Energy dependence study of directed flow in Au+Au collisions using an improved coalescence in a multiphase transport model

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The rapidity-odd component of directed flow (v_1) of identified hadrons $(\pi^{\pm}, K^{\pm}, K^0_S, p, \overline{p}, \phi, \Xi, \overline{\Xi}, \Lambda, \overline{\Lambda})$ and partons $(u, \overline{u}, d, \overline{d}, s, \overline{s})$ in Au+Au collisions at various beam energies $(\sqrt{s_{NN}} = 7.7, 11.5, 14.5, 19.6, 27,$ 39, 54.4, 62.4, 200 GeV) is analyzed using a multiphase transport model. A data driven approach (inspired from the experimental analysis) is performed here to distinguish the transported and produced quarks which are found to have different directed flow values at various collision beam energies. The coalescence sum rule (number of constituent quark scaling) violation is observed at lower energies where hadronic matters dominate. The strange quark (s) and ϕ meson slope (dv_1/dy) show a double sign change around 14.5 GeV unlike other partons and hadrons. It suggests that the strange quark is more sensitive to the softening of the equation of state.

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I. INTRODUCTION

The main goal of relativistic heavy-ion collision experiments is to understand the properties and evolution of strongly interacting matter, called the quark-gluon plasma, as well as to explore the hadron-quark phase transition. The rapidity-odd component of directed flow (v_1) is an important probe to study the in-medium dynamics as it is sensitive to the equation of state (EoS) of the produced medium. Directed flow is generated during the nuclear passage time $(2R/\gamma \approx 0.1 \text{ fm/}c \text{ at})$ 200 GeV) and it probes the onset of bulk collective dynamics in the early stage of the collision [1,2]. As a suggested signature of a first-order phase transition, directed flow is sensitive to the existence of the critical point and it plays an important role in the proposed beam energy scan program [3–8]. The first-order harmonic of the Fourier expansion in momentum distribution of emitted particles is characterized as directed flow.

$$v_1 = \langle \cos(\phi - \Psi_{\text{RP}}) \rangle,$$
 (1)

where ϕ and Ψ_{RP} are the azimuthal angle and reaction plane angle, respectively [9-11]. The v_1 contains both rapidityodd and rapidity-even components. Rapidity-odd component $[v_1^{\text{odd}}(y) = -v_1^{\text{odd}}(-y)]$ is referred to the sideward collective motion of emitted hadrons with respect to collision reaction plane. The rapidity-even component even $[v_1^{\text{even}}(y) =$ $v_1^{\text{even}}(-y)$] is unrelated to the reaction plane and it originates from event-by-event fluctuations in the initial colliding nuclei. In this paper, $v_1(y)$ implicitly refers to the odd component of directed flow. The transport and hydrodynamic models calculation suggested that the directed flow of baryon v_1 at midrapidity $(y \approx 0)$ is sensitive to the equation of state of the system [4,12]. Several hydrodynamic models calculation predict that the negative v_1 -slope near midrapidity called as wiggle or antiflow might be a possible QGP signature [13,14]. Number-of-constituent-quark (NCQ) scaling is an example of coalescence behavior among quarks. Because of the NCQ scaling, which is observed at the BNL Relativistic Heavy Ion Collider (RHIC) [15,16] and the CERN Large Hadron Collider (LHC) [17], the higher order flow harmonics like v_2 behaves as if it is developed at the partonic level [18-20]. There are recent experimental measurement of directed flow of various identified hadrons (π^{\pm} , K^{\pm} , K_{S}^{0} , p, $\bar{p}, \phi, \Lambda, \bar{\Lambda}$) from the STAR collaboration at RHIC over a wide range of colliding beam energies (7.7–200 GeV) [21]. Comprehensive v_1 measurement from STAR [21] supports the coalescence mechanism as the dominant process in particle formation dynamics. There are several studies in heavy-ion collisions to understand the hadron and nuclei formation via coalescence and also hadronization of quarks in heavy-ion collisions [22–29]. In recent articles the importance of coalescence mechanism and energy dependence directed flow are discussed [30-34] and an experimental review of v_1 can be found in Ref. [35].

The interplay between NCQ scaling and the transport of initial-state u and d quarks towards midrapidity during the collision offers possibilities for new insights [36]. The produced strange (s) and antistrange (\bar{s}) quark contribute in the resonance (ϕ) formation and hence also play vital role in understanding the particle formation mechanism. Understanding the strange quarks or particles are very important in order to understand the EoS, as the dv_1/dy of ϕ meson also shows a hints of sign change similar to baryons (p, Λ) [21].

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An approach to study v_1 performed in this paper is inspired from the STAR experiment at RHIC [21], where a comprehensive measurement of directed flow of identified hadrons are reported in a range of collision energies. The experimental paper verified the coalescence sum rule (NCQ scaling) using v_1 measurement although the NCQ scaling is well known in elliptic flow (v_2) measurement of identified hadrons at RHIC and LHC [15–17,37]. Our model calculation is also compared with the experimental results. The calculation can reasonably well describe the data for mesons over a range of energies. v_1 prediction for Ξ and $\overline{\Xi}$ baryons are also given along with the new energy 54.4 GeV for various hadron species.

The paper is organized in the following sections. Section II provides a brief description about the AMPT event generator [38]. The analysis details of calculating directed flow and the results which include the v_1 of partons and hadrons followed by the slope parameter (dv_1/dy) are discussed in the Sec. III. A summary with final remarks are given in Sec. IV.

II. THE AMPT MODEL

A multiphase transport model especially the string-melting version (AMPT-SM) is often used to understand the experimental heavy-ion collision results. The hot and dense matter formed due to relativistic heavy-ion collisions are expected to be in parton degrees of freedom and the AMPT-SM also evolves through the partonic medium, thus makes it a suitable model for interpreting the experimental results. The AMPT-SM version mainly consists of four parts. The initial conditions are taken from Heavy Ion Jet Interaction Generator (HIJING) [39]. Scatterings among partons are described by Zhang's parton cascade (ZPC) [40] model and for hadronization it uses the coalescence model. An extended relativistic transport (ART) model describes the final hadronic evolution [41]. HIJING model includes two-body nucleon-nucleon interactions to form excited strings and minijets via hard and soft processes. The minijet parton undergoes scattering before they fragment to partons and subsequently into hadrons. The partonic interaction in ZPC model is described by two body partonic elastic cross section (σ_p) as given in Eq. (2):

$$\sigma_p = \frac{9\pi\alpha_S^2}{2\mu^2}.\tag{2}$$

In this study the strong coupling constant (α_S) and parton screening mass (μ) are set to be 0.33 and 3.20 fm⁻¹, respectively, leading to $\sigma_p = 1.5$ mb. After partons freeze-out, the hadronization process in AMPT is described by a quark coalescence model. A meson is formed by combining a quark with a nearby antiquark. Similarly, three quarks (antiquarks) combine to form a baryon (antibaryon). Here, the formation process of mesons or baryons (antibaryons) is independent of the relative momentum among the coalescing partons. In this coalescence process, each number of baryons, antibaryons, and mesons in an event are conserved individually. However, in the present study an improved quark coalescence method has been used [42]. The constraint which forced separate conservation of the baryons, antibaryons, and mesons number via the quark coalescence has been removed in the new method. However, the net-baryons and net-strangeness

TABLE I. List of hadrons with their corresponding momentum cuts used in this analysis.

Hadron	p_T cut (GeV/ c)
$p,ar{p} \ \pi^\pm, K^\pm \ \Lambda, ar{\Lambda}, K^0_S, \Xi, \overline{\Xi} \ \phi$	$0.2 < p_T < 2.0$ $p_T > 0.2, p < 1.6$ $0.2 < p_T < 5.0$ $0.15 < p_T < 10.0$

numbers are still conserved for each event. In the new coalescence model, for a meson formation, any available quark searches all available antiquarks and records the closest relative distance (d_M) as the potential coalescence partner. The quark also searches all available quarks to find the closest one in distance as a potential coalescence partner to form a baryon, and then searches all other available quarks again to find the one that gives the smallest average relative distance (d_B) among these three quarks. The condition $d_B < d_M r_{\rm BM}$ has to be satisfied to form a baryon else a meson is formed. A new coalescence parameter $r_{\rm BM}$, controls the relative probability of a quark to form a baryon rather than meson. The limit of $r_{\rm BM} \to 0$ and $r_{\rm BM} \to \infty$ corresponds to no antibaryon formation (although to keep net-baryon number conservation, a minimum number of baryons would be formed) and almost no meson formation, respectively. Similar coalescence procedure is also applied to all antiquarks.

In this analysis the mean field is not included [43]. The new parameter $r_{\rm BM}$ which controls the relative probability to form a baryon via coalescence of a quark is set to 0.61. The popcorn parameter PARJ(5) value is changed to 0 from the default value 1.0 which controls the relative percentage of the $B\bar{B}$ and $BM\bar{M}$ channels. This $r_{\rm BM}$ parameter value is able to describe the dN/dy of proton yields at midrapidity in central Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV and central Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV as shown in the Ref. [42].

III. ANALYSIS AND RESULTS

In this study, an improved version of AMPT-SM model $\sigma_p = 1.5$ mb is used to study the directed flow of identified hadrons in midcentral (10-40%) Au+Au collisions at $\sqrt{s_{NN}}$ = 7.7, 11.5, 14.5, 19.6, 27, 39, 54.4, 62.4, 200 GeV, corresponding to RHIC beam energy scan program (BES-I). The centrality is determined using the charged particle multiplicity ($|\eta| < 0.5$). This analysis is inspired from the recent experimental measurement from STAR at RHIC, where a comprehensive v_1 measurement has been performed and coalescence sum rule is verified. The effect of hadronic interaction on directed flow is also studied by changing the hadron cascade time (t_{max}) in the AMPT-SM. The particles reported here are identified from their PYTHIA-id (PID). The particle selection cuts (e.g., momentum p, transverse momentum p_T) are listed in Table I, which is similar to the experimental data [21,44], in order to have a better comparison.

The directed flow is calculated by averaging the azimuthal angle (ϕ) using the formula $v_1 = \langle \cos(\phi - \Psi_{RP}) \rangle$ with respect to the reaction plane angle, Ψ_{RP} .

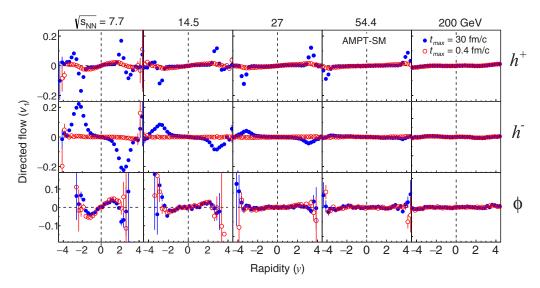


FIG. 1. Directed flow (v_1) as a function of rapidity (y) for hadron cascade time, $t_{\text{max}} = 30 \text{ fm/}c$ (solid marker), 0.4 fm/c (open marker). Upper, middle, and lower rows correspond to positively, negatively charged hadrons, and ϕ meson, respectively in 10–40 % centrality, Au+Au collisions at $\sqrt{s_{NN}} = 7.7$, 14.5, 27, 54.4, and 200 GeV using AMPT-SM.

Figure 1 shows the directed flow of charged hadrons and ϕ meson as a function of rapidity for $t_{\rm max}=0.4$ and 30 fm/c in 10–40% centrality, Au+Au collisions at 7.7, 14.5, 27, 54.4, and 200 GeV. The v_1 of negatively charged hadrons (and positively charged hadrons at higher energies) are found to be

not well developed for $t_{\rm max} = 0.4~{\rm fm/c}$, because the particles could not get enough time to have hadronic interactions unlike the case of $t_{\rm max} = 30~{\rm fm/c}$. However, positively charged hadrons at lower energies for $t_{\rm max} = 0.4~{\rm fm/c}$ have relatively significant v_1 as compared to higher energies because of the

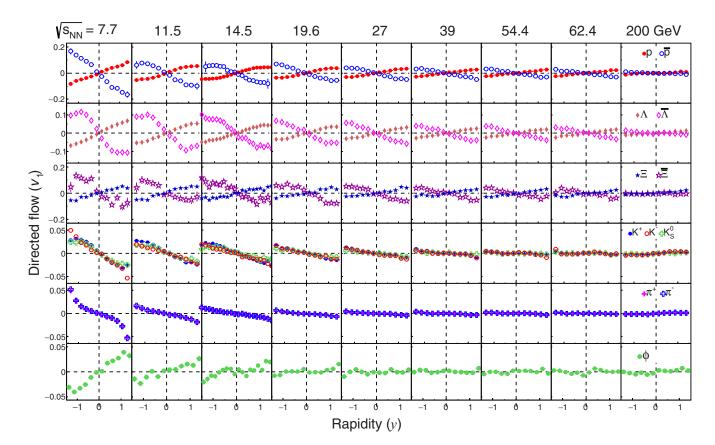


FIG. 2. Directed flow (v_1) as a function of rapidity (y) for hadron cascade time, $t_{\text{max}} = 30 \text{ fm/}c$ for different identified hadrons (rows) in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$, 11.5, 14.5, 19.6, 27, 39, 54.4, 62.4, and 200 GeV (columns).

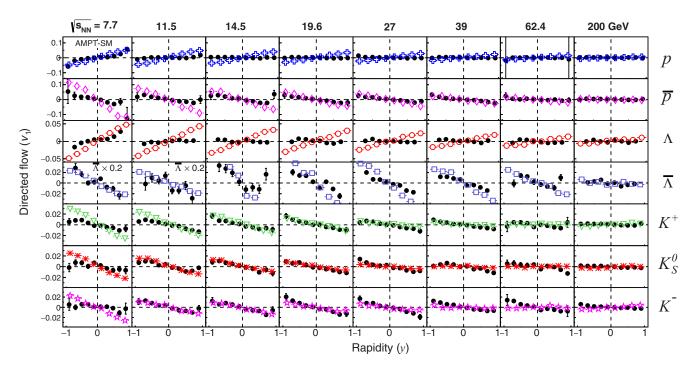


FIG. 3. Directed flow (v_1) as a function of rapidity (y) for different identified hadrons (rows) using AMPT-SM model (hadron cascade time, $t_{\text{max}} = 30 \text{ fm/}c$) is compared with the experimental data (solid circle) from STAR at RHIC [21] in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$, 11.5, 14.5, 19.6, 27, 39, 62.4, and 200 GeV (columns).

dominant transported quark at low energies. It is also observed that the hadronic interaction affects v_1 more at higher rapidity. The ϕ meson in our $t_{\rm max}=30$ fm/c results represent those which have not decayed by the time of 30 fm/c (i.e., ϕ meson that have survived to the time of 30 fm/c). However, ϕ meson life time is relatively large and hence they represent a majority of the total ϕ mesons. The v_1 of ϕ meson for $t_{\rm max}=0.4$ and 30 fm/c are found to be similar at higher energies, i.e., unaffected by hadronic interactions. This is because the ϕ meson has a small hadronic scattering cross section and long life time (\approx 42 fm/c), which thus leads to its decay mainly outside the fireball [45]. We also find that hadronic scatterings have little effect on the proton and antiproton v_1 within |y| < 1.5.

Figure 2 shows the directed flow of various identified hadrons (corresponding rows) as function of rapidity in semicentral (10–40%) Au+Au collisions at different collision beam energies (corresponding columns) using AMPT-SM, $t_{\rm max}=30~{\rm fm/c}$. The rapidity dependence of identified hadrons v_1 gets stronger with decreasing collision beam energy. At highest RHIC energy (200 GeV), particle and antiparticles v_1 values are found to be similar. The v_1 values of baryons and antibaryons have opposite trend and the difference increases with decrease with energy. The mesons like K^\pm and K_S^0 have similar v_1 values like π^+ and π^- over the measured beam energies. ϕ meson v_1 as a function of rapidity is observed to be similar to baryons (p, p, p, which have a strong positive slope at lower energy unlike other mesons (K^\pm , K_S^0 , π^\pm).

Figure 3 shows the comparison of directed flow as a function of rapidity between experimental data from STAR at RHIC [21] and AMPT-SM ($\sigma_p = 1.5 \text{ mb}$, $t_{\text{max}} = 30 \text{ fm/}c$)

calculation for different identified hadrons at various collision energies. The AMPT-SM model better describes the experimental data of mesons as compared to the baryons and antibaryons over the studied energy range.

The strength of directed flow signal at midrapidity is usually characterized by the linear term, F, in the equation $v_1(y) = Fy + F_3y^3$ [21] or by the slope (F') parameter of the fit function $v_1(y) = F'y + C$ [44]. Here, the slope parameter F' is denoted as dv_1/dy . By using the cubic fit function one can reduce sensitivity to the rapidity range in which the fitting is performed. However, in order to have a better comparison we have used the linear fit function similar to the experimental STAR result [21]. The fitting range for various hadron species are |y| < 0.8 for all measured particles except for ϕ meson which is fitted in the rapidity range, |y| < 0.6.

The collision beam energy dependence of the directed flow slope, dv_1/dy , for baryons $(p, \overline{p}, \Lambda, \Lambda, \Xi, \Xi)$ and mesons $(\pi^{-}, \pi^{+}, K^{+}, K^{-}, K_{S}^{0}, \phi)$ are shown in Fig. 4(a) and 4(b), respectively. The dv_1/dy of measured baryons such as p, Λ , and Ξ are found to have similar value and their antiparticles \bar{p} , $\bar{\Lambda}$, and $\bar{\Xi}$ have also similar slope within the uncertainty over the measured energy range. All the measured baryons have positive dv_1/dy whereas their antiparticles have negative slope values. In AMPT-SM, the sign change of baryon's (p and Λ) slope is not observed unlike observed in STAR experiment [21]. The dv_1/dy of π^- and π^- are similar; K^+ and $K^$ values are also similar except for lower energies (<19.6 GeV) and their average value corresponds to slope of K_s^0 meson. All the mesons except ϕ resonance have negative dv_1/dy below 39 GeV collision energy like the corresponding STAR results [21]. Overall magnitude of baryons and antibaryons dv_1/dy are larger than the mesons.

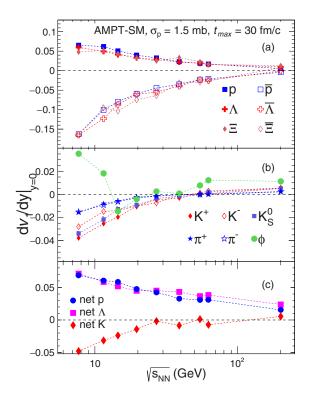


FIG. 4. The slope (dv_1/dy) of baryons, mesons and net-p, net- Λ , net-K are shown in upper, middle, and lower panels as a function of beam energy for 10–40% centrality, respectively. The dotted lines are smooth curves drawn here to guide the eye.

Figure 5 shows the ϕ meson slope calculated by using different fitting ranges for both linear and cubic function in Au+Au collisions from $\sqrt{s_{NN}} = 7.7$ –200 GeV. The dv_1/dy shows the sign change in between 11.5 to 27 GeV for the

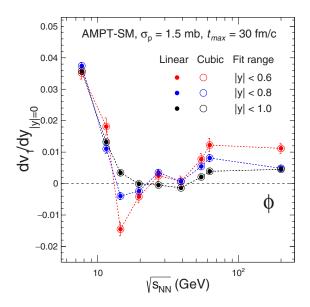


FIG. 5. Beam energy dependence of ϕ -meson slope parameter obtained using different fitting ranges and fit functions for 10–40% centrality. The dotted lines are smooth curves drawn here to guide the eye.

fitting range |y| < 0.6 and |y| < 0.8. The fitting range in the STAR measurement is |y| < 0.6 [21]. When the fitting range increases, the magnitude of negative slope decreases and for |y| < 1.0 the dv_1/dy becomes positive within uncertainty for all measured energies. There is a hint of slope change as observed in STAR [21] although the statistical significance is poor. The slope change of ϕ meson might be due to short range fitting (|y| < 0.6) of v_1 as a function of rapidity. There is no difference between the ϕ meson slope calculated using linear and cubic function even though different fitting ranges are considered. One can also observed that there is a sharp increase in the ϕ meson slope with decrease in energy (<11 GeV) which is similar to the STAR experimental results at RHIC [21].

The energy dependence of proton dv_1/dy receives contribution mainly in two ways: (i) v_1 of transported protons from the initial colliding beam rapidity toward the midrapidity and (ii) v_1 of protons from pair (particle and antiparticle) production near midrapidity. The importance of pair production increases with increase in colliding energy. The "net particle" is a measure of excess particles yield over its antiparticles. It is used to disentangle the transported quarks relative to that of produced in the collisions by using Eq. (3):

$$[v_1(y)]_p = r(y)[v_1(y)]_{\bar{p}} + [1 - r(y)][v_1(y)]_{\text{net}-p}, \quad (3)$$

where r(y) is the rapidity dependence of antiproton to proton ratio at each beam energy. The formulas for net-K and net- Λ are defined in the similar way as Eq. (3). Antiproton v_1 has been proposed as proxy of produced proton v_1 in Ref. [44] and net-p slope is also used to distinguish the transported baryonic matter and hydrodynamic effect [21,44]. There are also model calculation which suggests that the transported quarks (u and d from initial colliding nuclei) contribute more towards the formation of hadrons like p, Λ , and K^+ [36]. Figure 4(c) shows the net-p, net- Λ , and net-K dv_1/dy as a function of beam energy for midcentral (10–40%) Au+Au collisions from $\sqrt{s_{NN}} = 7.7$ –200 GeV. The net-p and net- Λ have positive and similar dv_1/dy unlike the net-K over the measured energy range.

In this analysis, there are several (12) hadrons which allow us to have a comprehensive study of constituent quark v_1 . The assumption like v_1 is developed in prehadronic stage, each type of quark has different directed flow and that hadrons are formed via quark coalescence can be tested here. The coalescence sum rule suggests that at smaller azimuthal anisotropy coefficient (v_n) , the detected hadron's v_n is sum of their constituent quark's v_n [21]. The popular example of NCQ scaling observed at RHIC and LHC are followed from the coalescence sum rule [15–17].

Figure 6 (upper panel) shows the comparison of $\bar{\Lambda}(\overline{uds})$ and $K^-(\bar{u}s)+\frac{1}{3}\bar{p}(\overline{uud})$ slope as a function of beam energy for 10–40% centrality in Au+Au collisions from 7.7–200 GeV. The example stated here is the most suitable to test coalescence sum rule because both $\bar{\Lambda}(\overline{uds})$ and $\bar{p}(\overline{uud})$ are produced unlike the u and d quarks which could be either produced or transported. However, by comparing these two cases we have assumed that s and \bar{s} have same flow. The scale factor $\frac{1}{3}$ is due to the assumption that \bar{u} and \bar{d} have the same v_1 . But we found that except for highest energy both of

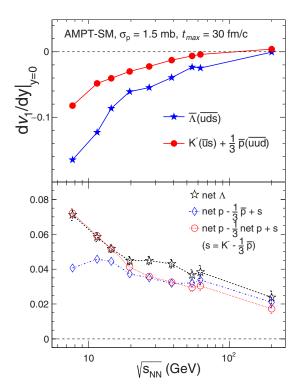


FIG. 6. Upper panel shows the comparison of $\bar{\Lambda}(\overline{uds})$ and sum rule test for produced quark $[K^-(\bar{u}s) + \frac{1}{3}\bar{p}(\overline{uud})]$ slope as a function of beam energy. Lower panel shows another set of sum rule test using net- Λ and net-p for 10–40% centrality in Au+Au collisions using AMPT-SM. The solid and dotted lines are smooth curves drawn here to guide the eye.

them are found to have different slope indicating violation of these assumptions. The dv_1/dy of s and \bar{s} are different except for highest RHIC energy as shown in Fig. 8 for AMPT-SM, $\sigma_p=1.5$ mb. As per the assumption, one can observe that \bar{u} and \bar{d} have similar slope as shown in Fig. 8. The STAR measurement at RHIC also found that the slope of $\bar{\Lambda}(uds)$ and

scaled $\overline{p(uud)}$ have different slope at lower energies [21] and this might be due to the assumption that s and \overline{s} have similar v_1 over all measured energy ranges which may not be valid for lower energies.

Figure 6 (lower panel) shows the first case of coalescence sum rule involving u and d quarks which are either transported or produced and hence it is cumbersome to distinguish them in general. However, one can naively expect that at lower beam energy, u and d quarks are mostly transported whereas these quarks are largely produced at high colliding beam energy. In this figure, two different coalescence sum rule scenarios are compared with the net- Λ (open star). First case is the net-p minus u plus s, where u and s quarks are obtained from $\bar{p}/3$ and $K^{-}(\bar{u}s) - \frac{1}{3}\bar{p}(uud)$, respectively, as represented by blue diamond symbol. Here, the produced u quark in net-p is replaced by s quark. However, we do not have the corresponding straight forward expression for representing the transported u and d quarks. The sum rule is found to be in a good agreement with net-Λ above 39 GeV and start deviating for lower energies. This observation suggests that the fraction of transported quarks in the constituent quarks assembly of net-p increases with decrease in collision beam energy, which imply that the assumption of produced u quark is removed by keeping the term (net- $p - \frac{1}{3}\bar{p}$) also starts deviating. The observation of getting the transported quark dominance at lower energy (≤ 39 GeV) in 10-40% centrality Au+Au collisions using AMPT-SM is qualitatively similar to that of observed in STAR experiment at RHIC [21].

The second case of coalescence sum rule, i.e., $(\frac{2}{3}\text{net}-p+s)$ is also shown in (red open circle marker) Fig. 6 (lower panel). In this sum rule, it is assumed that in the limit of low beam energy, the constituent quarks of net protons are dominated by transported quarks, and an s quark replaces one of the transported quarks. This assumption starts showing disagreement for beam energy greater than 19.6 GeV, i.e., the disagreement between black and red markers.

Figure 7 shows the directed flow of partons $(u, \bar{u}, d, \bar{d}, s, and \bar{s})$ systematic evolution in Au+Au collisions from low

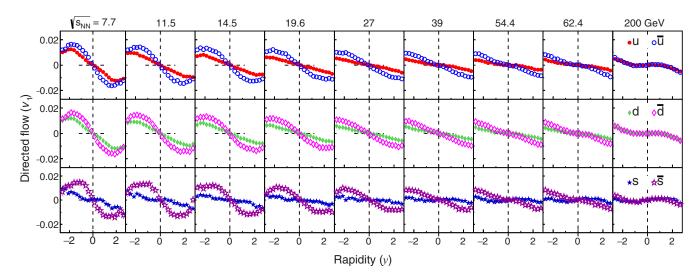


FIG. 7. Directed flow (v_1) as a function of rapidity (y) for different quark (antiquark) solid marker (open marker) are shown in corresponding rows for Au+Au collisions at $\sqrt{s_{NN}} = 7.7$, 11.5, 14.5, 19.6, 27, 39, 54.4, 62.4, and 200 GeV (columns) using AMPT-SM.

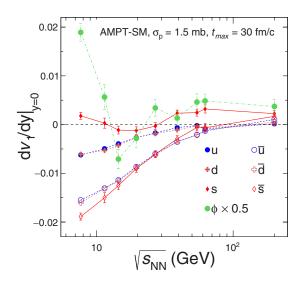


FIG. 8. Slope (dv_1/dy) as a function of $\sqrt{s_{NN}}$ for quark, antiquark, and ϕ meson using AMPT-SM for parton-parton cross section (σ_p) of 1.5 mb. The ϕ meson slope is divided with corresponding number of constituent quarks, i.e., 2.

to high energy ($\sqrt{s_{NN}} = 7.7$ to 200 GeV). All the antiquarks are produced unlike the u and d quarks which might be either transported or produced depending on the collisions beam energy. So, at highest RHIC energy both the quarks and antiquarks have same v_1 as these are expected to be mostly produced. However, with decrease in beam energy the v_1 difference between them increases and antiquarks shows larger directed flow than quarks. v_1 of quarks forming primordial proton have the opposite sign compared with v_1 of all quarks, and further study is needed to understand why this is the case for the quark coalescence in AMPT. Furthermore, the v_1 slope of quarks coalescing to primordial proton and the corresponding proton have similar slope. However, the final proton (Fig. 2) which includes decay contribution and hadronic interaction have similar positive slope like primordial proton but different in magnitude.

Directed flow slope parameter of quarks, antiquarks, and ϕ meson as a function of beam energy is shown in Fig. 8. The slope of u and d quarks is found to be similar and decreases with increase in beam energy unlike the s quark. All the light antiquarks $(\bar{u}, \bar{d}, \text{ and } \bar{s})$ have more negative slope than their corresponding quarks. However, there is a clear deviation of \bar{s} slope from the trend of s quark except at the highest energy. The ϕ meson slope does not scale with the s and \bar{s} quarks' slope.

IV. SUMMARY AND OUTLOOK

A comprehensive study of rapidity-odd component directed flow for charged and identified hadrons in Au+Au collisions (10–40% centrality) for a range of collision beam energies has been discussed using an improved coalescence AMPT-SM model. The coalescence sum rule or commonly known as NCQ scaling is tested using the directed flow measurement of identified hadrons. The analysis performed here is summarized in the following.

The effect of hadronic interaction on v_1 of charged hadrons and ϕ meson are reported. The v_1 of charged hadrons are found to be not well developed for $t_{\text{max}} = 0.4 \text{ fm/}c$, because the particles could not get enough time to have hadronic interactions unlike the case for $t_{\text{max}} = 30 \text{ fm/}c$, except for positively charged hadrons at lower energies where the transported quark effect is more dominant. However, the ϕ -meson v_1 is found to be unaffected by hadronic interactions because of its small hadronic cross section and also because it decays outside the fireball. The double sign change of ϕ meson slope in between 11.5 to 27 GeV is observed. This sign change is also found to be an artifact of small fitting ranges while extracting the v_1 slope. The sign change goes away making positive slope for all measured energies when the linear or cubic function is fitted in a larger rapidity range (|y| < 1.0). This observation emphasizes the crucial importance of fitting range while extracting the slope parameter (dv_1/dy) in the real data measurement. Predictions for directed flow as function of rapidity and slope parameter of various identified hadrons in semicentral (10-40%) Au+Au collisions at $\sqrt{s_{NN}} = 54.4$ GeV are reported. The v_1 calculation of Ξ and $\overline{\Xi}$ baryons is predicted for a range of energy and the values are found to be similar to protons and Λ baryons. The v_1 results at higher rapidity range are also shown here, which cover the event plane detector (EPD) pseudorapidity (η) range installed in STAR detector at RHIC for BES-II program.

We find that light quarks such as u and d have similar slope and it decreases with increase in beam energy unlike the s quark. The antiquarks $(\bar{u}, d, \text{ and } \bar{s})$ have more steeper negative slope than corresponding light quarks and are similar for the measured beam energy range. The s and \bar{s} quarks have different v_1 except for the highest energy. There is a clear indication that \bar{s} quark slope start deviating from the trend of s quark with the decrease in energy. The measured baryons $(p, \Lambda, \text{ and } \Xi)$ have similar positive slope that increases with decrease in beam energy unlike their corresponding antiparticles. The AMPT-SM model shows no sign change for p and Λ slope unlike that observed in STAR experiment at RHIC [21]. The slope of π^+ , π^- , K^+ , K^- , and \hat{K}^0_S mesons are positive at the highest RHIC energy then start decreasing and become negative with decrease in beam energy. The slope of K_s^0 is approximately the average of K^+ and K^- slope as observed by STAR at RHIC [21].

The test of coalescence sum rule using produced quarks is done by comparing the slope of $\bar{\Lambda}(uds)$ and $K^-(\bar{u}d) + \frac{1}{3}\bar{p}(uud)$. These are found to have different slopes and the departure increases with decrease in energy which might be due to a break down of the assumption that s and \bar{s} have same flow over the measured energy range. The slope of net-p and net- Λ are similar over the measured energy range. The sum rule (net- $p - \frac{1}{3}\bar{p} + s$) and net- Λ are found to be similar for energy higher than 39 GeV. The deviation at lower energy might be an indication that the assumption of produced u quarks effect can be removed by keeping the term $\frac{p}{3}$. This assumption does not hold at lower energies, which is similar to the observation by STAR at RHIC [21]. The sum rule $(\frac{2}{3}$ net-p + s) and net- Λ values start deviating

at energy higher than 19 GeV. This sum rule assumes that at lower energy the transported quarks dominate and one of the transported quark of net-*p* is replaced by *s* quark. Hence, this approximation breaks down in the limit of high beam energy which is qualitatively similar to the observation by STAR at RHIC [21].

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- [1] H. Sorge, Phys. Rev. Lett. 78, 2309 (1997).
- [2] N. Herrmann, J. P. Wessels, and T. Wienold, Annu. Rev. Nucl. Part. Sci. 49, 581 (1999).
- [3] D. H. Rischke et al., Heavy Ion Phys. 1, 309 (1995).
- [4] H. Stöcker, Nucl. Phys. A 750, 121 (2005).
- [5] Y. Nara, H. Niemi, A. Ohnishi, and H. Stöcker, Phys. Rev. C 94, 034906 (2016).
- [6] S. A. Bass, M. Belkacem, M. Bleicher *et al.*, Prog. Part. Nucl. Phys. 41, 255 (1998).
- [7] M. Bleicher et al., J. Phys. G 25, 1859 (1999).
- [8] B. A. Li and C. M. Ko, Phys. Rev. C 58, R1382 (1998).
- [9] S. Voloshin and Y. Zhang, Z. Phys. C 70, 665 (1996).
- [10] B. Alver and G. Roland, Phys. Rev. C 81, 054905 (2010); 82, 039903(E) (2010).
- [11] A. M. Poskanzer and S. A. Voloshin, Phys. Rev. C 58, 1671 (1998).
- [12] U. W. Heinz, in *Relativistic Heavy Ion Physics*, Landolt-Boernstein New Series edited by R. Stock (Springer Verlag, New York, 2010), Vol. I.23.
- [13] R. J. M. Snellings, H. Sorge, S. A. Voloshin, F. Q. Wang, and N. Xu, Phys. Rev. Lett. 84, 2803 (2000).
- [14] J. Brachmann, S. Soff, A. Dumitru, H. Stöcker, J. A. Maruhn, W. Greiner, L. V. Bravina, and D. H. Rischke, Phys. Rev. C 61, 024909 (2000).
- [15] J. Adams et al. (STAR Collaboration), Phys. Rev. Lett. 92, 052302 (2004); Phys. Rev. C 72, 014904 (2005); B. Abelev et al. (STAR Collaboration), ibid. 75, 054906 (2007); B. I. Abelev et al. (STAR Collaboration), Phys. Rev. Lett. 99, 112301 (2007).
- [16] S. S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. Lett. **91**, 182301 (2003); A. Adare *et al.* (PHENIX Collaboration), *ibid.* **98**, 162301 (2007); S. Afanasiev *et al.* (PHENIX Collaboration), *ibid.* **99**, 052301 (2007); A. Adare *et al.* (PHENIX Collaboration), Phys. Rev. C **85**, 064914 (2012).
- [17] K. Aamodt *et al.* (ALICE Collaboration), Phys. Rev. Lett. **105**, 252302 (2010); B. Abelev *et al.* (ALICE Collaboration), J. High Energy Phys. 06 (2015) 190.
- [18] J. H. Chen, Y. G. Ma, G. L. Ma, X. Z. Cai, Z. J. He, H. Z. Huang, J. L. Long, W. Q. Shen, C. Zhong, and J. X. Zuo, Phys. Rev. C 74, 064902 (2006).
- [19] Md. Nasim, Phys. Rev. C 89, 034909 (2014).
- [20] M. Nasim, B. Mohanty, and N. Xu, Phys. Rev. C 87, 014903 (2013).

- [21] L. Adamczyk et al. (STAR Collaboration), Phys. Rev. Lett. 120, 062301 (2018).
- [22] P. F. Kolb, L.-W. Chen, V. Greco, and C. M. Ko, Phys. Rev. C 69, 051901(R) (2004).
- [23] L. W. Chen and C. M. Ko, Phys. Rev. C 73, 044903 (2006).
- [24] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. C 93, 021903(R) (2016).
- [25] R. J. Fries, B. Müller, C. Nonaka, and S. A. Bass, Phys. Rev. Lett. 90, 202303 (2003).
- [26] R. J. Fries, V. Greco, and P. Sorensen, Annu. Rev. Nucl. Part. Sci. 58, 177 (2008).
- [27] J. Barrette et al., Phys. Rev. C 50, 1077 (1994).
- [28] S. Albergo et al., Phys. Rev. C 65, 034907 (2002).
- [29] V. Greco, C. M. Ko, and P. Lévai, Phys. Rev. Lett. 90, 202302 (2003).
- [30] J. Y. Chen, J. X. Zuo, X. Z. Cai, F. Liu, Y. G. Ma, and A. H. Tang, Phys. Rev. C 81, 014904 (2010).
- [31] C. Q. Guo, C. J. Zhang, and J. Xu, Eur. Phys. J. A 53, 233 (2017).
- [32] C. Q. Guo, H. Liu, and J. Xu, Phys. Rev. C 98, 024914 (2018).
- [33] M. Nasim and S. Singha, Phys. Rev. C 97, 064917 (2018).
- [34] C. Zhang, J. Chen, X. Luo, F. Liu, and Y. Nara, Phys. Rev. C 97, 064913 (2018).
- [35] S. Singha, P. Shanmuganathan, and D. Keane, Adv. High Energy Phys. 2016, 2836989 (2016).
- [36] J. C. Dunlop, M. A. Lisa, and P. Sorensen, Phys. Rev. C 84, 044914 (2011).
- [37] S. Shi, Adv. High Energy Phys. 2016, 1987432 (2016).
- [38] Z. W. Lin and C. M. Ko, Phys. Rev. C 65, 034904 (2002); L. W. Chen et al., Phys. Lett. B 605, 95 (2005); Z.-W. Lin, C. M. Ko, B. A. Li, B. Zhang, and S. Pal, Phys. Rev. C 72, 064901 (2005).
- [39] X. N. Wang and M. Gyulassy, Phys. Rev. D 44, 3501 (1991).
- [40] B. Zhang, Comput. Phys. Commun. 109, 193 (1998).
- [41] B. A. Li and C. M. Ko, Phys. Rev. C 52, 2037 (1995).
- [42] Y. He and Z. W. Lin, Phys. Rev. C 96, 014910 (2017).
- [43] J. Xu, L.-W. Chen, C. M. Ko, and Z.-W. Lin, Phys. Rev. C 85, 041901(R) (2012)
- [44] L. Adamczyk et al. (STAR Collaboration), Phys. Rev. Lett. 112, 162301 (2014).
- [45] K. A. Olive et al. (Particle Data Group), Chin. Phys. C 38, 090001 (2014).