Critical spin-flip scattering at the helimagnetic transition of MnSi

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(Received 6 March 2014; published 19 May 2014)

We report spherical neutron polarimetry (SNP) and discuss the spin-flip scattering cross sections as well as the chiral fraction *η* close to the helimagnetic transition in MnSi. For our study we have developed a miniaturized SNP device that allows fast data collection when used in small angle scattering geometry with an area detector. Critical spin-flip scattering is found to be governed by chiral paramagnons that soften on a sphere in momentum space. Carefully accounting for the incoherent spin-flip background, we find that the resulting chiral fraction *η* decreases gradually above the helimagnetic transition reflecting a strongly renormalized chiral correlation length with a temperature dependence in excellent quantitative agreement with the Brazovskii theory for a fluctuation-induced first order transition.

DOI: [10.1103/PhysRevB.89.180408](http://dx.doi.org/10.1103/PhysRevB.89.180408) PACS number(s): 75*.*25*.*−j*,* 75*.*10*.*−b*,* 75*.*50*.*−y

It has long been established that B20 transition metal compounds support a well-understood hierarchy of energy scales [\[1\]](#page-3-0) comprising of decreasing strength ferromagnetic exchange, Dzyaloshinsky-Moriya spin-orbit interactions, and, finally, magnetic anisotropies due to higher-order spin-orbit coupling. In turn, zero temperature spontaneous magnetic order in these materials appears essentially ferromagnetic on short distances with a long wavelength homochiral twist on intermediate distances that propagates along directions favored by the magnetic anisotropies on the largest length scales.

In recent years the nature of the associated helimagnetic transition at T_c , which makes contact with areas ranging from nuclear matter over quantum Hall physics to soft matter systems, has been the topic of heated scientific controversy since critical helimagnetic spin fluctuations are able to drive the transition first order. However, considerable differences exist as to the proposed character of these fluctuations. Based on a minimal description taking only into account the three scales mentioned above two scenarios may be distinguished. First, according to Bak and Jensen [\[2\]](#page-3-0) when the magnetic anisotropies are sufficiently strong, anisotropic critical paramagnons develop along the easy axis already at $T > T_c$. Second, in the opposite limit when the magnetic anisotropies are weak, critical spin fluctuations soften isotropically on the surface of a sphere in momentum space giving rise to a fluctuation-driven first order transition in the spirit of a proposal by Brazovskii [\[3\]](#page-3-0).

A third and completely different scenario by Rößler, Bogdanov, and Pfleiderer [\[4\]](#page-3-0) has generated great interest, which requires, however, an additional phenomenological parameter beyond the minimal model. In this model the formation of a skyrmion liquid phase at zero magnetic field between the paramagnetic and helimagnetic state is predicted, which implies an additional phase transition at a temperature T_{sk} > T_c . The formation of skyrmions at zero magnetic field is of great fundamental interest in the context of continuum models of matter as well as possible applications in spintronics.

Early experimental studies of the electrical resistivity [\[5\]](#page-3-0), specific heat [\[6\]](#page-3-0), thermal expansion [\[7,8\]](#page-3-0), ultrasound attenuation [\[9\]](#page-3-0), and neutron scattering [\[10\]](#page-3-0) in MnSi, as the most extensively studied B20 compound, were interpreted in terms of the scenario by Bak and Jensen. Comprehensive elastic neutron scattering, demonstrating critical fluctuations on the surface of a sphere, together with specific heat, susceptibility, and magnetization measurements, recently changed this view, providing quantitatively consistent evidence of a fluctuation-driven first order transition as proposed by Brazovskii [\[11–14\]](#page-4-0), see also Ref. [\[15,16\]](#page-4-0) for related work on $Cu₂OSeO₃$.

Following the discovery of a skyrmion lattice phase in small applied magnetic fields just below T_c [\[17–20\]](#page-4-0), several authors have argued that the specific heat and susceptibility provide evidence for further complex phases including a skyrmion liquid phase at zero field $[21-25]$. Most importantly, the observation of a chiral fraction $\eta \approx 1$ up to at least 1 K above T_c in a seminal SNP study in MnSi by Pappas *et al.* [\[26,27\]](#page-4-0) to date appears to provide the only microscopic evidence supporting the formation of skyrmions in zero magnetic field in a bulk material.

However, as explained in our Rapid Communication *η* provides a measure of the asymmetry of magnetic spin-flip scattering, assuming extreme values $\eta = \pm 1$ if one of the spin-flip scattering processes, i.e., \uparrow to \downarrow or vice versa, is forbidden. Hence *η* shows to what extent a magnetic system is homochiral. In contrast, by definition *η* is neither a direct measure of the topological winding as the defining new property of the skyrmion liquid, nor of the phase relationship of the underlying multi-*q* modulations determined recently in the skyrmion lattice phase in MnSi [\[28\]](#page-4-0). Moreover, to the best of our knowledge, a theoretical link between *η* and the formation of a skyrmion liquid phase has not been reported either.

Motivated by the broad interest in the helimagnetic transition of chiral magnets and the specific importance attributed in previous work of *η* we have revisited the entire issue from a more general point of view in an experimental and theoretical study of the critical spin-flip scattering in MnSi. We thereby do not find any evidence suggesting an additional phase transition

FIG. 1. (Color online) Schematic depiction of the miniaturized SNP device. (a) Schematic overview of the complete setup with cryostat, polarizer (P), analyzer (A), and detector (D). (b) Close-up view of the SNP device, as composed of the coil bodies (CB) with their neutron window (NW). The coils are surrounded by a μ -metal yoke (MM). (c) Orientation of the precession coils (blue arrows), local magnetic field (red arrow), and sample (S).

above T_c . As our main conclusion, our SNP results are in excellent *quantitative* agreement with the minimal model of a fluctuation-induced first order transition as predicted by Brazovskii, establishing also quantitative consistency with previous specific heat, magnetization, susceptibility, and elastic unpolarized neutron scattering studies $[11-14]$.

Carefully considering the experimental requirements for going beyond previous SNP studies, presented in detail below, revealed as the most prominent aspect the need to track incoherent signal contributions. To meet these requirements we have developed a versatile miniaturized SNP device [\[29\]](#page-4-0), which in its present version offers great flexibility at scattering angles up to 15◦. In particular, as opposed to the large size of SNP devices such as CryoPAD or MuPAD [\[30,31\]](#page-4-0) our entire setup (diameter 50 mm; height 120 mm) is integrated into a normal sample stick fitting a standard pulse-tube cooler. In turn this reduces the time required for setting up our SNP device to the time needed for a conventional sample change. Moreover, when combined with an area detector fast data collection at various sample orientations and momentum transfers are readily possible.

Shown in Fig. $1(a)$ is a schematic depiction of the miniaturized SNP setup developed for our study. Pairs of crossed precession coils (PC) before and after the sample generate Larmor precessions of the neutron polarization that permit to rotate the polarization of an incoming and scattered neutron beam in any arbitrary direction. The precession coils are wound on coil bodies (CB) with a large neutron window (NW), see Fig. $1(b)$. As its main advantage parasitic rotations of the polarization are minimized by the very compact geometry and the μ -metal yokes (MM) around the precession coils, which short circuit both external fields and the precession fields.

The measurements reported in this Rapid Communication were performed on the beam-line RESEDA [\[32\]](#page-4-0) at the Heinz Maier-Leibnitz Zentrum of the Technische Universitat¨

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München. Neutrons were polarized with a cavity providing a polarization of 95% and analyzed with a bender at an efficiency of 98%. Data were recorded at a neutron wavelength $\lambda = 4.5$ Å with a wavelength spread $\Delta\lambda/\lambda = 0.16$. Using a CASCADE [\[33\]](#page-4-0) area detector (PSD) with 200 mm \times 200 mm active area (128×128 pixels) at a distance to the sample of 1596 mm our setup corresponded effectively to small angle neutron scattering. The FWHM of the resolution and the wavelength band $\sigma_{\text{fCol}} = 3.6 \times 10^{-3} \text{ Å}^{-1}$ and $\sigma_{\text{W}} = 2.6 \times 10^{-3} \text{ Å}^{-1}$ 10^{-3} Å⁻¹, respectively, as considered at the wave vector of the helical modulation, $k = 0.039$ Å of MnSi, resulted in a momentum uncertainty $\Delta Q/Q \approx 10\%$.

For our study we used the MnSi single crystal investigated in Ref. [\[11\]](#page-4-0). The sample was oriented by x-ray Laue diffraction such that two $\langle 111 \rangle$ axes, in the following denoted by $\mathbf{k}_1 \parallel$ $[1\overline{1}\overline{1}]$ and $\mathbf{k}_2 \parallel [1\overline{1}1]$, were orthogonal to the neutron beam $\hat{n} \parallel [110]$, see Fig. [2\(a\).](#page-2-0) The combination of the large window of the PCs with a PSD detector and the possibility for rotations of the sample with respect to its vertical $[1\overline{1}2]$ axis by an angle = 0◦*,* 3◦*,* 25◦, see also Fig. 1(c), allowed us to track various points in reciprocal space going well beyond previous work.

For a physically transparent theoretical account of the spin-flip scattering in chiral magnets it is helpful to begin with the spin-resolved energy-integrated neutron scattering cross section $\sigma_{\hat{e}_{out}, \hat{e}_{in}}(\mathbf{Q}) \equiv \frac{d\sigma}{d\Omega}$, with momentum transfer **Q**, which describes the change of the spin eigenstate of the neutron from $|\hat{e}_{in}\rangle$ to $|\hat{e}_{out}\rangle$ with $\vec{\sigma}|\hat{e}_{\alpha}\rangle = \hat{e}_{\alpha}|\hat{e}_{\alpha}\rangle$. It comprises, in general, a nuclear and a magnetic contribution, as well as a nuclear-magnetic interference term [\[34,35\]](#page-4-0). The spin-polarization dependence of the cross section $\sigma(\mathbf{Q}, \hat{e}_{in}) = \sum_{\tau = \pm 1} \sigma_{\tau \hat{e}_{out}, \hat{e}_{in}}(\mathbf{Q})$ may then be resolved, when scattering a polarized neutron beam. The magnetic contribution to $\sigma(Q,\hat{e}_{in})$ consists thereby of a symmetric and an antisymmetric part $\sigma_{\text{mag}}(\mathbf{Q},\hat{e}_{\text{in}})$ = $\sigma_{\text{mag}}^S(\mathbf{Q}) + (\hat{Q}\hat{e}_{\text{in}})\sigma_{\text{mag}}^A(\mathbf{Q})$, where the latter is weighted by the scalar product $\hat{Q}\hat{e}_{in}$ with $\hat{Q} = Q/Q$. The so-called *chiral fraction η* is finally defined as $\eta(Q) = \sigma_{\text{mag}}^A(Q)/\sigma_{\text{mag}}^S(Q)$, measuring the chirality of the magnetic scattering.

Finite values of σ_{mag}^A (and thus η) originate in the polarization-dependent scattering characteristics of chiral magnetic systems like MnSi. In the helimagnetically ordered phase of MnSi, the magnetization may be described as $\vec{M}(\mathbf{r}) = M_{\text{hel}}(\hat{e}_1 \cos \mathbf{kr} + \hat{e}_2 \sin \mathbf{kr})$, with wave vector **k** and orthonormal unit vectors $\hat{e}_i \hat{e}_j = \delta_{ij}$, with $i, j = 1, 2$ orthogonal to **k**; for a left-handed helix $\hat{e}_1 \times \hat{e}_2 = -\hat{k}$. The magnetic Bragg scattering by a magnetic helix possesses equal magnitudes of symmetric and antisymmetric contributions $\sigma_{\text{mag}}^S(\mathbf{k})|_{\text{Bragg}} =$ $±\sigma_{mag}^{A}(\mathbf{k})|_{\text{Bragg}}$, where the sign indicates the handedness of the magnetic helix with $\eta_{\text{Bragg}}(\mathbf{k}) = 1$ being left and $\eta_{\text{Bragg}}(\mathbf{k}) =$ −1 being right handed. For left-handed helimagnetic order we thus expect $\sigma_{\text{mag}}(\mathbf{k}, \hat{e}_{\text{in}})|_{\text{Bragg}} \propto [1 + (\hat{k}\hat{e}_{\text{in}})] = 2 \cos^2(\delta\Omega/2),$ i.e., a sinusoidal dependence as a function of the angle $\delta\Omega$ enclosed by \hat{k} and \hat{e}_{in} .

In order to confirm the proper functioning of our SNP device this dependence was experimentally verified by rotating the polarization \hat{e}_{in} of the incident neutron in the plane perpendicular to the beam as depicted in Fig. $2(b)$, where the orientation is denoted by the angle $\Omega = \angle(\hat{e}_{in}, \mathbf{k}_1)$. Well below *T*^c weak cubic anisotropies align the helimagnetic wave vector

FIG. 2. (Color online) Operating principle of MiniMuPAD demonstrating the spin-selective Bragg scattering off helimagnetic order in MnSi. (a) Wave vectors \mathbf{k}_1 and \mathbf{k}_2 of two different helimagnetic domains are within the plane orthogonal to the neutron beam \hat{n} . (b) The angle Ω defines the relative orientation of the spin polarization of the incoming neutrons \hat{e}_{in} orthogonal to \hat{n} . (c)–(f) SANS patterns obtained for specific polarizations \hat{e}_{in} . (g) Intensity of magnetic Bragg peaks at $\pm \mathbf{k}_{1,2}$ as a function of \hat{e}_{in} , i.e., Ω , where $\Omega_1 = \arccos \frac{1}{3}$.

 \bf{k} along crystallographic $\langle 111 \rangle$ directions. In turn the intensity displays extrema whenever the polarization \hat{e}_{in} aligns with a wave vector of a helimagnetic domain $\pm \mathbf{k}_{1,2}$ for $\Omega_0 = 0$, $\Omega_1 =$ arccos $\frac{1}{3}$, $\Omega_2 = \pi$, and $\Omega_3 = \Omega_1 + \pi$. The associated intensity patterns at the PSD are shown in Figs. $2(c)$ through $2(f)$. Apart from a constant background, the variation of the intensity as a function of $\delta \Omega = \Omega - \Omega_n$, with $n = 0, 1, 2, 3$, shown in Fig. $2(g)$, may be fitted in perfect agreement with chiral Bragg scattering, underscoring the precision of the alignment of \hat{e}_{in} . This suggests, in particular, a negligible interference between magnetic and nuclear contributions. Differences of the maximum peak intensity are thereby due to a small

TABLE I. Polarization matrix in the helimagnetically ordered phase of MnSi at 6.5 K determined experimentally with our miniaturized SNP device.

$\mathbb{P}_{\alpha,\beta}$	$\beta = x$		$\beta = v$			$\beta = z$	
	$\alpha = x -0.9660 \pm 0.0004 -0.9692 \pm 0.0006 -0.9664 \pm 0.0006$ $\alpha = v$ 0.0115 ± 0.0013 0.0363 ± 0.0018 0.0321 ± 0.0018 $\alpha = z$ 0.0045 ± 0.0013 -0.079 ± 0.0018 0.0182 ± 0.0018						

misalignment of the **k**² crystallographic axis with respect to the SNP setup.

Using in addition an analyzer the fully spin-resolved scattering in the helical state may be discussed, where the polarization matrix is defined as

$$
\mathbb{P}_{\alpha,\beta}(\mathbf{Q}) = \frac{\sigma_{\hat{e}_{\alpha},\hat{e}_{\beta}}(\mathbf{Q}) - \sigma_{-\hat{e}_{\alpha},\hat{e}_{\beta}}(\mathbf{Q})}{\sigma_{\hat{e}_{\alpha},\hat{e}_{\beta}}(\mathbf{Q}) + \sigma_{-\hat{e}_{\alpha},\hat{e}_{\beta}}(\mathbf{Q})}
$$
(1)

and \hat{e}_{α} , with $\alpha = x, y, z$, forms a right-handed triad. Far below *T*^c Bragg scattering prevails close to the magnetic wave vector **k** thus allowing us to approximate $\sigma_{\hat{e}_{\text{out}}, \hat{e}_{\text{in}}}(\mathbf{k}) \approx \sigma_{\hat{e}_{\text{out}}, \hat{e}_{\text{in}}}^{\text{mag}}(\mathbf{k})|_{\text{Bragg}}.$ In this case the polarization matrix simplifies to $\mathbb{P}_{x,\beta}(\mathbf{k}) \approx$ -1 for all $\beta = x, y, z$ and zero otherwise, where we assumed $\hat{k} = \hat{e}_x$ without loss of generality. The experimental values for $\mathbb{P}_{\alpha,\beta}(\mathbf{k}_1)$ recorded at 6.5 K accounting for the analyzer efficiency are summarized in Table I. The deviation by ∼4% from purely magnetic Bragg scattering may be attributed to the flip efficiency of the precession coils and nonmagnetic scattering by the coating of the precession coils and aluminum in the beam.

This brings us finally to the spin-resolved scattering close to and above the helimagnetic transition at $T_c \approx 29$ K. Here the spin polarization of the in- and outgoing neutron beam were longitudinal to the transferred momentum $\sigma_{\tau_{\text{out}},\tau_{\text{in}}}^{\parallel}(\mathbf{Q}) =$ $\sigma_{\tau_{\text{out}},\hat{\mathcal{Q}},\tau_{\text{in}}\hat{\mathcal{Q}}}(\mathbf{Q})$, with $\tau_{\text{out}},\tau_{\text{in}} \in \{+1,-1\} \equiv \{\uparrow,\downarrow\}$. The spin-flip scattering $\sigma_{\pm,\mp}^{\parallel}(\mathbf{Q}) = \sigma_{\text{inc}}^S(\mathbf{Q}) + \sigma_{\text{mag}}^S(\mathbf{Q}) \mp \sigma_{\text{mag}}^A(\mathbf{Q})$ is then governed by critical chiral magnetism apart from an incoherent spin-flip background contribution σ_{inc}^S . From the theory of chiral magnets [\[36\]](#page-4-0) one expects the spin-flip scattering to assume the following simple form for $T > T_c$:

$$
\sigma_{\text{mag}}^{S}(\mathbf{Q}) \mp \sigma_{\text{mag}}^{A}(\mathbf{Q}) = \frac{\mathcal{A}k_{B}T}{(|\mathbf{Q}| \pm k)^{2} + \kappa^{2}(T)},
$$
 (2)

where k_B is the Boltzmann constant and A is a constant that depends on the form factor of MnSi. This expression applies to chiral systems tending to develop a left-handed helix with a pitch vector $k > 0$. The inverse correlation length $\kappa(T)$ represents the point of contact with the different theoretical proposals of the helimagnetic transition that motivated our study. In particular, for very weak cubic anisotropies in a Brazovskii scenario chiral paramagnons develop isotropically and become soft on a sphere in momentum space as $\kappa(T) \to 0$, see Eq. (2) [\[11\]](#page-4-0). These chiral paramagnons effectively display a one-dimensional character resulting in strong renormalizations of *κ*(*T*) according to the cubic equation $\kappa^2 / \kappa_{\rm Gi}^2 = \frac{T - T_{\rm MF}}{2} + \kappa_0 / \kappa$. For MnSi $T_{\rm dm} \approx 30.5$ K is the expected $\frac{T_{\text{MF}}}{T_0} + \kappa_{\text{Gi}}/\kappa$. For MnSi $T_{\text{MF}} \approx 30.5$ K is the expected mean-field transition temperature, where $\tilde{T}_0 \approx 0.5$ K and $\kappa_{\rm Gi} \approx$ 0.018 \AA^{-1} is the inverse Ginzburg length [\[11\]](#page-4-0). In turn, these strongly interacting paramagnons suppress the transition by

FIG. 3. (Color online) (a) Temperature dependence of the spinflip scattering cross section of MnSi $\sigma_{\pm,\mp}^{\parallel}(\mathbf{Q})$ with $|\mathbf{Q}| = k$ close to the critical temperature $T > T_c$, for different orientations specified by the angle $\phi = 0^{\circ}$, 3°, and 25°. For $\phi = 0$, $\hat{Q} \parallel [1\overline{1}\overline{1}]$ and other orientation are obtained by an anticlockwise rotation by *φ* around the [1¹2] axis, see also Fig. [1\(c\).](#page-1-0) Data for $\phi = 25^\circ$ was corrected for transmission. (b) The same data but on a different intensity scale; the dashed line is the fitted incoherent spin-flip background value σ_{inc}^S . (c) The chiral fraction *η* of Eq. (3). The solid lines are fits to Brazovskii theory, see text, predicting a turning point of $\eta(T)$ at T^* .

 $\Delta T = T_{\text{MF}} - T_c \approx 1.5$ K driving it first order, with characteristic signatures in physical quantities around T_{MF} .

Figures $3(a)$ and $3(b)$ display the temperature dependence of the spin-flip scattering $\sigma_{\pm\mp}^{\parallel}(\mathbf{Q})$ measured on a sphere with radius $|\mathbf{Q}| = k$ for different orientations $\hat{Q} = \mathbf{Q}/Q$. Approaching T_c , the chiral magnetic ordering indeed develops isotropically resulting in a negligible dependence on Q . Here σ_{-+}^{\parallel} reflects the strong *T* dependence of *κ* close to *T_c*, while σ_{+-}^{\parallel} is barely temperature dependent as it is suppressed by the

- [1] L. D. Landau and E. M. Lifshitz, *Course of Theoretical Physics, Vol. 8* (Pergamon, New York, 1980).
- [2] P. Bak and M. Jensen, [J. Phys. C](http://dx.doi.org/10.1088/0022-3719/13/31/002) **[13](http://dx.doi.org/10.1088/0022-3719/13/31/002)**, [L881](http://dx.doi.org/10.1088/0022-3719/13/31/002) [\(1980\)](http://dx.doi.org/10.1088/0022-3719/13/31/002).
- [3] S. A. Brazovskii, Sov. Phys. JETP **41**, 85 (1975).
- [4] U. Rößler, A. Bogdanov, and C. Pfleiderer, [Nature \(London\)](http://dx.doi.org/10.1038/nature05056) **[442](http://dx.doi.org/10.1038/nature05056)**, [797](http://dx.doi.org/10.1038/nature05056) [\(2006\)](http://dx.doi.org/10.1038/nature05056).
- [5] A. E. Petrova, E. D. Bauer, V. Krasnorussky, and S. M. Stishov, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.74.092401) **[74](http://dx.doi.org/10.1103/PhysRevB.74.092401)**, [092401](http://dx.doi.org/10.1103/PhysRevB.74.092401) [\(2006\)](http://dx.doi.org/10.1103/PhysRevB.74.092401).
- [6] S. M. Stishov, A. E. Petrova, S. Khasanov, G. K. Panova, A. A. Shikov, J. C. Lashley, D. Wu, and T. A. Lograsso, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.76.052405) **[76](http://dx.doi.org/10.1103/PhysRevB.76.052405)**, [052405](http://dx.doi.org/10.1103/PhysRevB.76.052405) [\(2007\)](http://dx.doi.org/10.1103/PhysRevB.76.052405).

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additional factor $4k^2$ in the denominator of Eq. [\(2\)](#page-2-0). Using the published results for $\kappa(T)$ and $k \approx 0.039 \text{ Å}^{-1}$ obtained in the same sample [\[11\]](#page-4-0), we are left with a single fitting parameter, namely the magnitude A in Eq. [\(2\)](#page-2-0), in addition to a temperature and Q-independent incoherent background σ_{inc}^S shown as the dotted line in Fig. 3(b). We find a remarkably good fit for both cross sections as shown by the solid lines.

Subtracting the incoherent background σ_{inc}^S determined experimentally, we obtain the chiral fraction

$$
\eta \equiv \frac{\sigma_{\text{mag}}^A(\mathbf{Q})}{\sigma_{\text{mag}}^S(\mathbf{Q})}\bigg|_{|\mathbf{Q}|=k} = \frac{1}{1 + \kappa^2(T)/(2k^2)}\tag{3}
$$

shown in Fig. $3(c)$. It is essential to note that the experimental values depend sensitively on σ_{inc}^S (likewise the error bars of η derive mainly from σ_{inc}^S). Within the error bars, however, the chiral fraction is in very good agreement with the Brazovskii theory of $\kappa(T)$. In particular, $\eta(T)$ displays a characteristic point of inflection at a temperature $T^* - T_c \approx 2$ K. It is finally instructive to note that $\eta(T)$ reported by Pappas *et al.* [\[26,27\]](#page-4-0) differs substantially up to \sim 2 K above T_c as shown in Fig. 3(c). Based on the information given in Refs. [\[26,27\]](#page-4-0) we strongly suspect that this difference is due to an overestimate of σ_{inc}^S .

In conclusion, we have investigated the critical spin-flip scattering with an emphasis on the chiral fraction *η* close to the helimagnetic transition in MnSi. For our study we have developed a miniaturized, low-cost SNP device for very fast experiments at scattering angles up to 15◦. Considering carefully the importance of incoherent background scattering we find excellent quantitative agreement of the temperature dependence of the chiral fraction *η* at various sample orientations with the Brazovski scenario of a fluctuationinduced first order transition. Our study thereby provides a quantitative connection of η with elastic neutron scattering as well as the magnetization, susceptibility, and specific heat [\[11\]](#page-4-0), completing a remarkably comprehensive account in a minimal model that does not require any additional phenomenological parameters such as those needed in a recent proposal [4] for the prediction of skyrmion formation at zero magnetic field in bulk chiral magnets.

We wish to thank A. Rosch, K. Pappas, and S. Grigoriev for helpful discussions, and M. Schulz and P. Schmakat for contributions at early stages of this work. J. K. acknowledges financial support through the TUM Graduate School. Financial support through DFG TRR80, FOR960, and ERC-AdG (291079 TOPFIT) are gratefully acknowledged.

- [7] S. M. Stishov, A. E. Petrova, S. Khasanov, G. K. Panova, [A. A. Shikov, J. C. Lashley, D. Wu, and T. A. Lograsso,](http://dx.doi.org/10.1088/0953-8984/20/23/235222) J. Phys.: Condens. Matter **[20](http://dx.doi.org/10.1088/0953-8984/20/23/235222)**, [235222](http://dx.doi.org/10.1088/0953-8984/20/23/235222) [\(2008\)](http://dx.doi.org/10.1088/0953-8984/20/23/235222).
- [8] A. E. Petrova, V. N. Krasnorussky, A. A. Shikov, W. M. Yuhasz, T. A. Lograsso, J. C. Lashley, and S. M. Stishov, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.82.155124) **[82](http://dx.doi.org/10.1103/PhysRevB.82.155124)**, [155124](http://dx.doi.org/10.1103/PhysRevB.82.155124) [\(2010\)](http://dx.doi.org/10.1103/PhysRevB.82.155124).
- [9] A. E. Petrova and S. M. Stishov, [J. Phys.: Condens. Matter](http://dx.doi.org/10.1088/0953-8984/21/19/196001) **[21](http://dx.doi.org/10.1088/0953-8984/21/19/196001)**, [196001](http://dx.doi.org/10.1088/0953-8984/21/19/196001) [\(2009\)](http://dx.doi.org/10.1088/0953-8984/21/19/196001).
- [10] S. V. Grigoriev, S. V. Maleyev, E. V. Moskvin, V. A. Dyadkin, P. Fouquet, and H. Eckerlebe, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.81.144413) **[81](http://dx.doi.org/10.1103/PhysRevB.81.144413)**, [144413](http://dx.doi.org/10.1103/PhysRevB.81.144413) [\(2010\)](http://dx.doi.org/10.1103/PhysRevB.81.144413).
- [11] M. Janoschek, M. Garst, A. Bauer, P. Krautscheid, R. Georgii, P. Böni, and C. Pfleiderer, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.87.134407) [87](http://dx.doi.org/10.1103/PhysRevB.87.134407), [134407](http://dx.doi.org/10.1103/PhysRevB.87.134407) [\(2013\)](http://dx.doi.org/10.1103/PhysRevB.87.134407).
- [12] A. Bauer and C. Pfleiderer, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.85.214418) **[85](http://dx.doi.org/10.1103/PhysRevB.85.214418)**, [214418](http://dx.doi.org/10.1103/PhysRevB.85.214418) [\(2012\)](http://dx.doi.org/10.1103/PhysRevB.85.214418).
- [13] A. Bauer, M. Garst, and C. Pfleiderer, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.110.177207) **[110](http://dx.doi.org/10.1103/PhysRevLett.110.177207)**, [177207](http://dx.doi.org/10.1103/PhysRevLett.110.177207) [\(2013\)](http://dx.doi.org/10.1103/PhysRevLett.110.177207).
- [14] S. Buhrandt and L. Fritz, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.88.195137) **[88](http://dx.doi.org/10.1103/PhysRevB.88.195137)**, [195137](http://dx.doi.org/10.1103/PhysRevB.88.195137) [\(2013\)](http://dx.doi.org/10.1103/PhysRevB.88.195137).
- [15] I. Živković, J. S. White, H. M. Rønnow, K. Prša, and H. Berger, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.89.060401) **[89](http://dx.doi.org/10.1103/PhysRevB.89.060401)**, [060401](http://dx.doi.org/10.1103/PhysRevB.89.060401) [\(2014\)](http://dx.doi.org/10.1103/PhysRevB.89.060401).
- [16] T. Adams, A. Chacon, M. Wagner, A. Bauer, G. Brandl, [B. Pedersen, H. Berger, P. Lemmens, and C. Pfleiderer,](http://dx.doi.org/10.1103/PhysRevLett.108.237204) Phys. Rev. Lett. **[108](http://dx.doi.org/10.1103/PhysRevLett.108.237204)**, [237204](http://dx.doi.org/10.1103/PhysRevLett.108.237204) [\(2012\)](http://dx.doi.org/10.1103/PhysRevLett.108.237204).
- [17] S. Muhlbauer, B. Binz, F. Jonietz, C. Pfleiderer, A. Rosch, ¨ A. Neubauer, R. Georgii, and P. Boni, [Science](http://dx.doi.org/10.1126/science.1166767) [323](http://dx.doi.org/10.1126/science.1166767), [915](http://dx.doi.org/10.1126/science.1166767) [\(2009\)](http://dx.doi.org/10.1126/science.1166767).
- [18] W. Münzer, A. Neubauer, T. Adams, S. Mühlbauer, C. Franz, F. Jonietz, R. Georgii, P. Böni, B. Pedersen, M. Schmidt et al., [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.81.041203) **[81](http://dx.doi.org/10.1103/PhysRevB.81.041203)**, [041203](http://dx.doi.org/10.1103/PhysRevB.81.041203) [\(2010\)](http://dx.doi.org/10.1103/PhysRevB.81.041203).
- [19] X. Z. Yu, Y. Onose, N. Kanazawa, J. H. Park, J. H. Han, Y. Matsui, N. Nagaosa, and Y. Tokura, [Nature \(London\)](http://dx.doi.org/10.1038/nature09124) **[465](http://dx.doi.org/10.1038/nature09124)**, [901](http://dx.doi.org/10.1038/nature09124) [\(2010\)](http://dx.doi.org/10.1038/nature09124).
- [20] X. Z. Yu, N. Kanazawa, Y. Onose, K. Kimoto, W. Z. Zhang, Y. Matsui, and Y. Tokura, [Nat. Mater.](http://dx.doi.org/10.1038/nmat2916) **[10](http://dx.doi.org/10.1038/nmat2916)**, [106](http://dx.doi.org/10.1038/nmat2916) [\(2011\)](http://dx.doi.org/10.1038/nmat2916).
- [21] A. Hamann, D. Lamago, T. Wolf, H. v. Löhneysen, and D. Reznik, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.107.037207) **[107](http://dx.doi.org/10.1103/PhysRevLett.107.037207)**, [037207](http://dx.doi.org/10.1103/PhysRevLett.107.037207) [\(2011\)](http://dx.doi.org/10.1103/PhysRevLett.107.037207).
- [22] S. S. Samatham and V. Ganesan, [Phys. Status Solidi \(b\)](http://dx.doi.org/10.1002/pssr.201206523) **[7](http://dx.doi.org/10.1002/pssr.201206523)**, [184](http://dx.doi.org/10.1002/pssr.201206523) [\(2013\)](http://dx.doi.org/10.1002/pssr.201206523).
- [23] H. Wilhelm, M. Baenitz, M. Schmidt, U. K. Rößler, A. A. Leonov, and A. N. Bogdanov, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.107.127203) **[107](http://dx.doi.org/10.1103/PhysRevLett.107.127203)**, [127203](http://dx.doi.org/10.1103/PhysRevLett.107.127203) [\(2011\)](http://dx.doi.org/10.1103/PhysRevLett.107.127203).
- [24] H. Wilhelm, M. Baenitz, M. Schmidt, C. Naylor, R. Lortz, U. K. Roessler, A. A. Leonov, and A. N. Bogdanov, J. Phys.: Condens. Matter **[24](http://dx.doi.org/10.1088/0953-8984/24/29/294204)**, [294204](http://dx.doi.org/10.1088/0953-8984/24/29/294204) [\(2012\)](http://dx.doi.org/10.1088/0953-8984/24/29/294204).
- [25] [L. Cevey, H. Wilhelm, M. Schmidt, and R. Lortz,](http://dx.doi.org/10.1002/pssb.201200632) *Phys. Status* Solidi (b) **[250](http://dx.doi.org/10.1002/pssb.201200632)**, [650](http://dx.doi.org/10.1002/pssb.201200632) [\(2013\)](http://dx.doi.org/10.1002/pssb.201200632).
- [26] C. Pappas, E. Lelièvre-Berna, P. Falus, P. M. Bentley, [E. Moskvin, S. Grigoriev, P. Fouquet, and B. Farago,](http://dx.doi.org/10.1103/PhysRevLett.102.197202) Phys. Rev. Lett. **[102](http://dx.doi.org/10.1103/PhysRevLett.102.197202)**, [197202](http://dx.doi.org/10.1103/PhysRevLett.102.197202) [\(2009\)](http://dx.doi.org/10.1103/PhysRevLett.102.197202).
- [27] C. Pappas, E. Lelièvre-Berna, P. Bentley, P. Falus, P. Fouquet, and B. Farago, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.83.224405) **[83](http://dx.doi.org/10.1103/PhysRevB.83.224405)**, [224405](http://dx.doi.org/10.1103/PhysRevB.83.224405) [\(2011\)](http://dx.doi.org/10.1103/PhysRevB.83.224405).
- [28] T. Adams, S. Mühlbauer, C. Pfleiderer, F. Jonietz, A. Bauer, A. Neubauer, R. Georgii, P. Böni, U. Keiderling, K. Everschor *et al.*, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.107.217206) **[107](http://dx.doi.org/10.1103/PhysRevLett.107.217206)**, [217206](http://dx.doi.org/10.1103/PhysRevLett.107.217206) [\(2011\)](http://dx.doi.org/10.1103/PhysRevLett.107.217206).
- [29] J. Kindervater, Diplomarbeit, Technische Universität München, 2010.
- [30] F. Tasset, [Phys. B: Condens. Matter](http://dx.doi.org/10.1016/0921-4526(89)90749-7) **[156](http://dx.doi.org/10.1016/0921-4526(89)90749-7)**, [627](http://dx.doi.org/10.1016/0921-4526(89)90749-7) [\(1989\)](http://dx.doi.org/10.1016/0921-4526(89)90749-7).
- [31] M. Janoschek, S. Klimko, R. Gähler, B. Roessli, and P. Böni, [Phys. B: Condens. Matter](http://dx.doi.org/10.1016/j.physb.2007.02.074) **[397](http://dx.doi.org/10.1016/j.physb.2007.02.074)**, [125](http://dx.doi.org/10.1016/j.physb.2007.02.074) [\(2007\)](http://dx.doi.org/10.1016/j.physb.2007.02.074).
- [32] W. Häußler, B. Gohla-Neudecker, R. Schwikowski, D. Streibl, and P. Böni, *[Phys. B: Condens. Matter](http://dx.doi.org/10.1016/j.physb.2007.02.086) [397](http://dx.doi.org/10.1016/j.physb.2007.02.086)*, [112](http://dx.doi.org/10.1016/j.physb.2007.02.086) [\(2007\)](http://dx.doi.org/10.1016/j.physb.2007.02.086).
- [33] W. Häußler, P. Böni, M. Klein, C. J. Schmidt, U. Schmidt, F. Groitl, and J. Kindervater, [Rev. Sci. Instrum.](http://dx.doi.org/10.1063/1.3571300) **[82](http://dx.doi.org/10.1063/1.3571300)**, [045101](http://dx.doi.org/10.1063/1.3571300) [\(2011\)](http://dx.doi.org/10.1063/1.3571300).
- [34] M. Blume, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRev.130.1670) **[130](http://dx.doi.org/10.1103/PhysRev.130.1670)**, [1670](http://dx.doi.org/10.1103/PhysRev.130.1670) [\(1963\)](http://dx.doi.org/10.1103/PhysRev.130.1670).
- [35] S. Maleyev, V. Baryakhtar, and R. Suris, Fiz. Tverd. Tela **4**, 3461 (1962).
- [36] S. V. Grigoriev, S. V. Maleyev, A. I. Okorokov, Y. O. Chetverikov, R. Georgii, P. Böni, D. Lamago, H. Eckerlebe, and K. Pranzas, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.72.134420) **[72](http://dx.doi.org/10.1103/PhysRevB.72.134420)**, [134420](http://dx.doi.org/10.1103/PhysRevB.72.134420) [\(2005\)](http://dx.doi.org/10.1103/PhysRevB.72.134420).