Detailed low-spin spectroscopy of 65Ni via neutron capture reaction

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An extended investigation of the low-spin structure of the ⁶⁵Ni nucleus was performed at the Institut Laue-Langevin, Grenoble, France, via the neutron capture reaction ⁶⁴Ni(*n*, γ)⁶⁵Ni, using the Fission Product Prompt Gamma-Ray Spectrometer High-Purity Ge array. The level scheme of ⁶⁵Ni was significantly expanded, with two new levels and 87 newly found transitions. Angular correlation analyses were also performed, allowing us to assign spins and parities for a number of states and to determine multipolarity mixing ratios for selected γ transitions. The low-energy part of the experimental level scheme (up to about 1.4 MeV) was compared with Monte Carlo shell-model calculations, which predict spherical shapes for all states, apart from the $9/2^+$ and the second excited 1/2[−] states of oblate deformation.

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I. INTRODUCTION

Neutron capture reactions, referred to below as (n, γ) , populate the nucleus at the neutron separation energy and rather low spin (i.e., the spin of the target nucleus coupled to the 1/2 spin of the neutron). Such system deexcites via γ -ray cascades towards the ground state. Using γ -coincidence techniques, a detailed level scheme of low-spin states, which is complementary to the ones reconstructed via fusion, deep-inelastic, or transfer reactions, can be obtained $[1-3]$. In neutron-rich Ni isotopes, detailed studies of the low-spin structures are particularly relevant, since they may shed light on the shape coexistence phenomenon. This phenomenon has been predicted, at low spin, in these nuclei by various theoretical calculations [\[4–9\]](#page-12-0) and also clearly observed in the even-even Ni isotopes with mass $A = 64, 66, 68,$ and 70 [\[10–15\]](#page-12-0).

In this context, the odd system 65 Ni is of special interest since it sits in between the even-even systems ⁶⁶Ni, where a shape-isomer-like structure (an extreme case of shape coexistence) has been recently observed $[12]$, and ⁶⁴Ni, which also shows coexistence of spherical, oblate, and prolate shapes [\[15\]](#page-12-0). A shape coexistence scenario is, therefore, expected also in 65 Ni, although with a more complex pattern due to its even-odd nature.

Our knowledge on the structure of ⁶⁵Ni comes mainly from (n, y) experiments performed about 40 years ago $[16–19]$, deep-inelastic-scattering experiments from the 1990s [\[20\]](#page-12-0), (d, p) reaction studies [\[21–23\]](#page-12-0), and β -decay studies [\[24\]](#page-12-0). The aim of this work is to perform a complete low-spin spectroscopic study of ⁶⁵Ni, by using a (n, γ) reaction, in order to significantly expand our understanding of its low-spin structure and search for the signatures of shape coexistence.

The experiment was performed at the reactor of Institut Laue-Langevin (ILL), Grenoble, France, using the array Fission Product Prompt Gamma-Ray Spectrometer (FIPPS). The level scheme of 65 Ni was significantly expanded, with two new levels and 87 newly found transitions. In addition, angular correlation studies allowed us to determine the multipolarity of several transitions and firmly assign spin and parity to a number of states.

The low-energy part of the experimental level scheme, up to around 1.4 MeV, was compared with the results from Monte Carlo shell-model (MCSM) calculations, which provide a description of the structure of the states in terms of wave-function compositions and associated intrinsic shapes [\[8,9,25\]](#page-12-0). A dominance of spherical configurations is observed in this energy range, with one possible excitation of oblate

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FIG. 1. Coincidence energy spectra gated on the 310- and 629-keV transitions, respectively, are shown in panels (a) and (b). Newly observed transitions are indicated by red circles, while known ones are indicated by blue triangles. Stars indicate escape peaks (EP), radiation from β^- decay of ⁶⁵Ni (β), random coincidences with ⁶⁵Ni itself (rc), and spurious peaks from Compton events (C), caused by the absence of Compton shields.

nature. The paper is organized as follows. Section Π contains the experimental details and results, including the construction of the level scheme (Sec. II A), the γ -transition intensity determination (Sec. [II B\)](#page-2-0), and the γ -angular correlation analysis (Sec. HC). In Sec. [III,](#page-11-0) a comparison between the experimental level scheme and MCSM predictions is presented. Section [IV](#page-11-0) contains the conclusions.

II. EXPERIMENTAL METHOD AND RESULTS

The experiment was performed at the ILL, using the FIPPS instrument. The FIPPS array consisted of eight High-Purity Ge (HPGe) clovers, each composed of four *n*-type HPGe crystals with a diameter of 50 mm (before shaping) and a length of 80 mm, which are front tapered [\[26\]](#page-12-0). The detectors were arranged symmetrically around the target position perpendicularly to the beam direction.

The ⁶⁵Ni nucleus was populated via a thermal neutron capture reaction on a highly isotopically enriched ⁶⁴Ni target. The neutron beam was delivered from the ILL reactor, collimated to a diameter of about 1.5 cm at target position and with a flux of 7.7×10^7 neutrons/(cm² s). The ⁶⁴Ni target (86.7 mg, enriched at 99.6%) was in powder form and was placed in a double-layered teflon bag at a distance of 9 cm from the front face of the HPGe clovers. The target holder consisted of a cylinder of about 50-mm diameter

and 50-cm length, made of lithium fluoride plastic (enriched in 6Li). This was used as a shielding for scattered neutrons, thus contributing to the suppression of beam-induced background.

The signals were treated with digital electronics and the data were collected in triggerless mode. The events were built using a 300-ns time window, validated by channel multiplicity 2. A total number of 1.3×10^{10} γ γ coincidences were collected and sorted into matrices.

A. Construction of the level scheme

To build the presented level scheme, the $\gamma\gamma$ coincidence technique was applied. A $\gamma\gamma$ matrix was constructed, considering a coincidence window of 300 ns. The energy calibration was performed using the neutron capture reaction ²⁷Al(*n*, γ)²⁸Al, providing reference prompt γ rays in an energy range from tens of keV to \approx 7 MeV.

The calibrated data were sorted into $\gamma\gamma$ matrices and coincidence gates on known transitions were used to obtain the new spectroscopic information. We observed 76 new transitions between known states. In Figs. $1(a)$ and $1(b)$ the spectra in coincidence with the 310-keV (3/2⁻ \rightarrow 5/2⁻) and 629-keV (3/2⁻ \rightarrow 1/2⁻) γ rays are shown. The newly found transitions are labeled in red, while the previously known ones are in blue.

FIG. 2. Examples of coincidence gated spectra for ⁶⁵Ni, obtained by setting gates on the primary γ rays of (a) 3084 keV, (b) 2097 keV, and (c) 1554 keV (see level scheme in Figs. [3–](#page-3-0)[5\)](#page-5-0). Escape peaks and peaks from random coincidences are marked with a star.

Two new excited levels were identified at 3014.43(14) and 4001.49(18) keV. Spectra gated on the 3084- and 2097-keV transitions populating these levels are presented in Figs. $2(a)$ and $2(b)$. They show peaks corresponding to different branches of their decay. In total, 11 new transitions populating or depopulating the two new levels were found. Figure $2(c)$ shows also the spectrum gated on the 1554-keV transition which populates the 4544-keV level.

The 65Ni level scheme established in this work is shown in Figs. [3](#page-3-0)[–5.](#page-5-0) Levels and transitions marked in black were previously known, while the ones in red are identified in the present investigation. Table [I](#page-6-0) summarizes the information about energy levels and γ transitions of ⁶⁵Ni, discussed in this paper.

B. Intensities

In order to determine the relative γ -ray intensities following the (n, γ) reaction, a detection efficiency curve ϵ_{γ} of the FIPPS array was carefully evaluated. Transitions from a ¹⁵²Eu calibration source, together with transitions of known intensities from the neutron capture reactions 48 Ti(*n*, γ)⁴⁹Ti and ²⁷Al(*n*, γ)²⁸Al, were used for this purpose. The efficiency curve was fit with the following phenomenological function:

$$
\epsilon_Y(E) = \exp\left(a_1 + a_2 E + a_3 \frac{\ln E/E_0}{E} + \frac{a_4}{E} + \frac{a_5}{E^2} + \frac{a_6}{E^3}\right)
$$
\n(1)

where $E = E_y/20$ and $E_0 = 1$ keV. The parameters resulting from the fit are $a_1 = -3.84(8)$, $a_2 = -0.0046(4)$ keV⁻¹, $a_3 = 16(2)$ keV, $a_4 = -23(4)$ keV, $a_5 = 43(9)$ keV², and $a_6 = -28(5) \text{ keV}^3$. The uncertainty on the efficiency curve was calculated imposing the condition that about 68% of the data points should be within one sigma from the fit, considering separately the ranges below and above 1 MeV.

Relative intensities were obtained from $\gamma\gamma$ coincidence spectra, by using different gating conditions. For each γ ray depopulating the neutron capture state (primary γ ray), gates were set on all the observed transitions from the level which is fed by the primary transition. It was assumed that they represent the full decaying intensity, owing to the high sensitivity of the FIPPS array. The relative intensity I_{γ} of the considered primary γ ray was obtained using Eq. (2):

$$
I_{\gamma} = N \sum_{i} \frac{A_{i}}{\epsilon_{\gamma} \epsilon_{\gamma_{i}}} \tag{2}
$$

where A_i is the primary transition peak area when gating on the transition *i*, and ϵ_{γ} and ϵ_{γ} are the singles photopeak efficiencies of the primary and gating γ rays. For secondary γ rays, gates were set on all the transitions feeding the level of interest and Eq. (2) was used. In this case ϵ_{γ} and ϵ_{γ} are the singles photopeak efficiencies of the considered secondary transition and the feeding γ rays, respectively. *N* is an overall normalization factor calculated by summing the intensities of $γ$ -γ pairs containing the 310-keV (3/2⁻ \rightarrow 5/2⁻) transition (i.e., the most intense one observed in the $\gamma\gamma$ matrix) and transitions feeding the 310-keV level.

Results of the intensity measurements are listed in Table [I.](#page-6-0)

C. Angular correlations, multipolarity determinations, and spin-parity assignments

Multipolarity mixing ratios of γ transitions and spin and parity assignments for a number of excited states in ⁶⁵Ni were established by employing a $\gamma\gamma$ angular correlation analysis.

Given two γ rays emitted in the same cascade at an angle θ between each other, the angular correlation $W(\theta)$ between

FIG. 3. First part of the proposed level scheme for ⁶⁵Ni. Newly found levels and γ transitions are marked in red. Intensities are reported in italic, above transition energies.

involved [\[28\]](#page-12-0):

them is generally expressed as a superposition of Legendre polynomials *Pn*:

$$
W(\theta) = A_0[1 + A_2 P_2(\cos \theta) + A_4 P_4(\cos \theta)]
$$
 (3)

where A_0 is a normalization coefficient and A_n ($n = 2, 4$) are the products of two coefficients, $A_n(1)$ and $A_n(2)$, depending on the character of the transitions and the spin of the levels

Here q_n indicates attenuation terms, which take into account the finite solid angle of the detectors.

 $A_n = q_n A_n(1) A_n(2)$, with $n = 2, 4$. (4)

The number of angle pairs of the FIPPS array considered in the present paper is 19. The attenuation coefficients q_n for the

6098.28

 $1/2^+$

FIG. 4. Second part of the proposed level scheme for ⁶⁵Ni. Newly found levels and γ transitions are marked in red. Intensities are reported in italic, above transition energies.

FIG. 5. Third part of the proposed level scheme for ⁶⁵Ni. Newly found levels and γ transitions are marked in red. Intensities are reported in italic, above transition energies.

apparatus were obtained considering different pairs of γ rays from a 152Eu source, with known character and multipolarity mixing ratios. The established values are $q_2 = 0.80(10)$ and $q_4 = 0.60(10).$

The multipolarity mixing ratio δ_{γ} for a γ transition is defined as follows [\[28\]](#page-12-0):

$$
\delta = \frac{\langle j_f ||\lambda'||j_i \rangle}{\langle j_f ||\lambda|| j_i \rangle} \tag{5}
$$

E_i (keV)	J_i^{π}	E_f (keV)	J^{π}_c	E_{ν} (keV)	$\mathbf{1}$				
		1920.41(8)	$5/2^+$	4177.8(8)	0.70(5)				
		1417.71(7)	$1/2^{-}$	4680.3(5)	17(2)				
		692.36(5)	$3/2^{-}$	5405.6(3)	46(6)				
		310.44(5)	$3/2^{-}$	5787.7(5)	86(12)				
		63.46(6)	$1/2^{-}$	6034.8(5)					

TABLE I. *(Continued).*

where λ is the lowest possible multipolarity of the photon γ emitted in the transition between the states with initial spin j_i and final spin j_f , and $\lambda' = \lambda + 1$. The mixing ratio value of a certain transition was extracted from the fit of angular correlation experimental data when the mixing ratio of the other transition of the considered pair was known, so that either $A_n(1)$ or $A_n(2)$ in Eq. [\(4\)](#page-3-0) is fixed, for each *n*. A fit of the experimental data was thus performed giving the experimental A_2^{exp} and A_4^{exp} values, which could then be compared to the corresponding theoretical analytical expressions [in the following indicated as $A_2(\delta)$ and $A_4(\delta)$], as a function of δ . The δ value was obtained minimizing the χ^2 function defined as

$$
\chi^{2} = \left(\frac{A_{2}^{\exp} - A_{2}(\delta)}{\Delta A_{2}^{\exp}}\right)^{2} + \left(\frac{A_{4}^{\exp} - A_{4}(\delta)}{\Delta A_{4}^{\exp}}\right)^{2} \tag{6}
$$

where ΔA_2^{exp} and ΔA_4^{exp} are the uncertainty of A_2^{exp} and A_4^{exp} .

We started the analysis by establishing the mixing ratios of the two most intense transitions in the decay scheme, namely, the 310-keV (3/2⁻ → 5/2⁻) and 629-keV (3/2⁻ → 1/2⁻) γ rays. These two transitions were subsequently used to determine the mixing ratios of other lines considered in pairs with them. A few pairs of nonconsecutive transitions were considered. In these cases, to estimate the uncertainty introduced by a possible deorientation caused by the intermediate transitions, the approach introduced in Ref. [\[2\]](#page-12-0) was applied. A deorientation was simulated by increasing the angular correlation attenuation in Eq. [\(4\)](#page-3-0), up to 50% for each cascade, and the corresponding variation of the resulting mixing ratio was adopted as a systematic uncertainty. When multiple pairs could be considered for the investigated transition (i.e., for the primary 3774- and 3952-keV γ rays), δ was taken as an average of the values resulting from the fit of individual pairs. In this case, the uncertainty was calculated taking into account the variation in δ values obtained from the analysis of the different pairs.

The information on angular correlation coefficients A_2^{exp} and A_4^{exp} and mixing ratios δ is summarized in Table [II](#page-9-0) for all the considered γ cascades. In the following paragraphs, we first discuss the angular correlation analysis performed to establish the multipolarity mixing ratios for the 310- and 629 keV transitions. Further, the spin assignments for a number of states are deduced.

1. Multipolarity mixing ratio of the 310- and 629-keV transitions

a. 310 keV, 3/2[−] → 5/2−*.* The 310-keV ground-state transition between $3/2^-$ and $5/2^-$ states is expected to have an $M1 + E2$ character. In order to determine its multipolarity

mixing ratio δ_{310} , the γ -ray cascade 5788-310 keV was considered. The angular correlation plot for this cascade is presented in panel (a) of Fig. [6.](#page-9-0) For the 5788-keV ($1/2^+ \rightarrow$ 3/2−) primary γ -ray transition, a dominant *E*1 character was assumed on the basis of angular momentum selection rules, with a maximum $M2$ mixing contribution of δ_{5788} < 0.05, which is consistent with the work of Ref. [\[1\]](#page-12-0). The δ_{310} value obtained with such assumption (i.e., averaging the solutions considering the lower and upper limits $\delta_{5788} = 0.00$ and 0.05) is 0.191(13).

b. 629 keV, 3/2⁻ → 1/2⁻. For the 629-keV transition, depopulating the 692-keV state, an $M1 + E2$ character was expected. To determine the multipolarity mixing ratio δ_{629} , the cascade 5406-629 keV was considered. Similarly to the previous case, for the 5406-keV primary transition a dominant *E*1 character was assumed, with a maximum *M*2 mixing contribution of δ_{5406} < 0.05. The two mixing ratio solutions for the 629-keV transition are $\delta_{629} = 0.052(11)$ and $-1.95(6)$. To constrain this result, the ≈ 0.7 -ps lifetime of the 692-keV level, provided by a one-neutron transfer reaction 64 Ni(13 C, 12 C)⁶⁵Ni study performed at sub-barrier Coulomb energy at Horia Hulubei National Institute of Physics and Nuclear Engineering (Bucharest, Romania) [\[30\]](#page-12-0), was used. For the mixing ratio $\delta_{629} = -1.95(6)$, the resulting reduced transition probability value would be $B(E2) \approx 480$ Weisskopf units (W.u.), which is unrealistically large. On the other hand, $\delta_{629} = 0.052(11)$ gives $B(E2) \approx 1$ W.u., which is a reasonable value. Having this result, we analyzed the angular correlation between the 629- and 1228-keV γ rays, which yielded a multipolarity mixing ratio for the 1228-keV transition $\delta_{1228} = -0.09(4)$, pointing to its $E1 + M2$ character.

2. 2324-keV level: $J^{\pi} = 3/2^{(-)}$ spin assignment

In order to assign spin and parity to the 2324-keV level, never determined before, the angular correlations between the primary 3774-keV γ ray feeding this level and the 310- and 629-keV lower-lying transitions were studied. The analysis pointed to a positive A_2^{exp} coefficient for both pairs, as shown, for example, for the 3774-629-keV case in panel (b) of Fig. [6.](#page-9-0) Given the spin of the capture state $J_{6098} = 1/2$, the most probable spin assignment for the 2324-keV level is 3/2, because $J_{2324} = 1/2$ would imply an isotropic angular correlation plot, while $J_{2324} = 5/2$ would give negative A_2 coefficients. Subsequently, the angular correlations for both pairs were used to establish the multipolarity mixing ratio of the 3774-keV transition, for which a dipole $+$ quadrupole character was expected. A solution common to both cascades is $\delta_{3774} =$ −0.1(2). We tentatively propose a negative parity for the

TABLE II. Experimental A_2^{exp} and A_4^{exp} coefficients and δ mixing ratios for each pair of γ rays $E_{\gamma 1}$ - $E_{\gamma 2}$ studied in this work, together with corresponding analytical A_2 and A_4 values. The spins of the levels involved are listed in the second column. In the case of nonconsecutive γ -ray cascades, all the four spins of the levels involved are given (see text for details). For the mixing ratio values, the sign convention defined by Krane and Steffen has been used [\[29\]](#page-12-0).

$E_{\gamma 1}$ - $E_{\gamma 2}$	$J_i^{\pi} \to J_m^{\pi} \to J_f^{\pi}$	Multipolarity	A_2^{\exp}	A_4^{\exp}	A ₂	A_4	δ_1	δ_2
1107-310	$1/2^- \rightarrow 3/2^- \rightarrow 5/2^-$	$M1(+E2)-M1(+E2)$	0.176(5)	0.001(5)	0.17	0.00	0.017(8)	0.191(13)
1228-629	$5/2^+ \rightarrow 3/2^- \rightarrow 1/2^-$	$E1(+M2)-M1(+E2)$	0.09(2)	$-0.02(3)$	0.08	0.00	$-0.09(4)$	0.052(11)
$2097 - 310a$	$1/2^+ \rightarrow (1/2^-)$, $3/2^- \rightarrow 5/2^-$	$E1-M1(+E2)$	0.02(2)	0.00(2)	0.00	0.00	0.00(5)	0.191(13)
$2589 - 629$ ^a	$1/2^+ \rightarrow 3/2^+$, $3/2^- \rightarrow 1/2^-$	$M1(+E2)-M1(+E2)$	0.15(2)	0.01(4)	0.14	0.00	$-0.1(2)$ or 2.1(7)	0.052(11)
$3774 - 310^a$	$1/2^+ \rightarrow 3/2^{(-)}$, $3/2^- \rightarrow 5/2^-$	$E1(+M2)-M1(+E2)$	0.08(2)	0.03(2)	0.10	0.00	$-0.1(2)^{b}$	0.191(13)
$3774 - 629$ ^a	$1/2^+ \rightarrow 3/2^{(-)}$, $3/2^- \rightarrow 1/2^-$	$E1(+M2)-M1(+E2)$	0.15(2)	0.01(3)	0.12	0.00	$-0.1(2)^{b}$	0.052(11)
$3952 - 310a$	$1/2^+ \rightarrow 3/2^-$, $3/2^- \rightarrow 5/2^-$	$E1(+M2)-M1(+E2)$	0.06(2)	0.05(3)	0.08	0.00	$-0.1(2)^{c}$	0.191(13)
$3952 - 629$ ^a	$1/2^+ \rightarrow 3/2^-$, $3/2^- \rightarrow 1/2^-$	$E1(+M2)-M1(+E2)$	0.13(2)	0.04(3)	0.11	0.00	$-0.1(2)^{c}$	0.052(11)
$5406 - 629$ ^d	$1/2^+ \rightarrow 3/2^- \rightarrow 1/2^-$	$E1-M1(+E2)$	0.213(10)	$-0.02(2)$	0.20	0.00	$0.00(5)^e$	0.052(11)
5406-692	$1/2^+ \rightarrow 3/2^- \rightarrow 5/2^-$	$E1-M1(+E2)$	0.066(10)	0.005(14)	0.07	0.00	$0.00(5)^e$	0.03(2)
5788-310 ^t	$1/2^+ \rightarrow 3/2^- \rightarrow 5/2^-$	$E1-M1(+E2)$	0.171(2)	$-0.004(4)$	0.16	0.00	$0.00(5)^{e}$	0.191(13)

^aNonconsecutive γ rays.

bValue determined averaging results from the transition pairs 3774-310 and 3774-629 keV.

c Value determined averaging results from the transition pairs 3952-310 and 3952-629 keV.

^dCascade used to determine the δ_{629} mixing ratio of the 629-keV transition.

eValue assumed on the basis of angular momentum selection rules (see text for details).

^fCascade used to determine the δ_{310} mixing ratio of the 310-keV transition.

2324-keV level, since the 3774-keV primary γ transition is more likely to have an $E1 + M2$ character $(1/2^+ \rightarrow 3/2^-)$ with low mixing, rather than $M1 + E2$ ($1/2^+ \rightarrow 3/2^+$), similarly to the other primary transitions reported in Table II .

3. 3509-keV level: J^π **= 3***/***2⁺** *spin assignment*

In earlier works, the 3509-keV level was reported as $3/2^+$ or $5/2$ ⁺ [\[27\]](#page-12-0). To restrict this assignment, the nonconsecutive γ -ray cascade 2589-629 keV was analyzed, where the 2589-keV transition is a primary γ ray populating the 3509keV level [see panel (c) of Fig. 6]. The other nonconsecutive cascade 2589-310 keV could not be used, since the 2589-keV transition is a doublet, with both the transitions of the doublet in coincidence with the 310-keV line. Only the J_{3509}^{π} = $3/2^+$ assignment is possible to obtain fit convergence. This spin-parity assignment implies an $M1 + E2$ character for the 2589-keV transition. By fitting the angular correlation data points, the two solutions $\delta_{2589} = 2.1(7)$ and $-0.1(2)$ are found for the mixing ratio of the 2589-keV transition. No other experimental information is available to further constrain the δ_{2589} value.

4. 4001-keV level: J^π **= (1***/***2−)** *spin assignment*

A tentative spin assignment for the newly observed 4001-keV level is proposed, through the study of the angular correlation of the nonconsecutive γ -ray cascade 2097-310 keV, where 2097 keV is the primary γ ray. The spins considered as possible options are $J_{4001} = 1/2$, 3/2, and 5/2.

The angular correlation appears quite isotropic [see panel (d) of Fig. 6], which is not compatible with $J_{4001} = 5/2$. In turn,

FIG. 6. Examples of angular correlations for pairs of γ rays in ⁶⁵Ni. The red solid line is the function $W(\theta)$ [see Eq. [\(3\)](#page-3-0)] fitted to the experimental points, with A_2^{exp} and A_4^{exp} values given in each panel. The blue dashed line is the curve calculated from the analytical expressions $A_2(\delta)$ and $A_4(\delta)$ by taking the δ mixing ratio which resulted from the analysis discussed above. Such curve reflects the quality of the mixing ratio fit, provided the spins of the states involved are fixed. Detector combinations at the smallest angles are not considered to avoid systematic errors caused by enhancement of the Compton background.

FIG. 7. (a) Excitation energy spectrum of ⁶⁵Ni, up to 1417 keV, from experiment and MCSM calculations. (b) MCSM predictions for the proton occupation number of the $5/2^-_1$ state. (c–j) Calculated differences with respect to the $5/2^-_1$ state for proton occupancy, for each excited state. (k) MCSM predictions for the neutron occupation number of the $5/2₁⁻$ state. (l–s) Calculated differences with respect to the $5/2₁⁻$ state for neutron occupancy, for each excited state. The excited states are indicated at the right side of the panels.

the spin-parity assignment of 1/2−, implying a dominant *E*1 character for the 2097-keV primary γ ray (with a maximum $M2$ mixing of δ_{2097} < 0.05, similarly to the cases discussed above), is consistent with the angular correlation data and the direct decay branch to the $5/2^-$ ground state. The spin $J_{4001} = 3/2$ cannot be firmly excluded, although it would require a large mixing for the primary 2097-keV γ ray. We adopt the tentative spin-parity assignment of $(1/2^-)$ for the 4001-keV level.

5. Other angular correlations considered

Other cascades of intense γ rays were considered in the angular correlation analysis. In particular, on the basis of spinparity assignments from earlier works [\[27\]](#page-12-0), the $1/2^+ \rightarrow 3/2^-$ 3952-keV transition is expected to have $E1(+M2)$ character. In our analysis, by considering the cascades 3952-310 and 3952-629 keV, the mixing ratio $\delta_{3952} = -0.1(2)$ was obtained, similarly to other primary γ rays considered, which have a dominant *E*1 character.

Furthermore, the cascade 5406-692 keV was analyzed. As discussed above, the primary 5406-keV transition is a pure *E*1 (with a maximum *M2* mixing contribution of δ_{5406} < 0.05), while the 692-keV ground-state transition depopulates the 692-keV state with a \approx 0.7-ps lifetime [\[30\]](#page-12-0). The solutions for the *M*1/*E*2 mixing ratio of the $3/2^ \rightarrow$ 5/2⁻ 692-keV transition are $\delta_{692} = 0.03(2)$ and $-5.3(4)$, which lead to *B*(*E*2) reduced transition probabilities of ≈ 0.06 and ≈ 67 W.u.,

respectively. Assuming a low collectivity for the low-lying states of ⁶⁵Ni, as predicted by shell-model calculations (later discussed), the first solution $\delta_{692} = 0.03(2)$ is adopted.

The same kind of argument was used to define the multipolarity mixing ratio for the $1/2^- \rightarrow 3/2^-$ (*M*1 + *E*2) 1107-keV transition, depopulating the 1/2[−] state at 1418 keV, with a lifetime upper limit of 300 fs, as determined in the complementary experiment of Ref. [\[30\]](#page-12-0). The angular correlation study of the 1107-310-keV pair gives the mixing ratio values $\delta_{1107} = 1.66(2)$ and 0.017(8), which correspond to $B(E2) > 72$ and $B(E2) > 0.03$ W.u., respectively. This points to $\delta_{1107} = 0.017(8)$ for the mixing ratio of the 1107-keV transition.

III. COMPARISON WITH MONTE CARLO SHELL-MODEL CALCULATIONS

The experimental level scheme of ⁶⁵Ni has been compared, up to 1.4 MeV, with MCSM results [\[8\]](#page-12-0). The calculations are performed considering an inert core of ⁴⁰Ca and the proton and neutron single-particle orbits $f_{7/2}$, $p_{3/2}$, $f_{5/2}$, $p_{1/2}$, $g_{9/2}$, and $d_{5/2}$ (i.e., the $pf-g_{9/2}-d_{5/2}$ model space).

Figure $7(a)$ shows the comparison between experimental results, in terms of energy, spin and parity of each state, and predictions from the MCSM. All excited states in the energy region here considered have negative parity, apart from the J^{π} =9/2⁺ state, located at an experimental excitation energy of 1017.1 keV. It is found that each experimental state has a theoretical counterpart, although with a discrepancy in energy up to 400 keV. An inversion between the ground state and the first excited state is also observed.

Information on the configurations of the calculated states can be obtained by inspecting the occupation numbers predicted by theory for the proton and neutron single-particle orbitals, as shown in Figs. $7(b)$ – $7(s)$. Figures $7(b)$ and $7(k)$ give the occupation numbers for the $5/2₁⁻$ state, which is the experimental ground state. In this case, while the protons are mainly in the $f_{7/2}$ orbital, the neutrons occupy the $f_{7/2}$, $p_{3/2}$, and $f_{5/2}$ single-particle states, with smaller contributions from $p_{1/2}$, $g_{9/2}$, and $d_{5/2}$. The associated intrinsic shape is predicted to be spherical. Figures $7(c)$ – $7(j)$ and $7(l)$ – $7(s)$ show the differences, in proton and neutron occupation numbers, between each excited state and the $5/2₁⁻$ ground state. All states, apart from $1/2₂⁻$ and $9/2₁⁺$, are characterized by small differences (less than one unit of occupation number), thus indicating similar wave-function compositions as the spherical $5/2₁⁻$ ground state. In contrast, in the case of the $1/2₂$ state, a sizable enhancement in the occupation of the neutron $g_{9/2}$ and proton $p_{3/2}$ and $f_{5/2}$ orbitals, with respect to $5/2₁⁻¹$, is found. For the $9/2₁⁺$ state, naturally the $g_{9/2}$ neutron orbital is occupied. In the neutron-rich Ni isotopes, such occupation pattern is characteristic of deformed structures [\[8\]](#page-12-0). A visualization of the deformation associated with a given MCSM state is offered by the so-called T-plot representation $[8,9]$, as shown in Fig. 8 for the $5/2^-_1$, $1/2^-_1$, $1/2^-_2$, and $9/2^+_1$ states of ⁶⁵Ni. Here, circles on the potential-energy surface indicate the projection of the MCSM eigenstates on the quadrupole moment coordinates. It is clearly seen that for the $5/2₁⁻$ ground state circles are concentrated around $Q_0 = Q_2 = 0$, indicating

FIG. 8. Potential-energy surfaces of ⁶⁵Ni, obtained via constrained Hartree-Fock calculations. Circles on the energy surface represent shapes of MCSM basis vectors according to a T-plot analysis for the states $5/2^-_1$, $1/2^-_1$, $1/2^-_2$, and $9/2^+_1$, shown in panels (a), (b), (c), and (d), respectively (see text for further details).

a spherical shape. The same feature is shown by the $1/2₁⁻$ state [Fig. $8(b)$], pointing to a spherical shape, as already suggested by the similarity of occupation numbers with the $5/2₁⁻$ experimental ground state discussed above. In turn, circles on the T plot corresponding to the $1/2₂⁻$ and $9/2₁⁺$ states are located around the oblate minimum.

IV. CONCLUSIONS

The low-spin structure of the 65 Ni nucleus was investigated in the neutron capture experiment ⁶⁴Ni(*n*, γ)⁶⁵Ni performed at ILL, with the FIPPS array. We observed 24 primary γ rays (15 new) and a total of 87 new γ -ray transitions were found, thus increasing considerably our knowledge on the low-spin structure of this nucleus. In total 28 discrete states (two new) were located below the 6098.3-keV neutron binding energy.

The lower excitation energy part of the level scheme, up to about 1.4 MeV, was compared to MCSM calculations, performed in the *pf* -*g*9/2-*d*5/² neutron and proton model space. The calculations indicate a spherical character of the considered states, apart from the second $1/2₂⁻$ excitation, at 1417.8 keV, and the first $9/2₁⁺$ excitation, at 1017.1 keV. In these cases, a significant occupation of the proton $p_{3/2}$ and *f*⁵/² and neutron *g*⁹/² orbitals is found, which is typical of deformed structure in the neutron-rich Ni region. A more detailed analysis of the state deformation, offered by the Tplot, points to an oblate shape for both the $1/2₂⁻$ and $9/2₁⁺$ states. According to MCSM calculations, prolate deformed structures are expected to occur at higher excitation energies (i.e., around 3 MeV), where the comparison between theory and experiment is more complex, due to the experimental uncertainty in spin-parity assignments.

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- [1] N. Cieplicka-Oryńczak, B. Fornal, S. Leoni, D. Bazzacco, A. Blanc, G. Bocchi, S. Bottoni, G. de France, M. Jentschel, U. Köster, P. Mutti, G. Simpson, T. Soldner, B. Szpak, C. Ur, and W. Urban, Phys. Rev. C **93**[, 054302 \(2016\).](https://doi.org/10.1103/PhysRevC.93.054302)
- [2] N. Cieplicka-Oryńczak, B. Szpak, S. Leoni, B. Fornal, D. Bazzacco, A. Blanc, G. Bocchi, S. Bottoni, G. de France, M. Jentschel, U. Köster, P. Mutti, G. Simpson, T. Soldner, C. Ur, and W. Urban, Phys. Rev. C **94**[, 014311 \(2016\).](https://doi.org/10.1103/PhysRevC.94.014311)
- [3] N. Cieplicka-Oryńczak, C. Michelagnoli, A. Gargano, B. Fornal, S. Leoni, G. Benzoni, A. Blanc, S. Bottoni, F. C. L. Crespi, Ł. W. Iskra, M. Jentschel, U. Köster, P. Mutti, [N. Pietralla, E. Ruiz-Martinez, V. Werner,](https://doi.org/10.1016/j.physletb.2020.135222) Phys. Lett. B **802**, 135222 (2020).
- [4] [M. Girod, J. P. Delaroche, D. Gogny, and J. F. Berger,](https://doi.org/10.1103/PhysRevLett.62.2452) Phys. Rev. Lett. **62**, 2452 (1989).
- [5] P. Bonche, S. Krieger, P. Quentin, M. Weiss, J. Meyer, M. [Meyer, N. Redon, H. Flocard, and P.-H. Heenen,](https://doi.org/10.1016/0375-9474(89)90426-0) Nucl. Phys. A **500**, 308 (1989).
- [6] P. Möller, A. J. Sierk, R. Bengtsson, H. Sagawa, and T. Ichikawa, Phys. Rev. Lett. **103**[, 212501 \(2009\).](https://doi.org/10.1103/PhysRevLett.103.212501)
- [7] P. Möller, A. Sierk, R. Bengtsson, H. Sagawa, and T. Ichikawa, [At. Data Nucl. Data Tables](https://doi.org/10.1016/j.adt.2010.09.002) **98**, 149 (2012).
- [8] Y. Tsunoda, T. Otsuka, N. Shimizu, M. Honma, and Y. Utsuno, Phys. Rev. C **89**[, 031301\(R\) \(2014\).](https://doi.org/10.1103/PhysRevC.89.031301)
- [9] T. Otsuka and Y. Tsunoda, J. Phys. G **43**[, 024009 \(2016\).](https://doi.org/10.1088/0954-3899/43/2/024009)
- [10] B. Crider, C. Prokop, S. Liddick, M. Al-Shudifat, A. Ayangeakaa, M. Carpenter, J. Carroll, J. Chen, C. Chiara, H. David, A. Dombos, S. Go, R. Grzywacz, J. Harker, R. Janssens, N. Larson, T. Lauritsen, R. Lewis, S. Quinn, F. Recchia, A. [Spyrou, S. Suchyta, W. Walters, and S. Zhu,](https://doi.org/10.1016/j.physletb.2016.10.020) Phys. Lett. B **763**, 108 (2016).
- [11] A. Gade and S. N. Liddick, J. Phys. G **43**[, 024001 \(2016\).](https://doi.org/10.1088/0954-3899/43/2/024001)
- [12] S. Leoni, B. Fornal, N. Marginean, M. Sferrazza, Y. Tsunoda, T. Otsuka, G. Bocchi, F. C. L. Crespi, A. Bracco, S. Aydin *et al.*, Phys. Rev. Lett. **118**[, 162502 \(2017\).](https://doi.org/10.1103/PhysRevLett.118.162502)
- [13] B. Olaizola, L. M. Fraile, H. Mach, A. Poves, F. Nowacki, A. Aprahamian, J. A. Briz, J. Cal-González, D. Ghița, U. Köster, W. Kurcewicz, S. R. Lesher, D. Pauwels, E. Picado, D. Radulov, [G. S. Simpson, and J. M. Udías,](https://doi.org/10.1103/PhysRevC.95.061303) Phys. Rev. C **95**, 061303(R) (2017).
- [14] M. Stryjczyk, Y. Tsunoda, I. G. Darby, H. De Witte, J. Diriken, D. V. Fedorov, V. N. Fedosseev, L. M. Fraile, M. Huyse, U. Köster, B. A. Marsh, T. Otsuka, D. Pauwels, L. Popescu, D. Radulov, M. D. Seliverstov, A. M. Sjödin, P. Van den Bergh, P. [Van Duppen, M. Venhart, W. B. Walters, and K. Wimmer,](https://doi.org/10.1103/PhysRevC.98.064326) Phys. Rev. C **98**, 064326 (2018).
- [15] N. Mărginean, D. Little, Y. Tsunoda, S. Leoni, R. V. F. Janssens, B. Fornal, T. Otsuka, C. Michelagnoli, L. Stan, F. C. L. Crespi, C. Costache, R. Lica, M. Sferrazza, A. Turturica, A. D. Ayangeakaa, K. Auranen, M. Barani, P. C. Bender, S. Bottoni, M. Boromiza, A. Bracco, S. Călinescu, C. M. Campbell, M. P. Carpenter, P. Chowdhury, M. Ciemała, N. Cieplicka-Orynczak, ´ D. Cline, C. Clisu, H. L. Crawford, I. E. Dinescu, J. Dudouet, D. Filipescu, N. Florea, A. M. Forney, S. Fracassetti, A. Gade,

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I. Gheorghe, A. B. Hayes, I. Harca, J. Henderson, A. Ionescu, L. W. Iskra, M. Jentschel, F. Kandzia, Y. H. Kim, F. G. Kondev, G. Korschinek, U. Köster, Krishichayan, M. Krzysiek, T. Lauritsen, J. Li, R. Mărginean, E. A. Maugeri, C. Mihai, R. E. Mihai, A. Mitu, P. Mutti, A. Negret, C. R. Niță, A. Olacel, A. Oprea, S. Pascu, C. Petrone, C. Porzio, D. Rhodes, D. ˘ Seweryniak, D. Schumann, C. Sotty, S. M. Stolze, R. Şuvăilă, S. Toma, S. Ujeniuc, W. B. Walters, C. Y. Wu, J. Wu, S. Zhu, and S. Ziliani, Phys. Rev. Lett. **125**[, 102502 \(2020\).](https://doi.org/10.1103/PhysRevLett.125.102502)

- [16] S. E. Arnell, R. Hardell, A. Hasselgren, C.-G. Mattson, and O. Skeppstedt, Phys. Scr. **4**[, 89 \(1971\).](https://doi.org/10.1088/0031-8949/4/3/001)
- [17] S. Cochavi and W. R. Kane, Phys. Rev. C **6**[, 1650 \(1972\).](https://doi.org/10.1103/PhysRevC.6.1650)
- [18] A. F. M. Ishaq, A. Robertson, W. V. Prestwich, and T. J. Kennett, Z. Phys. A **281**[, 365 \(1977\).](https://doi.org/10.1007/BF01408184)
- [19] [R. Vennink, W. Ratynski, and J. Kopecky,](https://doi.org/10.1016/0375-9474(78)90381-0) Nucl. Phys. A **299**, 429 (1978).
- [20] T. Pawłat, R. Broda, W. Królas, A. Maj, M. Ziębliński, H. Grawe, R. Schubart, K. Maier, J. Heese, H. Kluge, and M. Schramm, [Nucl. Phys. A](https://doi.org/10.1016/0375-9474(94)90247-X) **574**, 623 (1994).
- [21] T. Anfinsen, K. Bjorndal, A. Graue, J. Lien, G. Sandvik, L. [Tveita, K. Ytterstad, and E. Cosman,](https://doi.org/10.1016/0375-9474(70)90233-2) Nucl. Phys. A **157**, 561 (1970).
- [22] J. P. Schiffer, C. R. Hoffman, B. P. Kay, J. A. Clark, C. M. Deibel, S. J. Freeman, M. Honma, A. M. Howard, A. J. Mitchell, T. Otsuka, P. D. Parker, D. K. Sharp, and J. S. Thomas, Phys. Rev. C **87**[, 034306 \(2013\).](https://doi.org/10.1103/PhysRevC.87.034306)
- [23] G. Georgiev, I. Matea, D. L. Balabanski, J. M. Daugas, F. de Oliveira Santos, S. Franchoo, F. Ibrahim, F. Le Blanc, M. Lewitowicz, G. Lo Bianco, S. Lukyanov, V. Meot, P. Morel, G. Neyens, Y. E. Penionzhkevich, A. Saltarelli, O. Sorlin, M. Stanoiu, M. Tarisien, N. Vermeulen, D. Verney, and D. Yordanov, [Eur. Phys. J. A: Hadrons Nucl.](https://doi.org/10.1140/epja/i2006-10135-4) **30**, 351 (2006).
- [24] D. Pauwels, O. Ivanov, N. Bree, J. Büscher, T. E. Cocolios, M. Huyse, Y. Kudryavtsev, R. Raabe, M. Sawicka, J. Van de Walle, P. Van Duppen, A. Korgul, I. Stefanescu, A. A. Hecht, N. Hoteling, A. Wöhr, W. B. Walters, R. Broda, B. Fornal, W. Krolas, T. Pawlat, J. Wrzesinski, M. P. Carpenter, R. V. F. Janssens, T. Lauritsen, D. Seweryniak, S. Zhu, J. R. Stone, and X. Wang, Phys. Rev. C **79**[, 044309 \(2009\).](https://doi.org/10.1103/PhysRevC.79.044309)
- [25] [T. Togashi, Y. Tsunoda, T. Otsuka, and N. Shimizu,](https://doi.org/10.1103/PhysRevLett.117.172502) Phys. Rev. Lett. **117**, 172502 (2016).
- [26] C. Michelagnoli, A. Blanc, E. Ruiz-Martinez, A. Chebboubi, H. Faust, E. Froidefond, G. Kessedjian, M. Jentschel, U. Koester, [P. Mutti, and G. Simpson,](https://doi.org/10.1051/epjconf/201819304009) EPJ Web of Conferences **193**, 04009 (2018).
- [27] Evaluated Nuclear Structure Data File Search and Retrieval, [http://www.nndc.bnl.gov/ensdf/ensdf/ensdf.jsp.](http://www.nndc.bnl.gov/ensdf/ensdf/ensdf.jsp)
- [28] H. Morinaga and T. Yamazaki, *In-Beam Gamma-Ray Spectroscopy* (North-Holland, Amsterdam, 1976).
- [29] [K. S. Krane and R. M. Steffen,](https://doi.org/10.1103/PhysRevC.2.724) Phys. Rev. C **2**, 724 (1970).
- [30] M. Sferrazza, S. Bottoni, S. Leoni, B. Fornal, S. Fracassetti, N. Mărginean, C. Mihai, C. Costache, and A. Turturică (unpublished).