Searching for non-Fermi liquids under holographic light

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By expanding the set of background geometries beyond the commonly studied ones we identify those dual gravity models that may provide holographic descriptions for some prototypical non-Fermi liquid states of strongly correlated condensed matter systems. Specifically, we discuss prospective gravity duals of such iconic examples as the nonrelativistic fermions coupled to gauge fields and Dirac fermions with Coulomb interactions.

I. INTRODUCTION

Quantum theory of strongly correlated fermions has long been in a rather desperate need for nonperturbative techniques, the use of which could allow one to proceed beyond the customary (and often uncontrollable in the regime of interest) approximations when analyzing generic (nonintegrable) systems.

Despite all efforts, though, the overall progress has been quite limited. However, it has been argued that a recent proliferation of the ideas based on the hypothesis of holographic duality[1] may offer a possible way out of the stalemate, thus allowing the field to move towards a systematic classification of various "strange" metallic (compressible) states that are commonly (and often indiscriminately) referred to as non-Fermi liquids (NFLs).

The holographic correspondence postulates a connection between certain (non-Abelian, multi-(N) component, and supersymmetric) field theory models and their gravity duals living in one extra dimension, so for $N \gg 1$ the strong-coupling regime of the former can be mapped onto the weak coupling one of the latter (and vice versa). The support for this general idea is provided by a host of circumstantial evidence gathered from the "bona fide" theories of strings and hot QCD quark-gluon plasmas.

Nevertheless, despite a gradually building confidence in the validity of the original holographic hypothesis, the status of its recently proposed condensed matter applications remains, by and large, unknown.

Pursuing the phenomenological "bottom up" approach, the initial studies produced a number of rather baffling results, which include multiple Fermi surfaces (merging into one critical "Fermi ball," or, rather, a "flat band," in the extreme $N \to \infty$ limit), oscillatory frequency dependence and dispersionless poles of the fermion propagator, and so on.[2] Some of those early findings have already been interpreted as spurious artifacts of taking the limit $N \to \infty$. For one, no oscillatory dependence would arise in a more systematic "top-down" approach[3] [besides, the same effect can also be achieved by including a nonminimal (Pauli) fermion-gauge field coupling[4]].

Putting aside the central question about a general applicability of the original holographic hypothesis to (typically, neither non-Abelian/multicomponent, nor supersymmetric) condensed matter systems, one common limitation of the early studies was that their background metrics would typically be chosen from a handful of well-known solutions to the classical Einstein-Maxwell equations.

Obviously, such an "under the light" search for holographic NFLs lacks any physical input specific to a given strongly correlated system and, therefore, its chances of finding a gravity dual for that particular system appears to be rather hard to assess.

In essence, the early studies have so far found just one type of the NFL behavior, dubbed "semilocal criticality," whereby the fermion propagator features a nonanalytical (and, in general, oscillatory) frequency dependence, but only a nonsingular momentum one,

$$G(\omega,q) = \frac{1}{A(q) + B(q)\omega^{\nu_q}}. \tag{1}$$

Here the function $A(q)$ has simple zeros at each of the (potentially, multiple) "Fermi momenta" q_F^i, whereas $B(q)$ takes finite values at such points (see Ref. 2 for details).

This behavior suggests that, at long distances, the system effectively splits onto spatially uncorrelated "quantum impurities," each of which exhibits a characteristic $d = 0$ quantum-critical scaling. Although it was pointed out that the propagator (1) bears a certain resemblance to that expected in the context of some heavy fermion materials (see, e.g., Ref. 5 and references therein), in the absence of any solid agreement with experiment that would suggest otherwise, this might well turn out to be merely superficial.

In view of such uncertainty, the task of putting the holographic correspondence on a firm ground and ascertaining the status of its predictions more definitively would be best achieved if this technique were applied to those situations where prior insight has already been gained by some other means. To that end, it would be very helpful, first, to find the gravity duals of the already established (or, at least, suspected) NFL states.

In what follows, we demonstrate that accomplishing this task requires one to extend the class of metrics well beyond the commonly studied classical examples.

II. PROSPECTIVE GRAVITY DUALS OF CONDENSED MATTER SYSTEMS

The early applications of the holographic hypothesis yielding Eq. (1) utilized the standard Reissner-Nordstrom (RN) "black brane" solution, which minimizes the Einstein-

Maxwell action

$$S_g = \int \frac{1}{2\kappa^2}\left(\mathcal{R} + \frac{d(d+1)}{L^2}\right) - \frac{1}{2e^2}F_{\mu\nu}^2. \qquad (2)$$

Hereafter, $\int = \int dt\,dz\,d^d\vec{x}\sqrt{-\det g_{\mu\nu}}$ stands for a covariant $d+2$-dimensional volume integral, $\mathcal{R}$ is the scalar curvature, the second term represents a (negative) cosmological constant of the asymptotically anti-de Sitter space AdS_{d+2} of curvature radius L, κ^2 is the Newtonian coupling, and e is the charge of the $U(1)$ gauge field.

The corresponding metric (throughout this paper, the speed of light $c = 1$)

$$ds^2 = -f(z)dt^2 + g(z)dz^2 + h(z)d\vec{x}^2 \qquad (3)$$

has nonzero components $f(z)(z/L)^2 = (L/z)^2/g(z) = 1 - (1+\mu^2)(z/z_h)^{d+1} + \mu^2(z/z_h)^{2d}$, $h(z) = (L/z)^2$. Here μ is the (dimensionless) chemical potential of the fermion matter, which is assumed not to react back on the metric, and z_h is the inverse radius of the horizon determined from $f(z_h) = 0$.

The latter can remain finite even when the Hawking temperature $T = [d+1 - (d-1)\mu^2]/4\pi z_h$ vanishes, thereby giving rise to yet another artefact of the $N \to \infty$ limit, a seemingly nonvanishing entropy $S(T \to 0) \neq 0$. In the near-boundary $z \to 0$ (ultraviolet or UV) regime, Eq. (3) recovers the standard AdS_{d+2} form, $f(z) = g(z) = h(z) = (L/z)^2$.

In fact, the "locally critical" behavior (1) sets in at the extremal $T \to 0$ limit,[2] where the near-horizon geometry approaches $\mathrm{AdS}_2 \times R^d$. Moreover, a similar [albeit more physically sound, entropywise, $S(T) \sim T^{d/\eta}$] behavior was found for a variety of geometries which reduce to the one-parameter "Lifshitz" metric $f(z) = (L/z)^{2\eta}$, $g(z) = h(z) = (L/z)^2$ in the $z \gg 1$ (infrared or IR) regime.[6] It was also shown to result from an approximate (Thomas-Fermi) account of the fermions' back-reaction in the framework of the standard gravity (2), thus leading to the "electron star" scenario.[7]

In fact, the latter metric naturally emerges alongside a whole class of more general solutions in the so-called dilaton gravity whose Lagrangian includes an additional bulk scalar field[8]

$$S_{dg} = \int \frac{1}{2\kappa^2}\left[\mathcal{R} - \frac{(\partial\phi)^2}{2} + U(\phi)\right] - \frac{Z(\phi)}{2e^2}F_{\mu\nu}^2. \qquad (4)$$

In the minimal version, both the dilaton potential $U(\phi) = d(d+1)e^{\delta\phi}/L^2$ and the effective gauge coupling $Z(\phi) = e^{\gamma\phi}$ are given by simple exponential functions with the coefficients δ and γ.

In what follows, we focus on the $T = 0$ case and consider a still broader class of static and spherically symmetric metrics

$$f(z) = (L/z)^{2\eta}, \quad g(z) = (L/z)^{2\alpha}, \quad h(z) = (L/z)^{2\beta}, \qquad (5)$$

whereas at finite T one also has the freedom of altering the additional polynomial "emblackening factor," similar to that in the RN solution (3).

For any $\beta \neq 0$, Eq. (5) can be reduced to a two-parameter family of metrics known as the "hyperscaling violating" backgrounds.[9] The latter are characterized by the dynamical exponent ζ and hyperscaling violation parameter θ,

$$\zeta = \frac{\eta+1-\alpha}{1-\alpha+\beta}, \quad \theta = d\frac{1-\alpha}{1-\alpha+\beta}, \qquad (6)$$

which manifest themselves through the scaling properties of the excitation spectrum in the boundary theory: $\omega \to \lambda^\zeta\omega$ for $q \to \lambda q$ and that of the interval (3), $ds \to \lambda^{\theta/D}ds$.

In Ref. 9, a number of "top-down" string scenarios were presented, by which the "hyperscaling violating" geometries may arise. In that regard, the null energy criteria that must be obeyed by any physically sensible gravity dual with the metric (5) impose the inequalities

$$\beta(\eta - \beta + \alpha - 1) \geqslant 0, \quad (\eta - \beta)(1 - \alpha + \eta + d\beta) \geqslant 0. \qquad (7)$$

As in Ref. 9, the generalized metrics (5) do not cover the $z \to 0$ region and, therefore, require a proper UV completion. Thus, they should be viewed as gravity duals of some effective IR field theories residing at a finite z_0. Correspondingly, all the holographic propagators discussed in the rest of this paper pertain to the renormalized operators from such effective theories rather than those of the "microscopic" boundary ones. The latter can be obtained from the former by virtue of the matching procedure, akin to that of Ref. 2.

III. SEMICLASSICAL PROPAGATORS

In the holographic analyses, the bulk fermions of mass m couple to the metric and gauge fields in the minimal way,

$$S_f = \int \bar{\psi}\gamma_\mu\left(i\partial_\mu + \frac{i}{8}[\gamma_\lambda,\gamma_\nu]\omega_{\lambda\nu}^\mu + eA_\mu - m\right)\psi, \qquad (8)$$

where γ_μ are the γ matrices and $\omega_{\mu\nu}^\lambda$ is the spin connection.[1]

In the absence of explicit analytical solutions for the bulk fermion wave functions in generic gravitational backgrounds, one can still resort to the semiclassical approach. The equation for the Fourier-transformed wave function $\psi(z,\omega,q)$ features an effective single-particle potential[10]

$$V(z) = m^2 + \frac{q^2}{h(z)} + \frac{\omega^2}{f(z)}, \qquad (9)$$

which allows for two zero-energy solutions in the tunneling region $z_0 < z < z_t$,

$$\psi_\pm(z,\omega,q) \sim \frac{1}{V^{1/4}(z)}e^{\pm\int_{z_0}^{z_t}dz\sqrt{g(z)V(z)}}, \qquad (10)$$

where the turning point z_t is defined as $V(z_t) = 0$.

Using Eq. (10) one then finds the effective IR theory's Green function as the reflection coefficient for the wave incident at $z = z_0$[1]

$$G_{IR}(\omega,q) = \left.\frac{\psi_-(z,\omega,q)}{\psi_+(z,\omega,q)}\right|_{z\to z_0} \sim e^{-S(\omega,q)}, \qquad (11)$$

where

$$S(\omega,q) = 2\int_{z_0}^{z_t}dz\sqrt{g(z)V(z)}. \qquad (12)$$

Considering the metric (5) and focusing on the limit of a small fermion mass, one obtains the scaling behavior

$$S(\omega,q) \sim \left(\frac{q^{1-\alpha+\eta}}{\omega^{1+\beta-\alpha}}\right)^{\frac{1}{\eta-\beta}}, \qquad (13)$$

indicative of the underlying quasiparticle dispersion $\omega \sim q^\zeta$ governed by the dynamical exponent (6).

In the complementary limit of a large mass, the semiclassical analysis can be more conveniently employed directly in the real space.[11] In this regime, various quantum-mechanical amplitudes are dominated by the fermion paths that closely follow the classical trajectories (geodesics) obtained from the (imaginary-time) action

$$S(\tau,x) = m \int dz \sqrt{g(z) + f(z)(d\tau/dz)^2 + h(z)(dx/dz)^2}. \tag{14}$$

When evaluated on such a trajectory, Eq. (14) reads

$$S(\tau,x) = m \int_{z_0}^{z_t} dz \sqrt{\frac{g(z)}{r(z)}}, \tag{15}$$

where $r(z) = 1 - \Pi_x^2/h(z_t) - \Pi_\tau^2/f(z_t)$ is a function of the conjugate momenta $\Pi_x = \delta S/\delta(dx/dz)$ and $\Pi_\tau = \delta S/\delta(d\tau/dz)$ given by the integral equations

$$x = \Pi_x \int_{z_0}^{z_t} \frac{dz}{h(z)} \sqrt{\frac{g(z)}{r(z)}}, \quad \tau = \Pi_\tau \int_{z_0}^{z_t} \frac{dz}{f(z)} \sqrt{\frac{g(z)}{r(z)}} \tag{16}$$

and the turning point is obtained by solving the equation $r(z_t) = 0$. The minimal action (15) then controls the fermion propagator, $G(\tau,x) \sim e^{-S(\tau,x)}$.

While an explicit computation of Eq. (15) can be performed only in some special cases, determining simpler, one-parameter, dependencies $S(\tau)$ and $S(x)$ is possible for a broad variety of metrics. Specifically, for the metric (5) one obtains

$$S(x) \sim x^{\frac{1-\alpha}{1-\alpha+\beta}}, \quad S(\tau) \sim \tau^{\frac{1-\alpha}{1-\alpha+\eta}}. \tag{17}$$

Notably, both Eqs. (12) and (15) elucidate the role of the radial variable z as an energy-like renormalization scale parameter. However, a direct correspondence between the two can not be readily established, since the Fourier transformation relating $G(\omega,q)$ and $G(\tau,x)$ requires both functions, including their nonexponential prefactors, to be known across the entire ranges of their arguments. It might be possible, though, to relate their asymptotics by virtue of the saddle-point method, wherever applicable.

IV. FINITE DENSITY FERMIONS WITH SINGULAR INTERACTIONS

One important testing ground for the holographic hypothesis is provided by the theory of finite density fermions coupled to an Abelian gauge field. This problem has long been at the forefront of theoretical research where it was studied with a whole variety of techniques, although the case still remains unclosed. For instance, the recent results of Ref. 12, which revisited the attempts to obtain a self-consistent resummation to all orders in the spirit of the Eliashberg theory,[13] indicate that a naive $1/N$ expansion may not be as reliable as previously thought.

Furthermore, long-ranged and retarded ("singular") interactions that allow for a similar description are often associated with the onset of ground-state instabilities, and the

concomitant NFL behaviors might occur even in those systems whose microscopic Hamiltonians involve only short-ranged couplings.

Such interactions are mediated by gapless bosonic excitations of an emergent order parameter, and in all the diverse reincarnations of the problem, their gauge-like propagator conforms to the general expression

$$D(\omega,q) = \frac{1}{|\omega|/q^\xi + q^\rho}. \tag{18}$$

Important pertinent examples include anomalous electromagnetic skin effect in metals,[14] compressible quantum Hall states with screened repulsive interactions,[15] critical spin fluctuations in itinerant ferromagnets,[16] and density fluctuations in "quantum nematics,"[17] for all of which $\xi = 1$, $\rho = 2$. In contrast, normal skin effects and antiferromagnetic fluctuations would be described by $\xi = 0$, $\rho = 2$, whereas compressible quantum Hall states with the unscreened Coulomb interactions correspond to $\xi = 1$, $\rho = 1$.

The asymptotic IR behavior of the propagator of fermions coupled to a gaugelike bosonic mode can be evaluated by means of the eikonal-type procedure,[18] which reduces the former to the phase factor taken along the classical trajectory

$$G(\tau,x) \sim \left\langle \exp\left[i \int A_\mu(z=z_0)dx_\mu\right] \right\rangle_A = e^{-S}. \tag{19}$$

Here the averaging is performed over a (physical or effective) gauge field A_μ governed by the propagator $\langle A_\mu A_\nu \rangle = D(\omega,q)(\delta_{\mu\nu} - q_\mu q_\nu/q^2)$, thereby resulting in

$$S(\tau,x) = \frac{1}{2} \int \frac{d\omega d^d q}{(2\pi)^{d+1}} D(\omega,q) \frac{1 - \cos(\omega\tau - \mathbf{qx})}{(i\omega - \mathbf{vq})^2}, \tag{20}$$

where $\mathbf{v} \sim q_F \hat{\mathbf{x}}$ is the Fermi velocity in the direction of the vector $\mathbf{x}$.

To estimate the eikonal action (20) for a timelike interval

$$S(\tau) = \frac{1}{2} \int \frac{d\omega d^d q}{(2\pi)^{d+1}} D(\omega,q) \frac{1 - \cos\omega\tau}{(i\omega - \mathbf{vq})^2} \sim \tau^{\frac{\rho+1-d}{\xi+\rho}}, \tag{21}$$

we first perform the momentum and then the frequency integrations, thereby discovering that, at $\tau \to \infty$, the integral is dominated by the frequencies $\omega \sim \tau^{-1}$ and momenta $q \sim \tau^{-1/\xi+\rho}$.

Moreover, the kinematics of fermion scattering is such that, at small scattering angles (which is the regime amenable to the eikonal approximation), one finds $\omega \ll |\mathbf{vq}| \ll |\mathbf{v} \times \mathbf{q}|$ for all $2 - \xi < d < 1 + \rho$. Thus, the scattering momentum appears to be primarily directed along the Fermi surface and perpendicular to the local Fermi velocity.

With that observation in mind, one can also compute the integral (20) for a spacelike interval, thus obtaining the asymptotic behavior

$$S(x) \sim x^{\frac{\rho+1-d}{\xi+d-1}}. \tag{22}$$

It is then easy to see that Eqs. (17), (21), and (22) match, provided that the following relations hold:

$$\begin{aligned} \frac{1-\alpha}{1-\alpha+\beta} &= \frac{\rho+1-d}{\xi+d-1} \\ \frac{1-\alpha}{1-\alpha+\eta} &= \frac{\rho+1-d}{\rho+\xi}. \end{aligned} \tag{23}$$

The above results pertain to the propagators of the effective field theories which, unlike their underlying microscopic counterparts, become essentially universal after having undergone renormalization down to the IR scale z_0. As such, they need to be contrasted with the holographic Green functions computed at a finite z_0, rather than those at the original boundary $z = 0$.

Thus, in order to reproduce the effects of the singular interaction (18) with $\xi = 1$, $\rho = 2$ and $d = 2$ in the holographic setting, one needs to choose the relevant parameters as follows: $\eta = 2\beta = 2(1 - \alpha)$. For comparison, the case of $\xi = 1$, $\rho = 1$ can be covered by choosing $\eta = \beta$, $\alpha = 1$, whereas the case of $\xi = 0$, $\rho = 2$ requires $\eta = 1 - \alpha$, $\beta = 0$, which values would be unattainable within the class of hyperscaling violating metrics.

Notably, all of the above metrics comply with the criteria (7) for the existence of a consistent gravity dual, the second one being satisfied as a strict equality, $\eta = 1 - \alpha + \beta$ or $\zeta = 1 + \theta/d = [2(1 - \alpha) + \beta]/(1 - \alpha + \beta)$. Also, the corresponding values of the dynamical exponent ($\zeta = 3/2$, 1, and 2, respectively) agree with those inferred from contrasting the quasiparticle dispersion **vq** against the fermion self-energy[12–17]

$$\Sigma(\omega) = \int \frac{d\epsilon \, d^d q}{(2\pi)^{d+1}} \frac{D(\omega, q)}{i\omega + i\epsilon - \mathbf{vq}} \sim \omega^{\frac{d-1+\xi}{\xi + \rho}}, \quad (24)$$

which in comparison yields $\zeta = (\xi + \rho)/(d - 1 + \xi)$.

Obviously, the matching conditions (23) are only the necessary ones, so they may not always guarantee that the entire two-parameter functional dependence of the action $S(\tau, x)$ would be reproduced with this choice of parameters. Our discussion should then be viewed as merely suggestive of a possible holographic correspondence between the aforementioned theories, and, in order to further strengthen the case, more observables would need to be matched.

However, if the conditions (23) are not met, then no viable gravity duals of the aforementioned physically revelant NFL systems may be found among the generalized family of metrics (5). By this argument, one concludes that none of these such systems can be naturally mated with the classical AdS-RN metric considered in Ref. 2.

V. COULOMB INTERACTING DIRAC FERMIONS

The recent upsurge of interest in graphene and topological insulators—as well the earlier advent of 1D Coulomb metals (e.g., carbon nanotutes), gapless 2D high-T_c superconductors, and quasiparticle properties of 3D superfluid He^3—brought out the problem of (pseudo-)relativistic Dirac fermions with isolated Fermi points and potentially long-ranged (due to a lack of screening), albeit nearly instantaneous, interactions.

In the 1D case, the asymptotic dual geometry is AdS$_3$ and the corresponding (conformally invariant) boundary theory is that of chiral 1D fermions. Its proper UV completion is naturally achieved with the use of Eq. (3) for $\mu = 0$, known as the "nonrotating BTZ black hole,"[19] and the (exact) finite-temperature chiral fermion propagators read

$$G_\pm(\tau, x) = \left(\frac{\pi T}{\sinh \pi x_+ T} \right)^{2\Delta_+} \left[\frac{\pi T}{\sinh(x_- T)} \right]^{2\Delta_-}, \quad (25)$$

where $x_\pm = x \pm \tau$. In the $T \to 0$ limit Eq. (26) amounts to $G_\pm(\tau, x) = 1/x_+^{2\Delta_+} x_-^{2\Delta_-}$.

According to the holographic principle,[1] the boundary theory then must be strongly coupled, as manifested by the UV (left/right) fermion dimensions, $\Delta_\pm = mL/2 + 1/2 \pm 1/4$, which are necessarily large, $\Delta_+ + \Delta_- > 1$.[19] Notably, such dimensions cannot be obtained from any 1D theory with short-ranged repulsive couplings where the corresponding Luttinger parameter would be limited to the interval $1/2 \leqslant K \leqslant 1$, thereby resulting in $1/2 \leqslant \Delta_+ + \Delta_- = \frac{1}{4}(K + 1/K) \leqslant 5/8$.

In fact, the still-lower K values, $0 < K < 1/2$, can be attained only in the presence of long-ranged interactions, such as Coulomb, which endows the Luttinger parameter and the fermion dispersion with a slow momentum dependence,

$$K(q) = \frac{1}{\sqrt{1 + \sigma |\ln q|}}, \quad \epsilon_q = q\sqrt{1 + \sigma |\ln q|}, \quad (26)$$

where $\sigma = 2e^2/\pi$.

In the general case of the d-dimensional Dirac fermions with the 3D Coulomb interaction, $U_q \sim q^{1-d}$, the counterpart of Eq. (20) reads (after the frequency integration)

$$S(\tau, x) = \int \frac{d^d q \, U_q}{(2\pi)^d \epsilon_q} [1 - \cos(\epsilon_q \tau - qx)]. \quad (27)$$

In the 1D case, one then obtains the leading behavior,

$$S(\tau, x) \sim \sigma^{1/2} \ln^{3/2} |x - \tau|, \quad (28)$$

which gives rise to a faster-than-algebraic decay of the propagator $G(\tau, x) \sim e^{-S}$, thus implying that the system undergoes the 1D analog of the Mott transition.[20]

Making use of Eq. (15), one observes that capturing the behavior (28) in the holographic framework would require a logarithmic deformation of the asymptotic AdS$_3$ geometry

$$g(z) = (L/z)^2 \ln z/z_0, \quad f(z) = h(z) = (L/z)^2. \quad (29)$$

Although this ansatz may not be unique, it shows that a prospective gravity dual of the Coulomb-interacting Dirac fermions is likely to lie outside the family of the AdS-RN metrics utilized in Ref. 2. On the other hand, there exist solutions showing logarithmic behavior at intermediate values of z for some generalized dilaton potentials, as in Eq. (4), tuned to their degeneracy points.[8] Thus, one would have to tap into those resources if viable candidates to the role of the gravity dual of the 1D Coulomb metal were to be found.

VI. SUMMARY

In conclusion, we demonstrate that in order to ascertain the status of the holographic approach in the context of its applications to the hypothesized NFLs, one needs to venture out of the comfort zone of the customary gravitational backgrounds in search of new (self-consistent) solutions to the coupled equations for the metric, dilaton, gauge, and matter fields along the lines of Refs. 6–9.

To that end, we discuss the specific examples of non-relativistic fermions coupled to gaugelike fields and Dirac fermions with the Coulomb interactions, in both cases finding the prospective metrics to possibly belong to the solutions (alongside the hyperscaling-violating ones) of some generalized dilaton gravity models.

Based on these observations, we suggest that, within such a broad class of metrics, one would have better chances of "reverse engineering" the gravity duals of the already documented NFLs, thus putting the entire holographic machinery up to a decisive test. Then, after having affirmed that this approach might work in the already known cases, one can continue expanding the list of novel NFL states with greater confidence.

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