# <span id="page-0-0"></span>**Revealing frustrated local moment model for pressurized hyperhoneycomb iridate: Paving the way toward a quantum spin liquid**

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There have been tremendous experimental and theoretical efforts toward the discovery of a quantum spin-liquid phase in honeycomb-based-lattice materials with strong spin-orbit coupling. Here the bond-dependent Kitaev interaction between local moments provides strong magnetic frustration and, if it is the only interaction present in the system, it will lead to an exactly solvable quantum spin-liquid ground state. In all of these materials, however, the ground state is in a magnetically ordered phase due to additional interactions between local moments. Recently, it has been reported that the magnetic order in the hyperhoneycomb material, *β*-Li<sub>2</sub>IrO<sub>3</sub>, is suppressed upon applying hydrostatic pressure and the resulting state becomes a quantum paramagnet or possibly a quantum spin liquid. Using *ab initio* computations and strong-coupling expansion, we investigate the lattice structure and resulting local moment model in pressurized  $β$ -Li<sub>2</sub>IrO<sub>3</sub>. Remarkably, the dominant interaction under high pressure is not the Kitaev interaction or further neighbor interactions, but a different kind of bond-dependent interaction. This leads to strong magnetic frustration and may provide a platform for discovery of a new kind of quantum spin-liquid ground state.

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#### **I. INTRODUCTION**

Magnetic frustration is often regarded as a prominent route to realize quantum spin-liquid states, i.e., novel quantum paramagnetic states with fractionalized excitations [\[1\]](#page-5-0). In the Kitaev model on the honeycomb lattice, magnetic frustration is achieved by bond-dependent Ising interactions, where there exist a macroscopic number of classically degenerate ground states [\[2\]](#page-5-0). The quantum ground state can be solved exactly and is shown to be a quantum spin liquid. Recently, much effort has been put forward to realize the Kitaev interaction in honeycomb-based-lattice materials with strong spin-orbit coupling [\[3–5\]](#page-5-0), where the spin-orbit coupling and edgesharing octahedra structure allow such interactions [\[6\]](#page-5-0). This physics has been explored in two-dimensional honeycomb lattice systems such as Na<sub>2</sub>IrO<sub>3</sub> [\[7–15\]](#page-5-0),  $\alpha$ -Li<sub>2</sub>IrO<sub>3</sub> [\[9,16\]](#page-5-0), and RuCl3 [\[17–19\]](#page-5-0), as well as three-dimensional hyperhoneycomb  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub> [\[20–28\]](#page-5-0) and stripy honeycomb *γ* -Li<sub>2</sub>IrO<sub>3</sub> [\[29–](#page-5-0)[34\]](#page-6-0) systems. Here the local moments on Ir (or Ru) ions can be described by the pseudospin  $j_{\text{eff}} = 1/2$  degree of freedom, which is a spin-orbit entangled Kramers doublet [\[35,36\]](#page-6-0).

These materials, however, develop magnetic ordering at low temperatures, defying attempts to achieve quantum spinliquid ground states [\[7,18,24](#page-5-0)[,37\]](#page-6-0). It has been shown that such magnetic ordering occurs due to the presence of other interactions between  $j_{\text{eff}} = 1/2$  moments [\[31](#page-5-0)[,38\]](#page-6-0). On the other hand, the nature of the observed magnetic order is strongly dependent on the Kitaev interaction, which is an indirect evidence that the strength of the Kitaev interaction in these materials is significant. This suggests that if there is a way to control the relative strength of these interactions, one may be able to achieve a quantum spin-liquid ground state.

Very recently, hydrostatic pressure was applied to the hyperhoneycomb material,  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub>, and it was found that the

In this paper, we theoretically investigate the lattice structure and local moment model for  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub> under hydrostatic pressure using *ab initio* density functional theory (DFT) computations and strong-coupling expansion. It is shown that the dominant interaction between local moments in a highpressure structure is the so-called symmetric anisotropic (SA) interaction which depends on bond directions, as explained below. On the other hand, the usual Heisenberg and Kitaev interactions are generally suppressed and, in contrast to a naive expectation, further neighbor interactions are not so significant. If only the SA interaction is present, the classical version of the model is highly frustrated and there exists macroscopic degeneracy of classically degenerate ground states. Interestingly, the manifold of classically degenerate states in the SA model is very different from that of the Kitaev model [\[31\]](#page-5-0). This points to an interesting possibility that the quantum version of such a model may support a quantum spinliquid state that is distinct from the Kitaev spin-liquid state.

More specifically, we find that the space group of the *ab initio* optimized lattice structure remains unchanged (*Fddd*, SG. 70) under pressure at least up to 10.2 GPa, while the lattice parameters become more anisotropic compared to those at ambient pressure. The local moment model in the strongcoupling limit has the following general form [\[38\]](#page-6-0):

 $H = \sum$ -*ij* ∈*αβ*(*γ* )  $[J_{ij}S_i\cdot S_j + K_{ij}S_i^{\gamma}S_j^{\gamma} \pm \Gamma_{ij}(S_i^{\alpha}S_j^{\beta} + S_i^{\beta}S_j^{\alpha})],$ (1)

magnetic order disappears for sufficiently high pressure while the material remains insulating [\[39\]](#page-6-0). The nuclear magnetic resonance (NMR) and specific-heat measurements found no signature of any broken symmetry, which could be regarded as a sign of a possible quantum spin-liquid ground state. Hence the question is what kind of local moment interactions are present in the high-pressure phase and whether such interactions would lead to a quantum spin-liquid ground state.

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<span id="page-1-0"></span>

FIG. 1. Hyperhoneycomb network of Ir atoms in *β*-Li<sub>2</sub>IrO<sub>3</sub> shown in the conventional orthorhombic unit cell (depicted in black lines). Each site is connected to three nearest-neighbor sites by *X*- (green), *Y*- (blue), and *Z*- (red) type bonds. Note that the *X* and *Y* bonds are equivalent under  $C_2$  rotations. **a** and **b** lattice vectors, which are more sensitive than the **c** lattice vector to pressure, are depicted as black arrows.  $\hat{x}$ ,  $\hat{y}$ , and  $\hat{z}$  are local cubic axes.

where  $S_i$  is the  $j_{\text{eff}} = 1/2$  pseudospin at site *i*, the summation is over the nearest-neighbor (NN) bonds  $\langle i, j \rangle$  labeled by  $\gamma \in (X, Y, Z)$ , and  $\langle i, j \rangle \in \alpha \beta(\gamma)$  is shorthand for  $\langle i, j \rangle \in$  $\gamma, \alpha \neq \beta \neq \gamma$ . The  $\pm$  sign in front of  $\Gamma$  is a reminder that unlike the *J* and *K* terms, the  $\Gamma$  term can have relative minus signs on different bonds (the sign structures of  $\Gamma$  are explained in Ref.  $[31]$ ). Here, *J*, *K*, and  $\Gamma$  represent the Heisenberg, Kitaev, and SA interactions, respectively. At ambient pressure, the magnitudes of *J*, *K*, and  $\Gamma$  are uniform for all of the *X*, *Y*, and *Z*-type bonds (see Fig. 1 for NN bonds and Fig. [4](#page-3-0) for the magnitudes). When  $K$  is the dominant interaction with nonzero  $\Gamma$  and *J*, this model can explain the incommensurate counter-rotating spiral order observed in a resonant elastic x-ray scattering experiment [\[31\]](#page-5-0). A previous DFT computation shows that the material in the experimentally determined structure is indeed in this parameter regime [\[26\]](#page-5-0).

Upon increasing pressure, DFT results indicate that the bond lengths of the *X*- and *Y*-type bonds become shorter than that of the *Z*-type bond. The biggest change occurs in the  $t_{dd\sigma}$ -type hopping integral, which represents a direct overlap in the *σ* bonding channel between *d* orbitals at NN sites. As explained later, this change makes the Kitaev and Heisenberg interactions much smaller and these interactions on the *X*-, *Y*-, and *Z*-type bonds become anisotropic. The dominant interaction, however, is the SA interaction  $\Gamma$  while it becomes also bond anisotropic. In addition, further neighbor interactions are found to be, in general, less than 10% of the NN interactions. Hence it is clear that a good starting point for the local moment model at high pressure is the SA interaction model, which is highly frustrated at the classical level and holds promise for a quantum spin-liquid state.

## **II. COMPUTATIONAL DETAILS**

For the electronic structure calculations, we employ the Vienna *ab-initio* Simulation Package (VASP), which uses the projector-augmented wave (PAW) basis set [\[40,41\]](#page-6-0). 520 eV of plane-wave energy cutoff is used, and for *k*-point sampling, a  $9 \times 9 \times 9$  grid including  $\Gamma$  point is employed for the primitive cell. On-site Coulomb interaction is incorporated using the Dudarev's rotationally invariant DFT+*U* formalism [\[42\]](#page-6-0) with effective  $U_{\text{eff}} \equiv U - J = 2$  eV. We employ two different trial magnetic configurations, i.e., Néel-type and zigzag-type antiferromagnetic orders, which yield the same result. For each configuration with different cell volume and magnetism, structural optimization for the cell shape and internal coordinates is performed with a force criterion of  $1 \text{ meV}/\text{\AA}$  and without any symmetry constraints. A revised Perdew-Burke-Ernzerhof generalized gradient approximation (PBEsol) [\[43\]](#page-6-0) is used for structural optimizations and total-energy calculations, which yields the best agreement of calculated lattice parameters to the experimental ones in conjunction with spin-orbit coupling (SOC) and *U*eff [\[23\]](#page-5-0). Optimized structures are tabulated in Table [I](#page-3-0) of Appendix [A.](#page-5-0) After the structural optimizations, the hopping integrals between the Ir  $t_{2g}$  orbitals are computed by employing maximally localized Wannier orbital formalism (MLWF) [\[44,45\]](#page-6-0) implemented in the WANNIER90 package [\[46\]](#page-6-0), but without including  $U_{\text{eff}}$  and magnetism. The computed  $t_{2g}$  Wannier hopping integrals are presented in Table [II](#page-4-0) of Appendix [B.](#page-5-0) The VESTA  $[47]$  package was used to draw the crystal structure in Fig. 1.

It should be mentioned that structure optimizations in the absence of SOC or the Coulomb interaction lead to severe Ir-Ir dimerization of the *Z* bond, regardless of the choice of exchange-correlation functionals and other parameters. Since such dimerization has not been observed in experimental crystal structures [\[23,24\]](#page-5-0), we conclude that both SOC and Coulomb interaction are crucial in maintaining the observed hyperhoneycomb structure in  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub>. Note that a similar suppression of dimerization due to SOC was reported in the quasi-two-dimensional  $\alpha$ -RuCl<sub>3</sub>, which has the similar local geometry of edge-sharing metal-anion octahedra [\[48\]](#page-6-0).

# **III. EVOLUTION OF CRYSTAL STRUCTURE UNDER PRESSURE**

Figure 2 shows the evolution of the **a** and **c** lattice constants and NN bond lengths with respect to the hydrostatic pressure (more details about optimized crystal structures are in Table [I](#page-3-0)



FIG. 2. Pressure dependence of the ratios of lattice constants and NN bond lengths with respect to the experimental values  ${\bf a}_0, {\bf c}_0, X_0, Z_0$  at ambient pressure, reported in Ref. [\[23\]](#page-5-0). Note that  $\mathbf{a}/\mathbf{a}_0 \simeq \mathbf{b}/\mathbf{b}_0$  and  $d_\mathbf{X} = d_\mathbf{Y}$ .

of Appendix [A\)](#page-5-0). In Fig. [2,](#page-1-0) the volume of the unit cell is reduced from 103% to 91% with a decrement of 3%, where the largest and smallest volume correspond to the pressures of  $P = -3.9$  and 10.2 GPa in our calculations. Note that the ratio between the **a** and **c** lattice parameters becomes closer to the experimental  $\mathbf{a}_0/\mathbf{c}_0$  at  $P = -3.9$  GPa, and hence we take this pressure as a reference point. It is shown in the figure that the **a** (and **b**) lattice parameters are reduced by ∼1*.*5% more than the **c** parameter, implying that the *X* and *Y* bonds, forming the zigzag chains in the hyperhoneycomb structure (shown in Fig. [1\)](#page-1-0), are more compressed than the *Z* bonds. Indeed, the *X*-bond length, denoted as  $d_X$ , is two times more compressed than the *Z*-bond length  $d_Z$  at  $P = 10.2$  GPa;  $d_X^0 - d_X$  and  $d_Z^0 - d_Z$  are 3.4 and 1.7% of the experimental  $d_X^0$  and  $d_Z^0$  at ambient pressure, respectively [\[23\]](#page-5-0).

Compared to the lattice constants and the NN Ir-Ir bond lengths, the Ir-O bond lengths show smaller changes.  $d_{\text{Ir−O}}^Z$ and *d<sup>X</sup>* Ir−O, the Ir-O bond lengths participating in the NN *Z* and *X* bonds, are reduced by∼1*.*4% and 1.2%, respectively, when *P* is increased from −3*.*9 to 10.2 GPa. These changes are smaller compared to the ∼3 to 4.5% reduction of the NN Ir-Ir bond lengths. From this comparison, it can be deduced that the direct hopping channels due to the direct overlap of neighboring Ir  $t_{2g}$  orbitals, which are relevant to the Ir-Ir bond length, should be more affected by hydrostatic pressure than the oxygenmediated indirect channels, relevant to the Ir-O bond length. This is confirmed in the computation of the hopping integrals, as presented in the next section.

#### **IV.** *t***2***<sup>g</sup>* **HOPPING CHANNELS**

The hopping integrals between the NN Ir  $t_{2g}$  orbitals  $\{d_{xz}, d_{yz}, d_{xy}\}$  for the *X* and *Z* bonds, represented by 3 × 3 matrices, are as follows:

$$
\hat{T}^Z = \begin{pmatrix} t_1 & t_2 & t_i \\ t_2 & t_1 & -t_i \\ -t_i & t_i & t_3 \end{pmatrix}, \quad \hat{T}^X = \begin{pmatrix} t_3 & t_4 & t'_4 \\ t_4 & t_1 & t_2 \\ t'_4 & t_2 & t'_1 \end{pmatrix}, \quad (2)
$$

where the forms of  $\hat{T}^Z$  and  $\hat{T}^X$  are determined by the pointgroup symmetries at the *Z*- and *X*-bond centers [\[23,31\]](#page-5-0). Note that  $\hat{T}^{Y}$  can be obtained by applying twofold rotations to  $\hat{T}^{X}$ . Here the most dominant terms are  $t_1$  ( $t_1$ ),  $t_2$ , and  $t_3$ , which originate from  $t_{dd\delta}$ -like direct,  $t_{dpd\pi}$ -like indirect, and  $t_{dd\sigma}$ -like direct overlaps, respectively. The sign of the  $t_2$  term at the  $X$ bond flips when the twofold rotations along the  $\hat{z}$  and  $\mathbf{a} \parallel \hat{x} - \hat{y}$ axes are applied, and hence we show only the value of  $|t_2|$ hereafter. Other minor components,  $t_i$  and  $t_4$  (and  $t'_4$ ), come from trigonal distortions, where the antisymmetric  $t_i$  terms in  $\hat{T}^Z$  arise due to the absence of inversion at the *Z*-bond center. Detailed illustrations for such terms in  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub> are presented in Ref. [\[26\]](#page-5-0). Note that since Ir-O-Ir bond angles become closer to 90 $\degree$  when pressure is increased, the magnitudes of the  $t_i$ ,  $t_4$ , and  $t_4$  terms are reduced below 5% of that of the largest hopping term. The difference between  $t_1$  and  $t'_1$  at the *X* bond also reduces from  $\sim$ 20 to 6% of the average of *t*<sub>1</sub> and *t*<sup>'</sup><sub>1</sub> as *P* is increased from −3*.*9 to 10.2 GPa. Hence, hereafter we denote  $t_1$  as the averaged value of  $t_1$  and  $t_1'$  and present the evolution of  $t_{1,2,3}$  as a function of *P*.



FIG. 3. Pressure dependence of the ratios of three Ir  $t_{2g}$  hopping integrals for the *X* and *Z* bonds, with respect to  $t_{1,2,3}^{Z_0}$  denoting the  $t_{1,2,3}$ channels for the *Z* bond at  $P = -3.9$  GPa, respectively. Solid and dashed lines depict the evolution of hopping amplitudes in the *X* and *Z* bonds, respectively. Note that at *P* = −3*.*9 GPa, the NN hopping terms are isotropic.

Figure 3 shows the evolution of the ratios  $t_{1,2,3}/t_{1,2,3}^{Z_0}$  with respect to pressure, where  $t_{1,2,3}^{Z_0}$  are the values of the *Z*-bond hopping terms at  $P = -3.9$  GPa ( $t_1^{Z_0} = 80$ ,  $|t_2^{Z_0}| = 248$ , and  $t_3^{Z_0} = -139$  meV). As expected in the previous section, the  $t_{dd\sigma}$ -like  $t_3$  channel shows the largest enhancement of 260% at the *X* bond. Due to the larger compression of the *X* bond compared to the *Z* bond,  $t_3^X$  becomes 75% larger than  $t_3^Z$ . This huge enhancement makes  $t_3$  the dominant hopping term at *P* = 10*.*2 GPa; −365 and −248 meV for the *X* and *Z* bonds, respectively. The  $t_{dd\delta}$ -like  $t_1$  channel is also increased by the pressure, with smaller enhancement compared to *t*3. On the contrary, the  $t_2$  channel is almost unchanged with the small decrease of ∼4–8% at *P* = 10*.*2 GPa, due to the cancellation between the  $t_{dpd}$ -like indirect and  $t_{dd}$ -like direct overlaps within the  $t_2$  channel. As suggested in other systems with similar local crystal structure [\[19](#page-5-0)[,48\]](#page-6-0), these changes in NN hopping channels greatly affect the magnetic exchange interactions between the  $j_{\text{eff}} = 1/2$  pseudospins in the strongly correlated regime, as we will discuss in the following section.

It should be mentioned that compared to these huge changes in the NN channels, the second, third, and further-neighbor channels do not show any significant changes. For example, the largest NN hopping term  $(t_{NNN}^{\text{II}})$  in Ref. [\[26\]](#page-5-0)) is enhanced from 77 to 78 meV as *P* is increased from −3*.*9 to 10.2 GPa. The largest third-  $(t_{3NN}^{\text{II}})$  in Ref. [\[26\]](#page-5-0)) and fourth-neighbor terms, corresponding to  $-45$  and  $-31$  meV at  $P = -3.9$  GPa, respectively, are enhanced at most by 15 meV as *P* is increased. From these results, we conclude that the role of furtherneighbor terms is not significant in the pressure-induced paramagnetic phase of  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub>.

# **V. MAGNETIC EXCHANGE INTERACTIONS AT HIGH PRESSURE**

The huge changes in the NN hopping channels upon pressure affect the  $j_{\text{eff}} = 1/2$  NN exchange interactions <span id="page-3-0"></span>substantially, where each interaction term in the spin model written in Eq.  $(1)$  is represented as follows  $[31,38]$  $[31,38]$ :

$$
J = \frac{4}{27} \left[ \frac{(4J_H + 3U)(2t_1 + t_3)^2}{U^2} - \frac{16J_H(t_1 - t_3)^2}{(2U + 3\lambda)^2} \right], \quad (3)
$$

$$
K = \frac{32J_H}{9} \left[ \frac{(t_1 - t_3)^2 - 3t_2^2}{(2U - 3\lambda)^2} \right], \quad \Gamma = \frac{64J_H}{9} \frac{t_2(t_1 - t_3)}{(2U + 3\lambda)^2},\tag{4}
$$

where  $U$ ,  $J_H$ , and  $\lambda$  are the on-site Coulomb interaction, Hund's coupling, and Ir  $t_{2g}$ -orbital SOC, respectively. In this study, we employ  $U = 2.0$  eV,  $J_H/U = 0.2$ , and  $\lambda = 0.45$  eV. Note that apart from the overall energy scale, the ratios between the exchange interactions are almost insensitive to  $J_H/U$ when  $J_H/U > 0.05$ . In principle, the additional SA term  $\Gamma'$ is allowed to exist, which is proportional to  $t_4$  (and  $t_4$ ) as discussed in Ref. [\[49\]](#page-6-0), and the DM vector parallel to the bond direction at the *Z* bond is allowed as well. However, their magnitudes become insignificant as pressure is increased.

Figure  $4(a)$  shows the calculated values of the exchange interactions, where all the values are divided by the absolute mieractions, where an the values are divided by the absolute<br>magnitude of  $\sqrt{(J_2^0)^2 + (K_2^0)^2 + (\Gamma_2^0)^2}$  along the *Z* bond at *P* = −3*.*9 GPa and are shown as dimensionless numbers. Two notable features are found: (i) the SA term  $\Gamma$  is enhanced significantly by the pressure. At a relatively low pressure of  $\sim$  − 2.5 GPa, the SA term overcomes *K* and becomes the largest term. It becomes even larger under higher pressure; at  $P = 2.1$  GPa, the ratio between the magnitudes of the exchange interactions at the *X* bond is  $|J_X|: |K_X|: |\Gamma_X| = 0.43$ : 0.34 : 1, where all  $J, K, \Gamma < 0$ . As such, magnetic properties in the pressurized  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub> would be distinct from those of the Kitaev-dominated phases at ambient pressure [\[31\]](#page-5-0). (ii) While the *X* and *Z* bonds share almost the same values



FIG. 4. Pressure dependence of the exchange interactions for the  $j_{\text{eff}} = 1/2$  pseudospins. Note that dimensionless val- $\cos{\{\bar{J}, \bar{K}, \bar{\Gamma}\}}_{X,Z} \equiv \{J, K, \Gamma\}_{X,Z}/\sqrt{(J_Z^0)^2 + (K_Z^0)^2 + (\Gamma_Z^0)^2}$  are shown, where  $\{J, K, \Gamma\}_{Z}^{0}$  are the exchange interactions for the *Z* bond at *P* = −3*.*9 GPa.

of exchange interactions at  $P = -3.9$  GPa, the anisotropy between the *X* and *Z* bonds becomes significant in the highpressure regime of  $P > 5$  GPa with the sign flip of  $K$  on the  $X$ bonds. The anisotropy in the NN interactions becomes larger than the strength of further-neighbor exchange-interaction terms, and hence the anisotropy in the NN interactions would play a more significant role than further-neighbor interactions in the high-pressure phase of  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub> [\[50\]](#page-6-0).

As noticed in previous studies on  $\alpha$ -{Li,Na}<sub>2</sub>IrO<sub>3</sub> (Refs. [\[19,](#page-5-0)[51\]](#page-6-0)), the magnitude of exchange interactions is sensitive to the local Ir-O geometry, especially to the ratio between the NN Ir-Ir and Ir-O bond lengths. This ratio is in turn controlled by the Ir-O-Ir bond angle for a given Ir-Ir distance. Previous studies revealed that the ferromagnetic (FM) Kitaev term is suppressed when the Ir-O-Ir bond angle becomes smaller. Reference  $[19]$  found also that the  $\Gamma$  term is substantially enhanced as the bond angle is reduced. Since the reduced bond angle corresponds to the increased Ir-O distance and the reduced *p*-*d*-hopping amplitude, their results are consistent with our finding of reduced Kitaev and enhanced  $\Gamma$  terms under pressure.

Such pressure-induced effects on the anisotropic exchange interactions would manifest in the anisotropy of the magnetic susceptibility. For example, high-temperature expansion of Eq. [\(1\)](#page-0-0) yields the anisotropic Curie-Weiss temperatures, which satisfy  $\theta_{\mathbf{a}}^{\text{CW}} - \theta_{\mathbf{c}}^{\text{CW}} \simeq 2|\Gamma_Z|$  and  $\theta_{\mathbf{a}}^{\text{CW}} + \theta_{\mathbf{c}}^{\text{CW}} - 2\theta_{\mathbf{b}}^{\text{CW}} \simeq$  $2(K_Z - K_X)$ , where  $\theta_{\mathbf{a},\mathbf{b},\mathbf{c}}^{\mathbf{CW}}$  are the Curie-Weiss temperatures (multiplied by  $k_B$ ) with external field parallel to the  $a,b,c$ axes, respectively. Hence the change in anisotropic exchange interactions and bond anisotropy of the Kitaev term under pres-

TABLE I. Table of optimized lattice parameters and internal coordinates of pressurized *β*-Li<sub>2</sub>IrO<sub>3</sub> with *Fddd* (SG. 70, origin choice 2) space-group symmetry, where the internal coordinates for each inequivalent site are  $(1/8,1/8,z)$  for Ir and Li1/2,  $(x,1/8,1/8)$ for O1, and  $(x, y, z)$  for O2. *V* and  $V_0$  denote the cell volume for the optimized structure at the given pressure and that of an experimental one at the ambient pressure, respectively. In addition, Ir-Ir and Ir-O bond lengths and Ir-O-Ir bond angles in each NN bond are shown below.

	P(GPa) $V/V_0$	$-3.9$ 1.03	$-1.1$ 1.00	2.1 0.97	5.9 0.94	10.2 0.91
	a	5.964	5.908	5.848	5.790	5.729
	b	8.545	8.440	8.340	8.238	8.137
	$c(\AA)$	18.037	17.891	17.747	17.603	17.463
Ir $(16g)$	$\overline{z}$	0.7085	0.7085	0.7086	0.7087	0.7088
Li1 (16g)	$\overline{z}$	0.0441	0.0448	0.0454	0.0458	0.0460
Li2(16g)	$\overline{z}$	0.8769	0.8775	0.8779	0.8781	0.8781
O1 (16e)	$\boldsymbol{x}$	0.8561	0.8588	0.8614	0.8637	0.8658
O2(32h)	$\boldsymbol{x}$	0.6335	0.6320	0.6305	0.6289	0.6271
	y	0.3631	0.3654	0.3676	0.3698	0.3719
	Z.	0.0378	0.0384	0.0390	0.0397	0.0403
$d_{\rm Ir-Ir}$	Ζ	3.011	2.988	2.967	2.946	2.928
$\rm(\AA)$	X	3.005	2.973	2.940	2.907	2.874
$d_{\text{Ir}-\Omega}$	Z	2.041	2.035	2.028	2.020	2.012
(averaged)	X	2.034	2.029	2.023	2.017	2.010
$\theta_{\text{Ir}-\text{O-Ir}}$	Ζ	95.06	94.50	94.04	93.64	93.39
$(\text{deg})$	Χ	95.23	94.23	93.25	92.26	91.24

<span id="page-4-0"></span>sure can be detected from the anisotropy of high-temperature susceptibility data.

# **VI. DISCUSSION AND OUTLOOK**

In iridates with honeycomb or hyperhoneycomb lattices, the strong spin-orbit coupling and edge-sharing oxygen octahedra structure conspire to generate the celebrated Kitaev interaction, which provides magnetic frustration and an exactly solvable quantum spin-liquid ground state. Such physics has been one of the main driving forces for research on quantum spin-liquid phases in this class of materials. It is in contrast to a more conventional paradigm, where further-neighbor exchange interactions are used to engineer magnetic frustration in bipartite lattices such as the honeycomb or hyperhoneycomb lattices. In this work, we ask the question of whether the bond-dependent interaction or further-neighbor interaction is mainly responsible for the suppression of magnetic order or appearance of a correlated quantum paramagnetic state under high pressure as discovered in a recent experiment on pressurized  $\beta$ -Li<sub>2</sub>IrO<sub>3</sub> [\[39\]](#page-6-0). Remarkably, our analyses of *ab initio* computations with structure optimizations and strongcoupling expansion strongly suggest that the bond-dependent symmetric anisotropic interaction, which is *distinct* from the Kitaev interaction, is the dominant player in the magnetic frustration. Previous studies of the SA interaction on the honeycomb and hyperhoneycomb lattices have shown that there exists a macroscopically degenerate manifold of classical ground states [\[31](#page-5-0)[,38\]](#page-6-0). Hence it is conceivable that the quantum version of this model may support the emergence of a new kind of quantum spin-liquid ground state. This would be an excellent topic for future studies.

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TABLE II. A subset of Ir  $t_{2g}$  hopping terms  $\hat{T}_{ij}$  as representatives of each hopping channel up to third NN, where  $\mathcal{H}_{hop} = \sum_{ij} C_i^{\dagger} \cdot \hat{T}_{ij} \cdot C_j$ , and  $\{C^{\dagger}, C\}$  are the creation and annihilation operators for  $t_{2g}$  states, respectively. *d* is the approximate distance between Ir and O. Other hopping terms can be recovered by applying  $\hat{T}_{ji} = \hat{T}_{ij}^{\dagger}$ ,  $C_2^{\text{a,b,c}}$  rotations, and inversion. Values are in meV units, and terms smaller than 0.5 meV are not shown.



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## **APPENDIX A: OPTIMIZED CRYSTAL STRUCTURES UNDER THE PRESSURE**

As mentioned above, structure optimizations are carried out in the primitive unit cell without enforcing any symmetry constraints. However, the optimized structures show practically no deviation from the original *Fddd* space-group symmetry. The angles between the orthorhombic Bravais lattice vectors in the optimized structures do not deviate from the right angle  $(|\delta \theta| < 0.0004°)$ . The FINDSYM package [\[52\]](#page-6-0) is employed for refining the optimized structures, and the difference of internal coordinates between structures before and after the refinement is smaller than  $0.0002 \text{ Å}$  for each site. Therefore, we conclude that the optimized structure under pressure remains in *Fddd* symmetry without any symmetry lowering. The refined structures are presented in Table [I.](#page-3-0)

The pressure dependence of **b** deserves a comment: In Fig. [2](#page-1-0) in the main text, only the pressure dependence of **a** is presented. The compression of **b** is similar to that of **a**, where  $a/a_0$  and **b**/ **are 0.968 and 0.962, respectively, at**  $P = 10.2$  **GPa. <b>b** is slightly more compressed than **a**, but since the compression of **a** and **b** is similar and significantly larger than that of **c**, we present **a** as the representative.

# **APPENDIX B:** *t***2***<sup>g</sup>* **HOPPING INTEGRALS**

The Ir  $t_{2g}$  hopping integrals for two structures at  $P = -3.9$ and 10.2 GPa are shown in Table [II.](#page-4-0) The values are calculated without including  $U_{\text{eff}}$  and magnetism for each optimized structure.

- [1] L. Balents, [Nature \(London\)](https://doi.org/10.1038/nature08917) **[464](https://doi.org/10.1038/nature08917)**, [199](https://doi.org/10.1038/nature08917) [\(2010\)](https://doi.org/10.1038/nature08917).
- [2] A. Kitaev, [Ann. Phys.](https://doi.org/10.1016/j.aop.2005.10.005) **[321](https://doi.org/10.1016/j.aop.2005.10.005)**, [2](https://doi.org/10.1016/j.aop.2005.10.005) [\(2006\)](https://doi.org/10.1016/j.aop.2005.10.005).
- [3] [W. Witczak-Krempa, G. Chen, Y. B. Kim, and L. Balents,](https://doi.org/10.1146/annurev-conmatphys-020911-125138) Annu. Rev. Condens. Matter Phys. **[5](https://doi.org/10.1146/annurev-conmatphys-020911-125138)**, [57](https://doi.org/10.1146/annurev-conmatphys-020911-125138) [\(2014\)](https://doi.org/10.1146/annurev-conmatphys-020911-125138).
- [4] [J. G. Rau, E. K.-H. Lee, and H.-Y. Kee,](https://doi.org/10.1146/annurev-conmatphys-031115-011319) Annu. Rev. Condens. Matter Phys. **[7](https://doi.org/10.1146/annurev-conmatphys-031115-011319)**, [195](https://doi.org/10.1146/annurev-conmatphys-031115-011319) [\(2016\)](https://doi.org/10.1146/annurev-conmatphys-031115-011319).
- [5] [R. Schaffer, E. K.-H. Lee, B.-J. Yang, and Y. B. Kim,](https://doi.org/10.1088/0034-4885/79/9/094504) Rep. Prog. Phys. **[79](https://doi.org/10.1088/0034-4885/79/9/094504)**, [094504](https://doi.org/10.1088/0034-4885/79/9/094504) [\(2016\)](https://doi.org/10.1088/0034-4885/79/9/094504).
- [6] G. Jackeli and G. Khaliullin, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.102.017205) **[102](https://doi.org/10.1103/PhysRevLett.102.017205)**, [017205](https://doi.org/10.1103/PhysRevLett.102.017205) [\(2009\)](https://doi.org/10.1103/PhysRevLett.102.017205).
- [7] S. K. Choi, R. Coldea, A. N. Kolmogorov, T. Lancaster, I. I. Mazin, S. J. Blundell, P. G. Radaelli, Y. Singh, P. Gegenwart, K. R. Choi, S.-W. Cheong, P. J. Baker, C. Stock, and J. Taylor, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.108.127204) **[108](https://doi.org/10.1103/PhysRevLett.108.127204)**, [127204](https://doi.org/10.1103/PhysRevLett.108.127204) [\(2012\)](https://doi.org/10.1103/PhysRevLett.108.127204).
- [8] F. Ye, S. Chi, H. Cao, B. C. Chakoumakos, J. A. Fernandez-Baca, [R. Custelcean, T. F. Qi, O. B. Korneta, and G. Cao,](https://doi.org/10.1103/PhysRevB.85.180403) Phys. Rev. B **[85](https://doi.org/10.1103/PhysRevB.85.180403)**, [180403](https://doi.org/10.1103/PhysRevB.85.180403) [\(2012\)](https://doi.org/10.1103/PhysRevB.85.180403).
- [9] Y. Singh, S. Manni, J. Reuther, T. Berlijn, R. Thomale, W. Ku, S. Trebst, and P. Gegenwart, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.108.127203) **[108](https://doi.org/10.1103/PhysRevLett.108.127203)**, [127203](https://doi.org/10.1103/PhysRevLett.108.127203) [\(2012\)](https://doi.org/10.1103/PhysRevLett.108.127203).
- [10] I. I. Mazin, H. O. Jeschke, K. Foyevtsova, R. Valentí, and D. I. Khomskii, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.109.197201) **[109](https://doi.org/10.1103/PhysRevLett.109.197201)**, [197201](https://doi.org/10.1103/PhysRevLett.109.197201) [\(2012\)](https://doi.org/10.1103/PhysRevLett.109.197201).
- [11] J. Chaloupka, G. Jackeli, and G. Khaliullin, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.110.097204) **[110](https://doi.org/10.1103/PhysRevLett.110.097204)**, [097204](https://doi.org/10.1103/PhysRevLett.110.097204) [\(2013\)](https://doi.org/10.1103/PhysRevLett.110.097204).
- [12] [Y. Yamaji, Y. Nomura, M. Kurita, R. Arita, and M. Imada,](https://doi.org/10.1103/PhysRevLett.113.107201) *Phys.* Rev. Lett. **[113](https://doi.org/10.1103/PhysRevLett.113.107201)**, [107201](https://doi.org/10.1103/PhysRevLett.113.107201) [\(2014\)](https://doi.org/10.1103/PhysRevLett.113.107201).
- [13] V. M. Katukuri, S. Nishimoto, V. Yushankhai, A. Stoyanova, H. Kandpal, S. Choi, R. Coldea, I. Rousochatzakis, L. Hozoi, and J. van den Brink, [New J. Phys.](https://doi.org/10.1088/1367-2630/16/1/013056) **[16](https://doi.org/10.1088/1367-2630/16/1/013056)**, [013056](https://doi.org/10.1088/1367-2630/16/1/013056) [\(2014\)](https://doi.org/10.1088/1367-2630/16/1/013056).
- [14] S. Hwan Chun, J.-W. Kim, J. Kim, H. Zheng, C. C. Stoumpos, C. D. Malliakas, J. F. Mitchell, K. Mehlawat, Y. Singh, Y. Choi, T. Gog, A. Al-Zein, M. M. Sala, M. Krisch, J. Chaloupka, G. Jackeli, G. Khaliullin, and B. J. Kim, [Nat. Phys.](https://doi.org/10.1038/nphys3322) **[11](https://doi.org/10.1038/nphys3322)**, [462](https://doi.org/10.1038/nphys3322) [\(2015\)](https://doi.org/10.1038/nphys3322).
- [15] I. Rousochatzakis, J. Reuther, R. Thomale, S. Rachel, and N. B. Perkins, [Phys. Rev. X](https://doi.org/10.1103/PhysRevX.5.041035) **[5](https://doi.org/10.1103/PhysRevX.5.041035)**, [041035](https://doi.org/10.1103/PhysRevX.5.041035) [\(2015\)](https://doi.org/10.1103/PhysRevX.5.041035).
- [16] J. Knolle, G.-W. Chern, D. L. Kovrizhin, R. Moessner, and N. B. Perkins, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.113.187201) **[113](https://doi.org/10.1103/PhysRevLett.113.187201)**, [187201](https://doi.org/10.1103/PhysRevLett.113.187201) [\(2014\)](https://doi.org/10.1103/PhysRevLett.113.187201).
- [17] K. W. Plumb, J. P. Clancy, L. J. Sandilands, V. V. Shankar, Y. F. Hu, K. S. Burch, H.-Y. Kee, and Y.-J. Kim, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.90.041112) **[90](https://doi.org/10.1103/PhysRevB.90.041112)**, [041112\(R\)](https://doi.org/10.1103/PhysRevB.90.041112) [\(2014\)](https://doi.org/10.1103/PhysRevB.90.041112).
- [18] A. Banerjee, C. A. Bridges, J.-Q. Yan, A. A. Aczel, L. Li, M. B. Stone, G. E. Granroth, M. D. Lumsden, Y. Yiu, J. Knolle, S. Bhattacharjee, D. L. Kovrizhin, R. Moessner, D. A. Tennant, D. G. Mandrus, and S. E. Nagler, [Nat. Mater.](https://doi.org/10.1038/nmat4604) **[15](https://doi.org/10.1038/nmat4604)**, [733](https://doi.org/10.1038/nmat4604) [\(2016\)](https://doi.org/10.1038/nmat4604).
- [19] S. M. Winter, Y. Li, H. O. Jeschke, and R. Valentí, *[Phys. Rev. B](https://doi.org/10.1103/PhysRevB.93.214431)* **[93](https://doi.org/10.1103/PhysRevB.93.214431)**, [214431](https://doi.org/10.1103/PhysRevB.93.214431) [\(2016\)](https://doi.org/10.1103/PhysRevB.93.214431).
- [20] S. Mandal and N. Surendran, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.79.024426) **[79](https://doi.org/10.1103/PhysRevB.79.024426)**, [024426](https://doi.org/10.1103/PhysRevB.79.024426) [\(2009\)](https://doi.org/10.1103/PhysRevB.79.024426).
- [21] [E. K.-H. Lee, R. Schaffer, S. Bhattacharjee, and Y. B. Kim,](https://doi.org/10.1103/PhysRevB.89.045117) *Phys.* Rev. B **[89](https://doi.org/10.1103/PhysRevB.89.045117)**, [045117](https://doi.org/10.1103/PhysRevB.89.045117) [\(2014\)](https://doi.org/10.1103/PhysRevB.89.045117).
- [22] [S. B. Lee, E. K.-H. Lee, A. Paramekanti, and Y. B. Kim,](https://doi.org/10.1103/PhysRevB.89.014424) Phys. Rev. B **[89](https://doi.org/10.1103/PhysRevB.89.014424)**, [014424](https://doi.org/10.1103/PhysRevB.89.014424) [\(2014\)](https://doi.org/10.1103/PhysRevB.89.014424).
- [23] T. Takayama, A. Kato, R. Dinnebier, J. Nuss, H. Kono, L. S. I. Veiga, G. Fabbris, D. Haskel, and H. Takagi, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.114.077202) **[114](https://doi.org/10.1103/PhysRevLett.114.077202)**, [077202](https://doi.org/10.1103/PhysRevLett.114.077202) [\(2015\)](https://doi.org/10.1103/PhysRevLett.114.077202).
- [24] A. Biffin, R. D. Johnson, S. Choi, F. Freund, S. Manni, A. [Bombardi, P. Manuel, P. Gegenwart, and R. Coldea,](https://doi.org/10.1103/PhysRevB.90.205116) Phys. Rev. B **[90](https://doi.org/10.1103/PhysRevB.90.205116)**, [205116](https://doi.org/10.1103/PhysRevB.90.205116) [\(2014\)](https://doi.org/10.1103/PhysRevB.90.205116).
- [25] [R. Schaffer, E. K.-H. Lee, Y.-M. Lu, and Y. B. Kim,](https://doi.org/10.1103/PhysRevLett.114.116803) *Phys. Rev.* Lett. **[114](https://doi.org/10.1103/PhysRevLett.114.116803)**, [116803](https://doi.org/10.1103/PhysRevLett.114.116803) [\(2015\)](https://doi.org/10.1103/PhysRevLett.114.116803).
- [26] H.-S. Kim, E. K.-H. Lee, and Y. B. Kim, [Europhys. Lett.](https://doi.org/10.1209/0295-5075/112/67004) **[112](https://doi.org/10.1209/0295-5075/112/67004)**, [67004](https://doi.org/10.1209/0295-5075/112/67004) [\(2015\)](https://doi.org/10.1209/0295-5075/112/67004).
- [27] V. M. Katukuri, R. Yadav, L. Hozoi, S. Nishimoto, and J. Van Den Brink, [Sci. Rep.](https://doi.org/10.1038/srep29585) **[6](https://doi.org/10.1038/srep29585)**, [29585](https://doi.org/10.1038/srep29585) [\(2016\)](https://doi.org/10.1038/srep29585).
- [28] K. O'Brien, M. Hermanns, and S. Trebst, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.93.085101) **[93](https://doi.org/10.1103/PhysRevB.93.085101)**, [085101](https://doi.org/10.1103/PhysRevB.93.085101) [\(2016\)](https://doi.org/10.1103/PhysRevB.93.085101).
- [29] K. A. Modic, T. E. Smidt, I. Kimchi, N. P. Breznay, A. Biffin, S. Choi, R. D. Johnson, R. Coldea, P. Watkins-Curry, G. T. McCandless, J. Y. Chan, F. Gandara, Z. Islam, A. Vishwanath, A. Shekhter, R. D. McDonald, and J. G. Analytis, [Nat. Commun.](https://doi.org/10.1038/ncomms5203) **[5](https://doi.org/10.1038/ncomms5203)**, [4203](https://doi.org/10.1038/ncomms5203) [\(2014\)](https://doi.org/10.1038/ncomms5203).
- [30] I. Kimchi, J. G. Analytis, and A. Vishwanath, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.90.205126) **[90](https://doi.org/10.1103/PhysRevB.90.205126)**, [205126](https://doi.org/10.1103/PhysRevB.90.205126) [\(2014\)](https://doi.org/10.1103/PhysRevB.90.205126).
- [31] E. K.-H. Lee and Y. B. Kim, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.91.064407) **[91](https://doi.org/10.1103/PhysRevB.91.064407)**, [064407](https://doi.org/10.1103/PhysRevB.91.064407) [\(2015\)](https://doi.org/10.1103/PhysRevB.91.064407).
- <span id="page-6-0"></span>[32] I. Kimchi, R. Coldea, and A. Vishwanath, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.91.245134) **[91](https://doi.org/10.1103/PhysRevB.91.245134)**, [245134](https://doi.org/10.1103/PhysRevB.91.245134) [\(2015\)](https://doi.org/10.1103/PhysRevB.91.245134).
- [33] [B. Perreault, J. Knolle, N. B. Perkins, and F. J. Burnell,](https://doi.org/10.1103/PhysRevB.92.094439) *Phys.* Rev. B **[92](https://doi.org/10.1103/PhysRevB.92.094439)**, [094439](https://doi.org/10.1103/PhysRevB.92.094439) [\(2015\)](https://doi.org/10.1103/PhysRevB.92.094439).
- [34] E. K.-H. Lee, J. G. Rau, and Y. B. Kim, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.93.184420) **[93](https://doi.org/10.1103/PhysRevB.93.184420)**, [184420](https://doi.org/10.1103/PhysRevB.93.184420) [\(2016\)](https://doi.org/10.1103/PhysRevB.93.184420).
- [35] B. J. Kim, H. Jin, S. J. Moon, J.-Y. Kim, B.-G. Park, C. S. Leem, J. Yu, T. W. Noh, C. Kim, S.-J. Oh, J.-H. Park, V. Durairaj, G. Cao, and E. Rotenberg, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.101.076402) **[101](https://doi.org/10.1103/PhysRevLett.101.076402)**, [076402](https://doi.org/10.1103/PhysRevLett.101.076402) [\(2008\)](https://doi.org/10.1103/PhysRevLett.101.076402).
- [36] B. J. Kim, H. Ohsumi, T. Komesu, S. Sakai, T. Morita, H. Takagi, and T. Arima, [Science](https://doi.org/10.1126/science.1167106) **[323](https://doi.org/10.1126/science.1167106)**, [1329](https://doi.org/10.1126/science.1167106) [\(2009\)](https://doi.org/10.1126/science.1167106).
- [37] A. Biffin, R. D. Johnson, I. Kimchi, R. Morris, A. Bombardi, J. G. Analytis, A. Vishwanath, and R. Coldea, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.113.197201) **[113](https://doi.org/10.1103/PhysRevLett.113.197201)**, [197201](https://doi.org/10.1103/PhysRevLett.113.197201) [\(2014\)](https://doi.org/10.1103/PhysRevLett.113.197201).
- [38] J. G. Rau, E. K.-H. Lee, and H.-Y. Kee, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.112.077204) **[112](https://doi.org/10.1103/PhysRevLett.112.077204)**, [077204](https://doi.org/10.1103/PhysRevLett.112.077204) [\(2014\)](https://doi.org/10.1103/PhysRevLett.112.077204).
- [39] H. Takagi (unpublished).
- [40] G. Kresse and J. Furthmüller, *[Phys. Rev. B](https://doi.org/10.1103/PhysRevB.54.11169)* [54](https://doi.org/10.1103/PhysRevB.54.11169), [11169](https://doi.org/10.1103/PhysRevB.54.11169) [\(1996\)](https://doi.org/10.1103/PhysRevB.54.11169).
- [41] G. Kresse and D. Joubert, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.59.1758) **[59](https://doi.org/10.1103/PhysRevB.59.1758)**, [1758](https://doi.org/10.1103/PhysRevB.59.1758) [\(1999\)](https://doi.org/10.1103/PhysRevB.59.1758).
- [42] S. L. Dudarev, G. A. Botton, S. Y. Savrasov, C. J. Humphreys, and A. P. Sutton, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.57.1505) **[57](https://doi.org/10.1103/PhysRevB.57.1505)**, [1505](https://doi.org/10.1103/PhysRevB.57.1505) [\(1998\)](https://doi.org/10.1103/PhysRevB.57.1505).
- 
- [43] J. P. Perdew, A. Ruzsinszky, G. I. Csonka, O. A. Vydrov, G. E. [Scuseria, L. A. Constantin, X. Zhou, and K. Burke,](https://doi.org/10.1103/PhysRevLett.100.136406) Phys. Rev. Lett. **[100](https://doi.org/10.1103/PhysRevLett.100.136406)**, [136406](https://doi.org/10.1103/PhysRevLett.100.136406) [\(2008\)](https://doi.org/10.1103/PhysRevLett.100.136406).
- [44] N. Marzari and D. Vanderbilt, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.56.12847) **[56](https://doi.org/10.1103/PhysRevB.56.12847)**, [12847](https://doi.org/10.1103/PhysRevB.56.12847) [\(1997\)](https://doi.org/10.1103/PhysRevB.56.12847).
- [45] I. Souza, N. Marzari, and D. Vanderbilt, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.65.035109) **[65](https://doi.org/10.1103/PhysRevB.65.035109)**, [035109](https://doi.org/10.1103/PhysRevB.65.035109) [\(2001\)](https://doi.org/10.1103/PhysRevB.65.035109).
- [46] A. A. Mostofi, J. R. Yates, Y.-S. Lee, I. Souza, D. Vanderbilt, and N. Marzari, [Comput. Phys. Commun.](https://doi.org/10.1016/j.cpc.2007.11.016) **[178](https://doi.org/10.1016/j.cpc.2007.11.016)**, [685](https://doi.org/10.1016/j.cpc.2007.11.016) [\(2008\)](https://doi.org/10.1016/j.cpc.2007.11.016).
- [47] K. Momma and F. Izumi, [J. Appl. Crystallogr.](https://doi.org/10.1107/S0021889811038970) **[44](https://doi.org/10.1107/S0021889811038970)**, [1272](https://doi.org/10.1107/S0021889811038970) [\(2011\)](https://doi.org/10.1107/S0021889811038970).
- [48] H.-S. Kim and H.-Y. Kee, [Phys. Rev. B](https://doi.org/10.1103/PhysRevB.93.155143) **[93](https://doi.org/10.1103/PhysRevB.93.155143)**, [155143](https://doi.org/10.1103/PhysRevB.93.155143) [\(2016\)](https://doi.org/10.1103/PhysRevB.93.155143).
- [49] J. G. Rau and H.-Y. Kee, [arXiv:1408.4811.](http://arxiv.org/abs/arXiv:1408.4811)
- [50] We checked that the alternative expression for the exchange interactions, presented in Ref. [\[19\]](#page-5-0), yields consistent results with ours shown in Fig. [4,](#page-3-0) except for small quantitative difference as to slightly enhanced bond anisotropy.
- [51] S. Nishimoto, V. M. Katukuri, V. Yushankhai, H. Stoll, U. K. Rößler, L. Hozoi, I. Rousochatzakis, and J. van den Brink, Nat. Commun. **[7](https://doi.org/10.1038/ncomms10273)**, [10273](https://doi.org/10.1038/ncomms10273) [\(2016\)](https://doi.org/10.1038/ncomms10273).
- [52] H. T. Stokes and D. M. Hatch, [J. Appl. Crystallogr.](https://doi.org/10.1107/S0021889804031528) **[38](https://doi.org/10.1107/S0021889804031528)**, [237](https://doi.org/10.1107/S0021889804031528) [\(2005\)](https://doi.org/10.1107/S0021889804031528).