

## Solution of the sign problem for the half-filled Hubbard-Holstein model

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We show that, by an appropriate choice of auxiliary fields and exact integration of the phonon degrees of freedom, it is possible to define a “sign-free” path integral for the so-called Hubbard-Holstein model at half filling. We use a statistical method, based on an accelerated and efficient Langevin dynamics, for evaluating all relevant correlation functions of the model. Preliminary calculations at  $U/t = 4$  and  $U/t = 1$ , for  $\omega_0/t = 1$ , indicate a region around  $U \simeq \frac{g^2}{\omega_0}$  without either antiferromagnetic or charge-density-wave orders that is much wider compared to previous approximate calculations. The elimination of the sign problem in a model without explicit particle-hole symmetry may open different perspectives for strongly correlated models, even away from the purely attractive or particle-hole symmetric cases.

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**Introduction.** One of the most successful methods to obtain exact properties of strongly correlated models on a lattice is certainly the statistical (Monte Carlo) method based on the evaluation of a corresponding path integral defined in imaginary time. In particular, most successful applications are based on auxiliary fields  $\sigma_l^j$  introduced for each site  $j$  of the model and imaginary-time slice  $l$  of the path integral by means of the so-called “Hubbard-Stratonovich” transformation (HST) [1–4]. Since Hirsch’s seminal work in 1985 [5], several models have been studied, and their phase diagrams have been solved numerically for a large enough number  $N$  of sites, in very particular cases when the so-called sign problem does not affect the simulation of the corresponding partition function  $Z = \int [d\sigma_l^j] W(\{\sigma_l^j\})$  that is evaluated by standard statistical methods, as long as  $W(\{\sigma_l^j\}) \geq 0$ . The first example was the Hubbard model in the square lattice, displaying a trivial phase diagram for the insulating antiferromagnetic phase that turned out to be stable as soon as  $U > 0$ . More recently the method was extended to the honeycomb lattice displaying a less trivial transition at a critical value  $U_c$  between a semimetallic and an antiferromagnetic insulating or superconducting phase [6–9]. Other models are worth mentioning, such as the negative- $U$  model [10,11], spinless fermions with a repulsive nearest-neighbor interaction at half filling [12,13], the Anderson impurity model at half filling [14,15], the Kondo-lattice model at half filling [16,17], the Holstein model [18], and several others.

Most of the models so far solved without a sign problem are characterized by (i) an explicit spin-independent attractive interaction and/or (ii) a particular particle-hole symmetry of the electronic degrees of freedom, implying that the corresponding weight  $W(\{\sigma_l^j\})$  in the path-integral formulation can be written as the square of a quantity, and is therefore positive. All these models have been recently classified in Refs. [19,20]. For instance, in the Hubbard model with  $U > 0$ ,

the particle-hole transformation

$$c_{j\downarrow}^\dagger \rightarrow c_{j\downarrow} \quad (1)$$

( $c_{j\downarrow}^\dagger \rightarrow -c_{j\downarrow}$ ) for sites  $j$  in the  $A$  ( $B$ ) sublattice of a bipartite lattice, maps the positive- $U$  model to the negative- $U$  one with an equal number of spin-up and spin-down electrons, where  $c_{j,\sigma}^\dagger$  ( $c_{j,\sigma}$ ) creates (destroys) a fermion with spin  $\sigma = \uparrow, \downarrow$  at a given site  $j$ . In such a case the weight factorizes into two independent and identical contributions for different spins, thus  $W(\{\sigma_l^j\}) \geq 0$ .

The Hubbard-Holstein Hamiltonian is one of the simplest models describing the competition between an attractive interaction mediated by an optical phonon and the strong electron repulsion, defined by the Hubbard  $U$ , acting when two electrons of opposite spins occupy the same site. The Hubbard-Holstein model represents the key model to understand how the retarded interaction mediated by phonons can circumvent the strong electron-electron repulsion and give rise to superconductivity. It may be relevant not only to understand standard electron-phonon superconductivity, but also high-temperature superconductivity, because the isotope effect has been clearly detected [21] in cuprates, and the so-called kinks observed in photoemission experiments [22] clearly indicate the role of phonons, even in these strongly correlated materials.

The phase diagram of the model has been studied using several techniques, such as the Gutzwiller approximation [23], variational Monte Carlo (VMC) [24,25], dynamical mean-field theory (DMFT) [26–28], finite-temperature determinant quantum Monte Carlo (DQMC) [29–31], also in one dimension (1D) [32], but no unbiased zero-temperature calculation is known in two dimensions (2D).

In the present Rapid Communication we are able to establish ground-state benchmark results in the thermodynamic limit for this model, and some aspects of its zero-temperature

phase diagram, by using a determinantal method, which, as we are going to show, is not vexed by the so-called “sign problem.”

*Model and method.* The Hubbard-Holstein model is defined by the following Hamiltonian,

$$\begin{aligned}\mathcal{H} &= \mathcal{H}_K + \mathcal{H}_V, \\ \mathcal{H}_K &= K + \frac{\omega_0}{2} \sum_j \hat{P}_j^2, \\ \mathcal{H}_V &= \frac{U}{2} \sum_j (n_j - 1)^2 + g \sum_j \hat{X}_j (n_j - 1) + \frac{\omega_0}{2} \sum_j \hat{X}_j^2 \\ &= \frac{U}{2} \sum_j \left( n_j - 1 + \frac{g}{U} \hat{X}_j \right)^2 + \frac{\omega_0 - g^2/U}{2} \sum_j \hat{X}_j^2,\end{aligned}\quad (2)$$

where  $K = -t \sum_{(i,j),\sigma} c_{i,\sigma}^\dagger c_{j,\sigma} + \text{H.c.}$ ,  $n_{j\sigma} = c_{j,\sigma}^\dagger c_{j,\sigma}$  indicates the electron number with spin  $\sigma$  at the site  $j$ , and  $n_j = \sum_\sigma n_{j\sigma}$ .  $t$  is the hopping integral,  $U$  is the repulsive electron-electron interaction, whereas  $\omega_0$  is the phonon frequency,  $g$  is the electron-phonon coupling term, and  $\hat{X}_j$  and  $\hat{P}_j$  are the phonon position and momentum degrees of freedom, respectively.

At finite inverse temperature  $\beta$  the partition function of an electronic system described by the Hamiltonian  $\mathcal{H}$  is given by

$$\mathcal{Z} = \text{Tr}[e^{-\beta\mathcal{H}}] = \text{Tr}[(e^{-\Delta\tau\mathcal{H}})^T], \quad (3)$$

where  $\Delta\tau = \beta/T$  and the symbol  $\text{Tr}[O]$  associates a number to any operator  $O$  and is defined by

$$\text{Tr}[O] = \begin{cases} \text{Trace}[O] & \text{for the standard case,} \\ \langle \Psi | O | \Psi \rangle & \text{for the projection case,} \end{cases} \quad (4)$$

where  $\Psi$  is a chosen trial function that may be conveniently introduced for evaluating the trace (up to an irrelevant constant) in the zero-temperature limit, as long as  $\Psi$  has a nonzero overlap with the ground state of  $\mathcal{H}$ . The latter case is known as zero-temperature projection, which is adopted in all the forthcoming calculations. However, for the sake of generality, we derive the method in the general case, as it is defined also within the more conventional finite-temperature scheme.

The imaginary-time propagator  $e^{-\beta\mathcal{H}}$  can be written after Trotter decomposition as

$$e^{-\Delta\tau\mathcal{H}} = e^{-\Delta\tau\mathcal{H}_V} e^{-\Delta\tau\mathcal{H}_K} + O(\Delta\tau^2), \quad (5)$$

In order to derive the path integral, the phonon degrees of freedom, introduced for each site  $j$  and time slice  $l$ , are dealt with as in a conventional Feynmann path integral where the phonon positions  $X_l^j$  are changed at different time slices, just due to the phonon kinetic energy  $\frac{\omega_0}{2} \sum_j \hat{P}_j^2$ , with associated matrix elements,

$$\begin{aligned}\langle X_{l+1}^j | \exp\left(-\frac{\Delta\tau\omega_0}{2} \hat{P}_j^2\right) | X_l^j \rangle \\ \propto \exp\left[-\frac{1}{2\omega_0\Delta\tau} (X_{l+1}^j - X_l^j)^2\right].\end{aligned}\quad (6)$$

After that the operators  $\hat{X}_j$  turn onto classical real variables  $X_l^j$  to be integrated from  $-\infty$  to  $\infty$  in the corresponding path

integral. For the remaining interaction term in  $\mathcal{H}_V$  we can use a properly chosen HST coupled to the operator  $n_j - 1 + \frac{g}{U} X_l^j$ , namely,

$$e^{-\frac{U\Delta\tau}{2} (n_j - 1 + \frac{g}{U} X_l^j)^2} = \int_{-\infty}^{\infty} \frac{d\sigma_l^j}{\sqrt{2\pi}} e^{-\frac{1}{2}(\sigma_l^j)^2 + i\sqrt{U\Delta\tau}\sigma_l^j (n_j - 1 + \frac{g}{U} X_l^j)}, \quad (7)$$

where  $\sigma_l^j$  are indicating the auxiliary fields in the  $l$ th time slice. We thus get that the partition function can be expressed as a  $2N \times T$  dimensional integral over the classical real variables  $\sigma_l^j$  and  $X_l^j$ ,

$$\begin{aligned}\mathcal{Z} &= \int [dX] \int [d\sigma] \\ &\times \exp\left[-\frac{1}{2} \sum_{l,m,j} [A_{l,m} X_l^j X_m^j + \delta_{l,m} (\sigma_l^j)^2]\right] \\ &\times \text{Tr} \prod_{l=1}^T \left\{ \exp\left[\sum_j (i\sqrt{U\Delta\tau}\sigma_l^j) (n_j - 1 + \frac{g}{U} X_l^j)\right] \right. \\ &\left. \times \exp(-\Delta\tau K) \right\},\end{aligned}\quad (8)$$

where the product of noncommuting operators is meant from left to right with increasing  $l$ ,  $[dX] = \prod_{j,l} dX_l^j$ ,  $[d\sigma] = \prod_{j,l} d\sigma_l^j$ , and

$$\begin{aligned}A_{l,m} &= \frac{1}{\omega_0\Delta\tau} [2\delta_{l,m} - \delta_{l,m+1} - \delta_{m,l+1}] \\ &+ \Delta\tau(\omega_0 - g^2/U)\delta_{l,m}.\end{aligned}\quad (9)$$

Here, the boundary conditions for the phonon fields are  $X_{l+T}^j = X_l^j$  for the standard finite-temperature case (periodic in imaginary time) and  $X_0^j = X_{T+1}^j = 0$  for the projection case (open boundaries in imaginary time, within appropriate trial function  $\Psi$  [33]). In the path integral the dependence of the action on the fields  $X_l^j$  is just quadratic and determined by the matrix  $A$ . After simple inspection, the eigenvalues of  $A$  are given by

$$E_n = \frac{1}{\omega_0\Delta\tau} [2 - 2\cos(\omega_n\Delta\tau)] + \Delta\tau(\omega_0 - g^2/U), \quad (10)$$

where  $\omega_n\Delta\tau = \frac{2\pi n}{T}$  ( $\omega_n\Delta\tau = \frac{n\pi}{T+1}$ ) for the finite-temperature (projection) case and  $n = 1, \dots, T$ . Therefore this matrix  $A$  is positive definite ( $E_n > 0 \forall n$ ) if  $U > g^2/\omega_0$  and all the integrals in  $\{X_l^j\}$  can be carried out before the  $\{\sigma_l^j\}$  ones, as they are certainly converging,

$$\begin{aligned}\mathcal{Z} &\propto \int [d\sigma] \exp\left[-\frac{1}{2} \sum_{l,m,j} (P_{l,m}\sigma_l^j\sigma_m^j)\right] \\ &\times \text{Tr} \prod_l \left\{ \exp\left[i\sqrt{U\Delta\tau} \sum_j \sigma_l^j (n_j - 1)\right] \exp(-\Delta\tau K) \right\},\end{aligned}\quad (11)$$

where

$$P_{l,m} = \delta_{l,m} + \frac{\Delta\tau g^2}{U} [A^{-1}]_{l,m} \quad (12)$$

is real and providing a positive weight in Eq. (11). On the other hand, the remaining part contributing to the path-integral weight  $W(\{\sigma_i^j\})$ , and resulting from the Tr operation is certainly positive, because the spin-up and spin-down contributions factorize and, after the particle-hole transformation in Eq. (1), turn out to be complex-conjugate factors (with appropriate  $\Psi$  in the projection case). Thus we have finally determined a path integral for the Hubbard-Holstein model with a non-negative real weight  $W(\{\sigma_i^j\}) \geq 0$  for  $U > g^2/\omega_0$ .

A standard approach to evaluate correlation functions defined by the partition function  $\mathcal{Z}$  is the Monte Carlo (MC) method. Unfortunately, the standard technique with local updates is very inefficient in this case, due to the difficulty to sample the stiff harmonic part. A better method was recently introduced [30], including global updates of the phonon fields. Global updates are numerically very demanding as they require the computation of several determinants from scratch. Here, in order to define an efficient sampling, we will use the first-order accelerated Langevin dynamics [34], reviewed and generalized recently in Ref. [35]. The auxiliary fields at the  $n$ th Markov chain iteration are updated via the discretized Langevin equation,

$$\vec{\sigma}_{n+1} = \vec{\sigma}_n + \Delta_{\text{LD}} S^{-1} \vec{f}_n + \sqrt{2\Delta_{\text{LD}}} S^{-1/2} \vec{z}_n, \quad (13)$$

where  $\vec{\sigma} = \{\bar{\sigma}_i^j\}$  are shorthand notations for the fields represented as an  $N \times T$  dimensional vector,  $\Delta_{\text{LD}}$  is the Langevin-dynamics (LD) time step,  $S_{l,m}^{i,j}$  is the acceleration matrix that is chosen diagonal in the spatial indices  $S_{l,m}^{i,j} = P_{l,m} \delta_{i,j}$  and corresponding to the harmonic classical part of the partition function,  $\vec{z}_n$  are normally distributed random vectors, and  $\vec{f}_n = \{f_i^j\}_n$  are generalized forces with components,

$$\{f_i^j\}_n = \partial_{\sigma_i^j} \ln [W(\{\sigma_i^j\}_n)]. \quad (14)$$

It has been shown [36] that, within the complex auxiliary-field technique, the LD is free of ergodicity issues and we are able therefore to reproduce the results for the standard Hubbard model with  $g = 0$  (see Fig. 1), which represents the most difficult case in our approach as the acceleration matrix  $S$  turns out to be the trivial identity matrix. Better choices should be possible but this study is beyond the main purpose of this work.

**Results.** In order to access the information about the order parameters, we examine the charge and spin correlations of the model for different values of the Hubbard interaction and electron-phonon coupling on  $L \times L$  square lattices, at  $\omega_0/t = 1$ .

We adopt the recently proposed dynamic scaling [37]. We break the spin symmetry with a mean-field magnetic wave function  $\Psi$ , so that the antiferromagnetic order remains in the  $z$  direction for the chosen projection times  $\beta t = L$ . The thermodynamic limit  $\beta t = L = \infty$  remains unbiased, whereas the finite-size results do not recover the singlet finite- $L$  ground state, reachable only for much larger projection times. This technique has the considerable advantage to allow very stable simulations without the so-called ‘‘spikes’’ (samples of

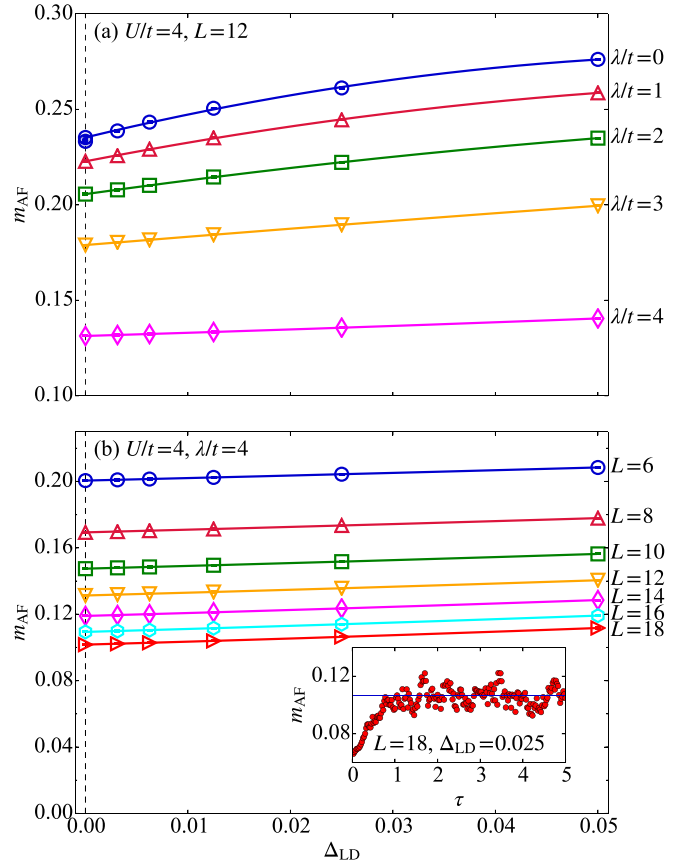


FIG. 1. Antiferromagnetic order parameters  $m_{\text{AF}}$  at  $U/t = 4$  as a function of the LD time step  $\Delta_{\text{LD}}$  (a) on a  $12 \times 12$  cluster at  $\beta t = 12$  and different values of the electron-phonon coupling strength  $\lambda$ . The  $\lambda = \Delta_{\text{LD}} = 0$  result (solid circle) in (a) is obtained with the standard Monte Carlo algorithm for the Hubbard model that is clearly consistent with the LD data, extrapolated to  $\Delta_{\text{LD}} \rightarrow 0$ . (b) Same as (a) for  $\lambda = 4$  with various system sizes  $N = L \times L$ , with  $\beta t = L$ . The inset shows the equilibration of  $m_{\text{AF}}$  to its average value (blue line) for the largest cluster as a function of the LD time  $\tau$ .

correlation functions much farther from their average values) implying infinite variance problems [38]. Within this setup  $m_{\text{AF}}$  can be computed as

$$m_{\text{AF}} = \frac{1}{N} \sum_i e^{i\mathbf{Q}\cdot\mathbf{r}_i} \langle S_i \rangle, \quad (15)$$

where  $S_i = \frac{1}{2}(n_{i\uparrow} - n_{i\downarrow})$  is the value of the spin at site  $i$  and the charge structure factor is given by

$$S_{\text{CDW}}(\mathbf{Q}) = \frac{1}{N} \sum_{i,j} e^{i\mathbf{Q}\cdot(\mathbf{r}_i - \mathbf{r}_j)} \langle n_i n_j \rangle, \quad (16)$$

where  $\langle \dots \rangle = \frac{\text{Tr}[\exp(-\beta\mathcal{H}/2) \dots \exp(-\beta\mathcal{H}/2)]}{\mathcal{Z}}$ , and the pitching vector  $\mathbf{Q} = (\pi, \pi)$ . We consider the evolution of these quantities as a function of the coupling  $\lambda = \frac{g^2}{\omega_0}$ . For  $U \gg \lambda$  we have a Mott insulator with a finite antiferromagnetic (AF) order parameter  $m_{\text{AF}} > 0$ . As it is shown in Fig. 1(b), the dependence of this quantity on the LD time step  $\Delta_{\text{LD}}$  is rather smooth and can be safely extrapolated to the unbiased  $\Delta_{\text{LD}} \rightarrow 0$  limit. Similar behavior is obtained for all the other

quantities considered in this work. The LD is particularly efficient just in the interesting region  $\lambda \simeq U$  where phase transitions or at least competitions between antiferromagnetic, charge-density-wave (CDW), or metallic and superconducting phases are expected [28]. In this case, the chosen acceleration matrix is particularly efficient because it allows short correlation times  $\simeq 1$  [see Fig. 1(b) and the inset] and a very weak time-step dependency, allowing large-scale simulations in this region.

As far as the systematic error implied by a finite Trotter time  $\Delta\tau$ , this becomes negligible provided measurements are evaluated at the middle of the kinetic-energy propagator  $\exp(-\Delta\tau K)$  [39], because in this way the error turns out quadratic in  $\Delta\tau$ . We have adopted  $\Delta\tau t = 0.1$  in all forthcoming calculations with an estimated error of less than 1% in all quantities studied.

We have performed a finite-size scaling of  $m_{\text{AF}}$  and  $S_{\text{CDW}}(\mathbf{Q})$  for  $U/t = 4$  and  $U/t = 1$  using clusters of sizes ranging from  $6 \times 6$  to  $18 \times 18$  for several couplings  $\lambda \leq U$  and report these quantities extrapolated to the thermodynamic limit in Fig. 2. As it is seen, the antiferromagnetic order drops continuously to much smaller values when we increase  $\lambda$  and suggests a continuous transition to a nonmagnetic phase at  $\lambda = \lambda_c^{\text{AF}} \simeq U$ . Within this assumption, and considering that the pure Holstein model for  $U = 0$  (i.e.,  $\lambda \gg U$ ) displays CDW order,  $S_{\text{CDW}}(\mathbf{Q})$  should diverge for  $\lambda \rightarrow \lambda_c^{\text{CDW}}$  from below, because the finite values reported in Fig. 2 imply no CDW order up to  $\lambda/U = 1$ . Quite interestingly, the results reported in Fig. 2 suggest that  $\lambda_c^{\text{CDW}}$  is significantly larger than  $\lambda_c^{\text{AF}}$ , because at small  $U/t$  there is no evidence of the  $S_{\text{CDW}}(\mathbf{Q})$  divergence, whereas for  $U/t = 4$ , despite the fit of the data being consistent with a very small critical exponent  $\theta$  that is very likely underestimated,  $\lambda_c^{\text{CDW}}$  is about 8% larger than  $\lambda_c^{\text{AF}}$ . Indeed, if we fix  $\theta$  to a larger value in the fit, we obtain an even larger value of  $\lambda_c^{\text{CDW}}$ .

*Conclusions.* In this Rapid Communication we have presented an original way to get rid of the sign problem in the Hubbard-Holstein model by using an appropriate auxiliary-field transformation combined with an exact integration of the phonon degrees of freedom. The Hubbard-Holstein model has been considered so far with algorithms affected by the sign problem, because, though at half filling, it does not satisfy the particle-hole transformation in the electronic degrees of freedom,

$$c_{j,\sigma}^\dagger = (-1)^j c_{j,-\sigma}, \quad (17)$$

where  $(-1)^j = 1 (-1)$  if  $j$  belongs to the  $A$  ( $B$ ) sublattice. This transformation leaves the model unchanged without electron-phonon coupling but changes its sign when present. The key idea of this work is to employ an exact integration of the phonon degrees of freedom that allows us to recover this property and get rid of the sign problem, at least in a relevant parameter region  $U > \lambda$ . As we have shown, the region  $\lambda \simeq U$  is important because it is close to the phase transition of the model, and was previously inaccessible by numerical methods due to very severe sign problems [30]. On the other hand, in realistic materials the Coulomb energy is much larger than the electron-phonon coupling, as well as the phonon frequency  $\omega_0$ , and therefore the region  $\lambda < U$

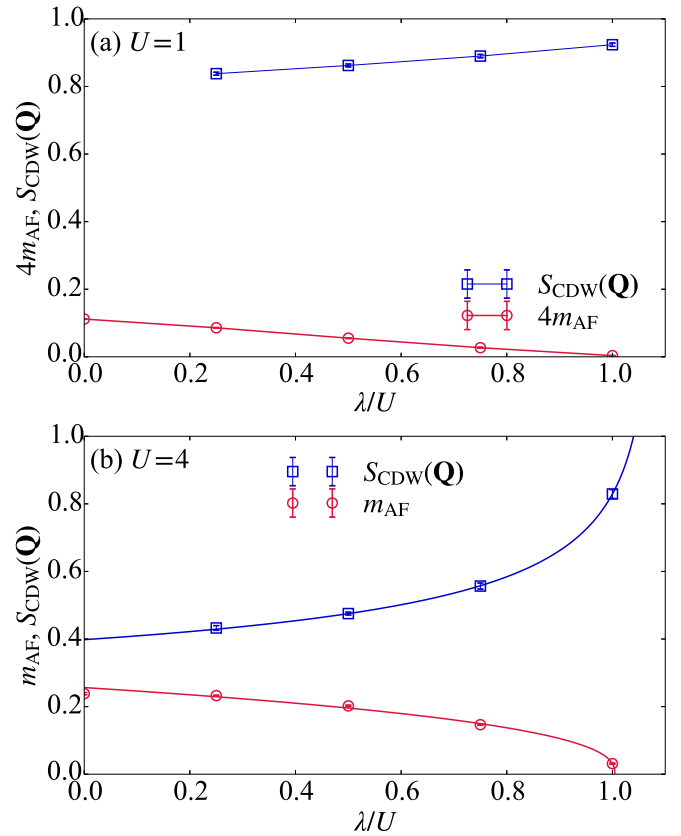


FIG. 2. Ground-state properties of the Hubbard-Holstein model obtained by extrapolating the antiferromagnetic order parameters to the thermodynamic limit for (a)  $U = 1$  and (b)  $U = 4$ . The solid lines in (b) are the fit to  $m_{\text{AF}}$  and  $S_{\text{CDW}}(\mathbf{Q})$  of the form  $\gamma(\frac{\lambda_c - \lambda}{U})^\theta$  with  $\gamma$ ,  $\theta$ , and  $\lambda_c$  being the fitting parameters determined by the least-squares method. The parameters are found to be  $\gamma = 0.256(8)$ ,  $\theta = 0.4$ , and  $\lambda_c = \lambda_c^{\text{AF}} = 4.02(2)$  for  $m_{\text{AF}}$  and  $\gamma = 0.407(2)$ ,  $\theta = -0.29$ , and  $\lambda_c = \lambda_c^{\text{CDW}} = 4.33(4)$  for  $S_{\text{CDW}}(\mathbf{Q})$ . The values of the exponent  $\theta$  are just indicative, as we do not have enough data close to the transitions.

that can be studied with the present technique is certainly the most important region for modeling realistic materials with the Hubbard-Holstein Hamiltonian.

Though the Hubbard-Holstein model is highly idealized, it is interesting to establish some benchmark results for the magnetic order parameter and the density structure factor (see Table I). We see that our estimated  $m_{\text{AF}}$  compares well

TABLE I. Values of  $m_{\text{AF}}$  and  $S_{\text{CDW}}(\mathbf{Q})$  in the thermodynamic limits, for different values of  $\lambda$ ,  $U/t$  at  $\omega_0 = t$ .

$\lambda/U$	$U/t = 1$		$U/t = 4$	
	$m_{\text{AF}}$	$S_{\text{CDW}}(\mathbf{Q})$	$m_{\text{AF}}$	$S_{\text{CDW}}(\mathbf{Q})$
0	0.0280(2)		0.238(3)	
0.25	0.0215(3)	0.838(4)	0.232(2)	0.433(7)
0.50	0.0138(3)	0.862(4)	0.202(4)	0.475(4)
0.75	0.0068(4)	0.890(5)	0.146(2)	0.557(9)
1	0.0009(1)	0.924(5)	0.031(2)	0.83(1)



with the established benchmarks [40] for  $\lambda = 0$ , and remains approximately the same for  $\lambda \ll U$ , but with no evidence of CDW order, because  $S_{\text{CDW}}(\mathbf{Q})$  is clearly finite. Thus, as soon as  $\lambda > 0$ , the electron-phonon coupling breaks the pseudo-SU(2) symmetry of the pure Hubbard model, and kills the CDW, leaving the AF order alone, in agreement with a rigorous theorem, recently proved [41]. This feature is reminiscent of the phase diagram of the negative- $U$  Hubbard model where the CDW order disappears immediately by a tiny amount of doping [10].

Finally, since the transition to a nonmagnetic phase is very close to  $\lambda \simeq U$ , we have been able to determine some aspects of its phase diagram, namely, that the transition is most likely continuous, at least up to  $U/t = 4$ , as no evidence of a first-order transition has been found for the  $U$  values studied so far. Also, rather unexpectedly, an intermediate phase  $\lambda_c^{\text{AF}} \leq \lambda \leq \lambda_c^{\text{CDW}}$  with no AF and CDW orders appears to be rather robust and wide, in contrast with previous DMFT and VMC results.

This technique can be possibly extended to many other models, so far affected by the sign problem, and may open the way to tackle other important models where the particle-hole symmetry is not satisfied, such as the Hubbard model at finite doping. Though we do not expect that the sign problem in this model can be definitively removed, this work certainly suggests that the sign of the weight  $W(\{\sigma_i^j\})$  can be very likely improved, being a property of the appropriate auxiliary field chosen, and the degrees of freedom selected in the path integral, where enormous freedom still needs to be explored.

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- [1] R. Blankenbecler, D. J. Scalapino, and R. L. Sugar, *Phys. Rev. D* **24**, 2278 (1981).
- [2] J. E. Hirsch, *Phys. Rev. B* **28**, 4059 (1983).
- [3] F. Becca and S. Sorella, *Quantum Monte Carlo Approaches for Correlated Systems* (Cambridge University Press, Cambridge, UK, 2017).
- [4] F. F. Assaad, *Phys. Rev. B* **71**, 075103 (2005).
- [5] J. E. Hirsch, *Phys. Rev. B* **31**, 4403 (1985).
- [6] S. Sorella and E. Tosatti, *Europhys. Lett.* **19**, 699 (1992).
- [7] S. Sorella, Y. Otsuka, and S. Yunoki, *Sci. Rep.* **2**, 992 (2012).
- [8] Y. Otsuka, S. Yunoki, and S. Sorella, *Phys. Rev. X* **6**, 011029 (2016).
- [9] Y. Otsuka, K. Seki, S. Sorella, and S. Yunoki, *Phys. Rev. B* **98**, 035126 (2018).
- [10] R. T. Scalettar, E. Y. Loh, J. E. Gubernatis, A. Moreo, S. R. White, D. J. Scalapino, R. L. Sugar, and E. Dagotto, *Phys. Rev. Lett.* **62**, 1407 (1989).
- [11] S. Karakuzu, K. Seki, and S. Sorella, *Phys. Rev. B* **98**, 075156 (2018).
- [12] E. F. Huffman and S. Chandrasekharan, *Phys. Rev. B* **89**, 111101 (2014).
- [13] Z.-X. Li, Y.-F. Jiang, and H. Yao, *Phys. Rev. B* **91**, 241117 (2015).
- [14] J. E. Hirsch and R. M. Fye, *Phys. Rev. Lett.* **56**, 2521 (1986).
- [15] M. Feldbacher, K. Held, and F. F. Assaad, *Phys. Rev. Lett.* **93**, 136405 (2004).
- [16] F. F. Assaad, *Phys. Rev. Lett.* **83**, 796 (1999).
- [17] S. Capponi and F. F. Assaad, *Phys. Rev. B* **63**, 155114 (2001).
- [18] J. E. Hirsch and E. Fradkin, *Phys. Rev. B* **27**, 4302 (1983).
- [19] L. Wang, Y.-H. Liu, M. Iazzi, M. Troyer, and G. Harcos, *Phys. Rev. Lett.* **115**, 250601 (2015).
- [20] Z.-X. Li, Y.-F. Jiang, and H. Yao, *Phys. Rev. Lett.* **117**, 267002 (2016).
- [21] J. Hofer, K. Conder, T. Sasagawa, G.-m. Zhao, M. Willemin, H. Keller, and K. Kishio, *Phys. Rev. Lett.* **84**, 4192 (2000).
- [22] A. Lanzara, P. V. Bogdanov, X. J. Zhou, S. A. Kellar, D. L. Feng, E. D. Lu, T. Yoshida, H. Eisaki, A. Fujimori, K. Kishio, J.-I. Shimoyama, T. Noda, S. Uchida, Z. Hussain, and Z.-X. Shen, *Nature (London)* **412**, 510 EP (2001).
- [23] P. Barone, R. Raimondi, M. Capone, C. Castellani, and M. Fabrizio, *Phys. Rev. B* **77**, 235115 (2008).
- [24] T. Ohgoe and M. Imada, *Phys. Rev. Lett.* **119**, 197001 (2017).
- [25] S. Karakuzu, L. F. Tocchio, S. Sorella, and F. Becca, *Phys. Rev. B* **96**, 205145 (2017).
- [26] P. Werner and A. J. Millis, *Phys. Rev. Lett.* **99**, 146404 (2007).
- [27] G. Sangiovanni, O. Gunnarsson, E. Koch, C. Castellani, and M. Capone, *Phys. Rev. Lett.* **97**, 046404 (2006).
- [28] J. Bauer and A. C. Hewson, *Phys. Rev. B* **81**, 235113 (2010).
- [29] E. A. Nowadnick, S. Johnston, B. Moritz, R. T. Scalettar, and T. P. Devereaux, *Phys. Rev. Lett.* **109**, 246404 (2012).
- [30] S. Johnston, E. A. Nowadnick, Y. F. Kung, B. Moritz, R. T. Scalettar, and T. P. Devereaux, *Phys. Rev. B* **87**, 235133 (2013).
- [31] M. Weber and M. Hohenadler, *Phys. Rev. B* **98**, 085405 (2018).
- [32] M. Hohenadler, F. F. Assaad, and H. Fehske, *Phys. Rev. Lett.* **109**, 116407 (2012).
- [33] Strictly speaking, in this case the diagonal elements  $A_{l,l}$  for  $l = 1, T$  have to be divided by two, but this change does not allow a simple diagonalization of the matrix and, on the other hand, corresponds only to an irrelevant change of  $|\Psi\rangle$ .
- [34] G. Parisi, in *Progress in Gauge Field Theory*, edited by G. 't Hooft *et al.*, NATO ASI Series B (Springer, Berlin, 1984), pp. 531–541.
- [35] G. Mazzola and S. Sorella, *Phys. Rev. Lett.* **118**, 015703 (2017).
- [36] S. Beyl, F. Goth, and F. F. Assaad, *Phys. Rev. B* **97**, 085144 (2018).
- [37] P. Weinberg and A. W. Sandvik, *Phys. Rev. B* **96**, 054442 (2017).
- [38] H. Shi and S. Zhang, *Phys. Rev. E* **93**, 033303 (2016).
- [39] In other words, measurements are done by inserting the operator  $O$  corresponding to a given correlation function, in the following way,  $\dots \exp(-\frac{\Delta\tau}{2}K)O \exp(-\frac{\Delta\tau}{2}K)\dots$ , where the dots indicate the remaining one-body propagators in  $W(\{\sigma_i^j\})$ .
- [40] M. Qin, H. Shi, and S. Zhang, *Phys. Rev. B* **94**, 085103 (2016).
- [41] T. Miyao, *Ann. Henri Poincaré* **19**, 2543 (2018).