

Large-scale Monte Carlo simulations of the isotropic three-dimensional Heisenberg spin glass

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We study the Heisenberg spin glass by large-scale Monte Carlo simulations for sizes up to 32^3 , down to temperatures below the transition temperature claimed in earlier work. The data for the larger sizes show more marginal behavior than that for the smaller sizes, indicating the lower critical dimension is close to, and possibly equal to, 3. We find that the spins and chiralities behave in a quite similar manner.

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I. INTRODUCTION

Following the convincing numerical work of Ballesteros *et al.*¹ there has been little doubt that Ising spin glasses in three dimensions have a finite temperature transition. In this paper we shall study a related model for which the existence of a finite temperature transition is still controversial: the isotropic Heisenberg spin glass, which is composed of classical spins with three components. Early work²⁻⁴ indicated a zero temperature transition, or possibly a transition at a very low but nonzero temperature. However, following the pioneering work of Villain,⁵ which emphasized the role of “chiralities” (Ising-like variables which describe the handedness of the noncollinear spin structures), Kawamura and Tanemura⁶ proposed, for the *XY* case (which have two-component spins), that the spin-glass transition only occurs at $T_{SG}=0$ and that a *chiral-glass* transition occurs at a finite temperature T_{CG} . This scenario requires that spins and chiralities decouple at long length scales. Kawamura and collaborators subsequently proposed⁷⁻⁹ that this “spin-chirality decoupling” scenario also holds for Heisenberg spin glasses.

However, the absence of a spin-glass transition in Heisenberg spin glasses has been challenged by Matsubara *et al.*,^{10,11} and Nakamura and Endoh¹² who argued that the spins and chiralities order at the same low but finite temperature. Recently Picco and Ritort¹³ also claimed a finite T_{SG} and inferred that probably $T_{SG}=T_{CG}$, though they did not investigate the chiralities directly. In earlier work,¹⁴ referred to as LY, we studied spin and chiral correlations on an equal footing, using the method of analysis that was the most successful for the Ising spin glass,¹ namely finite-size scaling of the correlation length. Considering a modest range of sizes, $L \leq 12$, we found that the behavior of spins and chiralities was quite similar and they both had a finite temperature transition, apparently at the same temperature.

However, quite recently Campos *et al.*¹⁵ were able to study larger sizes than LY, up to $L=32$. They agreed with LY that there is a single transition at which both spins and chiralities order, but they also found evidence for crossover, at the largest sizes, to a “marginal” behavior, reminiscent of that at the Kosterlitz-Thouless-Berezinskii^{16,17} (KTB) transition in the two-dimensional *XY* ferromagnet where there is a finite transition temperature T_c but no long-range order for

$T < T_c$. In fact the region below T_c is a *line* of critical points in the KTB theory. If a line of critical points also exists in the three-dimensional Heisenberg spin glass, then $d=3$ is the lower critical dimension d_l , below which there is no transition. However, given numerical uncertainties, Campos *et al.* cannot rule out the possibility that d_l is slightly less than 3, in which case there is spin-glass order below T_{SG} .

Hukushima and Kawamura⁹ studied sizes up to $L=20$ and found more marginal behavior when comparing $L=16$ and 20, than for the smaller sizes. However, they argued that this effect is greater for the spin correlation length than for the chiral correlation length, and hence concluded that, while T_{CG} is finite, T_{SG} is zero or possibly nonzero but less than T_{CG} , i.e., there is spin-chirality decoupling.

In this paper we perform Monte Carlo simulations of the Heisenberg spin glass, along the lines of LY, but for larger sizes. Our main motivation is to investigate the claim of Campos *et al.*¹⁵ that there is a line of critical points for $T < T_{SG}$. In order to test whether there is a critical *line*, as proposed by Campos *et al.*, or the usual situation of a single critical *point* at T_{SG} , it is necessary to investigate the behavior of the system *below* the estimated T_{SG} . Campos *et al.* were not able to do this and the evidence for the critical line was based on estimating corrections to scaling at T_{SG} . This is rather *indirect*, and perhaps not very reliable because the range of sizes and quality of the data, are not sufficient to disentangle the various corrections to scaling unambiguously. Very recently, the analysis of Campos *et al.* has been criticized by Campbell and Kawamura.¹⁸ Here, we do not rely on corrections to scaling but obtain data below T_{SG} and so *directly* find evidence for more marginal behavior at larger sizes.

As we shall see, the transition temperature in the Heisenberg spin glass is very low and so it is surprising to us that Campos *et al.*¹⁵ did not use the technique of “parallel tempering,”^{19,20} which is the commonly used approach to speed up simulations of spin glasses, especially at very low temperatures. Equilibrating lattices as large as 32^3 is very challenging, especially in the absence of parallel tempering. A second motivation of our study is therefore to use parallel tempering and to combine this with a very useful equilibration test described in the next section, to *ensure* that the data are equilibrated.

From our results, we see a strong crossover to much more marginal behavior for sizes $L \gtrsim 24$, in agreement with Campos *et al.*¹⁵ Whether there is a KTB-like critical line as proposed by Campos *et al.*, is, however, unclear. The lower critical dimension, d_l seems to be close to, or possibly equal to, 3. The behavior of the spins and chiralities is rather similar, and so, in contrast to Kawamura and collaborators,^{9,18} we do *not* believe that our data supports spin-chirality decoupling.

II. MODEL AND ANALYSIS

We use the standard Edwards-Anderson spin-glass model

$$\mathcal{H} = - \sum_{\langle i,j \rangle} J_{ij} \mathbf{S}_i \cdot \mathbf{S}_j, \quad (1)$$

where the \mathbf{S}_i are three-component classical vectors of unit length at the sites of a simple cubic lattice, and the J_{ij} are nearest-neighbor interactions with a Gaussian distribution with zero mean and standard deviation unity. Periodic boundary conditions are applied on lattices with $N=L^3$ spins.

The spin-glass order parameter, $q^{\mu\nu}(\mathbf{k})$, at wave vector \mathbf{k} , is defined to be

$$q^{\mu\nu}(\mathbf{k}) = \frac{1}{N} \sum_i S_i^{\mu(1)} S_i^{\nu(2)} e^{i\mathbf{k} \cdot \mathbf{R}_i}, \quad (2)$$

where μ and ν are spin components, and “(1)” and “(2)” denote two identical copies of the system with the same interactions. From this we determine the wave-vector-dependent spin-glass susceptibility $\chi_{\text{SG}}(\mathbf{k})$ by

$$\chi_{\text{SG}}(\mathbf{k}) = N \sum_{\mu,\nu} [\langle |q^{\mu\nu}(\mathbf{k})|^2 \rangle]_{\text{av}}, \quad (3)$$

where $\langle \dots \rangle$ denotes a thermal average and $[\dots]_{\text{av}}$ denotes an average over disorder. The spin-glass correlation length is then determined^{1,21} from

$$\xi_L = \frac{1}{2 \sin(k_{\min}/2)} \left(\frac{\chi_{\text{SG}}(0)}{\chi_{\text{SG}}(\mathbf{k}_{\min})} - 1 \right)^{1/2}, \quad (4)$$

where $\mathbf{k}_{\min} = (2\pi/L)(1, 0, 0)$.

For the Heisenberg spin-glass, Kawamura⁷ defines the local chirality in terms of three spins on a line as follows:

$$\kappa_i^\mu = \mathbf{S}_{i+\hat{\mu}} \cdot \mathbf{S}_i \times \mathbf{S}_{i-\hat{\mu}}. \quad (5)$$

The chiral glass susceptibility is then given by

$$\chi_{\text{CG}}^\mu(\mathbf{k}) = N [\langle |q_c^\mu(\mathbf{k})|^2 \rangle]_{\text{av}}, \quad (6)$$

where the chiral overlap $q_c^\mu(\mathbf{k})$ is given by

$$q_c^\mu(\mathbf{k}) = \frac{1}{N} \sum_i \kappa_i^{\mu(1)} \kappa_i^{\mu(2)} e^{i\mathbf{k} \cdot \mathbf{R}_i}. \quad (7)$$

We define the chiral correlation lengths $\xi_{c,L}^\mu$ by

$$\xi_{c,L}^\mu = \frac{1}{2 \sin(k_{\min}/2)} \left(\frac{\chi_{\text{CG}}(0)}{\chi_{\text{CG}}^\mu(\mathbf{k}_{\min})} - 1 \right)^{1/2}, \quad (8)$$

in which $\chi_{\text{CG}}(\mathbf{k}=0)$ is independent of μ . Note that $\xi_{c,L}^\mu$ will, in general, be different for $\hat{\mu}$ along \mathbf{k}_{\min} (the \hat{x} direction) and

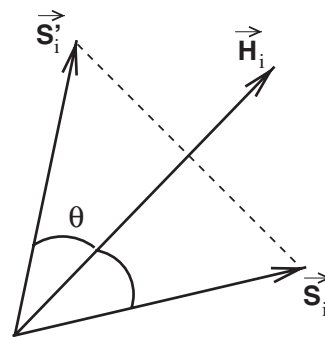


FIG. 1. \mathbf{H}_i is the local field on site i due to its neighbors. The spin at i is initially in direction \mathbf{S}_i . In a microcanonical (over-relaxation) move, the spin is reflected about \mathbf{H}_i according to Eq. (9), and so ends up in direction \mathbf{S}'_i .

perpendicular to \mathbf{k} , though we expect that this difference will vanish for large sizes. We denote these two lengths by $\xi_{c,L}^{\parallel}$ and $\xi_{c,L}^{\perp}$, respectively.

To equilibrate the system in as small a number of sweeps as possible, with the minimum amount of CPU time, we perform three types of Monte Carlo moves:

(1) “Microcanonical” sweeps²² (also known as “over-relaxation” sweeps). We sweep sequentially through the lattice, and, at each site, compute the local field on the spin, $\mathbf{H}_i = \sum_j J_{ij} \mathbf{S}_j$. The new value for the spin on site i is taken to be its old value reflected about \mathbf{H}_i , i.e.,

$$\mathbf{S}'_i = -\mathbf{S}_i + 2 \frac{\mathbf{S}_i \cdot \mathbf{H}_i}{H_i^2} \mathbf{H}_i, \quad (9)$$

see Fig. 1. These sweeps are microcanonical because they preserve energy. They are very fast because the operations are simple and no random numbers are needed. For reasons that are not fully understood, it also seems that they “stir up” the spin configuration very efficiently¹⁵ and the system equilibrates faster than if one only uses “heatbath” updates, described next.

(2) “Heatbath” sweeps.⁴ Since the microcanonical sweeps conserve energy they do not equilibrate the system.

We therefore also include some heatbath sweeps since these do change the energy, typically doing one after every 10 microcanonical sweeps. As for the microcanonical case, we sweep sequentially through the lattice. Referring to Fig. 1, we take the direction of \mathbf{H}_i , the local field on site i to be the polar axis, denote the polar angle as θ , and define the azimuthal angle ϕ such that $\phi=0$ for the old spin direction. The new spin direction \mathbf{S}'_i is characterized by angles θ and ϕ , relative to \mathbf{H}_i , as follows. The energy does not depend on the azimuthal angle, and so $\phi=2\pi r_1$, where r_1 is a random number chosen uniformly between 0 and 1. The polar angle is chosen such that, after the move, the spin is in local equilibrium with respect to the local field, i.e., if $x = \cos \theta$, then

$$P(x) = \frac{\beta H_i}{2 \sinh \beta H_i} e^{\beta H_i x}, \quad (10)$$

where $\beta=1/T$. To determine x with this probability, the procedure⁴ is to equate the cumulative distribution

TABLE I. Parameters of the simulations. N_{samp} is the number of samples, N_{equil} is the number of microcanonical Monte Carlo sweeps for equilibration for each of the $2N_T$ replicas for a single sample, and N_{meas} is the number of microcanonical sweeps for measurement. The number of heatbath sweeps is equal to 10% of the number of microcanonical sweeps. T_{min} and T_{max} are the lowest and highest temperatures simulated, and N_T is the number of temperatures used in the parallel tempering.

L	N_{samp}	N_{equil}	N_{meas}	T_{min}	T_{max}	N_T
4	500	5×10^3	10^4	0.0400	0.96	40
6	500	6×10^4	1.2×10^5	0.0400	0.96	40
8	536	8×10^4	1.6×10^5	0.0400	0.96	57
12	204	8×10^5	1.6×10^6	0.0400	0.61	88
16	202	8×10^5	1.6×10^6	0.1015	0.49	73
24	160	3.2×10^5	6.4×10^5	0.1200	0.49	122
32	56	9.6×10^5	1.28×10^6	0.1210	0.40	120

$$Q(x) = \int_{-1}^x P(x') dx' = \frac{e^{\beta H_i x} - e^{-\beta H_i}}{e^{\beta H_i} - e^{-\beta H_i}}, \quad (11)$$

to a second random number r_2 (also with a uniform distribution between 0 and 1) so

$$x = \frac{1}{\beta H_i} \ln[1 + r_2(e^{2\beta H_i} - 1)] - 1. \quad (12)$$

It is then necessary to convert the new spin direction \mathbf{S}'_i back to cartesian coordinates. Denoting the polar and azimuthal angles of \mathbf{H} by θ_H and ϕ_H relative to the Cartesian reference frame, and remembering that θ and ϕ are relative to \mathbf{H} , we have²³

$$S'_x = S''_x \cos \phi_H - S''_y \sin \phi_H, \quad (13)$$

$$S'_y = S''_x \sin \phi_H + S''_y \cos \phi_H, \quad (14)$$

$$S'_z = \cos \theta \cos \theta_H - \sin \theta \sin \theta_H \cos \phi, \quad (15)$$

where

$$S''_x = \cos \theta \sin \theta_H + \sin \theta \cos \theta_H \cos \phi, \quad (16)$$

$$S''_y = \sin \theta \sin \phi. \quad (17)$$

We see that the calculations in the heatbath moves are quite involved, which is why we do mainly microcanonical moves, just including *some* heatbath moves to change the energy and thereby ensure the algorithm is ergodic. Note, though, that the acceptance probability for the heatbath moves, and also for the microcanonical moves, is unity, so *no moves are wasted*.

(3) ‘‘Parallel tempering’’ sweeps. At low temperatures spin glasses are easily trapped in minima (valleys) of the free energy. In order to ensure that the system visits different minima with the correct Boltzmann weight during the time of the simulation we use the method of parallel tempering.^{19,20} One takes N_T copies of the system with the same bonds but at a range of different temperatures. The minimum temperature, $T_{\text{min}} \equiv T_1$, is the low temperature where one wants to investigate the system (below T_{SG} in our case), and the maximum, $T_{\text{max}} \equiv T_{N_T}$, is high enough that the system equilibrates very fast (well above T_{SG} in our case). A

parallel tempering sweep consists of swapping the temperatures of the spin configurations at a pair of neighboring temperatures, T_i and T_{i+1} , for $i=1, 2, \dots, T_{N_T-1}$ with a probability that satisfies the detailed balance condition. The Metropolis probability for this is¹⁹

$$P(T \text{ swap}) = \begin{cases} \exp(-\Delta\beta\Delta E) & (\text{if } \Delta\beta\Delta E > 0), \\ 1 & (\text{otherwise}), \end{cases} \quad (18)$$

where $\Delta\beta = 1/T_i - 1/T_{i+1}$ and $\Delta E = E_i - E_{i+1}$, in which E_i is the energy of the copy at temperature T_i . In this way, a given set of spins (i.e., a copy) performs a random walk in temperature space. Suppose that at some time in the simulation a copy is trapped in a valley at low T . Later on it will reach a high temperature where it randomizes quickly, so that when, still later, it is again at a low temperature, there is no reason for it to be in the same valley that it was in before. We do one sweep of temperature swaps after every 10 microcanonical sweeps.

Table I gives the parameters of the simulations. It will be seen that the number of temperatures is very large, and it is instructive to discuss the reason for this. The difference between two neighboring temperatures, ΔT , must be sufficiently small that there is an overlap between the energy distributions at those temperatures, so there is a reasonable probability that the same spin configuration occurs for both temperatures. Otherwise, the probability for accepting the temperature swap, Eq. (18), will be very small. Relating the width of the temperature distribution to the specific heat in the normal way, one finds that

$$\frac{\delta T}{T} \lesssim \frac{1}{\sqrt{CN}}, \quad (19)$$

where C is the specific heat per spin. For the Heisenberg spin glass C tends to a constant at low T because of the Gaussian (spin-wave) fluctuations about the local equilibrium positions. Hence, for a given size, we choose temperatures which decrease in a geometric manner. The required number of temperatures to cover a fixed range of T increases proportional to \sqrt{N} , and so, for large N , many copies are needed to cover even a factor of 2 in T . However, for the Ising spin glass with Gaussian interactions, the specific heat tends to zero as $T \rightarrow 0$ (presumably linearly in T). Hence much bigger

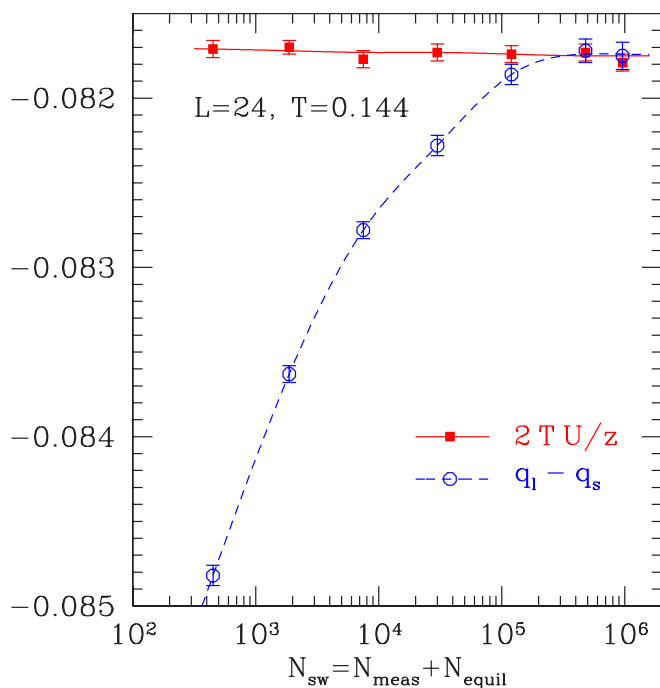


FIG. 2. (Color online) Equilibration plot, testing Eq. (20) for $L=24$ at the $T=0.144$. It is seen that the data for $2TU/z$ come together at about 3×10^5 total sweeps (equilibration plus measurement) and then stay at their common value indicating that equilibration has been achieved. The lines are guides to the eye. The figure shows that the energy comes close to its equilibrium value very quickly, whereas $q_l - q_s$ takes much longer.

steps in T can be taken at low T than for the Heisenberg case, leading to fewer temperatures being needed.

To test for equilibration²⁴ we require that data satisfy the relation²⁵

$$q_l - q_s = \frac{2}{z}TU, \quad (20)$$

which is valid for a Gaussian bond distribution. Here $U = -[\sum_{\langle i,j \rangle} J_{ij} \langle \mathbf{S}_i \cdot \mathbf{S}_j \rangle]_{\text{av}}$ is the average energy per spin, $q_l = (1/N_b) \sum_{\langle i,j \rangle} [\langle \mathbf{S}_i \cdot \mathbf{S}_j \rangle^2]_{\text{av}}$ is the “link overlap,” $q_s = (1/N_b) \sum_{\langle i,j \rangle} [\langle (\mathbf{S}_i \cdot \mathbf{S}_j)^2 \rangle]_{\text{av}}$, $N_b = (z/2)N$ is the number of nearest-neighbor bonds, and z ($=6$ here) is the lattice coordination number. Equation (20) is easily derived by integrating by parts the expression for the average energy with respect to J_{ij} , noting that the average $[\dots]_{\text{av}}$ is over a Gaussian function of the J_{ij} 's.

The spins are initialized in random directions so the energy, the right-hand side (RHS) of Eq. (20), is initially close to zero and decreases, presumably monotonically, to its equilibrium value as the length of the simulation increases. Hence the RHS of Eq. (20) will be too *large* if the simulation is too short to equilibrate the system. On the other hand, on the left-hand side (LHS) of Eq. (20), q_l will be too *small* if the simulation is too short because it starts off close to zero and then increases with MC time as the two replicas start to find the same local minima. The quantity q_s will be less dependent on Monte Carlo time than q_l since it is a local

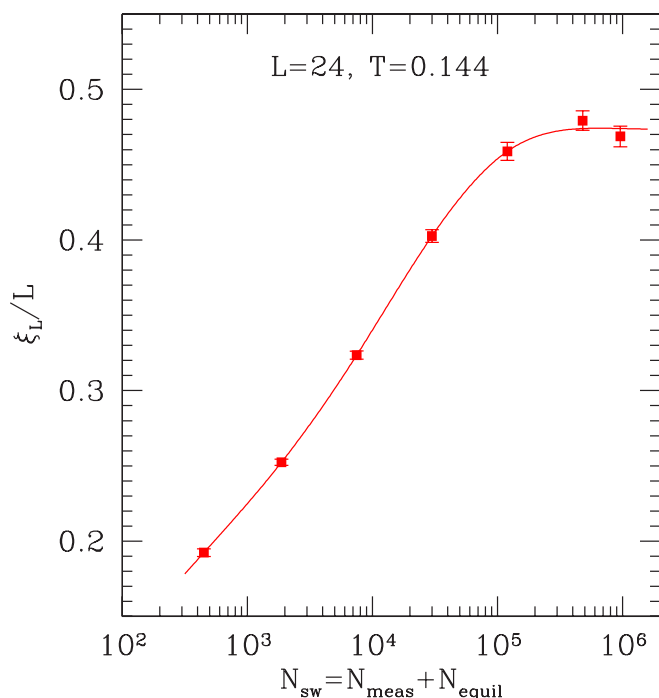


FIG. 3. (Color online) A plot of ξ_L/L as a function of the total number of sweeps for $L=24$ at $T=0.144$. The figure shows that the data flattens off at around 3×10^5 sweeps, the value where the two sets of data in Fig. 2 start to agree. This indicates that when the data in Fig. 2 agree within high precision, i.e., when Eq. (20) is satisfied, the correlation length has reached its equilibrium value. The line is a guide to the eye.

variable for a single replica. (For the Ising case it is just a constant.) Hence if the simulation is too short the LHS of Eq. (20) will be too low. In other words, the two sides of Eq. (20) approach the common equilibrium value from *opposite directions* as the length of the simulation increases. Only if Eq. (20) is satisfied within small error bars do we accept the results of a simulation.

Figure 2 shows a test to verify that Eq. (20) is satisfied at long times. For the parameters used, $L=24, T=0.144$, this occurs when the total number of sweeps ($N_{\text{sw}} = N_{\text{equil}} + N_{\text{meas}}$) is about 3×10^5 . Figures 3 and 4 show that the spin and chiral correlation lengths appear to become independent of N_{sw} , and hence are presumably equilibrated, when N_{sw} is larger than this *same* value. Hence, it appears that when Eq. (20) is satisfied to high precision, the data for the correlation lengths is equilibrated.

With the number of sweeps shown in Table I, Eq. (20) was satisfied for all sizes and temperatures. The error bars are made sufficiently small by averaging over a large number of samples. We are simulating system sizes which are very large by spin-glass standards (up to $N=32^3$), so it is crucial to have a stringent test like this for equilibration.

Since ξ_L/L is dimensionless, it has the finite size scaling form^{1,26,27}

$$\frac{\xi_L}{L} = \tilde{X}[L^{1/\nu}(T - T_{\text{SG}})], \quad (21)$$

where ν is the correlation length exponent. Note that there is no power of L multiplying the scaling function \tilde{X} , as there

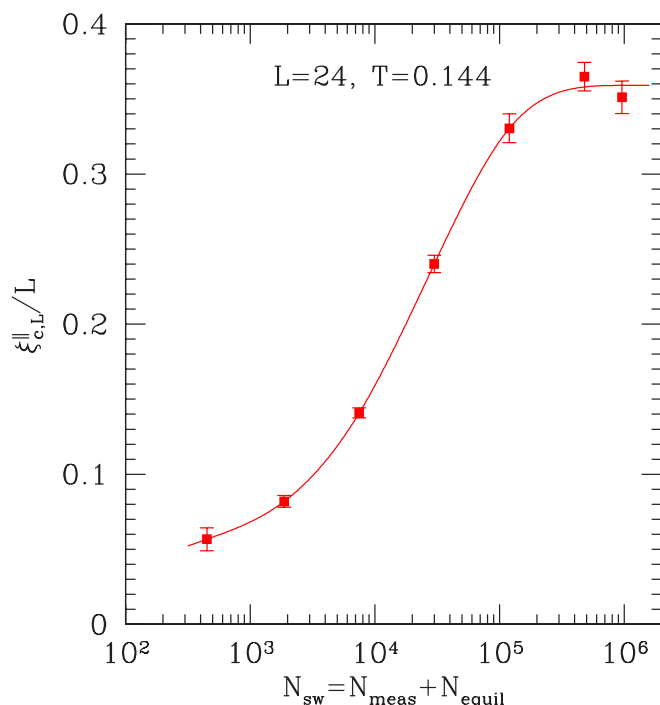


FIG. 4. (Color online) A plot of $\xi_{c,L}^{\parallel}/L$ as a function of the total number of sweeps for $L=24$ at $T=0.144$. As in Fig. 3, this figure shows that the data flattens off at around 3×10^5 sweeps, the value where the two sets of data in Fig. 2 start to agree. The line is a guide to the eye.

would be for a quantity with dimensions. There are analogous expressions for the chiral correlation lengths. From Eq. (21) it follows that the data for ξ_L/L for different sizes come together at $T=T_{SG}$. In addition, they are also expected to splay out again on the low- T side if there is spin glass order below T_{SG} . In a marginal situation, with a line of critical points as in the KTB transition, the data for different sizes would come together at T_{SG} and then stick together at lower T , see, for example, Fig. 3 of Ref. 1.

III. RESULTS

We studied sizes from $L=4$ to $L=32$, as shown in Table I. The CPU time involved to get this data is about 15 Mac G5 CPU years.

We start with the data for ξ_L/L shown in Fig. 5. As was found earlier by LY, the smaller sizes show a clear intersection, and also splay out at lower temperatures which would indicate spin-glass order. However, the situation for the largest sizes is less clearcut, with the $L=32$ and 24 data only coming together at a somewhat lower temperature than the temperature where the $L \leq 16$ data intersect. Furthermore the data for the largest sizes at the lowest temperatures does not depend strongly on size, indicating close to marginal behavior. This is qualitatively in agreement with Campos *et al.*¹⁵ An enlarged view of the important region is shown in Fig. 6.

Data for the parallel chiral-glass correlation length is shown in Fig. 7. The main features are the same as found for the spin-glass correlation length in Fig. 5. The smaller sizes

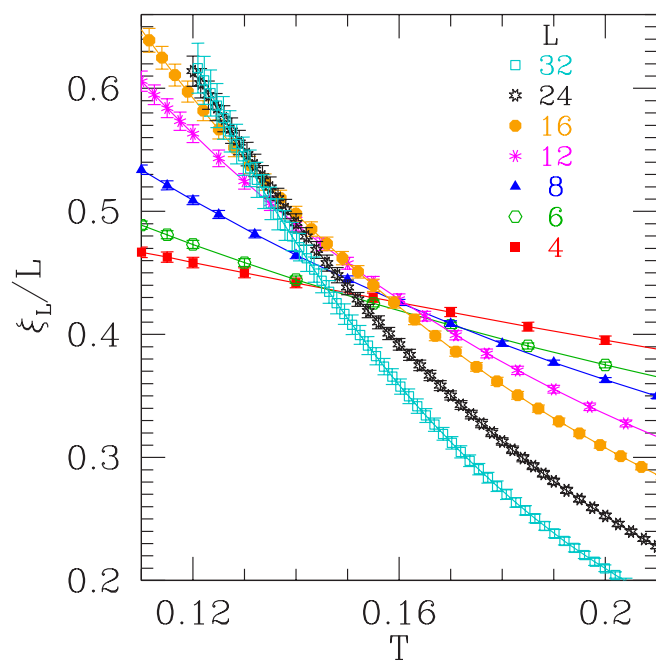


FIG. 5. (Color online) Data for ξ_L/L , the spin-glass correlation length divided by system size, as a function of T for different system sizes. Note that there are very many data points for the larger sizes. This is because a large number of temperatures are needed for the parallel tempering algorithm, as discussed in Sec. II.

show clear intersections, but the $L=32$ data lies lower, though perhaps not quite to the same extent as in Fig. 5. As for the spin-glass correlation length, the two larger sizes $L=32$ and 24 only come together at a lower temperature than the temperature (range) where the smaller sizes intersect, and do not splay out at still lower temperatures, at least in the range of T studied and within the error bars. An enlarged

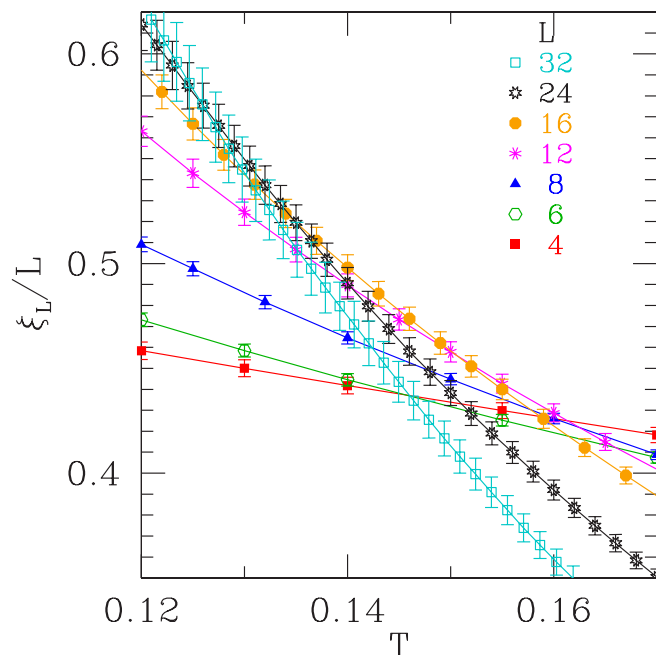


FIG. 6. (Color online) Enlarged view of a region of Fig. 5.

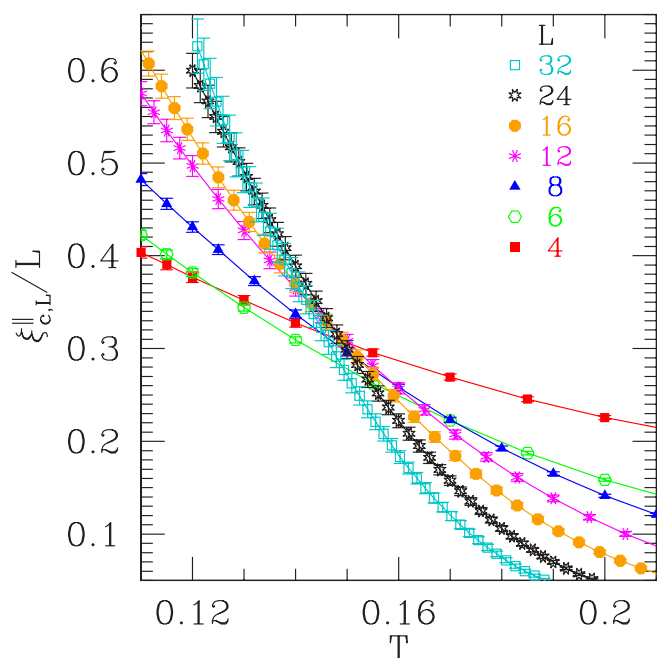


FIG. 7. (Color online) Data for $\xi_{c,L}^{\parallel}/L$, the “parallel” chiral-glass correlation length divided by system size, as a function of T for different system sizes.

view is shown in Fig. 8. In our view, the data in Figs. 5 and 7 are not very different; in particular, in both figures, the data for the largest *two* sizes merge at about the same temperature and stick together at lower temperatures. If we look at the largest *three* sizes there *is* a somewhat greater tendency for the chiral data to splay out at low T . Clearly larger sizes are needed to be sure of the trend in the thermodynamic limit.

With this data it is difficult to come to a firm conclusion about the nature of the transition. Clearly the smaller sizes have large corrections to scaling behavior, and so, at best, only the two largest sizes are in the asymptotic scaling region. The data is consistent with a line of critical points analogous to that in the KTB transition, as proposed by Campos *et al.*, in which the data for large sizes would merge at a single temperature and remain independent of size at lower temperatures. This is a scenario in which the lower critical dimension d_l is equal to 3. Another scenario consistent with our data, in which d_l is also equal to 3, is that $T_{SG}=T_{CG}=0$ but with an exponential divergence of the correlation lengths at $T=0$. In that case, it is likely that the data for large sizes would join a common curve, but the temperature at which the common curve is joined would decrease as the system size increased. This scenario is found in the two-dimensional Heisenberg ferromagnet, whereas the critical line of the KTB theory is found in the two-dimensional XY ferromagnet. Our data is also consistent with the possibility that, for the large sizes, the curves weakly intersect, implying a finite T_{SG} and a lower critical dimension slightly *less* than 3.

Unfortunately, because of the crossover effects in the data, it does not appear possible to give a meaningful estimate of the critical exponents, ν and η .

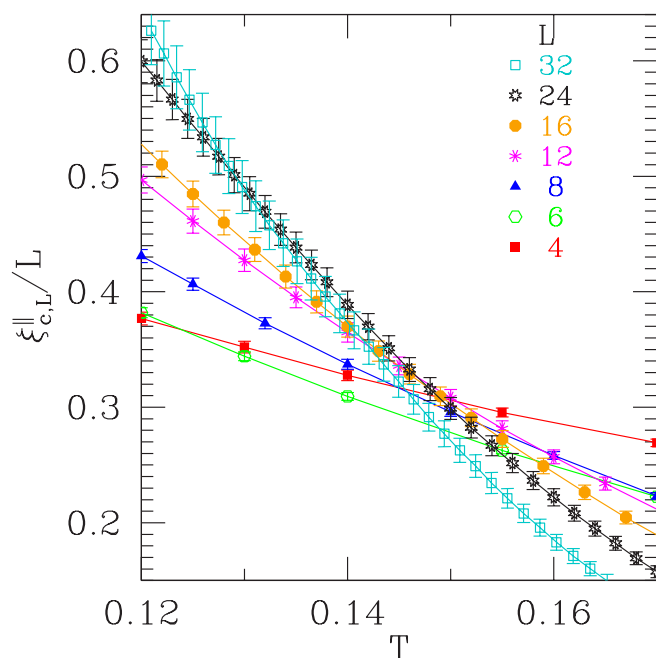


FIG. 8. (Color online) Enlarged view of a region of Fig. 7.

IV. CONCLUSIONS

In agreement with Campos *et al.*¹⁵ and Hukushima and Kawamura,⁹ we find a crossover in behavior for system sizes larger than about 16. Results from the larger sizes indicate a more marginal behavior than those from the smaller sizes. Unlike Kawamura and collaborators,^{9,18} we do not find that this effect is very different for the spins and chiralities. Compare, for example, Figs. 5 and 7. However, larger sizes would be needed to *confirm* that the asymptotic behaviors of the spin- and chiral-glass correlation lengths are indeed similar. Our data extends up to $L=32$, somewhat larger than the sizes ($L \leq 20$) studied by Hukushima and Kawamura. Our range of sizes is the same as that of Campos *et al.*, but we are able to go down to lower temperature, in particular *below* the putative spin-glass transition temperature. Whereas Campos *et al.* argue in favor of a critical line below T_{SG} , in our view, other possibilities cannot be ruled out, such as a transition at a lower value of T_{SG} or even an exponential divergence of the correlation length at $T=0$.

To distinguish between these scenarios would require a study of still larger sizes. It will be difficult to go to very much larger sizes without a better algorithm, since the present study used a quite substantial amount of computer time. An unfortunate feature of the present algorithm is that parallel tempering for vector spin models requires a large number of temperatures. The large number arises from the temperature independent specific heat at low temperatures, which, in turn, comes from a rather trivial feature: Gaussian spin-wave fluctuations. However, it is difficult to see how to eliminate their effect, and thereby reduce the number of temperatures.

It is interesting to note that we (and also Campos *et al.*¹⁵) have been able to study larger sizes for the Heisenberg spin glass ($L=32$) than has ever been done for the Ising spin

glass. For example, Katzgraber *et al.*²⁸ studied the Ising spin glass using considerable CPU time, but were still not able to equilibrate sizes larger than $L=16$ near T_{SG} for the case of Gaussian interactions. With $\pm J$ interactions, where “multi-spin coding” speeds the code up further, they could go up to $L=24$.

That one can equilibrate larger sizes in the Heisenberg case is surprising bearing in mind that (i) the Heisenberg algorithm is more complicated and so one sweep takes more CPU time than for the Ising case, and (ii) the transition temperature is significantly lower for the Heisenberg case. Noting that the mean-field transition temperature is $T_{SG}^{MF} = \sqrt{z}/m$, where z is the number of neighbors and m the number of spin components, if we take $T_{SG}=0.145$ for the Heisenberg case, the ratio T_{SG}/T_{SG}^{MF} is about 0.18. For the Ising case, one finds, e.g., Ref. 28, $T_{SG} \approx 0.95$ so $T_{SG}/T_{SG}^{MF} \approx 0.39$ which is more than 2 times the corresponding ratio for the Heisenberg spin glass.

The fact that, despite all this, one can study larger sizes in Heisenberg spin glasses than in Ising spin glasses indicates that barriers are smaller in the Heisenberg case. Another way of putting this is that the extra degrees of freedom in the Heisenberg model allow the system to find a way *around* barriers, which would have to be *gone over* for the Ising spin glass.

Overall, our results indicate that spins and chiralities behave in a quite similar manner, and that the lower critical dimension of the Heisenberg spin glass is close to, and possibly equal to, 3.

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²³See, for example, Secs. 3.2.2 and 3.4.3 of Ref. 29. In our opinion, the factor of $\sin \alpha^r \cos \beta^r \cos \phi$ in the second line of Eq. (3.62) in that reference should be $\sin \alpha^r \sin \beta^r \cos \phi$.

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