Kramers-Kronig constrained modeling of soft x-ray reflectivity spectra: Obtaining depth resolution of electronic and chemical structure

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Resonant x-ray scattering is a powerful technique for the study of electronic structure at the nanoscale. In common practice, the optical properties of the constituent components of a material must be known prior to modeling of the scattered intensity. We present a means of refining electronic structure, in the form of optical properties, simultaneously with physical structure, in a Kramers-Kronig consistent manner. This approach constitutes a sensitive and powerful extension of resonant x-ray scattering to materials where the optical properties are not sufficiently well known. The application of this approach to specular reflectivity from a single crystal of SrTiO₃ is presented as an example case, wherein we find evidence for both a nonresonant surface contaminant layer and a modified SrTiO₃ surface region. Extrapolating from this study we comment on the potential utility of this approach to resonant scattering studies in general.

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I. INTRODUCTION

Resonant x-ray scattering extends anomalous scattering¹ across atomic core levels to combine the element and state specific electronic structure available from absorption measurements with information on spatial correlations available from angle-resolved scattering. Such an approach can elucidate aspects of the spatial distribution of electronic or chemical structure not available away from resonance.^{2–7} It constitutes a growing application of tunable synchrotron radiation and promises access to ultrafast electron dynamics at short length scales in x-ray free-electron laser sources. Strong, sharp absorption features at the soft x-ray core levels make resonant x-ray scattering an attractive tool to study important aspects in magnetic, correlated electron, and even polymeric systems with nanometer-scale resolution and greater penetrating power than powerful electron spectroscopies.

In the near-visible spectral region several methods are widely used to obtain absolute optical constants, including their spectral and depth dependence, from reflectance data.^{8–10} Such methods include reflectometry, ellipsometry, and polarimetry, and generally aim to determine the real and imaginary parts of a complex index of refraction through measurements of reflected intensity, phase change, and/or polarization. At discrete photon energies, modeling angleresolved data is common and may or may not yield results consistent with the Kramers-Kronig (KK) dispersion relation. Full spectroscopic determinations require energy-resolved measurements and their analysis in a manner ensuring KK consistency. The same approaches and challenges apply in the soft x-ray range, although here it is much easier to measure scattered intensity than polarization or phase.

Resonant x-ray scattering intensities are generally described as the modulus squared of a complex optical contrast involving one or more complex refractive indices, $n = 1 - \delta - i\beta$. Applications often require accurate resonant x-ray absorption spectra (commonly referred to as XAS, NEXAFS, or XANES) to interpret results. Such spectra are obtained either from measurement or some form of model calculation, and are often adequate for the problem at hand.^{11,12} However, as resonant x-ray scattering applications become more refined, cases are encountered in which the spatial distribution of electronic structure itself is to be determined, and in such cases it is not obvious that measured or model optical spectra are sufficient. Rather, obtaining the spatial distribution of resonant optical properties directly by simultaneously modeling spectroscopy and structure in fitting experimental data would provide a useful alternative. Fixed-energy determinations of optical constants are common in the x-ray region [generally through fitting of angular resolved reflectivity, R(Q)].^{13–21} While this approach has been extended to fitting fixed energy reflectivity data at a number of energies spanning the atomic core level as a means of spectroscopy,²² full spectroscopic analysis is not generally applied to reflectivity data.

Here we present a flexible approach to model spectra and structure on equal footing in fitting reflected intensities, using a variational algorithm²³ to efficiently ensure KK consistency between δ and β as they refine along with structural parameters to yield a self-consistent model of the spatial distribution of electronic structure. This variational approach is designed to model spectra of arbitrary shape to provide accurate absolute values for optical spectra within the assumptions of the structural model, which may also be refined in the modeling process. This is an extension of an earlier approach to reflectivity modeling that used a small number of Lorentzian peaks to describe the absorption spectra.²⁴ The current approach is similar, but allows for much greater flexibility to fit the fine details of a measured reflectivity spectrum.

To illustrate this approach we apply it to determine the resonant optical properties across the Ti $L_{2,3}$ and O K edges in a single crystal of SrTiO₃ (001) (STO). This material was chosen because STO remains technologically important and because its near-surface region is known to exhibit subtle and complex behavior which remains a topic of active study.^{24–27} We find evidence that two surface layers, a contamination layer atop an STO-like layer departing from bulk properties, are needed to explain the specific set of experimental data that we model. Below, the variational approach and fitting strategy are

described, followed by its systematic application to reflected intensity spectra measured at several fixed incidence angles. Comments on its broader potential utility are given in the conclusion.

II. REFLECTIVITY MODELING

Our goal is to measure specular reflectivity as a function of x-ray energy across the resonant absorption edges of a material and then fit this with a model of both physical and electronic structure. The electronic structure is, for our purposes, that projected by resonant dipole transitions onto the optical properties as represented by the complex refractive index, $n = 1 - \delta - i\beta$. The real and imaginary parts of the index of refraction are directly related to the complex atomic scattering factors, $f(h\nu) = f_0 - f'(h\nu) - if''(h\nu)$, through

$$n(h\nu) = 1 - \delta(h\nu) - i\beta(h\nu) = 1 - \frac{r_e c^2 N_A}{2\pi\nu^2} \sum_i \frac{Z_i}{A_i} \rho_i f_i(h\nu),$$
(1)

where r_e is the classical electron radius, ρ_i is the density due to atom *i* such that $\sum \rho_i = \rho$, N_A is Avogadro's number, Z_i and A_i are the atomic number and weight of atom *i*, respectively, and $f_0 = Z - (\frac{Z}{82.5})^{2.37}$. We wish to determine the values of $f_i(hv)$ which give the best agreement to the measured reflectivity. The KK relation allows us to determine the real part of the atomic scattering factor from the imaginary part, or *vice versa*. We work with the imaginary part of the atomic scattering factor, as its features are generally well localized and necessarily positive, as opposed to the dispersive nature of the real part, which may also change sign near resonance. The relevant KK relation is then given by

$$f'(h\nu) = \frac{2}{\pi} \oint_0^\infty \frac{h\nu'}{(h\nu)^2 - (h\nu')^2} f''(h\nu') \, dh\nu'.$$
(2)

It is preferable to work with atomic scattering factors instead of the real and imaginary components of the index of refraction for a number of reasons. The simplest of these is that the KK transform is more easily written for atomic scattering factors. The appropriate KK transform relating δ and β involves the number density of each atom to properly determine the constant offset in δ . By working with atomic scattering factors, the KK transform is independent of considerations of density and composition, which are included only at the higher level where (1) is used to calculate the index of refraction, allowing those parameters to be more easily refined. By considering atomic scattering factors we may consider each species' contribution to the total index of refraction independently. This simplifies the assumption that the refined spectra should agree with the tabulated values away from resonance and leads to a natural means to deal with compound edges.

Having measured the reflectivity at a discrete set of energies $\{hv_j\}$, it is natural to then determine f'' at these same energies. We use the KK relation to obtain f' at these discrete energies to calculate the reflectivity for comparison with measurement. Unfortunately, the formula for specular reflectivity cannot, in general, be inverted to directly solve for the atomic scattering factors, so we must resort to iterative methods to solve for the

values of f''. Thus repeated KK transformations are required. This can be efficiently achieved using KK consistent functions to represent f'' and f', such as symmetrized Lorentzians or Gaussians for absorption lines²⁸ as we recently demonstrated for resonant reflectivity.²⁴ However, arbitrary f'' shapes are not generally compatible with the extended tails of Lorentzian absorption lines. Here we implement a much more flexible and general approach to avoid repeated KK evaluations, namely the variational approach of Kuzmenko.²³

A. The variational approach

In the variational approach, we write f'' as a sum of basis functions

$$f''(h\nu) = \sum_{j} C_{j} f_{j}^{\text{basis}}(h\nu).$$
(3)

In this way, we can uniquely define an arbitrary spectral shape through the set of coefficients C_j and the appropriate basis function. We assume the shape of f'' to be smooth within the resolution of our measurement. To ensure that our refined spectra have this feature, we select a basis function that includes point to point correlations. Namely, we choose a triangle function that extends to the next nearest neighbor, see Fig. 1,

$$f_{j}^{\text{basis}}(h\nu) = \Delta_{j}(h\nu) = \begin{cases} \frac{(h\nu - h\nu_{j-2})}{(h\nu_{j} - h\nu_{j-2})} & h\nu_{j-2} < h\nu < h\nu_{j}, \\ \frac{(h\nu_{j+2} - h\nu)}{(h\nu_{j+2} - h\nu_{j})} & h\nu_{j} < h\nu < h\nu_{j+2}, \\ 0 & \text{otherwise.} \end{cases}$$
(4)

Thus each point will depend on its nearest neighbors,

$$f''(h\nu_j) = C_{j-1}\Delta_{j-1}(h\nu_j) + C_j\Delta_j(h\nu_j) + C_{j+1}\Delta_{j+1}(h\nu_j),$$
(5)

so that there is an inherent point to point correlation in any result, ensuring local smoothness. While we have chosen this particular basis function for its convenience, there are many others that would also be suitable, as detailed more thoroughly in the work by Kuzmenko.²³ However, this ensemble of triangle functions allows for a locally smooth, continuous



FIG. 1. (Color online) A Gaussian function (dashed line), sampled at discreet points (circles), and approximated by a sum of basis triangle functions (gray), leading to a linear approximation to the original function (blue). The KK transform of a single triangle function (green) as well as the sum of all such functions representing the KK transform of the full Gaussian function (red) are also shown.

result from a large number of nearly independent parameters, avoiding the issue of overlap that would arise from any function with extended tails, such as Gaussian or Lorentzian functions.

The primary benefit to expressing f'' in this manner is that we can write an analytical expression for the KK transform of our basis function,

$$KK[\Delta_{j}(h\nu)] = \frac{1}{\pi} \left[\frac{g(h\nu, h\nu_{j-2})}{h\nu_{j} - h\nu_{j-2}} - \frac{(h\nu_{j+2} - h\nu_{j-2})g(h\nu, h\nu_{j})}{(h\nu_{j} - h\nu_{j-2})(h\nu_{j+2} - h\nu_{j})} + \frac{g(h\nu, h\nu_{j+2})}{h\nu_{j+2} - h\nu_{j}} \right]$$
(6)

with

$$g(x,y) \equiv (x+y)\ln|x+y| + (x-y)\ln|x-y|, \quad (7)$$

which allows us to write f' as the linear combination

$$f'(h\nu) = \sum_{j} C_{j} K K\{\Delta_{j}(h\nu)\}.$$
(8)

Evidently, as both $f'(h\nu)$ and $f''(h\nu)$ depend on the same set of coefficients C_j , changing a single point in $f''(h\nu)$ does not require the full KK transformation to update $f'(h\nu)$; instead a single term of the above sum is simply adjusted to update the previous spectrum and maintain KK consistency. This variational approach greatly reduces the computational burden while maximizing the flexibility to represent arbitrary spectral shapes.

B. Maximum entropy

One drawback to the variational approach to modeling reflectivity data is the large number of free parameters needed to define a spectral and physical model. Parameterizing the absorption spectra at each measured x-ray energy implies a roughly equal number of refinable parameters as data points, making the problem underconstrained. As a result, we cannot be certain to refine towards a stable solution representing the global minimum, nor that any such solution will be unique. We find a solution to this problem in the field of information theory, namely the principle of maximum entropy.²⁹ We define an entropy as

$$S = \sum_{j} C_j - m_j - C_j \ln \frac{C_j}{m_j},\tag{9}$$

where m_j are the default values for our parameters and represent the solution we would prefer in the absence of any data, which for practical purposes is our set of starting values. We consider the entropy as a constraint in our fit to the data which helps to stabilize the least-squares fitting for a free-form solution. Instead of considering just the set of parameters which give the best fit to the measured data, we will search for the set of parameters that give sufficient agreement with the data and simultaneously maximize the entropy. This can be accomplished through the use of a Lagrange multiplier, λ , in maximizing

$$\lambda S - \frac{\chi^2}{2},\tag{10}$$

where χ^2 represents the least-squares metric, or other appropriate figure of merit.

The maximum entropy solution is unique, and will deviate as little as necessary from the starting model while still fitting the data. While there are many ways of solving the maximum entropy equation (10),^{30–32} they will all, in general, maintain positivity for the refined parameters. As all of the parameters should be positive (representing things such as layer thickness, roughness, density, composition, or absorption, all of which may be refined³³), this is an advantageous feature which helps to ensure a physically meaningful result.

III. THE SrTiO₃ (001) SURFACE: AN EXAMPLE CASE

As an example of the method, we consider the Ti *L* edge and O *K* edge of a SrTiO₃ (100) single crystal measured using specular reflectivity data, R(hv), at several discrete incidence angles, θ . SrTiO₃ (STO) is a common substrate for thin film growth and an important dielectric material in its own right. STO represents an interesting test case which has been well studied and is suitably complex so as not to be trivial. In fact, the large number of studies on the STO surface are not all in agreement, possibly due to the dynamic nature of the surface. While this makes comparison of our results to those in the literature difficult, it also presents an opportunity for a unique look at this complex system.

The data presented here were measured from a commercially obtained (100) oriented substrate roughly 5 mm square. The sample had been exposed to x-rays under vacuum during previous measurements and stored under atmosphere for extended periods between measurement periods. As the surface state of STO has been observed to change,²⁴ the complex history of this sample makes its exact surface state uncertain at the time of data collection. Data were collected at beamline 6.3.2 of the Advanced Light Source at Lawrence Berkeley National Laboratory at incidence angles of 14°, 15°, 16°, 17°, and 30° in a single session and are considered internally self-consistent. Care was taken to correctly normalize all reflectivity spectra to the incident beam spectrum in order to obtain the best values for the absolute reflectivity. Refinements were done using our reflectivity analysis program written in PYTHON. The program uses the variational approach and maximum entropy to refine f'' values of resonant atoms and physical parameters of the model for the sample. Reflectivities are calculated using the Parratt formalism³⁴ with the inclusion of a Névot-Croce³⁵ roughness correction for each interface. Refinement details for each model considered can be found in Table I.

To start, we consider the simplest possible model, that of a semi-infinite slab of stoichiometric STO (model 1). In order to simplify and speed up the refinement procedure, we begin by considering only the data measured over the Ti edge at an incidence angle of $\theta = 14^{\circ}$. The refinement starts from a four-Lorentzian model spectrum for the Ti edge, similar to that used by Valvidares *et al.*²⁴ The results are shown in Fig. 2. The fit to the measured reflectivity is essentially perfect, indicating that the refinement has proceeded as intended. However, even a cursory inspection of the resulting f'' spectrum shows serious deficiencies. The dotted line is the tabulated data³⁶

TABLE I.	Refinement details for each of the models considered	 Those parameter 	values marked	with an asterisk (*	*) were held fixed	during
refinement.						

	Model 1	Model 2	Model 3a	Model 3b
Surface Layer Thickness (Å)		4.93	9.54	9.62
Surface Layer Roughness (Å)		3.30	3.93	3.48
Surface Layer Density (g/cm^3)		1.674	1.057	1.073
Surface Layer Composition		C^*	C^*	C^*
STO Surface Layer Thickness (Å)			21.65	21.24
STO Surface Layer Roughness (Å)			2.73	2.42
STO Surface Layer Density (g/cm ³)			5.13*	5.13*
STO Surface Layer Composition			Sr _{1.10} Ti _{0.88} O _{2.86}	Sr _{1.09} Ti _{0.88} O _{2.85}
STO Bulk Layer Thickness (Å)	∞^*	∞^*	∞^*	∞^*
STO Bulk Layer Roughness (Å)	10.60	5.42	9.35	6.75
STO Bulk Layer Density (g/cm ³)	5.13*	5.13*	5.13*	5.13*
STO Bulk Layer Composition	SrTiO ₃ *	SrTiO ₃ *	SrTiO ₃ *	SrTiO ₃ *
Figure of Merit $\left(\sum_{Y_{obs}} \left[\frac{Y_{obs} - Y_{calc}}{Y_{obs}}\right]^2\right)$	0.0368	3.258	5.344	3.518
Number of Parameters	221	302	307	607
Number of Data Points	224	1356	1695	1695

showing the continuum step, and the solid line is the result of the refinement. A physically meaningful absorption spectrum should approach the tabulated values away from resonance, and have a sharp onset of resonant features near the continuum step in the tabulated data. The refined spectrum has two obvious problems: It rises significantly at energies well below the onset of resonance, and it dips slightly below the pre-edge



FIG. 2. (Color online) Results of refinement of model 1. Reflectivity (a) across the Ti $L_{2,3}$ edge measured at $\theta = 14^{\circ}$ (black circles) and calculated from the refined model (red line). The resulting spectrum for the atomic scattering factor f is shown in (b), along with the step function from the tabulated data (dotted line) for that edge for comparison. Only the data shown were considered for the refinement of this model.

values at energies above resonance (\sim 468 eV). Both of these features are unphysical, indicating that this structural model is incorrect in some way.

As a modification to the naïve model 1, we include the possibility of a nonresonant surface layer, which we assume to be composed of carbonaceous material (model 2). This is the simplest modification of model 1 and represents the likely possibility that the sample surface is not perfectly clean. We fit model 2 to a larger subset of the available data, that at $\theta = 14^{\circ}, 15^{\circ}, 16^{\circ}$, and 17° over both the Ti and O edges. Note that this approach deals naturally with compound edges where absorption features of different atoms may overlap. We refine the thickness, density, and surface roughness of the nonresonant surface layer, the interfacial roughness, and the f'' spectra for Ti and O in the STO layer. We limit the refinable range for the absorption spectra to be between 440 eV and 500 eV for Ti and between 525 eV and 600 eV for O. The starting point for this model will be the result of the Lorentzian modeling for the Ti spectra and the absorption measured from appropriately normalized total electron yield for the O edge measured concomitantly with the $\theta = 15^{\circ}$ reflectivity with a 5 Å thick carbon top layer and all roughnesses being equal to 2 Å; these parameters are then refined. The results of model 2 are shown in Fig. 3.

Model 2 represents a significant improvement from the result of model 1. The fit to all of the refined data is good, though not perfect, but more importantly, the resulting f'' spectrum does not show the same severe unphysical behavior as in model 1. Examination of the physical structure shows plausible parameters as well (Table I). However, closer inspection of the reflectivity fits, especially at the Ti edge, shows a systematic variation as we move up in angle such that the model progresses towards an underestimation of the data on resonance, and an overestimation away from resonance. This systematic trend is extremely pronounced at $\theta = 30^\circ$, indicating again that structural model 2 is oversimplified. Understanding the nature of Keissig (thickness) fringes whereby interference minima move from high Q toward low Q as layer thickness increases, the strong θ dependence of fit quality suggests the



FIG. 3. (Color online) Results of refinement of model 2. Reflectivity (a) across the Ti $L_{2,3}$ and O K edges measured at $\theta = 14^{\circ}, 15^{\circ}, 16^{\circ}, 17^{\circ}$, and 30° (black circles) and calculated from the refined model (red line) is shown top to bottom. The Ti $L_{2,3}$ edge is shown in finer detail in (b), showing the systematic trend towards underfitting of the data at higher angle. Note that the reflectivity at $\theta = 30^{\circ}$ is shown here, but was not considered in the refinement of this model. The resulting spectra for the atomic scattering factors *f* are shown in (c) for Ti and (d) for O, along with the step functions from the tabulated data for that edge for comparison (dotted lines).

presence of a thicker layer in addition to the thin carbonaceous layer.

Clearly, the inclusion of a nonresonant surface layer is important, but not sufficient to explain our data. It is also clear that we benefit from the inclusion of as much data as possible in our model refinements, and so will refine our models against all available data henceforth. As a further modification of our STO model, we refer to the literature on this material, which suggests that there may exist a distinct surface region within the STO itself.^{24–27} This surface region may have a composition distinct from the bulk STO and/or may exhibit different spectral features for the resonant atoms. This suggests two possible models, one with distinct composition between the two STO regions, but common spectra for the resonant atoms (model 3a), and one with both distinct composition and spectra (model 3b).³⁷ We consider the simpler of these, model 3a, first.

The thickness of the resonant STO surface region is initially assumed to be 20 Å, in accord with recent results from Valvidares *et al.*²⁴ We make the starting point of the refinement of model 3a the endpoint of model 2 by giving the resonant surface region the same composition as that of the bulk, and setting the spectra for both to the refined result of model 2. In the interest of charge balance, we constrain the composition of the surface region in our model to be a mixture of SrO and TiO₂. This constraint allows for either an enrichment or depletion of Sr or Ti while adjusting the amount of O accordingly. For stoichiometric STO, such as we assume for the bulk, the ratio of SrO to TiO₂ will be 1:1. This is a much simpler calculation than for model 3b while still allowing for a distinct STO-like surface region to exist, the results of which can be seen in Fig. 4. The resonant surface layer refines to a thickness of 21.65 Å, in good accord with our starting value.

Model 3a is a significant improvement upon model 2. The reflectivity spectra are all well fitted, and the resulting spectra for the Ti and O edges show no unphysical behavior. There is still a small systematic discrepancy between measured and calculated reflectivity near the Ti edge, and some misfitting of the O edge in the $\theta = 30^{\circ}$ data, although that may be due to higher levels of noise in the data at that angle. All in all, this is a good model, and might well be considered sufficient for this set of data. Evidently there is a surface region of the STO which shows a Ti depletion and/or a corresponding enrichment of Sr which extends roughly 21 Å into the material. Such a depletion in Ti atoms need not affect the Ti $L_{2,3}$ resonant structure if the remaining atoms are still at the center of undistorted O octahedra, as they are in bulk STO. However, such a chemical change, extending over these length scales, has strong implications for understanding the properties of the STO near-surface region.

In the spirit of completeness, and to explore the limits of this technique, we consider model 3b, with distinct Ti and O spectra for the surface and bulk regions of the STO. The starting point for this refinement will be the result of the refinement for model 3a, and we let the surface and bulk STO spectra refine away from this point. The results of the refinement for model 3b are shown in Fig. 5. The thicknesses, roughnesses, and composition of the surface layers has changed only slightly



FIG. 4. (Color online) Results of refinement of model 3a. Reflectivity (a) across the Ti $L_{2,3}$ and O K edges measured at $\theta = 14^{\circ}, 15^{\circ}, 16^{\circ}, 17^{\circ}$, and 30° (black circles) and calculated from the refined model (red line) is shown top to bottom. The Ti $L_{2,3}$ edge is shown in finer detail in (b). All data shown were considered in the refinement of this model. The resulting spectra for the atomic scattering factors f are shown in (c) for Ti and (d) for O, along with the step functions from the tabulated data for that edge for comparison (dotted lines).

from those obtained for model 3a (Table I). The nonresonant surface layers in both models 3a and 3b appear to be nearly 10 Å thick, and only 1 g/cm³ in density. It should be noted that these two parameters are highly correlated, such that their product (giving the total amount of material present) constitutes the more meaningful parameter. This is simply because the effect of such a thin nonresonant surface layer is primarily an absorptive effect with very little contribution to the interference between the reflectivity from the other layers.

The improvement beyond model 3a is subtle, and best seen in the decrease in the figure of merit by roughly 35%. Given that the number of parameters being fitted has nearly doubled, it is unclear just how significant this improvement is. It is also clear that the differences in spectra between the resonant surface region and the bulk STO are subtle at best. Presumably questions as to the significance of these differences could be addressed with a more complete data set, including $R(h\nu)$ at a larger set of incidence angles as well as R(Q) at discrete hv; for the data considered here they likely fall within the range of uncertainty. While this method is capable of refining multiple spectra, even for the same element, it seems apparent that the distinct surface region in STO exhibits little, if any, spectral variation from the bulk, and is instead due primarily to compositional changes. As such it seems unnecessary to consider models of any greater complexity.

IV. DISCUSSION AND CONCLUSIONS

Variational fitting of reflectivity spectra is a powerful approach for extracting resonant optical constants for a material. We have taken as an example the case of a bare STO substrate. We find evidence for both a resonant STO-like surface region as well as a nonresonant contamination layer above that. The resonant surface region exhibits only subtle spectroscopic variation from the bulk, but does have a distinct composition with an enhanced SrO:TiO₂ ratio which provides optical contrast. This result is not in complete agreement with a previous study on the same material using the less sensitive Lorentzian line shape approach²⁴ which did find evidence for a resonant surface region of similar thickness, but did not consider the possibility of a nonresonant contamination layer. This may be due to the complex and dynamic nature of the STO surface, such that the samples measured may not exhibit the same surface behavior. It is also possible that the nature of the STO-like surface region appears differently when the model is constrained by a Lorentzian line shape in the absence of a nonresonant surface layer, as in that earlier work. Such constraints allow for a reasonable overall fit, but give poor agreement in the fine details of the measured spectra. The variational algorithm models this fine detail, which may possess much of the critical information about a sample.

This example illustrates some of the more important considerations in modeling of resonant reflectivity data. While an initial starting spectra is necessary, no specific assumptions about the details of the resonant spectra at a given edge need to be made for the method to converge. An optical spectrum will generally result, even if the structural model is inadequate, and in such cases the result must be considered as an *effective* optical spectrum within the context of a given model, such as for model 1 above.



FIG. 5. (Color online) Results of refinement of model 3b. Reflectivity (a) across the Ti $L_{2,3}$ and O K edges measured at $\theta = 14^{\circ}, 15^{\circ}, 16^{\circ}, 17^{\circ}$, and 30° (black circles) and calculated from the refined model (red line) is shown top to bottom. The Ti $L_{2,3}$ edge is shown in finer detail in (b). All data shown were considered in the refinement of this model. The resulting spectra for the atomic scattering factors f are shown in (c) for Ti and (d) for O, along with the step functions from the tabulated data for that edge for comparison (dotted lines).

It is thus essential to have criteria available to judge whether a given refined spectral-structural model is acceptable. Such criteria should include physical constraints on the resultant absorption spectra, which should be positive with proper limiting behavior in the pre- and post-edge regions. The uniqueness and accuracy of the final spectra ultimately rely on the self-consistency between the structural and spectral aspects of a model. Spectral self-consistency derives from explicit inclusion of KK consistency, and structural self-consistency from modeling spectral behavior at many angles, as done here, or through the explicit inclusion of angle-dependent scans at one or more energies. Provided that self-consistency is reached, the optical spectra can be considered as absolute and free from measurement artifacts that plague many soft x-ray spectroscopies.

In addition to self-consistency, results from modeling of reflectivity data should be consistent with other forms of measurement. For a suitably conductive sample, it is generally a simple matter to measure absorption spectra by total electron yield (TEY) concomitantly with reflectivity. While TEY measurements may be subject to saturation effects, which tend to reduce the observed intensity of strong absorption features as seen in Fig. 6, the refined spectra should have the same general features as the TEY spectra. Similar issues arise from self-absorption in fluorescence yield measurements (FY). Both TEY and FY measurements of absorption spectra suffer from ambiguities in normalization, especially the choice of energy at which to match the post-edge intensity. Such ambiguities are exacerbated in the case of overlapping edges, which are dealt with more naturally in the approach considered here as evidenced by the weakly overlapping Ti and O spectra. Scattered intensities are generally easy to measure so, by modeling scattering data, we obtain absorption spectra which may be considered exact to within the limitations of the model being considered. Such results should thus serve to provide more stringent constraints on theoretical calculations of spectra. While our method does not build in any form of consistency with, e.g., crystal field multiplet calculations of absorption spectra, results should be in agreement with such calculations, as are ours for Ti.³⁸



FIG. 6. (Color online) Comparison of the Ti $L_{2,3}$ f'' spectra obtained from the refinement of model 3a (blue) and those obtained from a total electron yield (TEY) measurement of the same sample measured at $\theta = 15^{\circ}$ (red). The tabulated values for the step function are shown for comparison (dotted line). Note the systematic depression of the peaks in the TEY measurement, a common result of saturation effects.

The approach presented here may be readily extended to any form of resonant scattering, not just specular reflectivity, provided a suitable expression for the scattering intensity is available for refinement. Reflectivity from stratified systems represents an obvious and broad class of problems amenable to this analysis, as demonstrated here. This approach of fitting reflectivity data, or indeed any resonant scattering data, solves a number of outstanding problems in spectroscopy. Spatial inhomogeneities are commonly averaged in soft x-ray spectroscopy, whereas this approach uses them to determine the spatial distribution of their spectral composition as a matter of course. Most importantly, this approach ensures a self-consistent description of spectroscopic and structural aspects of the sample.

The computational burden required to converge to acceptable tolerances is determined by the complexity of the structural model, both in the number of layers and the number of parameters, and the amount of data being fitted. The method is designed to be very efficient in this respect and can be scaled up to problems of arbitrary size and complexity. To this end, advances in computing power and developments in highly parallel processing will clearly be of interest to this approach in the future.

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- ¹Anomalous x-ray scattering typically refers to changes in the real part of atomic scattering factors as hv is tuned below the absorption edge.
- ²H. A. Dürr, E. Dudzik, S. S. Dhesi, J. B. Goedkoop, G. van der Laan, M. Belakhovsky, C. Mocuta, A. Marty, and Y. Samson, Science **284**, 2166 (1999).
- ³J. B. Kortright, S.-K. Kim, G. P. Denbeaux, G. Zeltzer, K. Takano, and E. E. Fullerton, Phys. Rev. B **64**, 092401 (2001).
- ⁴P. Abbamonte, L. Venema, A. Rusydi, G. A. Sawatzky, G. Logvenov, and I. Bozovic, Science **297**, 581 (2002).
- ⁵S. B. Wilkins, P. D. Hatton, M. D. Roper, D. Prabhakaran, and
- A. T. Boothroyd, Phys. Rev. Lett. **90**, 187201 (2003).
- ⁶K. J. Thomas, J. P. Hill, S. Grenier, Y.-J. Kim, P. Abbamonte, L. Venema, A. Rusydi, Y. Tomioka, Y. Tokura, D. F. McMorrow,
- G. Sawatzky, and M. van Veenendaal, Phys. Rev. Lett. **92**, 237204 (2004).
- ⁷S. Roy, M. R. Fitzsimmons, S. Park, M. Dorn, O. Petracic, I. V. Roshchin, Z.-P. Li, X. Batlle, R. Morales, A. Misra, X. Zhang, K. Chesnel, J. B. Kortright, S. K. Sinha, and I. K. Schuller, Phys. Rev. Lett. **95**, 047201 (2005).
- ⁸R. M. A. Azzam and N. M. Bashara, *Ellipsometry and Polarized Light* (North-Holland, Amsterdam, 1977).
- ⁹E. D. Palik, *Handbook of Optical Constants of Solids* (Academic Press, Inc., Orlando, Florida, 1985).
- ¹⁰H. G. Tompkins and W. A. McGahan, *Spectroscopic Ellipsometry* and *Reflectometry: A Users Guide* (John Wiley and Sons, New York, NY, 1999).
- ¹¹S. D. Brown, L. Bouchenoire, P. Thompson, R. Springell, A. Mirone, W. G. Stirling, A. Beesley, M. F. Thomas, R. C. C. Ward, M. R. Wells, S. Langridge, S. W. Zochowski, and G. H. Lander, Phys. Rev. B 77, 014427 (2008).
- ¹²L. Bouchenoire, A. Mirone, S. D. Brown, P. Strange, T. Wood, P. Thompson, D. Fort, and J. Fernández-Rodríguez, New J. Phys. **11**, 123011 (2009).
- ¹³J. M. Tonnerre, L. Seve, D. Raoux, G. Soullie, B. Rodmacq, and P. Wolfers, Phys. Rev. Lett. **75**, 740 (1995).
- ¹⁴M. Wormington, C. Panaccione, K. Matney, and D. Bowen, Philos. Trans. R. Soc. London A **357**, 2827 (1999).

- ¹⁵M. Sacchi, C. F. Hague, E. M. Gullikson, and J. H. Underwood, Phys. Rev. B **57**, 108 (1998).
- ¹⁶I. Diel, J. Friedrich, C. Kunz, S. DiFornzo, B. R. Muller, and W. Jark, Appl. Opt. **36**, 6376 (1997).
- ¹⁷J. E. Prieto, F. Heigl, O. Krupin, G. Kaindl, and K. Starke, Phys. Rev. B 68, 134453 (2003).
- ¹⁸A. van der Lee, F. Salah, and B. Harzallah, J. Appl. Crystallogr. 40, 820 (2007).
- ¹⁹S. M. Danauskas, D. Li, M. Meron, B. Lin, and K. Y. C. Lee, J. Appl. Crystallogr. **41**, 1187 (2008).
- ²⁰R. Soufli, A. L. Aquila, F. Salmassi, M. Fernanxez-Perea, and E. M. Gullikson, Appl. Opt. 47, 4633 (2008).
- ²¹M. Mezger, B. Jerome, J. B. Kortright, M. Valvidares, E. M. Gullikson, A. Giglia, N. Mahne, and S. Nannarone, Phys. Rev. B 83, 155406 (2011).
- ²²D. Ksenzov, T. Panzner, C. Schlemper, C. Morawe, and U. Pietsch, Appl. Opt. **48**, 6684 (2009).
- ²³A. B. Kuzmenko, Rev. Sci. Instrum. **76**, 083108 (2005).
- ²⁴S. M. Valvidares, M. Huijben, P. Yu, R. Ramesh, and J. B. Kortright, Phys. Rev. B **82**, 235410 (2010).
- ²⁵W. Meevasana, P. D. C. King, R. H. He, S. K. Mo, M. Hashimoto, A. Tamai, P. Songsiriritthigul, F. Baumberger, and Z. X. Shen, Nat. Mater. **10**, 114 (2011).
- ²⁶A. F. Santander-Syro, O. Copie, T. Kondo, F. Fortuna, S. Paihes, R. Weht, X. G. Qiu, F. Bertran, A. Niolaou, A. Taleb-Ibrahimi, P. L. Fevre, G. Herranz, M. Bibes, N. Reyren, Y. Apertet, P. Lecoeur, A. Barthelemy, and M. J. Rozenberg, Nature (London) **469**, 189 (2011).
- ²⁷K. Szot and W. Speier, Phys. Rev. B **60**, 5909 (1999).
- ²⁸C. D. Keefe, J. Mol. Spectrosc. **205**, 261 (2001).
- ²⁹D. S. Sivia and J. Skilling, *Data Analysis: A Bayesian Tutorial* (Oxford University Press Inc., New York, 2007).
- ³⁰S. F. Gull and G. J. Daniell, Nature (London) **272**, 686 (1978).
- ³¹J. Skilling and R. K. Bryan, Mon. Not. R. Astron. Soc. **211**, 111 (1984).
- ³²M. Sakata and M. Sato, Acta Crystallogr. Sect. A 46, 263 (1990).
- ³³We have implemented this algorithm in a PYTHON script, which can be made available upon request. The program is capable of implementing a number of algorithms for solving the maximum entropy equation (or performing a minimization without entropy

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considerations), as well as using a number of possible forms for the figure of merit. We generally prefer a modification of the algorithm developed by Gull and Daniell with a figure of merit of the form $(\frac{Y_{obs}-Y_{calc}}{Y_{obs}})^2$. ³⁴L. G. Parratt, Phys. Rev. **95**, 359 (1954).

- ³⁵L. Névot and P. Croce, Rev. Phys. Appl. (Paris) **15**, 761 (1980).
- ³⁶B. Henke, E. Gullikson, and J. Davis, At. Data Nucl. Data Tables 54, 181 (1993).
- ³⁷In models 3a and 3b the nonresonant surface layer is still present. We considered models with a resonant surface region in the absence of this nonresonant contamination layer, but found them to give poor fits. It was also noted that the refined spectra for the surface region contained very little resonant character, suggesting a preference for a nonresonant surface layer.
- ³⁸F. M. F. deGroot, J. C. Fuggle, B. T. Thole, and G. A. Sawatzky, Phys. Rev. B 41, 928 (1990).