Statistical properties of the well deformed ^{153,155}Sm nuclei and the scissors resonance

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(Received 14 November 2020; accepted 23 December 2020; published 12 January 2021)

The nuclear level densities (NLDs) and the γ -ray strength functions (γ SFs) of 153,155 Sm have been extracted from (d, $p\gamma$) coincidences using the Oslo method. The experimental NLD of 153 Sm is higher than the NLD of 155 Sm, in accordance with microscopic calculations. The γ SFs of 153,155 Sm are in fair agreement with QRPA calculations based on the D1M Gogny interaction. An enhancement is observed in the γ SF for both 153,155 Sm nuclei around 3 MeV in excitation energy and is attributed to the M1 scissors resonance (SR). Their integrated strengths were found to be in the range 1.3–2.1 and 4.4–6.4 μ_N^2 for 153 Sm and 155 Sm, respectively. The strength of the SR for 155 Sm is comparable to those for deformed even-even Sm isotopes from nuclear resonance fluorescence measurements, while that of 153 Sm is lower than expected.

DOI: 10.1103/PhysRevC.103.014309

I. INTRODUCTION

The stable samarium isotopic chain provides an excellent opportunity to systematically investigate the evolution of nuclear structure effects, from the semimagic and near spherical ¹⁴⁴Sm to the highly deformed ¹⁵⁴Sm isotope. As the nuclear shape changes, statistical quantities such as the nuclear level density (NLD) and γ -ray strength function (γ SF) are expected to change and provide evolutionary information across the isotopic chain. Furthermore, the behavior of resonance modes, such as the M1 scissors resonance (SR), E1 pygmy dipole resonance (PDR), and the low-energy enhancement (LEE), can be tracked.

The strength of the SR is sensitive to the ground state deformation [1–4]. The SR was first predicted in 1978 by Lo Iudice and Palumbo [5] before it was observed experimentally a few years later [6]. Even-even nuclei were initially considered to be the best experimental candidates to exhibit strong SR modes. However, it soon became apparent that this mode should also present in odd-even and odd-odd systems, although its intensity may be fragmented significantly, making

it more difficult to detect [7,8]. Since then, many heavy deformed even-even and odd-mass rare-earth nuclei have been systematically investigated. So far, the SR mode has been experimentally observed in vibrational and rotational [8], as well as in γ -soft nuclei [9,10] and has also been observed in the actinide region [11–17], as well as in the rare-earth mass region [4,18–27].

A range of different experimental techniques have been used to investigate the low-energy SR. These include ground state absorption experiments such as inelastic electron scattering [6], nuclear resonance fluorescence (NRF) [28], and average resonance capture (ARC) [29], as well as γ -decay experiments such as radiative neutron capture [22,25] and the Oslo method [30,31], which extracts information from ion scattering or transfer reactions. In general, γ -decay experiments have yielded larger SR strengths than the ground state absorption experiments. Summaries of experimental techniques to measure γ SFs can be found in Goriely *et al.* [32].

Recent measurements of the γ SF in the actinides [12,15,16] have uncovered that the SR exhibits a pronounced double-hump structure, seemingly independent of whether the nucleus has an even or odd number of neutrons. The splitting has also been observed, albeit weaker, in the transitional nucleus ¹⁸¹Ta [33,34]. It is suggested that the splitting of the SR may be due to the isovector spin-scissors mode [35], or due to triaxiality [36]. The splitting in the SR has not been reported for any of the rare-earth nuclei studied with the NRF

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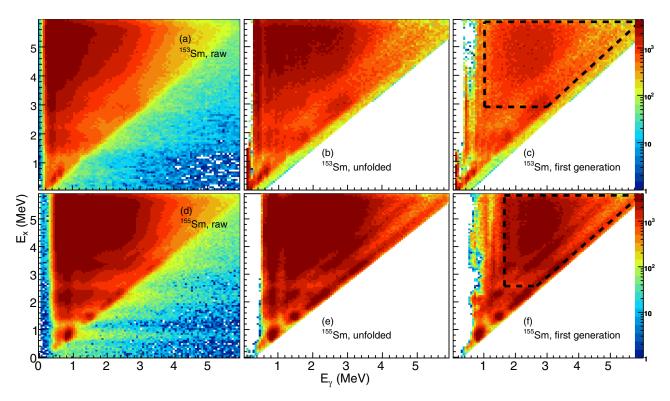


FIG. 1. The raw [panels (a), (d)], unfolded [panels (b), (e)], and first generation (primary) γ -ray matrix [panels (c), (f)] of 153,155 Sm. The dashed lines indicate the energy regions from which the NLDs and γ SFs were extracted.

technique [37], with the Oslo method [18–21,23,24,26,27], or other techniques [4,22,25].

The γ SFs of the isotopes ^{148,149}Sm were measured, and analyzed with the Oslo method, almost two decades ago, and a weak structure was identified to possibly be due to the SR, called a pygmy at the time [20]. With data already available on these two weakly deformed isotopes, together with the recent measurements of ^{151,153}Sm [26] (also analyzed with the Oslo method), it is interesting to extend the investigation towards more deformed Sm nuclei where the SR is expected to be more prominent.

In this paper, the measurement of the NLDs and γ SFs for the odd-even 153,155 Sm extracted with the Oslo method from the 152,154 Sm(d, $p\gamma$) reactions are reported. The integrated strengths of the SR in both isotopes are extracted and compared to previous experimental data. In Sec. II the experimental setup is presented and Sec. III provides a brief overview of the Oslo method and the normalization procedures used, as well as the measured NLDs and γ SFs. Section IV investigates the presence of the SR in Sm isotopes and its integrated strength. A discussion follows in Sec. V and a brief summary in Sec. VI.

II. EXPERIMENTAL SETUP

Two experiments were performed at the Oslo Cyclotron Laboratory (OCL) at the University of Oslo using self-supporting ¹⁵²Sm (enriched to 98.3%) and ¹⁵⁴Sm (enriched to 98.7%) targets with thicknesses of 2.9 and 3.2 mg/cm², respectively. Deuteron beams of 13.5 and 13 MeV were used

to populate excited states in 153,155 Sm. The SiRi particle telescope [38] and CACTUS scintillator [39] arrays were used to detect charged particles and γ rays in coincidence.

The ΔE -E SiRi particle telescope consisted of eight 130 μ m thin, segmented silicon ΔE detectors and eight 1550 μ m thick E silicon detectors. These were placed at backward angles to reduce detection of elastically scattered events and covered a polar angular range of $\theta_{lab}=126^{\circ}-140^{\circ}$ with respect to the beam direction. The energy resolution, as determined from the elastic peaks, is \sim 130 keV. The CACTUS array consisted of 26 and 24 NaI(Tl) detectors for the 153 Sm and 155 Sm measurements, respectively. The 5 in. \times 5 in. crystals were positioned 22 cm from the target, covering solid angles of 17% and 15% of 4π sr, respectively. CACTUS has an energy resolution of 7% FWHM for a 1.332 MeV γ -ray transition.

The E detectors provided a start signal and the NaI(Tl) detectors a stop signal for the time-to-digital converters, enabling event-by-event sorting for the γ -particle coincidence data. Calibration of the SiRi and CACTUS detectors was achieved using distinct γ -ray transitions of ²⁹Si obtained from ²⁸Si(d, $p\gamma$) calibration runs which provided well-resolved particle and γ -ray peaks. During offline analysis, charged-particle– γ coincidence events were extracted within a prompt time gate of 20 ns. Equivalently wide time gates were used to remove the majority of randomly correlated events from the prompt particle- γ events. The excitation energy (E_x) versus γ -ray energy (E_γ) matrices were constructed from the particle- γ coincidence events and are shown in Figs. 1(a) and 1(d). The Oslo method [30,31] is applied on these matrices

TABLE I. Parameters used for extraction and normalization of $\rho(E_x)$ and $\mathcal{T}(E_\gamma)$ in ^{153,155}Sm. The temperature parameter T_{CT} is used for the CT model extrapolation to the B_n .

Isotope	J_{π}	B_n (MeV)	$a $ (MeV $^{-1}$)	E1 (MeV)	T _{CT} (MeV)	$\sigma(B_n)_{ m RMI}$	D ₀ (eV)	$\rho_{\text{RMI}}(B_n)$ (10^6 MeV^{-1})	$ \rho_{\text{HFB}}(B_n) $ (10 ⁶ MeV ⁻¹)	$\langle \Gamma_{\gamma}(B_n) \rangle$ (meV)
¹⁵³ Sm	3/2 ⁺	5.868	18.5	-0.66 -0.56	0.57	6.0(6) ^a	41.0(28) ^b	1.77(36) ^a	2.52(49)	60.0°(37) ^d (110) ^e
¹⁵⁵ Sm	3/2 ⁻	5.807	18.0		0.55	6.1(6) ^a	112(15) ^b	0.66(16) ^a	0.80(17)	74°(11) ^d (13) ^e

^aCalculated with the rigid moment of inertia formula of von Egidy and Bucurescu [52].

to extract simultaneously the NLDs and the γ SFs up to the neutron-binding energies (B_n) through several iterative methods, discussed in the next section. The analysis in this work was performed using the Oslo method software, version 1.1.2 [40].

III. DATA ANALYSIS

A. The Oslo method

The γ -ray spectra, extracted for each E_x bin, were unfolded with the iterative procedure of Ref. [41] and then corrected for the known NaI(Tl) response functions remeasured in 2012 [42], in order to obtain the full-energy γ -ray spectra. At this point, the first-generation γ -ray method [43] is used to extract the primary γ rays from the decay cascades in each E_x bin. The raw matrices [(a) and (d)], the unfolded matrices [(b) and (e)], and the resulting first generation γ -ray matrices [(c) and (f)], $P(E_\gamma, E_x)$, are summarized in Fig. 1. The diagonals where $E_\gamma = E_x$ represent all direct decays to the ground state.

For 153 Sm, the region used for extraction of the NLD and γ SF was from $E_{\gamma}=0.96$ MeV to and including the $E_{x}=E_{\gamma}$ diagonal, and from $E_{x}=2.88$ MeV up to $E_{x}=5.69$ MeV. For 155 Sm, the limits were from $E_{\gamma}=1.65$ MeV to and including the $E_{x}=E_{\gamma}$ diagonal, and from $E_{x}=2.49$ MeV up to $E_{x}=5.73$ MeV. They were chosen to exclude regions characterized by discrete transitions at low excitation energies. The regions in Figs. 1(c) and 1(f) that correspond to $E_{\gamma}<1$ MeV have low statistics due to oversubtraction of discrete and strong γ -ray transitions. This energy region is therefore also excluded from further analysis.

The NLDs and γ SFs of 153,155 Sm were extracted simultaneously from the $P(E_{\gamma}, E_{x})$ matrix, using the ansatz [44,45]

$$P(E_x, E_y) \propto \rho(E_x - E_y) \mathcal{T}(E_y),$$
 (1)

where $\rho(E_x-E_\gamma)$ is the level density at the final E_x to which the nucleus decays. The parameter $\mathcal{T}(E_\gamma)$ is the γ -ray transmission coefficient, and, assuming the generalized Brink-Axel Hypothesis [46,47], it is only dependent on the γ -ray energy. It is also assumed in Eq. (1) that the γ -decay pattern from any initial excitation energy is independent of whether the nucleus was populated into this excitation energy directly from

a nuclear reaction or by γ -ray decays from higher-lying states [48,49]. A χ^2 minimization is performed, between the experimental $P(E_{\gamma}, E_x)$ and a theoretical $P_{theo}(E_{\gamma}, E_x)$ in which $\rho(E_x - E_{\gamma})$ and $\mathcal{T}(E_{\gamma})$ are treated as free parameters [30]:

$$\tilde{\rho}(E_x - E_\gamma) = A\rho(E_x - E_\gamma) \exp[\alpha(E_x - E_\gamma)],$$

$$\tilde{\mathcal{T}}(E_\gamma) = B\mathcal{T}(E_\gamma) \exp(\alpha E_\gamma).$$
(2)

The transformation parameters α , A, and B correspond to physical solutions and are deduced from external experimental data to get the solution to Eq. (1). At this point, the features in the NLDs and γ -ray transmission coefficients are fixed, except for the slopes and absolute values. Note that the resulting ρ and $\mathcal T$ functions do not depend on the initial parameters used in the iterative procedure.

B. Normalization of the NLDs

The extracted NLD is normalized to the known experimental discrete states [50] at $E_x \lesssim 1$ MeV and extrapolated to the NLD, $\rho(B_n)$, at the neutron-binding energy (B_n) , determining its slope and absolute value. The level density $\rho(B_n)$ is determined from the average s-wave neutron-resonance spacing D_0 [51], using Eq. (28) of Ref. [30]. The parameters used for the normalization are listed in Table I. Due to the unavailability of experimental data on the spin and parity J^{π} distribution at B_n , the rigid moment of inertia formula (RMI) which assumes equiparity [52], and the Hartree-Fock-Bogoliubov plus combinatorial (HFB+comb) [53] models were utilized to model the distributions. The RMI and HFB+comb spin distributions at B_n are shown in Fig. 2 for 153 Sm.

The HFB+comb model is a global microscopic approach to calculate the energy- and J^{π} -dependent NLD [53]. The HFB+comb model can be renormalized to match the known experimental discrete states and the average s-wave neutron-resonance spacing D_0 , as detailed in Ref. [53].

Figures 3(a) and 3(b) present the extracted NLDs for 153,155 Sm. At low E_x the NLDs follow closely the experimental discrete states [50]. The NLD bands represent the standard deviation of the level densities as a result of the statistical errors stemming from the Oslo method [30,31], as well as systematic errors. The systematic errors take into account the variations of slope and absolute value due to normalizing

^bTaken from Ref. [57].

^cWeighted average, \bar{x} , calculated from the $x_i \pm \sigma_i$ resonances listed in Ref. [57], using $\bar{x} = \sum_{i=1}^{N} w_i x_i / \sum_{i=1}^{N} w_i$, where $w_i = 1/\sigma_i^2$.

^dUncertainty propagation calculated using $\sigma_{\bar{x}} = \sqrt{1/\sum_{i=1}^{N} w_i}$.

eStandard deviation of the weighted average calculated using $\sigma_{\bar{x},std} = \sqrt{N' \sum_{i=1}^{N} w_i (x_i - \bar{x})^2 / (N' - 1) \sum_{i=1}^{N} w_i}$, where N' is the number of nonzero weights.

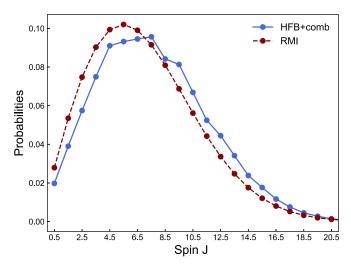


FIG. 2. Spin distributions at B_n estimated from HFB+comb (blue solid line) and RMI (red dashed line) models for 153 Sm.

using the HFB+comb (upper limit) and RMI (lower limit) as well as from varying D_0 within its uncertainties and varying a reduction factor which is used to scale the width of the spin distribution at B_n . The data points in Figs. 3(a) and 3(b) represent the average of the upper and lower limits with the statistical error bars. It is important to distinguish between statistical and systematic errors because the statistical errors are small and they limit the possibility of fluctuations between neighboring points. As can be seen in Fig. 3, the extracted experimental NLD is not available up to B_n due to the exclusion of transitions below the E_{γ} cutoff discussed in Sec. III A. In order to bridge the gap between the highest E_x data point and $\rho(B_n)$, an extrapolation is performed using the Fermi gas (upper limit) [52,54] and constant temperature (lower limit) [55,56] models.

In Fig. 3(c) the NLDs of 153,155 Sm are compared to the NLDs of $(p, d\gamma)^{151,153}$ Sm from Ref. [26] and are shown to be in reasonable agreement. Note that the level density is decreasing significantly from 153 Sm to 155 Sm. This will be discussed in Sec. V.

The $(d, p\gamma)$ reaction may populate a limited spin range due to its low-angular momentum transfer. This again influences the primary γ -ray spectra $P(E_x, E_\gamma)$. A slope correction of the γ SF might therefore be necessary in particular for sub-Coulomb barrier reactions [12–17,58]. To verify whether such a correction is necessary, the J^{π} distribution populated by the $(d, p\gamma)$ reaction has been investigated with the statistical nuclear reaction code TALYS (v1.95) [59] for the deuteron absorption compound reaction formation assuming isotropic emission. The nonelastic breakup, in which the neutron is absorbed by the target in a two-step (direct deuteron breakup + neutron absorption) mechanism, has been investigated as a function of the excitation energy of the residual nucleus following the Green's function transfer formalism of Refs. [60,61]. The results indicate low nonelastic breakup

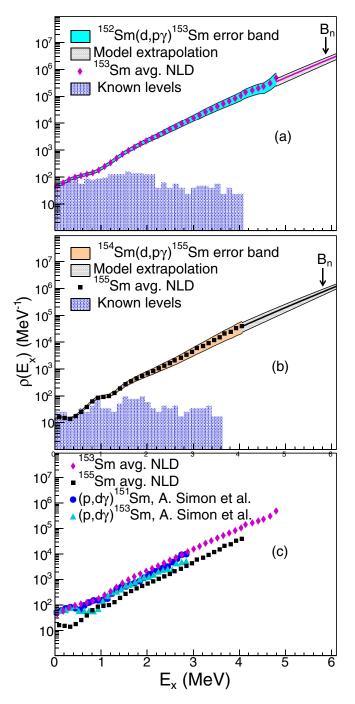


FIG. 3. Experimental NLDs of (a) 153 Sm and (b) 155 Sm from the present $(d, p\gamma)$ experiments, normalized using the RMI and HFB+comb models and extrapolated with the CT and Fermi gas models to match the estimated $\rho(B_n)$. The error bars on the NLD data points represent only statistical uncertainties, whereas the error band represent both statistical uncertainties as well as systematic uncertainties (see text for details). The error bands beyond the last data point (grey shaded) indicate the range of the extrapolation models and the solid line represent the average of the limits. The NLDs of 153,155 Sm are compared to the NLDs of $(p, d\gamma)$ 151,153 Sm from Ref. [26] in panel (c).

cross sections (\approx 2–3 mb/sr MeV) for $\theta_{lab} = 126^{\circ}$ –140°, and the deuteron-fusion proton-evaporation dominates, leading to a broad spin distribution. Therefore, the discrepancy in the populated spin range is considered small and no slope correction was performed.

C. Normalization of the ySFs

The γ -ray transmission coefficient, $\mathcal{T}_{XL}(E_{\gamma})$, of multipolarity L and electromagnetic character X (electric, E, or magnetic, M) is transformed to the total experimental γ SF through the relationship [51]

$$f(E_{\gamma}) \approx f_{E1}(E_{\gamma}) + f_{M1}(E_{\gamma})$$

$$\approx \frac{1}{2\pi E_{\gamma}^{3}} B[\mathcal{T}_{E1}(E_{\gamma}) + \mathcal{T}_{M1}(E_{\gamma})], \tag{3}$$

assuming dominance of dipole transitions for statistical γ -ray decays. This assumption is strongly supported by data; see, e.g., Ref. [63]. The absolute normalization parameter B in Eq. (3) is determined using the experimental average total radiative width $\langle \Gamma_{\gamma} \rangle_{\ell=0}$ at B_n . The corresponding parameters obtained and used for the normalization are summarized in Table I. The value of $\langle \Gamma_{\gamma}(B_n) \rangle_{\ell=0}$ was obtained by calculating the weighted average of the resonance widths listed in Ref. [57].

Using the total average radiative width, the parameter B is determined by [30,64]:

$$\langle \Gamma_{\gamma}(B_n) \rangle = \frac{D_0 B}{2\pi} \int_0^{B_n} dE_{\gamma} \mathcal{T}(E_{\gamma}) \times \rho(B_n - E_{\gamma})$$

$$\times \sum_{I=-1}^{J=1} g(B_n - E_{\gamma}, J_t + J \pm 1/2), \tag{4}$$

where J_t is the target spin (0 for 152,154 Sm) and the function $g(E_x, J)$ is the relative probability of a given spin at excitation energy E_x . In order to calculate the integral in Eq. (4), a log-linear function for the γSF is fitted for γ energies between 0 and $E_{\gamma,l}$ and from $E_{\gamma,h} \to B_n$, where $E_{\gamma,l}$ and $E_{\gamma,h}$ represent γ -ray energies for the lowest and highest data point, respectively. For the RMI, a function approximating the spin distribution is implemented [55,65]:

$$g(\sigma(E_x), J) = \frac{1}{2\sigma(E_x)^2} (2J + 1) \exp\left[\frac{-\left(J + \frac{1}{2}\right)^2}{2\sigma(E_x)^2}\right], \quad (5)$$

with

$$\sigma^{2}(E_{x}) = 0.0146A^{5/3} \frac{1 + \sqrt{1 + 4a(E_{x} - E_{1})}}{2a}, \quad (6)$$

where a is the level density parameter, E_1 is the total backshift parameter, and the spin-cutoff parameter, $\sigma(E_x)$, is a modeled variable related to the width of the distribution. However, in the HFB+comb model the explicit probability for each spin is given, as illustrated in Fig. 2. Therefore, the relevant spin probabilities required for Eq. (4) are directly obtained from the tables [59], while correcting for the excitation-energy shift used to normalize the NLDs.

The extracted experimental γ SFs of ^{153,155}Sm are shown in Fig. 4. As in the case of the NLDs, the γ SF bands include both

statistical and systematic errors. Here, the systematic errors also take into account the uncertainty of the $\langle \Gamma_{\nu} \rangle$ parameter. The averages of the limits and the statistical errors are shown as data points within the error bands. In Fig. 4, the ^{153,155}Sm experimental γ SFs are compared to the experimental giant electric dipole resonance (GEDR) data from 152,154 Sm(γ , n) photonuclear (γ -absorption) data [62], as there is no GEDR data on 153,155 Sm. The γ SFs from this work appear steep in comparison to the available GEDR data; possible explanations are discussed in Sec. V. The present data are further compared to E1 and M1 $(n, \gamma)^{155}$ Sm data measured in average resonance capture (ARC) experiments [29]. The cross sections $\sigma_{\nu}(E_{\nu})$ are transformed to γ SFs using the relation [47]

$$f(E_{\gamma}) = \frac{1}{3\pi^2 \hbar^2 c^2} \frac{\sigma_{\gamma}(E_{\gamma})}{E_{\gamma}},\tag{7}$$

where the factor $1/3\pi^2\hbar^2c^2 = 8.674 \times 10^{-8} \text{ mb}^{-1}\text{MeV}^{-2}$. Finally, in Fig. 4(c), the γ SFs of ¹⁵³Sm and ¹⁵⁵Sm are compared to each other as well as to the γSFs of $^{151,153}Sm$ from Ref. [26]. Several points of interest emerge from this comparison, such as the absolute value difference between the ¹⁵³Sm and ¹⁵⁵Sm from the current experiment. Previous measurements on close-lying nuclei give consistently similar absolute values for the γ SFs. This apparent deviation is discussed in Sec. V. Comparing the results for ¹⁵³Sm from Ref. [26] and from the present work, there are clear differences. Given that both data sets were analyzed using the Oslo method it is important to understand the differing features. First, there seems to be a large discrepancy at γ energies below 2 MeV, where the $(p, d\gamma)^{153}$ Sm data indicate a strong LEE, while the $(d, p\gamma)^{153}$ Sm keeps trending downward. Second, the SR appears to be significantly more pronounced in the $(p, d\gamma)^{153}$ Sm data. This will be explored below after the $B_{\rm SR}(M1)$ strengths are extracted and compared.

IV. THE SCISSORS RESONANCE

Experiments using the Oslo method can only extract the SR built on excited states in the quasicontinuum, whereas NRF measurements can only extract SR built on the ground state. From the γ SF, the integrated reduced transition strength for magnetic dipole transitions, $B_{SR}(M1)$, is obtained by determining the shape of the resonance and numerically integrating over the distribution using the standard Lorentzian function (SLo) [46,47,51], in the energy range relevant to the SR:

$$B_{\rm SR} = \frac{27(\hbar c)^3}{16\pi} \int f_{\rm SR}^{SLo}(E_{\gamma}) dE_{\gamma}, \tag{8}$$

where the factor $27(\hbar c)^3/16\pi = 2.5980 \times 10^8 \ \mu_N^2 \text{MeV}^2$.

Several fitting methods were explored, such as including the GEDR data and fitting a collection of SLo peaks, or making a model of an exponential background plus one SLo peak. However, none of the methods converged successfully and required many of the parameters to be fixed. Therefore a less sophisticated, but transparent method was chosen. To extract the strength of the SR an exponential function was fitted by two points in the experimental γ SFs, to approximate the background in the vicinity of the resonance, as shown

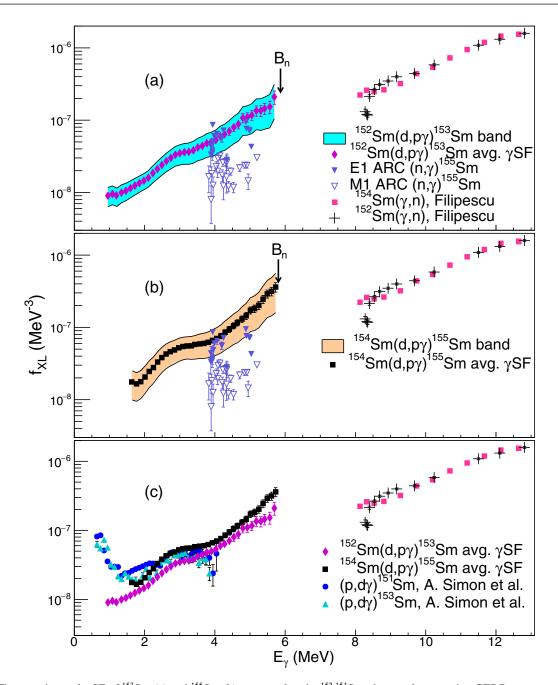


FIG. 4. The experimental γ SF of ¹⁵³Sm (a) and ¹⁵⁵Sm (b) compared to the ^{152,154}Sm photonuclear reaction GEDR measurements of [62], and E1 and M1 (n, γ) ¹⁵⁵Sm data measured in ARC experiments [29]. The γ SFs of ¹⁵³Sm and ¹⁵⁵Sm, which only include the systematic errors, are compared in panel (c). The γ SFs of ^{151,153}Sm from Ref. [26] are included for comparison.

in Fig. 5. It is deemed reasonable that the background is of exponential shape, as all conventional empirical models for the GEDR are of Lorentzian type, where the tail can be approximated by an exponential for a small energy interval. This background was then extracted from the data. An SLo was subsequently fitted to the extracted points, with statistical errors as weights in the fit. A fit was performed on the upper and lower limit of the γ SF, as well as for the average as shown in Fig. 5. The integrated $B_{SR}(M1)$ was found numerically from Eq. (8), and the resulting strengths are listed in Table II. Due to the extraction method, a range of the SR strength is given instead of a recommended value with uncertainties. The

energy centroid is consistent across the different fits and is also given with its small uncertainties in Table II.

V. DISCUSSION

In the current experiments, the NLD and γ SF of the deformed even-odd ¹⁵⁵Sm below B_n , were measured for the first time, as well as that of ¹⁵³Sm in the energy range $4 \leqslant E_{\gamma} \leqslant B_n$ MeV.

It is interesting to note in Fig. 3(c) that, even though the ^{153,155}Sm NLDs have similar slopes, the ¹⁵³Sm level density is higher than that of ¹⁵⁵Sm, which is counterintuitive. As

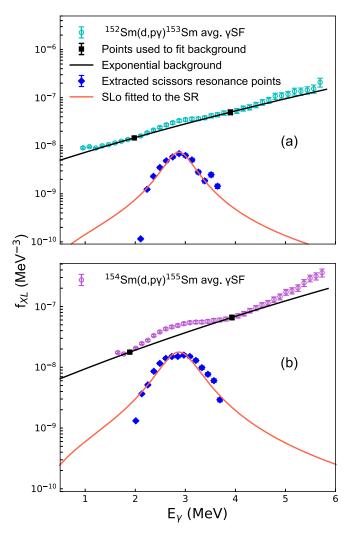


FIG. 5. The extracted SR of ¹⁵³Sm (a) and ¹⁵⁵Sm (b), fitted with the exponential background and SLo (see text for details).

more neutrons are added and deformation increases, it may be expected that there should be more levels in ¹⁵⁵Sm compared to ¹⁵³Sm. The same behavior has also been observed in neodymium isotopes [69,70] where the lighter, spherical or less deformed isotopes exhibit higher NLDs than the more deformed, heavier isotopes. This phenomenon is consistent with the microscopic description of the NLDs obtained within the combinatorial (HFB+combinatorial) [53] and statistical (HFBCS+statistical) [68] models which are compared to the experimental data for the even-odd ^{153,155}Sm isotopes, in Fig. 6.

The same feature of a lower NLD for 155 Sm is seen to be predicted by both microscopic models. Also shown in Fig. 6 are the energy-dependent level density parameters $a_{\rm stat}$ obtained within the microscopic statistical approach [68]. The a parameter is a measure of the single-particle level density at the Fermi surface and is consequently sensitive to shell and pairing effects. The difference between 153 Sm and 155 Sm NLDs is essentially due to stronger shell plus pairing effects in 155 Sm in comparison with 153 Sm, leading to a smaller single-particle level density in 155 Sm at the Fermi energy. This structure effect is also found in the ground-state microscopic energy predicted by most macroscopic-microscopic mass models [66,71] which give a maximum microscopic energy within the Sm neighboring isotopes for A=148-150 and a lower value for 155 Sm compared to 153 Sm.

The 153,155 Sm γ SFs show pronounced strength, which is localized at mean excitation energies of about $\omega_{SR} \approx 3$ MeV, a feature observed in most deformed rare-earth nuclei. The analytical technique used to extract the γ SFs, the Oslo method [30,31], cannot reveal fine structures in the γ spectra. However, the observed resonances at \sim 3 MeV are believed to be due to the M1 SR, which is consistent with other observations in this mass region. High resolution measurements (see Ref. [8] and references therein) have shown that these resonances are due to M1 transitions between high-j orbitals

TABLE II. Experimental integrated $B_{SR}(M1)$ strength for samarium isotopes extracted in the given energy ranges. The quadrupole deformation of the nuclear ground state δ^2 is taken from the FRDM12 database [66]. The quantity ω_{SR} denotes the energy centroid of the SR

Isotope ^A X	Deformation δ^2	Energy range (MeV)	$\omega_{ m SR} \ m (MeV)$	$B_{ m SR}(M1) \ (\mu_N^2)$	Ref.	Technique
¹⁵³ Sm	0.26	Full range ^a	2.89(0)	1.27-2.13	Present expt.	Oslo method
¹⁵³ Sm	0.26	2.0-4.0	2.89(0)	1.09-1.85	Present expt.	Oslo method
¹⁵⁵ Sm	0.27	Full range ^a	2.98(8)	4.40-6.44	Present expt.	Oslo method
¹⁵⁵ Sm	0.27	2.0-4.0	2.98(8)	3.58-5.30	Present expt.	Oslo method
¹⁴⁴ Sm	0.08	2.0-4.0	3.97(4)	0.28(0)	Ziegler et al. [67]	NRF
¹⁴⁸ Sm	0.18	2.0-4.0	3.07(3)	0.51(1)	Ziegler et al. [67]	NRF
¹⁵⁰ Sm	0.21	2.0-4.0	3.13(3)	0.97(10)	Ziegler et al. [67]	NRF
¹⁵¹ Sm	0.22	0.0-5.0	3.00(2)	7.80(340)	Simon <i>et al</i> . [26]	Oslo method
¹⁵² Sm	0.24	2.0-4.0	2.99(3)	2.35(20)	Ziegler et al. [67]	NRF
¹⁵³ Sm	0.26	0.0-5.0	3.00(2)	7.80(200)	Simon <i>et al.</i> [26]	Oslo method
¹⁵⁴ Sm	0.27	2.0-4.0	3.20(3)	2.65(30)	Ziegler et al. [67]	NRF

^aEquation (8) was integrated from 0 to 20 MeV. This equates to an unrestricted range, as including higher energies did not alter the tabulated value to a significant digit.

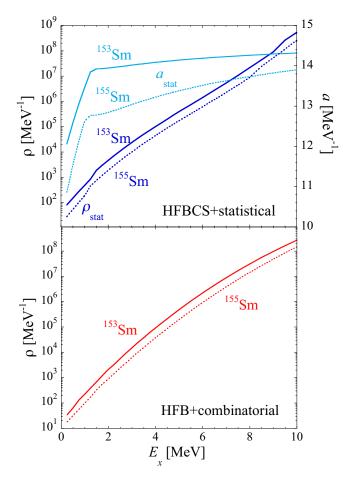


FIG. 6. Comparison between 153 Sm (solid lines) and 155 Sm (dashed lines) NLDs obtained within the combinatorial (red curves) [53] and statistical, ρ_{stat} , (dark blue curves) [68] models. Also shown are the energy-dependent level density parameters a_{stat} (light blue curves) obtained within the microscopic statistical approach [68].

and dominated by $J^{\pi}=1^+$ states when excited directly from a $J^{\pi}=0^+$ ground state [1]. The increase in the integrated SR strength from 153 Sm to 155 Sm is not entirely consistent with previous experimental findings and theoretical descriptions [8]. Considering that the deformation is comparable for the two nuclei the strengths should be comparable as well. A possible explanation might be that the reported calculated deformation for 155 Sm is underestimated, and therefore an experiment to measure both 153,155 Sm deformations is highly desirable.

In contrast to what was observed for the actinides [12,15,16], the SR in the odd-even rare-earth 153,155 Sm γ SF does not exhibit any double-hump structure.

Figure 7 compares the present integrated $B_{\rm SR}(M1)$ with the experimental NRF values and those of Ref. [26] extracted with the Oslo method analysis. Given the comparable deformation parameters in Table II, the $B_{\rm SR}(M1)$ for $(d,p\gamma)^{155}{\rm Sm}$ is in reasonable agreement with the measurements of Ziegler *et al.* [67] on $^{152,154}{\rm Sm}$ when extracted over the same energy region, while the $B_{\rm SR}(M1)$ for $(d,p\gamma)^{153}{\rm Sm}$ is lower than that of the neighboring nuclei. Also the $B_{\rm SR}(M1)$ for $(d,p\gamma)^{155}{\rm Sm}$ is in excellent agreement with measurements of the other rare-earth

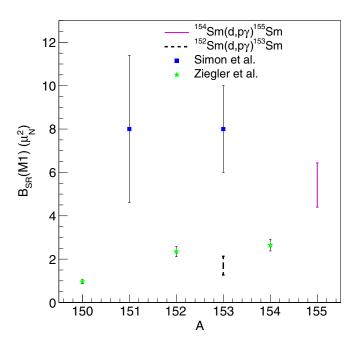


FIG. 7. The experimental low-lying M1 strength, $B_{SR}(M1)$, plotted against mass number A. The present measurements (lines) extracted over a full energy range are compared to 151,153 Sm data from Ref. [26] (blue squares) and 150,152,154 Sm from NRF measurements of Ref. [67] (green stars).

nuclei extracted with the Oslo method [18–21,23,24,26,27]. The results for ^{151,153}Sm [26] (also listed in Table II) are significantly higher than those from both the ^{150,152,154}Sm NRF values and the present measurements.

This follows predictably from the discussion in Sec. III C, where the SR appears much more pronounced in the γ SFs of $(p, d\gamma)^{151,153}$ Sm data. To explain this, and the LEE discrepancy, it is important to highlight the differences in the analyses performed. One difference between how the two data sets were analyzed is the region of extraction of the NLDs and γ SFs from the primary γ -ray matrices, as the $(p, d\gamma)^{153}$ Sm data were extracted for E_{ν} as low as 0.645 MeV and from $E_x = 2.525$ MeV to $E_x = 4.045$ MeV. Depending on the statistical nature of the nucleus at these energies, the resulting y SFs might be inconsistent. There might also be some residual transitions at low γ energies in the primary γ -ray matrices for the $(p, d\gamma)^{153}$ Sm data that lead to the differences in the γ SFs below 2 MeV. Importantly, the approaches for extracting the SR are different, for $(p, d\gamma)^{151,153}$ Sm a fit for the energy region $E_{\gamma} = 0$ –11 MeV is performed. Last, the highest data points for the γ SFs of $(p, d\gamma)^{153}$ Sm have high uncertainties and are located at γ energies which could give the appearance of a large SR and, depending on the reliability of the last data points, the slopes could be in sharp contrast.

In previous measurements, the $B_{SR}(M1)$ has proved to be less fragmented and stronger in even-even nuclei than in even-odd nuclei [8,37]. The strength seen in (γ, γ') experiments for odd-even cases can differ to a great extent. However, the unobserved strength can be estimated from a fluctuation analysis of the data; see Ref. [7]. If this is taken into account, the same accumulated strengths as in the even-even cases and

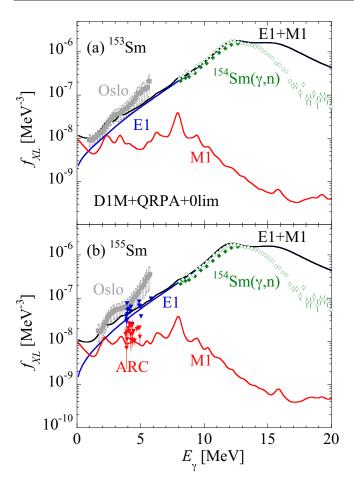


FIG. 8. Comparison between the D1M+QRPA+0lim [32,74] E1 (blue solid lines), M1 (red solid lines), and E1+M1 (black solid lines) with experimental data, (a) for 153 Sm and (b) for 155 Sm. The present Oslo data including systematic uncertainties correspond to gray squares and the ARC data to triangles (blue for E1 and red for M1). To give a fair approximation of the γ SF in the GDR region, the γ SF extracted from photoneutron cross section of the neighboring 154 Sm is shown by solid [62] and open triangles [75].

the same dependence on deformation is found. Different types of experiments and theoretical calculations that extract the $B_{\rm SR}(M1)$ over the same energy region yield similar strengths as shown for Dy isotopes [27]. The present ¹⁵⁵Sm measurement is in reasonable agreement whereas ¹⁵³Sm appears to contradict these findings. When the extraction performed here is limited by the same integration limits in Eq. (8), the ¹⁵³Sm B(M1) value is found to be significantly lower (higher limit of 1.9 μ_N^2) than measurements of the other rare-earth nuclei extracted with the Oslo method [27] and that of the neighboring even-even deformed ¹⁵⁴Sm isotope of Ref. [67].

neighboring even-even deformed 154 Sm isotope of Ref. [67]. A steep increase in the γ SF of 155 Sm above $E_{\gamma} > 4$ MeV might be an indication of a PDR. This excessive strength is also observed in the γ SF of 153 Sm, as shown in Fig. 5. However, it is not possible to determine the electromagnetic nature of the resonance from Oslo method-type experiments, and therefore information from other experiments such as NRF [28] or inelastic proton scattering measurements [72,73] is crucial. The LEE is not seen in the γ SFs of 153,155 Sm for

the energy range under investigation. This may be due to the current experimental conditions, which limit the extraction of useful data below ~ 1 and 1.6 MeV, respectively. This is in contrast to the $(d, p\gamma)^{151,153}$ Sm data shown in Fig. 4.

Finally, we compare in Fig. 8 the D1M+QRPA+0 $\lim E1$, M1 and E1 + M1 γ SF with available experimental data, i.e., the present Oslo data, ARC data known separately for E1 and M1 strengths in 155 Sm, and the γ SF extracted from photoneutron cross section of the neighboring ¹⁵⁴Sm isotope. The D1M+QRPA+0lim model corresponds to axially deformed Gogny-HFB plus quasi-particle random phase approximation (QRPA) predictions obtained with the D1M interaction. The model has been complemented by phenomenological shellmodel-inspired E1 and M1 LEE contributions to describe the deexcitation strength function [32,74]. As seen in Fig. 8, the D1M+QRPA+0lim predictions are in rather good agreement with data, especially for ¹⁵³Sm. In the ¹⁵⁵Sm case, the total calculated γ SF remains somewhat lower than the one found with the Oslo method though the agreement with ARC data is good. For both Sm isotopes, the large γ SF around 5 MeV cannot be explained by the D1M+QRPA+0lim model. Due to the phenomenological inclusion of an M1 LEE, the D1M+QRPA+0lim model can reproduce rather well the lowenergy points found by the Oslo method below and around 2 MeV.

VI. SUMMARY

The NLD and the γ SF of the deformed even-odd 153,155 Sm have been measured with the reaction $(d,p\gamma)$ below B_n at the Oslo Cyclotron Laboratory (OCL) using the Oslo method. A pronounced resonance, the SR, was observed for both nuclei. The SR integrated strengths in 153,155 Sm were determined to be in the ranges 1.3–2.1 and 4.4–6.4 μ_N^2 , respectively. These values are comparable to those of neighboring nuclei for 155 Sm and somewhat smaller in 153 Sm.

The experimental NLD of ¹⁵³Sm is found by the Oslo method to be larger than that of ¹⁵⁵Sm, a counterintuitive pattern that is confirmed by microscopic models and explained by stronger pairing plus shell effects in ¹⁵⁵Sm. QRPA calculations based on the D1M Gogny interaction are also found to predict the ^{153,155}Sm γ SF in fairly good agreement with the Oslo data, though the large strength around 5 MeV is not described by the model.

ACKNOWLEDGMENTS

The authors thank the cyclotron team at the University of Oslo for providing high-quality experimental conditions. This work is based on the research supported in part by the National Research Foundation of South Africa (Grants No. 118846, No. 92600, No. 90741, and No. 92789) and by the IAEA under Research Contract 20454. This work was partially supported by the Fonds de la Recherche Scientifique (FNRS), the Fonds Wetenschappelijk Onderzoek–Vlaanderen (FWO) under the EOS Project No. O022818F and the U.S. Department of Energy under Contract No. DE-AC52-07NA27344. The authors gratefully acknowledge funding from the Research Council of Norway (NFR) project Grant

No. 263030 (A.G., S.S., F.Z., and V.W.I.) and project Grant No. 262952 (G.M.T.), and by the Deutsche Forschungsgemeinschaft (DFG) under Grant No. SFB 1245, project

ID 279384907 (P.v.N.-C.). A.C.L. gratefully acknowledges support by the European Research Council through ERC-STG-2014 under Grant Agreement No. 637686.

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