## **Bulk modulus anomaly in**  $RCOO_3$  **(** $R = La$ **, Pr, and Nd)**

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In order to demonstrate the effect of hydrostatic pressure and chemical pressure on crystal structure and the spin-state transition in the perovskites  $RCOO<sub>3</sub>$  ( $R = La$ , Pr, and Nd), x-ray diffraction has been carried out under pressure up to 80 kbar. A sharp difference of the bulk modulus found between the higher-spin LaCoO<sub>3</sub> and  $\text{PrCOO}_3$  and the low-spin NdCoO<sub>3</sub> has been interpreted to reflect a pressure-induced spin-state transition in LaCoO<sub>3</sub> and PrCoO<sub>3</sub>. A change in the bandwidth of the  $\sigma$  bonding electrons due to the structural distortion has been shown to be the driving force for the spin-state transition caused by chemical pressure. On the other hand, the changes in this bandwidth must be overcompensated by the cubic-field splitting resulting from a shorter Co–O bond length in order to account for the spin-state transition under hydrostatic pressure.

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 $: 61.50$ .Ks,  $75.25.+z$ 

In the  $RCoO<sub>3</sub>$  ( $R$ =rare earth) perovskite family, the energy difference  $(\Delta_c - \Delta_{ex})$  between the cubic crystal-field splitting and the intra-atomic Hund exchange-field splitting at the octahedral-site Co(III) ions is small, i.e., comparable to *kT* at room temperature, which makes the spin state of the Co(III) ions extremely sensitive to temperature, chemical pressure, hydrostatic pressure, and subtle changes in the crystal structure. With increasing temperature,  $LaCoO<sub>3</sub>$  exhibits a progressive transition from the low-spin (LS) state Co(III):  $\hat{t}^6 e^{\hat{0}}$  to a higher-spin state  $t^{6-\delta} e^{\delta}$  ( $\delta > 0$ ), and the intermediate-spin (IS) state  $t^5e^1$  is dominant at 300 K.<sup>1–3</sup> Therefore, the observation of an unusually low bulk modulus for LaCoO<sub>3</sub> compared to that of other  $RMO<sub>3</sub>$  perovskites was interpreted to reflect a pressure-induced IS to LS transition.4 The bulk LS phase of  $LaCoO<sub>3</sub>$  was reported to be stabilized at room temperature in pressures  $P > 4$  GPa.

Substitution for  $La^{3+}$  of an  $R^{3+}$  ion of smaller ionic radius  $(IR)$  introduces a chemical pressure on the CoO<sub>3</sub> array, but a geometrical tolerance factor  $t = (R-O)/\sqrt{2(Co-O)} < 1$ , where  $(R-O)$  and  $(Co-O)$  are equilibrium bond lengths, allows cooperative  $CoO<sub>6/2</sub>$  site rotations that relieve the compressive stress on the Co–O bond. Consequently, the LS Co–O bond length changes little with IR. Nevertheless, the onset temperature for the transition from the LS state ( $\delta$ )  $= 0$ ) to a higher-spin state  $(\delta > 0)$  at the Co(III) ions increases with decreasing  $IR^{3,5}$  In this paper we argue that this apparent increase in  $(\Delta_c - \Delta_{ex})$  with decreasing IR cannot be attributed to a shorter Co–O bond length as occurs under hydrostatic pressure, so we are forced to look for an alternative explanation.

The cooperative  $MO_{6/2}$  site rotations in  $RMO_3$  perovskites bend the  $M-O-M$  bond angles from 180 $^{\circ}$  to  $(180^{\circ})$  $-2\omega$ ); this bending, which reduces the strength of the interatomic *M* –O–*M* interactions, can trigger an electronic transition as occurs in the  $RNiO<sub>3</sub>$  family.<sup>6</sup> The cooperative site rotations give rise to a sequence of structural symmetry changes with increasing IR from orthorhombic to rhombohedral to tetragonal to cubic with decrease in the bending angle ω. The tolerance factor *t* of the  $R^{3+}M^{3+}O_3$  perovskites increases under hydrostatic pressure.7 Therefore, hydrostatic pressure increases  $t$  whereas chemical pressure (smaller IR)

decreases *t*, but both pressures increase the effective  $(\Delta_c)$  $-\Delta_{ex}$ ) in the *R*CoO<sub>3</sub> family. In order to clarify this seeming paradox, we have carried out a room-temperature structural study of the  $RCoO_3$  perovskites under pressure for  $R = La$ , Pr, Nd. These selected members of the *RCoO*<sub>3</sub> family exhibit different concentrations of higher-spin Co(III) ions at room temperature, and there is a structural transition from the rhombohedral phase in  $LaCoO<sub>3</sub>$  to the orthorhombic phase in  $PrCoO<sub>3</sub>$ .

Powder samples of  $RCoO<sub>3</sub>$  were made by crushing singlecrystal ingots that had been grown in an infrared-heating image furnace. The oxygen stoichiometry of these samples was checked by measuring the thermoelectric power. Rectangular bars cut from these ingots were used previously to measure the thermal conductivity and magnetic susceptibility; $3$  these data showed the onset temperature from the LS state to a higher-spin state increases from  $\sim$ 35 K in LaCoO<sub>3</sub> to  $\sim$  200 K in PrCoO<sub>3</sub> and to  $\sim$  300 K in NdCoO<sub>3</sub>. The x-ray diffraction (XRD) under high pressure was carried out with a diamond-anvil cell;  $CaF<sub>2</sub>$  (Ref. 8) and a 4:1 mixture of methanol and ethanol were used, respectively, as the pressure manometer and the pressure medium. Application of a monocapillary collimator improved the beam intensity from a 2-kW fine-focus Mo anode tube. The XRD pattern was collected on a Fuji image plate that was scanned and digitized with a Fuji image-plate scanner BAS1800 II. The image profile was integrated to a two-column data of intensity vs  $2\theta$  with the software FIT2D, and lattice parameters were obtained by least-square fitting with the software JADE.

*Effect of pressure on the structure*. The distortion from cubic to orthorhombic symmetry in the  $RMO<sub>3</sub>$  perovskites is due to a cooperative rotation of the  $MO_{6/2}$  octahedra about the *b* axis in space group *Pbnm*, which makes  $b > a$ . However, we<sup>9</sup> have recently shown unambiguously that as the IR increases to beyond a critical value, a distortion of the  $MO_{6/2}$ sites from cubic symmetry is added to their cooperative rotation, and this site distortion inverts  $b > a$  to  $a > b$  before the perovskite transforms from orthorhombic *Pbnm* to rhombohedral  $R\overline{3}c$  symmetry; in the  $R\overline{3}c$  structure, the cooperative  $MO_{6/2}$ -site rotations are about the [111] axis. The  $RCoO_3$ family is one of the few that exhibits the *Pbnm* to  $R\overline{3}c$  cross-



FIG. 1. Pressure dependence of lattice parameters for rhombohedral  $LaCoO<sub>3</sub>$ . The hexagonal cell is used to index the rhombohedral phase. The pressure dependence of volume was fitted to the BM equation with fitting parameters labeled inside the figure.

over with increasing IR;  $LaCoO<sub>3</sub>$  has rhombohedral symmetry and a  $b > a$  in orthorhombic NdCoO<sub>3</sub> changes to  $a > b$  in orthorhombic PrCoO<sub>3</sub>. We first examine whether the  $RCoO<sub>3</sub>$ structure evolves with increasing tolerance factor under hydrostatic pressure in the same way it does under reducing chemical pressure.

As shown in Fig. 1, the rhombohedral phase of  $LaCoO<sub>3</sub>$  is stable to the highest pressures used in this work. Figure 2 shows the lattice-parameter order  $a > b$  in orthorhombic PrCoO<sub>3</sub> is retained until  $P \approx 45$  kbar, where a first-order transition to the rhombohedral phase takes place. Figure 3 shows a crossover from  $b > a$  to  $a > b$  at  $P \approx 10$  kbar in orthorhombic NdCoO<sub>3</sub> and an increase to  $P \approx 64$  kbar in the



FIG. 2. Pressure dependence of lattice parameters for orthorhombic PrCoO<sub>3</sub>. The two-phase region near  $P = 50$  kbar indicates the first-order character of the orthorhombic-rhombohedral phase transition.



FIG. 3. The same as Fig. 2 for NdCoO<sub>3</sub>.

pressure of the orthorhombic to rhombohedral transition. The structural evolution induced by hydrostatic pressure duplicates precisely what is found as the IR increases.

*Effect of pressure on the spin state*. Probing the spin-state transition on the Co(III) ions under chemical or hydrostatic pressure is not as straightforward as monitoring the structural changes since no specific structural change is associated with the spin-state transition<sup>10</sup> and this transition is progressive. However, thermal conductivity and magnetic susceptibility measurements<sup>3</sup> have shown that at room temperature the  $Co(III)$ -ion spin state is predominantly LS in NdCoO<sub>3</sub> and predominantly IS in  $LaCoO<sub>3</sub>$ ; PrCoO<sub>3</sub> has a smaller concentration of IS  $Co(III)$  than  $LaCoO<sub>3</sub>$ .

The schematic one-electron energy diagram of Fig.  $4(a)$ for the  $\pi$ -bonding *t* and  $\sigma$ -bonding *e* states in an octahedral site illustrates the subtle balance between the cubic-field splitting  $\Delta_c$  and the intra-atomic exchange splitting  $\Delta_{ex}$  at the Co(III) ions in the *R*CoO<sub>3</sub> perovskites. The difference in the effective  $(\Delta_c - \Delta_{ex})$  may be altered by the introduction of a Jahn-Teller site distortion, which stabilizes predominantly the IS state  $t^5e^1$  relative to the high-spin (HS) state  $t^4e^2$  in  $LaCoO<sub>3</sub>$  at 300 K. However, the bandwidth *W* resulting from the  $\sigma$ -bonding Co–O–Co interactions and the cubic-field splitting  $\Delta_c$  are the dominant factors to be considered in any comparison of the influences of chemical or hydrostatic pressure on the effective  $(\Delta_c - \Delta_{ex})$ , i.e., on  $(\Delta_c - W/2 - \Delta_{ex})$ . From Fig. 4(a), it is clear that broadening *W* favors stabilization of a higher-spin state; and the bandwidth is given by<sup>11,12</sup>  $W \sim \cos \omega / (C_0 - O),$ <sup>3,5</sup> where  $\omega$  is defined in the inset of Fig.  $4(a)$ . Figure  $4(b)$  shows that the  $(M-O)$  bond length varies little with IR in the perovskites  $RFeO<sub>3</sub>$  and  $RMnO<sub>3</sub>$ where there is no change in the spin state; it is the bending angle  $\omega$  that decreases systematically with increasing IR. The principle effect of chemical pressure is to buckle the  $MO_3$ array. These two examples show truly the effect of chemical pressure on the  $M$ –O bond length. In the  $RCoO<sub>3</sub>$  family, the fraction of higher-spin Co(III) at 300 K increases monotonically on going from  $R = Nd$  to Pr to La, which causes some



FIG. 4. (a) Schematic one-electron *d*-orbital energies for  $RCoO<sub>3</sub>$ , including a  $\sigma^*$  bandwidth *W* and a definition of the tilting angle  $\omega$ . (b) The tilting angle  $\omega$  and  $M$  –O bond length vs the ionic size of the  $R^{3+}$  ion (IR), which is that tabulated for nine coordination, for *RFeO<sub>3</sub>* (Refs. 21 and 22), *RMnO<sub>3</sub>* (Refs. 13 and 23), and  $RCoO<sub>3</sub>$  [ $R = La$  (Ref. 24), Pr (Ref. 25), the rest (Ref. 26)]. The Mn–O and Co–O bond lengths shown in the figure are the average value. The neutron diffraction data are available only for the  $RCoO<sub>3</sub>$  $(R = La, Pr)$ . Since they were obtained through the refinement of x-ray powder diffraction, the error bars and systematic error in the tilting angle  $\omega$  for the  $RCoO_3$  ( $R = Nd$ , Gd, Dy, and Ho) are so large that the real difference  $\Delta \omega$  between PrCoO<sub>3</sub> and NdCoO<sub>3</sub> cannot be resolved in the figure.

increase in the average (Co-O) bond length. However, we emphasize that this change in the equilibrium  $Co-O$  bond length is a consequence of the spin-state transition; it is *not* the driving force for this transition. We also point out that the bending angle  $\omega$  drops discontinuously across the transition from orthorhombic to rhombohedral symmetry.

In contrast to the lattice response to chemical pressure (smaller IR), hydrostatic pressure reduces both the  $(M-O)$ bond length and the bending angle  $\omega$  as has been shown in LaMnO<sub>3</sub>,<sup>13</sup> GdFeO<sub>3</sub>,<sup>14</sup> and PrNiO<sub>3</sub>.<sup>11</sup> Therefore, hydrostatic pressure not only increases  $\Delta_c$  by shortening the  $(M-O)$ bond length; it also broadens the bandwidth *W*, which would stabilize a higher-spin state. Therefore, we must look to experiments to determine how the system deals with this competition. An anomalously small room-temperature bulk modulus found for  $LaCoO<sub>3</sub>$  led Vogt *et al.*<sup>4</sup> to conclude that the LS state is stabilized by hydrostatic pressure in this compound; the higher-spin Co(III): *t*<sup>6−*δe*δ</sup> ions achieve a shorter equilibrium (Co-O) bond length by transferring their  $\sigma$ -bond  $e$  electrons to  $\pi$ -bonding *t* orbitals. On the other hand, an inverse pressure effect has been observed near the boundary of localized to itinerant electronic behavior in Sr-doped LaCoO<sub>3</sub>.<sup>15,16</sup> Sr doping in La<sub>1-x</sub>Sr<sub>x</sub>CoO<sub>3</sub> stabilizes a higherspin state  $t^6 \sigma^{*(1-x)}$  in which the *e* electrons occupy itinerantelectron states in a  $\sigma^*$  band of *e*-orbital parentage.<sup>17–19</sup> The transformation from localized  $-e$  to itinerant  $-\sigma^*$  states broadens significantly the bandwidth  $W$  of the  $\sigma$ -bonding  $e$ states, and this broadening overcomes any enlarged cubicfield splitting at the pressures employed.

Fitting the curve of volume versus pressure for  $LaCoO<sub>3</sub>$ , Fig. 1(b), to the Birch-Murnaghan (BM) equation

$$
P = \frac{3B_0}{2} \left[ \left( \frac{V_0}{V} \right)^{7/3} - \left( \frac{V_0}{V} \right)^{5/3} \right] \left\{ 1 + \frac{3}{4} (B' - 4) \left[ \left( \frac{V_0}{V} \right)^{2/3} - 1 \right] \right\}
$$

gives an even lower bulk modulus,  $B_o = 1220(30)$  kbar, than that reported by Vogt *et al.*<sup>4</sup> We found no clear evidence that the transition to all LS Co(III) is completed by 70 kbar. The continuous character of the transition from the higher-spin to the LS state under pressure makes the best fit to the  $V(P)$ curve with  $B' \approx 1$  in the BM equation, which deviates significantly from a value  $B'$ =4–6 typical of an elastic lattice.

 $PrCoO<sub>3</sub>$  has a smaller concentration of high-spin  $Co(III)$ at room temperature than  $LaCoO<sub>3</sub>$ . The increase in the effective  $(\Delta_c - \Delta_{ex})$  can be attributed to the larger bending of the Co–O–Co bonds in PrCoO<sub>3</sub>, which narrows the  $\sigma^*$  bandwidth  $W$ . Fitting the curve of Fig.  $2(b)$  to the BM equation with  $B' = 4$  gives  $B_0 = 1680(20)$  kbar in the orthorhombic (O) phase ( $P$ <45 kbar) and  $B_o$ =1650(40) kbar in the rhombohedral  $(R)$  phase  $(P > 55$  kbar). LaGaO<sub>3</sub> undergoes a similar  $O$ - $R$  transition under pressure; but in contrast, it has a higher  $B<sub>o</sub>$  in its  $R$  phase than in its  $O$  phase.<sup>20</sup> The bending angles  $\omega$  are reduced discontinuously on crossing from  $\mathcal O$  to  $\mathcal R$ symmetry. Therefore, the  $\sigma^*$  bandwidth *W* of PrCoO<sub>3</sub> is larger in the  $R$  phase, which decreases the effective  $(\Delta_c)$  $-\Delta_{ex}$ ) so as to increase the population of higher-spin Co(III). A greater population of higher-spin Co(III) lowers  $B<sub>o</sub>$  towards its value in  $LaCoO<sub>3</sub>$ .

NdCoO<sub>3</sub> contains few higher-spin Co(III) at room temperature, so we can expect a higher  $B<sub>o</sub>$  more in line with other  $RMO<sub>3</sub>$  perovskties as well as an increase in  $B<sub>o</sub>$  on crossing from the  $\mathcal O$  to the  $\mathcal R$  phase as in LaGaO<sub>3</sub>. Figure 3(b) shows that this expectation is indeed realized in NdCoO<sub>3</sub>. On the other hand, the increase in  $B<sub>o</sub>$  on going from isostructural  $PrCoO<sub>3</sub>$  to  $NdCoO<sub>3</sub>$  could be argued to be due to an intrinsic change caused by a larger bending angle  $\omega$ . Zhao *et al.*<sup>7</sup> have recently shown that the  $B_0$  of  $AMO_3$ perovskites depends primarily on the ratio of the compressibilities of the  $(A - O)$  and  $(M - O)$  bonds and only weakly on the elements *A* and *M*. They have further derived a relationship between this ratio, which is difficult to determine experimentally, and the ratio of bond-valence parameters that can be calculated from the crystal structure. Using their empirical relationship and calculated bond-valence parameters, we have obtained a maximum increase in  $B<sub>o</sub>$  between  $PrCoO<sub>3</sub>$  and NdCoO<sub>3</sub> due to the increased bond bending to be 3%, which is too small to account for the 15% jump in  $B_{\alpha}$ observed experimentally from Figs.  $2(b)$  and  $3(b)$ . This comparison confirms unambiguously that the unusually low values of  $B_0$  found in LaCoO<sub>3</sub> and PrCoO<sub>3</sub> are caused by the progressive pressure-induced transfer of *e* electrons to *t* orbitals in the  $Co(III)O<sub>3</sub>$  array. In conclusion, hydrostatic pressure increases the tolerance factor *t* and chemical pressure

(smaller IR) reduces it, but both stabilize the LS state relative to a higher-spin state by increasing the effective energy difference  $(\Delta_c - \Delta_{ex})$ , i.e.,  $(\Delta_c - W/2 - \Delta_{ex})$ . Chemical pressure increases the bending of the  $(180°-2\omega)$  Co–O–Co bond angle without changing significantly the LS Co(III)-O equilibrium bond length. In this case, the effective  $(\Delta_c - \Delta_{ex})$  is increased by a narrowing of the  $\sigma^*$  bandwidth *W* ~ cos  $\omega / (Co - O)^{3.5}$ . On the other hand, hydrostatic pressure decreases both the (Co–O) bond length and the bending of the Co–O–Co bond. Both effects increase *W*, but a decreased (Co–O) bond length increases  $\Delta_c$ . A bulk modulus  $B_o$  typical

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- <sup>1</sup> J. B. Goodenough, J. Chem. Phys. **6**, 287 (1958).
- 2K. Asai, A. Yoneda, O. Yokokura, J. M. Tranquada, G. Shirane, and K. Kohn, J. Phys. Soc. Jpn. 67, 290 (1998).
- <sup>3</sup> J.-Q. Yan, J.-S. Zhou, and J. B. Goodenough, Phys. Rev. B **69**, 134409 (2004).
- 4T. Vogt, J. A. Hriljac, N. C. Hyatt, and P. Woodward, Phys. Rev. B 67, 140401(R) (2003).
- <sup>5</sup> J. Baier, S. Jodlauk, M. Kriener, A. Reichl, C. Zobel, H. Kierspel, A. Freimuth, and T. Lorenz, Phys. Rev. B 71, 014443 (2005).
- <sup>6</sup> J.-S. Zhou and J. B. Goodenough, Phys. Rev. B **69**, 153105  $(2004).$
- <sup>7</sup> J. Zhao, N. L. Ross, and R. J. Angel, Acta Crystallogr., Sect. B: Struct. Sci. 60, 263 (2004).
- <sup>8</sup> R. J. Angel, J. Phys.: Condens. Matter 5, L141 (1993).
- <sup>9</sup> J.-S. Zhou and J. B. Goodenough, Phys. Rev. Lett. **94**, 065501  $(2005).$
- 10Sheng Xu and Yutaka Moritomo, J. Phys. Soc. Jpn. **70**, 3296  $(2001).$
- 11M. Medarde, J. Mesot, P. Lacorre, S. Rosenkranz, P. Fischer, and K. Gobrecht, Phys. Rev. B 52, 9248 (1995).
- 12W. A. Harrison, *The Electronic Structure and Properties of Solids* (Freeman, San Francisco, 1980).
- 13L. Pinsard-Gaudart, J. Rodriguez-Carvajal, A. Daoud-Aladine, I. Goncharenko, M. Medarde, R. I. Smith, and A. Revcolevschi,

of that for other  $RMO<sub>3</sub>$  perovskites has been found for NdCoO<sub>3</sub>, which has few higher-spin Co(III) at room temperature; but  $LaCoO<sub>3</sub>$  and  $PrCoO<sub>3</sub>$  have much smaller values of  $B<sub>o</sub>$ , and the reduction in  $B<sub>o</sub>$  increases with the population of higher-spin Co(III) at room temperature. Therefore, a pressure-induced relative stabilization of the LS state can be inferred from the compressibility. It follows that the increase in  $\Delta_c$  under hydrostatic pressure must be greater than the increase in *W*/2.

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Phys. Rev. B 64, 064426 (2001).

- 14N. L. Ross, J. Zhao, J. B. Burt, and T. D. Chaplin, J. Phys.: Condens. Matter 16, 5721 (2004).
- 15R. Lengsdorf, M. Ait-Tahar, S. S. Saxena, M. Ellerby, D. I. Khomskii, H. Micklitz, T. Lorenz, and M. M. Abd-Elmeguid, Phys. Rev. B 69, 140403(R) (2004).
- <sup>16</sup> I. Fita, R. Szymczak, R. Puzniak, I. O. Troyanchuk, J. Fink-Finowicki, Ya. M. Mukovskii, V. N. Varyukhin, and H. Szymczak, Phys. Rev. B 71, 214404 (2005).
- <sup>17</sup> J. B. Goodenough, Mater. Res. Bull. **6**, 967 (1971).
- 18M. A. Senaris-Rodriguez and J. B. Goodenough, J. Solid State Chem. 118, 323 (1995).
- 19D. Louca, J. L. Sarrao, J. D. Thompson, H. Roder, and G. H. Kwei, Phys. Rev. B 60, 10378 (1999).
- 20B. J. Kennedy, T. Vogt, C. D. Martin, J. B. Parise, and J. A. Hriljac, J. Phys.: Condens. Matter 13, L925 (2001).
- $^{21}$ M. Marezio and P. D. Dernier, Mater. Res. Bull. 6, 23 (1971).
- 22M. Marezio, J. P. Remeika, and P. D. Dernier, Acta Crystallogr. 26, 2008 (1970).
- <sup>23</sup> J. A. Alonso, M. J. Martinez-Lope, M. T. Casais, and M. T. Fernandez-Diaz, Inorg. Chem. 39, 917 (2000).
- 24P. G. Radaelli and S.-W. Cheong, Phys. Rev. B **66**, 094408  $(2002).$
- 25Y. Ren, J.-Q. Yan, J.-S. Zhou, S. Short, J. D. Jorgensen, and J. B. Goodenough (unpublished).
- <sup>26</sup>X. Liu and C. T. Prewitt, J. Phys. Chem. Solids **52**, 441 (1991).