

Effect of strong correlations on the disorder-induced zero bias anomaly in the two-site Anderson-Hubbard model

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Several recent exact-diagonalization calculations have established that the Anderson-Hubbard model has a disorder-induced zero bias anomaly (ZBA) (also called a disorder-induced pseudogap) in the density of states. In order to understand the physics of the ZBA, we study a simplified problem—an ensemble of two-site molecules with random site energies—for which analytical results are possible. For this ensemble, we examine how the ZBA forms in both the weakly correlated (mean field) and strongly correlated limits. In the weakly correlated case, the ZBA can be understood as the result of level repulsion between bonding and antibonding molecular orbitals. A similar level repulsion occurs in the strongly correlated case too but a larger contribution to the ZBA comes from the suppression of a triplet excitation mode. This inherently many-body mechanism does not have a counterpart in mean-field models.

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I. INTRODUCTION

A number of recent papers have shown the existence of a disorder-induced zero bias anomaly (ZBA) in the Anderson-Hubbard model (AHM) in one and two dimensions.^{1–5} These calculations have revealed that there is a V-shaped dip in the density of states at the Fermi energy ε_F . This dip is produced by the response of the inelastic self-energy to the disorder potential.^{1,5} Such a mechanism is well understood in conventional metals and insulators, where the effect was explained at the level of Hartree-Fock theory by Altshuler and Aronov.⁶ However, strong correlation effects are generally important in the AHM, and the Altshuler-Aronov mechanism is thus insufficient for this case.⁵

The AHM is the standard model for strongly correlated systems with disorder. Like the Hubbard model, electrons are assumed to move on a tight-binding lattice of atomiclike orbitals. A zero-range intraorbital Coulomb interaction U is included but longer-range interorbital interactions are neglected. Strong correlations are important when the intersite hopping matrix element t is small relative to U . The AHM differs from the Hubbard model by the addition of disorder, which is introduced by selecting the orbital energies ε_i from a random distribution of width Δ (the subscript i labels sites in the atomic lattice). The Hamiltonian is

$$H = -t \sum_{\langle ij \rangle, \sigma} c_{i\sigma}^\dagger c_{j\sigma} + \sum_i (\varepsilon_i \hat{n}_i + U \hat{n}_{i\uparrow} \hat{n}_{i\downarrow}), \quad (1)$$

where $\langle ij \rangle$ restricts the sum to nearest-neighbor sites, $\hat{n}_{i\sigma}$ is the number operator for site i and spin σ , $\hat{n}_i = \sum_\sigma \hat{n}_{i\sigma}$, and $\varepsilon_i \in [-\frac{1}{2}\Delta, \frac{1}{2}\Delta]$. (We use \hat{O} to indicate the operator form of an observable O ; thus $n_i = \langle \hat{n}_i \rangle$.) In this model, the ensemble-averaged density is $n=1$ (i.e., the band is half filled) for $\varepsilon_F = \frac{U}{2}$.

The conventional Altshuler-Aronov theory predicts that the Hartree and exchange self-energies make positive and negative contributions to the density of states at ε_F ,

respectively.⁶ The exchange self-energy is typically much larger than the Hartree self-energy, and the net result is a depletion of states at ε_F . However, the AHM has a zero-range interaction for which the exchange self-energy vanishes. Altshuler-Aronov theory predicts a peak in this case, which is illustrated by the nonmagnetic Hartree-Fock calculations in Fig. 1. This is in contrast to the V-shaped dip found in exact-diagonalization calculations.^{2–4} The Altshuler-Aronov prediction assumes a nonmagnetic ground state, and a number of unrestricted Hartree-Fock calculations have found a V-shaped dip at ε_F (Refs. 7–10) in the magnetic phase.⁹ While the unrestricted Hartree-Fock results are qualitatively similar to the exact-diagonalization results, there are some important differences. Notably, the ZBA in the unrestricted Hartree-Fock calculations grows with increasing U , eventually forming a broad soft gap when U is sufficiently

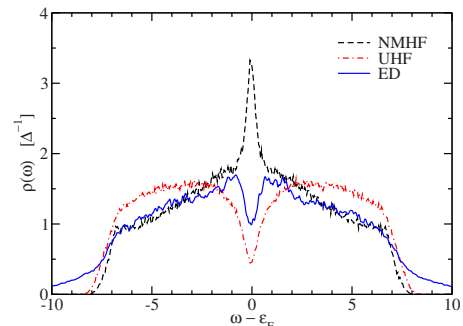


FIG. 1. (Color online) Comparison of densities of states for the Anderson-Hubbard model using different approximations. Results are shown for the nonmagnetic Hartree-Fock (NMHF), unrestricted Hartree-Fock (UHF), and exact diagonalization (ED) of small clusters. Hartree-Fock calculations are ensemble-averaged self-consistent calculations for 10×10 lattices and 1000 impurity configurations. Exact diagonalization calculations are for a 12-site lattice and 1000 impurity configurations. Model parameters are $\Delta = 20$, $U = 8$, $t = 1$, and $\varepsilon_F = U/2$, corresponding to half filling.

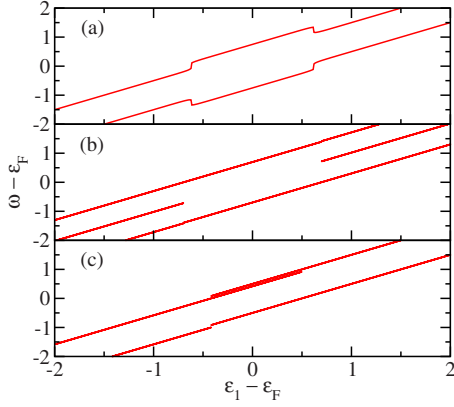


FIG. 2. (Color online) Low-energy excitation spectra as a function of site energy for the degenerate two-site model. Lines represent peak position ω in the tunneling density of states, plotted as a function of ϵ_1 for $t=0.5$. (a) Excitation spectrum for the mean-field Hamiltonian (3) with $V=0.5$ and $\epsilon_2=\epsilon_1$; (b) Excitation spectrum for exact-diagonalization calculations with $\epsilon_2=\epsilon_1-U$ and (c) $\epsilon_2=\epsilon_1$. Both (b) and (c) have $U=12$.

large. In contrast, the ZBA in exact-diagonalization calculations saturates for large U (provided $U < \Delta$; a Mott gap opens for $U \geq \Delta$), and empirically has a width $\sim t$.^{2,5} Densities of states based on the different approximations are illustrated in Fig. 1.

We note that the above discussion ignores the low-energy soft gap^{3,4,10} that has been inferred from exact diagonalization in one dimension, and found in unrestricted Hartree-Fock calculations in one and three dimensions. This gap appears on a scale $|\omega - \epsilon_F| \lesssim O(0.1t)$, and has been ascribed to long-range correlations.^{3,4} The current work examines the two-site AHM where long range correlations are absent, and there is no soft gap.

The advantage of the two-site AHM is that it is simple enough that analytical results are possible, and yet is sufficiently rich to explain much of the physics of the ZBA in larger systems.¹¹ Here, our goal is to compare the two-site AHM to a simple mean-field two-site model in order to answer the question: in what way is the ZBA in strongly correlated systems different from that in conventional metals?

Our main results are summarized in Fig. 2. In Fig. 2(a), we plot the low-energy excitation spectrum for a pair of sites with energies $\epsilon_1 = \epsilon_2$ using the mean-field model described in Sec. II. This model is meant to illustrate the conventional Altshuler-Aronov mechanism for the ZBA. Without interactions, the hybridization of atomic orbitals leads to a level splitting of $2t$ between bonding and antibonding orbitals. With interactions, there is a range of site energies near ϵ_F where the level splitting is larger than $2t$. This enhancement of level splitting (i.e., this level repulsion) shifts spectral weight away from ϵ_F and is the origin of the ZBA in this model. In Fig. 2(b), an alternative mechanism for shifting spectral weight away from ϵ_F is presented. In this case, exact results for the low-energy excitation spectrum of the two-site AHM are shown. We have taken $\epsilon_1 = \epsilon_2 + U$, which means that the lower Hubbard orbital of site 1 is degenerate with the upper Hubbard orbital of site 2. Here, the spectrum has three excitation poles, the middle of which is a triplet excitation.

The gap which is evident in the triplet spectrum is one of the main reasons for the pronounced ZBA in the two-site AHM, and is an inherently many-body mechanism that lies outside the mean-field Altshuler-Aronov paradigm. Finally, in Fig. 2(c), we show that interactions have little effect on the spectrum if we consider the case of degenerate orbitals $\epsilon_1 = \epsilon_2$. In this case, excitation spectra are shifted by $\sim O(t^2/U)$ from their noninteracting values.

The goal of this paper is to explore the physics behind these results. We discuss the mean-field mechanism for the ZBA in Sec. II, and emphasize in particular the role of level repulsion. We then derive, in Sec. III, an expression for the ensemble-averaged density of states for the two-site AHM. Finally, we discuss in Sec. IV the different mechanisms by which the ZBA found in Sec. III arises.

II. ZBA IN MEAN-FIELD THEORY

It is worth reviewing briefly how the ZBA arises in conventional metals. A variety of physical explanations for the Altshuler-Aronov ZBA have been given,^{6,12,13} and in this work we adopt the language of level repulsion.¹⁴

We consider an ensemble of two-site AHMs with randomly chosen site energies. Since we restrict ourselves to nonmagnetic solutions of the Hartree-Fock equations, a V-shaped ZBA is possible only if a nonlocal interaction is included. We therefore add a repulsive interaction $V\hat{n}_1\hat{n}_2$ to the Hamiltonian. In Hartree-Fock theory,

$$V\hat{n}_1\hat{n}_2 \rightarrow V(n_1\hat{n}_2 + n_2\hat{n}_1) - V \sum_{\sigma} \langle \langle c_{1\sigma}^{\dagger} c_{2\sigma} \rangle \rangle c_{2\sigma}^{\dagger} c_{1\sigma} + \text{H.c.} \quad (2)$$

The first and second terms are the Hartree and exchange contributions, respectively, and there is an additional Hartree contribution $\frac{1}{2}U\sum_i n_i\hat{n}_i$ from the on-site interaction. The Hartree contribution to the density of states is small for weak disorder⁶ but is central to the physics of the Coulomb gap for large disorder; the exchange contribution is largest for weak disorder, and underlies the Altshuler-Aronov mechanism for the ZBA. Because our goal is to contrast the Altshuler-Aronov mechanism with the physics of the AHM, we discuss only the exchange term.

Neglecting the Hartree contributions, we obtain the mean-field exchange Hamiltonian

$$H_X = \sum_i \epsilon_i \hat{n}_i - \tilde{t} \sum_{\sigma} (c_{1\sigma}^{\dagger} c_{2\sigma} + c_{2\sigma}^{\dagger} c_{1\sigma}), \quad (3)$$

where the renormalized hopping matrix element is $\tilde{t} = t + V\langle c_{1\sigma}^{\dagger} c_{2\sigma} \rangle$. The eigenenergies of H_X are

$$E_{X,\pm} = \frac{\epsilon_1 + \epsilon_2}{2} \pm \sqrt{\left(\frac{\epsilon_1 - \epsilon_2}{2}\right)^2 + \tilde{t}^2}, \quad (4)$$

and a straightforward calculation yields

$$\langle c_{1\sigma}^{\dagger} c_{2\sigma} \rangle = -\tilde{t} \frac{f(E_{X,+}) - f(E_{X,-})}{E_{X,+} - E_{X,-}}. \quad (5)$$

Equations (4) and (5) allow \tilde{t} to be determined self-consistently for each (ϵ_1, ϵ_2) pair. The ensemble-averaged

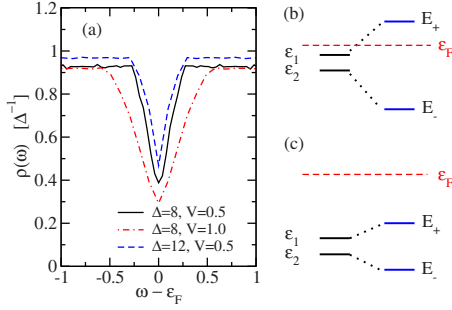


FIG. 3. (Color online) Zero bias anomaly for self-consistent solutions of the mean-field Hamiltonian (3). (a) The density of states for $t=1$ and different Δ and V . (b) The exchange self-energy is nonzero for configurations of ϵ_1 and ϵ_2 for which $E_+ > \epsilon_F > E_-$. This leads to an enhanced level repulsion relative to configurations, as in (c) where E_+ and E_- are on the same side of ϵ_F and the exchange self-energy vanishes.

density of states for this model exhibits a ZBA, as shown in Fig. 3.

The term “level repulsion” refers to the fact that the level spacing between molecular eigenenergies is greater than the level spacing between the original atomic energies, namely, $E_{X,+} - E_{X,-} > |\epsilon_1 - \epsilon_2|$. For a repulsive interaction V , $\tilde{t} > t$ and the level repulsion is enhanced by the exchange self-energy. This enhanced level repulsion, by itself, does not lead to a dip in the density of states; it is necessary that the amount of level repulsion depend on the values of $E_{X,\pm}$ relative to ϵ_F . At zero temperature, Eq. (5) shows that \tilde{t} is different from t only if $E_{X,+} > \epsilon_F > E_{X,-}$, as illustrated schematically in Fig. 3. This has the effect of pushing states away from ϵ_F , as shown numerically for the case $\epsilon_1 = \epsilon_2$ in Fig. 2(a). In this language, the ZBA in conventional metals is understood as level repulsion between filled and empty molecular orbitals near ϵ_F .

III. APPROXIMATE DIAGONALIZATION OF THE TWO-SITE AHM

We now turn to an approximate solution of the two-site AHM that preserves strong-correlation physics. We work in the strongly correlated limit $U \gg t$, where we can isolate terms that contribute to the density of states on the energy scale t . Higher-order terms, which contribute on the scale t^2/U , are neglected. We begin with a brief review of the atomic limit ($t=0$), where interactions already have a non-trivial effect on the density of states, and then show how the density of states is modified by a nonzero t .

A. Atomic limit

The density of states can be found exactly in the atomic limit $t=0$. Each site is independent, and the ground state $|G_i\rangle$ for the i th site is

$$|G_i\rangle = \begin{cases} |0\rangle, & \epsilon_F < \epsilon_i \\ |\uparrow\rangle, & \epsilon_F - U < \epsilon_i < \epsilon_F \\ |2\rangle, & \epsilon_i < \epsilon_F - U. \end{cases} \quad (6)$$

We have assumed a weak Zeeman splitting so that spin-up states are preferred when there is an odd number of elec-

trons. The spin-averaged retarded Green’s function for the i th site is

$$G_i(\omega) = \frac{1}{2} \sum_{m\sigma} \left[\frac{|\langle m|c_{i\sigma}|G_i\rangle|^2}{\omega^+ - E_{G_i} + E_m} + \frac{|\langle m|c_{i\sigma}^\dagger|G_i\rangle|^2}{\omega^+ + E_{G_i} - E_m} \right] = \frac{1 - n_i/2}{\omega^+ - \epsilon_i} + \frac{n_i/2}{\omega^+ - \epsilon_i - U}, \quad (7)$$

where $\omega^+ = \omega + i0$, $n_i = \sum_\sigma \langle \hat{n}_{i\sigma} \rangle$, $|m\rangle$ are a complete set of excited states with energies E_m , and E_{G_i} is the ground-state energy. The spin-averaged density of states at site i is thus

$$\rho_{\epsilon_i}(\omega) = -\frac{1}{\pi} \text{Im} G_i(\omega) \quad (8)$$

$$= \left(1 - \frac{n_i}{2}\right) \delta(\omega - \epsilon_i) + \frac{n_i}{2} \delta(\omega - \epsilon_i - U). \quad (9)$$

This equation shows that (i) strong correlations split the local spectrum at each site into a pair of poles at ϵ_i and $\epsilon_i + U$ and (ii) the weight of each pole depends on the electron density at that site. We refer to the poles at ϵ_i and $\epsilon_i + U$ as the lower Hubbard orbitals (LHO) and upper Hubbard orbitals (UHO), respectively. It is worth emphasizing that the energies of the LHO and UHO determine the total charge density at each site. From Eq. (6),

$$n_i = \begin{cases} 0, & \epsilon_F < \epsilon_i \text{ (LHO and UHO above } \epsilon_F) \\ 1, & \epsilon_i < \epsilon_F < \epsilon_i + U \text{ (LHO below; UHO above)} \\ 2, & \epsilon_i + U < \epsilon_F \text{ (LHO and UHO below } \epsilon_F). \end{cases} \quad (10)$$

At half filling ($\epsilon_F = U/2$), the ensemble-averaged density of states is

$$\rho(\omega) = \frac{1}{\Delta} \int_{-\Delta/2}^{\Delta/2} d\epsilon \rho_\epsilon(\omega) = \frac{1}{\Delta} \left[\Theta\left(\omega - U + \frac{1}{2}\Delta\right) \Theta\left(\frac{1}{2}\Delta - \omega\right) + \frac{1}{2} \Theta\left(\omega + \frac{1}{2}U\right) \Theta\left(\frac{3}{2}U - \omega\right) \right], \quad (11)$$

where $\Theta(x)$ is the step function. The result (11) is illustrated in Fig. 4. This figure explicitly shows the spectral weight contributed by the LHO and UHO in their different filling states. For this work, the most important aspects of the figure are (i) that both LHO and UHO contribute spectral weight at ϵ_F for $\epsilon_F \in [U - \frac{1}{2}\Delta, \frac{1}{2}\Delta]$, and (ii) that for this range of ϵ_F there is a “central plateau” where interactions enhance $\rho(\epsilon_F)$ relative to the noninteracting value Δ^{-1} .

B. Two-site case

1. Preliminary discussion

The results of exact numerical calculations of the density of states are shown in Fig. 5 for two cases: $U < \Delta$ and $U > \Delta$. We track the evolution of the density of states as a function of ϵ_F in both cases. When $U < \Delta$, there is a broad ZBA centered at ϵ_F for $\epsilon_F = 3$ and $\epsilon_F = 5$. However, the ZBA is unresolvable when ϵ_F is outside the central plateau. When

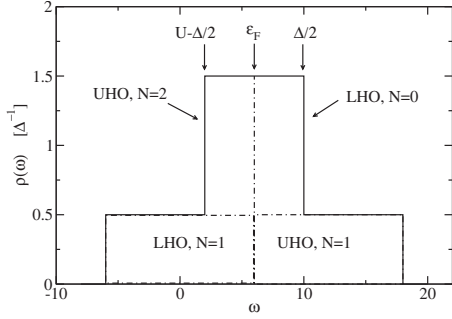


FIG. 4. Density of states in the atomic limit. The figure shows different contributions to the ensemble-averaged density of states in the limit $t=0$ for $U=12$ and $\Delta=20$. The spectral weights contributed by lower Hubbard orbitals (LHO) and upper Hubbard orbitals (UHO) in their different filling states are shown. For comparison, the noninteracting density of states is Δ^{-1} for $-\frac{1}{2}\Delta < \omega < \frac{1}{2}\Delta$. The density of states in the central plateau is $\frac{3}{2}\Delta^{-1}$ and is thus enhanced by interactions, relative to the noninteracting case. The central plateau extends over $[U-\frac{1}{2}\Delta, \frac{1}{2}\Delta]$ and is the region where the LHO and UHO coexist.

$U > \Delta$, there is a Mott gap at half filling, and a ZBA forms as one dopes away from half filling. This ZBA is qualitatively different from that found near half filling for $U < \Delta$ and has a width of order t^2/U . In this section, we focus on the large ZBA that appears near half filling for $U < \Delta$.

The approach we take is to calculate the density of states $\rho_{\epsilon_1, \epsilon_2}(\omega)$ for a single two-site AHM with site energies ϵ_1 and ϵ_2 . The density of states is then averaged over all possible configurations,

$$\rho(\omega) = \int_{-\Delta/2}^{\Delta/2} d\epsilon_1 \int_{-\Delta/2}^{\Delta/2} d\epsilon_2 \rho_{\epsilon_1, \epsilon_2}(\omega). \quad (12)$$

To simplify the analytic calculations, it is useful to partition the integration range $[-\frac{1}{2}\Delta, \frac{1}{2}\Delta]$ into subranges $\mathcal{A} = [-\frac{1}{2}\Delta, 0]$ and $\mathcal{B} = [0, \frac{1}{2}\Delta]$, as illustrated in Fig. 6. Sites whose UHO lies near ϵ_F belong to region \mathcal{A} while sites whose LHO lies near

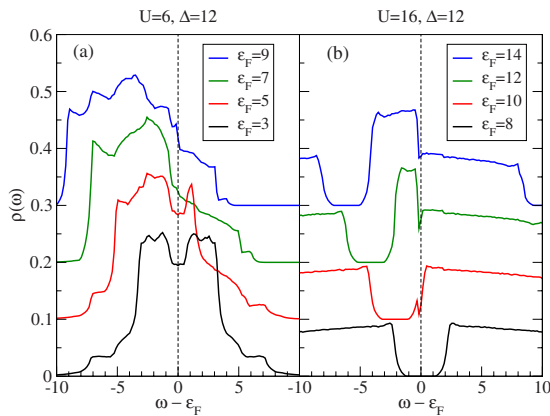


FIG. 5. (Color online) Exact numerical solution for the density of states. Density of states for (a) $U < \Delta$ and (b) $U > \Delta$ for different values of ϵ_F . In both panels, half filling corresponds to $\epsilon_F = U/2$. Curves are offset for clarity. Densities of states are averaged over 10^6 random disorder configurations.

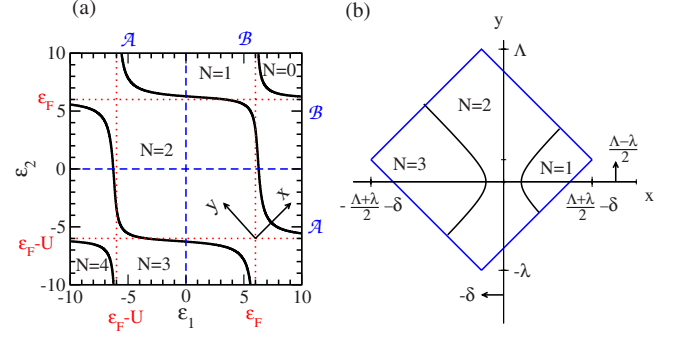


FIG. 6. (Color online) Phase diagram of the two-site AHM. (a) Filling states as a function of ϵ_1 and ϵ_2 for $\Delta=20$, $U=12$, $t=1$, and $\epsilon_F=U/2$, corresponding to half filling on average. Also shown are the regions \mathcal{A} and \mathcal{B} . Solid black lines indicate phase boundaries for t nonzero; dotted red lines at ϵ_F and $\epsilon_F - U$ indicate phase boundaries for $t=0$; dashed blue lines indicate boundaries of regions \mathcal{A} and \mathcal{B} . (b) Integration region $\mathcal{B}\mathcal{A}$ for $\epsilon_F = \frac{1}{2}U + \delta$, shown in terms of transformed coordinates.

ϵ_F belong to region \mathcal{B} . We have argued^{5,11} that the ZBA comes from level repulsion between LHO and UHO on neighboring sites, and it is indeed suggested by Figs. 2(b) and 2(c) that the important configurations have $\epsilon_1 \in \mathcal{A}$, $\epsilon_2 \in \mathcal{B}$ or $\epsilon_1 \in \mathcal{B}$, $\epsilon_2 \in \mathcal{A}$. The simplest approximation is to treat these configurations carefully while treating the other configurations in the atomic limit. As we show, this turns out to be sufficient to understand the essential physics of the ZBA.

We denote by $\rho_{\mathcal{X}\mathcal{Y}}(\omega)$ the density of states ensemble averaged over sites with $\epsilon_1 \in \mathcal{X}$ and $\epsilon_2 \in \mathcal{Y}$,

$$\rho_{\mathcal{X}\mathcal{Y}}(\omega) = \text{Im} \int_{\mathcal{X}} d\epsilon_1 \int_{\mathcal{Y}} d\epsilon_2 \rho_{\epsilon_1, \epsilon_2}(\omega). \quad (13)$$

$\rho_{\mathcal{A}\mathcal{A}}(\omega)$ and $\rho_{\mathcal{B}\mathcal{B}}(\omega)$ are evaluated in the atomic limit, using Eq. (9),

$$\begin{aligned} \rho_{\mathcal{A}\mathcal{A}+\mathcal{B}\mathcal{B}}(\omega) &\equiv \rho_{\mathcal{A}\mathcal{A}}(\omega) + \rho_{\mathcal{B}\mathcal{B}}(\omega) \\ &= \frac{1}{2\Delta} \left[\Theta\left(\frac{1}{2}\Delta - \omega\right) \Theta\left(\omega - U + \frac{1}{2}\Delta\right) \right. \\ &\quad \left. + \frac{1}{2} \Theta(U - |\omega - \epsilon_F|) \right]. \end{aligned} \quad (14)$$

For ω and ϵ_F near $U/2$ (half filling), $\rho_{\mathcal{A}\mathcal{A}+\mathcal{B}\mathcal{B}}(\omega) = \frac{3}{4}\Delta^{-1}$. Using $\rho_{\mathcal{B}\mathcal{A}}(\omega) = \rho_{\mathcal{A}\mathcal{B}}(\omega)$, the total density of states is

$$\rho(\omega) \approx \frac{3}{4\Delta} + 2\rho_{\mathcal{B}\mathcal{A}}(\omega). \quad (15)$$

A more careful derivation of $\rho_{\mathcal{A}\mathcal{A}+\mathcal{B}\mathcal{B}}(\omega)$ finds corrections to the atomic limit approximation on the energy scale $|\omega - \epsilon_F| < O(t^2/U)$.

The next step is to evaluate

$$\rho_{\mathcal{BA}}(\omega) = -\frac{1}{\pi\Delta^2} \int_B d\epsilon_1 \int_A d\epsilon_2 \text{Im} G_{\epsilon_1, \epsilon_2}(\omega)$$

with $G_{\epsilon_1, \epsilon_2}(\omega)$ the retarded Green's function averaged over sites and spins. It will be convenient to change integration variables to

$$x = \frac{\epsilon_2 + U + \epsilon_1}{2} - \epsilon_F, \quad (16a)$$

$$y = \frac{\epsilon_2 + U - \epsilon_1}{2}, \quad (16b)$$

and write

$$\rho_{\mathcal{BA}} = -\frac{2}{\pi\Delta^2} \int_{-\lambda}^{\Lambda} dy \int_{-x_y-\delta}^{x_y-\delta} dx \text{Im} G_{x,y}(\omega), \quad (17)$$

where the factor of 2 is the Jacobian for the transformation, and the integration limits are

$$x_y = \frac{\Lambda + \lambda}{2} - \left| y - \frac{\Lambda - \lambda}{2} \right| \quad (18)$$

and

$$\lambda \equiv \frac{\Delta - U}{2}; \quad \Lambda \equiv \frac{U}{2}. \quad (19)$$

The Fermi energy is written as

$$\epsilon_F = \frac{U}{2} + \delta. \quad (20)$$

This equation defines δ such that half filling corresponds to $\delta=0$. Figure 6 illustrates the integration region and gives the graphical meaning of λ , Λ , and δ .

The phase diagram Fig. 6 shows that there are three filling states in \mathcal{BA} , with $N=1, 2$, or 3 electrons. We now find the ground-state wave functions, energies, and phase boundaries for the different filling states.

2. Ground states in region \mathcal{BA}

We will determine the ground-state wave function in the region \mathcal{BA} using a truncated basis set that discards high-energy states. These high-energy states modify the ground-state wave functions and energies by $O(t^2/U)$, and our approximation is consequently valid for $U \gg t$.

In \mathcal{BA} , the one-electron ground state in the atomic limit is $|0\uparrow\rangle$ because $\epsilon_1 > \epsilon_2$. Making t nonzero mixes in a small amount of $|\uparrow 0\rangle$, proportional to $t^2/(\epsilon_1 - \epsilon_2)$. However, in \mathcal{BA} , $\epsilon_1 \sim \epsilon_2 + U$ so the mixing is of order t^2/U and is neglected in our approximation. The one-electron ground state is thus

$$|G1\rangle \approx |0\uparrow\rangle. \quad (21)$$

Similarly, the three-electron ground state is $|G3\rangle \approx |\uparrow 2\rangle$.

The two-electron ground state is found by diagonalizing the AHM Hamiltonian in the reduced basis $\{|s\rangle, |02\rangle\}$, where

TABLE I. Approximate N -electron ground states $|GN\rangle$ and their energies E_G for the two-site model with $\Delta > U \gg t$ in the region \mathcal{BA} . Variables x and y are defined in Eq. (16), and E_- is defined in Eq. (24).

N	$E_G - \epsilon_F N$	$ GN\rangle$	Ground state when
1	$\epsilon_2 - \epsilon_F$	$ 0\uparrow\rangle$	$x > \sqrt{y^2 + 2t^2}$
2	$E_- - 2\epsilon_F$	$\alpha_{1-} s\rangle + \alpha_{2-} 02\rangle$	$ x < \sqrt{y^2 + 2t^2}$
3	$\epsilon_1 + 2\epsilon_2 + U - 3\epsilon_F$	$ \uparrow 2\rangle$	$x < -\sqrt{y^2 + 2t^2}$

$$|s\rangle = \frac{1}{\sqrt{2}}(|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle) \quad (22)$$

is the singlet state. For $\epsilon_1 \sim \epsilon_2 + U \sim \epsilon_F$, each of these basis states has an energy $\sim 2\epsilon_F - U$. The discarded basis state $|20\rangle$ has an energy $\sim 2\epsilon_F + U$, and the amount of $|20\rangle$ mixed into the ground state by t is therefore $\sim O(t^2/U)$, which we ignore. The Hamiltonian matrix in the reduced basis is

$$H = \begin{bmatrix} \epsilon_1 + \epsilon_2 & -\sqrt{2}t \\ -\sqrt{2}t & 2\epsilon_2 + U \end{bmatrix} \quad (23)$$

which has eigenenergies

$$E_{\pm} = \frac{\epsilon_1 + 3\epsilon_2 + U}{2} \pm \sqrt{\left(\frac{\epsilon_2 + U - \epsilon_1}{2}\right)^2 + 2t^2} \quad (24)$$

and eigenstates

$$|\pm\rangle = \alpha_{1\pm}|s\rangle + \alpha_{2\pm}|02\rangle, \quad (25)$$

$$\alpha_{1\pm}^2 = \frac{\sqrt{y^2 + 2t^2} \mp y}{2\sqrt{y^2 + 2t^2}}; \quad \alpha_{2\pm}^2 = 1 - \alpha_{1\pm}^2, \quad (26)$$

where y is defined in Eq. (16). The two-electron ground state is $|G2\rangle = |-\rangle$. The different ground states, their energies, and the phase boundaries between them are tabulated in Table I. The next step is to calculate the density of states for each filling state.

3. Density of states for the three-electron ground state

First, we calculate the contribution to the density of states from the three-electron ground state. Throughout this work, we keep only terms with poles near ϵ_F , meaning that terms with poles near $\epsilon_1 + U$ or ϵ_2 are discarded. The spin- and site-averaged Green's function is then

$$G_{\epsilon_1, \epsilon_2}^{3e}(\omega) \approx \frac{1}{4} \left\{ \sum_{\pm} \frac{|\langle \pm | c_{1\uparrow} |\uparrow 2\rangle|^2 + |\langle \pm | c_{2\uparrow} |\uparrow 2\rangle|^2}{\omega^+ - (\epsilon_1 + 2\epsilon_2 + U) + E_{\pm}} + \frac{|\langle t | c_{2\uparrow} |\uparrow 2\rangle|^2 + |\langle t \uparrow | c_{2\downarrow} |\uparrow 2\rangle|^2}{\omega^+ - (\epsilon_1 + 2\epsilon_2 + U) + (\epsilon_1 + \epsilon_2)} \right\}, \quad (27)$$

where

$$|t\rangle = \frac{1}{\sqrt{2}}(|\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle) \quad (28)$$

is a triplet state. Using $\alpha_{2\pm}^2 = 1 - \alpha_{1\pm}^2$, we reduce Eq. (27) to

$$G_{\epsilon_1, \epsilon_2}^{3e}(\omega) = \frac{1}{4} \sum_{\pm} \frac{1 - \alpha_{1\pm}^2/2}{\tilde{\omega}^+ - x \pm \sqrt{y^2 + 2t^2}} + \frac{3}{8} \frac{1}{\tilde{\omega}^+ - x - y}, \quad (29)$$

where x and y are defined in Eq. (16) and $\tilde{\omega} = \omega - \epsilon_F$.

The ground state has three electrons for

$$-x_y - \delta < x < -\sqrt{y^2 + 2t^2}, \quad (30)$$

where the upper limit is the phase boundary between two- and three-electron states (cf. Table I), and the lower limit [cf. Eq. (18)] is the boundary of region $\mathcal{B}A$. Then

$$\begin{aligned} \rho_{\mathcal{B}A}^{3e} &= \frac{1}{2\Delta^2} \int_{-(\lambda+\delta)/2}^{(\Lambda+\delta)/2} dy \int_{-x_y-\delta}^{-\sqrt{y^2+2t^2}} dx \\ &\times \left[(1 - \alpha_{1+}^2/2) \delta(\tilde{\omega} - x + \sqrt{y^2 + 2t^2}) \right. \\ &\left. + (1 - \alpha_{1-}^2/2) \delta(\tilde{\omega} - x - \sqrt{y^2 + 2t^2}) + \frac{3}{2} \delta(\tilde{\omega} - x - y) \right]. \end{aligned} \quad (31)$$

The integration over x is straightforward because of the delta functions, which introduce the constraints $\tilde{\omega} < 0$ and

$$|y| < \frac{|\tilde{\omega}|}{2} \text{Re} \sqrt{1 - \frac{8t^2}{\tilde{\omega}^2}}, \quad \text{first term,}$$

$$-\frac{\lambda + \delta - |\tilde{\omega}|}{2} < y < \frac{\Lambda + \delta - |\tilde{\omega}|}{2}, \quad \text{second term,}$$

$$-\left(\frac{|\tilde{\omega}|}{2} - \frac{t^2}{|\tilde{\omega}|}\right) < y < \frac{\Lambda + \delta - |\tilde{\omega}|}{2}, \quad \text{third term.}$$

The result for the first term is valid for $0 > \tilde{\omega} > -\lambda - \delta$, i.e., for $\omega < \epsilon_F$ and in the central plateau. In deriving these results, we have neglected terms of order t^2/λ and t^2/Λ . We now integrate over y using

$$\int dy \left(1 - \frac{\alpha_{1\pm}^2}{2}\right) = \frac{1}{4} (3y \pm \sqrt{y^2 + 2t^2}), \quad (32)$$

to get the three electron contribution to the density of states,

$$\begin{aligned} \rho_{\mathcal{B}A}^{3e}(\tilde{\omega}) &= \frac{3\Theta(-\tilde{\omega})}{8\Delta^2} \left[\frac{2\lambda + \Lambda}{3} + \delta - |\tilde{\omega}| \left(1 - \text{Re} \sqrt{1 - \frac{8t^2}{\tilde{\omega}^2}}\right) \right. \\ &\left. + \Theta\left(-\tilde{\omega} - \frac{2t^2}{\Lambda + \delta}\right) \left(\Lambda + \delta - \frac{2t^2}{|\tilde{\omega}|}\right) \right]. \end{aligned} \quad (33)$$

To simplify the final expression, we have taken $\sqrt{\lambda^2 + 2t^2} \approx \lambda$ and $\sqrt{\Lambda^2 + 2t^2} \approx \Lambda$.

4. Density of states for the one-electron ground state

The derivation of the one-electron contribution to the density of states parallels that of the three-electron contribution. The Green's function is

$$\begin{aligned} G_{\epsilon_1, \epsilon_2}^{1e}(\omega) &\approx \frac{1}{4} \left\{ \sum_{\pm} \frac{|\langle \pm | c_{1\downarrow}^\dagger | 0 \uparrow \rangle|^2 + |\langle \pm | c_{2\downarrow}^\dagger | 0 \uparrow \rangle|^2}{\omega^+ + \epsilon_2 - E_{\pm}} \right. \\ &\left. + \frac{|\langle t | c_{1\downarrow}^\dagger | 0 \uparrow \rangle|^2 + |\langle t \uparrow | c_{1\uparrow}^\dagger | 0 \uparrow \rangle|^2}{\omega^+ + \epsilon_2 - (\epsilon_1 + \epsilon_2)} \right\} \\ &= \frac{1}{4} \sum_{\pm} \frac{1 - \alpha_{1\pm}^2/2}{\tilde{\omega}^+ - x \mp \sqrt{y^2 + 2t^2}} + \frac{3}{8} \frac{1}{\tilde{\omega}^+ - x + y}. \end{aligned} \quad (34)$$

The integration region is $x_y - \delta > x > \sqrt{y^2 + 2t^2}$ with x_y given by Eq. (18) and

$$\begin{aligned} \rho_{\mathcal{B}A}^{1e} &= \frac{1}{2\Delta^2} \int_{-(\lambda-\delta)/2}^{(\Lambda-\delta)/2} dy \int_{\sqrt{y^2+2t^2}}^{x_y-\delta} dx \\ &\times \left[(1 - \alpha_{1+}^2/2) \delta(\tilde{\omega} - x - \sqrt{y^2 + 2t^2}) \right. \\ &\left. + (1 - \alpha_{1-}^2/2) \delta(\tilde{\omega} - x + \sqrt{y^2 + 2t^2}) + \frac{3}{2} \delta(\tilde{\omega} - x + y) \right]. \end{aligned}$$

Letting $x \rightarrow -x$, this is the same as $\rho_{\mathcal{B}A}^{3e}(-\tilde{\omega})$ for $\delta \rightarrow -\delta$. Thus

$$\begin{aligned} \rho_{\mathcal{B}A}^{1e}(\tilde{\omega}) &= \frac{3\Theta(\tilde{\omega})}{8\Delta^2} \left[\frac{2\lambda + \Lambda}{3} - \delta - \tilde{\omega} \left(1 - \text{Re} \sqrt{1 - \frac{8t^2}{\tilde{\omega}^2}}\right) \right. \\ &\left. + \Theta\left(\tilde{\omega} - \frac{2t^2}{\Lambda - \delta}\right) \left(\Lambda - \delta - \frac{2t^2}{\tilde{\omega}}\right) \right]. \end{aligned} \quad (35)$$

5. Density of states for the two-electron ground state

Finally, the Green's function for the two-electron ground state is

$$\begin{aligned} G_{\epsilon_1, \epsilon_2}^{2e} &\approx \frac{1}{2} \left\{ \frac{|\langle 0 \downarrow | c_{1\uparrow} | G2 \rangle|^2 + |\langle 0 \downarrow | c_{2\uparrow} | G2 \rangle|^2}{\omega^+ - E_- + \epsilon_2} \right. \\ &\left. + \frac{|\langle \uparrow 2 | c_{1\uparrow}^\dagger | G2 \rangle|^2 + |\langle \uparrow 2 | c_{2\uparrow}^\dagger | G2 \rangle|^2}{\omega^+ + E_- - (\epsilon_1 + 2\epsilon_2 + U)} \right\} \end{aligned} \quad (36)$$

$$= \frac{1}{2} \sum_{\pm} \frac{1 - \alpha_{1\pm}^2/2}{\tilde{\omega}^+ - x \pm \sqrt{y^2 + 2t^2}} \quad (37)$$

Then,

$$\begin{aligned} \rho_{\mathcal{B}A}^{2e} &= \frac{1}{\Delta^2} \int_{-\lambda+|\delta|}^{\Lambda-|\delta|} dy \int_{x_1}^{x_2} dx \left(1 - \frac{\alpha_{1-}^2}{2}\right) \\ &\times \sum_{\pm} \delta(\tilde{\omega} - x \pm \sqrt{y^2 + 2t^2}), \end{aligned} \quad (38)$$

where $x_1 = -\min(x_y + \delta, \sqrt{y^2 + 2t^2})$, $x_2 = \min(x_y - \delta, \sqrt{y^2 + 2t^2})$. Performing the integrations over x and y gives

$$\begin{aligned} \rho_{\mathcal{B}A}^{2e}(\tilde{\omega}) &= \frac{3}{4\Delta^2} \left[\frac{2\lambda + \Lambda}{3} + |\tilde{\omega}| \left(1 - \text{Re} \sqrt{1 - \frac{8t^2}{\tilde{\omega}^2}}\right) \right] \\ &+ \frac{3\delta}{4\Delta^2} [\Theta(\tilde{\omega}) - \Theta(-\tilde{\omega})]. \end{aligned} \quad (39)$$

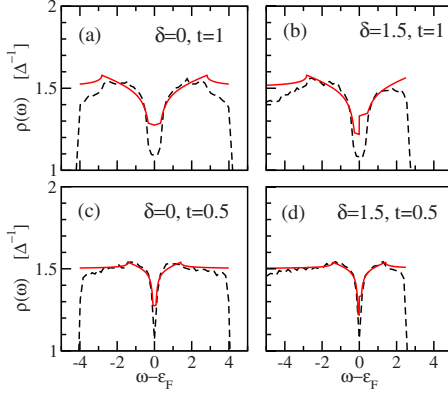


FIG. 7. (Color online) Density of states for the two-site ensemble-averaged Anderson-Hubbard model for different values of t and δ . Results are shown for exact numerical solution of the AHM (dashed black curves) and for the approximate result, Eq. (40) (solid red curves). Model parameters are $U=12$ and $\Delta=20$. Recall that $\varepsilon_F=U/2+\delta$ and that $\delta=0$ corresponds to half filling.

6. Total density of states

Putting the results of the different calculations together, we arrive at our final result for the total density of states (valid in the central plateau),

$$\begin{aligned} \rho(\omega) &= \rho_{AA+BB}(\omega) + 2\rho_{BA}^{1e}(\omega) + 2\rho_{BA}^{2e}(\omega) + 2\rho_{BA}^{3e}(\omega) \\ &= \frac{3}{2\Delta} + \frac{3}{2\Delta^2} \left[-\frac{U}{4} + \frac{\delta_s}{2} + \frac{|\tilde{\omega}|}{2} \left(1 - \text{Re} \sqrt{1 - \frac{8t^2}{\tilde{\omega}^2}} \right) \right. \\ &\quad \left. + \Theta \left(|\tilde{\omega}| - \frac{4t^2}{U-2\delta_s} \right) \left(\frac{U-2\delta_s}{4} - \frac{t^2}{|\tilde{\omega}|} \right) \right], \end{aligned} \quad (40)$$

where $\delta_s = \delta \text{sgn}(\tilde{\omega})$ and $\tilde{\omega} = \omega - \varepsilon_F$. The parameters λ , Λ , and δ are defined in Eqs. (19) and (20).

As a check of Eq. (40), we let $t \rightarrow 0$, in which case

$$\rho(\omega) \rightarrow \frac{3}{2\Delta}$$

which agrees with the previous atomic limit calculation for the central plateau. Equation (40) is plotted in Fig. 7 in comparison with the exact density of states determined from numerically diagonalizing the AHM.

IV. DISCUSSION

Figure 7 compares Eq. (40) to exact disorder-averaged numerical calculations for the density of states. The theory works well for $|\tilde{\omega}| > 4t^2/U$, out to the edges of the central plateau where it breaks down [for example, near $\tilde{\omega} = \pm 4$ in Fig. 7(a)]. The theory neglects terms of order t^2/U , and therefore fits the numerics better when t is smaller, as shown in Figs. 7(c) and 7(d). The fit for $|\tilde{\omega}| < 4t^2/U$ is not especially good but can be improved significantly by considering corrections of order t^2/U that were neglected in the previous section; we have not included these corrections because they complicate $\rho(E)$ significantly without adding physical insight. The focus of this discussion is therefore $|\tilde{\omega}| > 4t^2/U$.

The main qualitative idea that we emphasize in this section is that there are two distinct physical mechanisms that lead to the ZBA in Eq. (40). Both mechanisms occur for configurations where the LHO of one site and the UHO of the other site are nearly degenerate with ε_F , namely, for $\varepsilon_1 \sim \varepsilon_2 + U \sim \varepsilon_F$ or $\varepsilon_2 \sim \varepsilon_1 + U \sim \varepsilon_F$. The first mechanism is similar to that outlined in the mean-field calculation in Sec. II: level repulsion, caused by hybridization of many-body states, shifts states away from ε_F . The second mechanism does not have a mean-field counterpart: level repulsion gaps the spectrum of low-energy triplet excitations.

The first mechanism underlies the second last term in Eq. (40),

$$\frac{3|\tilde{\omega}|}{4\Delta^2} \left(1 - \text{Re} \sqrt{1 - \frac{8t^2}{\tilde{\omega}^2}} \right). \quad (41)$$

This term rises linearly from $\tilde{\omega}=0$ and is peaked at $\tilde{\omega} = \pm 2\sqrt{2}t$, which defines the width of the ZBA in Fig. 7. In our calculations, this term comes from transitions between two-electron singlet states and states with one or three electrons, and it is the level repulsion between the two-electron states that causes the ZBA. In the case, for example, where ε_1 and $\varepsilon_2 + U$ lie near ε_F , there are two nearly-degenerate two-electron singlets, $|s\rangle$ and $|02\rangle$; these hybridize as a result of the matrix element t to form bonding and antibonding many-body states with energies [from Eq. (24)],

$$E_{\pm} \approx 2\varepsilon_F - U \pm \sqrt{2}t. \quad (42)$$

Thus, the level repulsion between $|s\rangle$ and $|02\rangle$ shifts the many-body orbital energies up or down by $O(t)$. Starting from the two-electron ground state, with energy E_- , one has transitions

$$\begin{aligned} \alpha_{1-}|s\rangle + \alpha_{2-}|02\rangle &\xrightarrow{c_{1\sigma}^\dagger c_{2\sigma}^\dagger} |\sigma 2\rangle, \\ \alpha_{1-}|s\rangle + \alpha_{2-}|02\rangle &\xrightarrow{c_{1\sigma} c_{2\sigma}} |0\bar{\sigma}\rangle. \end{aligned} \quad (43)$$

We showed in Sec. III B that the three-electron energy is $\varepsilon_1 + 2\varepsilon_2 + U$ (which is approximately $3\varepsilon_F - U$), and the one-electron energy is ε_2 (approximately $\varepsilon_F - U$) so that the transition energies in Eq. (43) are

$$\omega_{\pm} \approx \varepsilon_F \pm \sqrt{2}t. \quad (44)$$

Because ω_{\pm} are shifted by $O(t)$ away from ε_F , the density of states at ε_F is reduced as t increases. As indicated above, this mechanism for depleting the low-energy density of states is similar to the mean-field mechanism discussed in Sec. II, where level repulsion between molecular states on opposite sides of ε_F increases the energy required to add or remove an electron. In this sense, the second-last term in Eq. (40) is Altshuler-Aronov-like.

The second mechanism does not have a mean-field counterpart, and results in the last term in Eq. (40),

$$\frac{3}{2\Delta^2}\Theta\left(|\tilde{\omega}| - \frac{4t^2}{U-2\delta_s}\right)\left(\frac{U-2\delta_s}{4} - \frac{t^2}{|\tilde{\omega}|}\right). \quad (45)$$

This term varies as $|\tilde{\omega}|^{-1}$ down to the low-energy cutoff at $|\tilde{\omega}| \sim 4t^2/U$, and makes the dominant contribution to the shape of the ZBA. The cutoff comes from the boundary between the region \mathcal{BA} and the region \mathcal{BB} in Fig. 6, where the approximate one- and two-electron wave functions used in deriving $\rho(E)$ cease to be valid.

In our calculations, Eq. (45) comes from transitions between one- or three-electron ground states, and two-electron triplet excitations. For the three-electron ground state, for example, these excitations have the form

$$|\uparrow 2\rangle \xrightarrow{c_{2\downarrow}} |\uparrow \uparrow\rangle, \quad |\uparrow 2\rangle \xrightarrow{c_{2\uparrow}} |t\rangle. \quad (46)$$

As mentioned above, the three-electron energies are nearly independent of t ; the triplet energies are also independent of t , however, so that the transition energies are not shifted by level repulsion. The mechanism for depleting the low-energy density of states in this term is therefore *not* that of Altshuler and Aronov.

Instead, it is the fact that a gap in the triplet spectrum opens as t increases that causes a depletion of states near ε_F (this gap was illustrated in Fig. 2). This gap occurs for configurations of $(\varepsilon_1, \varepsilon_2)$ that have one- or three-electron ground states in the atomic limit but two-electron ground states when t is nonzero. For example, when ε_1 and $\varepsilon_2 + U$ both lie slightly below ε_F , the atomic-limit ground state has three electrons and triplet excitations as in Eq. (46) are possible. When t is nonzero, the two-electron ground-state energy E_- is reduced by $O(t)$ while the three electron ground-state energy is reduced by $O(t^2/U)$. For sufficiently large t , the two-electron ground state has the lower energy and the triplet excitation is eliminated (i.e., the only possible tunneling processes are to one- or three-electron final states). In summary, the ZBA in the final term of Eq. (40) occurs because the phase space for low-energy triplet excitations is reduced when t increases.

As we discuss elsewhere,^{5,11} this calculation sheds light on the empirical observation, made for larger systems, that the width of the ZBA is of order t .² A naive argument based on the disorder-free Hubbard model would suggest that the ZBA might have a conventional Altshuler-Aronov form, but with an effective exchange interaction $V_{\text{eff}} = 4t^2/U$, so that the ZBA should grow with increasing t^2/U . As we have said

above, there are contributions to the density of states of this type; however, we have just shown that a much larger effect, of order t , comes from configurations with the LHO and UHO on neighboring sites degenerate.

We note that this explanation appears to contradict numerical evidence from the work of Chiesa *et al.*² on two-dimensional clusters that a large ZBA persists far from half filling and for large U since configurations with a degenerate LHO and UHO do not occur in these cases; as we have shown in Fig. 5, the ZBA vanishes rapidly (with increasing disorder) in the two-site model when ε_F is outside the central plateau. To check this, we have performed preliminary exact-diagonalization calculations for larger clusters (up to 12 sites). These calculations find that the width of the ZBA is not linear in t when ε_F is outside the central plateau, and suggest that the physics of the ZBA changes far from half filling. A more detailed study of how the ZBA evolves with doping needs to be undertaken.

V. CONCLUSIONS

In summary, we have found that the zero bias anomaly in the two-site Anderson-Hubbard model is the result of strong orbital hybridization in the two-electron ground state for configurations with $\varepsilon_1 \sim \varepsilon_2 + U \sim \varepsilon_F$ or with $\varepsilon_2 \sim \varepsilon_1 + U \sim \varepsilon_F$. Unlike in the conventional Hubbard model, this hybridization is not suppressed by the on-site interaction U , and leads to a level repulsion between molecular orbital energies of order t , rather than t^2/U .

The mechanism for the suppression of the tunneling density of states is, at least in part, different from in conventional mean-field models of interacting electrons. In mean-field theories, interactions cause a shift of molecular-orbital energies away from ε_F that leads directly to an increase in the energy required to remove or add an electron. This also occurs in the Anderson-Hubbard model; however, there is an additional depletion of low-energy spectral weight because the low-energy triplet excitation spectrum is gapped as a result of orbital hybridization. This mechanism is physically different from that of Altshuler and Aronov.

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