## Erratum: Evidence for a low-spin to intermediate-spin state transition in LaCoO<sub>3</sub> [Phys. Rev. B 66, 020402 (R)(2002)]

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The expression given for the scaling factor C between thermal expansion  $\alpha$  and susceptibility  $\chi$  in Eq. (4) is erroneous. The correct expression is

$$C = \frac{N_A g^2 \mu_B^2}{3k_B} \frac{S(S+1)}{d}.$$

Thus *C* does not depend explicitly on the orbital degeneracy  $\nu$ . Consequently, the values for the scenarios with  $\nu$ =3 have to be multiplied by 3. Moreover, a more detailed analysis of the susceptibility data (see Ref. 1) yields a somewhat smaller background susceptibility, which leads to a larger *g* factor of *g*=2.28 instead of *g*=2.13 in the respective fit. This causes a 15% increase of the scaling factors of all four scenarios and Table I should therefore read as noted below.

The smaller background susceptibility causes a similar increase for the experimental value  $C_{exp}$  of the scaling factor to  $C_{exp}=195 \text{ emuK/mole}$ . Therefore, our conclusion is unaffected because the very good agreement between the experimental result and the expected value for the LS/IS scenario with  $\nu=1$  remains unchanged and the deviation from the other scenarios is as large or even larger than before.

TABLE I. Parameters d and  $\Delta$  of the fits of the anomalous thermal expansion  $\Delta \alpha$  of LaCoO<sub>3</sub> (see Fig. 2) obtained for a LS/IS and for a LS/HS scenario with ( $\nu$ =3) and without ( $\nu$ =1) orbital degeneracy of the excited IS (HS) state. The respective scaling factors C of Eq. (4) are given in the last row. Experimentally we find  $C^{\exp}$ =195 emuK/mole.

|               | LS/IS: S=1 |             | LS/HS: S=2 |             |
|---------------|------------|-------------|------------|-------------|
|               | $\nu = 1$  | <i>v</i> =3 | $\nu = 1$  | <i>v</i> =3 |
| d (%)         | 0.66       | 0.44        | 0.55       | 0.38        |
| $\Delta$ (K)  | 185        | 265         | 205        | 256         |
| C (emuK/mole) | 190        | 290         | 690        | 1000        |

<sup>1</sup>J. Baier, S. Jodlauk, M. Kriener, A. Reichl, C. Zobel, H. Kierspel, A. Freimuth, and T. Lorenz, Phys. Rev. B 71, 014443 (2005).