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Electronic transport coefficients from *ab initio* simulations and application to dense liquid hydrogen

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Using Kubo's linear response theory, we derive expressions for the frequency-dependent electrical conductivity (Kubo-Greenwood formula), thermopower, and thermal conductivity in a strongly correlated electron system. These are evaluated within *ab initio* molecular dynamics simulations in order to study the thermoelectric transport coefficients in dense liquid hydrogen, especially near the nonmetal-to-metal transition region. We also observe significant deviations from the widely used Wiedemann-Franz law, which is strictly valid only for degenerate systems, and give an estimate for its valid scope of application toward lower densities.

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I. INTRODUCTION

An accurate description of dense hydrogen is of fundamental interest and has wide applications in astrophysics. Prominent examples are models of planetary interiors which consist predominantly of hydrogen and helium as in the case of Jupiter-like planets.^{1–3} Special attention has been paid on the high-pressure phase diagram of hydrogen and its isotopes. Major problems in this context are the slope of the melting line^{4–7} and the transition pressure to solid metallic hydrogen,^{8,9} which is expected near 4 Mbar¹⁰ at T = 0 K. Such extreme conditions are experimentally still not feasible yet.

Interestingly, metallic hydrogen has first been verified in the liquid at about 1.4 Mbar and few thousand kelvin.^{11,12} The problem of whether this transition is of first order has been discussed for decades. While numerous models within the chemical picture predict almost invariably a pronounced first-order phase transition with a critical point at about 10 000 K-15 000 K and 0.5 Mbar,¹³⁻¹⁸ only few *ab initio* simulations indicate such a behavior.^{19–23} A first-order liquid-liquid phase transition with a critical point near 2000 K and 120 GPa has been predicted from both quantum Monte Carlo (QMC) and finite-temperature density functional theory molecular dvnamics (FT-DFT-MD) simulations.²² These results were confirmed almost simultaneously by extensive FT-DFT-MD simulations,²³ which predict a critical point at about 1400 K and 132 GPa, that is, at a somewhat lower temperature. A detailed analysis of the changes in the structural and electronic properties with density and temperature clearly shows that the nonmetal-to-metal transition in dense liquid hydrogen drives this first-order phase transition.

The transport coefficients yield valuable information on the state of the strongly correlated liquid. For instance, the conductivity changes drastically along the nonmetal-tometal transition over many orders of magnitude which could already be verified experimentally.^{11,12} These measures have the potential to accurately characterize subtle changes of the electronic structure with density and temperature and, especially, alongside the liquid-liquid phase transition in hydrogen. We therefore calculate the complete set of thermoelectric transport coefficients (electrical and thermal conductivity, thermopower) via FT-DFT-MD simulations for a wide range of densities [$\rho = (0.05-20)$ g/cm³] and temperatures $[T = (1-50) \times 10^3 \text{ K}]$ and pay special attention to the nonmetal-to-metal transition region. The focus of our work is below 20 g/cm³ since higher densities are of primary importance for the physics of inertial confinement fusion and were studied previously.²⁴

The transport coefficients as well as their changes with density and temperature are again important for applications in astrophysics. For instance, a boundary between a nonconducting outer and a metallic inner envelope is usually assumed in interior models of gas giants such as Jupiter; its location should, in principle, be determined from the nonmetal-to-metal transition in hydrogen but is usually a free parameter.²⁵ Furthermore, it has been shown that demixing of helium from hydrogen occurs in H-He mixtures at megabar pressures due to the nonmetal-to-metal transition in the hydrogen subsystem.^{26,27} The treatment of this effect is essential in order to explain the excess luminosity of Saturn and its age.²⁸ Knowledge on the transport coefficients of dense liquid hydrogen is furthermore required for dynamo simulations of planetary magnetic fields.²⁹ The dynamo is driven by convection of conducting material deep in the interior so that the electrical and thermal conductivity are important input quantities. Another interesting problem is the formation of giant planets out of a protoplanetary disk. The conductivities change by many orders of magnitude during the accretion process as a consequence of the density and temperature increase, which could affect the radiation hydrodynamics of the collapsing disk and the luminosity of the young protoplanet strongly.

There are a number of theoretical models that predict the electrical and thermal conductivity with different assumptions about the electronic and ionic structure and their mutual interaction.^{16,30–38} Especially in a strongly coupled system it is difficult to calculate the transport properties accurately because the ionization degree and the effective two-particle scattering cross sections are not well defined. Strong ion-ion correlations, the dynamic nature of screening and exchange effects in the electron system as well as quantum effects such as Pauli blocking require a consistent quantum-statistical approach. Therefore, we calculate the transport coefficients within FT-DFT-MD simulations. This method has demonstrated its capacity to provide accurate data for the strongly correlated quantum regime.^{39–42}

For simple metals the relation between electrical and thermal conductivity is described by the famous Wiedemann-Franz law⁴³ using a fixed value, the Lorenz number L_0 , as proportionality constant. The benefit from this relation is the ability to obtain the thermal conductivity from the electrical conductivity very easily. For this procedure the Lorenz number $L(T,\rho)$ has to be known for a given density ρ and temperature T of the system. By calculating the Lorenz number *ab initio* the region where the original Wiedemann-Franz law is valid can be identified.

Our paper is organized as follows. We outline the FT-DFT-MD method in Sec. II. A generalization of the Kubo-Greenwood formula to calculate the Onsager coefficients L_{ik} is given in Sec. III. The results for the electrical and thermal conductivity, the thermopower, and the Lorenz number in dense liquid hydrogen are presented in Sec. IV for a wide range of densities and temperatures. In particular, the behavior predicted for the nonmetal-to-metal transition region is discussed. Conclusions are given in Sec. V.

II. THEORETICAL METHOD

We use the FT-DFT-MD framework, which combines classical molecular dynamics simulations for the ions with a quantum treatment of the electrons based on FT-DFT⁴⁴⁻⁴⁶ which is implemented in the VASP 5.2 program package.⁴⁷⁻⁴⁹ In the FT-DFT the Coulomb interactions between the electrons and the ions are treated using the projector-augmented wave method^{50,51} at densities below 9 g/cm³ with a converged energy cutoff of 800 eV. At higher densities it was necessary to perform all calculations with the Coulomb potential which required a substantial higher cutoff of 3000 eV.

The FT-DFT algorithm is used to derive the forces that act on the ions via the Hellmann-Feynman theorem at each MD step. This procedure is repeatedly performed in a cubic simulation box with periodic boundary conditions for several thousand MD time steps of 0.1 to 1 fs duration so that the total simulation time amounts up to 10 ps. The ion temperature is controlled with a Nosé thermostat.⁵²

Convergence was checked with respect to the particle numbers, which vary between 64 and 512 atoms depending on the density, the **k**-point sets used for the evaluation of the Brillouin zone, and the energy cutoff for the plane wave basis set. For the simulations we chose the Baldereschi mean value point,⁵³ which proved to yield well converged simulation runs.^{23,41} Several test calculations showed that higher efforts are necessary only in the vicinity of the phase transition.²³

The electronic transport coefficients are subsequently calculated by evaluating the respective transport formulas (see section below).⁵⁴ This is done for 10 to 20 ion configurations from the equilibrated MD simulation using Monkhorst-Pack **k**-point meshes⁵⁵ of $3 \times 3 \times 3$ to $6 \times 6 \times 6$ to reach the convergence.

The exchange-correlation functional is the most critical input in FT-DFT calculations. Here we use the approximation of Perdew, Burke, and Ernzerhof,⁵⁶ which is numerically efficient. This functional has been chosen in similar studies^{24,57,58} and reasonable results are, in general, expected in the metallic and high-temperature plasma regime.^{59,60} In the semiconducting region the obtained conductivities may be overestimated by

using this functional but will yet be useful for many practical applications.

Note that the ionic contribution to the transport coefficients is out of the scope of this work and is therefore consistently neglected.

III. TRANSPORT PROPERTIES AND ONSAGER COEFFICIENTS

In linear response theory (LRT) the response of an isotropic single component system of charge carriers (electrons with charge q = -e in our case) to an electric field **E** and a temperature gradient ∇T is expressed by the electric current \mathbf{J}_e and the heat current \mathbf{J}_q through

$$\langle \mathbf{J}_e \rangle = \frac{1}{q} \left(q L_{11} \mathbf{E} + \frac{L_{12} \nabla T}{T} \right), \tag{1}$$

$$\langle \mathbf{J}_q \rangle = \frac{1}{q} \left(q L_{21} \mathbf{E} + \frac{L_{22} \nabla T}{T} \right).$$
 (2)

The electrical conductivity σ , thermal conductivity λ , and thermopower α are then given by the Onsager coefficients L_{mn} in the following way:

$$\sigma = L_{11}, \quad \lambda = \frac{1}{T} \left(L_{22} - \frac{L_{12}^2}{L_{11}} \right), \qquad \alpha = \frac{L_{12}}{TL_{11}}.$$
 (3)

The Lorenz number is the ratio between electrical and thermal conductivity divided by the temperature,

$$L = \frac{e^2}{k_B^2 T} \frac{\lambda}{\sigma},\tag{4}$$

and is, according to the Wiedemann-Franz law,⁴³ constant in the limit of high density, where it can be calculated as $L_0 = \pi^2/3$ by means of the Sommerfeld expansion.⁶¹

Within the framework of Kubo's quantum-statistical LRT, which is described, for instance, in Refs. 62 and 63, the following expressions are obtained for the frequency-dependent Onsager coefficients $L_{mn}(\omega)$:

$$L_{mn}(\omega) = \frac{1}{3V} \langle \hat{\mathbf{J}}_m(t - i\hbar\tau); \hat{\mathbf{J}}_n \rangle_{\omega + i\varepsilon}.$$
 (5)

The current-current correlation functions are given as

$$\langle \hat{\mathbf{J}}_{m}(t-i\hbar\tau); \hat{\mathbf{J}}_{n} \rangle_{\omega+i\varepsilon}$$

$$= \lim_{\varepsilon \to 0} \int_{0}^{\infty} \mathrm{d}t \; e^{i(\omega+i\varepsilon)t} \int_{0}^{\beta} \mathrm{d}\tau \; \mathrm{Tr}\{\hat{\varrho}_{0}\hat{\mathbf{J}}_{m}\left(t-i\hbar\tau\right) \cdot \hat{\mathbf{J}}_{n}\},$$
(6)

where $\beta = (k_B T)^{-1}$ is the inverse thermal energy. The limit $\varepsilon \to 0$ has to be taken after the calculation of the thermodynamic limit, which is done by evaluating the trace. The statistical operator of the equilibrium $\hat{\varrho}_0$ contains the Kohn-Sham Hamilton operator $\hat{H}_{\rm KS}$.

The time-dependent current operators within the Heisenberg picture are defined as

$$\hat{\mathbf{J}}_m(t-i\hbar\tau) = e^{\frac{i}{\hbar}(t-i\hbar\tau)\hat{H}_{\mathrm{KS}}}\hat{\mathbf{J}}_m e^{-\frac{i}{\hbar}(t-i\hbar\tau)\hat{H}_{\mathrm{KS}}}.$$
(7)

The electric current operator $\hat{\mathbf{J}}_e = \hat{\mathbf{J}}_1$ and the heat current operator $\hat{\mathbf{J}}_q = \hat{\mathbf{J}}_2$ read in second quantization of spin degenerate Bloch states:

$$\hat{\mathbf{J}}_{m} = \sum_{\mathbf{k}\mathbf{k}'\nu\nu'} \langle \mathbf{k}\nu | \hat{\mathbf{j}}_{m} | \mathbf{k}'\nu' \rangle \hat{a}^{\dagger}_{\mathbf{k}\nu} \hat{a}_{\mathbf{k}'\nu'}.$$
(8)

Here **k** is the wave number and ν the band index. The electric and heat current operators are given as

$$\hat{\mathbf{j}}_1 = \frac{q}{m_e} \hat{\mathbf{p}},\tag{9}$$

$$\hat{\mathbf{j}}_2 = \frac{1}{m_e} \frac{\hat{H}\hat{\mathbf{p}} + \hat{\mathbf{p}}\hat{H}}{2} - h_e \hat{\mathbf{p}},\tag{10}$$

where h_e is the enthalpy per electron (see Ref. 64 for further information on the definition of the currents).

The eigenvalues of the Hamiltonian can be evaluated so that Eq. (8) can be simplified,

$$\hat{\mathbf{J}}_{m} = \frac{q^{2-m}}{m_{e}} \sum_{\mathbf{k}\mathbf{k}'\nu\nu'} \langle \mathbf{k}\nu | \hat{\mathbf{p}} | \mathbf{k}'\nu' \rangle \epsilon_{\mathbf{k}\nu\mathbf{k}'\nu'}^{m-1} \hat{a}_{\mathbf{k}\nu}^{\dagger} \hat{a}_{\mathbf{k}'\nu'}, \qquad (11)$$

where

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$$\epsilon_{\mathbf{k}\nu\mathbf{k}'\nu'} = \left(\frac{E_{\mathbf{k}\nu} + E_{\mathbf{k}'\nu'}}{2} - h_e\right) \tag{12}$$

is used. After inserting Eq. (11) into Eq. (6) and some operator algebra the trace can be evaluated according to Wick's theorem as

$$\operatorname{Tr}\{\hat{\varrho}_{eq}\hat{a}_{\mathbf{k}\nu}^{\dagger}\hat{a}_{\mathbf{k}'\nu'}\hat{a}_{\mathbf{p}\mu}^{\dagger}\hat{a}_{\mathbf{p}'\mu'}\} = f_{\mathbf{k}\nu}\delta_{\mathbf{k},\mathbf{k}'}\delta_{\nu,\nu'}f_{\mathbf{p}\mu}\delta_{\mathbf{p},\mathbf{p}'}\delta_{\mu,\mu'} + f_{\mathbf{k}\nu}\delta_{\mathbf{k},\mathbf{p}'}\delta_{\nu,\mu'}(1-f_{\mathbf{p}\mu})\delta_{\mathbf{p},\mathbf{k}'}\delta_{\mu,\nu'}.$$
(13)

The first term vanishes and the Fermi functions are defined as $f_{\mathbf{k}\nu} = (e^{\beta(E_{\mathbf{k}\nu}-\mu_e)}+1)^{-1}$. Altogether the trace reads

$$\Gamma \{ \hat{\varrho}_{0} \hat{\mathbf{J}}_{m}(t - i\hbar\tau) \cdot \hat{\mathbf{J}}_{n} \} = \frac{q^{4-m-n}}{m_{e}^{2}} \sum_{\mathbf{k}\mathbf{p}\nu\mu} e^{\frac{i}{\hbar}(t - i\hbar\tau)\Delta E} f_{\mathbf{k}\nu}(1 - f_{\mathbf{p}\mu}) \langle \mathbf{k}\nu | \hat{\mathbf{p}} | \mathbf{p}\mu \rangle \\ \times \langle \mathbf{p}\mu | \hat{\mathbf{p}} | \mathbf{k}\nu \rangle \epsilon_{\mathbf{k}\nu\mathbf{p}\mu}^{m-1} \epsilon_{\mathbf{p}\mu\mathbf{k}\nu}^{n-1},$$
(14)

with $\Delta E = E_{\mathbf{k}\nu} - E_{\mathbf{p}\mu}$. Now the τ integration can be performed:

$$\int_{0}^{\beta} \mathrm{d}\tau \; e^{\tau \Delta E} = \frac{e^{\beta \Delta E} - 1}{\Delta E}.$$
 (15)

In the same way the second integral can be solved:

$$\lim_{\varepsilon \to 0} \int_0^\infty \mathrm{d}t e^{(-\varepsilon + i\omega + \frac{i}{\hbar}\Delta E)t} = \hbar\pi \,\delta \left(\Delta E + \hbar\omega\right) + i\hbar \,\mathcal{P}\left(\frac{1}{\Delta E + \hbar\omega}\right). \tag{16}$$

Only the real part of this equation is considered here, since the imaginary part can be calculated more easily with a Kramers-Kronig relation. The Onsager coefficients now

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$$L_{mn} = \frac{\hbar\pi q^{4-m-n}}{3Vm_e^2} \sum_{\mathbf{k}\mathbf{p}\nu\mu} \frac{e^{\beta\Delta E} - 1}{\Delta E} f_{\mathbf{k}\nu} (1 - f_{\mathbf{p}\mu}) \\ \times \langle \mathbf{k}\nu | \hat{\mathbf{p}} | \mathbf{p}\mu \rangle \cdot \langle \mathbf{p}\mu | \hat{\mathbf{p}} | \mathbf{k}\nu \rangle \delta(\Delta E + \hbar\omega) \epsilon_{\mathbf{k}\nu\mathbf{p}\mu}^{m-1} \epsilon_{\mathbf{p}\mu\mathbf{k}\nu}^{n-1}.$$
(17)

In position representation the matrix elements have the following form:

$$\langle \mathbf{k}\nu | \hat{\mathbf{p}} | \mathbf{p}\mu \rangle = \delta_{\mathbf{p},\mathbf{k}} \bigg(\hbar \mathbf{k} \, \delta_{\nu,\mu} + \frac{1}{V} \int_{V} \mathrm{d}^{3} \mathbf{r} \, u_{\mathbf{k}\nu}^{*}(\mathbf{r}) \hat{\mathbf{p}} \, u_{\mathbf{k}\mu}(\mathbf{r}) \bigg). \quad (18)$$

Here V is the volume of the simulation box and the functions $u_{\mathbf{k},\nu}^{*}(\mathbf{r})$ and $u_{\mathbf{k},\mu}(\mathbf{r})$ are the Bloch factors. Because of the first Kronecker symbol the matrix elements are diagonal concerning the wave number, which eliminates the p sum. The spin summation leads to an additional factor of two. After using a relation between Fermi functions we finally arrive at the following expression for the Onsager coefficients:

$$L_{mn}(\omega) = \frac{2\pi q^{4-m-n}}{3Vm_e^2 \omega} \sum_{\mathbf{k}\nu\mu} \langle \mathbf{k}\nu | \hat{\mathbf{p}} | \mathbf{k}\mu \rangle \cdot \langle \mathbf{k}\mu | \hat{\mathbf{p}} | \mathbf{k}\nu \rangle$$
$$\times \epsilon_{\mathbf{k}\nu\mathbf{k}\mu}^{m+n-2} (f_{\mathbf{k}\nu} - f_{\mathbf{k}\mu}) \delta(E_{\mathbf{k}\mu} - E_{\mathbf{k}\nu} - \hbar\omega). \quad (19)$$

The Onsager coefficients (19) obey the symmetry relations $L_{mn}(\omega) = L_{mn}(-\omega)$ and $L_{mn}(\omega) = L_{nm}(\omega)$. The coefficient L_{11} is known as the frequency-dependent Kubo-Greenwood formula^{65,66} and has been widely applied in FT-DFT-MD simulations.39,67,68

Similar but not identical frequency-dependent formulas for L_{12} and L_{22} were given by Recoules *et al.*⁴⁰ but not formally derived in their work. The different formulations lead to deviations in the results for the thermopower α and in the thermal conductivity λ at nonzero frequencies. Furthermore, the heat current from Ref. 40 contains the chemical potential μ_e instead of the enthalpy h_e [see Eq. (10)]. However, it can be shown that the term proportional to the enthalpy per particle h_e has no influence on the thermal conductivity λ in one-component systems. As a consequence the numerical results of Refs. 24 and 40 could be reproduced by Eq. (19) in the limit of $\omega \to 0$ and are therefore not questioned by the current work. Differences, however, occur in the static results for the thermopower α .

The derivation in this chapter can be easily generalized to the spin-dependent form as well as to anisotropic systems.

IV. RESULTS

A. Electrical conductivity

Figure 1 shows isotherms of the electrical conductivity over a wide range of densities and temperatures. In previous work⁴¹ it could already be shown that these theoretical results agree well with data for the reflectivity and conductivity derived from shock-wave experiments. Here we concentrate on the general behavior of the theoretical curves on a large density and temperature scale and give results especially in the vicinity of the nonmetal-to-metal transition.



FIG. 1. (Color online) Electrical conductivity in dense hydrogen for different temperatures from the region of the liquid-liquid phase transition up to the dense plasma. At low densities the values increase with temperature and decrease with temperature at high densities.

In general, the conductivity rises in the whole range studied here with increasing density. On the one hand, this is the result of pressure ionization which increases the amount of conducting electrons, and on the other hand, even in the case of a fully ionized system, a rising density of electrons results in a further increase of the conductivity.

At T = 1000 K the liquid-liquid phase transition which was previously reported for dense hydrogen^{22,23} can be identified by the steep increase of the conductivity over several orders of magnitude in a small density interval, (0.7 - 0.9)g/cm³. At temperatures higher than 1500 K the transition is continuous and the increase spreads to a larger range in density. At densities below this transition the electrical conductivity increases with temperature which is caused by thermal ionization of neutral particles. The additional free charges contribute to the conductivity.

At densities above this transition the dependence on temperature is inverted: The electrical conductivity decreases with temperature, which is typical for metals. Increasing temperature broadens the Fermi function and therefore allows additional electron scattering processes which reduce their mobility. As a result, the conductivity is lower with increasing temperature.

This behavior is also illustrated in Fig. 2, which shows the electrical conductivity along isochores for different temperatures. The conductivity decreases for temperatures above 2000 K along the isochores for densities higher than 0.9 g/cm^3 , which are characteristic of the metallic phase. This indicates that most of the system is ionized and thus acts metal-like. Looking at lower densities the conductivity rises along the whole temperature range, which is due to thermal ionization. The isochores clearly show the general behavior of a rising conductivity with increasing density.

Experiments in copper plasmas⁶⁹ have indicated that the electrical conductivity becomes a function of only the coupling parameter Γ for values of $\Gamma \ge 10$. The plasma parameter Γ is defined by

$$\Gamma = \frac{e^2}{4\pi\varepsilon_0 k_B T} \left(\frac{4\pi n_e}{3}\right)^{\frac{1}{3}},\tag{20}$$



FIG. 2. (Color online) Electrical conductivity for different densities versus temperature.

where n_e is the number density of free electrons. Although such a behavior could not be confirmed later,⁷⁰ a simple functional form of the electrical conductivity at high values of Γ is still under discussion. To investigate whether such a simple scaling is valid in dense liquid hydrogen, the results for the electrical conductivity shown in Figs. 1 and 2 are plotted against Γ in Fig. 3. Only for temperatures higher than 10 000 K is the system strongly ionized. At lower temperatures the occurrence of partial ionization prevents a proper calculation of Γ within FT-DFT-MD, because the method does not distinguish between bound and free electrons. Therefore, we do not plot results for lower temperatures in Fig. 3.

The isotherms appear to be almost parallel in this logarithmic plot and are clearly separated. Even at the highest available values for $\Gamma \leq 60$ the isotherms do not tend to merge. We conclude that it is not possible to derive a simple temperature-independent relation for the conductivity that depends solely on the parameter Γ as it was proposed earlier for other metallic liquids.



FIG. 3. (Color online) Electrical conductivity as function of the coupling parameter Γ for different temperatures. The plotted data cover a density range of (0.05–20) g/cm³.



FIG. 4. (Color online) Thermal conductivity versus density for different temperatures. The values increase with temperature.

B. Thermal conductivity

The isotherms of the thermal conductivity are plotted in Fig. 4. They show a similar behavior as the electrical conductivity and indicate a sharp nonmetal-to-metal transition at a temperature below 1500 K. At this transition the thermal conductivity increases over several orders of magnitude in a narrow density range at about 0.9 g/cm³. The increase in the thermal conductivity is due to the growing number of delocalized electrons which are produced along the phase transition. Above the critical temperature this transition becomes broader and is caused by a combination of pressure and temperature ionization.

In contrast to the behavior of the electrical conductivity, the thermal conductivity does not decrease with temperature in the metallic phase. The isotherms increase systematically with temperature for all densities.

This is also shown along the isochores of the thermal conductivity that are plotted in Fig. 5. These curves depict likewise that the thermal conductivity rises invariably with increasing density and temperature.

C. Thermopower

The thermopower α characterizes the generation of an electric field as a response to a temperature gradient. For most systems this electric field has a direction opposite to the temperature gradient, which results in a negative thermopower. The thermopower is most sensitive to changes in the electronic structure since it can be expressed as the derivative of the logarithm of the electronic conductivity with respect to the energy at the Fermi surface.⁷¹ Such a relation, which is also known as Mott formula, follows from the Kubo-Greenwood equation (19) in the degenerate domain under strong scattering conditions.

Interestingly, large positive values for the thermopower were measured in fluid mercury^{72,73} near the liquid-vapor critical point, which is located at $T_c = 1751$ K and $\rho_c = 5.8$ g/cm³. In this region, isotherms of the electrical conductivity near T_c show a strong increase with the density, which is steepest just at the critical density ρ_c . This behavior was assumed to be related to fluctuations in the electron density, which are pronounced near the critical point due to critical fluctuations. In particular, a zero of the thermopower was observed exactly at the critical density. The interesting question arises as to whether a zero of the thermopower is a precursor of a first-order phase transition in dense liquids which undergo a nonmetal-to-metal transition. Wide regions with a positive thermopower have been predicted for dense hydrogen by a simple chemical model³⁴ but were not confirmed in an advanced chemical approach¹⁶ by *ab initio* simulations or by experiments yet.

In Fig. 6 isotherms of the thermopower α are plotted as function of the density. The symbols represent the results from the simulations and the error bars show the statistical uncertainties. As guide to the eye, polynomial functions were fitted to the numerical results. The thermopower is mostly negative and reaches a value of about zero at high densities. At lower densities the thermopower decreases. The low density limit of $\alpha = -60.60 \ \mu V/K$ is known from the Spitzer theory (see, e.g., Ref. 16). With higher temperatures the negative values become systematically larger. We expect that these values become smaller again at low densities to reach the Spitzer limit. The thermopower shows positive mean values below 20 000 K and between 0.2 g/cm³ and 0.5 g/cm³. These



FIG. 5. (Color online) Thermal conductivity versus temperature for different densities.



FIG. 6. (Color online) Thermopower versus density for different temperatures.



FIG. 7. (Color online) Thermopower versus temperature for different densities.

values are very small and, in fact, are about the size of the uncertainties of the respective mean values. We therefore cannot conclude definitely that significantly positive values occur within the accuracy of our ab initio calculations. To answer the question as to whether the zero of the thermopower occurs at the critical point of the liquid-liquid phase transition in dense hydrogen, which is predicted^{22,23} at about 0.8 g/cm³, the thermopower has to be evaluated for temperatures below 2000 K since $T_c = (1400-2000)$ K. For this region, reasonable results can be given for the electrical and thermal conductivity but not for the thermopower. This is because the values of the Onsager coefficients each decrease over orders of magnitude in systems with a majority of localized electrons, while especially the decrease of L_{12} is two orders of magnitude weaker than that of L_{11} . This behavior, on the one hand, causes the second contribution to the thermal conductivity in Eq. (3) to vanish, because L_{12} is squared in the numerator. As a result the thermal conductivity becomes proportional to L_{22} , which can be evaluated successfully in this case, as well as L_{11} . On the other hand, this leads to an enormous increase in the statistical fluctuations of the thermopower, which causes uncertainties that reach hundreds of $\mu V/K$.

In Fig. 7 isochores of the thermopower α are plotted. The thermopower decreases with higher temperature, which is more pronounced the lower the density is. Positive values appear at low temperatures, but the statistical error is too large to prove this result unambiguously.

D. Lorenz number

The Lorenz number Eq. (4) describes mainly the relation between thermal and electrical conductivity. For simple metals this relation is described via a constant and is known as the Wiedemann-Franz law. For high degeneracy this constant is $L_0 = \pi^2/3.^{61}$ This relation can be used to easily obtain the thermal conductivity for metals if the electrical conductivity is known. Here we calculate the Lorenz number for dense hydrogen in order to identify the region where the Wiedemann-Franz law is valid. High degeneracy occurs only at sufficiently high densities so that we expect deviations from the Wiedemann-Franz law at lower densities and higher temperatures.



FIG. 8. (Color online) Lorenz number versus density for different temperatures. The dashed black line displays the limiting value of $L_0 = \pi^2/3$.

The Lorenz number is shown in Fig. 8 for several temperatures. At densities above 1 g/cm³ the Lorenz number is almost constant and shows no temperature dependence, which indicates that the Wiedemann-Franz law is valid there. The value of $L_0 = \pi^2/3$, which is shown by a dashed black line, could be reproduced for each temperature at densities higher than (1–2) g/cm³ within the statistical uncertainties. This behavior is consistent with the metalliclike properties observed already for the conductivity at high densities.

At densities below (1-2) g/cm³ the Lorenz number shows strong deviations from the Wiedemann-Franz law with a pronounced temperature dependence. In particular, the Lorenz number rises strongly with decreasing density. This behavior becomes more pronounced at lower temperatures. On the other hand, at the lower temperatures the validity of the Wiedemann-Franz law extends to smaller densities, for example, down to 0.2 g/cm³ at 10 000 K and 15 000 K.

Along the isotherms of 30 000 K and 50 000 K a maximum can be identified. Thus, at lower densities the Lorenz number decreases again. At 30 000 K this maximum is found at a smaller density than at 50 000 K and shows a larger value. This indicates that with decreasing temperature the maximum shifts systematically to smaller densities and increases. For temperatures below 30 000 K the maximum is not displayed. It appears to be situated at densities below 0.05 g/cm³, which is the minimal density calculated here and which displays the current limit of our computational capabilities. However, we expect a maximum to appear for all temperatures because the Lorenz number is also known exactly at low densities from the Spitzer theory for fully ionized plasmas.³⁰ The latter predicts a constant low-density value of 1.5966 (see, e.g., Ref. 16).

Our results indicate that the deviations from the Wiedemann-Franz law might reach a full order of magnitude in certain regions of density and temperature. Further research will be necessary to investigate this important aspect in more detail. We also expect that certain deviations from the Wiedemann-Franz law occur in warm dense matter of arbitrary composition. In hydrogen, we predict that the Wiedemann-Franz is valid at densities above 2 g/cm³ for temperatures below 50 000 K (see Fig. 8).

V. CONCLUSIONS

We derived formulas for the Onsager transport coefficients L_{mn} within the Kubo theory for application in *ab initio* simulations. Using these expressions we calculated the electrical and thermal conductivity as well as the thermopower of liquid hydrogen for a wide range of temperatures and densities in the megabar range. In particular, we characterize the nonmetal-to-metal transition in hydrogen by observing a rapid increase in both the electrical and the thermal conductivity. The thermopower shows a trend toward positive values in a region where the critical point of the liquid-liquid phase transition is expected, similar to the behavior of liquid mercury. At low temperatures more accurate calculations will be necessary for this transport coefficient which are beyond the scope of the currently available computer capacity. The ab initio calculation of the Lorenz number shows, in addition, that the validity of the original Wiedemann-Franz law is limited to the metalliclike regime of hydrogen. The theoretical framework given here can be applied to bulk material calculations for arbitrary materials within FT-DFT-MD.

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