Light-induced magnetic precession in (Ga,Mn)As slabs: Hybrid standing-wave Damon-Eshbach modes

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Coherent oscillations associated with spin precessions were observed in ultrafast optical experiments on ferromagnetic (Ga,Mn)As films. Using a complete theoretical description of the processes by which light couples to the magnetization, values for the anisotropy constants and the spin stiffness were unambiguously determined from the data. Estimates for the hole-Mn exchange coupling are significantly larger than those previously reported. Results also reveal an important negative contribution to the energy due to surface anisotropy leading to excitations that are a mixture of bulk waves and surface modes.

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Following the pioneering work of Ohno *et al.* on $(Ga, Mn)As (Ref. 1)$ $(Ga, Mn)As (Ref. 1)$ $(Ga, Mn)As (Ref. 1)$ and $(In, Mn)As (Ref. 2)$ $(In, Mn)As (Ref. 2)$ $(In, Mn)As (Ref. 2)$, and the subsequent discovery of high-temperature ferromagnetism in $(Ga, Mn)N$,³ ferromagnets based on III-V compounds have been the subject of intense study prompted, in part, by the promise that their compatibility with semiconductor technology may lead to novel nonvolatile spintronic applications[.4](#page-3-4) Although it is now well established that the ferromagnetism in these alloys arises from the exchange interaction between holes and the manganese ions⁵ and, hence, that the Curie temperature (T_C) is determined mainly by the hole concentration, $6-8$ the relationship between transport, magnetic properties, and defect configurations is not yet fully understood and, moreover, important magnetic parameters remain poorly known. In this Brief Report, we report on the direct observation of coherent magnetization precessions in (Ga, Mn) As induced by subpicosecond laser pulses. $9-12$ We discuss a comprehensive model of the spin dynamics and the coupling of the magnetic eigenmodes to light, which allows one to gain accurate values of the exchange, bulk, and surface anisotropy constants from time-domain data, establishing a hitherto unrecognized relationship between nonlinear optics and ferromagnetic resonance (FMR).

We studied several films, grown by low-temperature molecular-beam epitaxy on semi-insulating (001) GaAs substrates. Their parameters are listed in Table [I.](#page-1-0) Using a Ti:sapphire laser that provided \sim 70 fs pulses of central wavelength of 800 nm at the repetition rate of 82 MHz, we performed pump-probe differential magnetic Kerr (DMK) experiments at \sim 4–6 K in the Voigt geometry. The laser beams, of energy density per pulse of $\sim 0.2 \ \mu J/cm^2$ (pump) and 0.03 μ J/cm² (probe), penetrated the slabs along the *z* axis [001]. The pump pulses induce coherent magnetic precessions modifying the reflection of the probe pulses that follow behind. Time-domain DMK measurements give the pumpinduced shift of the polarization angle of the reflected probe field, $\delta\theta$, as a function of the time delay between the two pulses.

To calculate the magnetic modes, we consider a (Ga, Mn) As film occupying the region $|z| \le L/2$. The Landau-Lifshitz equation of motion for the magnetization M is¹³

$$
\partial \mathbf{M}/\partial t = \gamma \mathbf{M} \times [\mathbf{H} - \nabla_{\mathbf{M}} F_A + D M^{-1} \nabla^2 \mathbf{M}], \tag{1}
$$

where **H** is the field strength and $F_A = -K_2 M_z^2 / M^2 - K_4 (M_x^4)$ $+ M_y^4$ /2 M^4 – $K_O(M_x - M_y)^2$ /2 M^2 is the magnetic anisotropy contribution to the free-energy density; terms αM_z^4 and of higher order give significantly smaller contributions.¹⁴ K_2 and K_4 are the out-of-plane and in-plane anisotropy constants. The associated terms are consistent with the tetragonal symmetry of our compressively strained films.¹⁴ The term αK_O accounts for a surface reconstruction of orthorhombic symmetry which leads to a monotonic temperature dependence of the orientation of the equilibrium magnetization M_0 from near [100] at $T=0$ to [110] for $T \ge 30$ K (at $H = 0$).^{[15,](#page-3-12)[16](#page-3-13)} While this behavior is ultimately the reason why light pulses can trigger coherent precessions (see below), the K_O term introduces a very small correction to the frequency and intensity of the eigenmodes and, thereby, it will be momentarily ignored. Assuming that both **H** and **M** depend only on *z*, we get using the Maxwell equations of magnetostatics $H_y = 0$ and $H_z + 4\pi M_z = 0$.¹⁷ In our experiments, the external field and, hence, M_0 are parallel to the easy axis which, for simplicity, we assume to be along $[100]$.^{[15](#page-3-12)} Decomposing the field into static and microwave components, $H = H_0$ **i** $+\mathbf{h}(\mathbf{r},t)$ with $|\mathbf{h}| \leq H_0$, we get from Eq. ([1](#page-0-0)) **M** \approx $(M_0, m_y \sin \Omega t, m_z \cos \Omega t)$, where $|m_{y,z}| \le M_0$ and

$$
(\Omega/\gamma)m_{y} = [H_{0} + 4\pi M_{0} + 2M_{0}^{-1}(K_{4} - K_{2}) - D\partial^{2}/\partial z^{2}]m_{z},
$$

$$
(\Omega/\gamma)m_{z} = (H_{0} + 2M_{0}^{-1}K_{4} - D\partial^{2}/\partial z^{2})m_{y}. \qquad (2)
$$

Thus, the solutions to Eq. (1) (1) (1) are either bulklike standing waves (sin αz and cos αz) or surface states (sinh αz and $\cosh(\alpha z)$ similar to those introduced by Damon and Eshbach (DE) for $F_A = D = 0.18$ $F_A = D = 0.18$ The eigenfrequencies are

$$
\Omega^2 = \Omega_C^2 \pm D\alpha^2 [\gamma^2 (H_0 + 2K_4/M_0) + \Omega_C^2/(H_0 + 2K_4/M_0)]
$$

+ $\gamma^2 D^2 \alpha^4$, (3)

where $\chi_c^2 = \gamma^2 [H_0 + 2K_4 / M_0] [H_0 + 4\pi M_0 + 2M_0^{-1} (K_4 - K_2)]$ and the plus (minus) sign corresponds to the bulk- (surface-) like modes. From Eq. (3) (3) (3) , it can be shown that the four solutions are all of the DE type for $\Omega < \Omega_C$, whereas two of

FIG. 1. (Color online) Precession eigenstates for the two lowest even modes of a 71-nm-thick (Ga, Mn)As film, centered at $z=0$, using the values of the parameters D , K_2 , and K_4 listed in Table [I](#page-1-0) and $2K_S/DM_0$ =−0.33, 0, and 1 nm⁻¹. The in-plane external field H_0 =0.1 T is along [100].

them are surfacelike and the other two bulklike for $\Omega > \Omega_c$. It follows that the precessions for $\Omega > \Omega_c$ are generally of mixed bulk-surface character. The actual eigenmodes are selected by the conditions at the boundaries which, in turn, depend on the surface energy density F_S . Here, we take F_S $= K_S M_z^2 / M_0^2$ (same for the two surfaces; therefore, the solutions are either even or odd with respect to $z=0$) and use the Rado-Weertman condition $\mathbf{M} \times (\nabla_{\mathbf{M}} F_S - D M_0^{-1} \partial \mathbf{M} / \partial n) = 0$,¹⁹ giving $\partial m_v / \partial z = 0$ and $\partial \ln m_z / \partial z = \pm 2K_s / DM_0$ at $z = \pm L/2$; **n** is a unit vector normal to the layers. We note that these conditions cannot be met by a standing wave. Hence, they are incompatible with the well-known Kittel's pinning condition, $m_v = m_z = 0$,²⁰ which only applies if M_0 is not in plane. 21 It is important to emphasize the crucial role played by the surface in determining the slab's magnetization pattern and the precession eigenfrequencies. In Fig. [1,](#page-1-1) we show calculations for the two lowest even modes, S_0 and S_2 which, as shown below, are the only ones that are relevant to our experiments) of a film of width $L=71$ nm. These results illustrate the progression from almost-pinned to unpinned standing waves, as we move from large positive values of K_S to zero anisotropy, as well as the DE (bulk-surface)

FIG. 2. (Color online) Results at $T=4$ K for the as-grown sample with *L*=120 nm. The top panel shows Voigt-geometry DMK data (open circles) at H_0 =0.17 T. The applied field is parallel to $[100]$. The black curve is the linear prediction fit, which gives two modes of period of 99.7 and 83.9 ps. Their contributions to the fitted signal are shown, respectively, by the blue and red curves. The bottom panel shows the Fourier transform of the fit and, in the inset, the calculated m_z component for the three lowest eigenmodes, using the parameter values listed in Table [I.](#page-1-0)

character of S_0 (S_2) for K_S <0. The hybrid nature of the modes brings into question previous analyses of the available FMR data, based on Kittel's²² or the unpinned $(K_S=0)$ condition, 23 which led to the conclusion that the magnetization is not distributed uniformly along the direction of growth.

The experimental results are summarized in Figs. [2](#page-1-2) and [3.](#page-2-0) The top graph in Fig. [2](#page-1-2) shows DMK data at *T*=4 K for the as-grown 120-nm-thick sample after subtraction of an exponentially decaying background which reflects the spin relax-

TABLE I. Manganese concentration (x) , thickness (L) , and Curie temperature (T_C) of the $(Ga_{1-x}, Mn_x)As$ samples used in this work and values of the spin stiffness (D) , surface (K_S) , and bulk $(K_2$ and $K_4)$ anisotropy constants inferred from the pump-probe data. Unless noted, the films are as-grown specimens. The effective Mn-Mn exchange constant is given by $J_{\text{Mn}} = Dg\mu_B N_{\text{Mn}}^{2/3} / 2S_{\text{Mn}}$, where $g \sim 2$ is the gyromagnetic factor of $\text{Mn}(d^5)$, with spin $S_{\text{Mn}}=5/2$, and N_{Mn} is the magnetic ion density.

Sample	\mathcal{X}	L (nm)	T_C (K)	D $(T \text{ nm}^2)$	$2K_s/M_0$ $(T \text{ nm})$	$2K_4/M_0$ (T)	$4\pi M_0 - 2K_2/M_0$ (T)
1 ^a	0.03	120	80	25 ± 4	-2.4 ± 0.6	0.07 ± 0.02	0.44 ± 0.05
2	0.03	120	70	18 ± 3	-2.0 ± 0.5	0.07 ± 0.02	0.42 ± 0.05
3	$0.05 - 0.06$	71	65	5.8 ± 1	-1.3 ± 0.4	0.08 ± 0.02	0.50 ± 0.05
$\overline{4}$	$0.05 - 0.06$	47	65	5.8 ± 1	-1.1 ± 0.3	0.09 ± 0.02	0.50 ± 0.05
.5	$0.05 - 0.06$	25	65			0.09 ± 0.04	0.50 ± 0.05

^aSample annealed at 280 °C for 1 h.

FIG. 3. (Color online) Measured magnetic-field dependence of the precession mode frequencies (top panel) and the ratio A_2 / A_0 (bottom panel). For clarity, results for the films with $L=120$ nm are not shown. Solid curves are fits using the theoretical expressions discussed in the text. The FMR spectrum of the 71-nm-thick sample is shown in the inset.

ation of photoexcited electrons. 24 The oscillations are assigned to the precession of the magnetization around M_0 . We used linear prediction methods to fit the time-domain data and, as shown in the bottom panel of Fig. [2,](#page-1-2) the Fourier transform of the fit reveals two modes, which are assigned to S_0 and S_2 . With the exception of the thinnest film (L) = 25 nm), these two modes were observed in the DMK spectra of all the samples. The dependence of their frequencies on the applied magnetic field is shown in Fig. [3,](#page-2-0) together with FMR results for the *L*=71 nm slab. The bottom graph gives the dependence of the ratio between the amplitudes of the two oscillations, A_2 / A_0 . As we shall see, this ratio is critical to the determination of the magnetic parameters, particularly the surface constant K_S .

The coupling between magnetic precessions and probe pulses is controlled by the antisymmetric spin-flip Raman susceptibility, which is closely related to that for Faraday rotation.²⁵ Raman selection rules were found to be strictly obeyed in all cases. In our Voigt geometry, the scattered probe field associated with the m_v component of the precession is polarized along the *z* axis and, therefore, gives no DMK signal. Using results for scattering by coherent vibrations, $26,27$ $26,27$ we get for m_z scattering

$$
\left(\nabla^2 - \frac{n_R^2}{c^2} \frac{\partial^2}{\partial t^2}\right)(e_x \pm i e_y) = \frac{4\pi \chi_M}{c^2} \frac{\partial^2}{\partial t^2} [m_z(z, t)(e_y \pm i e_x)],\tag{4}
$$

where $\mathbf{e} = (e_x, e_y, 0)$ is the probe electric field, n_R is the refractive index, $\chi_M = \partial \chi^{(0)}/\partial M$, and $\chi^{(0)}$ is the linear

FIG. 4. (Color online) Calculated surface-anisotropy dependence of the frequency and the ratio between the amplitudes of the two lowest even modes for $L=71$ nm at $H_0=0$. Values for other parameters are listed in Table [I.](#page-1-0)

susceptibility. Let us define the average $\langle m_z(t) \rangle$ $=(1/L)\int_{-L/2}^{+L/2} m_z(z, t) dz$. To lowest order, and provided (i) $Ln_R/c \ll 2\pi/\Omega$, (ii) the light penetration depth is $\gg L$, and (iii) multiple reflections can be ignored (these provisos are well obeyed in our experiments), it can be shown that coherent scattering is equivalent to a slowly varying modulation of the refractive index,

$$
\delta n_R(t) = \pm (2\pi \chi_M/n_R) \langle m_z(t) \rangle, \tag{5}
$$

with different signs for the two senses of circular polarization. Except for the constant factors, this expression is identical to that describing FMR. Because $\langle m_z(t) \rangle$ vanishes for odd precessions, Eq. (5) (5) (5) supports our contention that the two modes observed in the experimental data are S_0 and S_2 .

In order to extract the magnetic parameters from the experimental data, we still need to ascertain the mechanism by which photoexcitation leads to precessions. Unlike the sharp selection rules of Raman type observed in probe scattering, we find that the strength of the oscillations is nearly the same for pump pulses of arbitrary circular or linear polarization. Since the Raman coupling for the pump operates only if the beams are circularly polarized, $11,12$ $11,12$ such a mechanism cannot be the cause of the precessions in our case.²⁸ Instead, our results point toward a relatively simpler thermal origin which, as first identified in ferromagnetic metals, 10 relies on the temperature dependence of the anisotropy. Following the model proposed by van Kampen *et al.*, [10](#page-3-27) we attribute the magnetic coherence to a (in the scale of the magnon period) sudden deviation of the orientation of M_0 due to the temperature rise that follows the absorption of the light pulse.²⁹ Since the equilibrium orientation, initially along $[100]$, moves toward $|110|$ after the pulse hits, and the light penetration depth is $\geq L$ so that the excitation is uniform across the film, this gives an effective coupling proportional to $\langle m_y \rangle$. Combining this result with Eq. (5) (5) (5) , the amplitude of the DMK signal for a given precession is then $\delta\theta \propto \langle m_y \rangle$ $\times \langle m_z \rangle / \langle m_y^2 \rangle$. In Fig. [4,](#page-2-2) we show the surface-anisotropy dependence of the frequencies of the two lowest even modes and of A_2/A_0 . Because the amplitude ratio depends strongly on K_S but is far less sensitive to the other magnetic constants,

the observation of S_2 is, in itself, a strong indication that K_{S} < 0.

The parameters in Table [I](#page-1-0) were obtained from fits to the experimental data using the theoretical expression for $\delta\theta$ and Eq. ([3](#page-0-1)). Our values for the bulk anisotropy are in fairly good agreement with FMR results.^{14,[22](#page-3-19)[,23](#page-3-20)} An estimate of the Mnhole exchange interaction J_{pd} can be obtained from theoretical expressions for the transition temperature, which give $J_{pd} \sim (6\pi^2)^{1/3} \sqrt{DN_{\rm Mn}^{1/3}} / (\chi_h S_{\rm Mn})$, where χ_h is the hole spin susceptibility.³⁰ Using results of Hall measurements at room temperature to determine the hole concentration and, thus, χ_h , we get J_{pd} ~ 0.63 eV nm³ for the as-grown sample with

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- ¹H. Ohno, A. Shen, F. Matsukura, A. Oiwa, A. Endo, S. Katsumoto, and Y. Iye, Appl. Phys. Lett. **69**, 363 (1996).
- ²H. Ohno, Science **281**, 951 (1998).
- 3S. Sonoda, S. Shimizu, T. Sasaki, Y. Yamamoto, and H. Hori, J. Cryst. Growth 237-239, 1358 (2002).
- ⁴K. Ando, Science 312, 1883 (2006).
- 5T. Omiya, F. Matsukura, T. Dietl, Y. Ohno, T. Sakon, M. Motokawa, and H. Ohno, Physica E (Amsterdam) 7, 976 (2000); J. König, T. Jungwirth, and A. H. MacDonald, Phys. Rev. B **64**, 184423 (2001).
- ⁶ J. König, H. H. Lin, and A. H. MacDonald, Phys. Rev. Lett. **84**, 5628 (2000).
- 7T. Dietl, H. Ohno, and F. Matsukura, Phys. Rev. B **63**, 195205 $(2001).$
- 8G. Bouzerar, J. Kudrnovský, L. Bergqvist, and P. Bruno, Phys. Rev. B 68, 081203(R) (2003).
- 9^9 G. Ju, A. V. Nurmikko, R. F. C. Farrow, R. F. Marks, M. J. Carey, and B. A. Gurney, Phys. Rev. Lett. **82**, 3705 (1999).
- 10M. van Kampen, C. Jozsa, J. T. Kohlhepp, P. LeClair, L. Lagae, W. J. M. de Jonge, and B. Koopmans, Phys. Rev. Lett. **88**, 227201 (2002).
- 11A. V. Kimel, A. Kirilyuk, P. A. Usachev, R. V. Pisarev, A. M. Balbashov, and Th. Rasing, Nature (London) 435, 655 (2005).
- 12F. Hansteen, A. Kimel, A. Kirilyuk, and T. Rasing, Phys. Rev. B 73, 014421 (2006).
- ¹³L. Landau and E. Lifshitz, Phys. Z. Sowjetunion 8, 153 (1935).
- 14X. Liu, Y. Sasaki, and J. K. Furdyna, Phys. Rev. B **67**, 205204 $(2003).$
- 15U. Welp, V. K. Vlasko-Vlasov, X. Liu, J. K. Furdyna, and T. Wojtowicz, Phys. Rev. Lett. 90, 167206 (2003).
- 16K.-Y. Wang, M. Sawicki, K. W. Edmonds, R. P. Campion, S. Maat, C. T. Foxon, B. L. Gallagher, and T. Dietl, Phys. Rev. Lett. 95, 217204 (2005).

L=120 nm. This value is significantly larger than other experimental estimates, based on analyses of transport measurements in the paramagnetic region, which are in the range $0.054 - 0.15$ $0.054 - 0.15$ $0.054 - 0.15$ eV nm³ (Ref. 5).

In summary, we observed coherent precessions of combined bulk and surface character and determined the exchange and anisotropy constants of (Ga,Mn)As using a comprehensive model of the coupling of the magnetization to light.

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- ¹⁷These conditions do not apply to highly conducting metals; see N. S. Almeida and D. L. Mills, Phys. Rev. B 53, 12232 (1996), and references therein.
- 18R. W. Damon and J. R. Eshbach, J. Phys. Chem. Solids **19**, 308 $(1961).$
- 19G. T. Rado and J. R. Weertman, J. Phys. Chem. Solids **11**, 315 $(1959).$
- ²⁰C. Kittel, Phys. Rev. **110**, 1295 (1958).
- ²¹ R. F. Soohoo, Phys. Rev. **131**, 594 (1963).
- 22T. G. Rappoport, P. Redlinski, X. Liu, G. Zaránd, J. K. Furdyna, and B. Jankó, Phys. Rev. B 69, 125213 (2004).
- 23S. T. B. Goennenwein, T. Graf, T. Wassner, M. S. Brandt, M. Stutzmann, J. B. Philipp, R. Gross, M. Krieger, K. Zürn, P. Ziemann, A. Koeder, S. Frank, W. Schoch, and A. Waag, Appl. Phys. Lett. **82**, 730 (2003).
- 24A. V. Kimel, G. V. Astakhov, G. M. Schott, A. Kirilyuk, D. R. Yakovlev, G. Karczewski, W. Ossau, G. Schmidt, L. W. Molenkamp, and Th. Rasing, Phys. Rev. Lett. 92, 237203 (2004).
- ²⁵W. Hayes and R. Loudon, *Scattering of Light by Crystals* (Wiley, New York, 1978).
- ²⁶ R. Merlin, Solid State Commun. **102**, 207 (1997).
- 27T. E. Stevens, J. Kuhl, and R. Merlin, Phys. Rev. B **65**, 144304 $(2002).$
- 28 Recent ultrafast studies of orthoferrites and garnets (Refs. [11](#page-3-25) and [12](#page-3-9)) confirm the general theoretical prediction (Ref. [27](#page-3-24)) that the Raman mechanism (also referred to as the inverse Faraday effect in the case of magnons) is the dominant source of coherence in transparent media.
- ²⁹ An estimate based on the available data (Ref. [15](#page-3-12)) gives \sim 150 μ K for the change required at 4 K to shift this angle by 1 μ rad. These values are consistent with the parameters of our experiments and the thermal properties of low-temperature-grown GaAs.
- 30T. Jungwirth, J. Sinova, J. Mašek, J. Kučera, and A. H. Mac-Donald, Rev. Mod. Phys. **78**, 809 (2006).