Anisotropic Kosterlitz-Thouless transition induced by hard-wall boundaries

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The spatial dependence of the superfluid density is calculated for the Kosterlitz-Thouless transition in the presence of hard-wall boundaries, for the case of a single wall bounding the half-infinite plane, and for a superfluid strip bounded by two walls. The boundaries induce additional vortices that cause the superfluid density to become anisotropic, with the tensor component perpendicular to the wall falling to zero at the wall, whereas the component parallel to the wall remains finite. The effects of the boundaries are found to extend over all measured length scales, since the correlation length is infinite in the superfluid phase.

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The behavior of a superfluid near a hard wall is still not well understood, even though many experiments are conducted in the presence of container walls. The boundary condition at the wall is that the superfluid current normal to the wall should fall to zero. There have been many guesses that this can be satisfied by having the superfluid density itself fall to zero right at the wall, but this has never been rigorously verified except perhaps in the simplest Landau-Ginsburg models.¹ Dirichlet boundary conditions at a wall are often imposed on the order parameter in perturbative renormalization expansions,² but without any real justification, and since the superfluid density is not simply proportional to the order parameter in these theories its variation near the wall is often left undetermined.

Here we calculate directly the superfluid density near a hard wall for the case of the two-dimensional Kosterlitz-Thouless (KT) transition,³ by taking into account the image vortex pairs necessary to satisfy the boundary condition. This leads to a lowered energy of the pairs in the superfluid near the wall, and a consequent increase in the density of pairs. The results show that the superfluid density does indeed fall to zero at the wall, but only for the component measured in the direction perpendicular to the wall, while the component parallel to the wall remains finite. The presence of the wall thus causes the superfluid density to become an anisotropic tensor, and this perturbation is found to extend over all length scales from the wall that are being measured. We also consider the related problem of two parallel walls bounding a long superfluid strip, a geometry often used in experimental measurements. We note that several previous studies of the KT transition near hard-wall boundaries have been carried out,⁴ but none of these have incorporated the anisotropy of the superfluid density, which we find to be a major effect.

These results should be of interest to experimental studies of the KT transition. Although superfluid ⁴He films often completely wet their substrates and thus have no boundaries, it is now possible to micromachine channels in silicon to such a small thickness that near the bulk λ transition the helium in the channels undergoes a two-dimensional KT transition,⁵ and the geometry and dimensions of the bounding walls can easily be varied. The superfluid density can be measured with a small probing flow field in a given direction at frequency ω , and the length scale *L* over which the measurement is made is then the vortex diffusion length,⁶ which is thought to vary as $\omega^{-1/2}$. For helium films at kilohertz frequencies this length is known to be 4–5 orders of magnitude larger than the vortex core size.⁷

Consider a vortex pair whose center is a distance z from a wall which bounds a semi-infinite half plane of superfluid film. The pair have a separation r and the line joining their cores makes an angle θ with respect to the perpendicular to the wall. For convenience we will scale all of these lengths in the problem in units of the vortex core radius a_0 . Since the wall can be replaced by the opposite-sign image vortices, the energy of the pair, scaled by k_BT , is easily found to be ⁸

$$\frac{U(r,\theta,z)}{k_B T} = 2\pi K_0 \left[\ln r + \frac{1}{2} \ln \left(\frac{4z^2 - r^2 \sin^2 \theta}{4z^2 + r^2 \cos^2 \theta} \right) \right] + 2E_c$$
(1)

where $K_0 = \hbar^2 \sigma_s^0 / m^2 k_B T$ is the bare areal superfluid density in dimensionless form, and E_c is the core energy of a vortex. Since the energy is a function of the angle with the wall, the distribution of the thermally excited vortex pairs will be angle dependent, and this causes the superfluid density to become anisotropic, with tensor components K_{\parallel} and K_{\perp} . By following the arguments of Machta and Guyer⁹ for the anisotropic KT transition, it can be shown that the scaling relations for the components of the superfluid density are given by

$$\frac{\partial K_{\parallel}}{\partial \ell} = -4\pi^3 K_{\parallel} \sqrt{K_{\parallel} K_{\perp}} y_{\parallel}^2, \qquad (2)$$

$$\frac{\partial K_{\perp}}{\partial \ell} = -4\pi^3 K_{\perp} \sqrt{K_{\parallel} K_{\perp} y_{\perp}^2}, \qquad (3)$$

$$\frac{\partial y}{\partial \ell} = (2 - \pi \sqrt{K_{\parallel} K_{\perp}}) y, \qquad (4)$$

$$y_{\parallel,\perp}^2 = y^2 \frac{4}{\pi} \int_{\theta_{\min}}^{\pi/2} f(\theta) \left(\frac{4z^2 - e^{2\ell} \sin^2 \theta}{4z^2 + e^{2\ell} \cos^2 \theta} \right)^{-\pi\sqrt{K_{\parallel}K_{\perp}}} d\theta, \quad (5)$$

where $\ell = \ln r$, *y* is the vortex fugacity, and $f(\theta)$ equals $\cos^2 \theta$ for the parallel moment of the fugacity y_{\parallel} and $\sin^2 \theta$ for y_{\perp} . The lower limit in the integration of Eq. (5) arises when a



FIG. 1. (Color online) Superfluid density components in the half plane (a) perpendicular and (b) parallel to the wall, and (c) the corresponding fugacities y_{\perp} (dashed curves) and y_{\parallel} (solid curves), as a function of distance *z* from the wall. The curves are labeled by their values of $T/T_{\rm KT}$.

pair near the wall is at the angle θ_{\min} where one of the pair approaches the wall to within a core radius, and from the geometry this minimum value is $\theta_{\min} = \arccos[(2z-2)/e^{\ell}]$ when $2z-2 < e^{\ell}$, and $\theta_{\min}=0$ for $2z-2 \ge e^{\ell}$. The density of pairs n(r,z) with separation between r and r+dr at the distance z is given by $n(r,z)=(y_{\parallel}^2+y_{\perp}^2)/r^2$, and in the limit of large z this reduces to the infinite-plane result y^2/r^2 .

The recursion relations are iterated using variable-step Runge-Kutta techniques, as is the integration of Eq. (5). The iterations start at the bare scale $\ell=0$ and terminate at the scale over which the superfluid density is measured, ℓ_m =ln(L). The initial values at $\ell=0$ are $K_{\parallel}^0 = K_0^0 = K_0$, and spinwave effects are neglected by assuming $\sigma_s^0 = \sigma$, the liquid density. The initial fugacities are $y_{\parallel} = y_{\perp} = y = \exp(-E_c)$, where we assume the Villain form $E_c = \pi^2 K_0/2$. The critical value



FIG. 2. (Color online) Superfluid density components (a) perpendicular and (b) parallel to the walls of a strip of width W = 1000, and (c) the corresponding fugacities y_{\perp} (dashed curves) and y_{\parallel} (solid curves), as a function of distance *z* from the left wall. The curves are labeled by their values of T/T_{KT} .

of K_0 at T_{KT} is then K_{0c} =0.747 583, and the scaled temperature is given by $T/T_{\text{KT}}=K_{0c}/K_0$.

Figure 1 shows the components of the superfluid density and fugacity as a function of the distance z from the wall for the measuring scale $\ell_m = 12$ ($L = 1.63 \times 10^5$). The superfluid density component perpendicular to the wall is zero within a core radius of the wall, and then increases smoothly to the infinite-plane value for z > L/2. The perpendicular fugacity component y_{\perp} is a maximum at the wall, reflecting the increased density of pairs parallel to the wall that drives the superfluid density to zero, and then decreases smoothly with z. For distances larger than L/2 it then decreases to the infinite-plane value. At high temperatures and very close to the wall y_{\perp} can become larger than 1, a high-density regime where the recursion relations are not valid, but over most of the regime where the superfluid density varies near the wall y_{\perp} is less than 1 and the results should be accurate. In Fig. 1 both components of the superfluid density vary rapidly near T_{KT} , but there is not a sharp jump to zero because of finitesize broadening due to the cutoff of the recursion relations at the distance *L*. This is well known¹⁰ to give a temperature broadening above T_{KT} in the form $\Delta T \approx (b/\ln L)^2$, where for our parameters the value of the constant *b* is about 1.7.

If the distance over which the perpendicular superfluid density rises to one-half of its ultimate value at large z is denoted as $\xi_{1/2}$, it can be seen in Fig. 1 that this quantity varies rapidly with temperature. Fits to $\xi_{1/2}$ over the range $T/T_{\rm KT}$ between 0.75 and 1.0 are best characterized by the form $\xi_{1/2}=0.15L \exp[13.3/(1-T/T_{\rm KT})]$. This differs from the form $\exp[b/(1-T/T_{\rm KT})^{1/2}]$ that has been suggested as a "superfluid" coherence length for the infinite-plane case,⁶ but the reason for the difference in exponents is not clear; the precise form will require an analytic solution of the recursion relations of Eqs. (2)–(5).

The component of the superfluid density parallel to the wall remains finite at the wall, but with a rapid drop over 10-40 core radii depending on the temperature. This decrease is due to the point in the recursion relations where the pair separation is nearly 2z and one of the pairs approaches to within a core radius of the wall, effectively canceling with the image in the wall. This reduces the energy of the pair considerably, resulting in an increased vortex density that lowers the superfluid density. The phase space of such pairs is quite limited, however, since larger pairs are cut off at θ_{\min} , and the resulting fugacity component y_{\parallel} is relatively small right at the wall, has a tiny peak near 10 core radii for the higher temperatures (barely visible in Fig. 1), and then increases slowly with z as the phase space increases. For intermediate values of z the parallel superfluid density is relatively constant, but then as z approaches L/2 it increases sharply as the pair energy is reduced only for very small angles θ , and finally is not reduced at all past L/2 where the superfluid density quickly approaches the infinite-plane value. y_{\parallel} peaks just at L/2 where the phase space increases rapidly in conjunction with the lowered pair energy, and then past L/2 it decreases rapidly to the infinite-plane value of y.

If these calculations are repeated at different values of L, the results are nearly identical when plotted as a function of z/L, with the only differences occurring near $T_{\rm KT}$ due to the finite-L broadening of the transition noted above. The effect of the boundary extends over all length scales that are being measured, a consequence of the infinite Kosterlitz correlation length³ in the superfluid phase below $T_{\rm KT}$. The origin of the depression of the perpendicular superfluid density at macroscopic values of z can be understood by the excitation of pairs oriented parallel to the wall ($\theta = \pi/2$), which are primarily responsible for the reduction of the perpendicular component. It is found from Eq. (1) that the energy to excite such pairs becomes nearly linear in the variable z/r when r is greater than z. Hence even at very large z it becomes quite favorable to excite pairs with much larger values of r, and



FIG. 3. (Color online) Temperature dependence of the superfluid densities (a) perpendicular and (b) parallel to the walls of a strip of width W indicated by the numbers on the curves, at a distance from the wall z=W/4. (c) shows the corresponding fugacities y_{\perp} (dashed curves) and y_{\parallel} (solid curves. The dashed portions of the curves in (a) and (b) indicate where the fugacities become larger than 1, while the upper curve is the infinite-plane KT transition.

these strongly reduce the perpendicular superfluid density even at temperatures well below $T_{\rm KT}$. It is only when the maximum value of r affecting the superfluid density is limited by the measuring length L that the effects of the wall are finally curtailed.

The case of two hard walls defining a superfluid strip of width W has also been investigated using the same techniques, with the cutoff at L effectively acting as periodic boundary conditions at the ends of the strip. The image vortices for this geometry form two infinite lines of alternating-sign vortices, and the bare energy of a pair is given by

$$\frac{U(r,\theta,z)}{k_B T} = 2\pi K_0 \left[\ln r + \frac{1}{2} \ln \left(\frac{\left[\cosh(k\beta) - \cos(k\alpha) \right] \sqrt{\left[1 - \cos k(2z + \alpha) \right] \left[1 - \cos k(2z - \alpha) \right]}}{\left[\cosh(k\beta) - \cos(2kz) \right] (k^2 r^2/2)} \right) \right] + 2E_c \tag{6}$$

where $k = \pi/W$, $\alpha = r \cos \theta$, and $\beta = r \sin \theta$. With this potential the recursion relations are identical to Eqs. (2)–(4), while in Eq. (5) the term in the parentheses of the integral is replaced by the argument of the second logarithm in Eq. (6) with r $=\exp(\ell)$. Figure 2 shows the calculated superfluid densities and fugacities for a strip of width W=1000 and measuring scale $\ell_m = 12$. Shown are the values for z between 0 and W/2; the results are symmetric about W/2. Near z=0 the characteristics are quite similar to the single-wall result, with the perpendicular superfluid density rising from zero at the wall, and the parallel component remaining finite. The main difference comes when z approaches W/2. At this point both vortices in the pair can approach the walls, greatly lowering the energy and causing a sharp dip in both components of the superfluid density even at very low temperatures. An additional effect of the second wall is that the transition for the perpendicular component is greatly broadened out, and there is also a downward shift in the perpendicular transition temperature to well below the infinite-plane $T_{\rm KT}$.

When the calculation is repeated for different values of W, the curves are quite similar when plotted versus the scaled variable z/W, with the main differences being the broadening and T_c shift of the perpendicular component. Figure 3 shows the temperature dependence of the superfluid densities and fugacities at the point z=W/4, for a wide range of W. The perpendicular superfluid density shows the strong downward shift of T_c , which appears to be a logarithmic decrease with W, varying as $\ln(W^{-0.05})$. The transition is also greatly broadened, with a linear temperature decrease that seems to be independent of W. In contrast, the parallel component shows about a 15% decrease in T_c that is independent of W, and then a W-dependent finite-size temperature broadening varying approximately as $(b/\ln W)^2$, similar to the single-wall broadening but with L now replaced by W. These main results of anisotropy extending over macroscopic length scales should also apply to superconducting films that undergo the KT transition, since strip geometries are often used in the measurements. Finite-size effects at the KT transition are known to be important in low- T_c granular films,¹¹ high- T_c films,^{12,13} and Josephson-junction arrays.¹⁴ However, the superconducting case involves an additional length scale, the penetration depth λ ,¹⁵ and to compare with experiments it will be necessary to extend the above theory to incorporate that scale. A crude first approximation might be to replace the measurement length L in the above calculations with λ if it is smaller than L.

It also may be possible to extend the calculation to the three-dimensional superfluid transition by using similar vortex renormalization techniques ^{13,16} on the vortex loops near the wall. The images in the wall would again lower the energy to create loops, but unfortunately analytic solutions for the energy are only known for loops parallel to the wall,¹⁷ and numerical techniques will be necessary for the energies at other angles.¹⁸ We would expect again to find anisotropy in the superfluid density, but since in three dimensions (3D) the correlation length is finite this would only occur within a correlation length from the wall. The extension to the 3D case would be important for gaining a more complete understanding of the critical Casimir effect in helium films, where the boundary conditions at the upper and lower film surfaces strongly affect the magnitude of the Casimir force.^{2,13}

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- ¹A. A. Sobyanin, Sov. Phys. JETP **34**, 229 (1973).
- ²V. Dohm, Phys. Scr., T **49**, 46 (1993); M. Krech and S. Dietrich, Phys. Rev. A **46**, 1886 (1992); **46**, 1922 (1992).
- ³J. M. Kosterlitz and D. Thouless, J. Phys. C **6**, 1181 (1973); J. M. Kosterlitz, *ibid.* **7**, 1046 (1974); P. Minnhagen, Rev. Mod. Phys. **59**, 1001 (1987).
- ⁴ A. Holz and Chang-de Gong, Phys. Lett. A **120**, 469 (1974); M.
 B. Sobnack and F. V. Kusmartsev, Phys. Rev. Lett. **86**, 716 (2001); A. N. Artemov, J. Low Temp. Phys. **139**, 3 (2005).
- ⁵M. Diaz-Avila, M. O. Kimball, and F. M. Gasparini, J. Low Temp. Phys. **134**, 613 (2004).
- ⁶V. Ambegaokar, B. I. Halperin, D. Nelson, and E. Siggia, Phys. Rev. B **21**, 1806 (1980).
- ⁷D. McQueeney, G. Agnolet, and J. D. Reppy, Phys. Rev. Lett. **52**, 1325 (1984).
- ⁸H. Lamb, *Hydrodynamics* (Cambridge University Press, Cambridge, U.K., 1932).

- ⁹J. Machta and R. A. Guyer, J. Low Temp. Phys. 74, 231 (1989).
- ¹⁰M. N. Barber, in *Phase Transitions and Critical Phenomena* edited by C. Domb and J. L. Lebowitz (Academic, New York, 1983), Vol. 8.
- ¹¹S. J. Turneaure, T. R. Lemberger, and J. M. Graybeal, Phys. Rev. B **63**, 174505 (2001).
- ¹²D. R. Strachan, C. J. Lobb, and R. S. Newrock, Phys. Rev. B 67, 174517 (2003).
- ¹³G. A. Williams, Phys. Rev. Lett. **92**, 197003 (2004).
- ¹⁴J. Holzer, R. S. Newrock, C. J. Lobb, T. Aouaroun, and S. T. Herbert, Phys. Rev. B 63, 184508 (2001).
- ¹⁵S. W. Pierson and O. T. Valls, Phys. Rev. B **61**, 663 (2000).
- ¹⁶S. R. Shenoy and B. Chattopadhyay, Phys. Rev. B **51**, 9129 (1995).
- ¹⁷A. L. Fetter, Phys. Rev. A 6, 402 (1972).
- ¹⁸K. R. Shariff (private communication).