# **Magnetic mechanism of quasiparticle pairing in hole-doped cuprate superconductors**

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We have computed  $\alpha^2 F$ 's for the hole-doped cuprates within the framework of the one-band Hubbard model, where the full magnetic response of the system is treated properly. The *d*-wave pairing weight  $\alpha^2 F_d$  is found not only to contain a low-energy peak due to excitations near  $(\pi, \pi)$  expected from neutron-scattering data but also to display substantial spectral weight at higher energies due to contributions from other parts of the Brillouin zone as well as pair-breaking ferromagnetic excitations at low energies. The resulting solutions of the Eliashberg equations yield transition temperatures and gaps comparable to the experimentally observed values, suggesting that magnetic excitations of both high and low energies play an important role in providing the pairing glue in the cuprates.

DOI: [10.1103/PhysRevB.78.134513](http://dx.doi.org/10.1103/PhysRevB.78.134513)

PACS number(s): 74.20. - z, 71.10.Fd, 74.72. - h

## **I. INTRODUCTION**

Since the superconducting state in the cuprates evolves from the doping of a Mott insulator, it is natural to conjecture that the pairing is driven by magnetic fluctuations rather than by phonons. Quantum Monte Carlo (QMC) calculations provide evidence for  $d$ -wave pairing,<sup>1,[2](#page-4-1)</sup> where the pairing bosons reside predominantly in the (transverse) spin channel. Recent debate in this connection has centered on whether or not the magnetic-resonance peak is strong enough to account for the condensation energy. $3,4$  $3,4$  Although recent estimates seem to be affirmative,<sup>5</sup> they do not take into account competing pair-breaking effects which enter the Eliashberg equations.<sup>6</sup> Also, there are arguments that high-energy exci-tations play a role.<sup>7–[9](#page-4-7)</sup> Here we report a computation of  $\alpha^2 F$ 's for the hole-doped cuprates based on the one-band Hubbard model, where the full magnetic response of the system is included, and the Eliashberg equations are then solved selfconsistently to obtain the superconducting properties over a wide range of dopings and temperatures. The resulting transition temperatures and pairing gaps are found to be comparable to experimental values, showing clearly the viability of the magnetic mechanism in the cuprates. We find that excitations at both high and low energies are important.

Early calculations of magnetic pairing in the cuprates employed parametrized models of the susceptibility. The analy-sis of Radtke et al.<sup>[10](#page-4-8)</sup> [Radtke, Ullah, Levin, and Norman (RULN)] invoked neutron-scattering measurements, while that of Millis *et al.*<sup>[11](#page-4-9)</sup> [Millis, Monien, and Pines (MMP)] was based on NMR data. The model  $\alpha^2 F$ 's obtained lead to divergent predictions concerning the feasibility of magnetic mechanism.<sup>12</sup> Our *d*-wave pairing weight  $\alpha^2 F_d$  contains not only a low-energy peak (LEP) from near- $(\pi, \pi)$  scattering but also an additional high-energy feature (HEF) extending to  $\sim$  1.5 eV dominated by other regions of the Brillouin zone (BZ) as well as a significant pair-breaking contribution at low energies from ferromagnetic (FM) fluctuations. The HEF, which was missing in the RULN and MMP models, turns out to be crucially important in producing high transition temperatures and pairing gaps. The pair-breaking terms begin to dominate as the Fermi energy approaches the van Hove singularity (VHS) with increasing doping and can lead to the loss of superconductivity.

Our study bears on the recently discovered "waterfall" or high-energy kink (HEK) features observed over 0.3-1.2 eV range in the angle-resolved photoemission spectra (ARPES) of a number of cuprates. The magnetic susceptibility underlying our computation of  $\alpha^2 F$ 's yields self-energies and dispersions consistent with the waterfall effects, $13,14$  $13,14$  suggesting that the boson responsible for the waterfall effects is also a key player in generating significant pairing weight in  $\alpha^2 F_d$ and high condensation energy in the cuprates.

The calculations are based on a one-band Hubbard Hamiltonian extended to include pairing interaction. Specifically, in terms of susceptibility  $\chi_0$  and the Hubbard on-site repulsion  $U$ , we use the singlet pairing potential<sup>15</sup>

$$
V_s = \frac{U}{1 - U^2 \chi_0^2 (p' - p)} + \frac{U^2 \chi_0 (p' + p)}{1 - U \chi_0 (p' + p)}
$$
(1)

<span id="page-0-1"></span>and the mass-renormalization potential [Eq. 8 of Ref.  $15(a)$ ]

$$
V_z = \frac{U^2 \chi_0(p'-p)}{1 - U^2 \chi_0^2(p'-p)} + \frac{U^3 \chi_0^2(p'-p)}{1 - U \chi_0(p'-p)},
$$
 (2)

<span id="page-0-2"></span>where  $p$  and  $p'$  are the electron momenta, which are constrained to lie on the Fermi surface. Here  $V_z$   $[V_s]$  is the potential contributing to the normal [anomalous] part of the self-energy. These expressions have been found to give transition temperatures in good agreement with QMC results.<sup>2</sup> The resulting coupling constants in various pairing channels  $\alpha$  are

$$
\overline{\lambda}_{\alpha} = -\int \int d^2p d^2p' \overline{g}_{\alpha}(p) \overline{g}_{\alpha}(p') V(p, p', \omega = 0), \qquad (3)
$$

<span id="page-0-0"></span>where  $V=V<sub>s</sub>$  for the even-parity channels. The normalized weighting function  $\tilde{g}_g = g_a(p)/(N_0|v_p|)$ , where  $v_p$  is the Fermi velocity and  $N_0^2 = (2\pi)^3 \int g_\alpha(p)^2 d^2p / |v_p|$ . The  $g_\alpha$  are weighting functions of various symmetry, $15$  of which the most important are the lowest harmonics of *s* wave and  $d_{x^2-y^2}$ symmetry with  $g_s = 1$  and  $g_d = \cos(p_x a) - \cos(p_y a)$ . We also define the coupling constant  $\lambda_z$  via the *s*-wave version of Eq. ([3](#page-0-0)) with  $V=V_z$ . Then the effective BCS coupling is  $\lambda_{\alpha}$  $= \overline{\lambda}_{\alpha}/(1+\lambda_z)$ . The symmetrized Eliashberg functions then are

$$
\alpha^2 F_\alpha(\omega) = -\frac{1}{\pi} \int \int d^2p \, d^2p' \, \widetilde{g}_\alpha(p) \widetilde{g}_\alpha(p') V''(p, p', \omega), \quad (4)
$$

<span id="page-1-0"></span>where  $V''$  is the imaginary part of the corresponding  $V$ .

In the presence of strong magnetic fluctuations, the pseudogap is a manifestation that Migdal's theorem is not obeyed[.16,](#page-4-14)[17](#page-4-15) We have developed a relatively simple approximation scheme $13,18$  $13,18$  which can successfully reproduce the pseudogap and waterfall phenomena in the normal state of the cuprates over the full doping range. In the overdoped regime, this scheme reduces to calculating the self-energy in GW approximation using a reduced *U*=3.2*t* and dispersion renormalized via  $Z_0 = 2$ .<sup>19</sup> These values of *U* and  $Z_0$  yield self-energies in reasonable accord with the QMC results<sup>20</sup> and explain the recently observed waterfall effects in photoemission spectra of the cuprates. $13,14$  $13,14$  We therefore expect these parameters to be most appropriate near  $x=0.27$ , but to gain some understanding of how the band structure would affect superconductivity in the absence of pseudogap effects, we solved the Eliashberg equations over the full doping range  $x=0-0.4$  assuming  $Z_0$  and U to be doping independent. A more satisfactory procedure would be to let *U* increase in the underdoped regime. However, in the presence of a pseudogap, a tensor system of Eliashberg equations needs to be solved, and that is beyond the scope of the present calculation. In short, we proceed thus by solving Eqs. ([1](#page-0-1)) and ([2](#page-0-2)) using  $U=3.2t$  and  $\chi_0$  renormalized by  $Z_0$ . In particular, we neglect the additional modifications of Migdal's theorem in the superconducting state. Despite this limitation, our results provide a benchmark for the Eliashberg formulation in that we do not invoke empirical susceptibilities as has been the case in much of the existing literature.

Concerning technical details, we use a tight-binding parametrization of the dispersion of Bi2212 with the bilayer splitting neglected.<sup>21</sup>  $\chi_0$  is first computed within the randomphase approximation (RPA) scheme throughout the BZ for frequencies up to 2.88 eV.  $\alpha^2 F$ 's and the  $\lambda$ 's are then com-puted from Eqs. ([1](#page-0-1))–([4](#page-1-0)). Fermi surface restricted Eliashberg equations<sup>10</sup> are finally used to self-consistently obtain the gap  $\Delta(\omega)$  and renormalization  $Z(\omega)$  functions, with  $Z(0)$  $\equiv$  Z=1+ $\lambda_z$ <sup>[22](#page-4-20)</sup>

#### **II. PURE** *d***-WAVE SOLUTION**

Figure [1,](#page-1-1) which compares our typical *d*-wave pairing weights  $\alpha^2 F_d$  and  $\alpha^2 F_z$  with RULN and MMP models, highlights our key finding. Our  $\alpha^2 F_d$  (green line) in Fig. [1](#page-1-1)(a) displays two clear features, $^{23}$  a LEP around 40 meV and a broad humplike HEF extending from  $\sim 0.5$  to 1.0 eV [see also Fig.  $2(a)$  $2(a)$  below]. The LEP arises mainly from the magnetic response near  $(\pi, \pi)$ , but the HEF is connected with the response from other parts of the BZ particularly near  $(\pi,0)$ and  $(\pi/2, \pi/2)$ . Our LEP in Fig. [1](#page-1-1)(a) is similar to the weights assumed by RULN and MMP. This resemblance is not surprising since the RULN model<sup>10</sup> was designed to match neutron-scattering data near  $(\pi, \pi)$ , while the NMR data utilized by MMP (Ref.  $11$ ) are also most sensitive to weight in this part of the BZ. It has long been known that

<span id="page-1-1"></span>

FIG. 1. (Color online) Eliashberg functions  $\alpha^2 F_d$  and  $\alpha^2 F_z$  for hole doping  $x=0.30$  obtained in this work (green line) are compared with results of Refs.  $10$  (red-dotted line) and  $11$  (blue-dashed line).

neutron scattering near  $(\pi, \pi)$  accounts for only about 1/8th of the integrated spectral weight expected from a total scattering sum rule. $^{24}$  By basing their estimate solely on the neutron-scattering data near  $(\pi, \pi)$ , RULN severely underestimated the total *d*-wave pairing weight. The MMP analysis, based on NMR data, appears to have captured more of the weight—although still missing the HEF and thus underestimating the total weight. Note also from Fig.  $1(b)$  $1(b)$  that both models strongly underestimate the renormalization weight  $\alpha^2 F_z$ , which opposes the tendency for pairing.

The negative dip in Fig.  $1(a)$  $1(a)$  at energies below 20 meV deserves comment. This dip reflects pair-breaking magnetic scattering (PBS) near  $\Gamma$  and was overlooked in the phenom-

<span id="page-1-2"></span>

FIG. 2. (Color online) Eliashberg functions  $\alpha^2 F_d$  and  $\alpha^2 F_z$  over the doping range  $x=0.0-0.4$ . Lines of various colors refer to different dopings [see legend in (a)]. Left hand panels (a) and (b) give results over an extended frequency range of 0–1.5 eV, while righthand panels (c) and (d) highlight the low-energy region of 0–250 meV on an expanded energy scale.

<span id="page-2-0"></span>

FIG. 3. (Color online) Doping dependence of: (a)  $\lambda_z$ ; (b)  $\lambda_d$ ; and (c)  $\Delta_d(T=0)$  (left scale) compared to  $\lambda_d$  (right scale). In (a) three different computations of  $\lambda$ <sub>z</sub> are compared based on the full  $V$ <sub>z</sub> of Eq. ([2](#page-0-2)) (blue-solid line), a simplified  $V_{z0} = U^2 \chi_0$  (red-dotted line), and the estimate  $N(0)U$  (green-dashed line), where  $N(0)$  is the density of states at the Fermi energy.

enological RULN and MMP models. For simplicity, we will refer to these fluctuations as being FM, although this is strictly so only at  $\Gamma$ . This PBS is related to earlier indications of FM instabilities near a VHS[.25](#page-4-23)[,26](#page-4-24) A similar scenario of competing *d*-wave pairing vs pair-breaking effects has been discussed in the context of electron-phonon pairing.<sup>27</sup>

Figure [2](#page-1-2) shows how  $\alpha^2 F$ 's evolve with doping. In (a), the pairing weight in the high-energy feature of  $\alpha^2 F_d$  is seen to increase monotonically with increasing doping, displaying an approximate isosbestic point at  $\omega \sim 0.24$  eV. In the lowenergy region in (c), position of the peak in  $\alpha^2 F_d$  shifts to lower energies with increasing doping, and the negative pairbreaking peak grows dramatically consistent with the suggestion of Kopp *et al.*<sup>[26](#page-4-24)</sup> The nature of  $\alpha^2 F_d$  is seen to change quite substantially as the Fermi energy approaches the VHS at around  $x=0.39$ . Interestingly, by comparing (c) and (d), the low-energy peak in  $\alpha^2 F_z$  is seen to follow that in  $\alpha^2 F_d$  to lower energies with doping.

Figure [3](#page-2-0) shows the doping dependence of  $\lambda_z$ ,  $\lambda_d$ , and the low-temperature gap  $\Delta_d(T=0)$ . Three different estimates of  $\lambda_z$  are compared in (a) for illustrative purposes. Values based on using the bare susceptibility,  $V_{z0} = U^2 \chi_0$  (red-dashed line), are seen to be quite similar to the simple estimate  $N(0)U$ (green-dotted line), where  $N(0)$  is the density of states at the Fermi energy. The full  $V<sub>z</sub>$  (blue line) on the other hand yields a significant enhancement of  $\lambda_z$  over that obtained from  $\chi_0$ , especially near the region of the VHS peak, indicating that the system is close to a magnetic instability. Note that  $\lambda_d$  is positive for dopings less than  $\approx 0.4$ , but as the Fermi energy enters the region of the VHS with increasing doping,  $\lambda_d$  rapidly becomes large and negative due to FM fluctuations. Fig-ure [3](#page-2-0)(c) shows that this doping dependence of  $\lambda_d$  is well correlated with that of the pairing gap. We stress that these results hold for a *pure dx*2−*y*<sup>2</sup> order parameter. Harmonic content plays an important role as will be discussed below (Sec. III).

<span id="page-2-1"></span>

FIG. 4. (Color online) (a) Real part of the gap function  $\Delta(\omega, T)$ at doping *x*=0.10 as a function of frequency for a series of temperatures *T* (see legend). Thin black line is the plot of  $\Delta = \omega$  used to obtain the low-energy gap as discussed in text. (b) Computed temperature dependence of the low-energy gap  $\Delta_d(T)$  at various dopings  $x$  (see legend).

We turn now to discuss our solutions of the Eliashberg equations. Following common practice, we proceeded by discretizing the  $\alpha^2 F$ 's on the real frequency axis.<sup>28</sup> We find that our results are sensitive to the number  $N_m$  of points in the mesh. For the present calculations, based on a 768-point nonuniform mesh over 0–2.88 eV, the gap  $\Delta(\omega)$  is approximately converged in the low- $\omega$  regime allowing us to extract  $\Delta_d(T)$ . Figure [4](#page-2-1)(a) shows typical results for the real part of  $\Delta(\omega)$  for a range of temperatures at *x*=0.10. The prominent oscillations in  $\Delta(\omega)$  curves are the well-known consequence of discretizing  $\alpha^2 F$ 's in solving the Eliashberg equations.<sup>29</sup> We define the gap by taking the intersection of the  $\Delta(\omega)$  $= \omega$  line [thin-black line in Fig. [4](#page-2-1)(a)] with the  $\Delta(\omega)$  curve.

Figure [4](#page-2-1)(b) shows how the computed low-energy gap  $\Delta_d$ evolves with temperature at various dopings. Due to the difficulty of finding well-converged solutions when  $\Delta$  is small, we calculate  $\Delta_d(T)$  at a few low temperatures and use a fit to a *d*-wave BCS gap to estimate  $T_c$ . We find  $2\Delta_d(0)/k_B T_c$  $\sim$  3.2 for different dopings. The resulting  $T_c$ 's are somewhat smaller than QMC values, $<sup>2</sup>$  which is perhaps the effect of a</sup> finite *t*.

It is striking that the gap features in Fig.  $4(a)$  $4(a)$  extend to very high energies raising the obvious question as to how this high-energy tail would show up in tunneling spectra.<sup>30</sup> Insight in this regard is provided by Fig. [5,](#page-3-0) where we show a typical tunneling spectrum computed $31$  within our model. Tunneling spectra computed at other dopings are similar except that the features scale with  $\Delta_d$ .) The weight in Fig. [5](#page-3-0)

<span id="page-3-0"></span>

FIG. 5. (Color online) Typical computed superconductorinsulator-superconductor tunneling spectrum at *x*=0.10. Inset shows the high-energy tail on an expanded scale.

at energies above the peak-dip-hump feature is seen to be quite small with weak energy dependence (see inset) and would not be readily observable in the presence of an experimental background.

#### **III. LOW- VS HIGH-ENERGY PAIRING GLUE**

Within the present model, the LEP and HEF *both* play an essential role in generating large gaps. For example, at  $x=0.3$ , the HEF by itself produces a gap of only  $\sim 0.4$  meV while the LEP is virtually nonsuperconducting, even though the full  $\alpha^2 F_d$  yields a gap of 5.5 meV. (To be definite, we separate  $\alpha^2 F$  into LEP and HEF at the minimum in  $\alpha^2 F$  with  $\omega_{\text{min}}$ =0.3 eV.) Similarly, for *x*=0.1, LEP (HEF) by itself has  $a \sim 3$  (0.4) meV gap with a combined gap of  $\sim 17$  meV with  $\omega_{\text{min}}$ =0.16 eV.

This behavior can be readily understood from a two- $\lambda$ model.<sup>32</sup> Since this is a purely electronic mechanism, we use a modified Allen-Dynes formula, 33,[34](#page-4-32)

$$
T_c = \frac{\omega_{\text{ln}}}{1.2} \exp\left(\frac{-1.04(1+\lambda_z)}{\overline{\lambda}_d}\right) = \frac{\omega_{\text{ln}}}{1.2} \exp\left(\frac{-1.04}{\lambda_d}\right), \quad (5)
$$

 $\Delta(0)$ =3.54 $T_c$  with

$$
\bar{\lambda}_d = 2 \int_0^\infty \frac{\alpha^2 F(\omega)}{\omega} \tag{6}
$$

and

$$
\ln(\omega_{\text{ln}}) = \frac{2}{\overline{\lambda}_d} \int_0^\infty \ln(\omega) \frac{\alpha^2 F(\omega)}{\omega}.
$$
 (7)

The Allen-Dynes equation has a well-known limitation<sup>34</sup> that it predicts a maximum  $T_c = \omega_{\text{ln}} / 1.2$ , whereas the Eliashberg equations have a solution that grows without limit  $\sim \sqrt{\overline{\lambda}}$  as  $\overline{\lambda} \rightarrow \infty$ . We find that this leads to an underestimate of  $\Delta_{LEP}$ , while the model provides good estimates for the remaining gaps. For instance, for  $x=0.1$ ,  $\lambda_{\text{LEP}}=\lambda_{\text{HEF}}=0.15$ ,  $\omega_{\text{In,LEP}}$ =83 meV, and  $\omega_{\text{In,HEF}}$ =530 meV so  $\Delta_{\text{LEF}}$ =0.26 meV and  $\Delta_{\text{HEF}}=1.4$  meV. When both features are combined,  $\omega_{\text{ln}}$ =200 meV and  $\lambda_d$ =0.3 leading to  $\Delta_d$ =19 meV, which is in good agreement with the full calculation. While the Allen-

<span id="page-3-1"></span>

FIG. 6. (Color online) Doping dependence of:  $\lambda_d$  (blue line),  $\lambda_s$ (red line),  $\lambda_{dx}$  (green line),  $\lambda_p$  (violet line), and  $\lambda_{sp}$  (brown line) calculated from a  $15 \times 15$  harmonic matrix in each symmetry sector.

Dynes model is highly simplified, it does capture the observed trend that both peaks contribute significantly. Physically, the effective  $\lambda$  is in the weak-coupling regime,  $\lambda \sim \ll 1$ , so high  $T_c$  arises from the large  $\omega_{\text{ln}}$ , and the large boost from combining LEP and HEF arises since *e*<sup>−1/2</sup><sup> $\geq$ </sup> 2*e*<sup>−1/ $\lambda$ </sup>. Clearly, an electron-phonon coupling could play a similar role in further enhancing  $T_c$ .

### **IV. COMPETING ORDER PARAMETER SYMMETRIES**

The above calculations have been limited to a pure *d*-wave gap symmetry without harmonic content. In tetragonal symmetry there are five symmetry classes of superconducting gap and each class can involve higher harmonics of the given symmetry. $35$  While we have not solved the tensor Eliashberg equations, it is straightforward to generalize the  $\lambda$ calculations to include harmonic structure and to calculate the leading  $\lambda$  eigenvalue for each symmetry class. The results are shown in Fig. [6](#page-3-1) following the analysis of Ref. [36.](#page-4-34) We see that:  $(1)$  the pure- $d$  analysis of Sec. II holds in the low doping regime; (2) near the VHS, harmonic content stabilizes  $d$ -wave symmetry leading to the largest gaps; and  $(3)$ in this regime, other symmetries can become comparable to *d* wave. In particular, there is a tendency toward *s*-wave pairing in the overdoped case.

#### **V. CONCLUSIONS**

In summary, we have shown that within the present model the *d*-wave pairing weights  $\alpha^2 F_d$  and  $\alpha^2 F_z$  extend to very high energies of  $\sim$ 1 eV when the magnetic response of the system is properly taken into account. The associated superconducting gap is quite substantial being around 16 meV at low dopings.  $\alpha^2 F_d$  is found to contain not only the expected LEP below 200 meV but also a previously unrecognized HEF over 0.3–1.2 eV. We find that the LEP and HEF both play an important role in yielding a large gap in our model. This suggests that electron-phonon coupling could be important for further enhancing  $T_c$  as suggested by the isotope effect.<sup>37</sup> The gap spectrum  $\Delta(\omega)$  generally extends to the limit of  $\alpha^2 F$  (~1.5 eV) and may provide insight into a number of anomalous features connected with optical properties of the cuprates summarized, for example, in Ref. [38.](#page-4-36)

The scarcity of high- $T_c$  superconductors arises in part from the fact that when superconducting pairing is sufficiently strong, corresponding and stronger instabilities arise in other channels. We have for the most part neglected the effects of competing phases, but it is clear that they will be significant both in the underdoped regime and near the VHS. Near the VHS pair-breaking ferromagnetic scattering increases sharply, strongly suppressing a pure *d*-wave gap. While we find that inclusion of harmonic content could stabilize a *d*-wave superconductor even at the VHS, we have not accounted for a competing FM instability. Indeed, Storey *et al.*[39](#page-4-37) found that in Bi2212 the VHS induces strong pair breaking, suppressing superconductivity so that optimum  $T_c$ falls at a doping below the VHS. This is consistent with the evidence for strong FM pair breaking adduced by Kopp *et al.*[26](#page-4-24)

To conclude, we have demonstrated that when realistic  $\alpha^2 F$ 's are used to solve the Eliashberg equations, the magnetic mechanism is capable of producing transition temperatures and pairing gaps comparable in size to those observed PHYSICAL REVIEW B 78, 134513 (2008)

experimentally in the cuprates. The low values of these key superconducting properties found in earlier calculations are directly attributable to the fact that neutron scattering sees only a fraction of the total magnetic spectral weight in these materials.

## **ACKNOWLEDGMENTS**

It is a pleasure to acknowledge useful discussions with Mark Jarrell. This work was supported by the Basic Energy Sciences, Division of Material Science and Engineering, U.S. Department of Energy under Contract No. DE-FG02- 07ER46352 and benefited from the allocation of supercomputer time at NERSC and Northeastern University's Advanced Scientific Computation Center (ASCC).

- 1T. A. Maier, M. Jarrell, and D. J. Scalapino, Phys. Rev. B **74**, 094513 (2006).
- <span id="page-4-0"></span>2T. A. Maier, A. Macridin, M. Jarrell, and D. J. Scalapino, Phys. Rev. B 76, 144516 (2007).
- <span id="page-4-2"></span><span id="page-4-1"></span><sup>3</sup>E. Demler and S.-C. Zhang, Nature (London) **396**, 733 (1998); P. Dai, H. A. Mook, S. M. Hayden, G. Aeppli, T. G. Perring, R. D. Hunt, and F. Dogan, Science **284**, 1344 (1999).
- 4H.-Y. Kee, S. A. Kivelson, and G. Aeppli, Phys. Rev. Lett. **88**, 257002 (2002).
- <span id="page-4-4"></span><span id="page-4-3"></span>5H. Woo, P. Dai, S. M. Hayden, H. A. Mook, T. Dahm, D. J. Scalapino, T. G. Perring, and F. Dogan, Nat. Phys. **2**, 600  $(2006).$
- <sup>6</sup>G. M. Eliashberg, Sov. Phys. JETP 11, 696 (1960).
- <span id="page-4-5"></span><sup>7</sup> P. W. Anderson, Science **316**, 1705 (2007).
- <span id="page-4-6"></span>8T. A. Maier, D. Poilblanc, and D. J. Scalapino, Phys. Rev. Lett. 100, 237001 (2008).
- 9S. Chakraborty, D. Galanakis, and P. Phillips, arXiv:0807.2854 (unpublished).
- <span id="page-4-7"></span>10R. J. Radtke, S. Ullah, K. Levin, and M. R. Norman, Phys. Rev. B 46, 11975 (1992).
- <span id="page-4-8"></span>11A. J. Millis, H. Monien, and D. Pines, Phys. Rev. B **42**, 167  $(1990).$
- <span id="page-4-9"></span>12H.-B. Schüttler and M. R. Norman, Phys. Rev. B **54**, 13295  $(1996).$
- <span id="page-4-10"></span>13R. S. Markiewicz, S. Sahrakorpi, and A. Bansil, Phys. Rev. B 76, 174514 (2007).
- <span id="page-4-11"></span>14A. Macridin, M. Jarrell, T. Maier, and D. J. Scalapino, Phys. Rev. Lett. 99, 237001 (2007).
- <span id="page-4-12"></span>15D. J. Scalapino, E. Loh, Jr., and J. E. Hirsch, Phys. Rev. B **34**, 8190 (1986); 35, 6694 (1987).
- <span id="page-4-13"></span><sup>16</sup>P. Monthoux, Phys. Rev. B **68**, 064408 (2003).
- <span id="page-4-14"></span>17S. Moukouri, S. Allen, F. Lemay, B. Kyung, D. Poulin, Y. M. Vilk, and A.-M. S. Tremblay, Phys. Rev. B 61, 7887 (2000).
- <span id="page-4-15"></span><sup>18</sup>T. Das, R. S. Markiewicz, and A. Bansil, arXiv:0807.4257 (unpublished).
- <span id="page-4-16"></span><sup>19</sup>We use  $Z_0$  to distinguish the renormalization factor used in calculating  $\chi_0$ ; it is not identical to the doping dependent *Z* found from the Eliashberg equations.
- <span id="page-4-18"></span><span id="page-4-17"></span>20T. A. Maier, M. Jarrell, and D. J. Scalapino, Phys. Rev. B **75**, 134519 (2007).
- <sup>21</sup>The hopping parameters are taken from Ref. [13:](#page-4-11)  $(t, t', t'', t''')$  $=(360, -100, 35, 10)$  meV.
- <span id="page-4-19"></span> $22$ This BCS renormalization factor is the inverse of that normally used in Green function theory.
- <span id="page-4-20"></span><sup>23</sup>Calculated  $\alpha^2 F$ 's are mainly odd in  $\omega$ ; a small even component was subtracted out in the displayed results.
- <span id="page-4-22"></span><span id="page-4-21"></span>24H.-Y. Kee, S. A. Kivelson, and G. Aeppli, Phys. Rev. Lett. **88**, 257002 (2002); J. Lorenzana, G. Seibold, and R. Coldea, Phys. Rev. B 72, 224511 (2005); C. Stock, W. J. L. Buyers, Z. Yamani, Z. Tun, R. J. Birgeneau, R. Liang, D. Bonn, and W. N. Hardy, *ibid.* **77**, 104513 (2008).
- 25R. Hlubina, S. Sorella, and F. Guinea, Phys. Rev. Lett. **78**, 1343  $(1997).$
- <span id="page-4-23"></span>26A. Kopp, A. Ghosal, and S. Chakravarty, Proc. Natl. Acad. Sci. U.S.A. **104**, 6123 (2007).
- <span id="page-4-24"></span><sup>27</sup> N. Bulut and D. J. Scalapino, Phys. Rev. B **54**, 14971 (1996).
- <span id="page-4-25"></span>28A. E. Karakozov, E. G. Maksimov, and A. A. Mikhailovsky, Solid State Commun. **79**, 329 (1991).
- <span id="page-4-26"></span>29F. Marsiglio, M. Schossmann, and J. P. Carbotte, Phys. Rev. B **37**, 4965 (1988).
- <span id="page-4-27"></span><sup>30</sup> J. F. Zasadzinski, L. Ozyuzer, L. Coffey, K. E. Gray, D. G. Hinks, and C. Kendziora, Phys. Rev. Lett. **96**, 017004 (2006).
- <span id="page-4-28"></span> $31$  L. Coffey, arXiv:cond-mat/0103518 (unpublished).
- <span id="page-4-29"></span><sup>32</sup> F. Marsiglio and J. P. Carbotte, Phys. Rev. B **36**, 3937 (1987).
- <span id="page-4-30"></span><sup>33</sup> P. B. Allen and R. C. Dynes, Phys. Rev. B **12**, 905 (1975).
- <span id="page-4-31"></span><sup>34</sup> J. P. Carbotte, Rev. Mod. Phys. **62**, 1027 (1990).
- <sup>35</sup> R. Hlubina, Phys. Rev. B **59**, 9600 (1999).
- <span id="page-4-33"></span><span id="page-4-32"></span>36F. Guinea, R. S. Markiewicz, and M. A. H. Vozmediano, Phys. Rev. B 69, 054509 (2004).
- <span id="page-4-35"></span><span id="page-4-34"></span>37G.-H. Gweon, T. Sasagawa, S. Y. Zhou, J. Graf, H. Takagi, D.-H. Lee, and A. Lanzara, Nature (London) 430, 187 (2004); R. Khasanov, A. Shengelaya, D. Di Castro, D. G. Eshchenko, I. M. Savic, K. Conder, E. Pomjakushina, J. Karpinski, S. Kazakov, and H. Keller, Phys. Rev. B 75, 060505(R) (2007).
- <span id="page-4-36"></span>38H. J. A. Molegraaf, C. Presura, D. van der Marel, P. H. Kes, and M. Li, Science 295, 2239 (2002); A. F. Santander-Syro, R. P. S. M. Lobo, N. Bontemps, Z. Konstantinovic, Z. Z. Li, and H. Raffy, Europhys. Lett. **62**, 568 (2003).
- <span id="page-4-37"></span><sup>39</sup> J. G. Storey, J. L. Tallon, and G. V. M. Williams, Phys. Rev. B 76, 174522 (2007).