

Domain-wall free energy in Heisenberg ferromagnets

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We consider Gaussian fluctuations about domain walls embedded in one- or two-dimensional spin lattices. Analytic expressions for the free energy of one domain wall are obtained. From these, the temperature dependence of experimentally relevant spatial scales—i.e., the correlation length for spin chains and the size of magnetic domains for thin films magnetized out of plane—are deduced. Stability of chiral order inside domain walls against thermal fluctuations is also discussed.

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I. INTRODUCTION

The physics of magnetic domain walls (DWs) has experienced a sort of renaissance during the last decade. This was triggered by the perspective of employing DWs in spintronic devices [1–5] and by the improvement in spatial resolution with which magnetic textures could be resolved [6,7]. A lot of theoretical work has been done to investigate diverse physical properties of DWs in view of the novel applicative and experimental scenario [8–13]. However, the effect of thermal fluctuations within a DW as a single object—to our knowledge—has scarcely been investigated [14–19]. The basic theoretical formalism to tackle this problem analytically was developed between the 70's and the early 80's. Then, the thermodynamics of (not exactly solvable) one-dimensional (1D) classical-spin models was described through a dilute gas of noninteracting DWs, including their interplay with spin waves [20–24]. Some results [25,26] have been recently actualized in the context of magnonic applications [27–30]. Here, we focus on the free energy of a single DW, with particular regard to its dependence on temperature and on the system size. We discuss the implications on the physics of molecular spin chains [31–35], ferromagnetic films, and nanowires [36]. Remarkably, the profile of DWs embedded in all these systems is commonly described by the very same model at zero temperature: we consider the effect of thermally excited spin waves, i.e., Gaussian fluctuations [37], about this spin profile.

For the 1D case, the model is presented in Sec. II, where all the assumptions and analytic results are checked against numerical calculations on a discrete lattice (see also the appendices). The strategy followed to compute the DW free energy numerically is explained in details in Sec. III. In Sec. IV, we extend the analytic part to 2D systems. A central result is that Gaussian fluctuations suffice to explain

the floating of magnetic-domain patterns and the decrease of their characteristic period of modulation with increasing temperature, both facts being observed experimentally [38–40]. Within the same approximation, we conjecture the absence of chiral order within DWs interposed between saturated magnetic domains.

II. THE MODEL

We consider the following classical Heisenberg Hamiltonian:

$$\mathcal{H} = - \sum_{i=0}^L [J \vec{S}_i \cdot \vec{S}_{i+1} + D(S_i^z)^2] - D(S_{L+1}^z)^2, \quad (1)$$

where D represents the anisotropy energy and J is the exchange coupling. Each spin variable \vec{S}_i is a three-component unit vector associated with the i th site of the lattice. The two spins at opposite boundaries are forced to lie along the easy anisotropy axis either parallel ($\uparrow\uparrow$) or antiparallel ($\uparrow\downarrow$) to each other (see the sketch in Fig. 1). By computing the partition function for these two different boundary conditions (b.c.) the free-energy increase associated with the creation of a DW from a uniform ground state can be deduced. In this paper, we will focus on broad DWs [34], obtained for J significantly larger than D , so that the micromagnetic limit for Hamiltonian (1) is meaningful:

$$\mathcal{H} = \int_0^{L+1} dx \left[\frac{J}{2} |\partial_x \vec{S}|^2 - D(S^z(x))^2 \right] + \text{const}, \quad (2)$$

(unitary lattice constant is assumed). Ferromagnetic b.c. ($\uparrow\uparrow$) are obtained setting $\vec{S}(x=0) = \vec{S}(x=L+1) = (0,0,1)$, while antiferromagnetic b.c. ($\uparrow\downarrow$) correspond to $\vec{S}(x=0) = (0,0,1)$ and $\vec{S}(x=L+1) = (0,0,-1)$. Following the procedure presented in Refs. [34,41,42], $\vec{S}(x)$ is decomposed in two vector fields:

$$\vec{S}(x) = \vec{n}(x) \sqrt{1 - \vec{\phi}^2} + \vec{\phi}(x), \quad (3)$$

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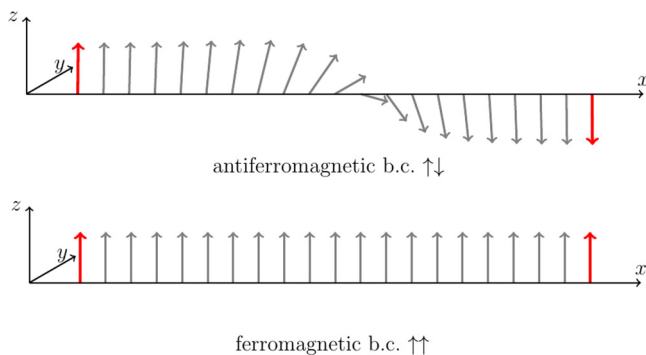


FIG. 1. (Color online) Sketch of the two boundary conditions considered in the manuscript with the corresponding minimum-energy profile $\vec{n}(x)$. For $\uparrow\uparrow$ b.c., we chose $\vec{n}(x) = (0, 0, 1)$, while for $\uparrow\downarrow$ b.c., $\vec{n}(x)$ is given by Eq. (19)—with arbitrary x_0 and φ_0 —and corresponds to the infinite-chain profile.

$\vec{\phi}(x)$ representing fluctuations and $\vec{n}(x)$ assumed to vary smoothly in space. If $|\vec{S}(x)| = 1$ and $|\vec{n}(x)| = 1$ are required, $\vec{n}(x) \cdot \vec{\phi}(x) = 0$. Therefore $\vec{\phi}(x)$ can be expressed on a *local*, two-dimensional basis orthogonal to $\vec{n}(x)$:

$$\vec{\phi}(x) = \sum_a \phi_a \vec{e}_a, \quad \text{with} \quad |\vec{e}_a(x)| = 1. \quad (4)$$

Through the decomposition given in Eq. (3), the Hamiltonian (2) can be expanded for small $\phi_a(x)$. If the expansion is truncated to quadratic terms, fluctuation amplitudes shall enter the partition function in the form of Gaussian integrals. This will be assumed in all the analytic derivations of the free energy and referred to as the *Gaussian approximation*. With the same meaning should be understood the expression Gaussian fluctuations [43]. By construction, the whole formalism is expected to hold only at low temperature. The original Hamiltonian thus splits in two contributions: $\mathcal{H} = \mathcal{H}[\vec{n}] + \mathcal{H}[\vec{\phi}] + \text{const.}$ $\mathcal{H}[\vec{n}]$ has the same form as Hamiltonian (2) provided that $\vec{S}(x)$ is substituted with $\vec{n}(x)$. The fluctuation Hamiltonian reads

$$\mathcal{H}[\vec{\phi}] = \sum_a \int_0^{L+1} dx \phi_a(x) \hat{H}_a \phi_a(x), \quad (5)$$

in which \hat{H}_a acts as a Schrödinger-like operator that takes a different form depending on the chosen slow-varying profile $\vec{n}(x)$. After having solved the eigenvalue problem

$$\hat{H}_a \psi(x) = \varepsilon \psi(x), \quad (6)$$

each component (labeled by a) of the fluctuating field can be expanded on eigenfunctions of \hat{H}_a :

$$\phi_a(x) = \sum_\nu a_{a,\nu} \Psi_{a,\nu}(x) + \sum_q a_{a,q} \Phi_{a,q}(x). \quad (7)$$

In our notation, we associate the greek index ν with (possible) bound states and q with free states. By free states we mean functions that are delocalized throughout the spin chain, also when $L < \infty$. As a consequence of the finite size and of our choice of boundary conditions, such “free” states actually correspond to the wave functions of a free particle in a box when $\vec{n}(x)$ is assumed uniform (lower sketch in Fig. 1). In our vocabulary, there is no bound state in this case. When $\uparrow\downarrow$

b.c. are assumed, instead, Eq. (6) admits one bound state per component a . As it takes just one value, the label ν will be dropped henceforth from eigenfunctions $\Psi_{a,\nu}(x)$, coefficients $a_{a,\nu}$, and eigenvalues $\varepsilon_{a,\nu}$ (quantities associated with free states will still be denoted by the label q). If $\Psi_a(x)$ and $\Phi_{a,q}(x)$ are normalized properly, the expansion (7) allows rewriting the fluctuation Hamiltonian as

$$\mathcal{H}[\vec{\phi}] = \sum_a \left[\varepsilon_a |a_a|^2 + \sum_q \varepsilon_{a,q} |a_{a,q}|^2 \right], \quad (8)$$

where ε_a and $\varepsilon_{a,q}$ are the bound-state and free-state eigenvalues of Eq. (6), respectively.

A partition function that depends parametrically on the slow-varying field $\vec{n}(x)$ is obtained integrating over fluctuations, namely,

$$\mathcal{Z}[\vec{n}] = e^{-\beta\mathcal{H}[\vec{n}]} \int \mathcal{D}[\vec{\phi}] e^{-\beta\mathcal{H}[\vec{\phi}]}, \quad (9)$$

where $\int \mathcal{D}[\phi]$ stands for functional integral and $\beta = 1/T$ ($k_B = 1$ henceforth). This partition function can be rewritten as product of Gaussian integrals by making use of Eqs. (7) and (8):

$$\begin{aligned} \mathcal{Z}[\vec{n}] &= e^{-\beta\mathcal{H}[\vec{n}]} \left[\int_a \prod da_a e^{-\beta\varepsilon_a |a_a|^2} \right]^{n_{\text{dw}}} \\ &\times \int_{a,q} \prod da_{a,q} e^{-\beta\varepsilon_{a,q} |a_{a,q}|^2}. \end{aligned} \quad (10)$$

For antiparallel b.c. ($\uparrow\downarrow$) $n_{\text{dw}} = 1$, meaning that one DW is present in the spin chain, while $n_{\text{dw}} = 0$ for parallel b.c. ($\uparrow\uparrow$). All the free states have positive energy, which makes their Gaussian integrals convergent. After this integration, Eq. (10) reads

$$\begin{aligned} \mathcal{Z}[\vec{n}] &= e^{-\beta\mathcal{H}[\vec{n}]} \left[\int_a \prod da_a e^{-\beta\varepsilon_a |a_a|^2} \right]^{n_{\text{dw}}} \\ &\times \exp \left[-\frac{1}{2} \sum_{a,q} \ln \left(\frac{\beta\varepsilon_{a,q}}{\pi} \right) \right]. \end{aligned} \quad (11)$$

The integration over bound-state amplitudes da_a needs more care. We will see that bound states are associated with vanishing energy so that their contribution to the partition function cannot be evaluated through a standard Gaussian integration. The remedy to handle this divergence is presented in details in Ref. [21]. As the explicit form of the bound-state eigenfunctions $\Psi_a(x)$ is required, we prefer to postpone this discussion. Further on, we will approximate the summation over free states $\sum_{a,q}$ in Eq. (11) with an integral. This approximation requires the knowledge of the density of states [44], which also needs to be determined previously by solving the eigenvalue problem in Eq. (6). From $\mathcal{Z}[\vec{n}]$, the free energy corresponding to the b.c. $\uparrow\uparrow$ and $\uparrow\downarrow$ sketched in Fig. 1 can be computed. We will focus on the difference between those free energies $\Delta F = F_{\uparrow\downarrow} - F_{\uparrow\uparrow} = -T \ln(\mathcal{Z}_{\uparrow\downarrow}/\mathcal{Z}_{\uparrow\uparrow})$, which we identify with *one* DW free energy. A similar problem is discussed for a bulk magnet in Ref. [17]; there, a good qualitative agreement between Langevin simulations and mean-field approach is found. A crucial difference with respect to the case investigated here is the occurrence of long-range

order. Though the Mermin-Wagner theorem [45] does not apply to the 1D model considered by us—because of the presence of uniaxial anisotropy—magnetic order is anyway destroyed by thermally-excited DWs in the thermodynamic limit. Still in this limit, physics should be independent of the boundaries. However, as long as the system size is smaller than the correlation length, physics *does* depend on boundary conditions; specifically those sketched in Fig. 1 mimic the situation described for a bulk magnet in Ref. [17]. Since the correlation length typically decreases as temperature is increased, choosing $\uparrow\uparrow$ boundary conditions does not warrant the absence of DWs. More precisely, we expect that DWs start forming spontaneously only when ΔF approaches zero. So we can reasonably think that ΔF is a good estimate for the free energy of *one* DW as long as it is positive.

A. Schrödinger-like eigenvalue problem

Using the decomposition in Eq. (3), the terms entering Hamiltonian (2) can be expanded to second order in $\phi_a(x)$, which gives

$$\begin{aligned} [\partial_x \vec{S}(x)]^2 &= (1 - \vec{\phi}^2)[\partial_x \vec{n}(x)]^2 + \sum_a (\partial_x \phi_a)^2 + \sum_a c_a^2 \phi_a^2, \\ [S^z(x)]^2 &= [n^z(x)]^2(1 - \vec{\phi}^2) + \sum_a \phi_a^2 (e_a^z)^2. \end{aligned} \quad (12)$$

c_a are the components of $\partial_x \vec{n}$ on the basis \vec{e}_a (remember that $|\vec{n}(x)| = 1$ so that $\partial_x \vec{n} \perp \vec{n}$). The fluctuation Hamiltonian reads

$$\begin{aligned} \hat{H}_a &= -\frac{J}{2} \partial_x^2 + V_a(x), \quad \text{with} \\ V_a(x) &= \frac{J}{2} [c_a^2 - (\partial_x \vec{n})^2] + D[(n^z)^2 - (e_a^z)^2]. \end{aligned} \quad (13)$$

With our choice of boundaries, one has the equivalence

$$\int_0^{L+1} dx (\partial_x \phi)^2 = - \int_0^{L+1} dx \phi \partial_x^2 \phi, \quad (14)$$

which yields the second derivative in Eq. (13). The “potential” $V_a(x)$ will be specified by the choice of profile $\vec{n}(x)$. The latter will be chosen as the minimal-energy profile consistent with ferromagnetic and antiferromagnetic b.c. These two cases are discussed separately in the following.

1. Uniform profile

When ferromagnetic ($\uparrow\uparrow$) b.c. are chosen, the ground state is given by a uniform profile. Thus we set $\vec{n}(x) = (0, 0, 1)$ or, equivalently, $\vec{n} = \vec{e}_z$ and \vec{e}_a given by the two vectors $\vec{e}_{x,y}$ ($a = x, y$ for this case). Accordingly, the potential $V_a(x)$ takes the form

$$V_a = D + \begin{cases} 0, & x \in (0, L+1), \\ \infty, & \text{else,} \end{cases} \quad (15)$$

so that the eigenvalue problem is formally equivalent to that of a particle in a box whose normalized eigenfunctions are given by

$$\Phi_{a,q}(x) = \sqrt{\frac{2}{L+1}} \sin(qx), \quad \text{with } a = x, y, \quad (16)$$

$q = l\pi/(L+1)$, $l = 1, 2, 3, \dots$, and eigenvalues

$$\varepsilon_{a,q} = \frac{J}{2} q^2 + D. \quad (17)$$

2. Domain-wall profile

In the case of antiferromagnetic ($\uparrow\downarrow$) b.c., \vec{n} is chosen to be

$$\begin{aligned} n^x(x) &= \cos \varphi_0 \operatorname{sech}[c(x - x_0)], \\ n^y(x) &= \sin \varphi_0 \operatorname{sech}[c(x - x_0)], \\ n^z(x) &= -\tanh[c(x - x_0)], \end{aligned} \quad (18)$$

where $c = \sqrt{2D/J}$ is the inverse DW width. Rigorously, the spin profile in Eq. (18) minimizes the energy of an infinite chain with $\uparrow\downarrow$ b.c., but here we will use it to build the Schrödinger-like equation for a finite chain (see Appendix F for further details). As we set no anisotropy in the hard plane (xy), nor we consider magnetostatic interaction, DWs parametrized by different angles φ_0 have the same energy (e.g., Bloch and Néel DWs). One of the two vectors \vec{e}_a is the tangent vector \vec{e}_{x_0} :

$$\begin{aligned} e_{x_0}^x(x) &= \cos \varphi_0 \tanh[c(x - x_0)], \\ e_{x_0}^y(x) &= \sin \varphi_0 \tanh[c(x - x_0)], \\ e_{x_0}^z(x) &= \operatorname{sech}[c(x - x_0)], \end{aligned} \quad (19)$$

that is proportional to $\partial_x \vec{n}$:

$$\begin{aligned} \partial_x \vec{n} &= -c \operatorname{sech}[c(x - x_0)] \vec{e}_{x_0}, \quad \text{with} \\ c_{x_0} &= -c \operatorname{sech}[c(x - x_0)]. \end{aligned} \quad (20)$$

The other one can be found from the vector product of \vec{e}_{x_0} and \vec{n} : $\vec{e}_{\varphi_0} = \vec{e}_{x_0} \times \vec{n} = (-\sin \varphi_0, \cos \varphi_0, 0)$ (note that $c_{\varphi_0} = 0$). For our choice of profile and b.c., this vector coincides with the DW chirality [46], that is,

$$\frac{1}{\pi} \int_0^{L+1} \vec{n} \times \partial_x \vec{n} dx = -\vec{e}_{\varphi_0}. \quad (21)$$

The motivation for labeling the vectors \vec{e}_a with $a = x_0, \varphi_0$ will be clear at the end of this paragraph and—we believe—will facilitate reading what follows. The components of the fluctuation field $\vec{\phi}(x)$ on the basis vectors \vec{e}_{x_0} and \vec{e}_{φ_0} will be labeled accordingly: ϕ_{x_0} and ϕ_{φ_0} . The spatial dependence of $\vec{n}(x)$ and $\vec{e}_{x_0}(x)$ propagates to the potentials entering the Schrödinger-like equation associated with the two independent components of $\vec{\phi}(x)$:

$$V_{x_0}(x) = V_{\varphi_0}(x) = D\{2 \tanh^2[c(x - x_0)] - 1\}. \quad (22)$$

In the general case in which an intermediate anisotropy axis is present, which breaks the φ_0 degeneracy in the xy plane, the potentials $V_{x_0}(x)$ and $V_{\varphi_0}(x)$ are different. When the potentials in Eq. (22) are inserted into the eigenvalue problem (6), one obtains an equation that is a special case of the more general one:

$$\begin{aligned} \hat{H}_m \psi(x) &= \tilde{\varepsilon} \psi(x), \quad \text{with} \\ \hat{H}_m &= -\partial_\eta^2 + m(m+1) \tanh^2 \eta, \end{aligned} \quad (23)$$

where the change of variable $\eta = c(x - x_0)$ has been performed, the energy $\tilde{\varepsilon}$ is adimensional in units of D , and $m \in \mathbb{N}$.

In terms of the lowering and raising operators

$$\begin{aligned}\hat{a}_m &= \partial_\eta + m \tanh \eta, \\ \hat{a}_m^\dagger &= -\partial_\eta + m \tanh \eta,\end{aligned}\quad (24)$$

one has $\hat{H}_m = \hat{a}_m^\dagger \hat{a}_m + m$. In Appendix A, we recall a demonstration [37] that if $|\psi\rangle_m$ is an eigenstate of \hat{H}_m then $\hat{a}_{m+1}^\dagger |\psi\rangle_m$ is an eigenstate of \hat{H}_{m+1} . This property allows us to construct the eigenstates of \hat{H}_1 —which corresponds to the case treated in this paragraph ($\uparrow\downarrow$)—from the knowledge of the eigenstates of $\hat{H}_0 = -\partial_\eta^2$. Operatively, this means that we can obtain the solution to the eigenvalue problem associated with the potentials in Eq. (22) by applying the raising operator \hat{a}_1^\dagger to the eigenfunctions of a free-particle in a box $\psi_0(q, x) = A \cos(qx) + B \sin(qx)$, A and B being constants to be specified by b.c. Details of the calculations are reported in Appendix B. To proceed in the derivation of the DW free energy, we only need to know the density of states $\rho(q)$ and the eigenvalues $\varepsilon_{a,q}$. The last few have the same dependence on q as in Eq. (17) but the allowed values of q are different, which eventually yields a different $\rho(q)$ with respect to the case of ferromagnetic ($\uparrow\uparrow$) b.c. In addition to these free states, the Schrödinger-like Eq. (6) [with the potentials given in Eq. (22)] also admits one bound state [20,25,27,37]:

$$\Psi_a(x) = \sqrt{\frac{c}{2}} \operatorname{sech}[c(x - x_0)] \quad (25)$$

with eigenvalue $\varepsilon_a = 0$ (remember that we drop the index ν because there is just one bound state per component $a = x_0, \varphi_0$). When an intermediate-anisotropy term $-D_x (S_i^x)^2$ is added to Hamiltonian (1), the energy of the bound state $\Psi_{\varphi_0}(x)$ becomes positive. This suggests that the vanishing of ε_{φ_0} be related to the degeneracy of the profile in Eq. (18) with respect to the angle φ_0 , when $D_x = 0$. In fact, $\partial_{\varphi_0} \vec{n} = \operatorname{sech}[c(x - x_0)] \vec{e}_{\varphi_0}$, meaning that the vector \vec{e}_{φ_0} indicates the direction along which the minimum-energy profile is deformed in response to a variation $\delta\varphi_0$. In other words, any rotation on the xy plane of the chiral (vector) degree of freedom—defined in Eq. (21)—does not affect the DW energy if $D_x = 0$; while only reflections of the chirality vector, $\varphi_0 \rightarrow \varphi_0 + \pi$, leave the energy unchanged when $D_x \neq 0$ (e.g., left- and right-handed Bloch DWs are degenerate). On the contrary, the energy ε_{x_0} remains zero also when an intermediate-anisotropy axis exists ($D_x \neq 0$). This is related to the degeneracy [21,37] with respect to x_0 , as confirmed by the equivalence $\partial_{x_0} \vec{n} = c \operatorname{sech}[c(x - x_0)] \vec{e}_{x_0}$, obtained from Eq. (20) with the substitution $x \leftrightarrow x_0$.

B. Domain-wall free energy

In the case in which $\vec{n}(x)$ represents a DW profile ($\uparrow\downarrow$ b.c.), the considerations exposed above furnish a recipe to replace the integration over the amplitudes da_{x_0} and da_{φ_0} by integrals over the DW center x_0 and the angle φ_0 , respectively. This makes it possible to evaluate the contribution of bound states to the partition function (11). From the relations between \vec{e}_{x_0, φ_0} , $\partial_{x_0} \vec{n}$, and $\partial_{\varphi_0} \vec{n}$, the required Jacobian can be deduced [20,21]:

$$\begin{aligned}|\partial_{x_0} \vec{n}| dx_0 &= \Psi_{x_0}(x) da_{x_0} & da_{x_0} &= dx_0 \sqrt{2c}, \\ |\partial_{\varphi_0} \vec{n}| d\varphi_0 &= \Psi_{\varphi_0}(x) da_{\varphi_0} & da_{\varphi_0} &= d\varphi_0 \sqrt{2/c}.\end{aligned}\quad (26)$$

The integrals we are interested in reduce to

$$\int_{a=x_0, \varphi_0} \Pi da_a = 2 \int_0^L dx_0 \int_0^{2\pi} d\varphi_0 = 4\pi L, \quad (27)$$

where the fact that $\varepsilon_{x_0} = \varepsilon_{\varphi_0} = 0$ has been used and the factor 2 comes from the Jacobian deduced in Eq. (26). Strictly speaking, it would be more appropriate to let the integral over dx_0 range from $1/(2c)$ to $L + 1 - 1/(2c)$. However, since this numerical factors do not affect the results significantly, we prefer to keep analytic expressions as simple as possible.

When a uniform $\vec{n}(x)$ is assumed ($\uparrow\uparrow$ b.c.), the eigenvalue problem in Eq. (6) does not admit bound-state solutions, thus we can formally set the integral in Eq. (27) equal to one. We approximate the remaining summation on the free states in Eq. (11) with an integral:

$$\sum_q \rightarrow \frac{1}{2} \int_{-\infty}^{\infty} \rho(q) dq. \quad (28)$$

The density of states has the form $\rho(q) = \rho_{\uparrow\uparrow} - \gamma(q)$, with

$$\gamma(q) = \frac{2c \tanh[c(L+1)/2]}{\pi \{c^2 \tanh^2[c(L+1)/2] + q^2\}} \quad (29)$$

and $\rho_{\uparrow\uparrow} = (L+1)/\pi$. The density of states for these linear excitations superimposed to a DW profile $\vec{n}(x)$ is derived in Appendix B. As defined in Eq. (29), $\gamma(q)$ corresponds to placing one DW in the middle of the chain [$x_0 = (L+1)/2$]. The general formula would depend parametrically on the coordinate of the DW center [see Eq. (B9)], which renders the calculation of the partition function unnecessarily complicated. In the limit $L \gg 1/c$, Eq. (29) can be further simplified setting $\tanh^2[c(L+1)/2] \approx 1$ [see Eq. (B13)]. The resulting expression of $\rho(q)$ is plotted in Fig. 2 as a (blue) short-dashed line, finding good agreement with the same quantity determined numerically (red squares). Details about the numerical calculation are also given in Appendix B. For large values of q , $\rho(q)$ approaches the density of states obtained with uniform b.c. $\rho_{\uparrow\uparrow} = (L+1)/\pi$ (green, solid line). Consistently with Eq. (29), the partition functions for

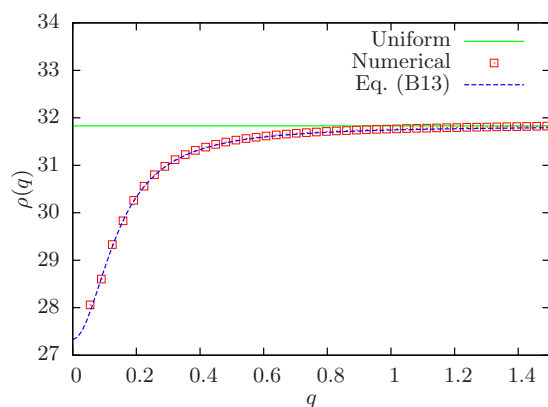


FIG. 2. (Color online) Density of states $\rho(q) = \partial n / \partial q$ obtained analytically (blue, short-dashed line) and numerically (red squares) for $D/J = 0.01$ and $L = 99$. For small values of q , the density is decreased with respect to $\rho_{\uparrow\uparrow} = 100/\pi$, expected for a uniform \vec{n} , due to DW-spin wave interaction.

the two b.c. sketched in Fig. 1 read

$$\begin{aligned} \mathcal{Z}_{\uparrow\uparrow} &= \exp \left[-\frac{1}{2} \int_{-\infty}^{\infty} \ln \left(\frac{\beta \varepsilon_{a,q}}{\pi} \right) \rho_{\uparrow\uparrow} dq \right], \\ \mathcal{Z}_{\uparrow\downarrow} &= 4\pi L e^{-\beta \mathcal{E}_{\text{dw}}} \\ &\quad \times \exp \left[-\frac{1}{2} \int_{-\infty}^{\infty} \ln \left(\frac{\beta \varepsilon_{a,q}}{\pi} \right) (\rho_{\uparrow\uparrow} - \gamma(q)) dq \right], \end{aligned} \quad (30)$$

where $\mathcal{E}_{\text{dw}} = 2\sqrt{2JD}$ represents the DW energy with respect to the uniform ground state, while $\varepsilon_{a,q} = Jq^2/2 + D$ consistently with Eqs. (B2) and (17). Equations (30) are obtained from Eq. (11) by setting $n_{\text{dw}} = 0, 1$ for $\uparrow\uparrow$ and $\uparrow\downarrow$ b.c., respectively. We are now in the position to give an expression for the DW free energy:

$$\begin{aligned} \Delta F &= -T \ln(\mathcal{Z}_{\uparrow\downarrow}/\mathcal{Z}_{\uparrow\uparrow}) \\ &= \mathcal{E}_{\text{dw}} - T \ln(4\pi L) - \frac{T}{2} \int_{-\infty}^{\infty} \ln \left[\frac{\beta J}{2\pi} (q^2 + c^2) \right] \gamma(q) dq. \end{aligned} \quad (31)$$

Using the fact that

$$\int_{-\infty}^{\infty} \frac{\ln(q^2 + c^2)}{q^2 + c_L^2} dq = \frac{2\pi}{c_L} \ln(c + c_L) \quad (32)$$

with $c_L = c \tanh[c(L+1)/2]$, after some algebra, one obtains

$$\Delta F = \mathcal{E}_{\text{dw}} - T \ln(4\beta DL \{1 + \tanh[c(L+1)/2]\}^2). \quad (33)$$

Note that $\tanh(c(L+1)/2)$ approaches one in the limit $L \gg 1/c$. In Fig. 3, the free-energy difference ΔF computed numerically with the transfer-matrix technique (see Sec. III) is compared with analytic predictions. Two possible system sizes were considered $L = 20, 100$. The theoretical curve given in Eq. (33) and numerical points literally overlap at low temperature. At intermediate temperatures, it turns out that a better agreement is achieved when the bare values of D and c appearing in the argument of the logarithm are replaced by the corresponding renormalized quantities

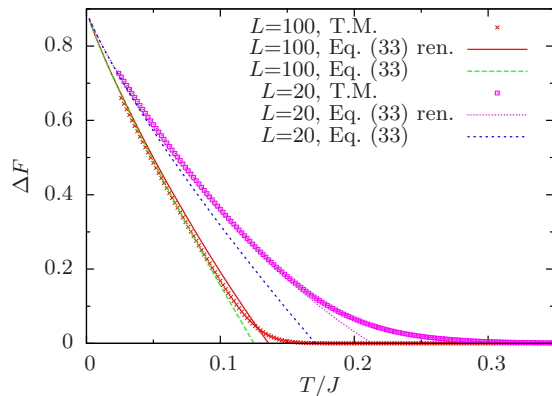


FIG. 3. (Color online) Domain-wall free energy ΔF as a function of temperature for different system sizes L and $D/J = 0.1$. Symbols (crosses $L = 100$ and squares $L = 20$) are numerical values obtained by means of transfer-matrix calculations (T.M. in the legend). Lines represent the prediction of Eq. (33), either using bare constants D and c or renormalized ones $D(\lambda_c)$ and $c(\lambda_c)$ (abbreviated as “ren.” in the legend), as described in the main text.

$D(\lambda_c) = D Z^3(\lambda_c)$ and c by $c(\lambda_c) = c Z(\lambda_c)$. The derivation of the temperature-dependent $Z(\lambda)$, describing Polyakov-renormalization flux [34,41,42], is reposed for the reader’s convenience in Appendix C. λ_c is the cutoff length at which renormalization stops. By increasing L , the temperature at which $\Delta F \simeq 0$ systematically lowers, as expected from the logarithmic dependence on L in Eq. (33). For this reason, the renormalization effects are less severe and already the Gaussian approximation works very well for $L = 100$. However, for $L = 20$, the prediction of Eq. (33) with bare constants [labeled with “Eq. (33)” in the figure legend] deviates more significantly from the exact numerical results. In particular, the transfer-matrix calculation displays a concavity not shown by the theoretical curve obtained with bare values of c and D . The concavity is, instead, reproduced at intermediate temperatures by Eq. (33) if renormalized constants $D(\lambda_c)$ and $c(\lambda_c)$ are used [labeled with “Eq. (33) ren.” in the figure legend].

C. Susceptibility and correlation length

In an infinite spin chain, pair-spin correlations decay with the distance as $\langle S^\alpha(x+r)S^\alpha(x) \rangle = \langle (S^\alpha(x))^2 \rangle e^{-|r|/\xi}$, where ξ is the correlation length and $\alpha = x, y, z$. Henceforth, we will focus on their z component and on the magnetic susceptibility along the easy axis, to which they are related through the equation

$$\begin{aligned} \chi &= -\frac{1}{L} \frac{\partial^2 F}{\partial B_z^2} = \frac{(g\mu_B)^2}{T} \langle (S^z(x))^2 \rangle \sum_{r=-\infty}^{+\infty} e^{-|r|/\xi} \\ &= \frac{(g\mu_B)^2}{T} \langle (S^z(x))^2 \rangle \coth \left(\frac{1}{2\xi} \right) \end{aligned} \quad (34)$$

with μ_B Bohr magneton and g Landé factor ($k_B = 1$). In the high-temperature limit, ξ vanishes and $\langle (S^\alpha(x))^2 \rangle$ becomes independent of $\alpha = x, y, z$ so that the Curie law is recovered [with the substitution $\langle (S^z(x))^2 \rangle \rightarrow S(S+1)/3$]. In the low-temperature limit, $\langle (S^z(x))^2 \rangle \rightarrow S^2$, $\xi \gg 1$ and $\chi T \sim \xi$. The last relation is normally employed to extract information about the spin-Hamiltonian parameters directly from experimental susceptibility data. For the Ising chain $\xi \sim e^{\beta \mathcal{E}_{\text{dw}}}$ throughout a wide range of temperatures. Therefore it is generally believed that the thermodynamics of anisotropic spin chains may be described with the Ising model just fitting the exchange interaction to the barrier that controls the exponential divergence of χT (Δ_ξ in the literature). In a previous work [34], we showed that this is not appropriate for the model described by Hamiltonian (1) when DWs extend over several lattice units. This is the regime we address in the present work [the use of Hamiltonian (2) is legitimate only in this limit, $c \ll 1$]. Existing expansions of ξ often disagree among them and their applicability is limited to some undefined low-temperature region. Starting from Eq. (33) and Polyakov renormalization, the following low-temperature expansion is derived in Appendix D:

$$\xi \simeq \frac{1}{4e c \beta \mathcal{E}_{\text{dw}}} \frac{e^{\beta \mathcal{E}_{\text{dw}}}}{Z^3(\lambda_c)}. \quad (35)$$

The validity of Eq. (35) was checked against numerical results obtained with the transfer-matrix technique and compared

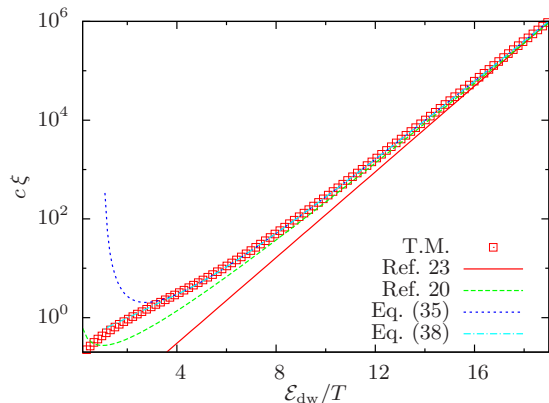


FIG. 4. (Color online) Correlation length in units of DW width $c\xi$ computed for $D/J = 0.1$. Symbols are obtained with the transfer-matrix algorithm for an infinite chain (T.M. in the legend). Lines represent analytic expressions given in the literature, Eqs. (36) and (37), and obtained by us. The dash-dotted (light-blue) curve [“Eq. (38)” in the legend] is obtained by continuing Eq. (35) with Eq. (38) for $\mathcal{E}_{\text{dw}}/T < 4.262$. The different theoretical curves have been rescaled by constant factors to match the numerical points at low temperatures.

with other two expansions reported in the literature. More specifically, in Refs. [22,23] it was proposed

$$\xi \sim \frac{1}{c} e^{\beta\mathcal{E}_{\text{dw}}}, \quad (36)$$

which differs by a temperature-dependent prefactor from the result obtained in Ref. [20]:

$$\xi \sim \frac{1}{c\beta\mathcal{E}_{\text{dw}}} e^{\beta\mathcal{E}_{\text{dw}}}. \quad (37)$$

In Fig. 4, all analytic results have been rescaled by a constant in order to overlap the numerical points at low temperature. Compared with the other two low-temperature expansions, Eq. (35) reproduces the numerical points up to higher temperatures, but it fails as well for $T > \mathcal{E}_{\text{dw}}/4$. More dramatic is the divergence of Eq. (35) as T approaches \mathcal{E}_{dw} , associated with the vanishing of $Z(\lambda_c)$. We understand this spurious effect as due to the fact that the cutoff λ_c at which renormalization stops becomes of the order of ξ itself. Tentatively, we tried to tune this artifact by assuming that the contributions arising from spin waves (including Polyakov renormalization) somehow freeze when λ_c becomes of the order of the actual number of spins aligned along the same direction, which happens for $\beta\mathcal{E}_{\text{dw}} = 4.262$ (see Appendix D). For higher temperatures, we thus suggest to replace Eq. (35) by

$$\xi = \frac{A_0}{c} e^{b\beta\mathcal{E}_{\text{dw}}}, \quad (38)$$

with constants $A_0 = 0.3282$ and $b = 0.5496$ determined by requiring the continuity of ξ and its derivative with respect to $\beta\mathcal{E}_{\text{dw}}$. Note that in all Eqs. (35)–(38), the product $c\xi$ only depends on $\beta\mathcal{E}_{\text{dw}}$. In other words, the correlation length in units of DW width should be a *universal, scaling* function of the same variable, as highlighted numerically in Ref. [34]. Of course, different expansions for ξ yield a different scaling function,

whose validity is eventually confirmed by comparison with numerical results and experiments. The phenomenological curve obtained using Eq. (35) for $\beta\mathcal{E}_{\text{dw}} > 4.262$ and Eq. (38) otherwise is plotted with a dash-dotted (light-blue) line in Fig. 4. Though it is not fully justified on rigorous footing, the curve reproduces well the *universal* behavior of $c\xi$ versus $\beta\mathcal{E}_{\text{dw}}$ till the correlation length becomes of the order of the DW width ($c\xi \simeq 1$). At such short scales, the notion of correlation length itself becomes somewhat meaningless.

Remarkably, the different expansions (35)–(38) disagree in the temperature range in which ξ can be accessed experimentally. In fact, in real systems, defects and impurities limit the chain size to 10^2 – 10^4 consecutive, interacting spins [32,33]. For typical values [47] of J and D , one has $c \simeq 0.1$. Therefore the behavior of the correlation length can be studied experimentally only for $T > \mathcal{E}_{\text{dw}}/6$, for the purest spin chains, or $T > \mathcal{E}_{\text{dw}}/4$, in worst cases. When the latter scenario occurs, Eq. (38) suggests that the barrier Δ_ξ , defined as $\xi \sim e^{\beta\Delta_\xi}$, should approach half of the DW energy \mathcal{E}_{dw} . Indeed, this has been observed in Mn(III)-TCNE molecular spin chains [48–50], whose main thermodynamic features should be captured by the model Hamiltonian (1). In a recently published review article [51], it was pointed out that Eq. (36) does not succeed in reproducing the measured susceptibility of this family of spin chains (though Eq. (36) was implicitly attributed to Ref. [34] instead of Ref. [23]). For completeness, we mention that preliminary results indicate that averaging over different directions of the easy axis—with respect to the applied field—also affects the value of the effective barrier Δ_ξ when the latter is extracted from powder measurements.

The temperature dependence of $\langle (S^z(x))^2 \rangle$ can also affect the magnetic susceptibility at intermediate temperatures, $T > \mathcal{E}_{\text{dw}}/4$ [see Eq. (34)]. According to Eq. (C3) and the following equations, also this quantity should depend—to a large extent—only on $\beta\mathcal{E}_{\text{dw}}$. In passing, we note that within our classical model the correct Curie constant cannot be recovered at high temperature: a fully quantum-mechanical calculation would be needed. This question lies beyond the scope of the present work.

III. NUMERICAL CHECKS

In this section, we describe and critically comment the numerical calculations that were made in order to check robustness and limitations of our analytic results. The derivation of Eq. (33) strongly relies on the solution to Eq. (6) and on the correctness of the density of states $\rho(q)$ given in Eq (29). All these intermediate passages were checked diagonalizing the equivalent discrete-lattice problem, as explained in details in Appendix B. The transfer-matrix algorithm and how finite size constrains the minimal-energy spin profile will be discussed in separate sections.

A. Finite-size transfer matrix

The thermodynamic properties of the model described by Hamiltonian (1) and—more generally—of any classical-spin chain with nearest-neighbor interactions can be computed with the transfer-matrix technique. This approach has been largely used to derive analytic [22,23,52] and numerical

results [53,54] for infinite chains. Less frequently it has been employed to study finite systems [55], for which more flexible methods exist [56,57]. As far as equilibrium properties are concerned, integration of stochastic Landau-Lifshitz-Gilbert equation, Monte-Carlo and transfer-matrix algorithms should produce the same results (for a comparison between the first two approaches see Refs. [56], [57], and [34] for the last two). For our specific problem of computing the free-energy difference between the configurations sketched in Fig. 1, the transfer-matrix technique turned out to be extremely efficient. Let us separate the bulk interactions from those characterizing the spins at boundaries:

$$\mathcal{H} = - \sum_{i=1}^{L-1} V_{\text{bulk}}(\vec{S}_i, \vec{S}_{i+1}) - V_{\text{bc}}(\vec{S}_1) - V_{\text{bc}}(\vec{S}_L). \quad (39)$$

Defining

$$\begin{aligned} V_{\text{exch}}(\vec{S}_i, \vec{S}_{i+1}) &= J \vec{S}_i \cdot \vec{S}_{i+1}, \\ V_{\text{ss}}(\vec{S}_i) &= D (S_i^z)^2, \end{aligned} \quad (40)$$

we have

$$\begin{aligned} V_{\text{bulk}}(\vec{S}_i, \vec{S}_{i+1}) &= V_{\text{exch}}(\vec{S}_i, \vec{S}_{i+1}) + \frac{1}{2}(V_{\text{ss}}(\vec{S}_i) + V_{\text{ss}}(\vec{S}_{i+1})), \\ V_{\text{bc}}(\vec{S}_1) &= \frac{1}{2} V_{\text{ss}}(\vec{S}_1) + J S_1^z, \\ V_{\text{bc}}(\vec{S}_L) &= \frac{1}{2} V_{\text{ss}}(\vec{S}_L) \pm J S_L^z, \end{aligned} \quad (41)$$

the last term being plus for $\uparrow\uparrow$ b.c. and minus for $\uparrow\downarrow$ b.c.. V_{exch} can possibly describe anisotropic exchange, Dzyaloshinskii-Moriya interaction, etc., and V_{ss} any type of single-spin interaction (other anisotropy terms, Zeeman energy, etc.). With these conventions, the transfer-matrix kernel [58] takes the form

$$\mathcal{K}_{\text{TM}}(\vec{S}_i, \vec{S}_{i+1}) = \exp[\beta V_{\text{bulk}}(\vec{S}_i, \vec{S}_{i+1})]. \quad (42)$$

The latter defines the following eigenvalue problem:

$$\int \mathcal{K}_{\text{TM}}(\vec{S}_i, \vec{S}_{i+1}) W_m(\vec{S}_{i+1}) d\Omega_{i+1} = \lambda_m W_m(\vec{S}_i) \quad (43)$$

whose eigenvalues may typically be ordered from the largest to the smallest one:

$$\lambda_0 > \lambda_1 > \lambda_2 > \dots$$

The eigenvalue problem (43) can be solved analytically only in few fortunate cases [59]. Generally, it can be converted into a linear-algebra problem and solved numerically by discretizing the unitary sphere [60–62]. For each temperature, the number of points used to sample the solid angle was increased until the desired precision in the free-energy calculation was reached. Once that eigenvalues and eigenvectors are known, the partition function is given by

$$\begin{aligned} \mathcal{Z}_{\text{TM}} &= \sum_m \lambda_m^{L-1} w_m^1 w_m^L \quad \text{with} \\ w_m^1 &= \int \exp\left[\frac{1}{2}\beta V_{\text{bc}}(\vec{S}_1)\right] W_m(\vec{S}_1) d\Omega_1, \\ w_m^L &= \int \exp\left[\frac{1}{2}\beta V_{\text{bc}}(\vec{S}_L)\right] W_m(\vec{S}_L) d\Omega_L. \end{aligned} \quad (44)$$

The calculation of $\mathcal{Z}_{\uparrow\downarrow}$ and $\mathcal{Z}_{\uparrow\uparrow}$ just requires to change one sign in the definition of $V_{\text{bc}}(\vec{S}_L)$ given in Eq. (41). In this way, $\Delta F = -T \ln(\mathcal{Z}_{\uparrow\downarrow}/\mathcal{Z}_{\uparrow\uparrow})$ was computed also numerically (with two different methods described in Appendix E) to check the validity of Eq. (33). Results are plotted in Fig. 3. Further details concerning the calculation of $\langle (S^z(x))^2 \rangle$ shown in Fig. 6 may be found in Ref. [55], while details about how to compute the correlation length for the infinite chain (see Fig. 4) are given in Ref. [34].

B. General spin profile

The whole theoretical treatment developed in the previous and in the next sections strongly depends on the profile $\vec{n}(x)$ assumed in Eq. (18). Strictly speaking, this profile minimizes the DW energy of an infinite chain when the micromagnetic limit is legitimate (i.e., for $c \ll 1$). Still working with broad DWs, analytic expressions can be derived for the case of $L < \infty$, which are recalled for the reader's convenience in Appendix F. The accuracy of those analytic results was—in turn—checked against a discrete-lattice calculation based on nonlinear maps [63–65]. The most remarkable fact is that when the DW width becomes of the order of the system size (L) the dependence of $n^z(x)$ on the spatial coordinate crosses over from hyperbolic tangent to cosine (see Fig. 8). This imposes a natural range of validity to our model, which depends on the size of the characteristic length involved in a specific calculation. For the correlation length, e.g., meaningful analytic results can be derived only as long as $\xi \gtrsim 1/c$. We already commented on this while deriving Eq. (38). In the next section, we will discuss the temperature dependence of the period of modulation of 2D striped domain patterns. Also in this case, our considerations are expected to hold as long as the separation between two successive DWs is much larger than $1/c$. Since the period of modulation is found to decrease with increasing temperature, this self-consistency requirement eventually sets a high-temperature threshold above which our theory cannot be applied. Remarkably, the mean-field approach prescribes the low-temperature stripe phase to evolve into a single-cosine modulation at the *mean-field* Curie temperature [40,66]. However, since the mean-field approximation is known to overestimate the Curie temperature, in realistic systems it may happen that long-range magnetic order is lost before single-cosine modulation is attained. Summarizing, if for *some* reason the period of modulation is constrained to be of the order of $1/c$, a single-cosine modulation is expected. Whether this is necessarily the case when the true Curie temperature of a magnetic film is approached remains—to the best of our knowledge—not clear.

IV. DOMAIN WALLS IN THIN FILMS

The same treatment presented up to now can be extended to DWs described by the spin profile in Eq. (18) embedded in 2D or 3D magnetic lattices. Here we focus on 2D systems, realized in thin ferromagnetic films or (flat) nanowires. More concretely, we address systems whose thickness, τ , (defined along the z direction) is smaller than the DW width: $\tau < 1/c$. The generalization of Hamiltonian (2) to two

dimensions reads

$$\mathcal{H} = \int_0^{L_y+1} dy \int_0^{L_x+1} \left[\frac{J}{2} |\nabla \vec{S}|^2 - D(S^z)^2 \right] dx, \quad (45)$$

where L_x and L_y denote the lateral extensions (in lattice units) of the system along x and y directions, respectively. Exchange and anisotropy constants are related to the atomic counterparts by the thickness: $J = \tau J_{\text{atom}}$ and $D = \tau D_{\text{atom}}$ (τ being the number of monolayers in the film). The vector field defining the local spin direction is a function of two spatial variables $\vec{S} = \vec{S}(x, y)$, but it can still be decomposed as in Eq. (3). The major difference is that the fluctuation field becomes a function of x and y , $\vec{\phi}(x, y)$, while $\vec{n}(x)$ is assumed to depend only on x . Strictly speaking, the assumption of a DW profile being described by Eq. (18) does not allow considering complex magnetic-domain patterns, such as bubbles or labyrinths. Those patterns are not really the focus of the present paper, which rather deals with thermal properties of DWs. Being fully determined by $\vec{n}(x)$, the same basis set \vec{e}_a defined for the 1D case can be employed to study Gaussian fluctuations about a DW hosted in a 2D lattice. Generally, the corresponding Schrödinger-like operator is obtained by adding a free-particle term, acting along the y direction, to the operators \hat{H}_a defined previously:

$$\hat{H}_a^{2D} = -\frac{J}{2} \partial_y^2 + \hat{H}_a. \quad (46)$$

For the uniform case, $\vec{n} = \vec{e}_z$, the eigenfunctions of \hat{H}_a^{2D} are the same as those of a free particle in a box:

$$\Phi_{a,q_x,q_y}(x, y) = \frac{2 \sin(q_x x) \sin(q_y y)}{\sqrt{(L_x + 1)(L_y + 1)}}, \quad (47)$$

with $q_\alpha = l_\alpha \pi / (L_\alpha + 1)$, $l_\alpha = 1, 2, 3, \dots$ ($\alpha = x, y$), and eigenvalues

$$\varepsilon_{a,q_x,q_y} = \frac{J}{2} (q_x^2 + q_y^2) + D. \quad (48)$$

When $\vec{n}(x)$ is chosen as in Eq. (18), the free-state eigenfunctions are given by

$$\Phi_{a,q_x,q_y}(x, y) = \sqrt{\frac{2}{L_y + 1}} \sin(q_y y) \Phi_{a,q_x}(x), \quad (49)$$

with the same eigenvalues as in Eq. (48), $\Phi_{a,q_x}(x)$ being the eigenfunctions defined in Appendix B. The bound-state contribution reads

$$\Psi_{a,q_y}(x, y) = \sqrt{\frac{2}{L_y + 1}} \sin(q_y y) \Psi_a(x), \quad (50)$$

with $\Psi_a(x)$ defined in Eq. (25) and eigenvalues $\varepsilon_{a,q_y} = Jq_y^2/2$ (for both components $a = x_0, \varphi_0$). By analogy with the 1D case, each component of the fluctuating field $\vec{\phi}(x, y)$ can be expanded on eigenfunctions of the operator \hat{H}_a^{2D} defined in Eq. (46). Finally, the fluctuation Hamiltonian can be written as

$$\mathcal{H}[\vec{\phi}] = \sum_a \left[\sum_{q_y} \varepsilon_{a,q_y} |b_{a,q_y}|^2 + \sum_{q_x, q_y} \varepsilon_{a,q_x,q_y} |b_{a,q_x,q_y}|^2 \right], \quad (51)$$

where b_{a,q_y} and b_{a,q_x,q_y} are the projections of a generic $\phi_a(x, y)$ on the basis functions $\Psi_{a,q_y}(x, y)$ and $\Phi_{a,q_x,q_y}(x, y)$, respectively. Using the same conventions as in Eq. (11), the partition function is now given by

$$\mathcal{Z}[\vec{n}] = e^{-\beta \mathcal{H}[\vec{n}]} \left[\int \prod_{a,q_y} db_{a,q_y} e^{-\beta \varepsilon_{a,q_y} |b_{a,q_y}|^2} \right]^{n_{\text{dw}}} \times \exp \left[-\frac{1}{2} \sum_{a,q_x,q_y} \ln \left(\frac{\beta \varepsilon_{a,q_x,q_y}}{\pi} \right) \right] \quad (52)$$

(the label ν has been dropped because it takes just one value for each component a).

Let us consider first the contribution of the bound state associated with translation invariance of DW centers, x_0 . We can formally use the substitution defined in Eq. (26) and replace each amplitude variable by the corresponding coordinate of DW center $b_{x_0,q_y} = x_0(q_y) \sqrt{2c}$ so that

$$\mathcal{Z}_{x_0} = \int \prod_{x_0,q_y} db_{x_0,q_y} e^{-\beta \varepsilon_{x_0,q_y} |b_{x_0,q_y}|^2} = (2c)^{L_y/2} \int \prod_{q_y} dx_0(q_y) \exp \left[-\frac{\beta \mathcal{E}_{\text{dw}}}{2} q_y^2 |x_0(q_y)|^2 \right], \quad (53)$$

where the equivalence $\mathcal{E}_{\text{dw}} = 2Jc$ has been used. The dependence on q_y in $x_0(q_y)$ is better understood by thinking that in the ground state the DW develops as a straight line along y . Therefore $x_0(q_y)$ actually describes the displacement field associated with a corrugation of the DW induced by thermal fluctuations. In fact, a 1D Hamiltonian for elastic displacements (with “stiffness” \mathcal{E}_{dw}),

$$\mathcal{H}_{\text{el}} = \frac{\mathcal{E}_{\text{dw}}}{2} \sum_y [x_0(y+1) - x_0(y)]^2, \quad (54)$$

would yield a partition function equal to Eq. (53) for $q_y \sim 0$, namely, in the continuum-limit formalism. However, if the lattice discreteness is reintroduced, undesired divergences are avoided. From textbooks it is known that Hamiltonian (54) reads

$$\mathcal{H}_{\text{el}} = 2\mathcal{E}_{\text{dw}} \sum_{q_y} \sin^2 \left(\frac{q_y}{2} \right) |x_0(q_y)|^2 \quad (55)$$

in the appropriate Fourier space with $q_y = l_y \pi / (L_y + 1)$ and $1 \leq l_y \leq L_y$. When the Boltzmann weights in Eq. (53) are replaced with $\beta \mathcal{H}_{\text{el}}$, after Gaussian integrations with respect to dx_0 , one obtains

$$\mathcal{Z}_{x_0} = (2c)^{L_y/2} \exp \left\{ -\frac{1}{2} \sum_{q_y} \ln \left[\frac{2\beta \mathcal{E}_{\text{dw}}}{\pi} \sin^2 \left(\frac{q_y}{2} \right) \right] \right\}. \quad (56)$$

For $L_y \gg 1$, one has

$$\begin{aligned} \sum_{q_y} \ln \left[\sin \left(\frac{q_y}{2} \right) \right] &\simeq \frac{L_y + 1}{\pi} \int_0^\pi \ln \left[\sin \left(\frac{q_y}{2} \right) \right] dq_y \\ &= -(L_y + 1) \ln 2, \end{aligned} \quad (57)$$

which leads to

$$\mathcal{Z}_{x_0} = \left(\frac{2\pi}{\beta J} \right)^{L_y/2} \quad (58)$$

(within the present approximation $L_y + 1 \simeq L_y$). A detailed derivation of the contribution due to the bound state associated with the degeneracy with respect to φ_0 is given in Appendix G. At low temperature, this term turns out to be the same as Eq. (58), i.e., $\mathcal{Z}_{\varphi_0} = \mathcal{Z}_{x_0}$.

As for 1D case, we focus on the ratio between the partition functions obtained for $\uparrow\downarrow$ and $\uparrow\uparrow$ b.c.. For one DW embedded in a thin film, this is

$$\begin{aligned} \frac{\mathcal{Z}_{\uparrow\downarrow}}{\mathcal{Z}_{\uparrow\uparrow}} &= e^{-\beta L_y \mathcal{E}_{\text{dw}}} \mathcal{Z}_{x_0} \mathcal{Z}_{\varphi_0} \\ &\times \exp \left\{ \frac{1}{2} \sum_{q_y} \int_{-\infty}^{\infty} \ln \left[\frac{\beta J (q_x^2 + q_y^2 + c^2)}{2\pi} \right] \gamma(q_x) dq_x \right\}. \end{aligned} \quad (59)$$

Note that *only* the summation over the q_x label has been substituted with an integral [following the same conventions as in Eqs. (28) and (29)]. Moreover, the fact that the energies ε_{a,q_x,q_y} in Eq. (48) actually do not depend on the label a has been used. Exploiting again Eq. (32), the argument of the exponential in Eq. (59) can be approximated as

$$\begin{aligned} &\sum_{q_y} \left[\ln \left(\frac{\beta J}{2\pi} \right) + 2 \ln \left(c_L + \sqrt{q_y^2 + c^2} \right) \right] \\ &\simeq L_y \ln \left(\frac{\beta J c_L^2}{2\pi} \right) \\ &+ \frac{2(L_y + 1)}{\pi} \int_0^\pi \ln \left[1 + \sqrt{\left(\frac{q_y}{c_L} \right)^2 + \left(\frac{c}{c_L} \right)^2} \right] dq_y, \end{aligned} \quad (60)$$

where we considered $L_y \gg 1$ and $c_L = c \tanh[c(L_x + 1)/2]$ now. The integral in the last line of Eq. (60) can be solved analytically:

$$\begin{aligned} \int_0^\pi \dots dq_y &= \pi \left[\ln(1 + \tilde{c} \cosh t) \right] + c_L t - \pi \\ &- 2c_L \sqrt{\tilde{c}^2 - 1} \arctan \left[\sqrt{\frac{\tilde{c} - 1}{\tilde{c} + 1}} \tanh \left(\frac{t}{2} \right) \right] \\ &\stackrel{L_x \gg 1}{\simeq} \pi \ln \left[1 + \sqrt{1 + \left(\frac{\pi}{c} \right)^2} \right] + c \operatorname{arcsinh} \left(\frac{\pi}{c} \right) - \pi \end{aligned} \quad (61)$$

with $\sinh t = \pi/c$ and $\tilde{c} = c/c_L$. The latter approaches one when $L_x \gg 1$ is considered. To the leading order and at low temperature, the free energy per unit length of a DW embedded in a thin film, whose linear dimensions are much larger than

the lattice spacing, reads

$$\begin{aligned} \mathcal{F}_{\text{dw}} &= -\frac{T}{L_y} \ln(\mathcal{Z}_{\uparrow\downarrow}/\mathcal{Z}_{\uparrow\uparrow}) \\ &= \mathcal{E}_{\text{dw}} - 2T \left[\frac{c}{\pi} \operatorname{arcsinh} \left(\frac{\pi}{c} \right) - 1 \right] \\ &- 2T \ln \left\{ c \left[1 + \sqrt{1 + \left(\frac{\pi}{c} \right)^2} \right] \right\} - \frac{T}{L_y} \ln(4\pi L_x). \end{aligned} \quad (62)$$

The last entropic term arises from two contributions neglected in \mathcal{Z}_{x_0} and \mathcal{Z}_{φ_0} , namely a rigid translation of the DW as a whole and a rigid rotation of the chirality vectors (see Appendix G for details). Those contributions have been already computed for the 1D case in Eq. (27), for which y takes just one value. Even if this entropic term practically does not affect the free energy for $L_x, L_y \gg 1$, it is important in relation to the underlying degeneracies with respect to x_0 and φ_0 . In fact, these degeneracies eventually determine the (i) *floating* of magnetic-domain patterns and (ii) the vanishing of critical current for DW motion in nanowires with no intermediate anisotropy ($D_x = 0$).

A. Positional order of domain walls

The fact that both \mathcal{Z}_{φ_0} and \mathcal{Z}_{x_0} arise from elasticlike Hamiltonians has important implications that deserve some further comment. As already pointed out, \mathcal{Z}_{x_0} is associated with corrugation of the DW as a function of y (see the sketch in Fig. 5). In the ideal case considered here, this instability destroys positional order (defined by the profile \bar{n}) at finite temperatures. In real films, positional order may be stabilized by some substrate anisotropy, pinning or dipolar interaction. The last one is responsible for the emergence of magnetic-domain patterns in films magnetized out of plane. In this context, Eq. (62) may be applied to study the temperature dependence of the optimal period of modulation for a stripe pattern, known to be the ground state [67–70] (gray and white

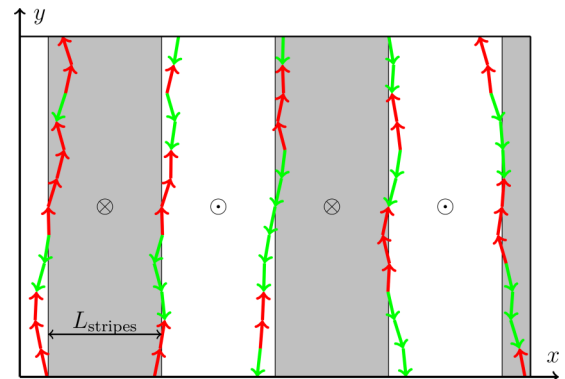


FIG. 5. (Color online) Schematic view of stripe pattern in a film with out-of-plane anisotropy. Gray and white areas represent domains with opposite magnetization in the ground state. Colored (online) arrows give a pictorial representation of DW corrugation and short-range chiral order at finite temperature. Note that Bloch DWs have been assumed and disorder in the arrows direction [$\varphi_0(y)$ variable] has been exaggerated to help visualization.

domains in Fig. 5). At zero temperature, a modulated phase results from the competition between magnetostatic energy (which favors in this case antiparallel alignment of spin pairs) and the energy cost to create DWs. Several approaches lead to the same scaling of the characteristic period of modulation [67,68,71]: $L_{\text{stripes}} \sim e^{\mathcal{E}_{\text{dw}}/(4\Omega\tau^2)}$, with $\Omega = \mu_0 M_s^2 a^3 / 4\pi$ being the strength of dipolar interaction, τ is the dimensional film thickness, a is the *dimensional* lattice constant, and M_s is the saturation magnetization (remember that in this section, it is $\mathcal{E}_{\text{dw}} = 2\tau\sqrt{2J_{\text{atom}}D_{\text{atom}}}$). A first-order estimate of thermal effects is obtained by replacing \mathcal{E}_{dw} in the expression of L_{stripes} with \mathcal{F}_{dw} defined in Eq. (62): $L_{\text{stripes}} \sim e^{\mathcal{F}_{\text{dw}}/(4\Omega\tau^2)}$. The latter is consistent with a decrease of the period of modulation with increasing temperature, as observed in experiments [38,39] and predicted by mean-field approach [40,71]. Note that in deducing the DW free energy \mathcal{F}_{dw} , dipolar interaction was totally overlooked. For instance, this interaction is known to produce an effective intermediate anisotropy that stabilizes Bloch type of DWs in films magnetized out of plane. Strictly speaking, since we assumed $D_x = 0$, our treatment should not apply to stripped patterns emerging in those films. A more accurate study of this specific phenomenon—which we intend to address in a forthcoming paper—shall possibly produce a different entropic contribution to the DW free energy. Nevertheless, the main message of this paragraph is expected to hold true: Gaussian fluctuations around a DW profile suffice to account for a decrease of L_{stripes} with increasing temperatures.

Still referring to domain patterns in films with dominant out-of-plane anisotropy, elastic deviations from the ideal stripe phase are known to yield anisotropic decay of correlations [72] along x and y . In the naive description of the stripe pattern given above, we ideally froze the elastic modes associated with compression along x . Yet, some disorder is expected to arise from the corrugation modes, developing along y and associated with \mathcal{Z}_{x_0} . Without entering the details, this fact already suggests the decay of spatial correlations to be more severe along y rather than along x . Consistently with this picture, several theoretical works [72–75] predict that the stripped ground state should evolve into a 2D smectic or nematic phase [76] at finite temperature. The degeneracy with respect to x_0 of a single DW embedded in a thin film propagates to more complex patterns and leads to the *floating-solid* description [the specific choice of $\vec{n}(x)$ being dependent on x *only* does not allow for rotational invariance on the xy plane, which in physical systems also occurs].

Usually, in real films all these effects are hindered by pinning. As related to the latter, the assumption—stated by Hamiltonian (54)—that DWs behave as elastic interfaces is the starting point for describing depinning and creep dynamics [77–79]. However, at relatively high temperatures, pinning may become negligible, thus restoring the idealized theoretical picture sketched above. The observation of stripe mobility in Fe/Cu(001) films, indeed, supports this scenario [38].

B. Chiral order of domain walls

The instability with respect to φ_0 is, instead, related to DW chirality. As pointed out in Appendix G, \mathcal{Z}_{φ_0} may be thought of as arising from the 1D XY Hamiltonian given in Eq. (G2) in the limit of small misalignment between neighboring

chirality vectors $\vec{e}_{\varphi_0}(y)$. It is well known that the 1D XY model—with nearest-neighbor interactions—can only sustain short-range order. Therefore our picture suggests that just short-range chiral order should develop along the y direction when only uniaxial anisotropy and exchange interaction are considered. An intermediate anisotropy $D_x \neq 0$ —for instance, of magnetostatic origin—stabilizes two possible values of φ_0 . Eventually, this drives the original XY Hamiltonian \mathcal{H}_{xy} in Eq. (G2), which describes the *effective* coupling between chirality vectors, towards the Ising universality class. Neither in this case long-range chiral order along y is expected to be stable. From a snapshot taken at finite temperature we would rather expect domains of opposite chirality to alternate *randomly* along, e.g., a Bloch DW [for which $\vec{e}_{\varphi_0} = (\pm 1, 0, 0)$]. This phenomenon was recently observed on Fe/Ni/Cu(001) films [6], consisting of ten monolayers of Ni and 1.3 of Fe. Our conjecture is sketched pictorially in Fig. 5. Arrows represent the alternating magnetization direction along DWs, instead of the \vec{e}_{φ_0} vectors, to facilitate the comparison with experiments reported in Ref. [6]. A scenario consistent with chiral order of DWs requires either a significant film thickness (so that dimensional crossover may occur) or the presence of a Dzyaloshinskii-Moriya interaction [80]. The latter, still in Ref. [6], was observed to stabilize homochiral Néel type of DWs, with $\vec{e}_{\varphi_0} = (0, 1, 0)$, in samples with thinner Ni interlayer.

Our considerations about short-range chiral order of DWs seem in striking contrast with Villain’s conjecture [81] (confirmed by experiments on Gd-based spin chains [82]). This prescribes that in spin chains that develop short-range chiral order and are packed in a 3D crystal long-range chiral order should set in at higher temperature than magnetic ordering. In our mindset, the “information” about chirality of DWs in a 2D stripe-domain pattern cannot propagate along x , from one DW to the next because they are separated by regions in which all spins are aligned along the easy axis. Some misalignment between neighboring spins is, instead, needed to have a finite chiral order parameter. This ceases to hold true, e.g., close to the spin-reorientation transition where a canted-stripe phase was recently observed in 2D simulations [66,83].

One should not forget that the arguments provided here strictly rely on thermodynamic equilibrium. Homochirality of DWs may be observed in experiments [84–86] and simulations as a result of slow dynamics [87–89], similarly to what happens in superparamagnetic nanoparticles [57,90,91], for which long-range ferromagnetic order would be forbidden by equilibrium thermodynamics.

Chirality is also related to adiabatic spin transfer torque (STT), through which an electric current may displace a DW hosted in a ferromagnetic nanowire. Translation of the DW, i.e., a variation of the x_0 parameter, is in this case necessarily accompanied by a precession of the φ_0 angle, which produces a periodic change of the DW structure between Bloch and Néel type. The corresponding Landau-Lifshitz-Gilbert equation reads

$$\frac{d\vec{S}}{dt} = \gamma_0 \vec{H}_{\text{eff}} \times \vec{S} + \alpha_G \vec{S} \times \frac{d\vec{S}}{dt} - u \partial_x \vec{S}, \quad (63)$$

where the effective field is defined from the Hamiltonian \mathcal{H} in Eq. (45) (possibly modified to include the magnetostatic

energy or an intermediate anisotropy) as

$$\vec{H}_{\text{eff}} = -\frac{1}{g\mu_B} \frac{\delta\mathcal{H}}{\delta\vec{S}}, \quad (64)$$

γ_0 is the gyromagnetic ratio, α_G is the Gilbert damping, and the last term accounts for the adiabatic STT with

$$u = \frac{g\mu_B J_e P}{2eM_s a}, \quad (65)$$

where J_e represents the electrical current density and P is the polarization factor of the current [the lattice unit a has been added because the derivative in Eq. (63) is considered dimensionless]. The adiabatic STT term in Eq. (63) can be regarded as a precessional contribution about an effective “field” $\vec{H}_{\text{STT}} = -u(\vec{S} \times \partial_x \vec{S})/\gamma_0$, which is proportional to the local chirality vector [46]. Due to its close relation with chirality, we believe that reconsidering adiabatic STT in the perspective of short-range chiral order may contribute to shed some light on the puzzling scenario of DW motion induced by electric currents. For instance, adiabatic STT seems to catch the main physics of Co/Ni nanowires magnetized out of plane [92,93], while it largely fails for prototypical Permalloy nanowires (magnetized in plane) [94–96]; the critical current for DW motion is observed to depend strongly on temperature [97,98] in some samples and not in others [99]. Even if the extension to 2D of the sketch in Fig. 1 naturally leads to films or nanowires magnetized out-of-plane, our calculation can be adapted to samples magnetized in plane by a permutation of coordinates [36,100]. Within our picture, the ratio between the correlation length characterizing short-range chiral order (ξ_{φ_0}) and the actual transverse size of a sample (L_y) should discriminate two regimes: for $\xi_{\varphi_0} \ll L_y$ the response of a DW to adiabatic STT is expected to depend strongly on temperature; for $\xi_{\varphi_0} \gg L_y$ this dependence should be much less dramatic. Also in this context, it is worth remarking that thermally assisted DW depinning [16,101] and possible nonhomogeneous mechanisms of precession [102] have been neglected in our considerations.

V. CONCLUSIONS

We considered the effect of thermalized linear excitations about a DW profile. Expressions for the free energy of a DW embedded in 1D or 2D lattices were derived as a function of temperature and the system size. This was achieved by rephrasing in the language of Polyakov renormalization [34,41,42] some known results, obtained from linearization of the Landau-Lifshitz equation [25–30]. Our approach is equivalent to the steepest-descent approximation of functional integrals [20]. It has, in our opinion, the advantage of allowing for an easier generalization to 2D. Moreover, it provides a better insight on the role of fast- and slow-varying degrees of freedom, while keeping track of the nonhomogeneity within the fluctuation field. For instance, it is straightforward to realize that fluctuations associated with bound states shall be localized at the DW center. This information might be relevant to the aim of accounting efficiently for thermal fluctuations in the Landau-Lifshitz-Gilbert equation [103], beyond the mean-field level (Landau-Lifshitz-Bloch equation [14,18]).

From the knowledge of the DW free energy, we provided a phenomenological expansion for the correlation length that may be used to fit the susceptibility of single-chain magnets (slow-relaxing spin chains [31–35]). The last ones are often realized by creating a preferential path for the exchange interaction between anisotropic magnetic units—consisting of transition metals or rare earths—through an organic radical. Thus, to some extent, Eq. (1) can be considered a reference Hamiltonian for single-chain magnets in general [34,35].

In a previous work [40], we explained the shrinking of magnetic domains, observed in films with out-of-plane anisotropy, within a mean-field approach and assuming a nonhomogeneous spin profile. Retaining the last feature for the unperturbed profile, in this paper, we showed that Gaussian fluctuations also lead to a qualitatively similar result.

As a further implication, our model suggests that long-range chiral order cannot occur within DWs interposed between saturated domains. The robustness of this conjecture certainly deserves to be checked *beyond* the Gaussian approximation. Yet, it seems consistent with recent experiments on Fe/Ni/Cu(001) films [6]. This softening of chiral order may acquire some relevance also in view of DW manipulations by means of spin-transfer torque.

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APPENDIX A: LOWERING AND RAISING OPERATORS

In the next two appendices, we provide some details about the solution of the eigenvalue problem (6) in the presence of $\uparrow\downarrow$ b.c.. In terms of the lowering and raising operators

$$\begin{aligned} \hat{a}_m &= \partial_\eta + m \tanh \eta, \\ \hat{a}_m^\dagger &= -\partial_\eta + m \tanh \eta, \end{aligned} \quad (A1)$$

the generalized fluctuation Hamiltonian given in Eq. (23) reads $\hat{H}_m = \hat{a}_m^\dagger \hat{a}_m + m$. Using the commutator relation

$$[\hat{a}_m^\dagger, \hat{a}_m] = -\frac{2m}{\cosh^2 \eta}, \quad (A2)$$

one finds the alternative representation $\hat{H}_m = \hat{a}_{m+1} \hat{a}_{m+1}^\dagger - (m+1)$. We now show that if $|\psi\rangle_m$ is an eigenstate of \hat{H}_m then $\hat{a}_{m+1}^\dagger |\psi\rangle_m$ is an eigenstate of \hat{H}_{m+1} . Let $|\psi\rangle_m$ be an eigenstate of $\hat{a}_{m+1} \hat{a}_{m+1}^\dagger$ with eigenvalue λ_m , i.e., $\hat{a}_{m+1} \hat{a}_{m+1}^\dagger |\psi\rangle_m = \lambda_m |\psi\rangle_m$, then

$$\hat{a}_{m+1}^\dagger (\hat{a}_{m+1} \hat{a}_{m+1}^\dagger |\psi\rangle_m) = \hat{a}_{m+1}^\dagger \lambda_m |\psi\rangle_m \quad (A3)$$

$$= \lambda_m (\hat{a}_{m+1}^\dagger |\psi\rangle_m), \quad (A4)$$

which means that $\hat{a}_{m+1}^\dagger |\psi\rangle_m$ is an eigenstate of $\hat{a}_{m+1}^\dagger \hat{a}_{m+1}$ with the same eigenvalue λ_m . With the above representations of \hat{H}_m and \hat{H}_{m+1} , it can easily be checked that if $|\psi\rangle_m$ is an eigenstate

of \hat{H}_m with eigenvalue \mathcal{E}_m then $\hat{a}_{m+1}^\dagger |\psi\rangle_m$ is an eigenstate of \hat{H}_{m+1} with eigenvalue $\mathcal{E}_m + 2(m+1)$. Thus the Hamiltonians \hat{H}_m and \hat{H}_{m+1} share the same spectrum, except for the fact that \hat{H}_{m+1} has the additional eigenstate $|0\rangle_{m+1}$, defined by $\hat{a}_{m+1}|0\rangle_{m+1} = 0$, which is not present in the spectrum of \hat{H}_m . This property allows constructing eigenstates of the Hamilton operator \hat{H}_m recursively. For the case of our interest, we stop the iteration at the first step $m = 1$, namely we just apply \hat{a}_1^\dagger to the free-particle eigenstates (see the following appendix for the explicit calculation). The missing *vacuum* state $|0\rangle_1$ is nothing but the bound state represented through Eq. (25) in real space.

APPENDIX B: DERIVATION OF THE DENSITY OF STATES

In the finite system, the solutions of the free-particle Hamiltonian are $\psi_0(q, x) = A \cos(qx) + B \sin(qx)$, where A, B are constants to be determined from the boundary and normalization conditions. The *free-state* solutions of the Schrödinger-like Eq. (6) obtained for antiferromagnetic ($\uparrow\downarrow$) b.c. [namely, when $\vec{n}(x)$ describes a DW profile] can be obtained by applying the raising operator \hat{a}_1^\dagger defined in Eq. (24) to $\psi_0(q, x)$:

$$\begin{aligned} \Phi_{a,q}(x) = & A \{q \sin(qx)/c + \tanh[c(x - x_0)] \cos(qx)\} \\ & + B \{\tanh[c(x - x_0)] \sin(qx) - q \cos(qx)/c\}, \end{aligned} \quad (\text{B1})$$

with $a = x_0, \varphi_0$. Independently of the determination of A and B , it is straightforward to show that $\Phi_{a,q}(x)$ are solutions of the Schrödinger-like Eq. (6) with eigenvalues

$$\varepsilon_{a,q} = \frac{J}{2} q^2 + D, \quad (\text{B2})$$

$$q(L+1) = \alpha + p\pi \quad \text{with } p \in \mathbb{Z}, \quad (\text{B7})$$

$$\alpha = \arctan \left[\frac{q \{\tanh(cx_0) + \tanh[c(L+1-x_0)]\}}{c \{\tanh(cx_0) \tanh[c(L+1-x_0)] - q^2/c^2\}} \right],$$

which can be solved for p to obtain

$$p = \frac{q(L+1)}{\pi} - \frac{1}{\pi} \arctan \left[\frac{q \{\tanh(cx_0) + \tanh[c(L+1-x_0)]\}}{c \{\tanh(cx_0) \tanh[c(L+1-x_0)] - q^2/c^2\}} \right]. \quad (\text{B8})$$

The density of states [44] is defined as

$$\begin{aligned} \rho(q) = \frac{dp}{dq} = & \frac{L+1}{\pi} - \frac{c}{\pi} \left\{ \tanh(cx_0) \tanh[c(L+1-x_0)] + \frac{q^2}{c^2} \right\} \frac{\tanh(cx_0) + \tanh[c(L+1-x_0)]}{\Delta(q, x_0)}, \quad \text{with} \\ \Delta(q, x_0) = & c^2 \left\{ \tanh(cx_0) \tanh[c(L+1-x_0)] - \frac{q^2}{c^2} \right\}^2 + q^2 \{\tanh(cx_0) + \tanh[c(L+1-x_0)]\}^2. \end{aligned} \quad (\text{B9})$$

In order to compute the DW free energy analytically, we assume the DW to be centered in the system, i.e., $x_0 = (L+1)/2$. In this case, Eq. (B5) reads

$$\tan[q(L+1)] = \frac{2q \tanh[c(L+1)/2]}{c \{\tanh^2[c(L+1)/2] - q^2/c^2\}}. \quad (\text{B10})$$

formally equal to those obtained starting from a uniform profile ($\uparrow\uparrow$ b.c.). We show in the following that the allowed values of q are not the same, which affects the density of states.

The constants A and B in Eq. (B1) have to be determined from the boundary conditions $\Phi_{a,q}(0) = \Phi_{a,q}(L+1) = 0$ and the normalization condition

$$\int_0^{L+1} dx |\Phi_{a,q}(x)|^2 = 1. \quad (\text{B3})$$

Though deriving an analytic expression for A and B is computationally demanding, the density of states can easily be obtained. Nontrivial solutions ($A \neq 0$ and $B \neq 0$) exist only if the determinant of the system defined by the boundary conditions $\Phi_{a,q}(0) = \Phi_{a,q}(L+1) = 0$ vanishes:

$$\begin{aligned} & \tanh(cx_0) \tanh[c(L+1-x_0)] \tan[q(L+1)] \\ & - q \tanh(cx_0)/c - q^2 \tan[q(L+1)]/c^2 \\ & - q \tanh[c(L+1-x_0)]/c = 0, \end{aligned} \quad (\text{B4})$$

which gives the following transcendental equation for the determination of the possible q values:

$$\tan[q(L+1)] = \frac{q \{\tanh(cx_0) + \tanh[c(L+1-x_0)]\}}{c \{\tanh(cx_0) \tanh[c(L+1-x_0)] - q^2/c^2\}}. \quad (\text{B5})$$

This equation can be solved using, e.g., graphical methods. More significantly, it allows deriving an analytic expression for the density of states. To this purpose, let us set

$$\tan \alpha = \frac{q \{\tanh(cx_0) + \tanh[c(L+1-x_0)]\}}{c \{\tanh(cx_0) \tanh[c(L+1-x_0)] - q^2/c^2\}}. \quad (\text{B6})$$

The solution of Eq. (B5) is

Setting $\tan \alpha = q/c \tanh[c(L+1)/2]$, Eq. (B10) takes the form

$$\tan[q(L+1)] = \tan(2\alpha). \quad (\text{B11})$$

The density of states simplifies to

$$\rho(q) = \frac{L+1}{\pi} - \frac{2c \tanh[c(L+1)/2]}{\pi \{c^2 \tanh^2[c(L+1)/2] + q^2\}}, \quad (\text{B12})$$

and for $L \gg 1/c$, $\tanh^2[c(L+1)/2] \approx 1$,

$$\rho(q) = \frac{L+1}{\pi} - \frac{2c}{\pi(c^2 + q^2)}. \quad (\text{B13})$$

In order to check the validity of Eq. (B12), we performed a numerical diagonalization of Eq. (6) for a finite, discrete lattice (the excellent agreement is summarized in Fig. 2).

Diagonalization of the discrete eigenvalue problem. The discrete version of Eq. (6) is given by

$$\sum_{j,k} \psi(j) \left[-\frac{J}{2} \Delta_{j,k} + V(k) \delta_{j,k} \right] \psi(k) = \varepsilon, \quad (\text{B14})$$

where $\Delta_{j,k}$ is the discrete Laplace operator

$$\Delta_{j,k} = \begin{cases} -2, & j = k, \\ 1, & k = j \pm 1, \\ 0, & \text{else,} \end{cases} \quad (\text{B15})$$

the potential is

$$V(k) = D[2 \cos^2(\theta_k) - 1], \quad \text{for } \uparrow\downarrow \text{ b.c.,} \quad (\text{B16})$$

$$V(k) = D, \quad \text{for } \uparrow\uparrow \text{ b.c.,}$$

and θ_k corresponds to the discrete profile computed, e.g., with the nonlinear map method (see Appendix F). A complete orthonormal system fulfilling the boundary conditions $\psi(0) = \psi(L+1) = 0$ is

$$u_n(j) = \sqrt{\frac{2}{L+1}} \sin(q_n j), \quad q_n = \frac{\pi}{L+1} n \quad (\text{B17})$$

with $0 \leq j \leq L+1$ and $1 \leq n \leq L$. The functions ψ are expanded in this basis,

$$\psi(j) = \sum_n w_n u_n(j), \quad (\text{B18})$$

and inserted into Eq. (B14) to get

$$\sum_{n,m} w_n \{ J[1 - \cos(q_n)] \delta_{n,m} + M_{n,m} \} w_m = \varepsilon, \quad (\text{B19})$$

with

$$M_{n,m} = \sum_j u_n(j) V(j) u_m(j). \quad (\text{B20})$$

This is equivalent to solving the eigenvalue problem

$$\sum_m \{ J[1 - \cos(q_m)] \delta_{n,m} + M_{n,m} \} w_m = \varepsilon_n w_n. \quad (\text{B21})$$

It is worth noting that in a finite discrete system the possible number of eigenfunctions L is finite. As a consequence, every time a domain wall is added, a free state is ‘‘lost’’ and a bound state is ‘‘gained.’’ However, this procedure is correct only when the distance between DWs is large enough to treat them independently. In our case, we just considered one or no DW in the system. When no DW is present, i.e., for $\uparrow\uparrow$ b.c., $u_n(j)$ are solutions to the eigenvalue problem in Eq. (B14) with eigenvalues

$$\varepsilon_{\uparrow\uparrow,n} = J[1 - \cos(q_n)] + D. \quad (\text{B22})$$

As already mentioned, the dispersion relation is expected to depend on q in the same way for both choices of b.c., but the

allowed values of q may, generally, be different. The spectrum in Eq. (B22) was compared with that obtained by solving the eigenvalue problem (B21) numerically: ε_n obtained in both cases were plotted against the eigenvalue index n . Indeed, the two dispersion relations turned out to overlap. Therefore the ‘‘shifted’’ $q_{\uparrow\downarrow,n}$, corresponding to $\uparrow\downarrow$ b.c., could be deduced through the following formula:

$$q_{\uparrow\downarrow,n} = \arccos \left(\frac{J + D - \varepsilon_{\uparrow\downarrow,n}}{J} \right). \quad (\text{B23})$$

Eventually, the density of states plotted with symbols in Fig. 2 was obtained with the discrete derivative

$$\rho(q) = \frac{\partial n}{\partial q} \simeq \frac{1}{q_{\uparrow\downarrow,n+1} - q_{\uparrow\downarrow,n}}. \quad (\text{B24})$$

APPENDIX C: POLYAKOV RENORMALIZATION

When ferromagnetic b.c. ($\uparrow\uparrow$) are considered, the expansion in Eq. (7) involves only free states $\Phi_{a,q}(x)$, those given in Eq. (16). Thermal averages of the coefficients $a_{a,q}$ and their products are momenta of the Gaussian integrals entering the partition function $\mathcal{Z}[\vec{n}]$ and, therefore, fulfill the relations

$$\langle a_{a,q} \rangle = 0, \quad (\text{C1})$$

$$\langle a_{a,q} a_{a,q'} \rangle = \frac{T}{2 \varepsilon_{a,q}} \delta_{q,q'}$$

(equipartition theorem). From this, it follows that thermal averages of the fluctuation-field components read

$$\begin{aligned} \langle \phi_a^2 \rangle_{\uparrow\uparrow} &= \frac{T}{L+1} \sum_q \frac{\sin^2(qx)}{\varepsilon_{a,q}} \\ &\simeq \frac{T}{2(L+1)} \int_{-\infty}^{\infty} \frac{\sin^2(qx)}{\varepsilon_{a,q}} \rho_{\uparrow\uparrow} dq \\ &= \frac{T}{2\pi J} \int_{-\infty}^{\infty} \frac{1 - \cos(2qx)}{q^2 + c^2} dq \\ &= \frac{T}{\mathcal{E}_{\text{dw}}} [1 - e^{-2cx} - e^{-2c(L+1-x)}], \end{aligned} \quad (\text{C2})$$

the first two terms of the last line come directly from integration on the complex plane, while the third one has been added by hand for symmetry with respect to the boundaries $x = 0$ and $x = L+1$. This contribution is lost when the summation \sum_q is approximated with an integral (first line). Starting from Eq. (12), it can be shown [34,42] that the following relations hold for thermal averages:

$$\langle (\partial_x \vec{S})^2 \rangle = (1 - \langle \phi_a^2 \rangle_{\uparrow\uparrow}) (\partial_x \vec{n})^2 + \sum_a \langle (\partial_x \phi_a)^2 \rangle_{\uparrow\uparrow}, \quad (\text{C3})$$

$$\langle (S^z(x))^2 \rangle = (1 - 3 \langle \phi_a^2 \rangle_{\uparrow\uparrow}) (n^z)^2 + \langle \phi_a^2 \rangle_{\uparrow\uparrow}.$$

In Fig. 6, we compare $\langle (S^z(x))^2 \rangle$ obtained by inserting the analytic formula for $\langle \phi_a^2 \rangle_{\uparrow\uparrow}$ given in Eq. (C2) with numerical, transfer-matrix calculations. The analytic expression agrees very well with numerics at low temperature. For an infinite system, the Gaussian approximation would give $\langle \phi_a^2 \rangle_{\infty} = T/\mathcal{E}_{\text{dw}}$, which is indeed recovered away from the boundaries (not shown). The more refined, Polyakov approach consists

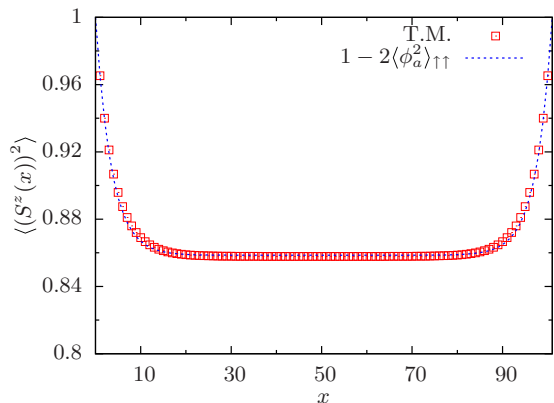


FIG. 6. (Color online) $\langle (S^z(x))^2 \rangle$ obtained for $\uparrow\uparrow$ b.c.: transfer-matrix calculation (symbols, T.M. in the legend) and analytic result (line) obtained combining Eq. (C2) with Eq. (C3): $\langle (S^z(x))^2 \rangle = 1 - 2\langle \phi_a^2 \rangle_{\uparrow\uparrow}$. Calculation parameters are $D/J = 0.01$, $T/J = 0.02$, $L = 100$ spins.

in integrating the free-state degrees of freedom progressively, starting from fluctuations associated with shorter spatial scales. Consistently with Eq. (C3), this leads to renormalization of the spin-Hamiltonian parameters [41,42]:

$$\begin{aligned} J &\rightarrow J(1 - \langle \phi_{dq}^2 \rangle), \\ D &\rightarrow D(1 - 3\langle \phi_{dq}^2 \rangle), \end{aligned} \quad (\text{C4})$$

with

$$\langle \phi_{dq}^2 \rangle = \frac{T}{2\pi J} \frac{1}{q^2 + c^2} dq. \quad (\text{C5})$$

The term c^2 at the denominator is usually neglected to facilitate the analytic treatment; the error due to this approximation will be compensated by a proper choice of cutoff at which the renormalization procedure stops. It is convenient to pass to the real space through the substitution $\lambda = 1/q$ so that the renormalization flux takes the form

$$\begin{aligned} J(\lambda + d\lambda) &= J(\lambda) \left(1 - \frac{T}{2\pi J(\lambda)} d\lambda \right), \\ D(\lambda + d\lambda) &= D(\lambda) \left(1 - 3 \frac{T}{2\pi J(\lambda)} d\lambda \right). \end{aligned} \quad (\text{C6})$$

The integration of the first differential equation is straightforward and gives $J(\lambda) = JZ(\lambda)$, with $Z(\lambda) = 1 - T\lambda/(2\pi J)$ and J being the original, bare exchange constant appearing in Hamiltonian (1). Taking into account that $dZ(\lambda) = -T d\lambda/(2\pi J)$, the second equation takes the form

$$\frac{dD(\lambda)}{D(\lambda)} = 3 \frac{dZ(\lambda)}{Z(\lambda)} \Rightarrow D(\lambda) = DZ^3(\lambda), \quad (\text{C7})$$

where, again, D is the original, bare anisotropy constant and $D(\lambda)$ the renormalized one. The cut-off λ_c at which the renormalization stops can be chosen in order that the Gaussian result for an infinite system, $\langle \phi_a^2 \rangle_\infty = T/\mathcal{E}_{\text{dw}}$, is recovered at low temperature, namely, $\lambda_c = 2\pi J/\mathcal{E}_{\text{dw}} = \pi/c$. The whole renormalization is meaningful as long as $Z(\lambda_c) > 0$, that is for $T < \mathcal{E}_{\text{dw}}$. This is not a problem because the requirement that

the DW free energy be positive, $\Delta F \gtrsim 0$, is generally more strict.

APPENDIX D: SCALING OF THE CORRELATION LENGTH

In this appendix, we describe the way in which a semi-analytical expression for the correlation length was deduced. As mentioned in the main text and shown in Fig. 3, a better agreement between ΔF given in Eq. (33) and numerical results is obtained by introducing the renormalized parameters $D(\lambda_c)$ and $c(\lambda_c)$ *only* in the argument of the logarithm. Considering that $D(\lambda_c) = DZ^3(\lambda_c)$ and $c(\lambda_c) = cZ(\lambda_c)$, after this substitution, the DW free energy takes the form

$$\Delta F = \mathcal{E}_{\text{dw}} - T \ln(\beta \mathcal{E}_{\text{dw}} c L Z^3 \{1 + \tanh[cZ(L+1)/2]\}^2). \quad (\text{D1})$$

We are interested in the value of $L = \bar{L}$ for which $\Delta F(\bar{L}) = 0$. Let us set $\xi = c\bar{L}$ and $\zeta = \beta \mathcal{E}_{\text{dw}}$ so that our problem reduces to an implicit equation in the variables ξ and ζ . Then consider the following map in the ξ variable:

$$\begin{aligned} \Xi_{it} &= \frac{Z(\zeta)}{2} (\xi_{it} + c), \\ \xi_{it+1} &= \frac{e^\zeta}{\zeta Z^3(\zeta)} [1 + \tanh(\Xi_{it})]^{-2}, \end{aligned} \quad (\text{D2})$$

which is parametrically dependent on ζ , with $it = 1, 2, \dots$ representing the iteration index. From its definition, it follows that $Z(\lambda_c) = Z(\zeta) = 1 - 1/\zeta$. The initial condition of the map (D2) is set to

$$\xi_1 = \frac{e^\zeta}{4\zeta Z^3(\zeta)}. \quad (\text{D3})$$

The fixed point ξ_∞ gives the sought for solution $\Delta F(\bar{L}) = 0$. For $\xi_1 \gg 1$, it is $\Xi_1 \gg 1$ and $\tanh(\Xi_1) \simeq 1$, consequently. In this case, the map converges already at the first iteration, namely $\xi_1 \simeq \xi_\infty$. Numerically, one finds that the latter equivalence holds within 1.4% for any $D/J > 0.1$.

To extend the validity of Eq. (D3) to high temperatures, we used the trick—not fully justified—that Polyakov renormalization somehow stops when λ_c becomes of the order of the actual number of spins aligned along the same direction: $\bar{L}/2 - 3/2c$. This number is evaluated placing, ideally, the DW in the middle of a segment of length \bar{L} ; the contribution of $1/c$ misaligned spins lying roughly on half DW is subtracted as well as that of $1/(2c)$ “blocked” spins at one boundary [consistently with Eq. (C2)]. In terms of the reduced variable entering the map (D2), $\lambda_c = \bar{L}/2 - 3/2c$ when $\xi_\infty = 2\pi + 3$, which occurs at $\zeta = 4.262$. We are now in the position to give an expression for the scaling function $f(\zeta)$ describing the *universal* behavior of the product $c\xi$, which is expected to scale like $c\bar{L}$, as a function of the scaling variable $\zeta = \beta \mathcal{E}_{\text{dw}}$:

$$f(\zeta) = \begin{cases} \frac{e^\zeta}{4\zeta Z^3(\zeta)} & \text{for } \zeta > 4.262, \\ A_0 e^{b\zeta+1} & \text{otherwise,} \end{cases} \quad (\text{D4})$$

with constants $A_0 = 0.3282$ and $b = 0.5496$ determined by requiring the continuity of ξ_1 and its derivative with respect to ζ . The exact relation between the scaling function $f(\zeta)$ and ξ computed numerically can only be obtained by fitting

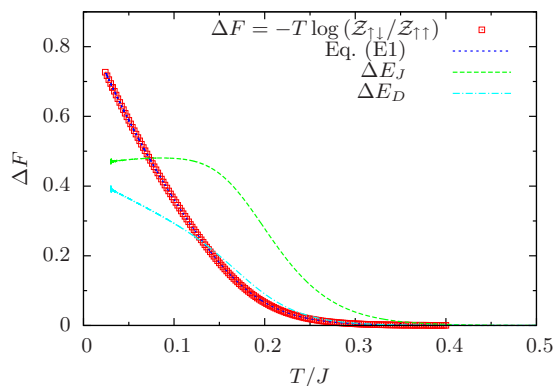


FIG. 7. (Color online) Transfer-matrix calculation. Domain-wall free energy as a function of temperature obtained with $\Delta F = -T \ln(\mathcal{Z}_{\uparrow\downarrow}/\mathcal{Z}_{\uparrow\uparrow})$ (squares) or by means of Eq. (E1) (blue, short-dashed). The averaged differences ΔE_J (exchange energy, upper green curve) and ΔE_D (anisotropy energy, lower magenta curve) are also plotted. Computation parameters are $L = 20$ and $D/J = 0.1$.

a constant prefactor (see Fig. 3); this turns out to be of the order of the Euler constant, therefore we can reasonably set the equivalence $c\xi \simeq f(\zeta)/e$.

APPENDIX E: ALTERNATIVE CALCULATION OF THE DW FREE ENERGY IN A SPIN CHAIN

Given the general relation $\partial(\beta F)/\partial\beta = \langle \mathcal{H} \rangle$, the free energy is usually evaluated as

$$F(\beta) = \frac{\beta_0}{\beta} F(\beta_0) + \frac{1}{\beta} \int_{\beta_0}^{\beta} \langle \mathcal{H} \rangle d\beta \quad (\text{E1})$$

in Monte Carlo [104] or stochastic-LLG simulations [17,18]. As suggested in Ref. [17], the DW free energy can be computed with the same approach provided that integration is started from a temperature at which $F(\beta_0)$ is independent of b.c. In order to prove the results presented in Fig. 3, we repeated the calculation of ΔF following this approach. The initial temperature (in J units) was chosen as $1/\beta_0 = 0.6$, for which the difference between the average energy $\langle \mathcal{H} \rangle$ computed for $\uparrow\uparrow$ and $\uparrow\downarrow$ b.c. is about $2 \times 10^{-5} J$. The free energy computed integrating $\langle \mathcal{H} \rangle_{\uparrow\downarrow} - \langle \mathcal{H} \rangle_{\uparrow\uparrow}$ for $L = 20$ is displayed in Fig. 7 as a (blue) short-dashed line. It clearly overlaps the points shown in Fig. 3 and replotted here for convenience. It is worth mentioning that while the accuracy of Eq. (E1) is bound to repeating the computation at sufficiently small temperature steps ($d\beta = 0.01$ for the present case), the formula $\Delta F = -T \ln(\mathcal{Z}_{\uparrow\downarrow}/\mathcal{Z}_{\uparrow\uparrow})$ can be used independently of any constraint. In Fig. 7 the two independent contributions to $\langle \mathcal{H} \rangle_{\uparrow\downarrow} - \langle \mathcal{H} \rangle_{\uparrow\uparrow}$ coming from the exchange (ΔE_J) and anisotropy (ΔE_D) energies are also plotted. The irregular behavior observed at the lowest temperatures is associated with numerical instabilities, typical of the transfer-matrix technique in the limit $T \rightarrow 0$.

APPENDIX F: SPIN PROFILE IN LATERALLY CONSTRAINED DOMAIN WALLS

Here, we propose again an analytic derivation of the DW profile for finite systems [46,105], which employs a continuum

formalism, and check its results against numerical calculations on a discrete lattice. We first consider the DW profile for an infinite chain with $\uparrow\downarrow$ b.c.. Since for $T = 0$ one has $\vec{n}(x) = \vec{S}(x)$, the minimum-energy profile given in Eq. (18) can be obtained from Hamiltonian (2). The latter in polar coordinates $\vec{S} \equiv (\sin\theta \cos\varphi, \sin\theta \sin\varphi, \cos\theta)$ reads

$$\begin{aligned} \mathcal{H} = & \int \frac{J}{2} [(\partial_x \theta)^2 + \sin^2(\theta) (\partial_x \varphi)^2] dx \\ & - \int [D \cos^2(\theta) + D_x \sin^2(\theta) \cos^2(\varphi)] dx + \text{const.} \end{aligned} \quad (\text{F1})$$

The intermediate anisotropy D_x such that $0 < |D_x| < |D|$ has been introduced for the sake of generality. The profile that minimizes the functional in Eq. (F1) with respect to $\theta(x)$ and $\varphi(x)$ is the solution to the following Euler–Lagrange equations:

$$\begin{aligned} J \partial_x^2 \theta &= J \sin\theta \cos\theta (\partial_x \varphi)^2 + 2\tilde{D} \cos\theta \sin\theta, \\ J \sin^2(\theta) \partial_x^2 \varphi &+ 2J \sin\theta \cos\theta \partial_x \theta \partial_x \varphi \\ &= 2D_x \sin^2(\theta) \sin\varphi \cos\varphi \end{aligned} \quad (\text{F2})$$

[where $\tilde{D} = D - D_x \cos^2(\varphi)$], compatible with $\uparrow\downarrow$ b.c., that is,

$$\begin{aligned} \cos[\theta(x)] &= -\tanh[c(x - x_0)], \\ \varphi(x) &= \varphi_0 = 0, \frac{\pi}{2}, \pi, \frac{3\pi}{2}, \end{aligned} \quad (\text{F3})$$

with the more general $c = \sqrt{2\tilde{D}J}$ than in the main text. The choice of φ_0 depends on the sign of the intermediate anisotropy. For $D_x > 0$, one has $\varphi_0 = 0, \pi$ (Néel DW), while for $D_x < 0$, it is $\varphi_0 = \pi/2, 3\pi/2$ (Bloch DW). Note that in both cases there exist two degenerate solutions that correspond to opposite chirality of the DW [46]. The energy associated with the profile (F3)—with respect to a uniform ground state—is $\mathcal{E}_{\text{dw}} = 2\sqrt{2\tilde{D}J} - D_x$.

A helpful, well-established [21] analogy consists in interpreting x as “time” and θ as “spatial coordinate.” Then the first Euler-Lagrange equation in Eq. (F2) describes the motion of a classical particle of “mass” J moving in a potential $V(\theta) = D \cos^2(\theta) + D_x \sin^2(\theta) \cos^2(\varphi_0)$ ($\varphi = \varphi_0$ has been assumed). The “particle energy”

$$\varepsilon = \frac{J}{2} (\partial_x \theta)^2 + \frac{\tilde{D}}{2} \cos(2\theta) \quad (\text{F4})$$

is a constant of integration and it is univocally defined by the b.c.

$$\begin{aligned} \theta(0) &= 0, \\ \theta(L+1) &= \pi, \end{aligned} \quad (\text{F5})$$

consistent with Fig. 1. For the infinite chain $\theta(-\infty) = 0$ and $\theta(\infty) = \pi$, which yields $\varepsilon = D/2$, namely, $\partial_x \theta = 0$ at the boundaries. For finite chains, instead, one expects $\varepsilon > D/2$, i.e., a nonvanishing derivative at the boundaries. We will see that this condition is necessary to develop a continuum model for finite chains [it guarantees the convergence of the elliptic integrals in Eqs. (F6) and (F7)]. Integration of Eq. (F4) gives

the following implicit equation for the DW profile:

$$x = \sqrt{\frac{J}{\tilde{D}}} \int_0^\theta \frac{d\theta'}{\sqrt{\frac{2\varepsilon}{\tilde{D}} - \cos(2\theta')}}}, \quad (\text{F6})$$

where the b.c. $\theta(0) = 0$ has been used. The second b.c. in Eq. (F5) implies

$$L + 1 = 2\sqrt{\frac{J}{2\varepsilon + \tilde{D}}} \mathcal{K}\left(\sqrt{\frac{2\tilde{D}}{2\varepsilon + \tilde{D}}}\right), \quad (\text{F7})$$

\mathcal{K} being the complete elliptic integral of the first kind. This equation can be solved numerically to deduce ε . The whole model holds for broad DWs, $c \ll 1$, when the continuum limit is appropriate. Numerical instabilities are expected for $\varepsilon \simeq \tilde{D}/2$, when integrals diverge. Assuming that the DW is centered at $(L + 1)/2$ with $\theta[(L + 1)/2] = \pi/2$, the spin profile can also be computed as

$$x = \frac{L + 1}{2} - \sqrt{\frac{J}{\tilde{D}}} \int_\theta^{\pi/2} \frac{d\theta'}{\sqrt{\frac{2\varepsilon}{\tilde{D}} - \cos(2\theta')}}}, \quad (\text{F8})$$

which solves the possible numerical divergence at the boundaries. Though Eq. (F8) only applies to $x \leq (L + 1)/2$, the other half of the profile can be deduced from the property $\theta(L + 1 - x) = \pi - \theta(x)$. For very short chains, the constant of integration is approximately $\varepsilon = J(\partial_x \theta)^2/2$ [meaning that in Eq. (F4) the derivative dominates over the cosine term also at boundaries]. As a consequence, the condition (F7) takes the form

$$L + 1 = \sqrt{\frac{J}{2\varepsilon}} \int_0^\pi d\theta, \quad (\text{F9})$$

which gives $\varepsilon = J(\pi/(L + 1))^2/2$, i.e., the derivative $\partial_x \theta$ is constant over the whole chain and takes the value of the smallest wave vector q allowed in the reciprocal space. The spin profile can be computed analytically:

$$\theta(x) = \frac{\pi}{L + 1}x \Rightarrow S^z(x) = \cos\left(\frac{\pi}{L + 1}x\right). \quad (\text{F10})$$

The above profile corresponds to a single harmonic. In Fig. 8, the spin profile obtained with different methods (at $T = 0$) for two different chain lengths is shown. The harmonic approximation is satisfactory for $L = 6$ but, as expected, does not reproduce the discrete-lattice calculation for $L = 150$ (not shown).

Discrete-lattice calculation. In the case of a discrete chain, a recursive formula (nonlinear map) for the computation of the spin profile was proposed in Refs. [63–65]. The starting point is the DW Hamiltonian

$$\mathcal{H}_{\text{dw}} = J \sum_{i=0}^L [1 - \cos(\theta_{i+1} - \theta_i)] + \sum_{i=0}^{L+1} D \sin^2(\theta_i), \quad (\text{F11})$$

where the decomposition

$$\vec{S}_i \cdot \vec{S}_j = \cos(\theta_i) \cos(\theta_j) + \sin(\theta_i) \sin(\theta_j) \cos(\varphi_i - \varphi_j) \quad (\text{F12})$$

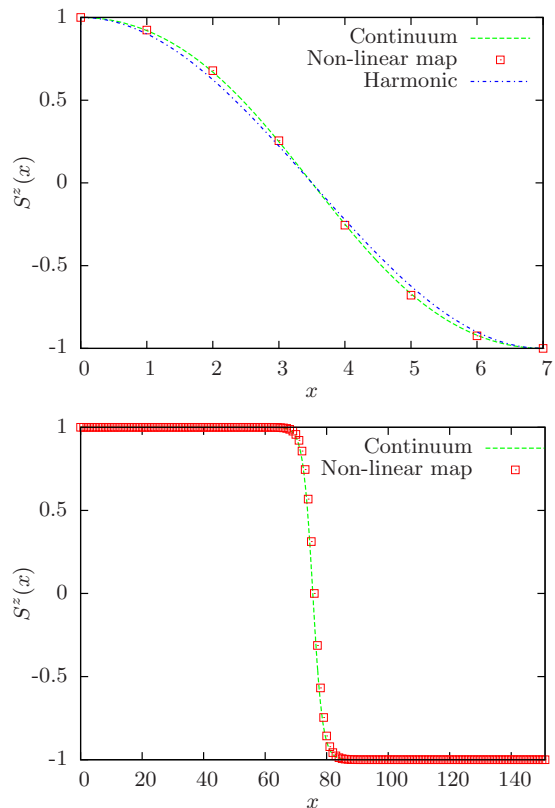


FIG. 8. (Color online) DW spin profile obtained with different methods for $D/J = 0.05$, yielding $c^{-1} \approx 3.16$. Nonlinear map and continuum approach produce the same spin profile for both $L = 6, 150$. For $L = 6$, the harmonic approximation given by Eq. (F10) is also accurate.

with $\varphi_i = \varphi_0$ was used. To find the profile, \mathcal{H}_{dw} has to be minimized:

$$\frac{\partial \mathcal{H}_{\text{dw}}}{\partial \theta_i} = 0, \quad 0 < i < L + 1, \quad (\text{F13})$$

which yields

$$J \sin(\theta_i - \theta_{i-1}) - J \sin(\theta_{i+1} - \theta_i) + D \sin(2\theta_i) = 0. \quad (\text{F14})$$

The nonlinear map is built as follows:

$$\tilde{s}_{i+1} = \sin(\theta_{i+1} - \theta_i), \quad (\text{F15})$$

so that

$$\tilde{s}_{i+1} = \tilde{s}_i + \frac{D}{J} \sin(2\theta_i) \quad (\text{F16})$$

and

$$\theta_{i+1} = \theta_i + \arcsin(\tilde{s}_{i+1}). \quad (\text{F17})$$

The procedure gives a unique solution when the boundary conditions are inserted

$$\theta_0 = 0, \quad \theta_{L+1} = \pi, \quad (\text{F18})$$

and the first step is given, $\theta_L = \pi - \delta$. The difficulty of the procedure is to find the correct value of δ that satisfies the boundary conditions (F18). In practice, one tries different values of δ until $|\theta_0|$ is smaller than a preset threshold.

However, this procedure gets numerically unstable for long chains. For L odd, one can circumvent the problem by letting the iteration start at the DW center, where $\theta_{(L+1)/2} = \pi/2$ and $\theta_{(L+3)/2} = \pi/2 + \delta$. The more dramatic misalignment between two adjacent spins occurs at the DW center. Therefore here the iteration step δ can be defined more easily. After a certain iteration \bar{i} , all the spins characterized by a lattice index $i > \bar{i}$ are assumed to point along the same direction as the boundary, $\theta_i = \pi$. The configuration of spins lying on the other side of the DW is obtained from the condition

$$\theta_{L+1-i} = \pi - \theta_i, \quad (L+1)/2 < i \leq L+1. \quad (\text{F19})$$

Figure 8 shows that the results obtained with the nonlinear map (for a discrete lattice) and those computed with the continuum formalism [by means of Eqs. (F7) and (F8)] are not distinguishable.

APPENDIX G: \mathcal{Z}_{φ_0} CONTRIBUTION AND DW CHIRALITY

With a similar strategy followed to obtain Eq. (58), the contribution due to the bound state with vanishing energy associated with degeneracy in φ_0 can be evaluated. Still from Eq. (26) it follows that $b_{\varphi_0, q_y} = \varphi_0(q_y)\sqrt{2/c}$. The corresponding factor in the partition function is transformed accordingly:

$$\begin{aligned} \mathcal{Z}_{\varphi_0} &= \int \prod_{\varphi_0, q_y} db_{\varphi_0, q_y} e^{-\beta \varepsilon_{\varphi_0, q_y} |b_{\varphi_0, q_y}|^2} \\ &= \left(\frac{2}{c}\right)^{L_y/2} \int \prod_{q_y} d\varphi_0(q_y) \exp\left[-\frac{\beta J}{c} q_y^2 |\varphi_0(q_y)|^2\right]. \end{aligned} \quad (\text{G1})$$

Establishing the same analogy as before, with a ‘‘stiffness’’ $2J/c$, a contribution equivalent to Eq. (58) is recovered, i.e., $\mathcal{Z}_{\varphi_0} = \mathcal{Z}_{x_0}$. On a more physical basis, the Boltzmann weight

in Eq. (G1) may be thought of as arising from an XY-model Hamiltonian,

$$\begin{aligned} \mathcal{H}_{xy} &= -\frac{2J}{c} \sum_y \cos[\varphi_0(y+1) - \varphi_0(y)] + \frac{2J}{c} L_y \\ &= -\frac{2J}{c} \sum_y \vec{e}_{\varphi_0}(y+1) \cdot \vec{e}_{\varphi_0}(y) + \frac{2J}{c} L_y, \end{aligned} \quad (\text{G2})$$

involving the φ_0 angles of different arrays of spins, labeled by y . From the second line, it is clear that \mathcal{H}_{xy} describes the effective coupling between chirality vectors of neighboring chains. Along this line, Eq. (G1) can equivalently be written as

$$\begin{aligned} \mathcal{Z}_{\varphi_0} &\simeq \left(\sqrt{\frac{2}{c}} e^{-2\beta J/c}\right)^{L_y} \int_{-\pi}^{\pi} d\varphi_0(1) \dots d\varphi_0(L_y) \\ &\times \exp\left\{\frac{2\beta J}{c} \sum_y \cos[\varphi_0(y+1) - \varphi_0(y)]\right\}. \end{aligned} \quad (\text{G3})$$

In terms of relative angles $\delta_y = \varphi_0(y+1) - \varphi_0(y)$ (and assuming $L_y + 1 \simeq L_y$), the above expression takes the form

$$\begin{aligned} \mathcal{Z}_{\varphi_0} &= \left(\sqrt{\frac{2}{c}} e^{-2\beta J/c}\right)^{L_y} \left[\int_{-\pi}^{\pi} d\delta_y e^{\frac{2\beta J}{c} \cos(\delta_y)}\right]^{L_y} \\ &= \left[2\pi \sqrt{\frac{2}{c}} e^{-2\beta J/c} \mathcal{I}_0\left(\frac{2\beta J}{c}\right)\right]^{L_y}, \end{aligned} \quad (\text{G4})$$

where $\mathcal{I}_0(\kappa)$ is the modified Bessel function of the first kind. For $2\beta J/c \gg 1$ (low temperatures), it is $\mathcal{I}_0(\kappa) \simeq e^\kappa / \sqrt{2\pi\kappa}$ and $\mathcal{Z}_{\varphi_0} = \mathcal{Z}_{x_0}$ is recovered again. For the sake of simplicity, this equivalence has been assumed in the calculation of the DW free energy for the 2D case [see Eq. (62)].

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