Pressure screening in the interior of primary shells in double-wall carbon nanotubes

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The pressure response of double-wall carbon nanotubes has been investigated by means of Raman spectroscopy up to 10 GPa. The intensity of the radial breathing modes of the outer tubes decreases rapidly but remain observable up to 9 GPa, exhibiting a behavior similar (but less pronounced) to that of single-wall carbon nanotubes, which undergo a shape distortion at higher pressures. In addition, the tangential band of the external tubes broadens and decreases in amplitude. The corresponding Raman features of the internal tubes appear to be considerably less sensitive to pressure. All findings lead to the conclusion that the outer tubes act as a protection shield for the inner tubes whereas the latter increase the structural stability of the outer tubes upon pressure application.

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Carbon nanotubes have attracted intense scientific interest due to their fascinating essentially one-dimensional electronic and vibrational band structure, their unique mechanical properties, as well as the prospect for numerous applications. Raman spectroscopy has become a widespread tool for the analysis and characterization of carbon nanotubes and numerous high-pressure Raman scattering studies, on singlewall carbon nanotubes (SWCNTs) and multiwall carbon nanotubes, have made important contributions towards the understanding of the physical properties of these materials. $1-3$ A discontinuous reduction in the intensity of the low frequency radial breathing modes (RBMs) near 2 GPa has been observed in SWCNTs, accompanied, in some cases, by changes in the pressure coefficients of the tangential modes to lower values. These experimental findings have been attributed to a pressure induced hexagonal¹ or oval² distortion of the cylindrically shaped cross section of the bundled nanotubes. In addition, high-pressure x-ray diffraction (XRD) measurements together with theoretical calculations suggest a structural distortion at \sim 1.5 GPa, which is also associated with a pressure-induced nanotube polygonization.4 Raman spectroscopy at ambient pressure has been also successfully employed in the study of the more recently observed⁵ and synthesized in bulk quantities⁶ double-wall carbon nanotubes (DWCNTs), suggesting that the outer tubes provide an unperturbed environment to their interior⁷ and that the interaction in a DWCNT bundle is stronger than the inner-outer tube interaction.⁸ In this work, we study the effect of high-pressure on DWCNTs by means of Raman spectroscopy in order to investigate their structural stability and compare it with that of SWCNTs, elucidating the differences induced by the inner-outer tube interaction.

Two main processes are employed to synthesize DWCNTs. The direct process by catalytic chemical vapor deposition $(CCVD)^9$ and the peapod conversion process⁶ used for the preparation of the DWCNT material studied here. The starting raw SWCNT material was generated by the pulsed laser vaporization of a carbon rod with Ni and Co catalysts in a furnace operated at 1473 or 1523 K,¹⁰ while the peapods were prepared by a reaction of the purified uncapped SWCNTs with C_{60} vapor.^{11,12} The SWCNTs and the resulting peapods have been extensively characterized by transmission electron microscopy (TEM), XRD, and Raman measurements, revealing a purity of the SWCNTs higher than 90% with a diameter distribution from 1.25 to 1.47 nm and a filling of 85% for C_{60} in peapods.¹² At the final step of the preparation process, the peapods were converted into bundled DWCNTs by heating for 5 h at 1473 K in vacuum, following Bandow's procedure.^{6,13} Raman spectra of the DWCNTs were recorded in the backscattering geometry using a micro-Raman, triple grating system (DILOR XY) equipped with a cryogenic charge coupled device (CCD) detector. The spectral resolution of the system was \sim 2 cm⁻¹. High pressure Raman measurements were carried out using a Mao-Bell type diamond anvil cell (DAC). The 4:1 methanolethanol mixture was used as pressure transmitting medium and the ruby fluorescence technique was used for pressure calibration. For excitation, the 514.5 nm line of an Ar⁺ laser was focused on the sample by means of a $20\times$ objective, while the laser power was kept below 2.5 mW—measured directly before the cell—in order to eliminate laser-heating effects on the probed material and the concomitant softening of the observed Raman peaks. $14,15$ The phonon frequencies were obtained by fitting Lorentzian functions to the experimental peaks, whereas numerical integration after background subtraction was used for the calculation of integrated intensities of the RBM bands.

Raman spectra of the DWCNT material at room temperature and various pressures up to 10.3 GPa are illustrated in Fig. 1 along with a spectrum taken after pressure release to ambient conditions (top panel). The Raman spectrum of the starting SWCNTs at ambient pressure is also included for

FIG. 1. Raman spectra of the DWCNTs at room temperature and various pressures, recorded upon pressure increase and after total pressure release (top spectrum). The Raman spectrum of the starting SWCNT material at ambient conditions is also included for comparison (bottom spectrum). The low frequency region in all spectra and the *D* peak region in the spectra recorded at ambient pressure have been suitably enhanced in order to improve visibility. The vertical lines mark the main DWCNTs peaks followed with pressure, while the asterisk marks a band due to the methanol-ethanol mixture.

comparison (bottom panel), revealing the transformation to DWCNTs. Note that the Raman spectrum for the intermediate peapod product (not shown in Fig. 1) is almost identical to that of the SWCNTs apart from small RBM frequency shifts to lower values and the appearance of the characteristic fullerene peak attributed to the $A_g(2)$ C_{60} intramolecular mode at \sim 1465 cm⁻¹ (very weak). Two frequency regions are displayed: (i) $100-500$ cm⁻¹, containing the RBMs of the carbon nanotubes and (ii) 1300–1800 cm⁻¹, where the tangential modes of the rolled graphene sheets are located. It is well-known that in SWCNTs the frequencies of the RBMs, $\omega_{\rm RBM}$ are inversely proportional to the diameter, $d_{\rm t}$ of the tubes,¹⁶ following the general expression $\omega_{\text{RBM}}(cm^{-1})$ $=$ *A*/ d _t(nm) + *B*. For a rough estimation of the tube diameters in the DWCNT material, we have used the values *A* $=$ 234 cm⁻¹ nm and *B*=10 cm⁻¹, applied previously for $SWCNT^{17}$ and $DWCNT^{7}$ bundles.

Three main radial bands are observed at ambient conditions, each comprising of several individual Raman peaks, reflecting tubes with different chiral vector and the innerouter tube interaction.^{7,8} Two peaks located at 175 and 186 cm⁻¹ (labeled, as R_1 and R_2 in Table I) constitute the first

TABLE I. The phonon frequencies and their pressure coefficients for the well resolved Raman peaks in DWCNTs.

		Parabolic fitting		Linear fitting
	ω_i	$\partial \omega_i / \partial P$	$\partial^2 \omega_i / \partial P^2$	$\partial \omega_i / \partial P$
Mode	(cm^{-1})	(cm^{-1}/GPa)	(cm^{-1}/GPa^2)	(cm^{-1}/GPa)
R_1	175			5.8
R_2	186	7.17	-0.16	5.8
R_3	267			2.2
R_4	323			1.5
R_5	384			1.1
D_1	1320			
D_2	1348	10.71	-0.36	7.1
G_1	1503	1.51	0.06	2.1
G_2	1524	1.80	0.04	2.2
G_3	1536	1.75	0.14	3.1
G_4	1567			
G_5	1579	2.65	0.07	3.3
G_6	1592	7.50	-0.14	6.1

RBM band. By means of the above-mentioned expression, these peaks are associated with carbon tubes of relatively large diameters, in the range $1.33-1.42$ nm (outer or primary tubes). The second radial band extends from 300 to 350 cm−1 comprising of at least three peaks, with the strongest located at \sim 323 cm⁻¹ (R_4), whereas the higher energy RBM band spans the frequency range $370-400$ cm⁻¹. In the latter RBM band, four Raman peaks can be clearly resolved, with the one at \sim 384 cm⁻¹ (R_5) being intense and extremely narrow. Both bands containing the strong peaks R_4 and R_5 are attributed to the inner (secondary) nanotubes. Their frequencies suggest diameters in the range of 0.66– 0.77 and 0.61–0.64 nm, respectively. The difference in the mean diameter of the primary and secondary tubes is \sim 0.7 nm, marginally larger than the double of the turbostratic constraint of graphite at room temperature (0.344 nm) . These results are compatible with the XRD studies of a DWCNT material prepared with exactly the same method, revealing that the mean primary tube diameter is \sim 1.38 nm with an inner-outer tube separation of 0.36 nm.¹³ In addition to the main three RBM bands, two weak and broad peaks are also resolved in the low frequency region at \sim 106 and \sim 267 cm⁻¹. The former could be assigned to carbon nanotubes with a very large diameter $(\sim 2.44 \text{ nm})$, while the latter corresponds to tubes of ~ 0.91 nm in diameter.

In the high frequency region of the Raman spectrum for the DWCNTs two main bands are observed at ambient conditions. The weaker band marked by "*D*" is attributed to a disorder-induced mode, $3,18$ which also appears in graphite.¹⁹ In our DWCNT material it is comprised mainly of two peaks located at 1320 (D_1) and 1348 cm⁻¹ (D_2) compared to the single corresponding peak in SWCNTs at 1347 cm^{-1} (bottom spectrum). As the D band frequency exhibits a downshift with decreasing nanotube diameter,²⁰ the D_1 and D_2 peaks should be ascribed to secondary tubes and primary tubes,

FIG. 2. Pressure dependence of the Raman modes in DWCNTs. In the low frequency region (left panel) only the stronger and wellresolved peaks are plotted. The open (solid) symbols denote data taken for increasing (decreasing) pressure while solid lines are least square fittings.

respectively. The stronger Raman band marked by "*G*" is related to the E_{2g} mode of graphite¹⁹ and corresponds to in-plane carbon stretching vibrations in nanotubes (tangential band).^{3,16} In SWCNTs, the tangential band contains two main components resulting from the carbon displacements parallel and perpendicular to the tube axis, usually labeled as *G*⁺ and *G*[−], respectively.^{3,17} Moreover, according to theoretical calculations, 18 the tangential band is expected to show a redshift for sufficiently small nanotube diameters, providing an additional splitting of this band in SWCNTs and DWCNTs.²¹ In the DWCNT material investigated here, at least six components (G_1-G_6) are resolved at ambient pressure and their frequencies are tabulated in Table I. Based on previous Raman studies of SWCNTs and DWCNTs,^{3,21} we interpret the G_6 strong Raman peak and the G_5 shoulderlike peak as the $G⁺$ component of the carbon nanotubes, reflecting the existence of the primary and secondary tubes, respectively. The remaining G_1 - G_4 peaks are attributed to the G^- component of the various tubes having different size, keeping in mind that the lower energy *G*[−] peaks are associated with carbon nanotubes of smaller diameter.²²

Upon pressure application all the observed Raman peaks shift towards higher energies, while at the same time significant relative intensity changes take place. The pressure dependence of the most characteristic Raman lines is illustrated in Fig. 2, while their pressure coefficients (parabolic when applicable) are given in Table I. With increasing pressure, the RBM band of the outer tubes $(R_1 \text{ and } R_2)$ displays strong intensity attenuation. Above 3 GPa the R_1 shoulderlike peak

FIG. 3. (a) Integrated intensities of the R_1 , R_2 (circles, outer tubes) and R_4 (squares, large inner tubes) radial bands normalized to the higher frequency RBM band $(R₅$ peak region, small inner tubes) as a function of pressure. (b) Pressure dependence of the FWHM of the two strongest tangential modes G_5 and G_6 , corresponding to inner and outer nanotubes, respectively. The open (solid) symbols denote data taken for increasing (decreasing) pressure in both panels while solid lines are drawn to guide the eye.

cannot be resolved from R_2 , which disappears completely for pressures higher than 9 GPa. On the other hand, the RBM bands of the secondary tubes are hardly affected by the pressure, especially the R_5 peak of the small inner tubes, which remains narrow up to 10.3 GPa. This effect is quantitatively illustrated in Fig. $3(a)$, where the integrated intensities of the outer (R_1-R_2) and the larger inner tubes (R_4) RBM bands normalized to that of the smaller ones (R_5) are plotted against pressure. The relative integrated intensity of the R_4 band remains almost unaffected up to the highest pressure attained, in contrast to that of the outer tube RBM band, which decreases rapidly by an order of magnitude up to \sim 2.5 GPa. However, as already mentioned above, the R_2 peak, although weak, persists for pressures up to 9 GPa in contrast to the situation encountered in SWCNTs studies, where the RBM bands disappear above 1.5 or 1.7 GPa ^{1,2}

It becomes evident from Fig. 2, that the R_2 peak exhibits a small sublinear behavior, similar to that predicted theoretically by Venkateswaran *et al.* for the RBM band in SWCNTs under high pressure.¹ According to their model, there is no penetration of the pressure transmitting medium into the interstitial channels of the nanotube bundles and the applied pressure causes a hexagonal distortion of the tube cross section, eliminating the radial band. This description can be also adopted for the outer tubes in DWCNTs, although in this case the distortion of the outer tubes is expected to be smaller (the $R₂$ band persists with pressure), possibly due to the presence of the inner tubes. This assumption is further supported by the smaller pressure slopes of the R_1 and the R_2 peaks (primary tubes) in comparison to those reported in the literature for the RBM band of SWCNTs and compiled in Ref. 3. Moreover, the pressure coefficient for the Raman peaks associated with the outer tubes is much larger than those corresponding to the inner ones (Fig. 2, Table I). The overall behavior of the RBM bands under pressure indicates that the outer carbon nanotubes are, by far, more vulnerable to pressure application than the inner tubes in line with the proposed stronger inter-DWCNT interaction than that between inner and outer tube at ambient conditions.⁸ It seems logical to suggest that the existence of the primary tubes results in a screening of the applied pressure on the secondary tubes, while the latter provide structural support against pressure induced deformation of the outer tubes. This assumption is consistent with recent results of the high pressure Raman study of C_{60} peapods by Rafailov *et al.* where a pressure shielding of C_{60} inside the tubes is also observed.²³ Finally, the puzzling pressure response of the R_3 radial peak, corresponding to tubes of an intermediate diameter, must be noted. Namely, the peak intensity decreases with increasing pressure and its pressure evolution cannot be followed beyond 3 GPa, in close analogy to what is observed for the RBM band of the primary tubes. However, the very small pressure coefficient of the R_3 peak—comparable to those of the secondary tubes—prevents an unambiguous assignment.

The pressure response of the tangential band is also of great importance, further supporting the above considerations. The G_6 peak assigned to the G^+ band of the outer tubes exhibits a much larger pressure coefficient than that of $G₅$ attributed to the inner tubes, in agreement with our proposed assignment and the pressure screening effect inside the primary tubes. The different pressure coefficients result in a more clear separation of the two peaks at elevated pressures $(Fig. 1)$. At the same time, a significant broadening and amplitude drop of the G_6 peak take place with pressure. In Fig. 3(b), the full width at half maximum (FWHM) of the G_5 and *G*⁶ peaks is plotted as a function of pressure. It is evident that G_6 broadens much faster than G_5 peak, reflecting again the larger deformation of the outer tubes and the pressure screening for the inner ones. Another noticeable point is the sublinear pressure dependence of the G_6 peak position (similar to that of the corresponding radial band, R_2) in contrast to the superlinear behavior of G_5 . This can be understood by assuming that with increasing pressure the inner-outer tube interaction becomes progressively stronger supporting the primary tubes (reduced slope), while at the same time the secondary tubes are increasingly affected by pressure (increased slope). The G_4 shoulder, assigned to the G^- component of the larger primary tubes, shifts swiftly with pressure (like the G^+ component of these tubes) and merges with the G_5 and G_6 peaks above \sim 1 GPa. The rest of the *G* peaks, associated with smaller nanotubes, display pressure coefficients considerably smaller than that of the G_6 peak and comparable with that of $G₅$. Their superlinear trend with pressure further supports their assignment to the secondary tubes.

The pressure dependence of the *D* band could not be followed at low pressures due to the overlap with the strong Raman signal of the diamond in the DAC around 1332 cm^{-1} . Only above 6 GPa, a weak and broad peak (D_2) appears in the measured spectral window. Its pressure behavior appears again to be sublinear in agreement with our tentative assignment of this peak to the outer shells. As peaks associated with inner tubes have much smaller slopes, the D_1 peak does not appear in our spectral window up to 10.3 GPa. Note that the broad and weak band observed at \sim 1455 cm⁻¹ inside the DAC (asterisk in Fig. 1), is absent in spectra taken outside the cell. This peak is attributed to the hydrostatic pressure medium of methanol-ethanol.²⁴

Although the pressure-induced shifts of the Raman peaks in DWCNTs are fully reversible, this is not the case for the relative intensities of certain bands. Namely, the integrated intensity of the R_1 - R_2 band and the amplitude of the G_6 peak do not fully recover after total pressure release. Moreover, the intensity of the *D* bands after pressure release remains somewhat larger to that initially recorded at ambient conditions. These divergences suggest the existence of residual pressure-induced deformations of the primary tubes, in analogy to those observed in SWCNTs.¹

Summarizing, our high pressure Raman study on the DWCNTs show that the application of pressure initially causes the deformation of the primary tubes, which actually shield the inner tubes against pressure. At higher pressure, the increased interaction between outer and inner shells acts as to provide structural support against the deformation of the outer tubes. During the course of this work, a high pressure Raman study (excitation with 633 nm) of a DWCNT material prepared by the CCVD method has been published.²⁵ The results of this study, focused on the pressure response of the two high frequency tangential peaks attributed to the inner and outer tubes, are in good agreement with the frequency shifts and the linewidths pressure evolution of the G_5 and G_6 peaks presented here.

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10 928 (1999).

- ²M. J. Peters, L. E. McNeil, J. P. Lu, and D. Kahn, Phys. Rev. B **61**, 5939 (2000).
- ¹U. D. Venkateswaran, A. M. Rao, E. Richter, M. Menon, A. Rinzler, R. E. Smalley, and P. C. Eklund, Phys. Rev. B **59**,
- 3 I. Loa, J. Raman Spectrosc. **34**, 611 (2003).
- ⁴ J. Tang, L. C. Qin, T. Sasaki, M. Yudasaka, A. Matsushita, and S. Iijima, Phys. Rev. Lett. **85**, 1887 (2000).
- ⁵B. W. Smith, M. Monthioux, and D. E. Luzzi, Nature (London) **396**, 323 (1998).
- 6S. Bandow, M. Takizawa, K. Hirahara, M. Yudasaka, and S. Iijima, Chem. Phys. Lett. **337**, 48 (2001).
- 7R. Pfeiffer, H. Kuzmany, Ch. Kramberger, Ch. Schaman, T. Pichler, H. Kataura, Y. Achiba, J. Kurti, and V. Zolyomi, Phys. Rev. Lett. **90**, 225501 (2003).
- 8R. Pfeiffer, Ch. Kramberger, F. Simon, H. Kuzmany, V. N. Popov, and H. Kataura, cond-mat/0406670 (unpublished).
- 9E. Flahaut, R. Bacsa, A. Peigney, and C. Laurent, Chem. Commun. (Cambridge) 12, 1442 (2003).
- 10H. Kataura, Y. Kumazawa, Y. Ohtsuka, S. Suzuki, Y. Maniwa, and Y. Achiba, Synth. Met. 103, 2555 (1999).
- 11H. Kataura, Y. Maniwa, T. Kodama, K. Kikuchi, H. Hirahara, K. Suenaga, S. Iijima, S. Suzuki, Y. Achiba, and W. Kraetschmer, Synth. Met. 121, 1195 (2001).
- 12H. Kataura, Y. Maniwa, M. Abe, A. Fujiwara, T. Kodama, K. Kikuchi, H. Imahori, Y. Misaki, S. Suzuki, and Y. Achiba, Appl. Phys. A: Mater. Sci. Process. A74, 349 (2002).
- 13M. Abe, H. Kataura, H. Kira, T. Kodama, S. Suzuki, Y. Achiba, K. Kato, M. Takata, A. Fujiwara, K. Matsuda, and Y. Maniwa, Phys. Rev. B 68 , 041405(R) (2003).
- 14P. V. Huong, R. Cavagnat, P. M. Ajayan, and O. Stephan, Phys. Rev. B 51, 10 048 (1995).
- ¹⁵ Y. Ando, X. Zhao, and H. Shimoyama, Carbon **39**, 569 (2001).
- 16A. M. Rao, E. Richter, S. Bandow, B. Chase, P. C. Eklund, K. A. Williams, S. Fang, K. R. Subbaswamy, M. Menon, A. Thess, R. E. Smalley, G. Dresselhaus, and M. S. Dresselhaus, Science **275**, 187 (1997).
- 17A. Jorio, M. A. Pimenta, A. G. Souza Filho, R. Saito, G. Dresselhaus, and M. S. Dresselhaus, New J. Phys. 5, 139 (2003).
- 18W. S. Basca, D. Ugarte, A. Chatelain, and W. A. de Heer, Phys. Rev. B 50, 15 473 (1994).
- ¹⁹R. J. Nemanich and S. A. Solin, Phys. Rev. B **20**, 392 (1979).
- 20A. G. Souza Filho, A. Jorio, Ge. G. Samsonidze, G. Dresselhaus, M. A. Pimenta, M. S. Dresselhaus, A. K. Swan, M. S. Unlu, B. B. Goldberg, and R. Saito, Phys. Rev. B **67**, 035427 (2003).
- 21S. Bandow, G. Chen, G. U. Sumanasekera, R. Gupta, M. Yudasaka, S. Iijima, and P. C. Eklund, Phys. Rev. B **66**, 075416 (2002) .
- 22A. Jorio, A. G. Souza Filho, G. Dresselhaus, M. S. Dresselhaus, A. K. Swan, M. S. Unlu, B. B. Goldberg, M. A. Pimenta, J. H. Hafner, C. M. Lieber, and R. Saito, Phys. Rev. B **65**, 155412 $(2002).$
- 23P. M. Rafailov, C. Thomsen, and H. Kataura, Phys. Rev. B **68**, 193411 (2003).
- 24X. B. Wang, Z. X. Shen, S. H. Tang, and M. H. Kuok, J. Appl. Phys. 85, 8011 (1999).
- 25P. Puech, H. Hubel, D. J. Dunstan, R. R. Bacsa, C. Laurent, and W. S. Bacsa, Phys. Rev. Lett. 93, 095506 (2004).