev. Lett. **117**, 162501 (2016).
- [21] S. Nishimura, Prog. Theor. Exp. Phys. 2012, 03C006 (2012).
- [22] P.-A. Söderström, S. Nishimura, P. Doornenbal, G. Lorusso, T. Sumikama, H. Watanabe, Z. Xu, H. Baba, F. Browne, S. Go et al., Nucl. Instrum. Methods B 317, 649 (2013).
- [23] J. Taprogge, A. Jungclaus, H. Grawe, S. Nishimura, P. Doornenbal, G. Lorusso, G. S. Simpson, P.-A. Söderström, T. Sumikama, Z. Y. Xu et al., Phys. Rev. C 91, 054324 (2015).
- [24] A. I. Morales, A. Algora, B. Rubio, K. Kaneko, S. Nishimura, P. Aguilera, S. E. A. Orrigo, F. Molina, G. de Angelis, F. Recchia et al., Phys. Rev. C 95, 064327 (2017).
- [25] H. Bateman, Proc. Cambridge Philos. Soc. 15, 423 (1910).
- [26] C. Dossat, N. Adimi, F. Aksouh, F. Becker, A. Bey, B. Blank, C. Borcea, R. Borcea, A. Boston, M. Caamano *et al.*, Nucl. Phys. A 792, 18 (2007).
- [27] J. Honkanen, M. Kortelahti, K. Eskola, and K. Vierinen, Nucl. Phys. A 366, 109 (1981).

- [28] Y. Arai, E. Tanaka, H. Miyatake, M. Yoshii, T. Ishimatsu, T. Shinozuka, and M. Fujioka, Nucl. Phys. A 420, 193 (1984).
- [29] M. J. López Jiménez, B. Blank, M. Chartier, S. Czajkowski, P. Dessagne, G. de France, J. Giovinazzo, D. Karamanis, M. Lewitowicz, V. Maslov *et al.*, Phys. Rev. C 66, 025803 (2002).
- [30] M. Wang, G. Audi, F. G. Kondev, W. Huang, S. Naimi, and X. Xu, Chin. Phys. C 41, 030003 (2017).
- [31] C. Mazzocchi, Z. Janas, J. Döring, M. Axiotis, L. Batist, R. Borcea, D. Ott, E. Caurier, G. de Angelis, E. Farnea *et al.*, Eur. Phys. J. A 12, 269 (2001).
- [32] G. Audi, F. G. Kondev, M. Wang, W. Huang, and S. Naimi, Chin. Phys. C 41, 030001 (2017).
- [33] E. P. Wigner, in Proceedings of the Robert A. Welch Foundation Conference on Chemical Research, Houston, Texas, edited by A. Milligan (Robert A. Welch Foundation, Houston, Texas, 1957), p. 67.
- [34] W. Benenson and E. Kashy, Rev. Mod. Phys. **51**, 527 (1979).
- [35] G. Audi, O. Bersillon, J. Blachot, and A. Wapstra, Nucl. Phys. A 729, 3 (2003).
- [36] M. D. Santo, Z. Meisel, D. Bazin, A. Becerril, B. A. Brown, H. Crawford, R. Cyburt, S. George, G. F. Grinyer, G. Lorusso et al., Phys. Lett. B 738, 453 (2014).
- [37] E. Browne and J. K. Tuli, Nucl. Data Sheets 114, 1849 (2013).
- [38] E. Grodner, A. Gadea, P. Sarriguren, S. M. Lenzi, J. Grebosz, J. J. Valiente-Dobón, A. Algora, M. Górska, P. H. Regan, D. Rudolph *et al.*, Phys. Rev. Lett. **113**, 092501 (2014).
- [39] H. M. David, P. J. Woods, G. Lotay, D. Seweryniak, M. Albers, M. Alcorta, M. P. Carpenter, C. J. Chiara, T. Davinson, D. T. Doherty *et al.*, Phys. Lett. B **726**, 665 (2013).
- [40] T. W. Henry, M. A. Bentley, R. M. Clark, P. J. Davies, V. M. Bader, T. Baugher, D. Bazin, C. W. Beausang, J. S. Berryman, A. M. Bruce *et al.*, Phys. Rev. C 92, 024315 (2015).
- [41] G. Audi, A. Wapstra, and C. Thibault, Nucl. Phys. A 729, 337 (2003).
- [42] D. Rudolph, C. Andreoiu, J. Ekman, C. Fahlander, M. N. Mineva, S. M. Lenzi, E. Farnea, C. Rossi-Alvarez, C. A. Ur, M. Axiotis *et al.*, Phys. Rev. C 69, 034309 (2004).

Correction: The omission of a support statement in the Acknowledgments has been fixed.