Linear response theory and the universal nature of the magnetic excitation spectrum of the cuprates

M. R. Norman

Materials Science Division, Argonne National Laboratory, Argonne, Illinois 60439, USA (Received 29 January 2007; revised manuscript received 20 March 2007; published 16 May 2007)

Linear response theory, commonly known as the random-phase approximation (RPA), predicts a rich magnetic excitation spectrum for *d*-wave superconductors. Many of the features predicted by such calculations appear to be reflected in inelastic neutron-scattering data of the cuprates. In this paper, I will present results from RPA calculations whose input is based on angle-resolved photoemission data, and discuss possible relevance to inelastic neutron-scattering data of $La_{2-x}Sr_xCuO_4$ (LSCO), YBa₂Cu₃O_{6+x} (YBCO), and Bi₂Sr₂CaCu₂O_{8+x} (Bi2212) in their superconducting and nonsuperconducting phases. In particular, the question of the universality of the magnetic excitation spectrum will be addressed.

DOI: [10.1103/PhysRevB.75.184514](http://dx.doi.org/10.1103/PhysRevB.75.184514)

PACS number(s): 74.25.Ha, 25.40.Fq, 74.25.Jb, 74.72.-h

I. INTRODUCTION

The nature of the magnetic excitation spectrum of the cuprates is an important topic in its own right. However, it has many implications beyond this as well. It is of special relevance to those theories that propose a "magnetic" origin for cuprate superconductivity.¹ And it is of particular interest to the current debate whether magnetic excitations or phonons are responsible for certain strong-coupling features observed in tunneling, infrared conductivity, and angle-resolved photoemission spectra (ARPES).^{2-[4](#page-7-2)}

Although there are a large number of theories for explaining the observed magnetic excitation spectrum, to a first approximation, these theories can be collapsed into two groups. The first are linear response calculations based on a twodimensional (2D) Fermi surface and the presence of a *d*-wave energy gap. In this group are included calculations which go beyond random-phase approximation (RPA) [such as FLEX $(Ref. 5)$ $(Ref. 5)$ $(Ref. 5)$] and the generalization to the particleparticle channel of the $SO(5)$ approach.⁶ The other group are those based on coupled spin ladders and related approaches. $7-11$ They assume phase segregation of the material into undoped antiferromagnetic domains separated by one-dimensional "stripes" containing the doped holes. They have the advantage of treating the full quantum-mechanical nature of the spins but have the disadvantage that fermionic degrees of freedom are neglected.

It is not the intent here to discuss the relative merits of one approach versus the other. In some sense, both theories are different limits of a more complex theory which properly treats spin and charge degrees of freedom. Rather, I wish to discuss what success linear response calculations have in regard to describing magnetic excitation properties of the cuprates, but as well discuss their limitations. In particular, I wish to address the question of whether there is a "universal" behavior of the spin excitation spectrum based on the RPA results. Ultimately, I hope these results will have some relevance to a final understanding of the magnetic behavior of these fascinating materials.

II. METHODOLOGY

The RPA expression for the interacting susceptibility is

$$
\chi(q,\omega) = \frac{\chi_0(q,\omega)}{1 - U\chi_0(q,\omega)}.\tag{1}
$$

In this equation, χ_0 is the polarization bubble constructed from bare Green's functions, and *U* is the effective Hubbard interaction which results from projection onto the single band subspace (that is, the band that is the antibonding combination of copper $d_{x^2-y^2}$ and oxygen p_x and p_y orbitals¹²).

It is often stated that RPA is a weak-coupling approach, applicable only to heavily overdoped materials which exhibit Fermi-liquid-like behavior.¹³ However, it was shown by Schrieffer *et al.*^{[14](#page-7-9)} that RPA reproduces the superexchange \overline{J} and the resulting spin-wave dispersion of the undoped material. Obviously, there are important quantum corrections to these results, but these do not impact the overall correctness of the approach. One important point to remark is that the calculation of Ref. [14](#page-7-9) was self-consistent in the sense that the gapless nature of the spin excitations at $q = (\pi, \pi)$ is a direct consequence of the mean-field equation for the Hubbard gap. Such self-consistency is difficult to implement in the doped case because of the added complication of having to calculate the screening caused by the doped holes.

There are two ways to proceed for the doped case. Continue to use bare Green's functions and replace *U* by an effective U_{eff} which accounts for screening.^{15[,16](#page-7-11)} Or take the infinite *U* limit, replace the Green's functions by those of the t-J model, and replace U by J_{eff} , where J_{eff} takes into account the reduction of *J* by the introduction of doped holes in the t-J approach, one is forced to scale *J* down to prevent a magnetic instability from occurring¹⁷). An important additional remark is that the screened interaction is now dependent on transferred momentum *q*. This is particularly clear in the t-J model, where $J(q) = -J[\cos(q_x a) + \cos(q_y a)]/2$ due to exchange between near-neighbor copper sites.

One limitation of the RPA approach is the use of bare Green's functions as opposed to dressed ones. However, it is well known that using dressed Green's functions means that vertex corrections must be included.¹⁶ The neglect of such vertex corrections typically leads to worse results than the RPA.

The approach taken in this paper will be somewhat more phenomenological. The bare Green's functions will be based

on dispersions derived from angle-resolved photoemission. The resulting *Ueff* will be taken as an adjustable constant, designed to reproduce the location of the *S*= 1 resonance mode at $q = (\pi, \pi)$. Calculations have been performed as well assuming that U_{eff} has the same *q* dependence as $J(q)$. In a more complete theory, one might work instead within the context of dynamical mean-field theory and its cluster generalization.¹⁸

The heart of the RPA calculation turns out to be the structure of the bare polarization bubble χ_0 . For a superconductor, this is 19

$$
\chi_0(q,\omega) = \sum_{k} \left[\frac{1}{2} \left(1 + \frac{\epsilon_k \epsilon_{k+q} + \Delta_k \Delta_{k+q}}{E_k E_{k+q}} \right) \frac{f(E_{k+q}) - f(E_k)}{\omega - (E_{k+q} - E_k) + i\delta} + \frac{1}{4} \left(1 - \frac{\epsilon_k \epsilon_{k+q} + \Delta_k \Delta_{k+q}}{E_k E_{k+q}} \right) \frac{1 - f(E_{k+q}) - f(E_k)}{\omega + (E_{k+q} + E_k) + i\delta} + \frac{1}{4} \left(1 - \frac{\epsilon_k \epsilon_{k+q} + \Delta_k \Delta_{k+q}}{E_k E_{k+q}} \right) \frac{f(E_{k+q}) + f(E_k) - 1}{\omega - (E_{k+q} + E_k) + i\delta} \right],
$$
\n(2)

where $E_k = \sqrt{\epsilon_k^2 + \Delta_k^2}$. The coherence factors play a critical role in the results.²⁰ For an s -wave superconductor, the coherence factor [in term 3 of Eq. (2) (2) (2)] vanishes on the Fermi surface. The net result is that there are no spin collective modes in this case. On the other hand, for an order parameter satisfying the condition $\Delta(k+q) = -\Delta(k)$, the coherence factor becomes maximal on the Fermi surface (equal to 2). As a consequence, the imaginary part of the bubble has a step discontinuity at threshold, and thus the real part has a logarithmic divergence. This divergence guarantees the appearance of a collective mode below threshold. This collectivemode energy at $q = (\pi, \pi)$ will be denoted as the resonance energy.

There is a common misconception that the appearance of the collective mode requires the presence of a *d*-wave superconducting gap. In fact, the only requirement is that there is a gap and the coherence factors do not vanish on the Fermi surface. As an example, imagine that the pseudogap phase is a phase disordered version of a superconductor. To a first approximation, we can then set $\langle \Delta_k \Delta_{k+q} \rangle = 0.^{21}$ $\langle \Delta_k \Delta_{k+q} \rangle = 0.^{21}$ $\langle \Delta_k \Delta_{k+q} \rangle = 0.^{21}$ In this case, the coherence factor is now unity on the Fermi surface independent of the value of *k*, and one finds a collective mode below threshold as in the case of a *d*-wave superconductor. As the coherence factor is now half that of the superconductor, then larger values of *Ueff* are needed to obtain similar collective-mode energies. In this paper, a *d*-wave order parameter of the form $\Delta(k) = \Delta_0 [\cos(k_x a) - \cos(k_y a)]/2$ is assumed, although for some calculations the phase disordered approximation was invoked as well.

Finally, some technical remarks are in order. Equation ([2](#page-1-0)) is solved by replacing δ by some finite Γ and doing a simple sum over the Brillouin zone. The smaller Γ is, the more *k* points are needed in the sum. Results have been generated for Γ ranging between 0.1 and 2 meV, with zone meshes ranging from 400×400 to 4000×4000 ; but in this paper, results are shown for $\Gamma = 2$ meV (400 × 400 mesh). All calculations were performed for a temperature of 1 meV. All

FIG. 1. Re χ_0 in the normal state at zero energy along the symmetry axes of the zone for (a) tb1, (b) tb2, (c) tb3, and (d) tb4. For all figures, the zone notation is in π units and the susceptibility is in units of states per eV/f.u.

susceptibilities quoted here are states per eV per $CuO₂$ formula unit and should be multiplied by the matrix element $\sum_{\sigma} g^2 \mu_B^2 / \sigma |S_z| \sigma^2 = 2\mu_B^2$ to compare to neutron-scattering data.

III. RESULTS: NORMAL STATE

The most important input into the RPA calculations is the assumed form of the fermionic dispersion. For instance, a well studied case is a tight-binding model with just nearneighbor hopping[.6](#page-7-4) When considering the response at $q=(\pi,\pi)$, the threshold for the polarization bubble corresponds to exciting from the node of the *d*-wave order parameter, k_N , to an unoccupied state, $k_N + q$, along the zone diagonal. As such, the threshold is independent of the *d*-wave energy gap. The Fermi surfaces observed by angle-resolved photoemission, though, differ considerably from those predicted by a near-neighbor-tight binding model. In particular, the observed Fermi surfaces are characterized by "hot spots," that is, points on the Fermi surface which satisfy the condition $\epsilon_k = \epsilon_{k+q} = 0$. In this case, the threshold of the bubble at $q = (\pi, \pi)$ is determined by twice the energy gap at the hot spots.

When considering the response for a general *q*, there is even more sensitivity to the assumed fermionic dispersion. To illustrate this point, in Fig. [1,](#page-1-1) I show the real part of the normal state $(\Delta = 0)$ bubble at $\omega = 0$ for four different dispersions along the three symmetry axes of the 2D Brillouin zone a summary of the properties of these four tight-binding dis-persions is listed in Table [I](#page-2-0)). The first dispersion $(tb1)$ is that based on an early tight-binding fit to ARPES spectra on Bi22[1](#page-1-1)2.²² Two prominent features are observed [Fig. 1(a)]. First, there is a broad maximum around $q = (\pi, \pi)$. Second, there are incommensurate peaks along the bond and diagonal directions which correspond to scattering between the antin-

TABLE I. Tight-binding dispersions based on angle-resolved photoemission data. The first four columns list the coefficient c_i of each term (eV), that is, $\epsilon(\vec{k}) = \sum c_i \eta_i(\vec{k})$. The last column lists the basis functions (the lattice constant a is set to unity). Below this are several properties of these dispersions: the node (k_N, k_N) , the antinode $(1, k_{AN})$ (in π/a units), the nodal velocity v_N , the antinodal velocity v_{AN} (in eV Å), the energy at $(\pi, 0)$, ϵ_M (eV), the bandwidth W (eV), and the value of U needed for resonance at 40 meV in the superconducting state.

tb1	tb ₂	tb ₃	tb4	$2\eta_i(\vec{k})$
0.1305	0.1960	0.1197	0.0801	$\overline{2}$
-0.5951		-0.6798 -0.5881 -0.7823		$\cos k_x + \cos k_y$
0.1636	0.2368	0.1461	0.0740	2 cos k_x cos k_y
-0.0519	-0.0794		$0.0095 -0.0587$	$\cos 2k_x + \cos 2k_y$
-0.1117				0.0343 -0.1298 -0.1398 $\cos 2k_x \cos k_y + \cos k_x \cos 2k_y$
0.0510	0.0011		$0.0069 - 0.0174$	2 cos $2k_x \cos 2k_y$
0.365	0.365	0.414	0.440	k_N
0.180	0.180	0.100	0.180	k_{AN}
1.63	1.50	1.14	1.77	v_N
0.48	1.50	0.24	0.91	v_{AN}
-0.034	-0.119	-0.010	-0.070	ϵ_M
1.414	1.291	1.436	1.844	W
0.274	0.563	0.214	0.332	U

odal sections of the Fermi surface near the $(\pi, 0)$ points. As can be seen from the [2](#page-2-1)D plot in Fig. $2(a)$, this scattering forms a square box around the zone center, with maximum intensity at the box corners. These peaks in the zero Γ limit are logarithmically divergent.²³ Although potentially important for the charge response, they are probably not so for the spin response [since $J(q)$ has opposite sign in this region of the zone].

FIG. 2. (Color online) 2D zone plots of Fig. [1.](#page-1-1) The dynamic range is the same as in Fig. [1.](#page-1-1)

The other noticeable feature of Fig. $1(a)$ $1(a)$ is the broad maximum around $q=(\pi,\pi)$. As a consequence, in the gapped case, the dominant behavior below the resonance energy is commensurate. $24,25$ $24,25$ However, this behavior is actually the exception rather than the rule. At this point, a discussion on how the fermionic dispersion is chosen is in order.²² A fit is performed using a six parameter tight-binding model which includes real-space lattice vectors of types $(0,0)$, $(1,0)$, $(1,1)$, $(2,0)$, $(2,1)$, and $(2,2)$. In the original analysis, the fitting variables were (1) the position of the Fermi surface along $(0,0)$ – (π, π) (the node), (2) the Fermi velocity at the node, (3) the Fermi surface along $(\pi, 0) - (\pi, \pi)$ (the antinode), (4) the energy at $(\pi, 0)$, (5) the curvature of the dispersion at $(\pi, 0)$ (along the k_x direction), and (6) the energy at (π, π) (this last condition is simply for stability of the fit). The behavior around $q = (\pi, \pi)$ in Fig. [1](#page-1-1)(a) is largely a consequence of the fact that the saddle-point energy at $(\pi, 0)$, −34 meV, is close to the Fermi energy. For dispersions where this is displaced deeper in energy, a minimum is found in Re χ_0 at $q=(\pi,\pi)$ rather than a maximum.

To illustrate this, a tight-binding analysis has been performed on the latest high-resolution ARPES measurements on optimal doped Bi2212. 26 These data reveal an isotropic Fermi velocity with a value of 1.5 eV Å. Fixing this condition forces the energy at $(\pi, 0)$ to lie at -119 meV, as compared to -34 meV in the earlier fit.²⁷ This difference (dispersion tb2) has a profound effect on the real part of the bubble, as can be seen in Fig. $1(b)$ $1(b)$. Now, incommensurate peaks are observed around $q = (\pi, \pi)$ as well, and as can be seen in Fig. $2(b)$ $2(b)$, they form a diamond-shaped pattern. This structure is actually discernible in Fig. $1(a)$ $1(a)$ as kinks, but it is obscured by the dominant maximum around $q = (\pi, \pi)$. The behavior of Fig. $1(b)$ $1(b)$ is also found for the t, t' dispersion used in Ref. [28,](#page-8-0) as the $(\pi, 0)$ point in that case is at -129 meV.

The incommensurate effects can also be strengthened by flattening the Fermi surface around the node.^{17[,24](#page-7-19)} A dispersion which incorporates this effect, as well as the presence of a flat quasiparticle band near $(\pi, 0)$, has the Fermi surface illustrated in Fig. $3(c)$ $3(c)$ (this is dispersion tb3 of Ref. [25](#page-7-20) and is denoted as tb3 here as well). In this case, the effect of the enhanced incommensurability and the tendency to have a maximum at $q = (\pi, \pi)$ due to the small energy of the $(\pi, 0)$ point (-10 meV) approximately balance, leading to a net flat behavior around $q = (\pi, \pi)$, as can be seen in Fig. [1](#page-1-1)(c).

One issue is that there is not much inelastic neutronscattering (INS) data on Bi2212 because of the small sample sizes, though this is starting to change.²⁹ A large amount of data is present for YBCO, but the electronic structure of this material is complicated because of its strong orthorhombicity and the presence of metallic chain layers. Instead, I consider LSCO. This material has the advantage of having extensive inelastic neutron-scattering data and a simpler electronic structure (no bilayer splitting). Recently, Tranquada et al.^{[30](#page-8-2)} have pointed out that the inelastic magnetic response of nonsuperconducting $La_{1.875}Ba_{0.125}CuO₄$ (LBCO) is very similar to that of superconducting YBCO, 31 which makes this material of particular interest. Extensive ARPES data exist on LSCO. Recently, an in depth study was done on underdoped LSCO (Ref. [32](#page-8-4)) and revealed a quite striking Fermi surface

FIG. 3. Fermi surfaces corresponding to Fig. [1.](#page-1-1)

[Fig. $3(d)$ $3(d)$]. The Fermi surface is characterized by two straight sections, one centered about the node, the other about the antinode. This Fermi surface, and the observed dispersion (denoted as tb4), is fitted using the condition (1) node at $(0.44, 0.44) \pi$, (2) velocity at the node $(1.8 \text{ eV} \text{Å})$, (3) curvature of the Fermi surface at the node (zero), (4) antinode at $(1, 0.18)$ π , (5) $(\pi, 0)$ energy at −70 meV, and (6) (π, π) energy at 1 eV (for fit stability). The result for the zero-frequency real response is shown in Figs. $1(d)$ $1(d)$ and $2(d)$ $2(d)$. Again, note the strong "boxlike" structure around the zone center, and the pronounced incommensurate behavior around (π, π) due to the flat Fermi surface near the node. In this context, it should be noted that the boxlike structure near *q* $(0,0)$ is due to nesting of the antinodal regions, but the diamondlike structure near $q = (\pi, \pi)$ is due to nesting of the nodal regions. This is in contrast with the alternate stripes model, where the incommensurability around these two *q* vectors has the same origin.

A number of other tight-binding dispersions besides the four discussed here have been analyzed as well, including some that account for bilayer splitting.³³ The results fall within the range of behavior discussed for the four dispersions here. In addition, bilayer splitting leads to some extra details in the even (optic) spin response due to differing momentum locations of features from bonding-bonding and antibonding-antibonding responses. 34 For simplicity, these details are not discussed here as they do not occur in the dominant odd (acoustic) spin channel which involves only the bonding-antibonding response.

IV. RESULTS: SUPERCONDUCTING STATE

The presence of a *d*-wave energy gap leads to a polarization bubble which now becomes gapped for all *q* vectors in the first zone but two (these vectors being the ones that connect the nodes of the *d*-wave order parameter). Because of

FIG. 4. (Color online) Particle-hole continuum edge along the symmetry axes of the zone. The maximum *d*-wave gap is 30 meV. The dashed line in (d) corresponds to the process involving k_N and k_N+q .

the strong anisotropy of the *d*-wave gap, the lower edge of the particle-hole continuum also has a strong *q* dependence. This continuum edge is plotted along the zone symmetry axes in Fig. [4](#page-3-1) for the four dispersions studied here. The threshold at (π, π) corresponds to twice the energy gap at the hot spots. As one moves away from this wave vector, the threshold splits^{28,[35](#page-8-7)} into several ones: two for \vec{q} along (π, x) and three for \vec{q} along (x, x) . Along (x, x) , the zero threshold at $\vec{q}_N = 2\vec{k}_N$ obviously corresponds to two nodal points, the analogous minimum along (π, x) involving only one nodal point [the other being at $\vec{k}_N + (\pi, x)$].

As briefly discussed above, the effect of an energy gap has a profound impact on the polarization bubble. The presence of an energy gap, along with the fact that the coherence factors at threshold do not generally vanish in the *d*-wave case, leads to a step jump in the imaginary part of the bubble at threshold. By Kramers-Kronig, such a step jump causes a logarithmic divergence in the real part of the bubble at threshold. As a consequence, any finite value of *U* will lead to the presence of a collective mode below threshold. The exception is in those cases where Δ_k and Δ_{k+q} have the same sign, which in the *d*-wave case can occur for *q* vectors significantly displaced from (π, π) . In addition, for a pole to occur, then U must have a positive sign (this is relevant to near-neighbor exchange models, where the positive sign only occurs for $q_x+q_y > \pi$).

The value of U needed to have a pole at $q = (\pi, \pi)$ at a particular energy varies significantly among the dispersions looked at here (the values for 40 meV are also listed in Table [I](#page-2-0)). This is a combination of the different Fermi surfaces shown in Fig. [3,](#page-3-0) the different energies of the van Hove singularity at $(\pi, 0)$ for these dispersions, and the different overall bandwidths. This variation of *U* involves a rather subtle issue concerning whether one uses a bare dispersion based LINEAR RESPONSE THEORY AND THE UNIVERSAL...

FIG. 5. (Color online) 2D zone plots of Im χ for tb1 at (a) 10 meV, (b) 30 meV, (c) 40 meV, and (d) 50 meV. The maximum d -wave gap is 30 meV, and U (0.274 eV) is chosen to give a resonance at 40 meV. Scale is such that the lowest is 0 and the highest are (a) 0.7, (b) 4.4, (c) 112, and (d) 21.3.

on band theory, or a renormalized (quasiparticle) dispersion, for ϵ_k in Eq. ([2](#page-1-0)) (and also the issue of bilayer splitting). Again, the purpose here is to simply look at a range of dispersions and contrast their behaviors.

Figures [5–](#page-4-0)[8](#page-4-1) show the variation of the imaginary part of the RPA susceptibility $[Eq. (1)]$ $[Eq. (1)]$ $[Eq. (1)]$ with U tuned to yield a pole at $q=(\pi,\pi)$ at 40 meV (50 meV for dispersions tb2 and tb4-the reason for this difference will be discussed below).

FIG. 6. (Color online) 2D zone plots of Im χ for tb2 at (a) 10 meV, (b) 30 meV, (c) 50 meV, and (d) 100 meV. The maximum d -wave gap is 30 meV, and U (0.508 eV) is chosen to give a resonance at 50 meV. Scale is such that the lowest is 0 and the highest are (a) 6.3, (b) 20.6, (c) 22.6, and (d) 7.0.

FIG. 7. (Color online) 2D zone plots of Im χ for tb3 at (a) 10 meV, (b) 30 meV, (c) 40 meV, and (d) 50 meV. The maximum d -wave gap is 30 meV, and $U(0.214 \text{ eV})$ is chosen to give a resonance at 40 meV. Scale is such that the lowest is 0 and the highest are (a) 4.0, (b) 101, (c) 112, and (d) 10.6.

U was treated as a constant, though calculations were also performed for $U = J(q)$. A maximum energy gap of 30 meV was assumed.

Figure [5](#page-4-0) shows results for tb1. At low energies, one sees weak intensity near \vec{q}_N corresponding to the nodes. As the energy is raised to exceed the minimum threshold along (π, x) , this incommensurate pattern rotates 45°. The preference for this rotated pattern was discussed by Schulz $17,36$ $17,36$ and

FIG. 8. (Color online) 2D zone plots of Im χ for tb4 at (a) 10 meV, (b) 30 meV, (c) 50 meV, and (d) 100 meV. The maximum d -wave gap is 30 meV, and $U(0.276 \text{ eV})$ is chosen to give a resonance at 50 meV. Scale is such that the lowest is 0 and the highest are (a) 1.4, (b) 10.6, (c) 27.1, and (d) 5.5.

FIG. 9. (Color online) Fermi surface for tb4 (dashed line) and its *q* translated image (solid line) for (a) $q=(1, 0.76)\pi$ and (b) $q = (0.88, 0.88) \pi.$

is a consequence of the fact that translation by (x, x) only brings the Fermi surface into coincidence in one quadrant of the zone, but translation by (π, x) brings it into coincidence in two quadrants (Fig. [9](#page-5-0)). In addition, though, one sees a global maximum at $q = (\pi, \pi)$ which is a consequence of the global maximum seen in the normal-state response at zero energy discussed earlier. Above the resonance energy (40 meV), the response becomes incommensurate, with the displacement of the maxima away from (π, π) increasing in magnitude (but the response decreasing in strength) with increasing energy. For a range of energies above resonance $(\sim 12 \text{ meV})$, the incommensurate response corresponds to damped poles, and mostly has maxima along (π, x) , but for some energies the maxima are along (x, x) . This is due to the anisotropy of the splitting of the particle-hole threshold as one moves away from (π, π) . The overall response is best appreciated by plotting the intensity as a function of energy along the (x, x) direction in momentum space as shown in Fig. $10(a)$ $10(a)$. In the pole region $(40-52 \text{ meV})$, one sees a magnonlike (quadratic) dispersion. Above this region, the response rapidly loses strength and the dispersion becomes more steep.

I now contrast this behavior with that from the next dispersion, tb2. This dispersion, based on ARPES fits to the bonding Fermi surface in Bi2212, has (1) an isotropic velocity around the Fermi surface and (2) an energy for the $(\pi, 0)$ point which is much deeper than for tb1 $(-119 \text{ meV}$ as compared to -34 meV). As a consequence, this dispersion yields a bubble whose real part has a relatively weak momentum dependence, and is very typical of many of the dispersions used in the theoretical literature. In fact, the momentum dependence is so weak that with a constant *U*, the condition to yield a resonance at 40 meV at (π, π) implies the presence of long-range order (that is, there is a q value at zero energy where the real part of the bubble exceeds $1/U$). This forces us to move the resonance condition close to the edge of the continuum (where the real part of the bubble has a peak) in order to avoid this problem. And obviously, because of the weak momentum dependence, the value of *U* needed to obtain a resonance condition is significantly larger than for tb1. The results are plotted in Figs. 6 and $10(b)$ $10(b)$. Below resonance, despite the lack of a commensurate response, the incommensurate response is similar to tb1, being dominated by node-node processes at low energies, and then rotating to a

FIG. 10. (Color online) Plots of Im χ versus energy and $(x, x)\pi$ for (a) $tb1$, (b) $tb2$, (c) $tb3$, and (d) $tb4$. Same conditions as Figs. [5–](#page-4-0)[8.](#page-4-1) The intensity is on a logarithmic grid ranging from 0.1 to 100.

bond centered response once the minimum of the continuum edge along (π, x) is exceeded. The response becomes commensurate at resonance (50 meV), then one has a weak incommensurate response above resonance whose maxima generally sit along the diagonals. Although this 45° rotation of the incommensurate maxima is indeed what is seen in YBCO (Ref. [31](#page-8-3)) and LBCO, 30 the incommensurate response above resonance still has an overall diamond shape to it unlike the experiment. This shape is a consequence of $2k_F$ scattering which can be seen as well in the normal state [Fig. $2(b)$ $2(b)$]. The overall behavior is best appreciated in Fig. $10(b)$ $10(b)$, where in contrast to dispersion tb1, a very clear "reverse magnon" dispersion is evident for the resonance mode. Above resonance, the incommensurate response is almost dispersionless and seems to follow the steep particle-hole response involving scattering along the zone diagonal [dashed red line in Fig. $4(d)$ $4(d)$]. This steepness is due to the high Fermi velocity along the nodal direction and unlike the rest of the continuum response, it is independent of the energy gap.

The next dispersion, tb3, had been designed to enhance the incommensurability by flattening the Fermi surface around the node. In Fig. [7,](#page-4-3) one can see the strong diamondshaped incommensurate pattern below resonance that is a result of the enhanced Fermi surface nesting in the nodal region. The evolution of the pattern is again a diagonal response at very low energies due to node-node processes, a rotation by 45° to a bond centered pattern once the threshold along (π, x) is exceeded, commensurability at resonance (40 meV), and then again a complicated incommensurate response above resonance due to the anisotropic splitting of the continuum edge as one moves away from (π, π) . The overall dispersion of the magnetic response is most visible in Fig.

FIG. 11. (Color online) Plots of Im χ versus energy and $(x, x)\pi$ for tb4 for (a) superconducting state with a resonance at 40 meV $(U=0.332$ eV) and (b) pseudogap state with a resonance at 50 meV $(U=0.343$ eV). The intensity is on a logarithmic grid ranging from 0.1 to 100.

[10](#page-5-1)(c). Note the pronounced downward dispersion of the resonance which represents a true pole in the RPA susceptibility in this case (this response is below the continuum edge). One then sees a weak intensity gap near the nodenode vector due to crossing into the continuum (the so-called silent band³⁷). Beyond this, a new pole appears²⁸ on the other side of the continuum (the edge of which is again controlled by the high Fermi velocity along the nodal direction). One then reenters the continuum, and the response rapidly loses intensity and the damped polelike response is lost above 55 meV.

I now turn to the last dispersion, tb4. This is based on a fit to underdoped LSCO ARPES data.³² The Fermi surface in this case is characterized by strong nesting, both in the nodal region and also the antinodal one. Putting the resonance condition at 40 meV yields the same reversed magnon dispersion, silent band effect around q_N (node-node vector), and second mode behavior for $q < q_N$ that was so prominent for tb3 [Fig. $11(a)$ $11(a)$]. The one contrast is at higher energies, the incommensurability is much better defined than for the other dispersions due to the strong nesting [this high-energy dispersion follows the steep dispersion along the nodal direction shown in Fig. $4(d)$ $4(d)$].

However, for most of the results presented here, I choose to show instead the case where the resonance condition is at 50 meV. This is interesting for two reasons: (1) this energy is where the maximum response was seen in LBCO, 30 and (2) this energy corresponds to the continuum edge at (π, π) (and thus where there is a maximum in the real part of the bubble). U in this case is adjusted so that it is equal to the inverse of this maximum value.

The results are shown in Figs. 8 and $10(d)$ $10(d)$. In this case (as well as when the resonance is at 40 meV), the incommensurate response below resonance is much cleaner than for the other dispersions, with very clear maxima along the (π, x) direction once the continuum edge energy along this direction is exceeded, with x at the lowest energies being given by the condition $\vec{k}_N = (\pi/2 - x/4, \pi/2 - x/4)$. One again sees a clear rotation of the pattern by 45° as one crosses through the commensurate resonance energy, but as with tb2, the response above resonance has an overall diamond shape. Unlike the other dispersions, a well defined incommensurate

FIG. 12. (Color online) Plots of Im χ versus energy and $(x, x)\pi$ for tb4 for (a) superconducting state and (b) normal state. *U* is 0.276 eV. The intensity is on a logarithmic grid ranging from 0.02 to 20.

pattern with maxima along the diagonal direction persists to much higher energies, remaining clearly visible up to 140 meV. Above this energy, the response is less well defined but still incommensurate. Note the very well defined incommensurability at 100 meV, which is similar to that of the real part of the bubble (indicating polelike behavior). The interesting point about this dispersion, as commented above, is that the high-energy response typically follows the outer branch of the dispersive response corresponding to scattering processes along the nodal direction [dashed red line in Fig. $4(d)$ $4(d)$]. This nodal velocity is set by the renormalized Fermi velocity along the node, which in RVB-type models is proportional to the superexchange *J*. This velocity is also essentially constant with doping, 38 suggesting a connection between the universal behavior observed in both ARPES and INS data.

For this dispersion, calculations have also been performed in the pseudogap approximation $[\Delta_k \Delta_{k+q} = 0$ in Eq. ([2](#page-1-0))]. This is shown in Fig. $11(b)$ $11(b)$, and as can be seen, are virtually identical to the superconducting results shown in Fig. $10(d)$ $10(d)$. This indicates that the *d*-wave phasing relation is not necessary to obtain the results shown in this paper (just an energy gap—though we again note that for the *s*-wave superconducting case, there is no resonance effect). In this connection, most of the dramatic findings in INS data are in the underdoped regime where a pseudogap persists to very high temperatures.

It is also amusing to present results for this dispersion when the energy gap is set to zero. This is shown in Fig. $12(b)$ $12(b)$ for the same *U* as used in the superconducting case in Fig. $10(d)$ $10(d)$. The response is always incommensurate and is bond oriented for $\omega < \omega_c$ and diagonal oriented for $\omega > \omega_c$ (Fig. [13](#page-7-22)). ω_c is the energy where the constant energy contour for the dispersion switches from being holelike to electronlike. Besides the lack of a commensurate resonance response, and a spin gap due to the energy gap, there are interesting similarities with the superconducting results.

V. CONCLUSIONS

In conclusion, RPA calculations show a remarkably rich behavior, in particular, in the evolution of the momentum response as a function of energy. This behavior becomes

FIG. 13. (Color online) 2D zone plots of Im χ for tb4 in the normal state at (a) 10 meV, (b) 30 meV, (c) 50 meV, and (d) 100 meV. *U* is 0.276 eV. Scale is such that the lowest is 0 and the highest are (a) 4.8 , (b) 6.8 , (c) 7.2 , and (d) 3.8 .

even richer in the superconducting and pseudogap phases due to the strong momentum anisotropy of the particle-hole continuum edge due to the anisotropy of the energy gap. There is a striking similarity of many of these findings to experimental INS data in LSCO, YBCO, and Bi2212, in particular, the reversed magnon behavior of the resonance mode, and the finding for several dispersions of a rotation by 45° of the incommensurate response when passing through the resonance energy, which argue for universality in the magnetic response as has been commented on in regard to experimental INS data.

On the other hand, there are some differences that need to be kept in mind. For instance, the incommensurate response above resonance is generally more variable in the calculations than indicated by experiment. The fact that this is apparent as well in calculations where the bare Green's functions are replaced by experimental Green's functions³⁹ indicates that this is a general issue. The author speculates that the RPA calculations and alternate ones based on coupled spin ladders are different limits of a more complete theory that properly includes the full quantum-mechanical nature of both the spin and charge degrees of freedom. The development of such a theory should help shed more light on the relation of the magnetic fluctuation spectrum and the existence of *d*-wave superconductivity.

ACKNOWLEDGMENTS

I would like to thank Andrey Chubukov, Steve Hayden, and John Tranquada for discussions, and a special thanks to Herb Mook for his encouragement to initiate this project back in 1999. The work was supported by the U.S. Department of Energy, Office of Science, under Contract No. DE-AC02-06CH11357.

- 1A. V. Chubukov, D. Pines, and J. Schmalian, *The Physics of Superconductors*, edited by K. H. Bennemann and J. B. Ketterson (Springer, Berlin, 2003), Vol. 1, p. 495.
- 2A. Damascelli, Z. Hussain, and Z.-X. Shen, Rev. Mod. Phys. **75**, 473 (2003).
- ³D. N. Basov and T. Timusk, Rev. Mod. Phys. **77**, 721 (2005).
- ⁴ J. Lee, K. Fujita, K. McElroy, J. A. Slezak, M. Wang, Y. Aiura, H. Bando, M. Ishikado, T. Matsui, J.-X. Zhu, A. V. Balatsky, H. Eisaki, S. Uchida, and J. C. Davis, Nature (London) 442, 546 $(2006).$
- ⁵N. E. Bickers, D. J. Scalapino, and S. R. White, Phys. Rev. Lett. 62, 961 (1989).
- 6E. Demler, W. Hanke, and S. C. Zhang, Rev. Mod. Phys. **76**, 909 $(2004).$
- 7 M. Vojta and T. Ulbricht, Phys. Rev. Lett. **93**, 127002 (2004); M. Vojta, T. Vojta, and R. K. Kaul, *ibid.* 97, 097001 (2006).
- 8G. S. Uhrig, K. P. Schmidt, and M. Gruninger, Phys. Rev. Lett. 93, 267003 (2004).
- ⁹G. Seibold and J. Lorenzana, Phys. Rev. Lett. 94, 107006 (2005).
- 10 D. X. Yao, E. W. Carlson, and D. K. Campbell, Phys. Rev. Lett. 97, 017003 (2006).
- 11B. M. Andersen and O. F. Syljuasen, Phys. Rev. B **75**, 012506 $(2007).$
- ¹²W. E. Pickett, Rev. Mod. Phys. **61**, 433 (1989).
- 13 J. T. Tranquada, arXiv:cond-mat/0512115 (unpublished).
- ¹⁴ J. R. Schrieffer, X. G. Wen, and S. C. Zhang, Phys. Rev. B **39**, 11663 (1989).
- ¹⁵ N. Bulut and D. J. Scalapino, Phys. Rev. B **53**, 5149 (1996).
- ¹⁶ Y. M. Vilk and A. M. S. Tremblay, J. Phys. I 7, 1309 (1997).
- ¹⁷ J. Brinckmann and P. A. Lee, Phys. Rev. Lett. **82**, 2915 (1999); Phys. Rev. B 65, 014502 (2001).
- 18A.-M. S. Tremblay, B. Kyung, and D. Senechal, Low Temp. Phys. 32, 424 (2006).
- ¹⁹ J. R. Schrieffer, *Theory of Superconductivity* (Benjamin, Reading, MA, 1964).
- 20H. F. Fong, B. Keimer, P. W. Anderson, D. Reznik, F. Dogan, and I. A. Aksay, Phys. Rev. Lett. **75**, 316 (1995).
- 21 L. B. Ioffe and A. J. Millis, Science 285, 1241 (1999).
- 22M. R. Norman, M. Randeria, H. Ding, and J. C. Campuzano, Phys. Rev. B 52, 615 (1995).
- 23 C. Li, S. Zhou, and Z. Wang, Phys. Rev. B 73 , 060501(R) (2006).
- ²⁴ M. R. Norman, Phys. Rev. B **61**, 14751 (2000).
- ²⁵ M. R. Norman, Phys. Rev. B **63**, 092509 (2001).
- 26A. Kaminski, H. M. Fretwell, M. R. Norman, M. Randeria, S. Rosenkranz, U. Chatterjee, J. C. Campuzano, J. Mesot, T. Sato, T. Takahashi, T. Terashima, M. Takano, K. Kadowaki, Z. Z. Li, and H. Raffy, Phys. Rev. B 71, 014517 (2005).
- 27 The dispersion being fitted corresponds to the bonding (bilayer) one.
- 28 I. Eremin, D. K. Morr, A. V. Chubukov, K. H. Bennemann, and M. R. Norman, Phys. Rev. Lett. 94, 147001 (2005).
- 29B. Fauque, Y. Sidis, L. Capogna, A. Ivanov, K. Hradil, C. Ulrich, A. I. Rykov, B. Keimer, and P. Bourges, arXiv:cond-mat/ 0701052 (unpublished).
- ³⁰ J. M. Tranquada, H. Woo, T. G. Perring, H. Goka, G. D. Gu, G. Xu, M. Fujita, and K. Yamada, Nature (London) 429, 534 $(2004).$
- 31S. M. Hayden, H. A. Mook, P. Dai, T. G. Perring, and F. Dogan, Nature (London) 429, 531 (2004).
- 32X. J. Zhou, T. Yoshida, D.-H. Lee, W. L. Yang, V. Brouet, F. Zhou, W. X. Ti, J. W. Xiong, Z. X. Zhao, T. Sasagawa, T. Kakeshita, H. Eisaki, S. Uchida, A. Fujimori, Z. Hussain, and Z.-X. Shen, Phys. Rev. Lett. **92**, 187001 (2004).
- 33A. Kaminski, S. Rosenkranz, H. M. Fretwell, M. R. Norman, M. Randeria, J. C. Campuzano, J.-M. Park, Z. Z. Li, and H. Raffy, Phys. Rev. B 73, 174511 (2006).
- ³⁴ I. Eremin, D. K. Morr, A. V. Chubukov, and K. H. Bennemann, arXiv:cond-mat/0611267 (unpublished).
- 35A. P. Schnyder, A. Bill, C. Mudry, R. Gilardi, H. M. Ronnow, and J. Mesot, Phys. Rev. B **70**, 214511 (2004).
- ³⁶ H. J. Schulz, Phys. Rev. Lett. **64**, 1445 (1990).
- 37S. Pailhes, Y. Sidis, P. Bourges, V. Hinkov, A. Ivanov, C. Ulrich, L. P. Regnault, and B. Keimer, Phys. Rev. Lett. **93**, 167001 $(2004).$
- 38X. J. Zhou, T. Yoshida, A. Lanzara, P. V. Bogdanov, S. A. Kellar, K. M. Shen, W. L. Yang, F. Ronning, T. Sasagawa, T. Kakeshita, T. Noda, H. Eisaki, S. Uchida, C. T. Lin, F. Zhou, J. W. Xiong, W. X. Ti, Z. X. Zhao, A. Fujimori, Z. Hussain, and Z.-X. Shen, Nature (London) 423, 398 (2003).
- 39U. Chatterjee, D. K. Morr, M. R. Norman, M. Randeria, A. Kanigel, M. Shi, E. Rossi, A. Kaminski, H. M. Fretwell, S. Rosenkranz, K. Kadowaki, and J. C. Campuzano, arXiv:cond-mat/ 0606346 (unpublished).